Dark Photon Search at A Circular e^+e^- Collider

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Motivation

- 2 The dark photon model
- 3 Dark photon at a circular e^+e^- collider
- Sensitivities at CEPC and FCC-ee circular colliders



Motivation

• The Standard Model has demonstrated huge successes in providing experimental predictions!



Z=-4 Fre FMV + i \py + h.c. + 4: Yii 4: + h. c. $D_{\mu}\phi l^2 - V(\phi)$

Motivation

• But it leaves some phenomena unexplained. Evidences for DM.



(a) Galaxy rotation curves





(c) Gravitational lens

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Bullet cluster

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(d)

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- $SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_{A'}$ a new gauge group added
- Lagrangian with kinetic mixing of $U(1)_Y$ and $U(1)_{A'}$

$$L_{\text{kinetic}} = -\frac{1}{4} B_0^{\mu\nu} B_{0,\mu\nu} - \frac{1}{2} \sigma F_{0,\mu\nu}' B_0^{\mu\nu} - \frac{1}{4} F_{0,\mu\nu}' F_0^{\prime\mu\nu} . \tag{1}$$

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$$B_0 = c_W A_0 - s_W Z_0$$
: $U(1)_Y$ gauge field
• $B_{0,\mu\nu} = \partial_\mu B_{0,\nu} - \partial_\nu B_{0,\mu}$: $U(1)_Y$ gauge field strength tensor
• A'_0 : dark photon field
• $F'_{0,\mu\nu} = \partial_\mu A'_{0,\nu} - \partial_\nu A'_{0,\mu}$: dark photon field strength tensor
• σ : mixing parameter

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$$c_W = \cos \theta_W$$
, $s_W = \sin \theta_W$

• After $U(1)_{A'}$ symmetry breaking, A'_0 receives a mass $m_{A'}$.

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• Redefine the fields to get rid of mixing terms.

$$\begin{pmatrix} A_{0} \\ Z_{0} \\ A'_{0} \end{pmatrix} = \begin{pmatrix} 1 & 0 & -\frac{c_{W}\sigma}{\sqrt{1-\sigma^{2}}} \\ 0 & 1 & \frac{s_{W}\sigma}{\sqrt{1-\sigma^{2}}} \\ 0 & 0 & \frac{1}{\sqrt{1-\sigma^{2}}} \end{pmatrix} \begin{pmatrix} A \\ Z \\ A' \end{pmatrix} .$$
(2)

• The mass matrix for the redefined fields A, Z, A' is in the form

$$M = \begin{pmatrix} 0 & 0 & 0 \\ 0 & m_Z^2 & \frac{\sigma s_W}{\sqrt{1 - \sigma^2}} m_Z^2 \\ 0 & \frac{\sigma s_W}{\sqrt{1 - \sigma^2}} m_Z^2 & \frac{1}{1 - \sigma^2} m_{A'}^2 + \frac{\sigma^2 s_W^2}{1 - \sigma^2} m_Z^2 \end{pmatrix} .$$
(3)

• We need to further diagonalize this mass matrix.

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• After diagonalizing the mass matrix, we get the mass eigenstates A_1 , Z_1 and A'_1 , also their masses. Restriction: $m_Z - m_{A'} \gg \sigma^2 m_Z$. In this work, $m_{A'}$: $1 \sim 60$ GeV. σ : $10^{-3} \sim 10^{-2}$.

$$m_{A_1}^2 = 0,$$
 (4)

$$m_{Z_1}^2 \approx m_Z^2 + \frac{m_Z^4 s_W^2 \sigma^2}{m_Z^2 - m_Z^2},$$
 (5)

$$m_{A_1'}^2 \approx m_{A'}^2 + \frac{(c_W^2 m_Z^2 - m_{A'}^2)m_{A'}^2 \sigma^2}{m_Z^2 - m_{A'}^2}$$
(6)

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• Using notation $-c_W \sigma = \epsilon$, the effective Lagrangian concerning A_1 , Z_1 and A'_1 interaction with SM currents is in the following form

$$L_{\text{int}} = J_{em}^{\mu} A_{1\mu} + J_{Z}^{\mu} Z_{1\mu} + \epsilon J_{em}^{\mu} A_{1\mu}' + \underbrace{\frac{m_{A'}^{2} s_{W} \epsilon}{(m_{Z}^{2} - m_{A'}^{2}) c_{W}} J_{Z}^{\mu} A_{1\mu}'}_{(m_{Z}^{2} - m_{A'}^{2}) c_{W}} J_{Z}^{\mu} A_{1\mu}'$$
(7)

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• $m_{A'} < 1$ MeV, $A' \rightarrow 3\gamma$, Landau-Yang theorem. The dark photon can be cosmologically stable, and can be the candidate of dark matter.



Pospelov, Ritz, Voloshin 2008

Figure: $V \sim A'$ and $\kappa \sim \sigma$

• $m_{A'} > 1$ MeV, $A' \rightarrow e^+e^-$. The dark photon decays fast and can be the mediator of the dark force.

$$\Gamma_{A'} = \sum_{f} \Gamma(A' \to \bar{f}f) = \sum_{f} \epsilon^2 \frac{Q_f^2 \alpha_{em} m_{A'}}{3} (1 + \frac{2m_f^2}{m_{A'}^2}) \sqrt{1 - \frac{4m_f^2}{m_{A'}^2}} .$$
(8)

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- What a circular e^+e^- collider can do for the dark photon?
- We study the prodcution and search for dark photon using porcess $e^+e^- o \gamma A'^* o \gamma \mu^+\mu^-$



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• Parameters for some future e^+e^- colliders.



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• Feynman diagrams for $e^-e^+ \rightarrow \gamma \mu^- \mu^+$ process. $s = (p_1 + p_2)^2$, $s_3 = (k_1 + k_2)^2$



 $d\sigma_{\gamma(\gamma,Z)}/ds_3$

$$= \frac{4\alpha_{em}^{3}(s^{2} + s_{3}^{2})}{3s^{3}s_{3}(s - s_{3})} \left(s(\ln(s/m_{e}^{2}) - 1) + s_{3}(\ln(s_{3}/m_{\mu}^{2}) - 1)\right) \\ + \frac{\alpha_{em}^{3}(8\sin^{4}\theta_{W} - 4\sin^{2}\theta_{W} + 1)^{2}}{48\sin^{4}\theta_{W}\cos^{4}\theta_{W}} \frac{s^{2} + s_{3}^{2}}{s^{2}(s - s_{3})} \\ \times \left(\frac{s_{3}(\ln(s/m_{e}^{2}) - 1)}{(s_{3} - m_{Z}^{2})^{2}} + \frac{s(\ln(s_{3}/m_{\mu}^{2}) - 1)}{(s - m_{Z}^{2})^{2}}\right) \\ - \frac{\alpha_{em}^{3}(1 - 4\sin^{2}\theta_{W})^{2}}{16\sin^{4}\theta_{W}\cos^{4}\theta_{W}} \frac{s + s_{3}}{s^{2}(s - s_{3})} \frac{ss_{3}}{(s - m_{Z}^{2})(s_{3} - m_{Z}^{2})} \\ + \frac{\alpha_{em}^{3}(1 - 4\sin^{2}\theta_{W})^{2}}{6\sin^{2}\theta_{W}\cos^{2}\theta_{W}} \frac{s^{2} + s_{3}^{2}}{s^{2}(s - s_{3})} \left(\frac{\ln(s/m_{e}^{2}) - 1}{s_{3} - m_{Z}^{2}} + \frac{\ln(s_{3}/m_{\mu}^{2}) - 1}{s - m_{Z}^{2}}\right) \\ - \frac{\alpha_{em}^{3}}{4\sin^{2}\theta_{W}\cos^{2}\theta_{W}} \frac{s + s_{3}}{s^{2}(s - s_{3})} \left(\frac{s_{3}}{s_{3} - m_{Z}^{2}} + \frac{s}{s - m_{Z}^{2}}\right)$$
(9)

$$\frac{d\sigma_{\gamma A'}}{ds_3} \approx \frac{4\alpha_{em}^3 \epsilon^4 s_3 (s^2 + s_3^2)}{3s^2 (s - s_3) ((s_3 - m_{A'}^2)^2 + \Gamma_{A'}^2 m_{A'}^2)} \left(\ln(s/m_e^2) - 1 \right)$$
(10)

• Cross section near $m_{A'}$ ($\Gamma \ll \sigma_{\mu\mu} \ll m_{A'}$)

$$\sigma_{\gamma A'}^{m_{\mu\mu}} \left(\sigma_{\gamma(\gamma,Z)}^{m_{\mu\mu}} \right) = \int_{(m_{A'} - \sigma_{\mu\mu})^2}^{(m_{A'} + \sigma_{\mu\mu})^2} \frac{d\sigma_{\gamma A'}}{ds_3} \left(\frac{d\sigma_{\gamma(\gamma,Z)}}{ds_3} \right) ds_3 \tag{11}$$



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• The SM contributions $N_{\gamma(\gamma,Z)}$ and dark photon contribution $N_{\gamma A'}$

$$N_{\gamma(\gamma,Z)} = \sigma_{\gamma(\gamma,Z)}^{m_{\mu\mu}} LT , \quad N_{\gamma A'} = \sigma_{\gamma A'}^{m_{\mu\mu}} LT .$$
 (12)

- Take the resolution of the invariant mass measurement similar to LHCb, $\sigma_{\mu\mu} = 0.5\% m_{A'}$.
- $N_{\gamma A'}$: signal
- $\sqrt{N_{\gamma(\gamma,Z)}}$: the statistic sensitivity for SM background
- $\chi = N_{\gamma A'} / \sqrt{N_{\gamma(\gamma,Z)}}$: indicator how well one can obtain constraints on ϵ^2 and $m_{A'}$

• Cross section for SM background and dark photon contribution for different dark photon masses.



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• One-year running events number for different colliders.



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- We have studied the possibility of searching for dark photon at a circular e^+e^- collider through the process $e^+e^- \rightarrow \gamma A'^* \rightarrow \gamma \mu^+\mu^-$.
- The CEPC and FCC-ee e^+e^- colliders can provide good sensitivity especially for dark photon mass at range 20 \sim 60GeV.
- In the range of 20 GeV to 60 GeV for $m_{A'}$, the smallest $\sigma_{\mu\mu}$ is 100 MeV which is reachable at the CEPC and FCC-ee.
- Still working on the detector simulation.

Backup

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$$\sigma_{\it NLO}/\sigma_{\gamma(\gamma,Z)}=5 imes10^{-5}$$

Table: ratio of interference term to dark photon contribution at $\epsilon = 10^{-2}$

$\frac{\sigma_{int}}{\sigma_{out}}(10^{-5})$ $\sqrt{s}(\text{GeV})$			
	160	240	350
$m_{A'}(\text{GeV})$			
1	3.1	3.1	3.1
30	6.0	5.4	5.2
60	12	9.3	8.1

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Table: ratio of additional dark photon contribution to old dark photon contribution at $\epsilon = 10^{-2}$

$\frac{\sigma_{\gamma A'}^{new} - \sigma_{\gamma A'}}{\sigma_{\gamma A'}} \sqrt{s} (\text{GeV})$ $m_{A'} (\text{GeV})$	160	240	350
1	$3.7 imes 10^{-11}$	$3.7 imes10^{-11}$	$3.7 imes10^{-11}$
30	$4.0 imes 10^{-5}$	$4.0 imes10^{-5}$	$4.0 imes10^{-5}$
60	$5.2 imes 10^{-3}$	$5.2 imes10^{-3}$	$5.2 imes10^{-3}$

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