Theoretical review on the prospect for new physics in charm sector

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Weizmann Institute



The 4th Int. Workshop on Charm Physics:

Charm 2010

Friday, October 22, 2010

Outline

Brief introduction, importance of up flavor physics.

Model independent information (effective field theory).

Charm physics & alignment models.

• $D - \overline{D}$ mixing, the connection with tFCNC.

Model dependent information: (MFV, SUSY, RS).

The potential of
$$D^0 \to \mu^+ \mu^-$$
.

Conclusions.

Friday, October 22, 2010

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Absence (?) of deviation from SM predictions implies severe bound on new physics (NP).

• Most of precise information involves K, B mesons, linked to down type FCNC.

Output the severe hierarchy problem is induced by the top sector, which is indeed extended in most of natural NP models.



What do we know about New Phys. flavor sector, model independently?



Generic bounds via effective theory

- $\Delta F = 2$ processes among the cleanest.
- In the SM proceed at loop and highly suppressed.
- To leading order beyond the SM:

$$\frac{\left(\bar{q}_i q_j\right)\left(\bar{q}_i q_j\right)}{\Lambda_{\rm NP}^2}$$

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What are the bounds on $\Lambda_{\rm NP}$

for different flavor transitions?

Isidori, Nir & GP, Ann. Rev. Nucl. Part. Sci. (10)

Operator	Bounds on	$\Lambda \text{ in TeV } (c_{ij} = 1)$	Bounds on a	$c_{ij} \ (\Lambda = 1 \text{ TeV})$	Observables
	Re	Im	Re	Im	
$(\bar{s}_L \gamma^\mu d_L)^2$	9.8×10^2	1.6×10^4	9.0×10^{-7}	3.4×10^{-9}	$\Delta m_K; \epsilon_K$
$(\bar{s}_R d_L)(\bar{s}_L d_R)$	1.8×10^4	$3.2 imes 10^5$	6.9×10^{-9}	2.6×10^{-11}	$\Delta m_K; \epsilon_K$
$(\bar{c}_L \gamma^\mu u_L)^2$	1.2×10^3	2.9×10^3	5.6×10^{-7}	1.0×10^{-7}	$\Delta m_D; q/p , \phi_D$
$(\bar{c}_R u_L)(\bar{c}_L u_R)$	6.2×10^3	1.5×10^4	5.7×10^{-8}	1.1×10^{-8}	$\Delta m_D; q/p , \phi_D$
$(ar{b}_L \gamma^\mu d_L)^2$	5.1×10^2	$9.3 imes 10^2$	3.3×10^{-6}	$1.0 imes 10^{-6}$	$\Delta m_{B_d}; S_{\psi K_S}$
$(\bar{b}_R d_L)(\bar{b}_L d_R)$	1.9×10^3	3.6×10^3	5.6×10^{-7}	1.7×10^{-7}	$\Delta m_{B_d}; S_{\psi K_S}$
$(ar{b}_L\gamma^\mu s_L)^2$	1	$.1 \times 10^2$	7.6	$\times 10^{-5}$	Δm_{B_s}
$(ar{b}_Rs_L)(ar{b}_L s_R)$	3	0.7×10^2	1.3	$\times 10^{-5}$	Δm_{B_s}
$(\bar{t}_L \gamma^\mu u_L)^2$					

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<u> </u>		Re	Im	Re	Im	
X	$(\bar{s}_L \gamma^\mu d_L)^2$	9.8×10^2	$1.6 imes 10^4$	9.0×10^{-7}	3.4×10^{-9}	$\Delta m_K; \epsilon_K$
	$(ar{s}_R d_L)(ar{s}_L d_R)$	1.8×10^4	$3.2 imes 10^5$	6.9×10^{-9}	2.6×10^{-11}	$\Delta m_K; \epsilon_K$
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	$(\bar{b}_L \gamma^\mu d_L)^2$	$5.1 imes 10^2$	$9.3 imes 10^2$	3.3×10^{-6}	1.0×10^{-6}	$\Delta m_{B_d}; S_{\psi K_S}$
($(ar{b}_Rd_L)(ar{b}_Ld_R)$	1.9×10^3	$3.6 imes 10^3$	5.6×10^{-7}	1.7×10^{-7}	$\Delta m_{B_d}; S_{\psi K_S}$
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SS	$(ar{b}_R d_L)(ar{b}_L d_R)$	$1.9 imes 10^3$	$3.6 imes 10^3$	5.6×10^{-7}	1.7×10^{-7}	$\Delta m_{B_d}; S_{\psi K_S}$
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D-system falls only behind the K-one

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Friday, October 22, 2010

Huge recent progress in measurement of mass splitting & CP violation (CPV) in the *D* system:

System parameters roughly determined (HFAG):

$$x = (1.00 \pm 0.25) \times 10^{-2},$$

$$y = (0.77 \pm 0.18) \times 10^{-2},$$

$$1 - |q/p| = +0.06 \pm 0.14,$$

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Absence of *D* CPV
a SM victory!

SM: D system is controlled by 2 gen' physics \Rightarrow CP conserving

> Bottom contribution is down by: $\mathcal{O}\left(\frac{m_c^2}{m_b^2} \times \frac{V_{ub}V_{cb}^*}{V_{us}V_{cs}^*}\right) = 10^{-4} \text{ (see talk by Lenz)}$

The power of CPV in the D system



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$$\begin{split} y_{12} &\equiv |\Gamma_{12}|/\Gamma, \qquad x_{12} \equiv 2|M_{12}|/\Gamma, \qquad \phi_{12} \equiv \arg(M_{12}/\Gamma_{12}). \\ x_{12}^{\text{NP}} &\lesssim x_{12}^{\text{exp}} \sim 0.012, \qquad x_{12}^{\text{NP}} \sin \phi_{12}^{\text{NP}} \lesssim x_{12}^{\text{exp}} \sin \phi_{12}^{\text{exp}} \sim 0.0022, \end{split}$$

If x is due to NP then it missed a chance to revealed itself in $\mathcal{O}(1)$ CPV. $|x_{12}^{NP}/x|$ Gedalia, et. al (09).



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What do we conclude ?

Physics



Sort and list Nobel Prizes and Nobel Laureat

What do we conclude ?



Resulting bounds are too strong to allow for generic TeV-scale

NP - tension with solving the fine tuning problem.

http://nobelprize.org/nobel_prizes/physics/laureates/2008/

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Resulting bounds are too strong to allow for generic TeV-scale

NP - tension with solving the fine tuning problem.

Hint for underlying structure of microscopic laws of nature.

http://nobelprize.org/nobel_prizes/physics/laureates/2008/

Page 1 of 1

What kind of NP survives?

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uFCNC data, a crucial test of alignment

Down type flavor violation can be shut off via alignment, where anarchic NP is diagonal in the down mass basis.



Yasmin & Gilad Perez <jasgilperez@gmail.com>

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careful domino allignment to reference your reservation. **Priority Club Rewards:** Your Priority Club Rewards number applies to this reservation.

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MR GILAD PEREZ

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What if down alignment is at work ?



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	Re	Im	Re	Im	
$(s_L\gamma, a_L)$	102×102	1.6×10^{4}	9.0×10^{-7}	$3/1 \times 10^{-9}$	$_{K}, _{K}$
$(\overline{L}_{L})(\overline{L}_{W}K)$	1.0×10^{11}	$3.2 imes 10^5$	6.9×10^{-9}	2.6 imes 10	
$(ar{c}_L \gamma^\mu u_L)^2$	1.2×10^3	$2.9 imes 10^3$	5.6×10^{-7}	1.0×10^{-7}	$\Delta m_D; q/p , \phi_D$
$(\bar{c}_R u_L)(\bar{c}_L u_R)$	6.2×10^3	$1.5 imes 10^4$	5.7×10^{-8}	1.1×10^{-8}	$\Delta m_D; q/p , \phi_D$
	5.1×10^{2}	$9.3 imes 10^2$	3.3×10^{-6}	1.0×10^{-6}	$\Delta_d, \mathcal{S}_{\psi}K_S$
$(\overline{b}, d)(\overline{b}, \overline{b})$	1.0 / 10	$3.6 imes 10^3$	5.6×10^{-1}	1.7×10^{-7}	
	1.	1×10^2	7.6	$\times 10^{-5}$	
$(\overline{h}, \overline{c})(\overline{L}, \overline{c})$	J.	7×10^{2}	1.3	X 10	Ame
$(\bar{t}_L \gamma^\mu u_L)^2$					

What if down alignment is at work ?



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	Re	Im	Re	Im	
$(s_L\gamma, a_L)$	2.0×10^{2}	1.6×10^{4}	9.0×10^{-7}	3.4×10^{-9}	$\underline{-}$
$\left(\frac{1}{2} \right) \left(\frac{1}{2} \right) $	1.0 × 101	$3.2 imes 10^5$	6.9×10^{-9}	2.6 imes 10	
$(ar{c}_L \gamma^\mu u_L)^2$	1.2×10^3	$2.9 imes 10^3$	5.6×10^{-7}	1.0×10^{-7}	$\Delta m_D; q/p , \phi_D$
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$(\overline{h}, d)(\overline{l}, h)$	1.0 / 10	3.6×10^{3}	5.6×10^{-1}	1.1 × 10	
	1.	1×10^{2}	7.6	$\times 10^{-5}$	•••• <i>D</i> _s
$(\overline{b} - \alpha)(\overline{b} - n)$	J.	7×10^{2}	1.3	X 10	Amp
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u-FCNC data remove immunities!

Friday, October 22, 2010

Up sector







Friday, October 22, 2010

The power of CPV in D mixing & how it kills alignment models





Wednesday, October 20, 2010
2-gen' effective theory for $\Delta F = 2$

Robust model independent bounds:

(*i*) robust (*ii*) LLRR - stronger, but model dependent.

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 $\frac{1}{\Lambda_{\rm NP}^2} \left[z_1^K (\overline{d_L} \gamma_\mu s_L) (\overline{d_L} \gamma^\mu s_L) + z_1^D (\overline{u_L} \gamma_\mu c_L) (\overline{u_L} \gamma^\mu c_L) + z_4^D (\overline{u_L} c_R) (\overline{u_R} c_L) \right].$

[More info' in Δc =1, Golowich, et. al (09), Kagan & Sokolof (09)]

2-gen' effective theory for $\Delta F = 2$

Robust model independent bounds:



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$$\frac{1}{\Lambda_{\rm NP}^2} (\overline{Q_{Li}}(X_Q)_{ij} \gamma_\mu Q_{Lj}) (\overline{Q_{Li}}(X_Q)_{ij} \gamma^\mu Q_{Lj}),$$

One cannot eliminate the constraint from K & D systems simultaneously!

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One cannot eliminate the constraint from K & D systems

simultaneously!



When effects of $SU(2)_L$ breaking are small, the terms that lead to z_1^K and z_1^D have the form

$$\frac{1}{\Lambda_{\rm NP}^2} (\overline{Q_{Li}}(X_Q)_{ij} \gamma_\mu Q_{Lj}) (\overline{Q_{Li}}(X_Q)_{ij} \gamma^\mu Q_{Lj}),$$

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One cannot eliminate the constraint from K & D systems



Implications of CPV in $D^0 - \bar{D}^0$ mixing

(i) Model independent;
(ii) General minimal flavor violation (GMFV);
(iii) SUSY;
(iv) Randall-Sundrum (RS).

Ciuchini, et al. (07); Csaki, et al. (08); Kagan, et al. (09); Gedalia, et al. (09,10,10); Blum, et al. (09); Buras et.al.; Csaki, et al. (09); Bauer, et al. (09); Bigi, et al. (09); Altmannshofer, et al. (09,10); Blanke, et al. (09); Crivellin & Davidkov (10).

$$\begin{array}{c|c} & \hat{v}_1 \\ & \Delta m_D \\ & Y_u Y_u^{\dagger} \\ \hline & 2\theta_C \\ 2\theta_C \\ 2\theta_d \\ \hline & Y_d Y_d^{\dagger} \\ & \Delta m_K \\ & \hat{v}_3 \end{array}$$

$$L = |X_Q| = \left(X_Q^2 - X_Q^1\right)/2$$

Constraining the eigenvalue difference of flavor violation source, indep' of it's direction!







CPV in D: Model Dependent Implications



(i) MFV (exciting #1); (ii) SUSY; (iii) Randall-Sundrum (RS).

General MFV (GMFV) vs. Linear MFV (LMFV)

Kagan, GP, Volanksy & Zupan, PRD (09); Gedalia, Grossman, Nir & GP, PRD (09).

Comparable NP contributions from strange & bottom (unlike SM)

$$r_{sb} \equiv \frac{y_s^2}{y_b^2} \left| \frac{V_{us}^{\text{CKM}} V_{cs}^{\text{CKM}}}{V_{ub}^{\text{CKM}} V_{cb}^{\text{CKM}}} \right| \sim 0.5,$$

$$C_1^{cu} \propto \left[y_s^2 \left(V_{cs}^{\text{CKM}} \right)^* V_{us}^{\text{CKM}} + \left(1 + r_{\text{GMFV}} \right) y_b^2 \left(V_{cb}^{\text{CKM}} \right)^* V_{ub}^{\text{CKM}} \right]^2$$

$$r_{\text{GMFV} result of resummation \sum_n y_b^n}$$

General MFV (GMFV) vs. Linear MFV (LMFV)

Kagan, GP, Volanksy & Zupan, PRD (09); Gedalia, Grossman, Nir & GP, PRD (09).

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Comparable NP contributions from strange & bottom (unlike SM)

 $C_1^{cu} \propto \left[y_s^2 \left(V_{cs}^{\text{CKM}} \right)^* V_{us}^{\text{CKM}} + \left(1 + \left(\int_{\text{GMFV}} y_b^2 \left(V_{cb}^{\text{CKM}} \right)^* V_{ub}^{\text{CKM}} \right)^2 \right]^2$ $|x_{12}^{NP}/\mathbf{x}|$ $r_{\rm GMFV}$ result of Determining what "phase" describes nature yield microscopic info'. Well beyond the LHC reach! Improvement via BESIII threshold measurements; Within the reach of LHCb & maybe Tevatron; Looking forward for Exp' talks...

 $r_{sb} \equiv \frac{y_s^2}{y_b^2} \left| \frac{V_{us}^{\rm CKM} V_{cs}^{\rm CKM}}{V_{vb}^{\rm CKM} V_{cb}^{\rm CKM}} \right| \sim 0.5 \,,$

SUSY+RS

SUSY (doom of alignment)

Gedalia, et. al (09).

Robust

 $\frac{m_{\tilde{Q}_2} - m_{\tilde{Q}_1}}{m_{\tilde{Q}_2} + m_{\tilde{Q}_1}} \le \begin{cases} 0.034 & \text{maximal phases} \\ 0.27 & \text{vanishing phases} \end{cases}$

$$\frac{m_{\tilde{u}_2} - m_{\tilde{u}_1}}{m_{\tilde{u}_2} + m_{\tilde{u}_1}} \lesssim 0.02 - 0.04.$$

Generic

squark doublets, 1TeV;

average of the doublet & singlet mass splitting.

RS (constraining alignment)

Csaki, Falkowski & Weiler, PRD (09); Gedalia, et. al (09).

Robust

$$m_{\rm KK} > 2.1 f_{Q_3}^2 \,{\rm TeV} \,,$$

 f_{Q_3} is typically in the range of 0.4- $\sqrt{2}$.

$$m_{\mathrm{KK}} > rac{4.9\,(2.4)}{y_{5D}}\,\mathrm{TeV}$$
 IR (bulk) Higgs

Generic

 $\frac{1}{2} \lesssim y_{5D} \lesssim \frac{2\pi}{N_{KK}}$ for brane Higgs; $\frac{1}{2} \lesssim y_{5D} \lesssim \frac{4\pi}{\sqrt{N_{KK}}}$ for bulk Higgs,

charming top Phys. @ the LHC



 $\Delta F = 2, \left[(\bar{t}, \bar{b})_L X_Q(u, d)_L \right]^2$

Gedalia, Mannelli & GP, PLB; JHEP (10).

\blacklozenge Signal is in same sign tops: ~uu ightarrow tt

Operator	Bounds on Λ in TeV $(c_{ij} = 1)$		Bounds on c_{ij} ($\Lambda = 1$ TeV)		Observables
	Re	Im	Re	Im	
$(\bar{s}_L \gamma^\mu d_L)^2$	9.8×10^2	1.6×10^4	$9.0 imes 10^{-7}$	3.4×10^{-9}	$\Delta m_K; \epsilon_K$
$(\bar{s}_R d_L)(\bar{s}_L d_R)$	1.8×10^4	$3.2 imes 10^5$	$6.9 imes 10^{-9}$	2.6×10^{-11}	$\Delta m_K; \epsilon_K$
$(\bar{c}_L \gamma^\mu u_L)^2$	1.2×10^3	2.9×10^3	$5.6 imes10^{-7}$	$1.0 imes 10^{-7}$	$\Delta m_D; q/p , \phi_D$
$(\bar{c}_R u_L)(\bar{c}_L u_R)$	6.2×10^3	1.5×10^4	$5.7 imes 10^{-8}$	1.1×10^{-8}	$\Delta m_D; q/p , \phi_D$
$(\bar{b}_L \gamma^\mu d_L)^2$	$5.1 imes 10^2$	$9.3 imes10^2$	$3.3 imes 10^{-6}$	$1.0 imes 10^{-6}$	$\Delta m_{B_d}; S_{\psi K_S}$
$(\bar{b}_R d_L)(\bar{b}_L d_R)$	1.9×10^3	$3.6 imes10^3$	$5.6 imes10^{-7}$	$1.7 imes 10^{-7}$	$\Delta m_{B_d}; S_{\psi K_S}$
$(\bar{b}_L \gamma^\mu s_L)^2$	$1.1 imes 10^2$		$7.6 imes 10^{-5}$		Δm_{B_s}
$(\bar{b}_Rs_L)(\bar{b}_Ls_R)$	$3.7 imes 10^2$		$1.3 imes 10^{-5}$		Δm_{B_s}
$(\bar{t}_L \gamma^\mu u_L)^2$?		?		?

 $\Delta F = 2, \left[(\bar{t}, \bar{b})_L X_Q(u, d)_L \right]^2$

Projected LHC bound, same sign tops.

Operator	Bounds on Λ in TeV $(c_{ij} = 1)$		Bounds on c_{ij} ($\Lambda = 1$ TeV)		Observables
	Re	Im	Re	Im	
$(\bar{s}_L \gamma^\mu d_L)^2$	9.8×10^2	$1.6 imes 10^4$	9.0×10^{-7}	3.4×10^{-9}	$\Delta m_K; \epsilon_K$
$(\bar{s}_R d_L)(\bar{s}_L d_R)$	1.8×10^4	3.2×10^5	6.9×10^{-9}	2.6×10^{-11}	$\Delta m_K; \epsilon_K$
$(\bar{c}_L \gamma^\mu u_L)^2$	1.2×10^3	2.9×10^3	5.6×10^{-7}	1.0×10^{-7}	$\Delta m_D; q/p , \phi_D$
$(\bar{c}_R u_L)(\bar{c}_L u_R)$	$6.2 imes 10^3$	$1.5 imes 10^4$	5.7×10^{-8}	1.1×10^{-8}	$\Delta m_D; q/p , \phi_D$
$(ar{b}_L \gamma^\mu d_L)^2$	5.1×10^2	$9.3 imes 10^2$	3.3×10^{-6}	1.0×10^{-6}	$\Delta m_{B_d}; S_{\psi K_S}$
$(ar{b}_Rd_L)(ar{b}_L d_R)$	1.9×10^3	$3.6 imes 10^3$	5.6×10^{-7}	1.7×10^{-7}	$\Delta m_{B_d}; S_{\psi K_S}$
$(ar{b}_L \gamma^\mu s_L)^2$	1.1×10^2		7.6×10^{-5}		Δm_{B_s}
$(ar{b}_Rs_L)(ar{b}_L s_R)$	$3.7 imes 10^2$		1.3×10^{-5}		Δm_{B_s}
$(\bar{t}_L \gamma^\mu u_L)^2$	12		7.1×10^{-3}		$uu \rightarrow tt$

What gives strongest bound on $[(\bar{t}, \bar{b})_L X_Q(u, d)_L]^2$? Despite $\mathcal{O}(\lambda_C^5)$ suppression: Define: $L \equiv |X_Q^{\Delta F=2}|$

 $uu \rightarrow tt \text{ (LHC projected):} \ L < 12\left(\frac{\Lambda_{\text{NP}}}{1 \text{ TeV}}\right); \ \Lambda_{\text{NP}} > 0.08 \text{ (1) TeV},$

Wednesday, October 20, 2010

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$$[(\bar{t}, \bar{b})_L X_Q(u, d)_L]^2$$
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$$D^0 - \overline{D}^0 \text{ (present):} \quad L < 1.8 \left(\frac{\Lambda_{\text{NP}}}{1 \text{ TeV}}\right); \quad \Lambda_{\text{NP}} > 0.57 (7.2) \text{ TeV},$$

Wednesday, October 20, 2010

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Despective (A) suppression:
Define: $L \equiv |X_Q^{\Delta F=2}|$

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Despite $O(\lambda_{1,0}^{5})$ suppression, CPV in D mixing is more powerful in constraining 3rd gen' FCNC!

Friday, October 22, 2010

On the potential power of $D^0 \rightarrow \ell^+ \ell^-$

Practically no SM short distance:

Burdman, Golowich, Hewett & Pakvasa PRD (02).



$$\mathcal{B}r_{D^0 \to \mu^+ \mu^-}^{(\gamma\gamma)} \simeq 2.7 \times 10^{-5} \mathcal{B}r_{D^0 \to \gamma\gamma} \sim 10^{-13}$$

Donnerstag, 9. September 2010

On the potential power of $D^0 \rightarrow \ell^+ \ell^-$

Practically no SM short distance:





$$\mathcal{B}r_{D^0 o \mu^+ \mu^-}^{(\gamma \gamma)} \simeq 2.7 imes 10^{-5} \mathcal{B}r_{D^0 o \gamma \gamma} \sim 10^{-13}$$

 $\mathcal{H}_{\mathcal{V}} = \bar{u}_L \gamma^\mu c_L V_\mu + \bar{\ell} \gamma^\mu \ell V_\mu + \dots$ Golowich, Hewett, Pakvasa & Petrov, PRD (09)

On the potential power of $D^0 \rightarrow \ell^+ \ell^-$



Friday, October 22, 2010

Conclusions

• uFCNC is playing important role in learning about the microscopic world.

Charm phys. remove "immunities" => constrains alignments.

\diamond CPV in $D - \overline{D}$ mixing extremely powerful:

(i) disfavors SUSY alignment; (ii) constraining RS alignment;(iii) approaching 1TeV MFV models (factor of a few away).

Constraining 3rd generation physics.

 $\diamond D^0 \rightarrow \mu^+ \mu^-$ particularly interesting, promising future.



Robust bounds for $\Delta t = 1$



 3-gen' case the structure is much richer (8 Gell-Mann matrices), a "covariant" treatment is necessary.
 Simplification: @ LHC light quark jets look the same.

Approximate U(2) Limit of Massless Light Quarks

LHC projected bound



Friday, October 22, 2010

Flavor @ the LHC, spectrum/couplings very important



Grossman et al. (09); Gedalia & Perez (10)

Parametric solutions to the RS little CP problem & some LHC implications.





Friday, October 22, 2010

U-anarchy - constrained by D phys.

Generic warped models (up-type anarchy): Agashe, et. al (04,06).

Observable	M_G^{\min}	[TeV]	$y_{5\mathrm{D}}^{\mathrm{min}}$ or $f_{Q_3}^{\mathrm{max}}$		
	IR Higgs	$\beta = 0$	IR Higgs	eta=0	
$ ext{CPV-}B_d^{LLLL}$	$12f_{Q_{3}}^{2}$	$12f_{Q_{3}}^{2}$	$f_{Q_3}^{\rm max} = 0.5$	$f_{Q_3}^{\rm max} = 0.5$	
$ ext{CPV-}B_d^{LLRR}$	$4.2/y_{5D}$	$2.4/y_{5D}$	$y_{5\mathrm{D}}^{\mathrm{min}} = 1.4$	$y_{\rm 5D}^{\rm min}=0.82$	
$CPV-D^{LLLL}$	$0.73 f_{Q_3}^2$	$0.73 f_{Q_3}^2$	no bound	no bound	
$CPV-D^{LLRR}$	$4.9/y_{5{ m D}}$	$2.4/y_{5D}$	$y_{5\mathrm{D}}^{\mathrm{min}} = 1.6$	$y_{5\mathrm{D}}^{\mathrm{min}}=0.8$	
ϵ_K^{LLLL}	$7.9 f_{Q_3}^2$	$7.9f_{Q_3}^2$	$f_{Q_3}^{\max} = 0.62$	$f_{Q_3}^{\max} = 0.62$	
ϵ_{K}^{LLRR}	$49/y_{5D}$	$24/y_{5D}$	above (6.7)	$y_{5\mathrm{D}}^{\mathrm{min}} = 8$	
	$CPV-B_{d}^{LLLL}$ $CPV-B_{d}^{LLRR}$ $CPV-D^{LLLL}$ $CPV-D^{LLRR}$ ϵ_{K}^{LLLL}	$\begin{array}{c} & \mbox{IR Higgs} \\ & \mbox{IR Higgs} \\ & \mbox{CPV-}B_d^{LLLL} & 12f_{Q_3}^2 \\ & \mbox{CPV-}B_d^{LLRR} & 4.2/y_{5D} \\ & \mbox{CPV-}D^{LLLL} & 0.73f_{Q_3}^2 \\ & \mbox{CPV-}D^{LLRR} & 4.9/y_{5D} \\ & \mbox{ϵ_K^{LLLL}} & 7.9f_{Q_3}^2 \end{array}$	$\begin{array}{c c} & \text{IR Higgs} & \beta = 0 \\ \hline & \text{IR Higgs} & \beta = 0 \\ \hline & \text{CPV-}B_d^{LLLL} & 12f_{Q_3}^2 & 12f_{Q_3}^2 \\ \hline & \text{CPV-}B_d^{LLRR} & 4.2/y_{5D} & 2.4/y_{5D} \\ \hline & \text{CPV-}D^{LLLL} & 0.73f_{Q_3}^2 & 0.73f_{Q_3}^2 \\ \hline & \text{CPV-}D^{LLRR} & 4.9/y_{5D} & 2.4/y_{5D} \\ \hline & \epsilon_K^{LLLL} & 7.9f_{Q_3}^2 & 7.9f_{Q_3}^2 \end{array}$	$\begin{array}{c c} \mbox{IR Higgs} & \beta = 0 & \mbox{IR Higgs} \\ \hline \mbox{CPV-}B_d^{LLLL} & 12f_{Q_3}^2 & 12f_{Q_3}^2 & f_{Q_3}^{\max} = 0.5 \\ \hline \mbox{CPV-}B_d^{LLRR} & 4.2/y_{5D} & 2.4/y_{5D} & y_{5D}^{\min} = 1.4 \\ \hline \mbox{CPV-}D^{LLLL} & 0.73f_{Q_3}^2 & 0.73f_{Q_3}^2 & \mbox{no bound} \\ \hline \mbox{CPV-}D^{LLRR} & 4.9/y_{5D} & 2.4/y_{5D} & y_{5D}^{\min} = 1.6 \\ \hline \mbox{ϵ_{K}^{LLLL}} & 7.9f_{Q_3}^2 & 7.9f_{Q_3}^2 & f_{Q_3}^{\max} = 0.62 \end{array}$	

edalia, et. al (09); sidori, et. al (10).

U-anarchy - constrained by D phys.

Generic warped models (up-type anarchy): Agashe, et. al (04,06).

(09);(10).

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ϵ_K^{LLLL}		$7.9 f_{Q_3}^2$	$7.9 f_{Q_3}^2$	$f_{Q_3}^{\max} = 0.62$	$f_{Q_3}^{\max} = 0.62$	Gedalia, et. al
ϵ_{K}^{LLRR}		$49/y_{5D}$	$24/y_{5D}$	above (6.7)	$y_{5\mathrm{D}}^{\mathrm{min}} = 8$	Isidori, et. al

RS alignment (via shining):
$$y_{5D}^d \gtrsim 3y_{5D}^u$$
_{Csaki, et. al (09).}
 $\frac{1}{2} \lesssim y_{5D} \lesssim \frac{2\pi}{N_{\text{KK}}}$ for brane Higgs; $\frac{1}{2} \lesssim y_{5D} \lesssim \frac{4\pi}{\sqrt{N_{\text{KK}}}}$ for bulk Higgs,

Factor of few improvement exclude models.