

Shapes and energy spectra of hypernuclei in the Skyrme Hartree-Fock method

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Outline

- Introduction
- Extended Skyrme Hartree-Fock
- Properties of hypernuclei shapes of hypernuclei enegery spetra







Why to study hypernuclei?

Nucleon-nucleon interaction Hyperon-nucleon interaction Impurity of nuclear system

Multistrange system: Neutron star, ...



Theoretical studies about hypernuclei

Energy spectrum Decay properties The effect of hyperon(s) Shapes



Theoretical studies about the shapes of hypernuclei

Studies based on spherical symmetry:

- 1. Relativistic mean-field model (RMF)
- 2. Skyrme Hartree-Fock model (SHF)
- 3. Woods-Saxon potential + YN interaction
- 4. Few-body theory



Calculations considering deformation:

Deformed HF with nonrealistic YN interaction:

T. H. Ho and A.Volkov, Phys. Lett. B30, 303, 1969.W. H. Bassichis, A. Gal, Phys. Rev. C1, 28, 1970.J. Zofka, Czech, J. Phys. B30, 95, 1980.

Nilsson Model:

assume the same deformation for core and hypernuclei: K. Hagino, Phys. Rev. C63, 044318, 2001



Deformed SHF with Microscopic YN int.

X.-R. Zhou, H.-J. Schulze, H. Sagawa, et al., Phys. Rev. C76, 034312(2007)

Relativistic mean-field model (RMF): Myaing Thi Win, H. Hagino, et al., Phys. Rev. C 78, 054311 (2008)

Triaxial SHF with Skyrme-like YN interaction: Myaing Thi Win, H. Hagino, et al., Phys. Rev. C **83**, 014301 (2011)

Antisymmetrized molecular dynamics (AMD): M. Isaka,, et al., Phys. Rev. C **83**, 044323 (2011)

Triaxial RMF: Bing-Nan Lu (吕炳楠), S.-G.Zhou, Phys. Rev. C **84**, 014328 (2011)



Superdeformation in hypernuclei



Bing-Nan Lu, E. Hiyama, H. Sagawa, and S.-G. Zhou, Phys. Rev. C 89, 044307 (2014)

M. Isaka, K. Fukukawa, M. Kimura, E. Hiyama, H. Sagawa, and Y. Yamamoto, Phys. Rev. C 89, 024310 (2014) 9



Why to study deformations of hypernuclei

Many p-shell and sd-shell nuclei are deformed.

For example, experimentally, ¹⁰B and ¹¹C have large quadrupole moments.

F. Ajzenberg-Selove, Nucl. Phys. A490, 1 (1988); A506, 1(1990).

Also, ⁸Be is known to be strongly deformed due to its double-a structure.



Several models for deformed nuclei

Alpha-model

Projected shell model (PSM)

Deformed Skyrme Hartree-Fock (DSHF)

Relativistic mean-field model (RMF)

Antisymmetrized molecular dynamics (AMD)

Microscopic hyperon-nucleon interaction

for deformed hypernuclei



Comparison of different hyperon-nucleon potential



H.-J. Schulze PHYSICAL REVIEW C 88, 024322 (2013)



Extended DSHF including hyperon-nucleon interaction

DSHF + YN interaction (NSC89):

Total energy of a hypernucleus in extended DSHF:

$$E = \int d^3 r \,\epsilon_{\rm SHF}(r)$$

where the energy density

 $\boldsymbol{\varepsilon}_{\mathsf{SHF}} = \boldsymbol{\varepsilon}_{N}[\rho_{n}, \rho_{p}, \tau_{n}, \tau_{p}, \boldsymbol{J}_{n}, \boldsymbol{J}_{p}] + \boldsymbol{\varepsilon}_{\Lambda}[\rho_{n}, \rho_{p}, \rho_{\Lambda}, \tau_{\Lambda}]$

Due to the YN force,

$$\epsilon_{\Lambda} = \frac{\tau_{\Lambda}}{2m_{\Lambda}} + \epsilon_{N\Lambda}(\rho_N, \rho_{\Lambda}) + \left(\frac{m_{\Lambda}}{m_{\Lambda}^*(\rho_n, \rho_p, \rho_{\Lambda})} - 1\right) \frac{\tau_{\Lambda} - C\rho_{\Lambda}^{5/3}}{2m_{\Lambda}}$$



The energy density functional $\varepsilon_{N\Lambda}$ is obtained from a fit *to the* binding energy per baryon, $B/A(\rho_n, \rho_p, \rho)$, of asymmetric hypermatter, as generated by BHF calculations.

$$\varepsilon_{N\Lambda} = (\rho_n + \rho_p + \rho_\Lambda) \frac{B}{A} (\rho_n, \rho_p, \rho_\Lambda)$$
$$- (\rho_n + \rho_p) \frac{B}{A} (\rho_n, \rho_p, 0) - \frac{C\rho_\Lambda^{5/3}}{2m_\Lambda}$$

Effective mass of hyperon

$$\frac{m_{\Lambda}^{*}}{m_{\Lambda}} = \left[1 + \frac{U_{\Lambda}\left(k_{F}^{(\Lambda)}\right) - U_{\Lambda}(0)}{k_{F}^{(\Lambda)^{2}}/2m_{\Lambda}}\right]^{-1}$$



In practice we use the following parametrizations:

$$\varepsilon_{N\Lambda} \approx -[368 - (1717 + 268\alpha - 920\alpha^2)\rho_N + (2932 - 776\alpha + 2483\alpha^2)\rho_N^2]\rho_N\rho_\Lambda + (449 - 2470\rho_N + 5834\rho_N^2)\rho_N\rho_\Lambda^{5/3},$$

 $\frac{m_{\Lambda}^*}{m_{\Lambda}} \approx 1 - (1.58 + 0.12\alpha - 0.12\alpha^2 + 0.54y - 0.14y^2)\rho_N$

$$+ (4.11 + 2.11\alpha + 2.88\alpha^{2} + 0.35y + 1.17y^{2})\rho_{N}^{2}$$
$$- (4.03 + 7.08\alpha + 5.18\alpha^{2} - 0.93y + 3.27y^{2})\rho_{N}^{3},$$



Extended SHF equation

Minimizing the total energy of the hypernucleus, one arrives with extended SHF equation

$$\left[-\nabla \cdot \frac{1}{2m_q^*(r)}\nabla + V_q(r) - i\nabla W_q(r) \cdot (\nabla \times \boldsymbol{\sigma})\right]\phi_q^i(\boldsymbol{r}) = -e_q^i\phi_q^i(\boldsymbol{r})$$

with the modified mean field by hyperon:

$$V_{N} = V_{N}^{\text{SHF}} + \frac{\partial \epsilon_{N\Lambda}}{\partial \rho_{N}} + \frac{\partial}{\partial \rho_{N}} \left(\frac{m_{\Lambda}}{m_{\Lambda}^{*}(\rho_{N})}\right) \frac{\tau_{\Lambda} - C\rho_{\Lambda}^{5/3}}{2m_{\Lambda}}$$
$$V_{\Lambda} = \frac{\partial \epsilon_{N\Lambda}}{\partial \rho_{\Lambda}} - \left(\frac{m_{\Lambda}}{m_{\Lambda}^{*}(\rho_{N})} - 1\right) \frac{5}{3} \frac{C\rho_{\Lambda}^{2/3}}{2m_{\Lambda}},$$

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Pairing interaction

We take a density-dependent delta pairing

$$V_q(\boldsymbol{r}_1, \boldsymbol{r}_2) = V'_q \left[1 - \frac{\rho_N(r)}{\rho_0} \right] \delta(\boldsymbol{r}_1 - \boldsymbol{r}_2)$$

For light nuclei,

Nucl. Phys. **A551**, 434 (1993)

 $V'_q = -410 \,\mathrm{MeV fm^3}$

Nucl. Phys. A722, c183, 2003

For medium-mass and heavy nuclei,

$$V'_p = -1146 \,\mathrm{MeV fm^3}$$

 $V'_n = -999 \,\mathrm{MeV fm^3}$
Euro. Phys. J. A8, 59, 2000



1. Shapes of Hypernuclei



Binding energies vs deformations



X.-R. Zhou, H.-J. Schulze, H. sagawa et.al, PRC76, 034312(2007)

Binding energies vs deformations



Superdeformation of hypernucei in SHF



X.-R. Zhou, E. Hiyama, in preparation

Density distribution of superdeformed hypernuclei



RMF

AMD





2. Energy spectra of hypernulcei

Limitations of extended SHF+BCS method:

- The SHF models can just give the single-particle energies and ground state of Λ hypernuclei in intrinsic frame of reference.
- The conservation of particle number is destroyed by BCS method.
- The study of the gamma spectra and electromagnetic transitions needs symmetry restoration.

Angular momentum and particle-number projection (AMP&PNP) are needed !

Projected SHF+BCS Model

The projected mean-field state

$$\left|\phi, JMK\right\rangle = P_{MK}^{J} P^{N} P^{Z} \left|\phi\right\rangle$$

The projection operator

$$P_{MK}^{J} = \frac{2J+1}{8\pi^{2}} \int d\Omega D_{MK}^{J}(\Omega) R(\Omega)$$
$$\hat{P}^{N} = \frac{1}{2\pi} \int e^{i\varphi(\hat{N}-N)} d\varphi.$$

Energies with angular momentum and E2 transitions

$$E_{K}^{J} = \frac{\langle JMK, \phi | H | \phi, JMK \rangle}{\langle JMK, \phi | \phi, JMK \rangle}$$
$$B(E\lambda, I_{i} \longrightarrow I_{f}) = \frac{1}{2I_{i} + 1} |\langle I_{f}K_{f} || \hat{Q}_{\lambda} |I_{i}K_{i} \rangle|^{2}$$

nergy potential surface of ${}^{12}C$ and ${}^{13}_{\Lambda}C$



ergy potential surface of 20 Ne and $^{21}_{\Lambda}$ Ne



nergy potential surface of ^{24}Mg and ^{25}Mg



nergy potential surface of 26 Si and $^{27}_{\Lambda}$ Si

 β =0.234(²⁶Si) β =0.221(²⁷_{Λ}Si)



nparison of NSC89(upper) and Skyrme-type(lower) Interactions















Energy spectra of ^{12}C and $~~^{13}_{\Lambda}\text{C}$



Energy spectra of ²⁰Ne and $^{21}_{\Lambda}$ Ne



Energy spectra of $~^{24}\text{Mg}$ and $~^{25}_{\Lambda}\text{Mg}$



Energy spectrum of ^{28}Si and $^{29}_{\Lambda}\text{Si}$



E2 transition probabilities

<i>B</i> (<i>E</i> 2)	¹² C	2	$^{13}_{\Lambda}\mathrm{C}$	²⁰ N	e	$^{21}_{\Lambda}{ m Ne}$	²⁴ M	g	$^{25}_{\Lambda} Mg$	²⁸ S	i	$^{29}_{\Lambda}{ m Si}$
$I_i \rightarrow I_f$	exp	cal	cal	exp	cal	cal	exp	cal	cal	exp	cal	cal
$+(\frac{3^{+}/2 \rightarrow 1^{+}/2}{5^{+}/2 \rightarrow 1^{+}/2})$	4.65(26)	6.42	$\frac{5.61}{5.69}$	20.3(10)	19.49	$\frac{17.72}{17.72}$	21.5(10)	24.30	$\tfrac{22.91}{23.14}$	13.2(5)	14.23	$\tfrac{13.89}{13.61}$
$+\left(\frac{7^{+}/2 \rightarrow 3^{+}/2}{9^{+}/2 \rightarrow 5^{+}/2}\right)$		10.11	$\frac{8.23}{8.84}$	22.0(2)	28.39	$\frac{23.09}{25.73}$	39.0(4)	35.45	$\frac{30.53}{33.94}$	16.4(18)	21.46	$\tfrac{18.60}{20.88}$
$+(\frac{11^{+}/2 \rightarrow 7^{+}/2}{13^{+}/2 \rightarrow 9^{+}/2})$	—		_	20.0(3)	32.52	$\tfrac{28.45}{29.46}$	—	41.15	$\frac{37.52}{39.22}$	10.6(10)	25.75	$\tfrac{24.09}{24.86}$

 Exp B(E2) values come from http://www.nndc.bnl.gov

evaluate the change of size the following quantity is introduced .

$$_{B} = \frac{B(E2, I_{ic}^{+} \rightarrow I_{fc}^{+})_{H}}{B(E2, I_{i}^{+} \rightarrow I_{f}^{+})}$$

TABLE III. Calculated Γ_B values.

Γ_B	¹² C	²⁰ Ne	²⁴ Mg	²⁸ Si
$2^+ \rightarrow 0^+$	0.885	0.903	0.957	0.964
$4^+ { ightarrow} 2^+$	0.898	0.913	0.955	0.973
$6^+ \rightarrow 4^+$	_	0.908	0.956	0.970





- 1.The DSHF was extended to hypernuclei by including a microscopically derived hyperon-nucleon interaction.
- 2.The calculated core nuclei and the corresponding hypernuclei have similar deformations with the same sign when the core nuclei are well deformed. The superdeformation of hypernuclei is found in SHF.
- 3. The projected SHF+BCS model gives reasonable results of energy spectra and E2 transition rates for well-deformed sd-shell nuclei and hypernuclei.



Cooperators

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