

HH pair searches

combination with $wwyy$, $bbyy$, $bb\tau\tau$, $bbbb$

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and people from $bb\tau\tau$, $bbbb$, $bbyy$

CDS entry

[$wwyy$]<https://cds.cern.ch/record/1967498>

[$bb\tau\tau$]<https://cds.cern.ch/record/1967500>

[combination]<https://cds.cern.ch/record/1984111/>

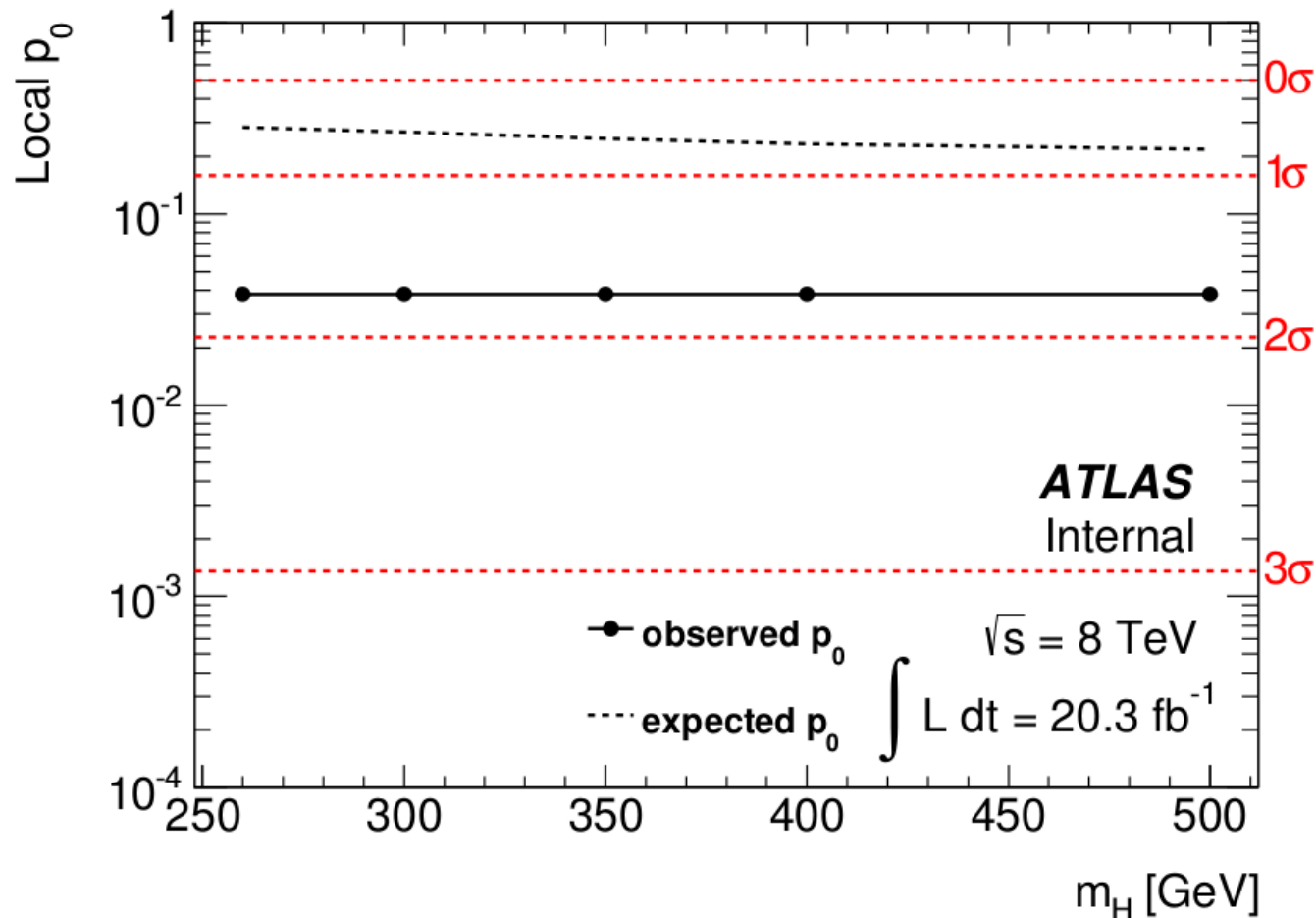
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17-04-2015
IHEP

Higgs approval:

<https://indico.cern.ch/event/387805/>

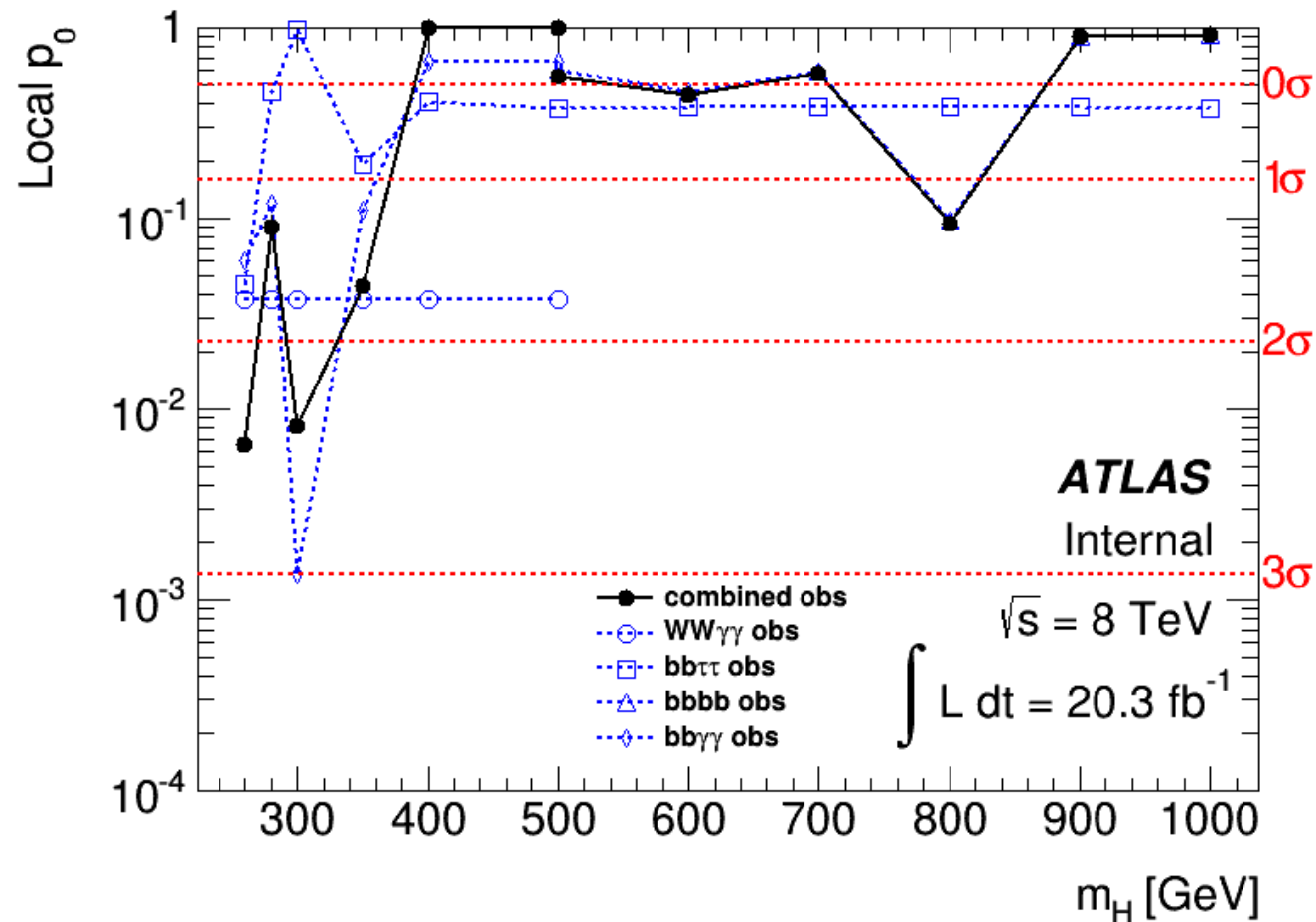
wwyy p0 updates [wwyy int note]

- inject 1pb signal for each mass before any cuts
- due to an increasing signal efficiency as a function of mass, the expected p0 decreases along the mass



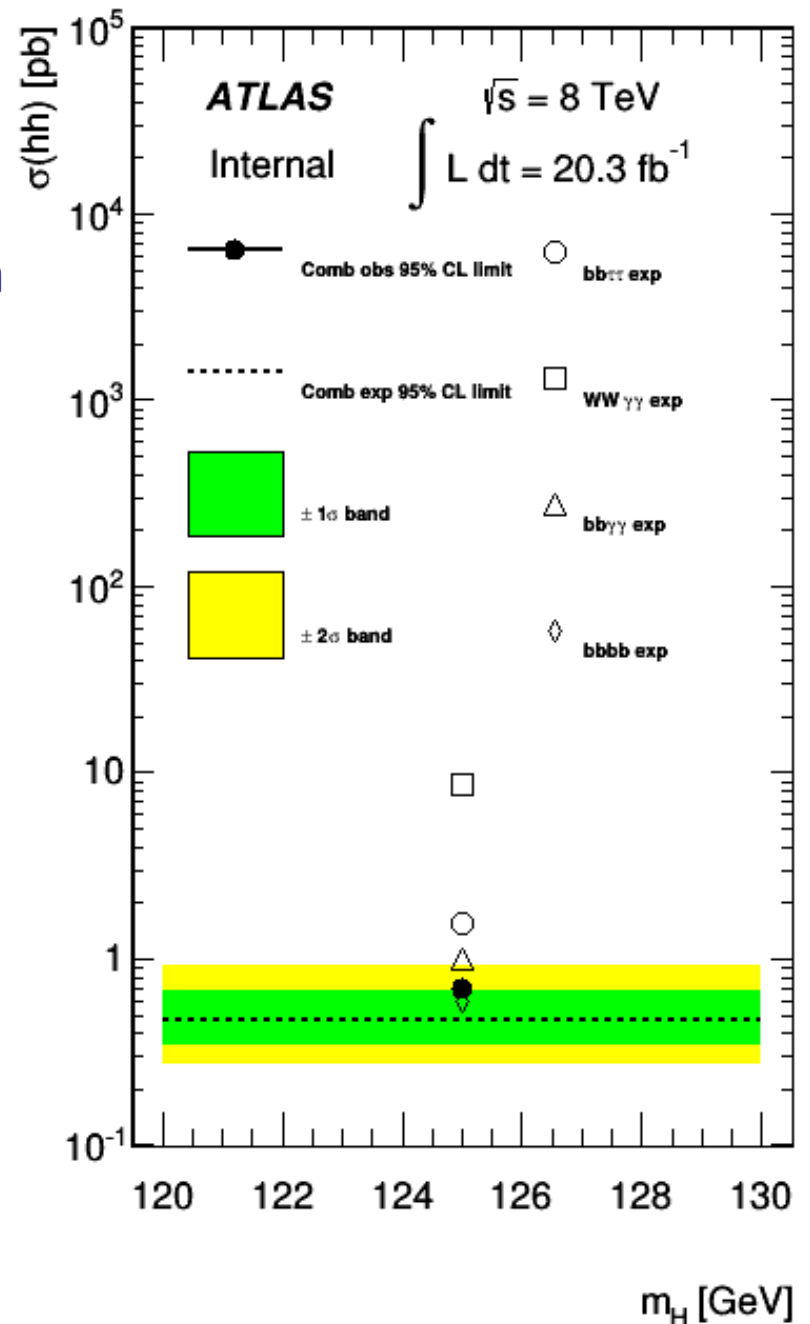
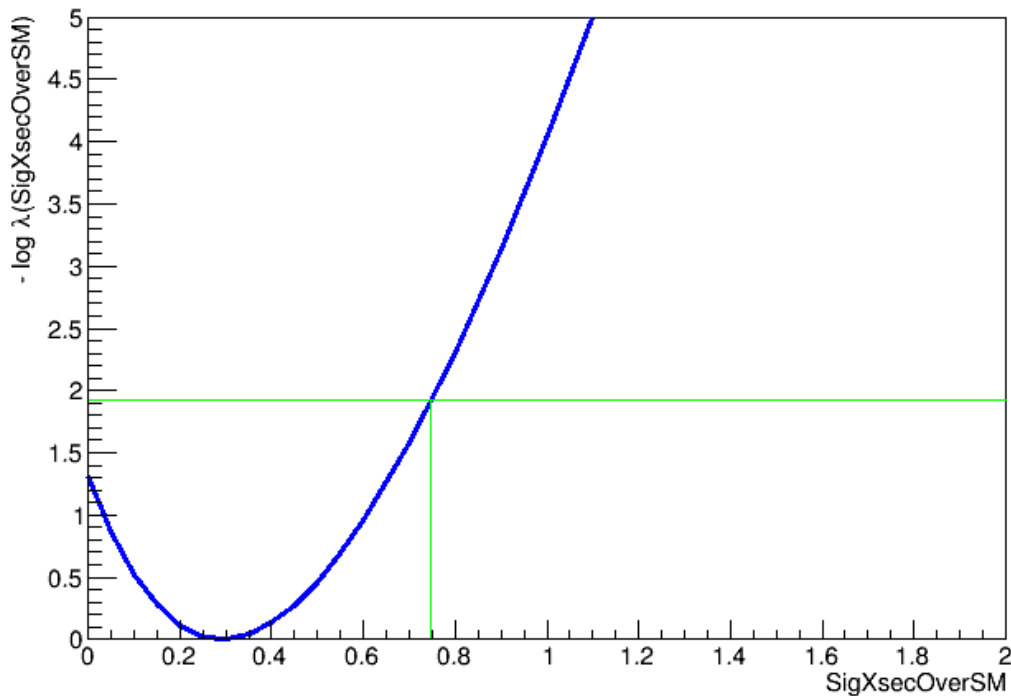
comb p0 updates [hhcomb int note]

- newly added mass point of **280**GeV
 - combining bbtatau+bbyy@280 and wwy@260
- 280 limitis also finished on grid, to-be-added



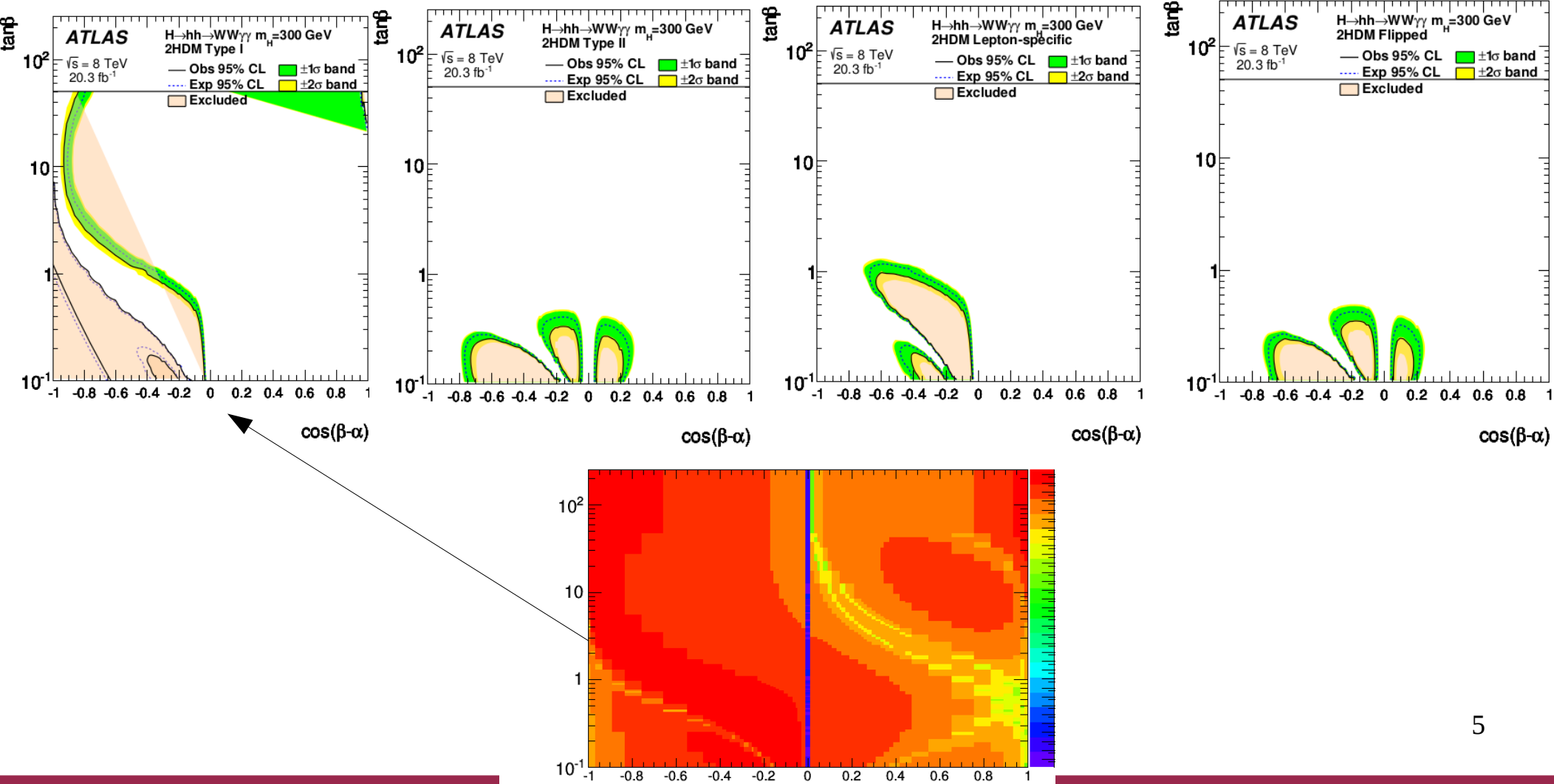
Combined limits [nonres]

- For non-resonance (toys), the contributions are mainly from $bbbb$, $bb\gamma\gamma$ followed by $bb\tau\tau$ and $ww\gamma\gamma$
- 9.55fb is predicted in SM hh production [arXiv:1309.6594] @NNLO
- obs limit is **72** times the SM hh
- obs mu = $x_{s_measured} / x_{s_SM} = \mathbf{30}$



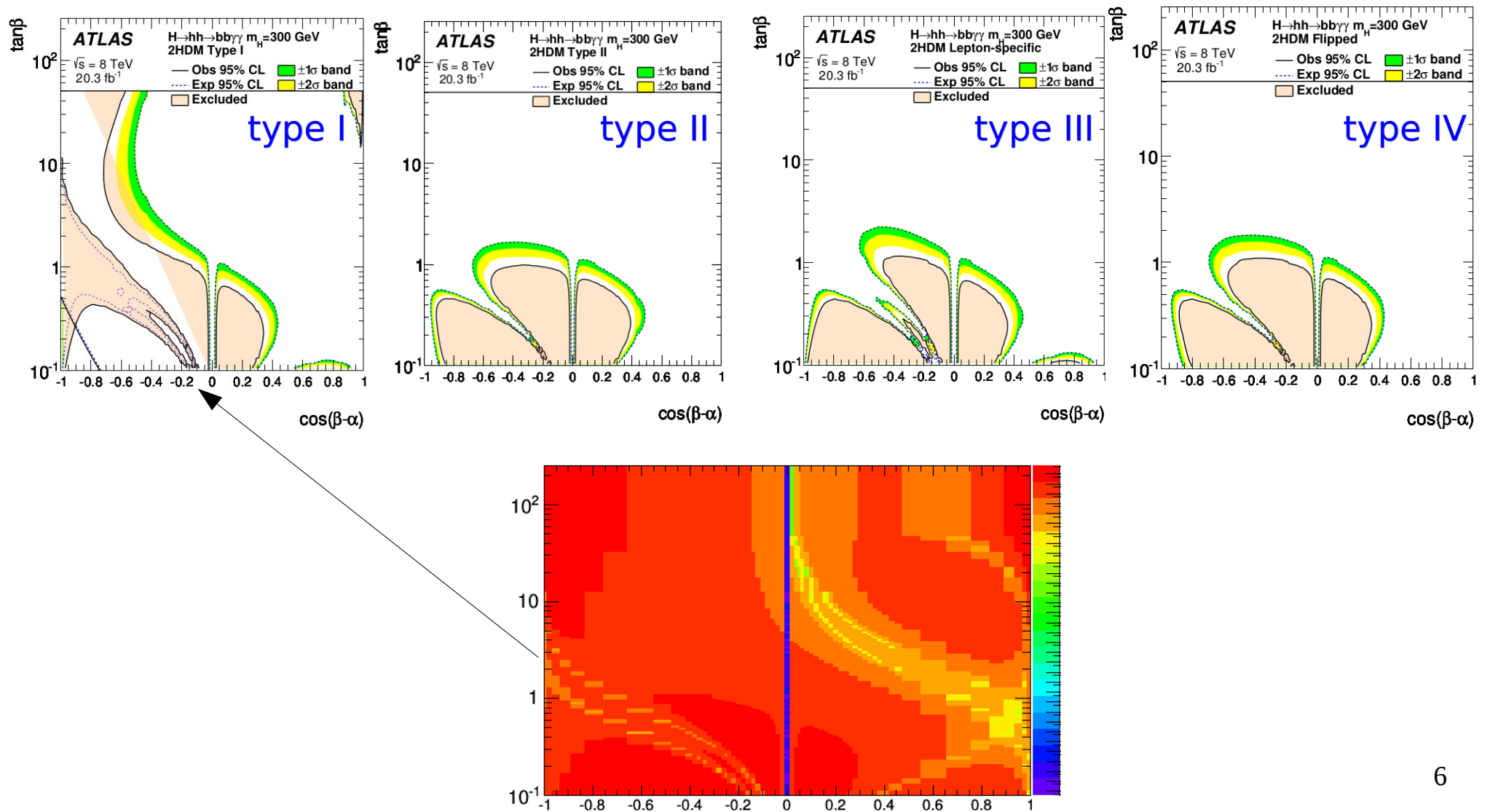
2HDM

- $w\bar{w}\gamma\gamma$ does not contribute a lot considering the limit setting in combination and the exclusion areas
- can simply ignore $w\bar{w}\gamma\gamma$ channel in 2HDM interpretation



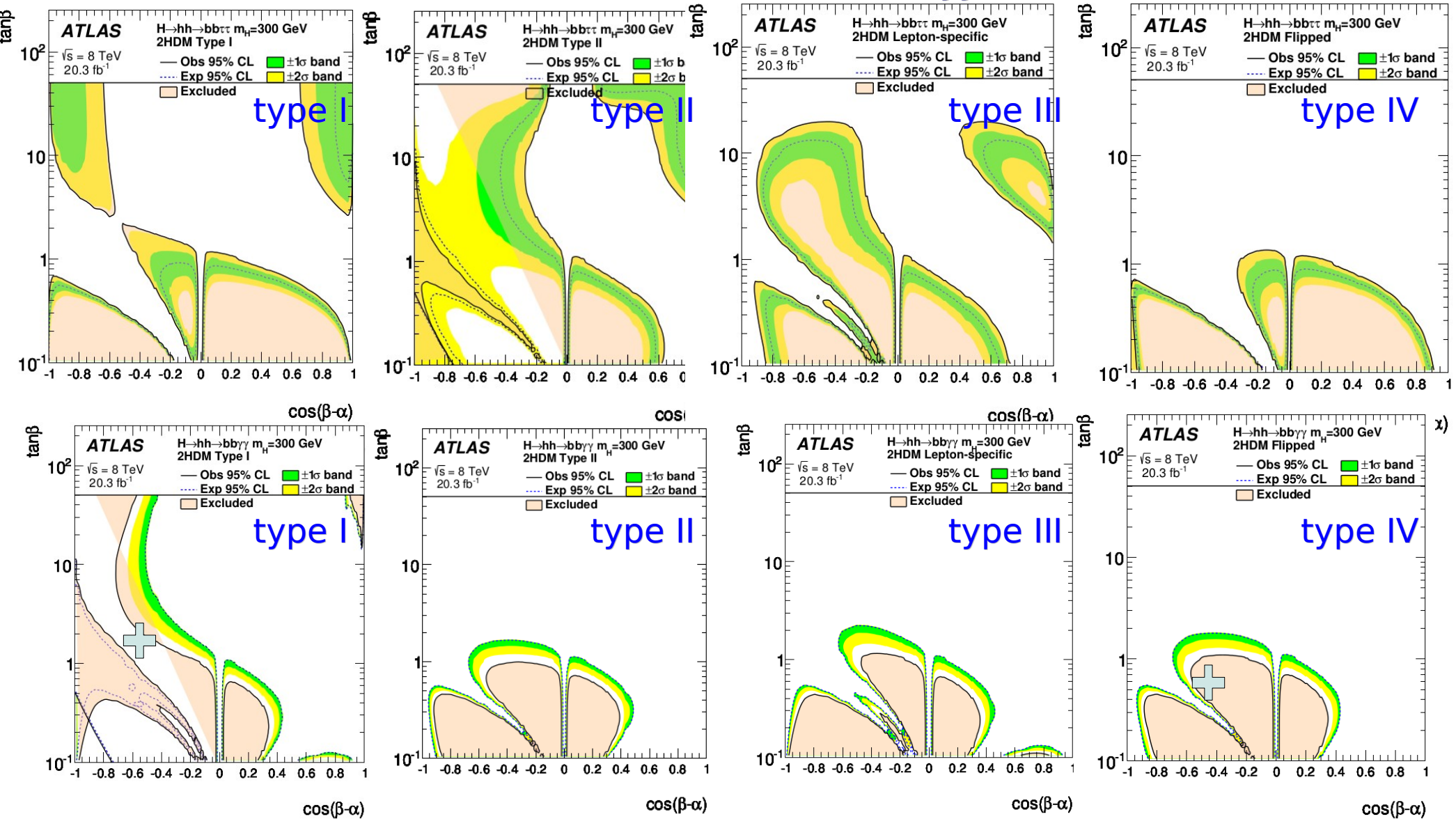
2HDM

- $b\bar{b}\gamma\gamma$



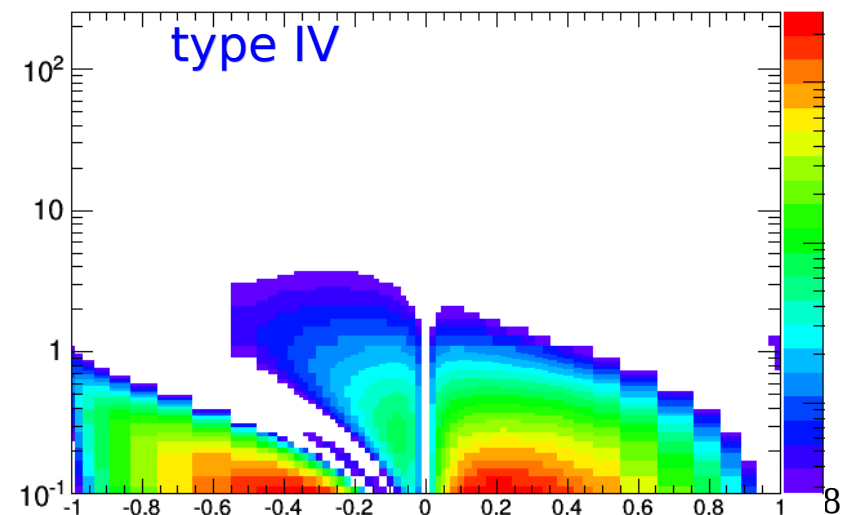
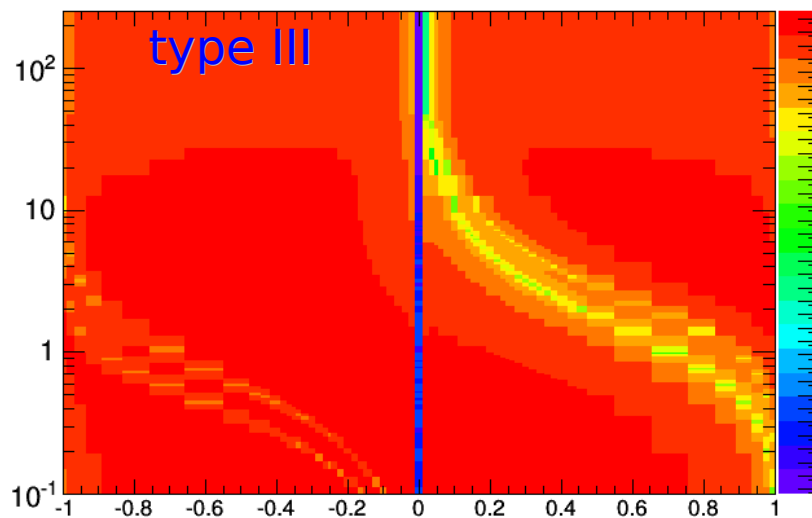
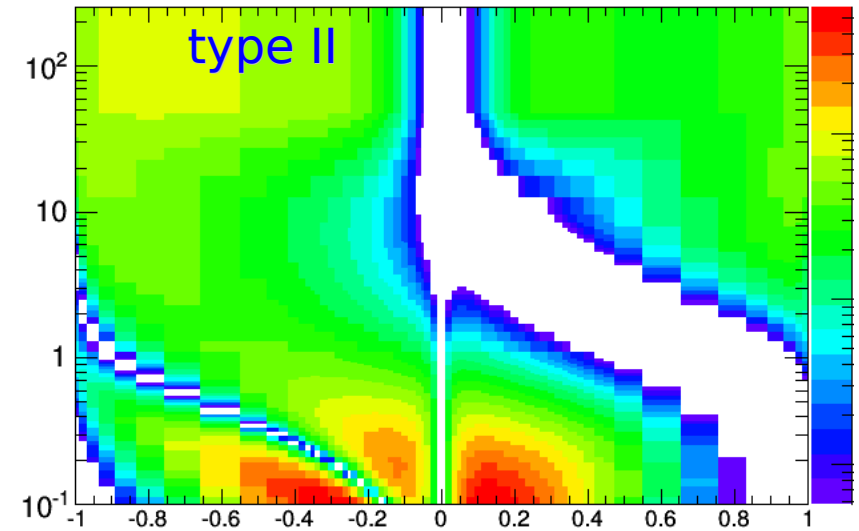
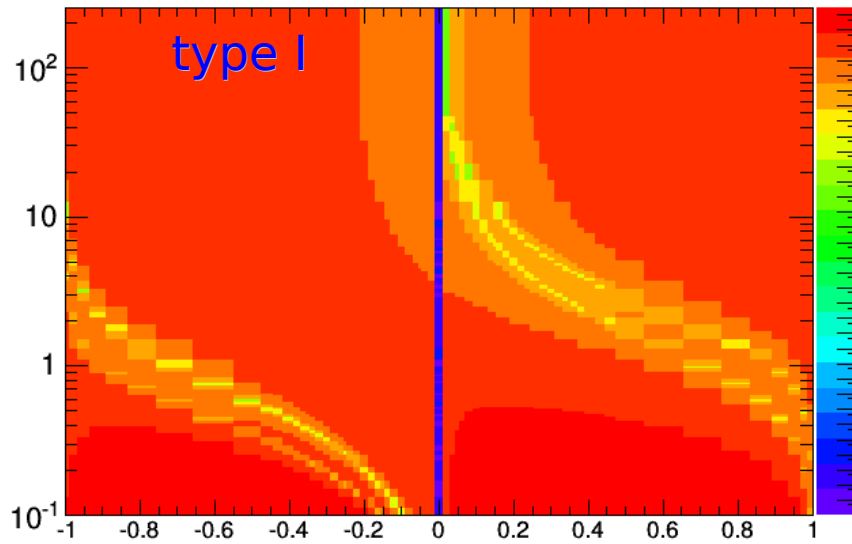
2HDM

- $b\bar{b}\tau\tau$ contributes a lot
- $b\bar{b}\tau\tau$ covers most of the exclusions from $b\bar{b}\gamma\gamma$



2HDM

- `bbautau XS*BR` contour check to make sure the exclusion plots are correctly produced in previous page



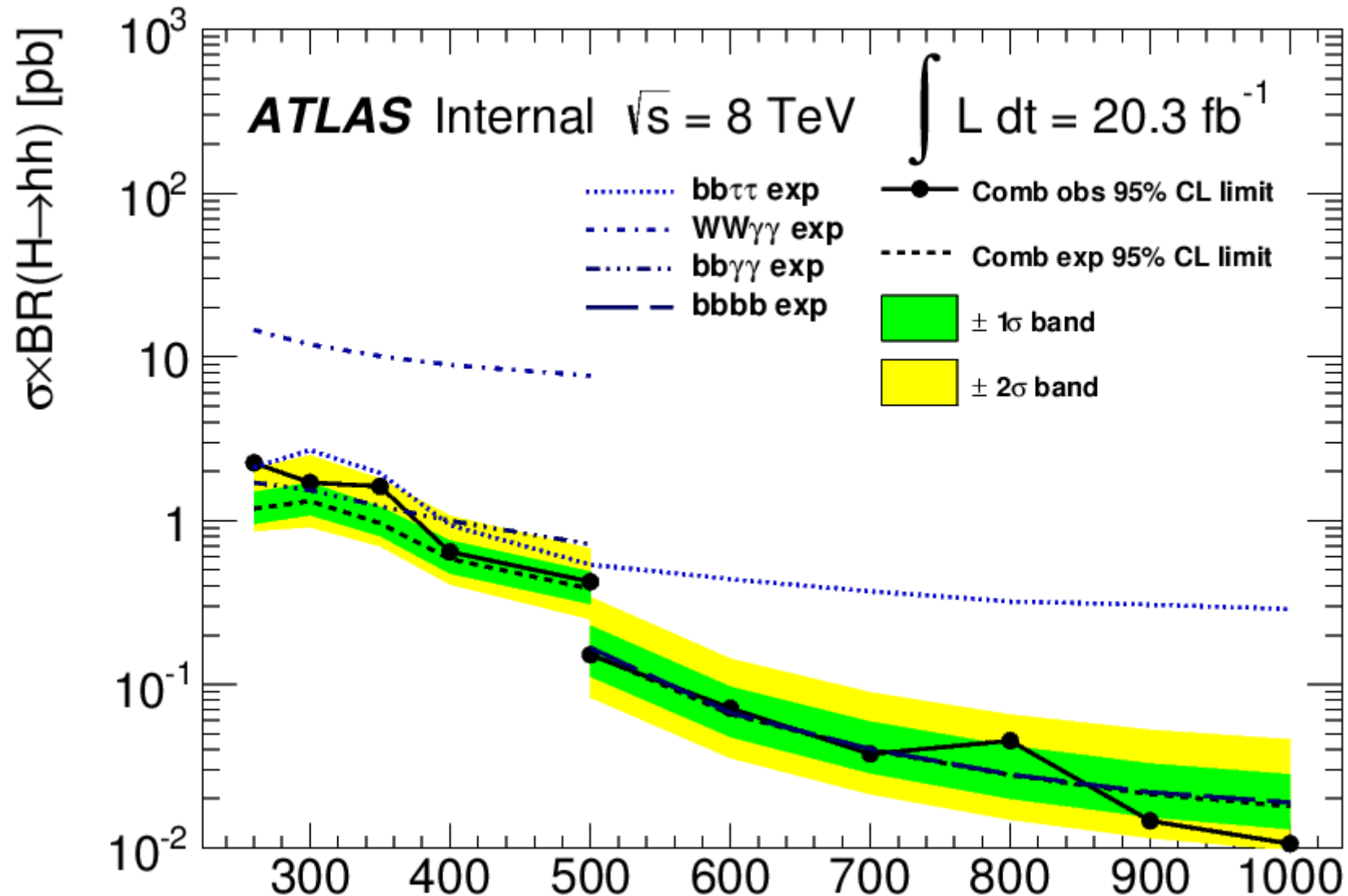
Summary

- interpolation in bbtatau
 - low mass region follow mass points of bbyy
 - high mass region, dominated by bbbb, needed?
- interpolation in wwyw can be simplified into an interpolation of signal efficiency as a function of mass, due to the same bkg estimates independent of mass
- 2HDM
 - 1. overlay exclusions from bbtb and bbyy in transparent style
 - 2. take into account of bbyy exclusions only if width < X%
- All CDS comments up to now have been answered
- a check on non-res limit with all systematics shutdown ... ongoing, any idea of switching off nuisance parameters from the workspace level?

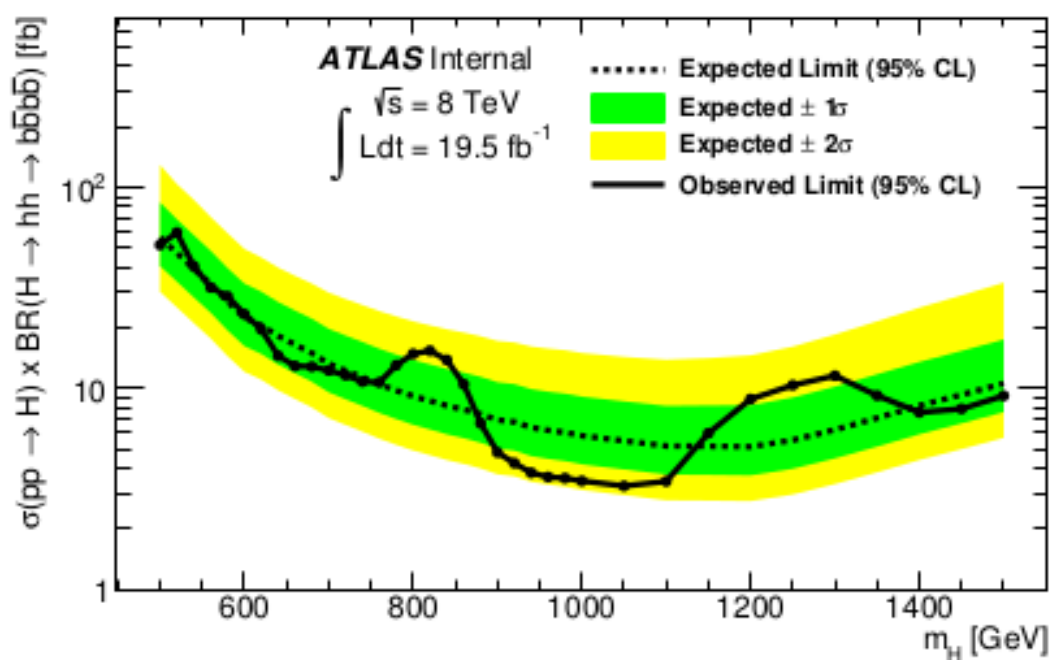
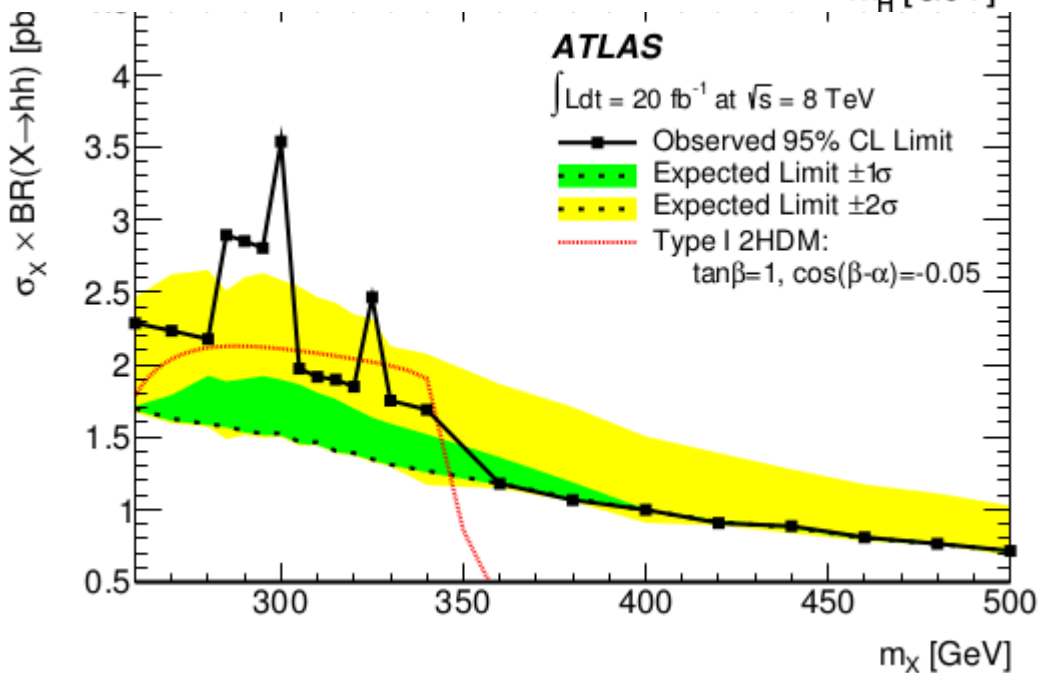
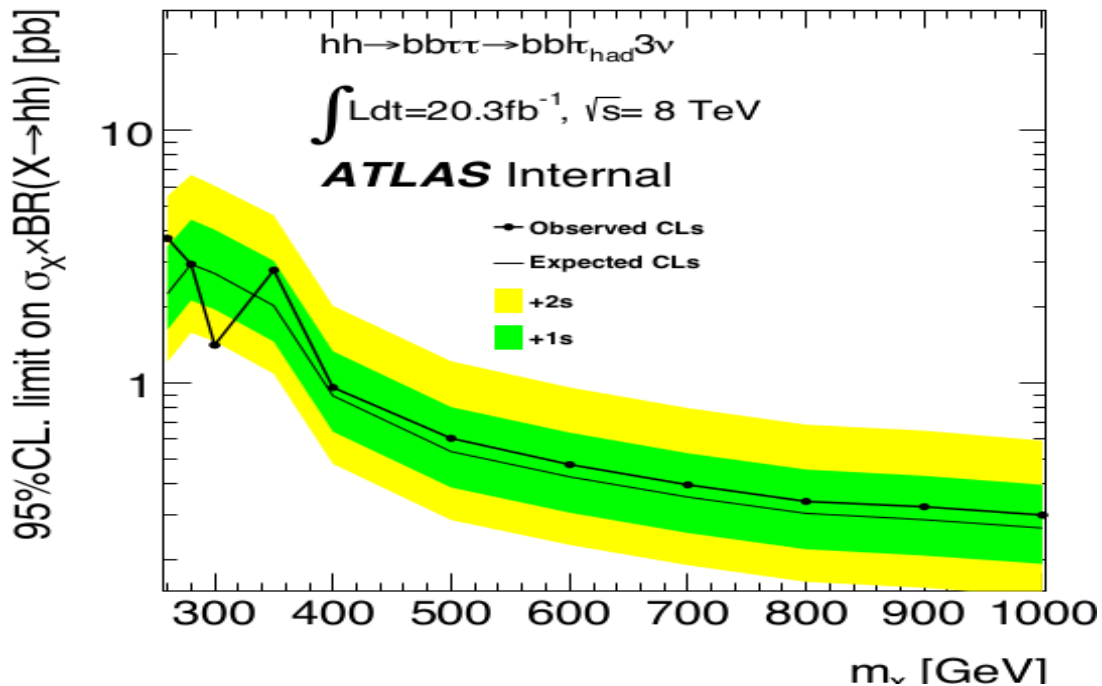
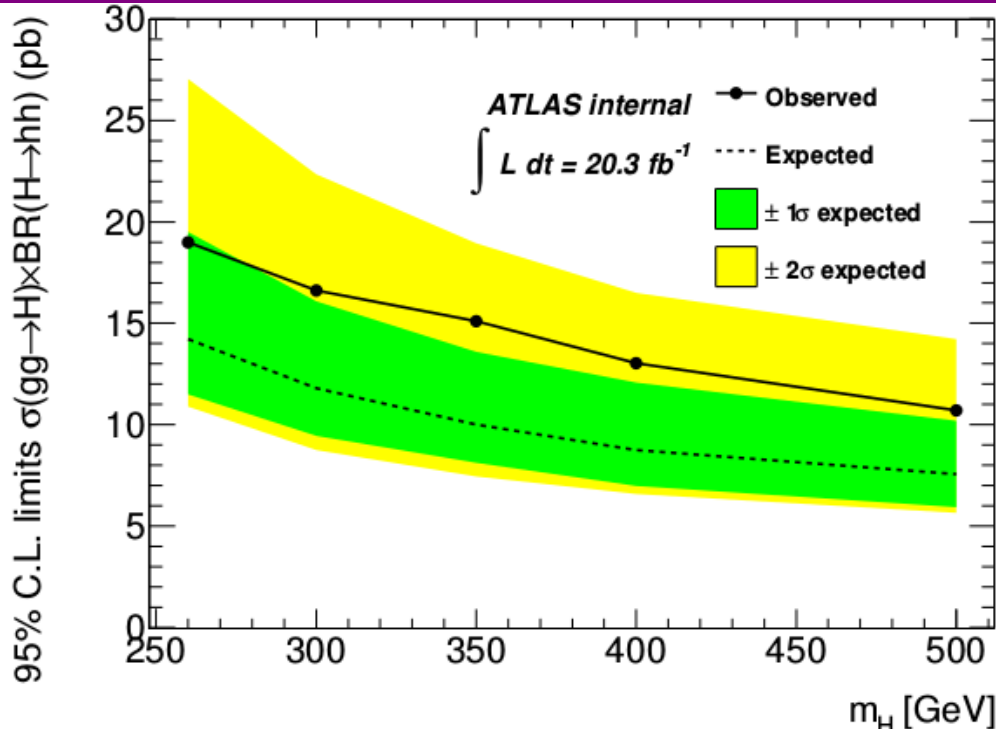
backup

comb observed limits [res]

- [260,500] from toys; [500,1000] from asymptotics
- more points **interpolated** in high mass region



sub-channel limits [res]



comb observed limits [nonres]

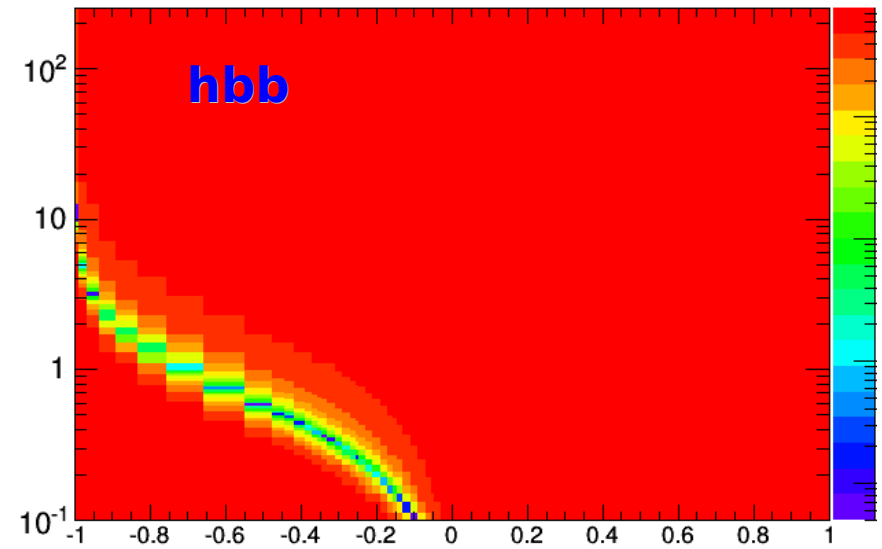
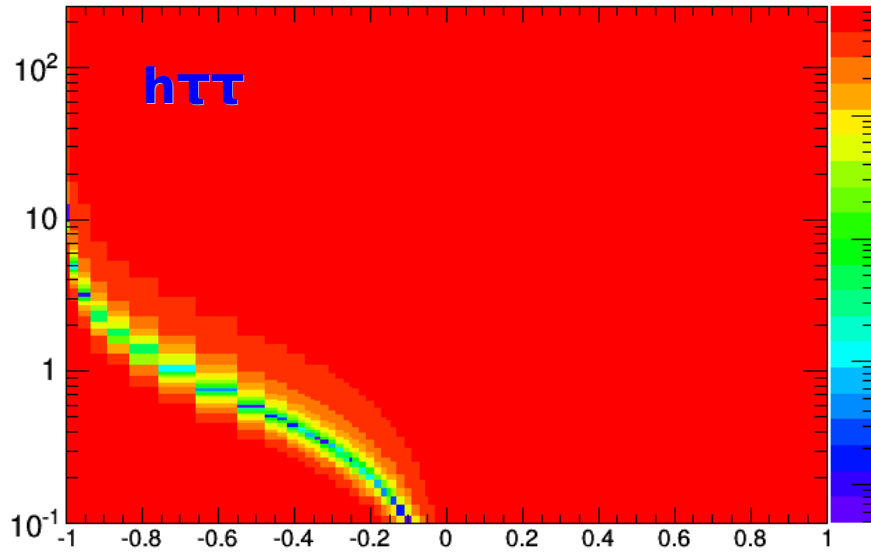
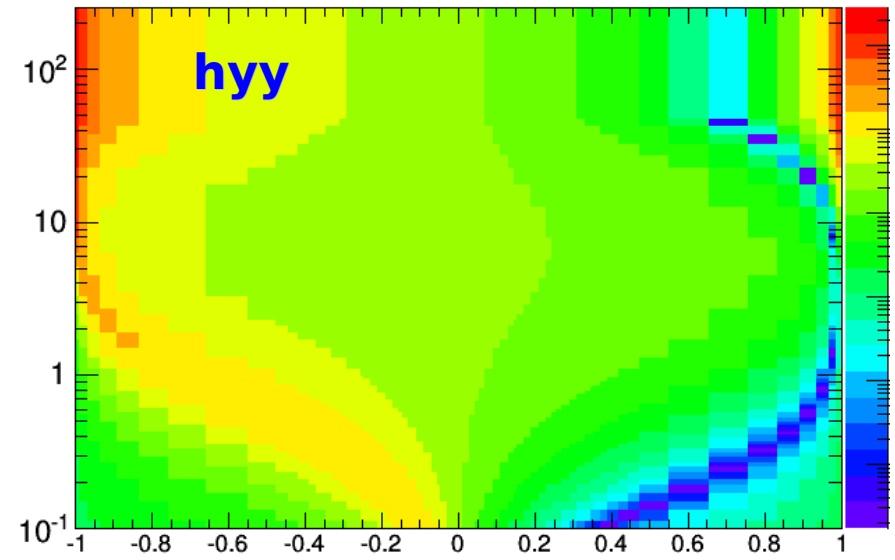
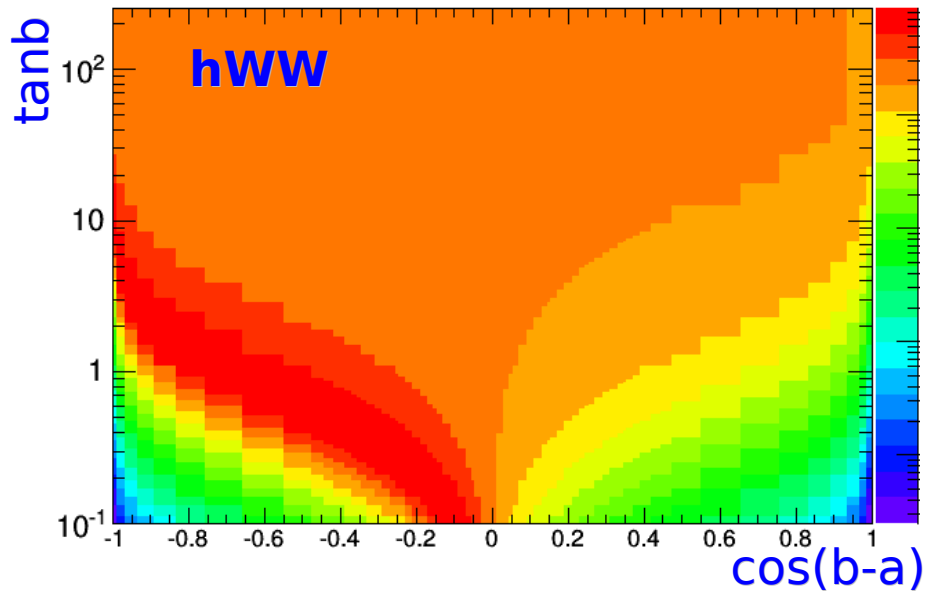
	OBS	EXP	+2sig	+1sig	-1sig	-2sig
bbbb	<u>0.59388</u>	0.59457	1.22329	0.85430	0.42842	0.31912
bbyy	2.20000	1.00546	2.39049	1.52719	0.72449	0.53966
wwyy	12.6259	8.81794	16.297	11.9078	7.10667	6.70651
bb $\tau\tau$	1.29582	1.53934	3.32687	2.26839	1.10918	0.82620
combine	0.69188	0.47370	0.93053	0.66888	0.34834	0.27022

Hhh interpretation

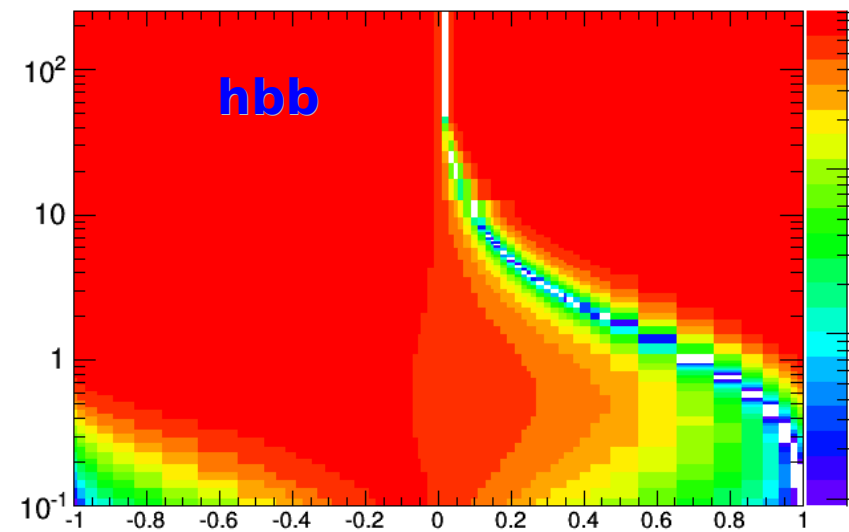
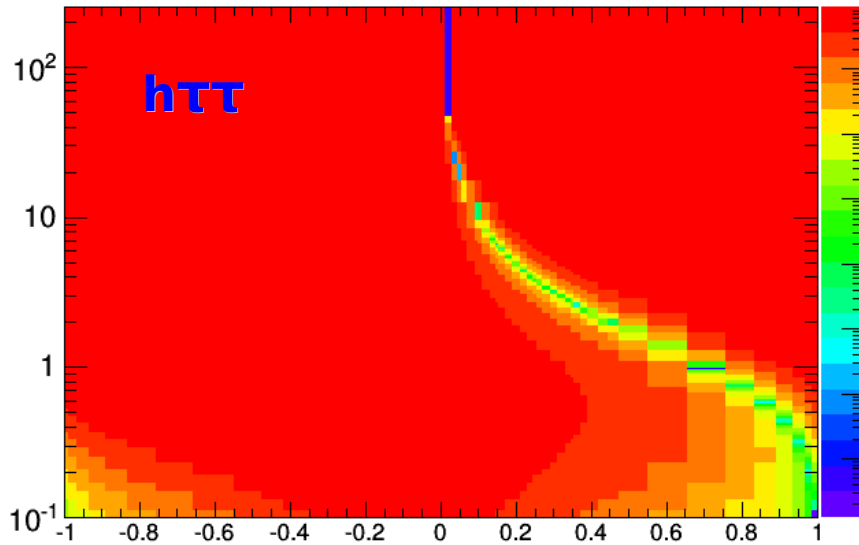
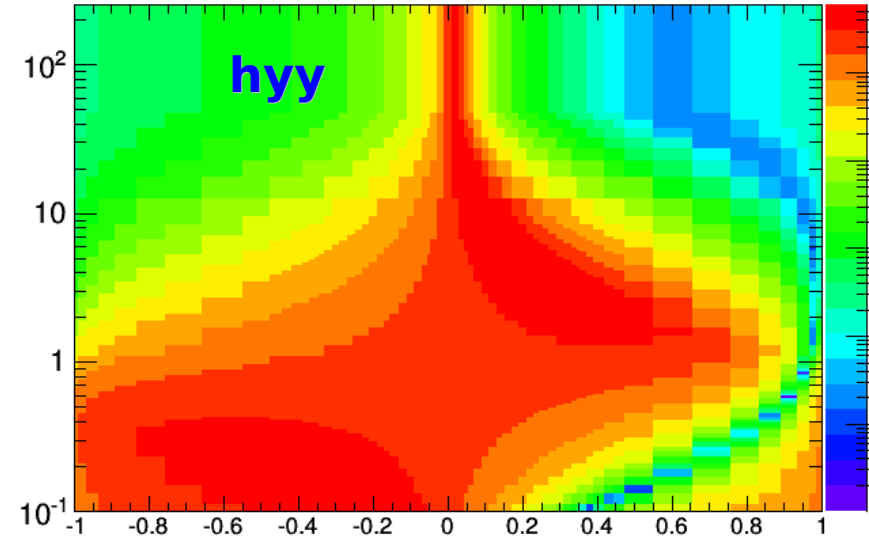
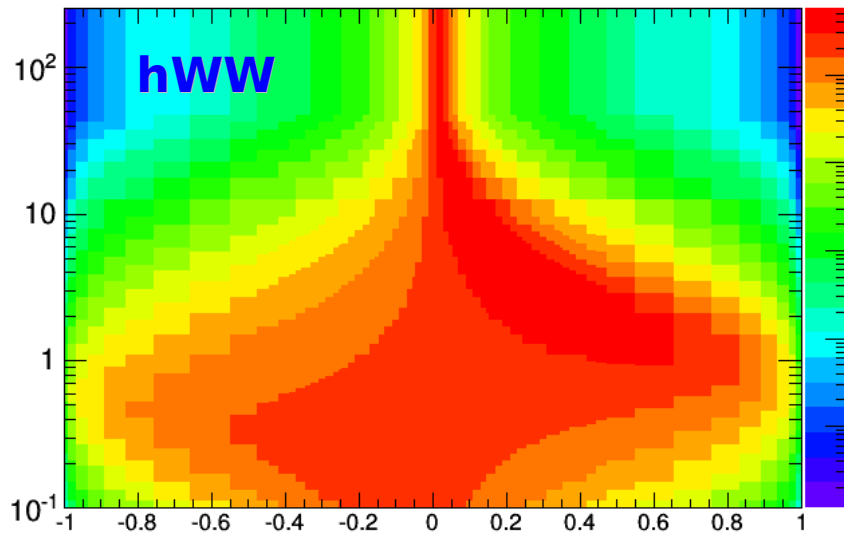
- 4 channels are included
 - res-low-mass: $b\bar{b}y\bar{y}, b\bar{b}\tau\bar{\tau}, w\bar{w}y\bar{y} \sim$ for example 300GeV
 - res-high-mass: $b\bar{b}b\bar{b}, b\bar{b}\tau\bar{\tau}$
- open discussion

	Type I	Type II	Type III	Type IV
ξ^V_{Sh}	$\sin(\beta - \alpha)$	$\sin(\beta - \alpha)$	$\sin(\beta - \alpha)$	$\sin(\beta - \alpha)$
ξ^u_{Sh}	$\frac{\cos \alpha}{\sin \beta}$	$\frac{\cos \alpha}{\sin \beta}$	$\frac{\cos \alpha}{\sin \beta}$	$\frac{\cos \alpha}{\sin \beta}$
ξ^d_{Sh}	$\frac{\cos \alpha}{\sin \beta}$	$-\frac{\sin \alpha}{\cos \beta}$	$\frac{\cos \alpha}{\sin \beta}$	$-\frac{\sin \alpha}{\cos \beta}$
ξ^l_{Sh}	$\frac{\cos \alpha}{\sin \beta}$	$-\frac{\sin \alpha}{\cos \beta}$	$-\frac{\sin \alpha}{\cos \beta}$	$\frac{\cos \alpha}{\cos \beta}$

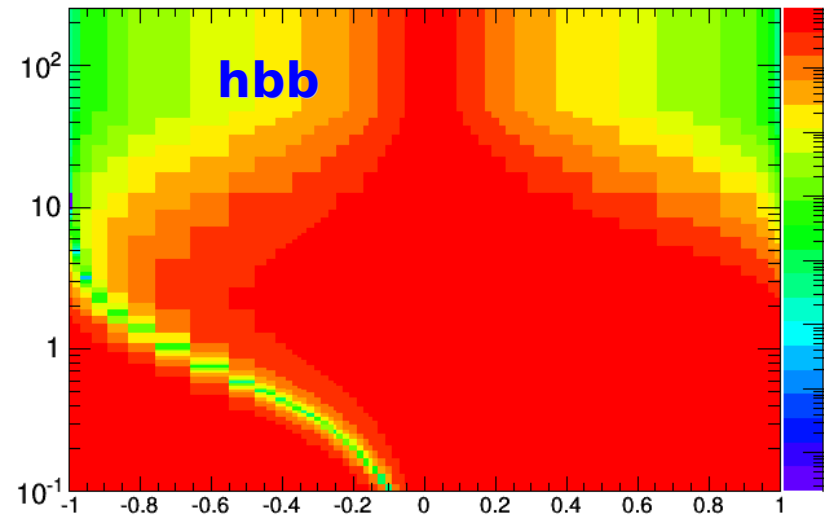
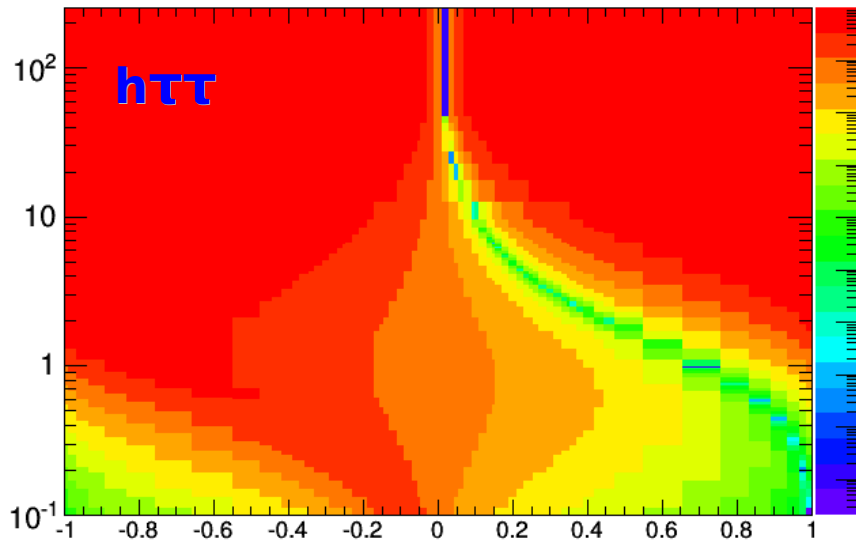
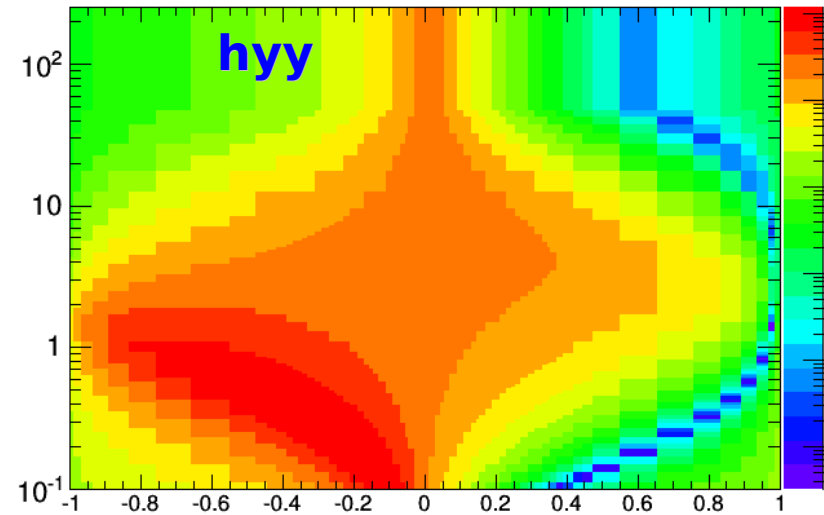
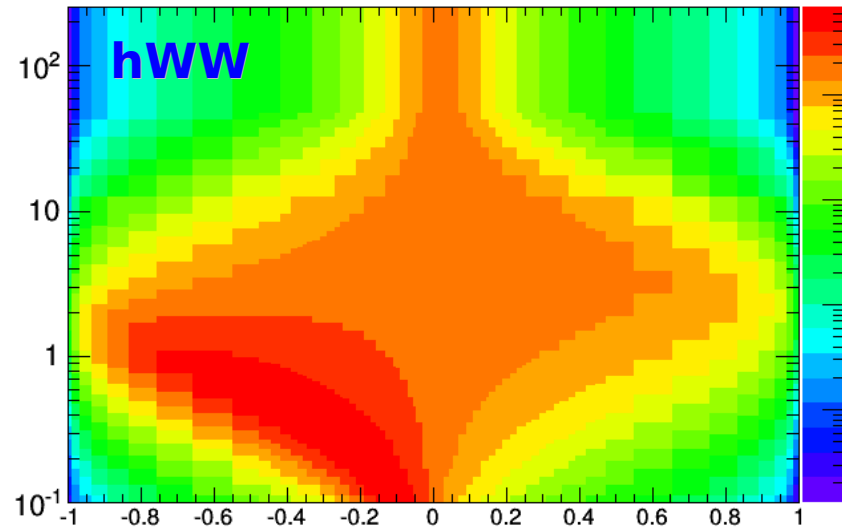
2HDM type I @ $m_H = 300$ GeV



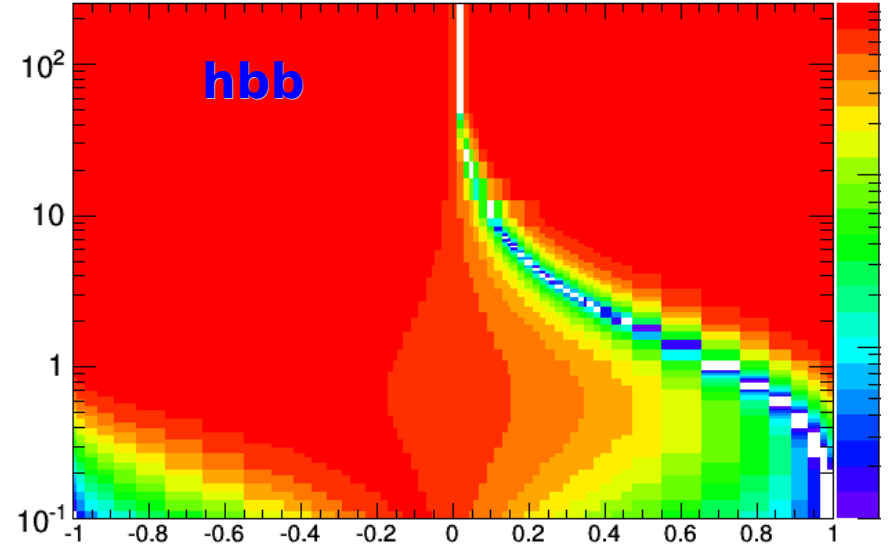
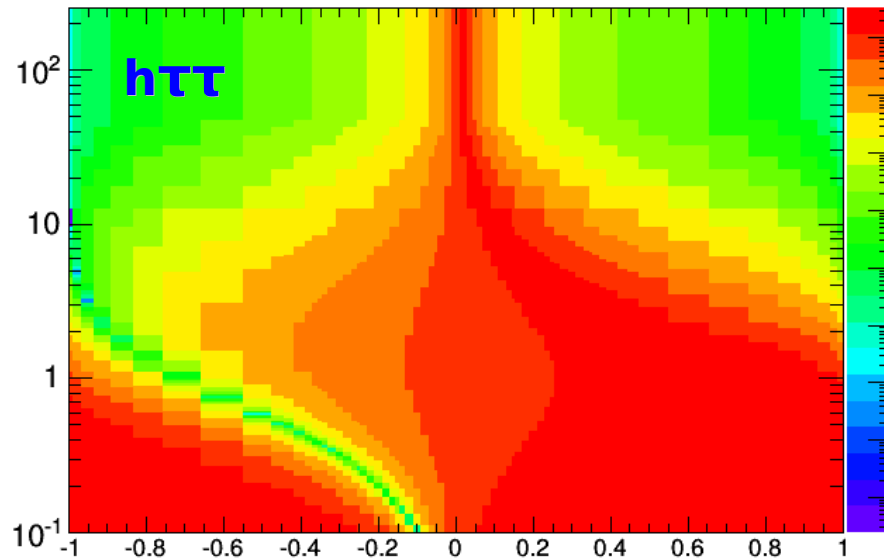
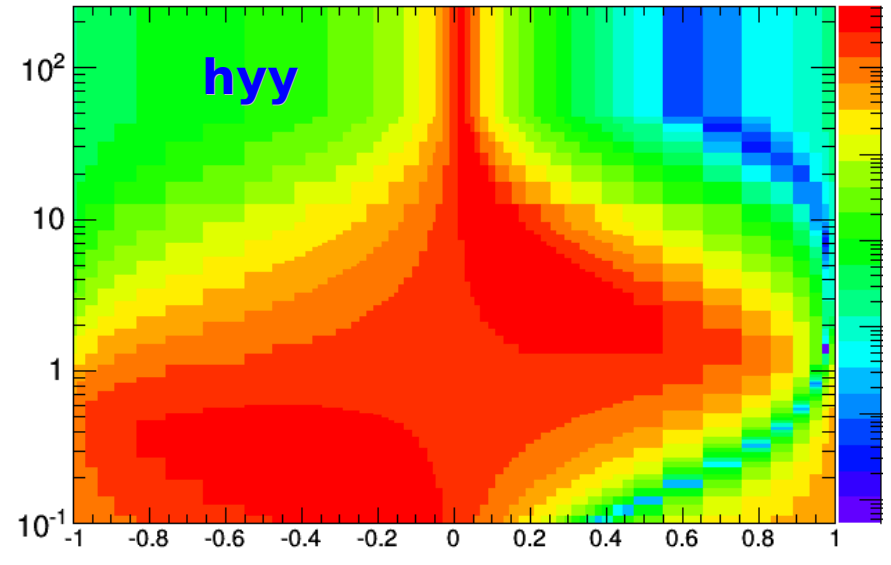
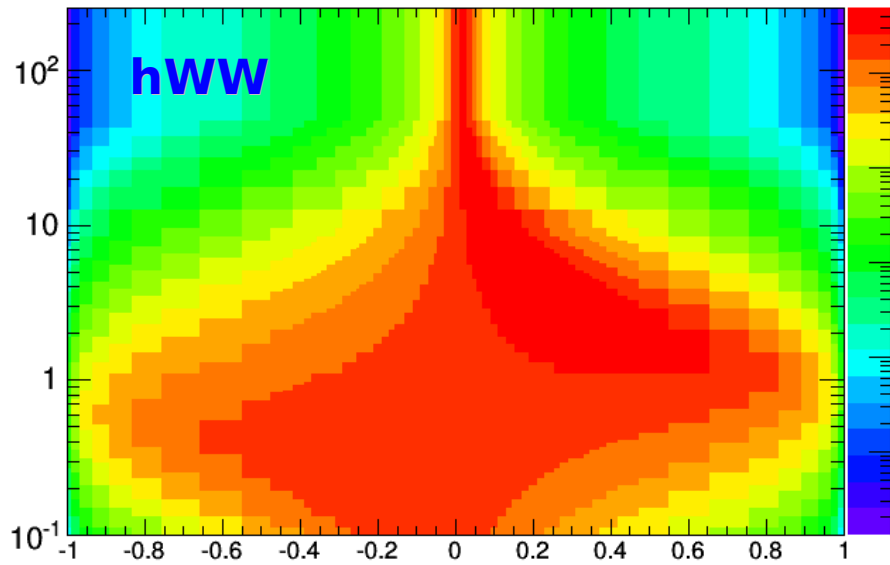
2HDM type II @ $m_H = 300$ GeV



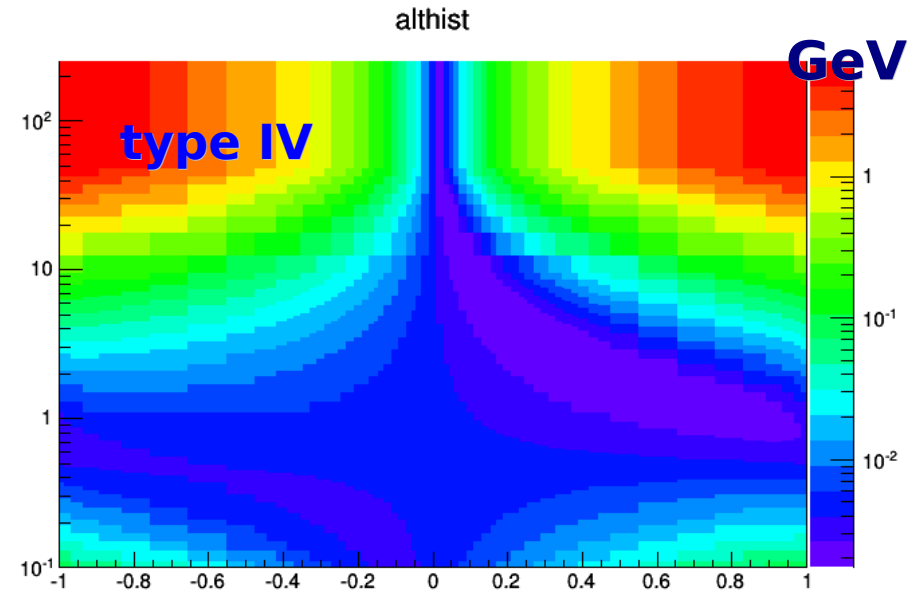
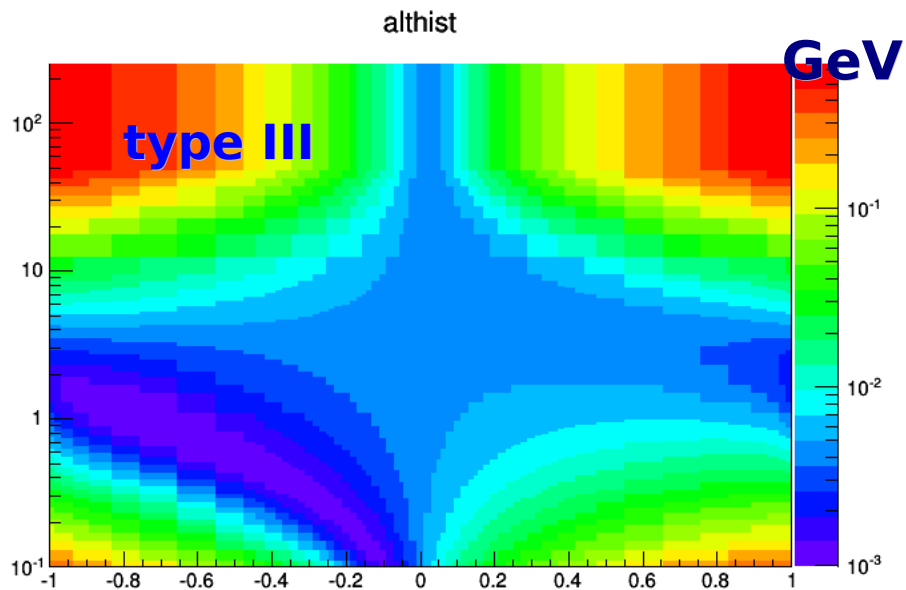
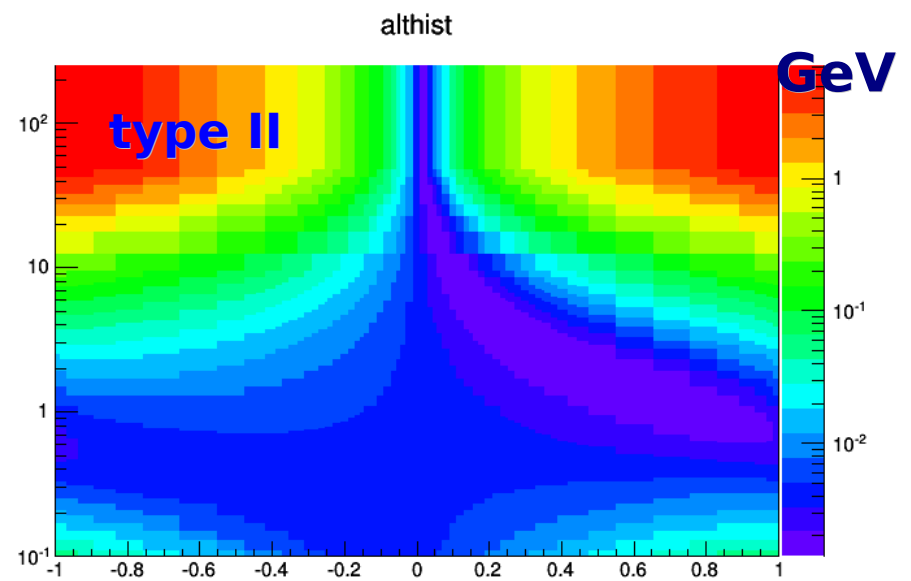
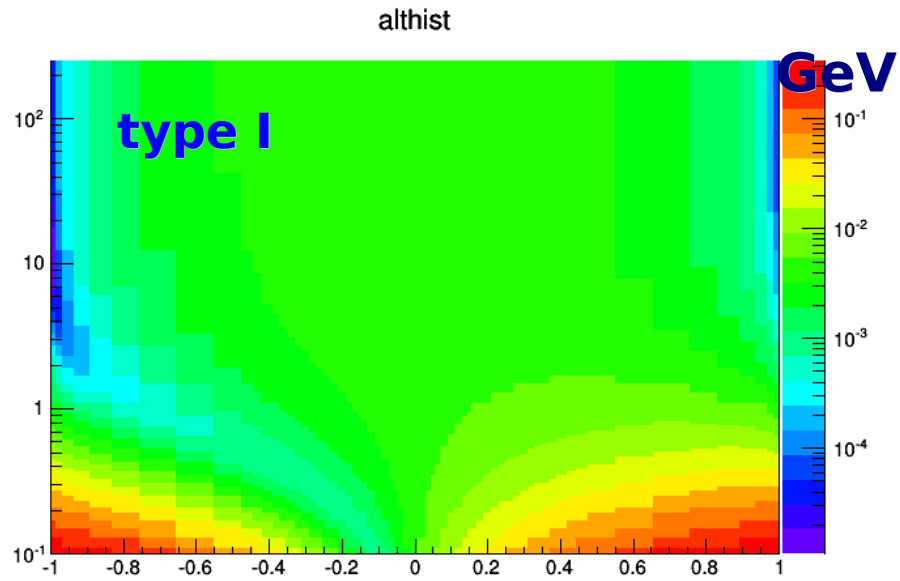
2HDM type III @ $m_H = 300$ GeV



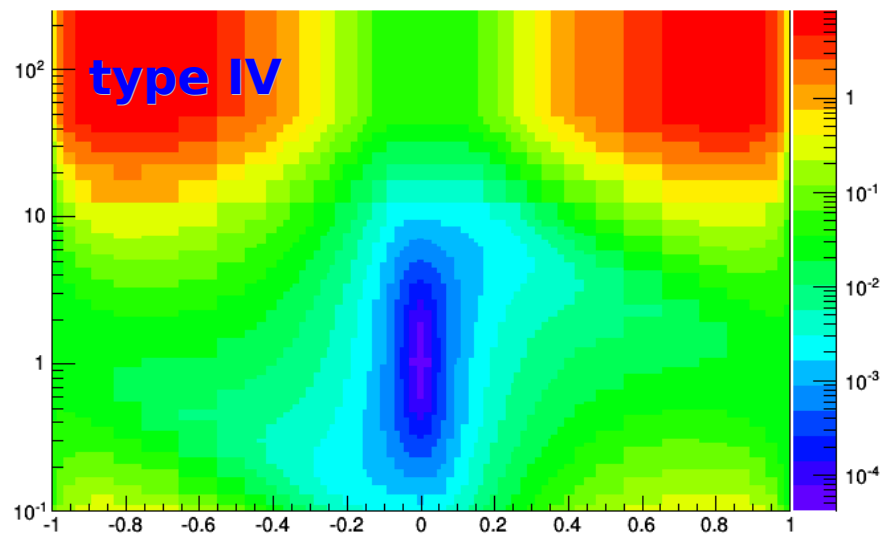
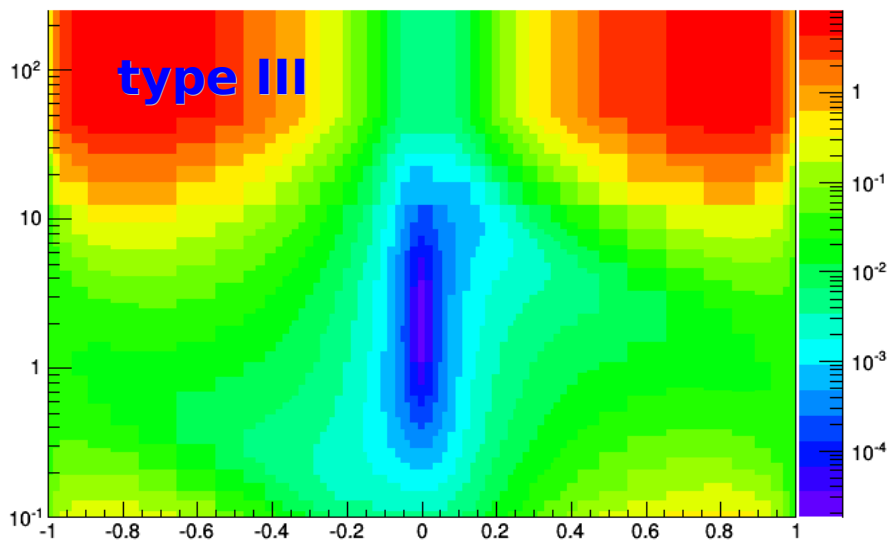
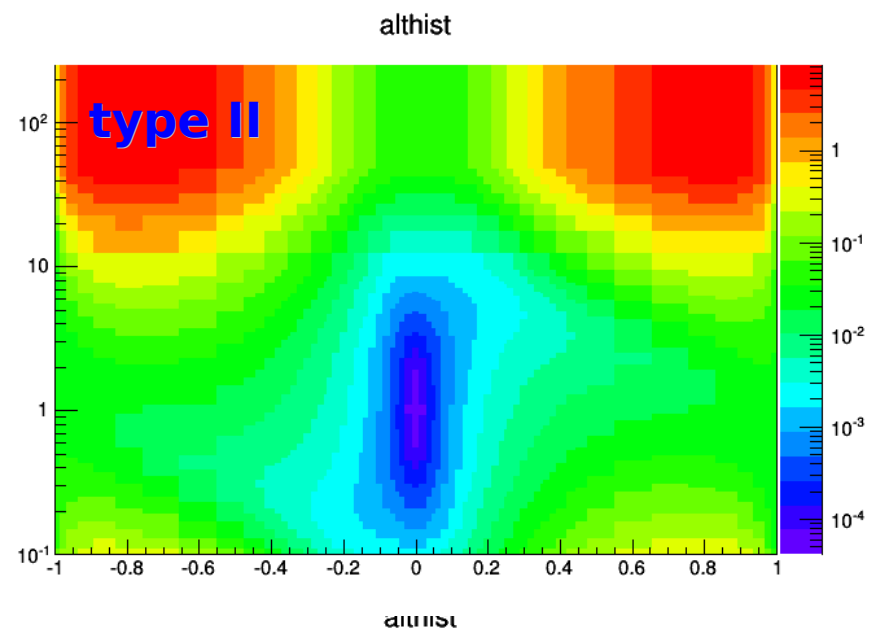
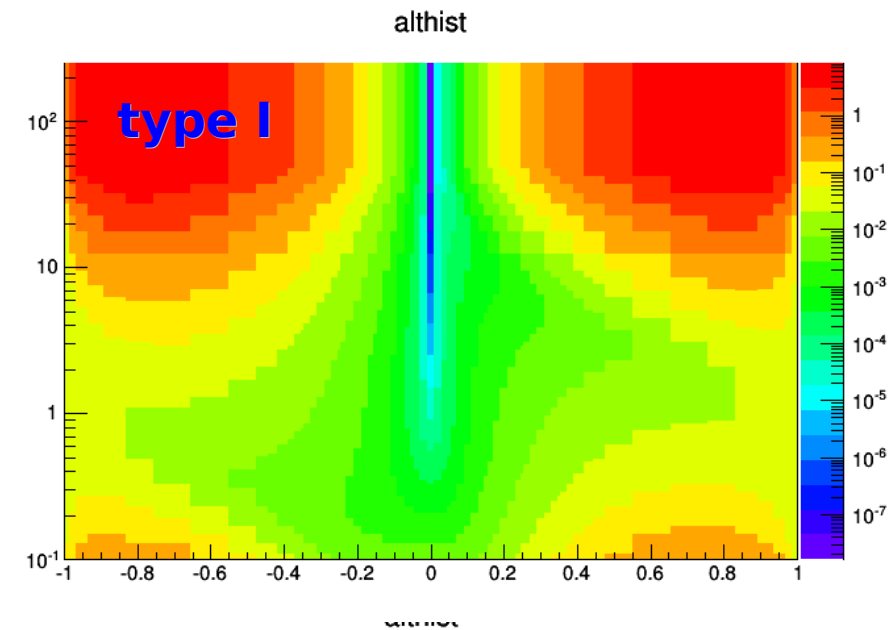
2HDM type IV @ $m_H = 300$ GeV



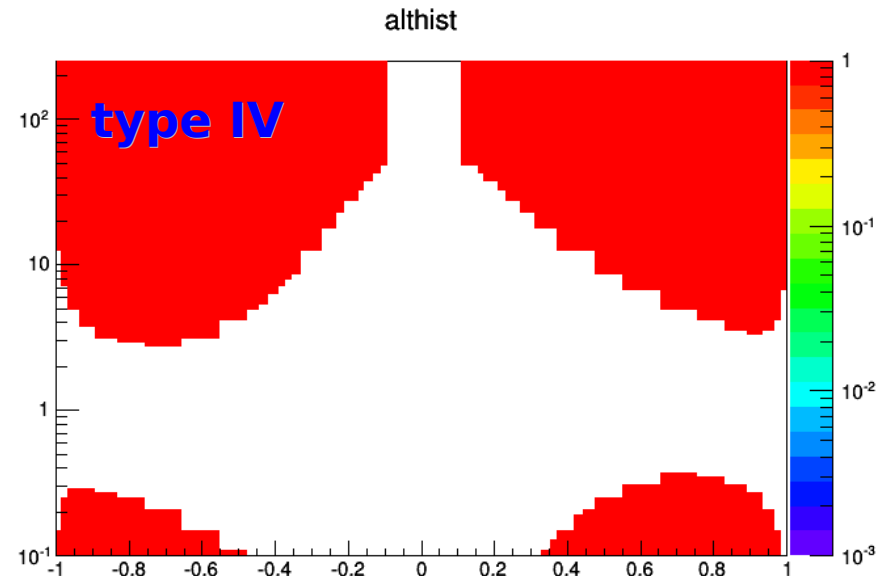
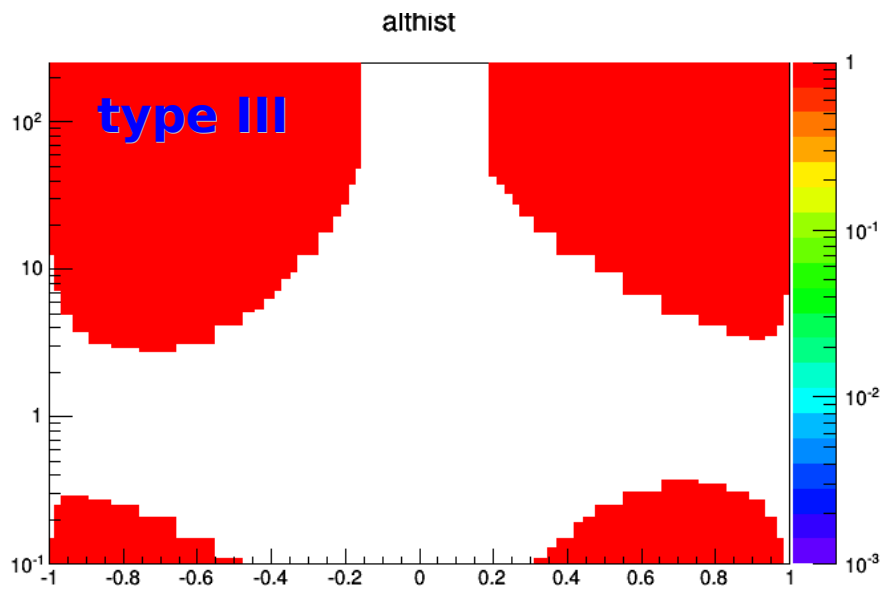
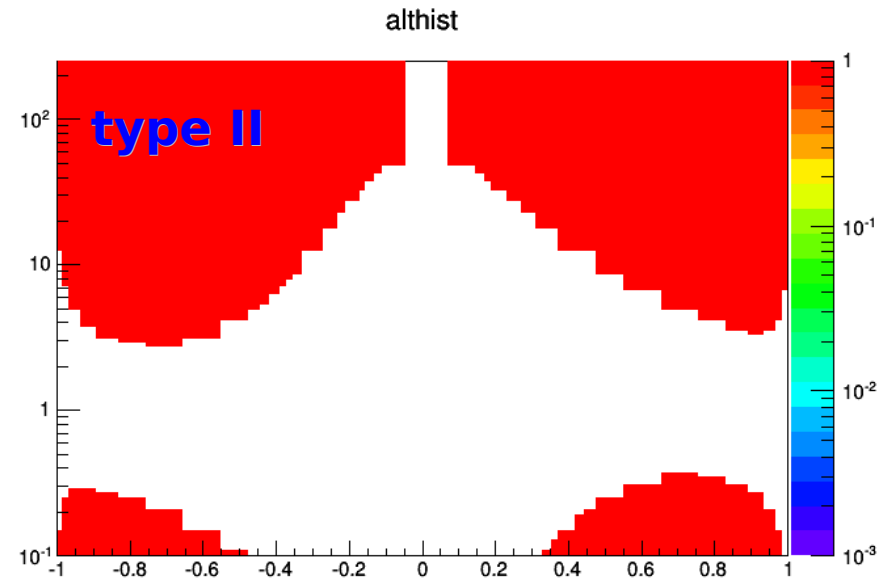
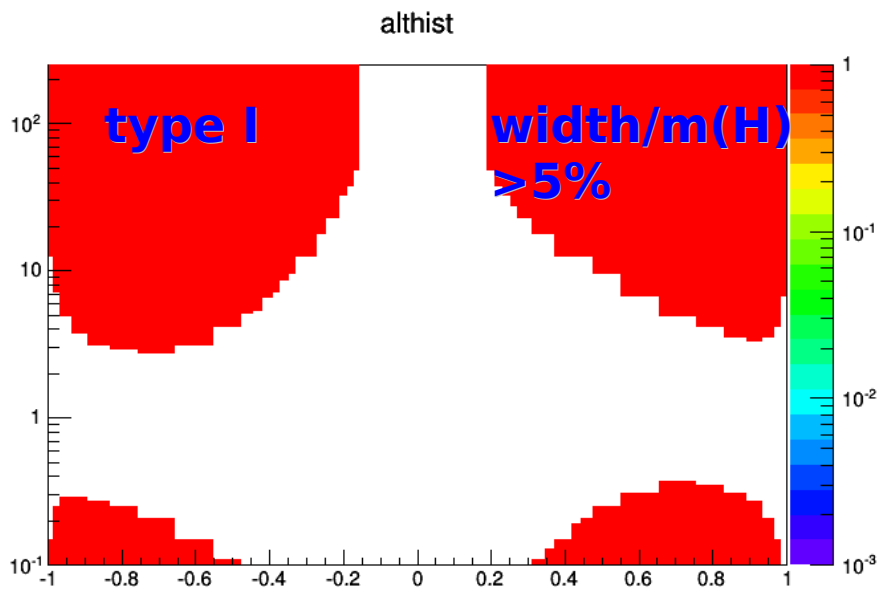
2HDM with h @ $m_H = 300$ GeV



2HDM witdh_H / mH @ 300 GeV

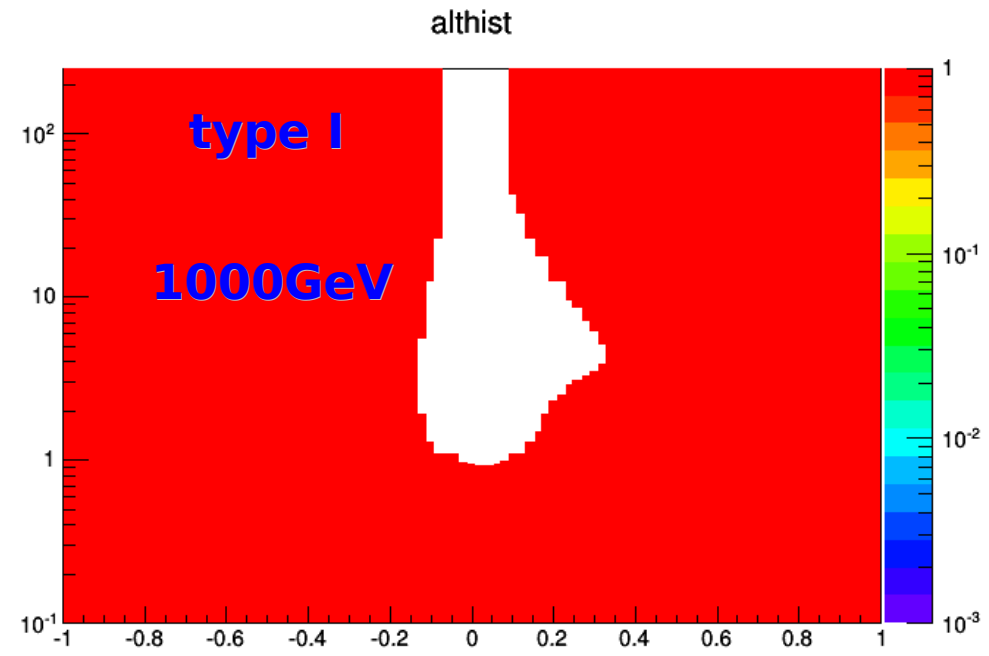
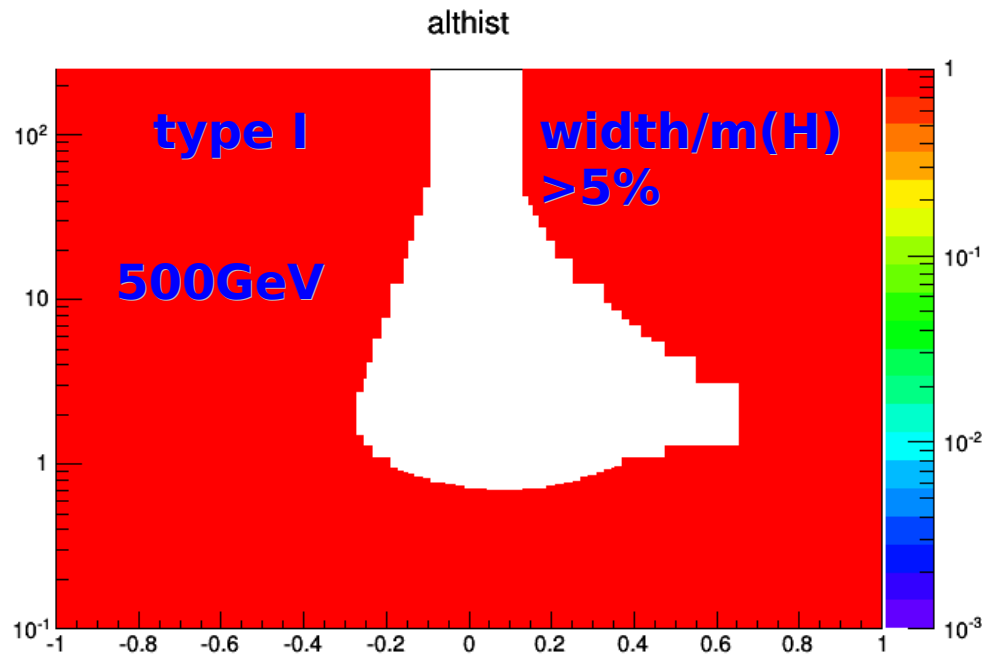


2HDM width_H / m_H @ $m_H = 300 \text{ GeV}$



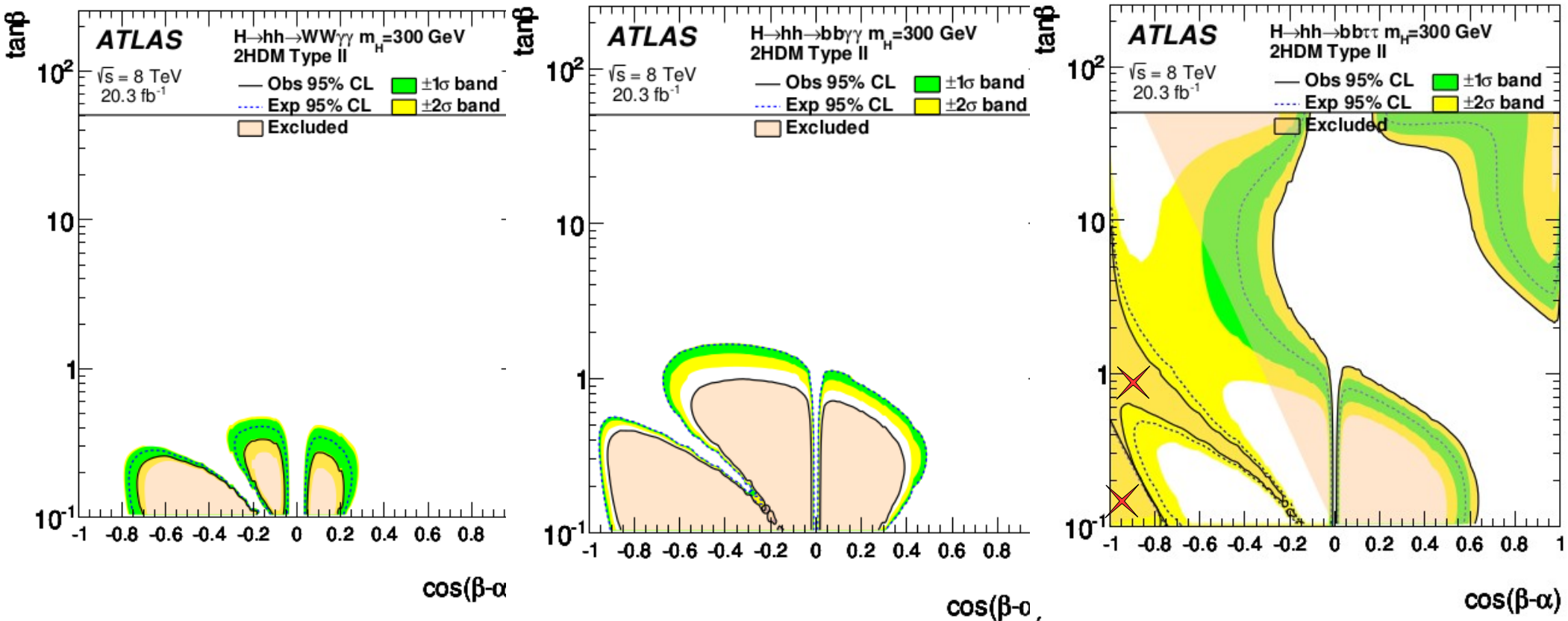
in general, width effect should be considered more than AZh case²¹

2HDM width_H / mH @ higher mH

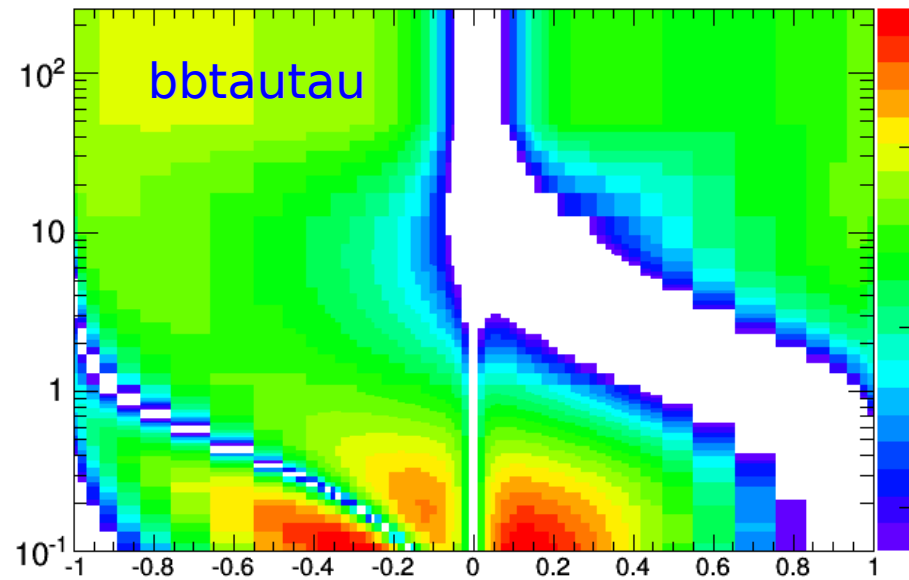
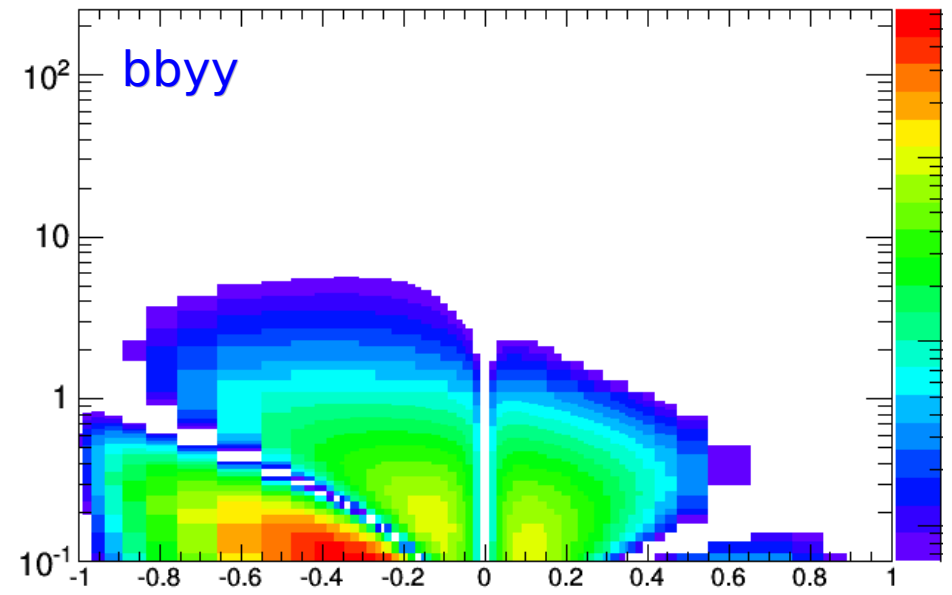
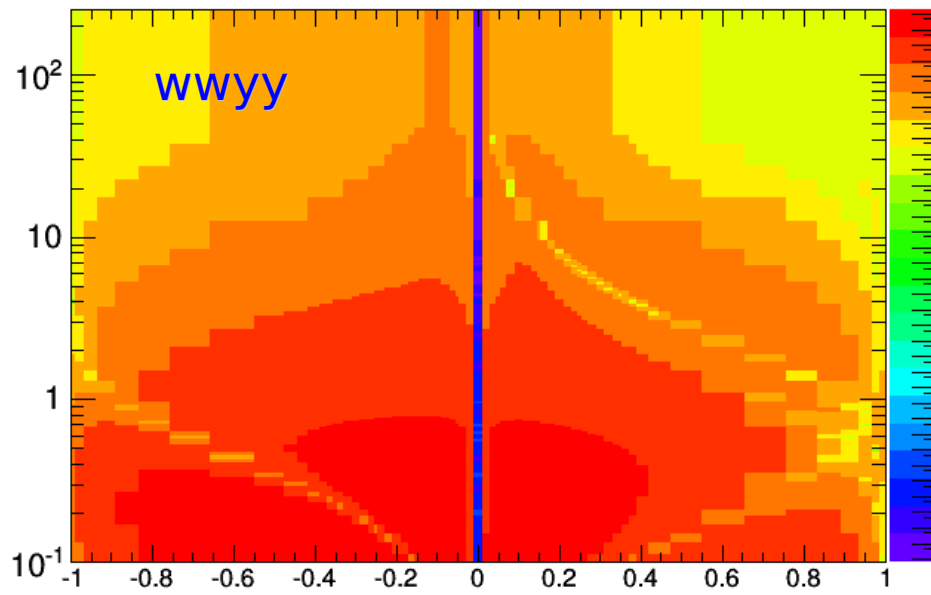


Looking at subchannel 2HDM exclusion

- Too many coupling, not easy to do parameterization for comb limits, Carl proposed to exclude with the most powerful limit at each point == overlay limits from all ch
- make 2HDM type II exclusion @ 300GeV $w\gamma\gamma$ / $b\gamma\gamma$ / $b\tau\tau$
- $\tan\beta$ should be larger than 0.5 (theo stability, we did it in Azh)
 - but $w\gamma\gamma$ does not exclude anything if you start from $t\beta=0.5$

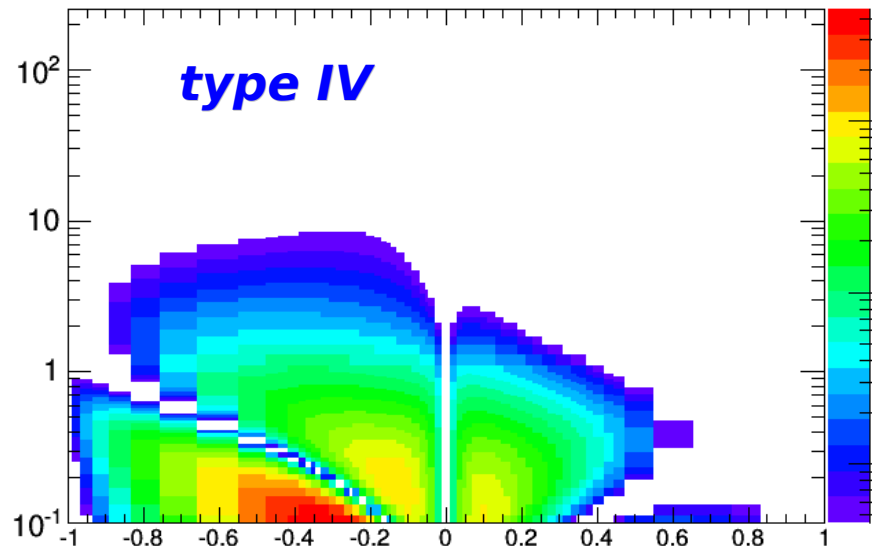
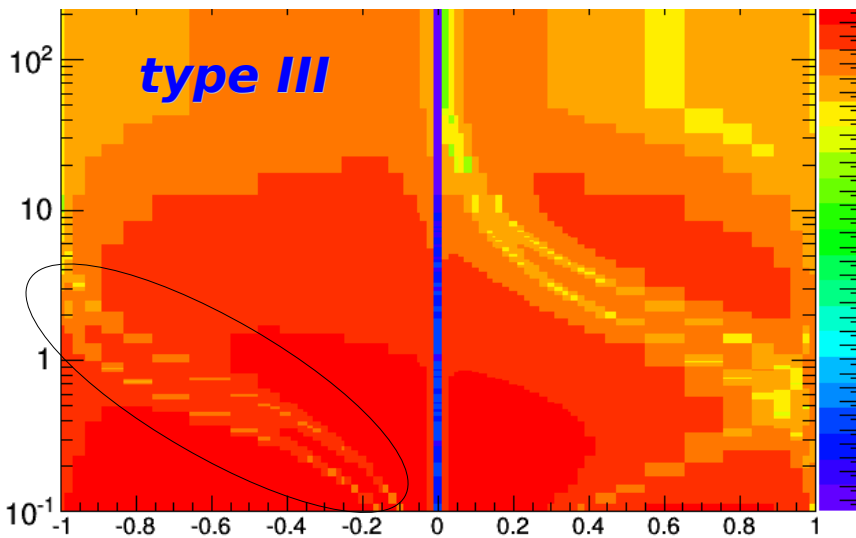
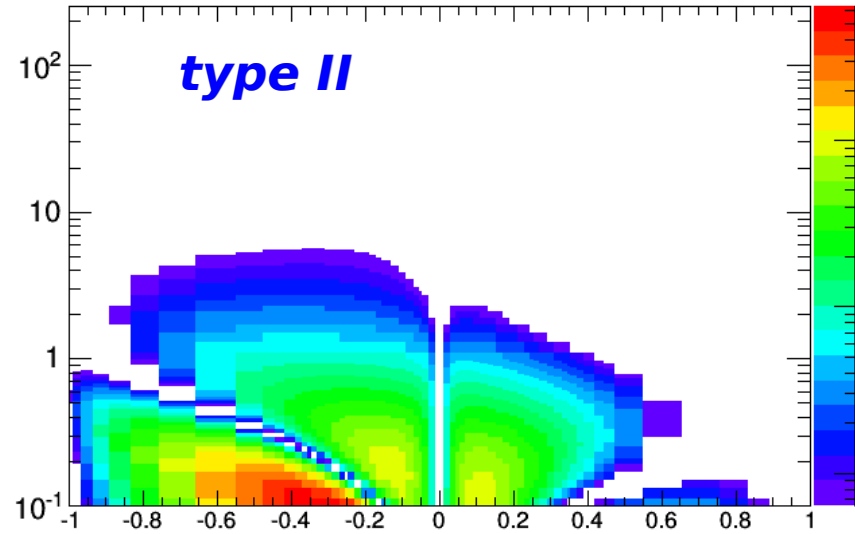
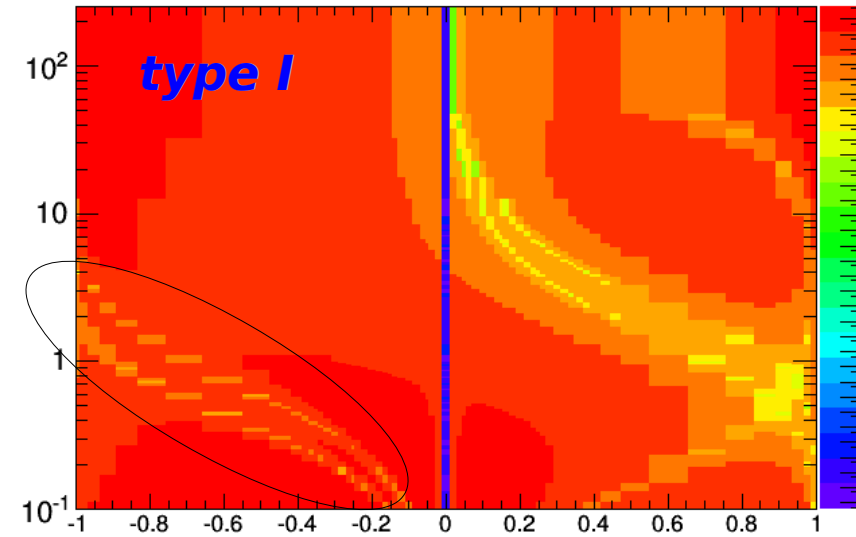


Check subchannel 2HDM xsbr contours



2HDM

- ***XS * BR contours in bby***



double valley structure will be difficult to deal with interpolation in plotting ...