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Antimagnetic rotation in nuclei: A microscopic description

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Rotation bands in nuclei

- Substantial quadrupole deformation
- Strong electric quadrupole (E2) transitions
- Rotational bands with $\Delta I = 2$
- Coherent collective rotation of many nucleons

Bohr PR1951



Gamma-ray energy (keV)



Twin PRL1986

Magnetic Rotation (MR)

- near-spherical or weakly deformed nuclei
- strong MI and very weak E2 transitions
- rotational bands with $\Delta I = I$
- shears mechanism







Experiment: MR

Magnetic rotation: 85 nuclei



Antimagnetic Rotation (AMR)

Magnetic rotation $\langle \longrightarrow \rangle$ Ferromagnet



- \checkmark rotational bands with $\Delta I = I$
- \checkmark near spherical nuclei; weak E2 transitions
- ✓ strong MI transitions
- \checkmark B(MI) decrease with spin
- \checkmark shears mechanism
- Antimagnetic rotation $\langle \longrightarrow \rangle$ Antiferromagnet
- \checkmark rotational bands with $\Delta I = 2$
- \checkmark near spherical nuclei; weak E2 transitions
- \checkmark no MI transitions
- \checkmark B(E2) decrease with spin
- ✓ two "shears-like" mechanism





Experiment: AMR



Small B(E2) Decrease tendency

Large J(2)/B(E2) Increase tendency



Simons PRL2003; Simons PRC2005

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Theory

✓ Semiclassical particle plus rotor model Clark ARNPS2000 simple geometry for the energies and transition probabilities

✓ Pairing-plus-quadrupole tilted axis cranking (TAC) model a schematic Hamiltonian Frauendorf NPA1993; Frauendorf NPA2000

A fully self-consistent microscopic investigation?



DFT: Cranking version

TAC based on Covariant Density Functional Theory

Meson exchange version:

3-D Cranking: *Madokoro, Meng, Matsuzaki, Yamaji, PRC 62, 061301 (2000)*2-D Cranking: *Peng, Meng, Ring, Zhang, PRC 78, 024313 (2008)*<u>Point-coupling version:</u> Simple and more suitable for systematic investigations
2-D Cranking: *PWZ, Zhang, Peng, Liang, Ring, Meng, PLB 699, 181 (2011)*

TAC based on Skyrme Density Functional Theory

3-D Cranking: Olbratowski, Dobaczewski, Dudek, Płóciennik, PRL 93, 052501(2004) 2-D Cranking: Olbratowski, Dobaczewski, Dudek, Rzaca-Urban, Marcinkowska, Lieder, APPB 33, 389(2002)

Fully self-consistent microscopic investigations

Fully taken into account polarization effects

self-consistently treated the nuclear currents

> without any adjustable parameters for rotational excitations

Tilted axis cranking CDFT



Transformed to the frame rotating with the uniform velocity

$$\Omega = (\Omega_x, 0, \Omega_z) = (\Omega \cos \theta_\Omega, 0, \Omega \sin \theta_\Omega)$$
$$x^{\alpha} = \begin{pmatrix} t \\ \mathbf{x} \end{pmatrix} \rightarrow \tilde{x}^{\alpha} = \begin{pmatrix} \tilde{t} \\ \tilde{\mathbf{x}} \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & \mathbf{R} \end{pmatrix} \begin{pmatrix} t \\ \mathbf{x} \end{pmatrix}$$



Koepf NPA1989; Kaneko PLB1993; Madokoro PRC1997

Equation of motion

Dirac equation

$$\begin{array}{ccc} m + \mathbf{S} + \mathbf{V} - \mathbf{\Omega} \cdot \mathbf{J} & \sigma(p - \mathbf{V}) \\ \sigma(p - \mathbf{V}) & -m - \mathbf{S} + \mathbf{V} - \mathbf{\Omega} \cdot \mathbf{J} \end{array} \left| \begin{pmatrix} f \\ g \end{pmatrix} = \varepsilon \begin{pmatrix} f \\ g \end{pmatrix}$$

$$V(r) = \alpha_{V}\rho_{V} + \gamma_{V}\rho_{V}^{3} + \delta_{V}\Delta\rho_{V} + \tau_{3}\alpha_{TV}\rho_{TV} + \tau_{3}\delta_{TV}\Delta\rho_{TV} + e\frac{1-\tau_{3}}{2}A$$
$$V(r) = \alpha_{V}\mathbf{j}_{V} + \gamma_{V}\mathbf{j}_{V}^{3} + \delta_{V}\Delta\mathbf{j}_{V} + \tau_{3}\alpha_{TV}\mathbf{j}_{TV} + \tau_{3}\delta_{TV}\Delta\mathbf{j}_{TV} + e\frac{1-\tau_{3}}{2}A$$
$$S(r) = \alpha_{S}\rho_{S} + \beta_{S}\rho_{S}^{2} + \gamma_{S}\rho_{S}^{3} + \delta_{S}\Delta\rho_{S}$$

V(r) vector potential time-like V(r) vector potential space-like

S(r) scalar potential

Observables

Binding energy

$$E_{\text{tot}} = \sum_{k=1}^{A} \epsilon_{k} - \int d^{3}r \left\{ \frac{1}{2} \alpha_{S} \rho_{S}^{2} + \frac{1}{2} \alpha_{V} j_{V}^{\mu} (j_{V})_{\mu} \right. \\ \left. + \frac{1}{2} \alpha_{TV} j_{TV}^{\mu} (j_{TV})_{\mu} + \frac{2}{3} \beta_{S} \rho_{S}^{3} + \frac{3}{4} \gamma_{S} \rho_{S}^{4} \right. \\ \left. + \frac{3}{4} \gamma_{V} \left(j_{V}^{\mu} (j_{V})_{\mu} \right)^{2} + \frac{1}{2} \delta_{S} \rho_{S} \Delta \rho_{S} + \frac{1}{2} \delta_{V} (j_{V})_{\mu} \Delta j_{V}^{\mu} \right. \\ \left. + \frac{1}{2} \delta_{TV} j_{TV}^{\mu} \Delta (j_{TV})_{\mu} + \frac{1}{2} e j_{p}^{0} A_{0} \right\} + \sum_{k=1}^{A} \langle k | \boldsymbol{\Omega} \, \hat{\boldsymbol{J}} | k \rangle$$

 $+ E_{c.m.}$

Angular momentum

$$J = \sqrt{\langle \hat{J}_x \rangle^2 + \langle \hat{J}_z \rangle^2} \equiv \sqrt{I(I+1)}$$

Observables

Binding energy

$$E_{\text{tot}} = \sum_{k=1}^{A} \epsilon_{k} - \int d^{3}r \left\{ \frac{1}{2} \alpha_{S} \rho_{S}^{2} + \frac{1}{2} \alpha_{V} j_{V}^{\mu} (j_{V})_{\mu} \right. \\ \left. + \frac{1}{2} \alpha_{TV} j_{TV}^{\mu} (j_{TV})_{\mu} + \frac{2}{3} \beta_{S} \rho_{S}^{3} + \frac{3}{4} \gamma_{S} \rho_{S}^{4} \right. \\ \left. + \frac{3}{4} \gamma_{V} (j_{V}^{\mu} (j_{V})_{\mu})^{2} + \frac{1}{2} \delta_{S} \rho_{S} \Delta \rho_{S} + \frac{1}{2} \delta_{V} (j_{V})_{\mu} \Delta j_{V}^{\mu} \right. \\ \left. + \frac{1}{2} \delta_{TV} j_{TV}^{\mu} \Delta (j_{TV})_{\mu} + \frac{1}{2} e j_{p}^{0} A_{0} \right\} + \sum_{k=1}^{A} \langle k | \boldsymbol{\Omega} \, \hat{\boldsymbol{J}} | k \rangle$$

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$$J = \sqrt{\langle \hat{J}_x \rangle^2 + \langle \hat{J}_z \rangle^2} \equiv \sqrt{I(I+1)}$$

B(M1) and B(E2) transition probabilites

$$B(M1) = \frac{3}{8\pi} \mu_{\perp}^2 = \frac{3}{8\pi} (\mu_x \sin \theta_J - \mu_z \cos \theta_J)^2,$$

$$B(E2) = \frac{3}{8} \left[Q_{20}^p \cos^2 \theta_J + \sqrt{\frac{2}{3}} Q_{22}^p (1 + \sin^2 \theta_J) \right]^2,$$

AMR in ¹⁰⁵Cd

First odd-A nucleus with antimagnetic rotation



Choudhury et al, PRC 82,061308 (2010)

Numerical Details

- ✓ Harmonic oscillator shells: Nf = 10
- ✓ Effective interaction: PC-PK1
- ✓ Configurations: $\nu [h_{11/2}(g_{7/2})^2] \otimes \pi [(g_{9/2})^{-2}]$
- ✓ Polarizations:



PWZ, Peng, Liang, Ring, Meng PRL 107, 122501(2011)

Single particle routhians



✓ Time reversal symmetry broken: energy splitting

- \checkmark For proton, two holes in the top of $g_{9/2}$ shell
- ✓ For neutron, one particle in the bottom of h_{11/2} shell, the other six are distributed over the (gd) shell with strong mixing
- ✓ This configuration is similar to $\nu[h_{11/2}(g_{7/2})^2] \otimes \pi[(g_{9/2})^{-2}]$, but not exactly

Energy and angular momentum



 \checkmark The energy and total angular momentum agree well with the data.

- ✓ The spin increase linearly with frequency / nearly constant moment of inertia.
- ✓ Without polarization, a much smaller frequency is needed to reach the same angular momentum.
- ✓ Without polarization, there is a maximal angular momentum of roughly 17ħ

B(E2) and deformation



- ✓ The calculated B(E2) values are in excellent agreement with the data.
- ✓ The B(E2) values decrease with the increasing spin / two "shears-like" mechanism.
- ✓ Without polarization, the B(E2) values are reduced to only ~60% of the self-consistent results, and dropped to zero when the frequency Ω≥0:5 MeV.
- ✓ It is of importance to emphasize that polarization effects play a very important role in the self-consistent microscopic description of AMR bands, especially for theE2 transitions.

Two shears mechanism



- The two proton angular momentum are pointing opposite to each other and are nearly perpendicular to the neutron angular momentum. They form the blades of the two shears.
- Increasing Ω, the two proton blades towards to each other and generates the total angular momentum.

Summary

Covariant density functional theory has been extended to describe rotational excitations including MR and AMR.

≻¹⁰⁵Cd: AMR

reproduce well the AMR pictures, E, I, and B(E2) values in a fully self-consistent microscopic way for the first time

In collaboration with

Haozhao Liang Jie Meng Jing Peng Peter Ring Shuangquan Zhang

Thank You!

Single-particle angular momentum



- ✓ For the protons, only the two holes in the $g_{9/2}$ shell contribute.
- $\checkmark\,$ For the neutrons, only the particles above the N= 50 shell contribute.
- ✓ Angular momentum results from the alignment of proton holes and the mixing within the neutron orbitals.
- ✓ Due to the strong mixing between neutrons, a core in the phenomenological model cannot be well defined.