Topical Seminar on Frontier of Particle Physics 2007-2008

Neutrino Physics and Astrophysics

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Neutrino Mass Models part I

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Caveats:

- there is not an established theory of neutrino masses and mixing angles
- the existing models (too many!) reflect on the one side our prejudices and on the other side the lack of a unifying principle

part I (general ideas)

- 1. How to extend the SM to incorporate neutrino masses
- 2. Purely Dirac neutrino masses
- 3. Neutrino masses from D=5 operator
- 1. D=5 operator and see-saw mechanism
- 2. Additional tests of D=5 operator
- 3. Flavour symmetry (I): the hierarchy puzzle

Part II: much more speculative!
Flavour symmetry (II): the lepton mixing puzzle

Beyond the Standard Model

a non-vanishing neutrino mass is the first evidence of the incompleteness of the Standard Model [SM]

in the SM neutrinos belong to SU(2) doublets with hypercharge Y=-1/2 they have only two helicities (not four, as the other charged fermions)

$$l = \begin{pmatrix} v_e \\ e \end{pmatrix} = (1, 2, -1/2)$$

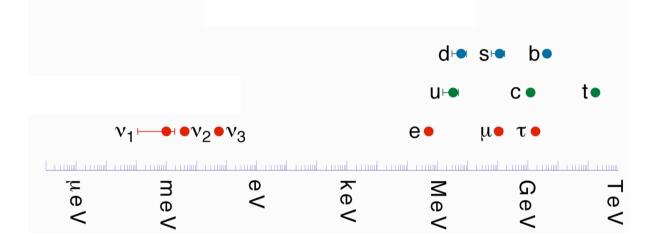
the requirement of invariance under the gauge group $G=SU(3)\times SU(2)\times U(1)$ forbids pure fermion mass terms in the lagrangian. Charged fermion masses arise, after electroweak symmetry breaking, through gauge-invariant Yukawa interactions

not even this term is allowed for SM neutrinos, by gauge invariance

Questions

how to extend the SM in order to accommodate neutrino masses?

why neutrino masses are so small, compared with the charged fermion masses?



why lepton mixing angles are so different from those of the quark sector?

$$U_{PMNS} = \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0\\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}}\\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} + \text{corrections}$$

$$U_{PMNS} = \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0\\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}}\\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} + \text{corrections} \quad V_{CKM} \approx \begin{pmatrix} 1 & O(\lambda) & O(\lambda^4 \div \lambda^3)\\ O(\lambda) & 1 & O(\lambda^2)\\ O(\lambda^4 \div \lambda^3) & O(\lambda^2) & 1 \end{pmatrix}$$

$$\lambda \approx 0.22$$

How to modify the SM?

the SM, as a consistent QFT, is completely specified by

- 0. invariance under local transformations of the gauge group $G=SU(3)\times SU(2)\times U(1)$ [plus Lorentz invariance]
- 1. particle content three copies of (q,u^c,d^c,l,e^c) one Higgs doublet Φ
- 2. renormalizability (i.e. the requirement that all coupling constants g_i have non-negative dimensions in units of mass: d(g_i)≥0. This allows to eliminate all the divergencies occurring in the computation of physical quantities, by redefining a finite set of parameters.)

(0.+1.+2.) leads to the SM Lagrangian, L_{SM} , possessing an additional, accidental, global symmetry: (B-L)

O. We cannot give up gauge invariance! It is mandatory for the consistency of the theory. Without gauge invariance we cannot even define the Hilbert space of the theory [remember: we need gauge invariance to eliminate the photon extra degrees of freedom required by Lorentz invariance]!

We could extend G, but, to allow for neutrino masses, we need to modify 1. (and/or 2.) anyway...

First possibility: modify (1), the particle content

there are several possibilities one of the simplest one is to mimic the charged fermion sector

Example 1 $\begin{cases} \text{add (three copies of)} & v^c \equiv (1,1,0) \\ \text{right-handed neutrinos} \end{cases} v^c \equiv (1,1,0) \quad \text{full singlet under } G=SU(3)\times SU(2)\times U(1) \\ \text{ask for (global) invariance under B-L} \\ \text{(no more automatically conserved as in the SM)} \end{cases}$

the neutrino has now four helicities, as the other charged fermions, and we can build gauge invariant Yukawa interactions giving rise, after electroweak symmetry breaking, to neutrino masses

$$L_{Y} = d^{c} y_{d} (\Phi^{+} q) + u^{c} y_{u} (\tilde{\Phi}^{+} q) + e^{c} y_{e} (\Phi^{+} l) + v^{c} y_{v} (\tilde{\Phi}^{+} l) + h.c.$$

$$m_{f} = \frac{y_{f}}{\sqrt{2}} v \qquad f = u, d, e, v$$

with three generations there is an exact replica of the quark sector and, after diagonalization of the charged lepton and neutrino mass matrices, a mixing matrix U appears in the charged current interactions

$$-\frac{g}{\sqrt{2}}W_{\mu}^{-}\overline{e}\sigma^{\mu}U_{PMNS}\nu+h.c.$$

 $-\frac{g}{\sqrt{2}}W_{\mu}^{-}e\sigma^{\mu}U_{PMNS}v + h.c.$ U_{PMNS} has three mixing angles and one phase, like V_{CKM}

a generic problem of this approach

the particle content can be modified in several different ways in order to account for non-vanishing neutrino masses (additional right-handed neutrinos, new SU(2) fermion triplets, additional SU(2) scalar triplet(s), SUSY particles,...). Which is the correct one?

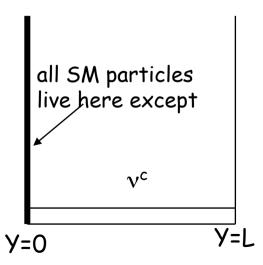
a problem of the above example

if neutrinos are so similar to the other fermions, why are so light?

$$\frac{y_v}{y_{top}} \le 10^{-12}$$

Quite a speculative answer:

neutrinos are so light, because the right-handed neutrinos have access to an extra (fifth) spatial dimension



neutrino Yukawa coupling

$$v^{c}(y=0)(\tilde{\Phi}^{+}l)$$
 = Fourier expansion
$$= \frac{1}{\sqrt{L}}v_{0}^{c}(\tilde{\Phi}^{+}l) + \dots \text{ [higher modes]}$$

if L>>1 (in units of the fundamental scale) then neutrino Yukawa coupling is suppressed

Second possibility: abandon (2) renormalizability

A disaster?

$$L = L_{d \le 4}^{SM} + \frac{L_5}{\Lambda} + \frac{L_6}{\Lambda^2} + \dots$$

a new scale Λ enters the theory. The new (gauge invariant!) operators L_5 , L_6 ,... contribute to amplitudes for physical processes with terms of the type

$$\frac{L_5}{\Lambda} \to \frac{E}{\Lambda} \qquad \frac{L_6}{\Lambda^2} \to \left(\frac{E}{\Lambda}\right)^2 \qquad \dots$$

the theory cannot be extrapolated beyond a certain energy scale $E \approx \Lambda$. [at variance with a renormalizable (asymptotically free) QFT]

If E<< Λ (for example E close to the electroweak scale, 10^2 GeV, and $\Lambda \approx 10^{15}$ GeV not far from the so-called Grand Unified scale), the above effects will be tiny and, the theory will *look like* a renormalizable theory!

$$\frac{E}{\Lambda} \approx \frac{10^2 \, GeV}{10^{15} \, GeV} = 10^{-13}$$

an extremely tiny effect, but exactly what needed to suppress m_{ν} compared to m_{top} !

Worth to explore. The dominant operators (suppressed by a single power of $1/\Lambda$) beyond L_{SM} are those of dimension 5. Here is a list of all d=5 gauge invariant operators

$$\frac{L_5}{\Lambda} = \frac{\left(\tilde{\Phi}^+ l\right)\left(\tilde{\Phi}^+ l\right)}{\Lambda} =$$
$$= \frac{v}{2} \left(\frac{v}{\Lambda}\right) vv + \dots$$

a unique operator!
[up to flavour combinations]
it violates (B-L) by two units

 $= \frac{v}{2} \left(\frac{v}{\Lambda} \right) vv + \dots$ it is suppressed by a factor (v/ Λ) with respect to the neutrino mass term of Example 1: $v^{c}(\tilde{\Phi}^{+}l) = \frac{v}{\sqrt{2}}v^{c}v + \dots$

it provides an explanation for the smallness of m_{ν} :

the neutrino masses are small because the scale Λ , characterizing (B-L) violations, is very large. How large? Up to about 10^{15} GeV

from this point of view neutrinos offer a unique window on physics at very large scales, inaccessible in present (and probably future) man-made experiments.

since this is the dominant operator in the expansion of L in powers of $1/\Lambda$, we could have expected to find the first effect of physics beyond the SM in neutrinos ... and indeed this was the case!

L_5 represents the effective, low-energy description of several extensions of the SM

Example 2: see-saw

add (three copies of) $v^c \equiv (1,1,0)$

full singlet under $G=SU(3)\times SU(2)\times U(1)$

this is like Example 1, but without enforcing (B-L) conservation

$$L(v^{c}, l) = v^{c} y_{v} (\tilde{\Phi}^{+} l) + \frac{1}{2} v^{c} M v^{c} + h.c.$$

mass term for right-handed neutrinos: G invariant, violates (B-L) by two units.

the new mass parameter M is independent from the electroweak breaking scale v. If M>>v, we might be interested in an effective description valid for energies much smaller than M. This is obtained by "integrating out" the field v^c

$$L_{\rm eff}(l) = -\frac{1}{2}(\tilde{\Phi}^+ l) \Big[y_{\nu}^T M^{-1} y_{\nu} \Big] (\tilde{\Phi}^+ l) + h.c. + \dots \text{ terms suppressed by more powers of M-1}$$

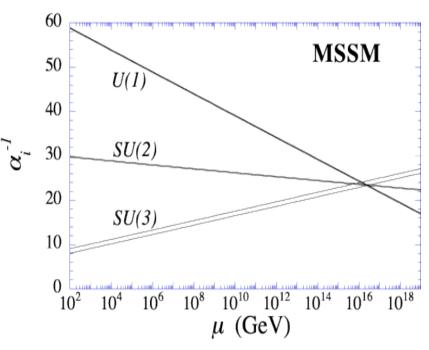
this reproduces L_5 , with M playing the role of Λ . This particular mechanism is called (type I) see-saw.

Theoretical motivations for the see-saw

 $\Lambda \approx 10^{15}$ GeV is very close to the so-called unification scale M_{GUT} .

an independent evidence for M_{GUT} comes from the unification of the gauge coupling constants in (SUSY extensions of) the SM.

such unification is a generic prediction of Grand Unified Theories (GUTs): the SM gauge group G is embedded into a simple group such as SU(5), SO(10),...



Particle classification: it is possible to unify all SM fermions (1 generation) into a single irreducible representation of the GUT gauge group. Simplest example: G_{GUT} =SO(10)

$$16 = (q, d^c, u^c, l, e^c, v^c)$$
 a whole family plus a right-handed neutrino!

quite a fascinating possibility. Unfortunately, it still lacks experimental tests. In GUT new, very heavy, particles can convert quarks into leptons and the proton is no more a stable particle. Proton decay rates and decay channels are however model dependent. Experimentally we have only lower bounds on the proton lifetime.

2 additional virtues of the see-saw

The see-saw mechanism can enhance small mixing angles into large ones

$$m_{v} = -\left[y_{v}^{T} M^{-1} y_{v}\right] v^{2}$$

Example with 2 generations

$$y_{v} = \begin{pmatrix} \delta & \delta \\ 0 & 1 \end{pmatrix} \qquad \begin{array}{c} \delta <<1 \\ \text{small mixing} \end{array}$$

$$M = \begin{pmatrix} M_{1} & 0 \\ 0 & M_{2} \end{pmatrix} \text{ no mixing}$$

$$y_{v}^{T} M^{-1} y_{v} = \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \frac{\delta^{2}}{M_{1}} + \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix} \frac{1}{M_{2}}$$
$$\approx \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \frac{\delta^{2}}{M_{1}} \quad \text{for } \frac{M_{1}}{M_{2}} << \delta^{2}$$

The (out-of equilibrium, CP-violating) decay of heavy right-handed neutrinos in the early universe might generate a net asymmetry between leptons and anti-leptons. Subsequent SM interactions can partially convert it into the observed baryon asymmetry

$$\eta = \frac{(n_B - n_{\overline{B}})}{s} \approx 6 \times 10^{-10}$$

weak point of the see-saw

full high-energy theory is difficult to test

$$L(v^{c}, l) = v^{c} y_{v} (\tilde{\Phi}^{+} l) + \frac{1}{2} v^{c} M v^{c} + h.c.$$

depends on many physical parameters:

- 3 (small) masses + 3 (large) masses
- 3 (L) mixing angles + 3 (R) mixing angles
- 6 physical phases = 18 parameters

the double of those describing $(L_{SM})+L_5$: 3 masses, 3 mixing angles and 3 phases

few observables to pin down the extra parameters: η ,... [additional possibilities exist under special conditions, e.g. Lepton Flavor Violation at observable rates]

easier to test the low-energy remnant L_5

[which however is "universal" and does not implies the specific see-saw mechanism of Example 2]

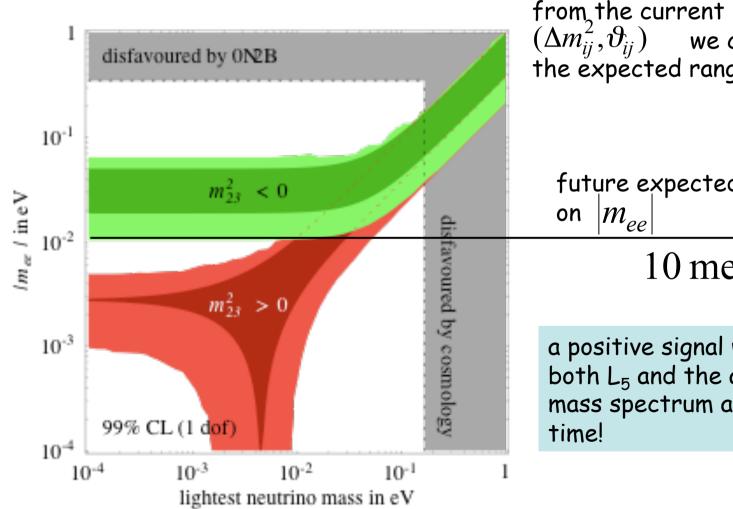
look for a process where B-L is violated by 2 units. The best candidate is $0\nu\beta\beta$ decay: $(A,Z)-\lambda(A,Z+2)+2e^{-1}$ this would discriminate L₅ from other possibilities, such as Example 1.

The decay in $0\nu\beta\beta$ rates depend on the combination $\left|m_{ee}\right| = \left|\sum U_{ei}^2 m_i\right|$

$$\left| m_{ee} \right| = \left| \sum_{i} U_{ei}^{2} m_{i} \right|$$

$$|m_{ee}| = |\cos^2 \vartheta_{13} (\cos^2 \vartheta_{12} \ m_1 + \sin^2 \vartheta_{12} e^{2i\alpha} \ m_2) + \sin^2 \vartheta_{13} e^{2i\beta} \ m_3|$$

[notice the two phases α and β , not entering neutrino oscillations]



from the current knowledge of $(\Delta m_{ij}^2, \vartheta_{ij})$ we can estimate the expected range of $|m_{ee}|$

future expected sensitivity

10 meV

a positive signal would test both L_5 and the absolute mass spectrum at the same

Flavor symmetries I (the hierarchy puzzle)

hierarchies in fermion spectrum

$$\frac{g}{m_t} = \frac{m_u}{m_t} << \frac{m_c}{m_t} << 1 \qquad \frac{m_d}{m_b} << \frac{m_s}{m_b} << 1 \qquad |V_{ub}| << |V_{cb}| << |V_{us}| \equiv \lambda < 1$$

$$\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2} = (0.025 \div 0.049) \approx \lambda^2 << 1 \quad (2\sigma)$$

$$|U_{e3}| < 0.18 \le \lambda \quad (2\sigma)$$

call ξ_i the generic small parameter. A modern approach to understand why ξ_i <-1 consists in regarding ξ_i as small breaking terms of an approximate flavour symmetry. When ξ_i =0 the theory becomes invariant under a flavour symmetry F

Example: why $y_e << y_{top}$? Assume $F=U(1)_F$

F(t)=F(t^c)=F(h)=0
$$y_{top}(h+v)t^ct$$
 allowed
$$F(e^c)=p>0 \ F(e)=q>0 \ y_e(h+v)e^ce \ \text{breaks U(1)}_F \ \text{by (p+q) units}$$
 if $\xi=\langle \phi \rangle/\Lambda <1$ breaks U(1) by one negative unit $y_e \approx O(\xi^{p+q}) << y_{top} \approx O(1)$

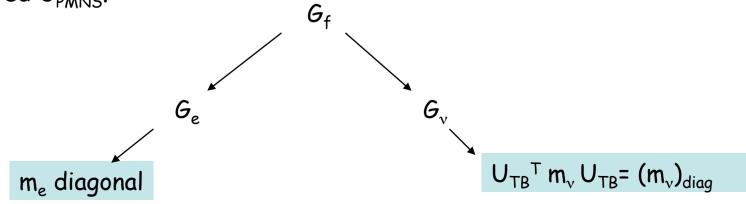
provides a qualitative picture of the existing hierarchies in the fermion spectrum

Flavor symmetries II (the lepton mixing puzzle)

$$\text{why} \quad U_{PMNS} \approx U_{TB} \equiv \begin{bmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0\\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}}\\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{bmatrix} ?$$
 [TB=TriBimaximal]

$$U_{PMNS} = U_e^+ U_v$$

Consider a flavor symmetry G_f such that G_f is broken into two different subgroups: G_e in the charged lepton sector, and G_v in the neutrino sector. m_e is invariant under G_e and m_v is invariant under G_v . If G_e and G_v are appropriately chosen, the constraints on m_e and m_v can give rise to the observed U_{PMNS} .



The simplest example is based on a small discrete group, $G_f = A_4$. It is the subgroup of SO(3) leaving a regular tetrahedron invariant. The elements of A_4 can all be generated starting from two of them: S and T such that

$$S^2 = T^3 = (ST)^3 = 1$$

S generates a subgroup Z_2 of A_4 T generates a subgroup Z_3 of A_4

simple models have been constructed where G_e = Z_3 and G_v = Z_2 and where the lepton mixing matrix U_{PMNS} is automatically U_{TB} , at the leading order in the SB parameters. Small corrections are induced by higher order terms.

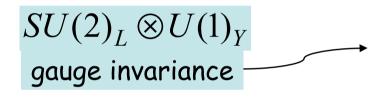
the generic predictions of this approach is that θ_{13} and $(\theta_{23}-\pi/4)$ are very small quantities, of the order of few percent: testable in a not-so-far future.

Conclusion (theory)

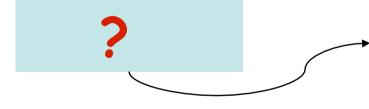
theory of neutrino masses

it does not exist! Neither for neutrinos nor for charged fermions. We lack a unifying principle.

like weak interactions before the electroweak theory



all fermion-gauge boson interactions in terms of 2 parameters: g and g'



Yukawa interactions between fermions and spin 0 particles: many free parameters (up to 22 in the SM!)

only few ideas and prejudices about neutrino masses and mixing angles

caveat: several prejudices turned out to be wrong in the past!

- m_v≈10 eV because is the cosmologically relevant range
- solution to solar is MSW Small Angle
- atmospheric neutrino problem will go away because it implies a large angle