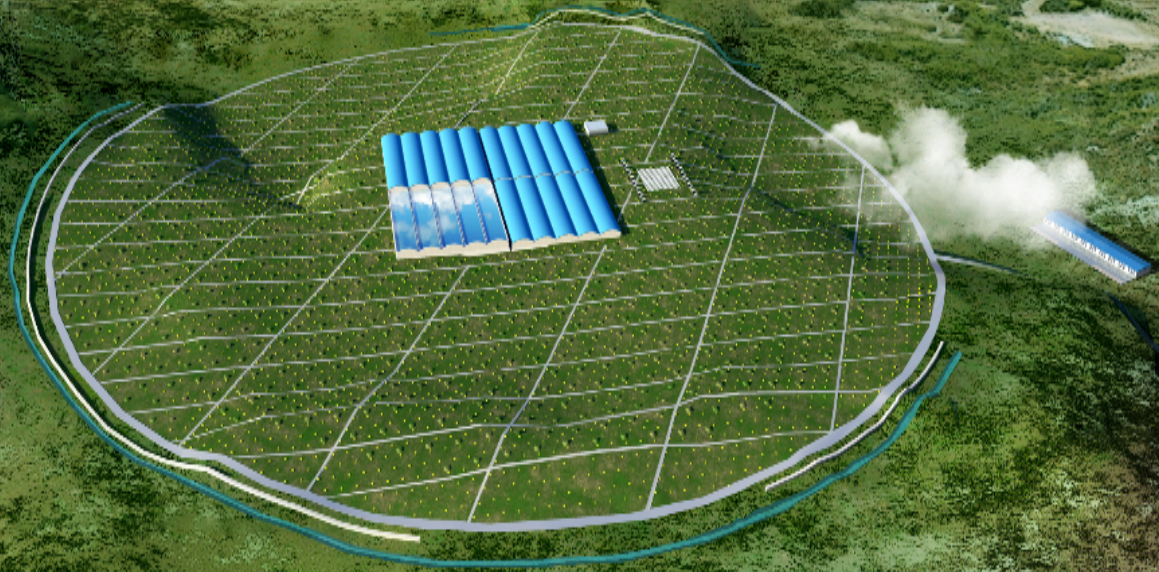


Survey of the TeV gamma ray sky by the LHAASO-WCDA experiment



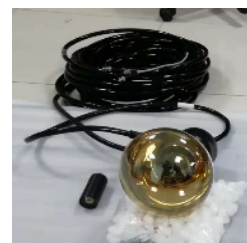
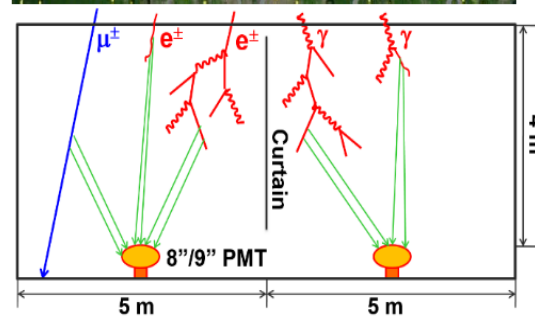
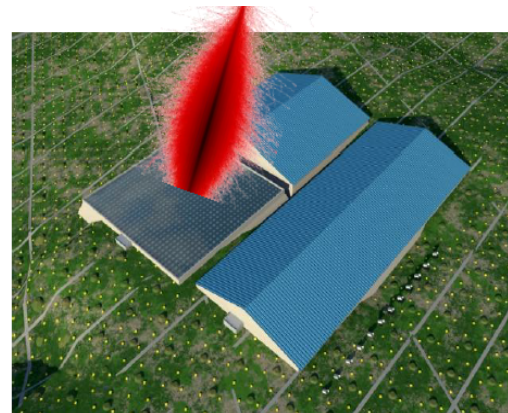
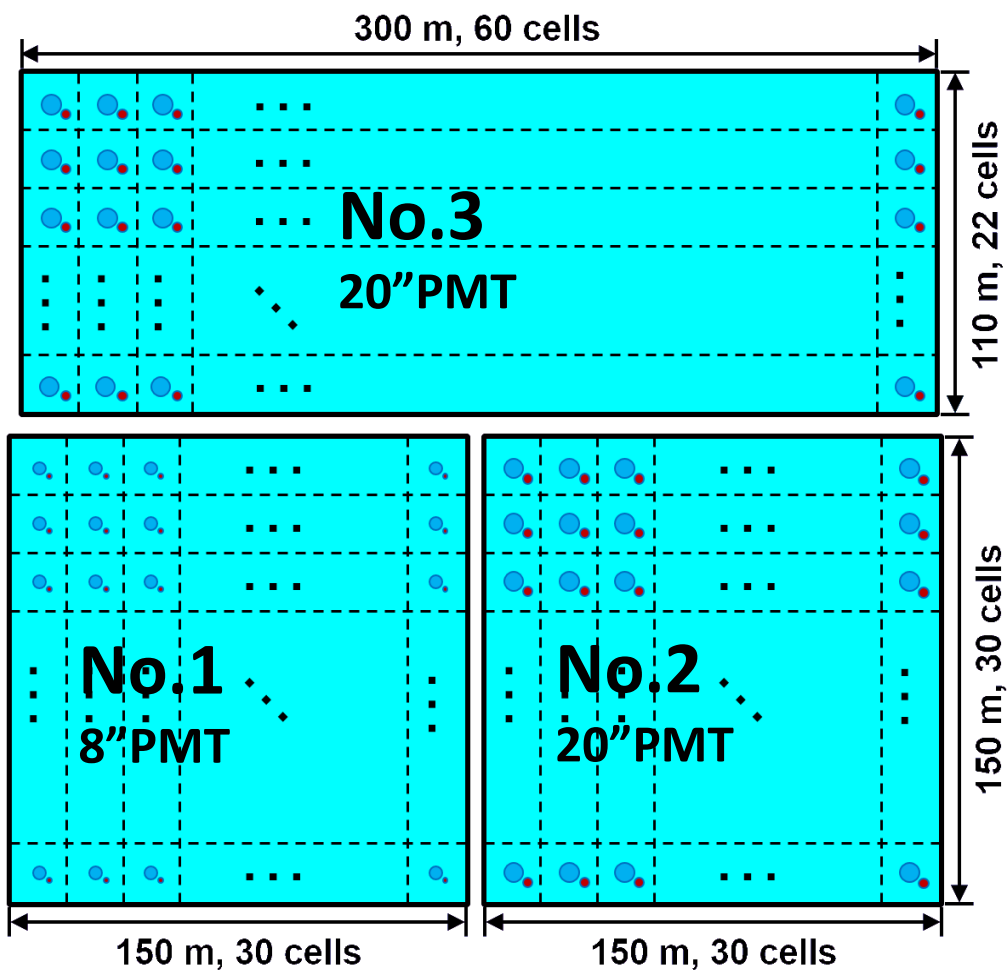
Min Zha @IHEP

For the LHAASO collaboration

outline

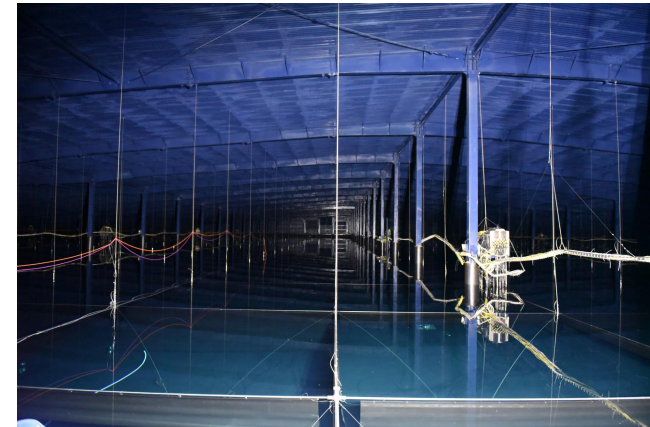
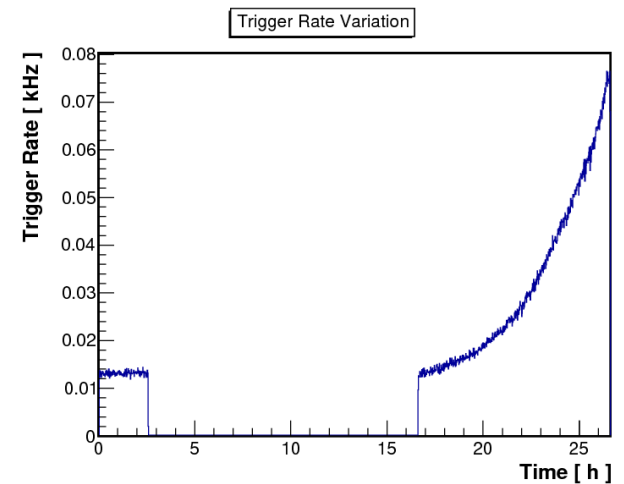
- Water Cherenkov Detector Array status
- Preliminary results from WCDA-1
- Summary and outlook

LHAASO-WCDA



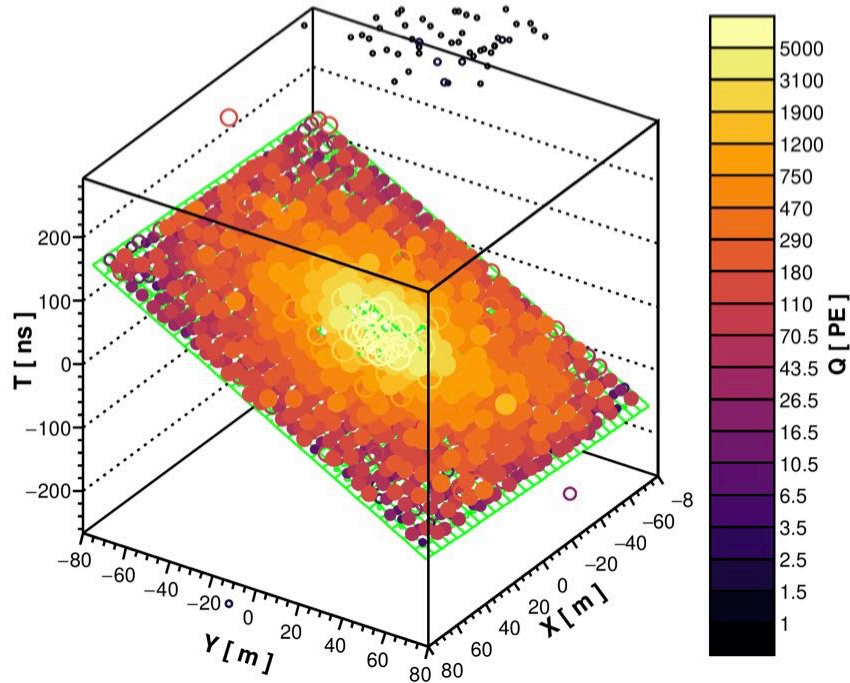
Timeline of WCDA#1

- 2018/10/22: start deployment;
- 2019/01/10: deployment complete;
- 2019/01/11: start dry data-taking mode;
- 2019/02/25: DAQ with water ;
- 2019/04/07: achieve full water depth;
- 2019/04/26: start scientific data-taking .
- Running mode @ test water running mode
 - 2 hours + multi-mode looping
 - S3:90s E2:30m S3:90s SH4:20m S3:90s E2:30m S3:90s E2:30m
- Normal running mode from 20190511
 - EM17
 - 17 hits within 300 ns;
- GRB + shower mode from 20190604
 - GRBM17
 - To now: 36 GRBs
- B-pmt + S-pmt joint mode from 20191009
 - P1GRBM20M

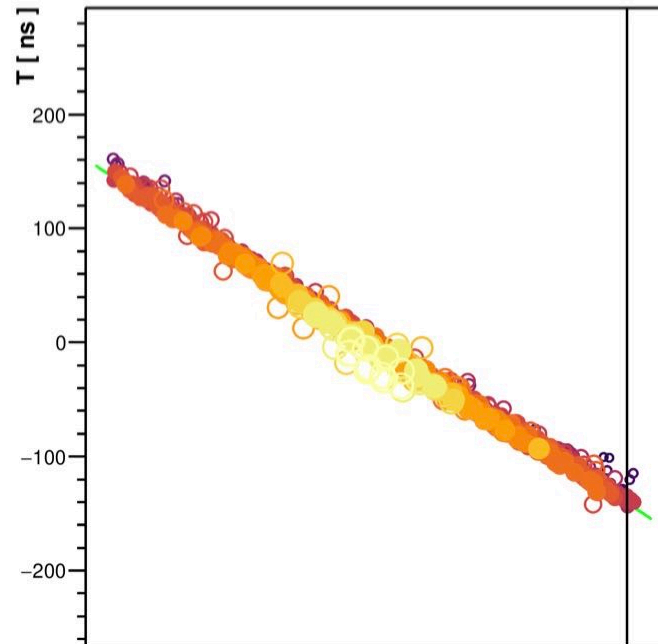


Valentine day event

20190214/173639/0.180367688: nTrig=0, $\theta=34.43\pm 0.02^\circ$, $\phi=93.67\pm 0.03^\circ$

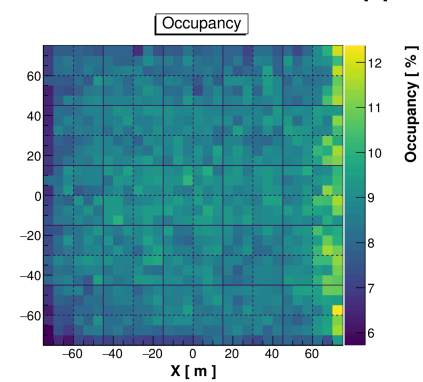
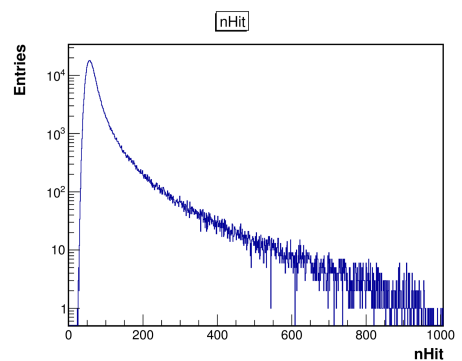
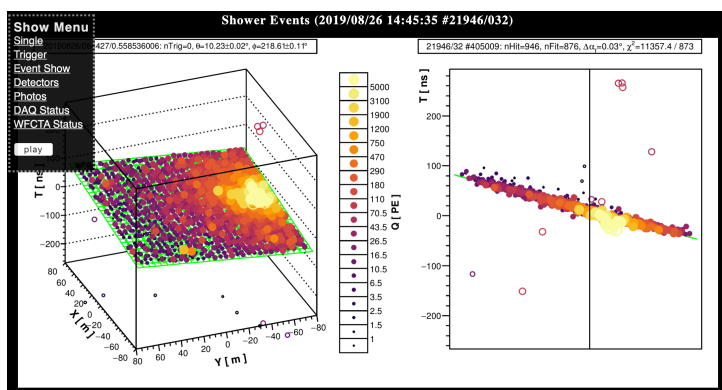
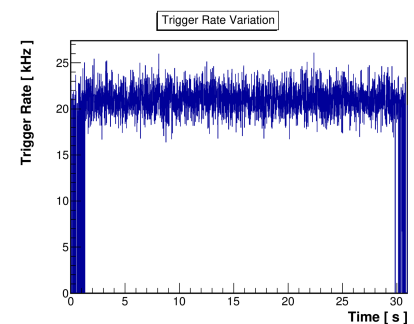
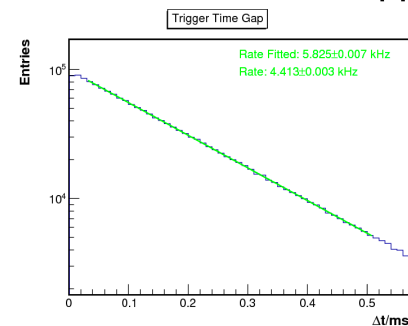
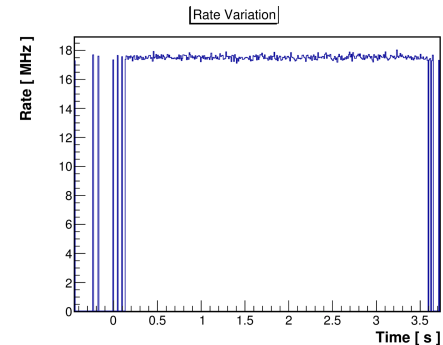
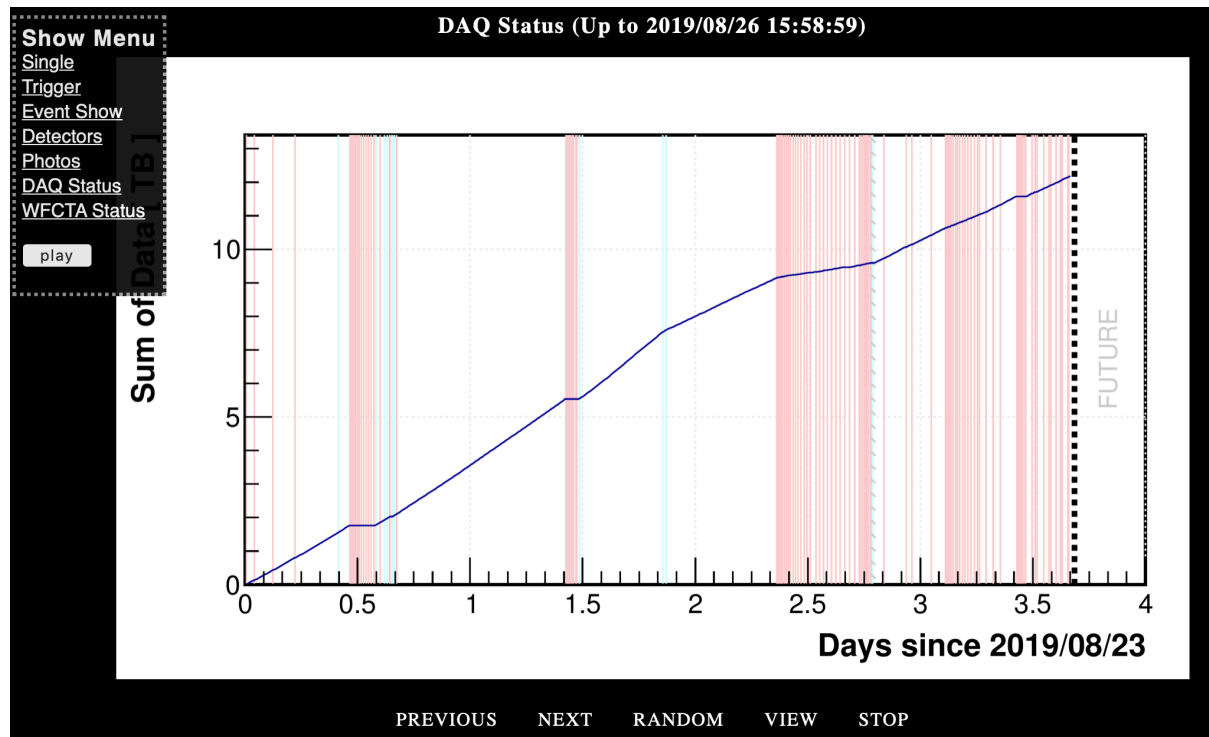


10001/7 #85620: nHit=1687, nFit=784, $\Delta\alpha_r=0.03^\circ$, $\chi^2=6127.2 / 781$

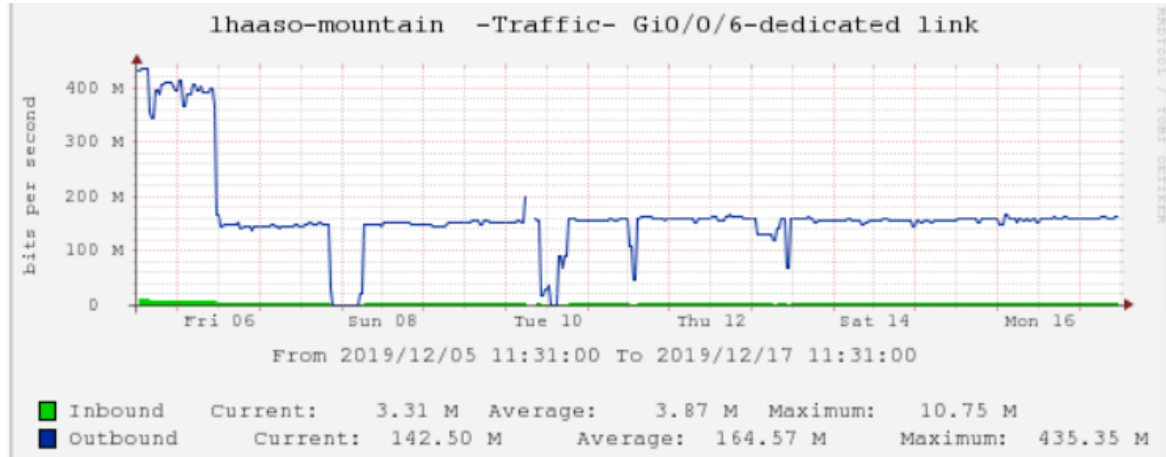


Data on-line monitor

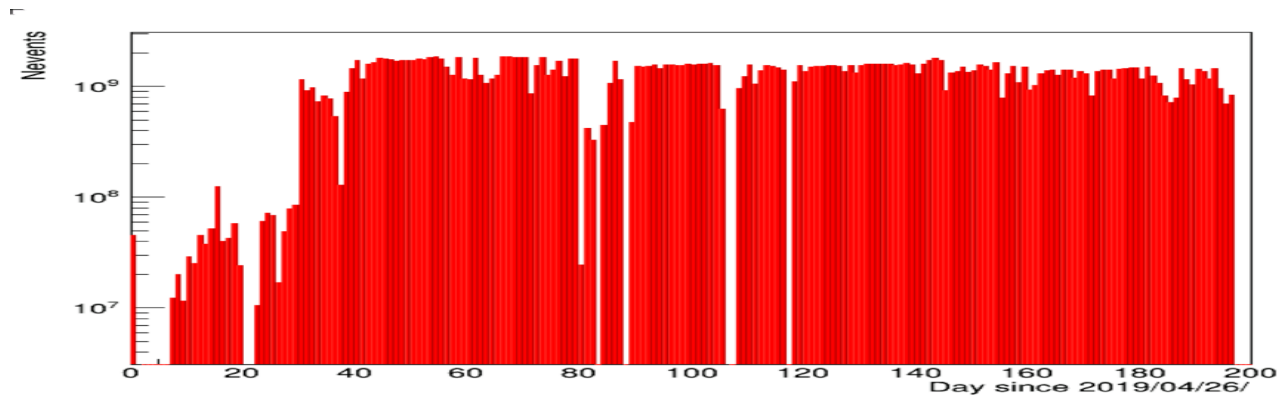
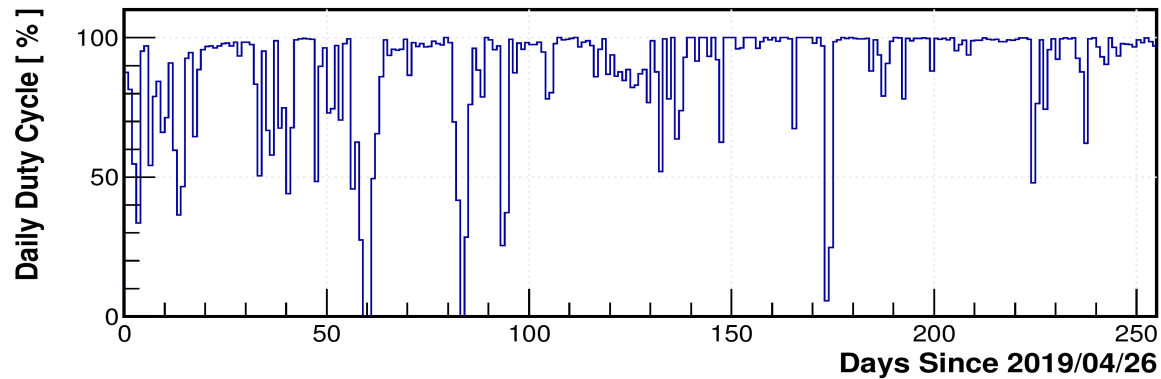
http://202.38.128.103:9019/wcda_detectors/



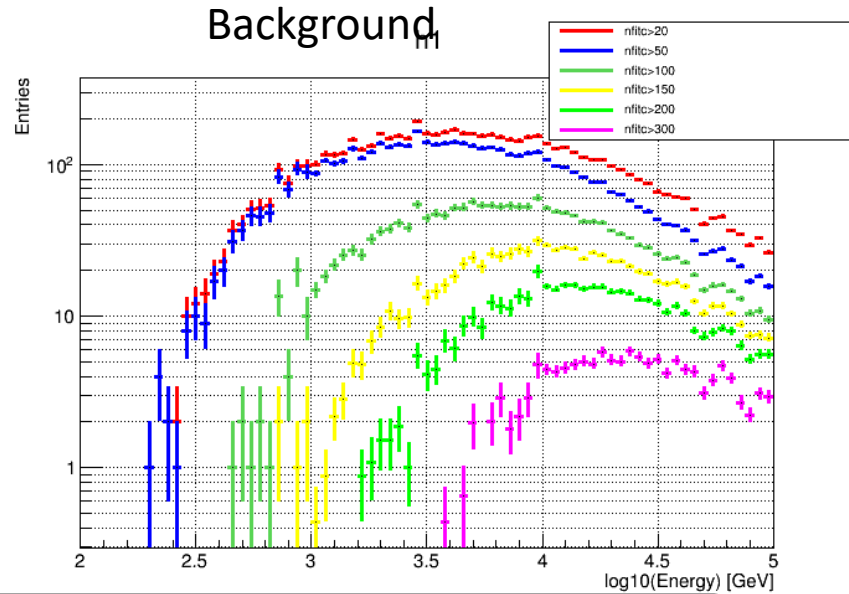
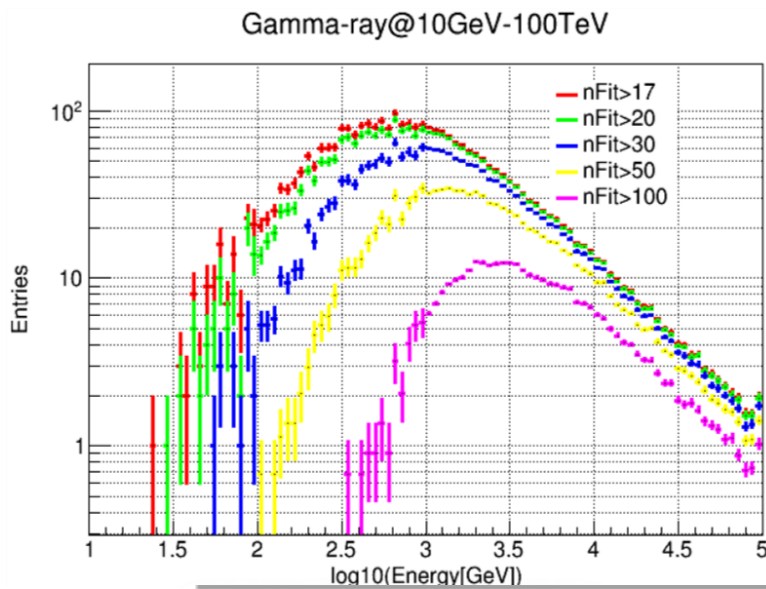
WCDA-1 data taking



Overall Duty Cycle = 88.12%



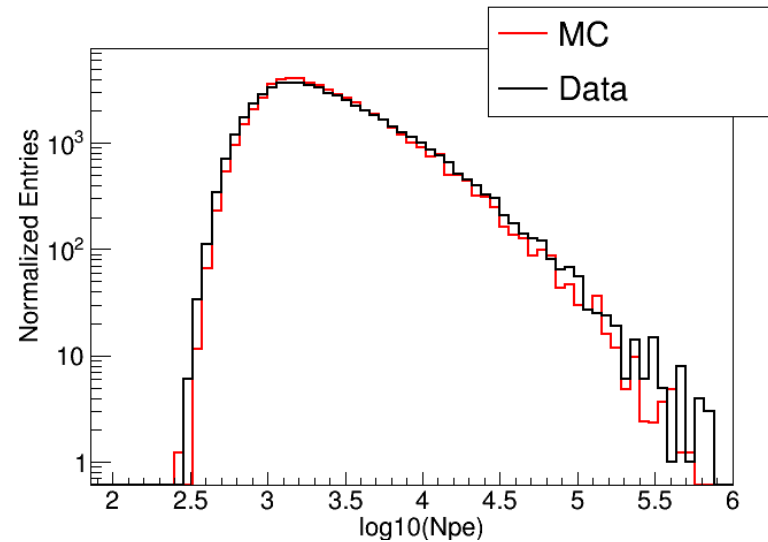
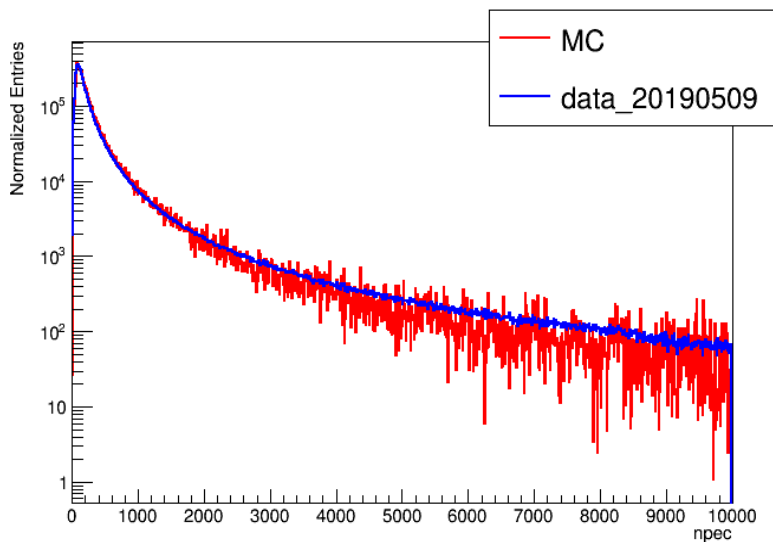
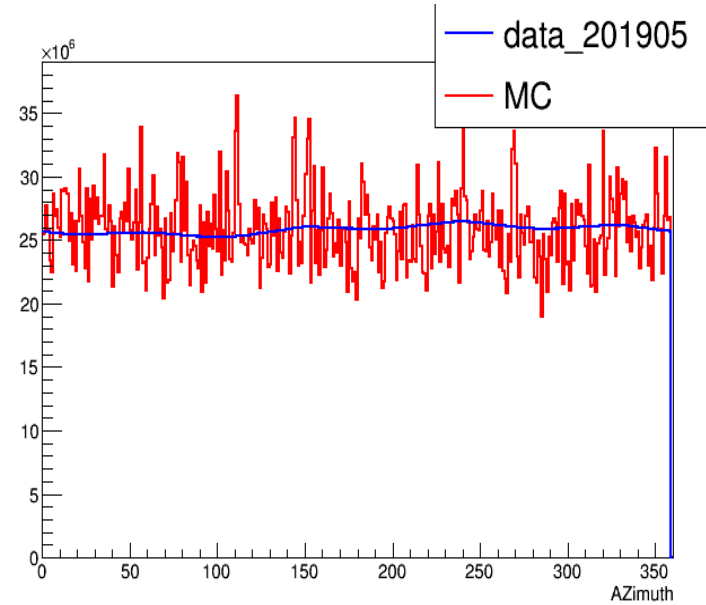
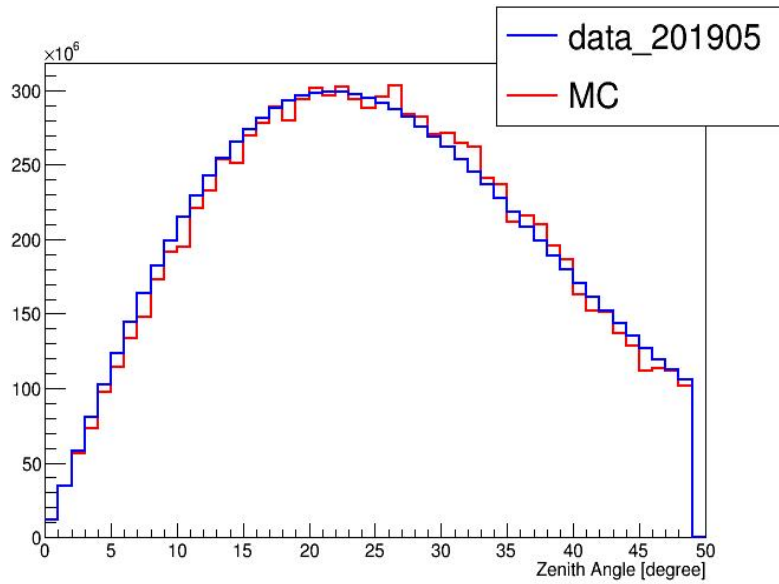
Data .vs. MC



CR	>17 (NHit)	>20 (NHit)	>30 (NHit)	>50 (NHit)	>100 (NHit)
P	13.2 kHz	9.7 kHz	4.7 kHz	1.9 kHz	0.6 kHz
He	4.9 kHz	3.5 kHz	1.7 kHz	0.8 kHz	0.3 kHz
CNO	1.3 kHz	0.9 kHz	0.4kHz	0.2 kHz	0.07 kHz
MgAlSi	0.6 kHz	0.4 kHz	0.2 kHz	0.08 kHz	0.03 kHz
Fe	0.5 kHz	0.3 kHz	0.16 kHz	0.07 kHz	0.02 kHz
All (8m)	20.5 kHz	15.0 kHz	7.2 kHz	2.9 kHz	1.0 kHz

- ◆ The data trigger is about 22.1 KHz and the M.C. is 20.5 KHz, thus the M.C. trigger rate is consistent with data within 8% level.

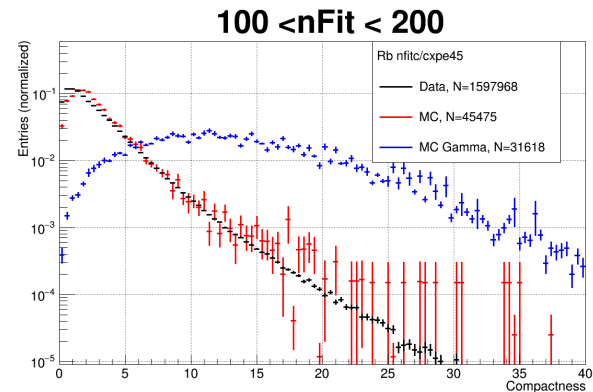
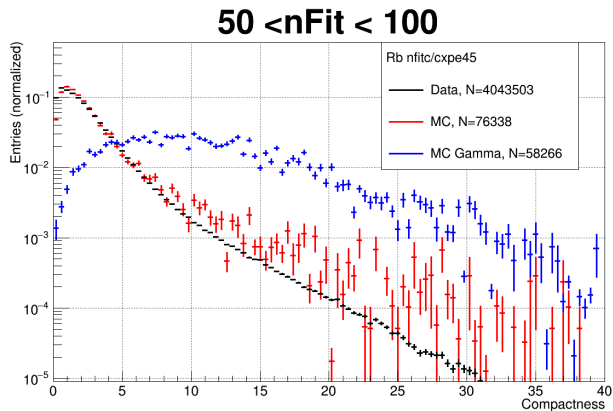
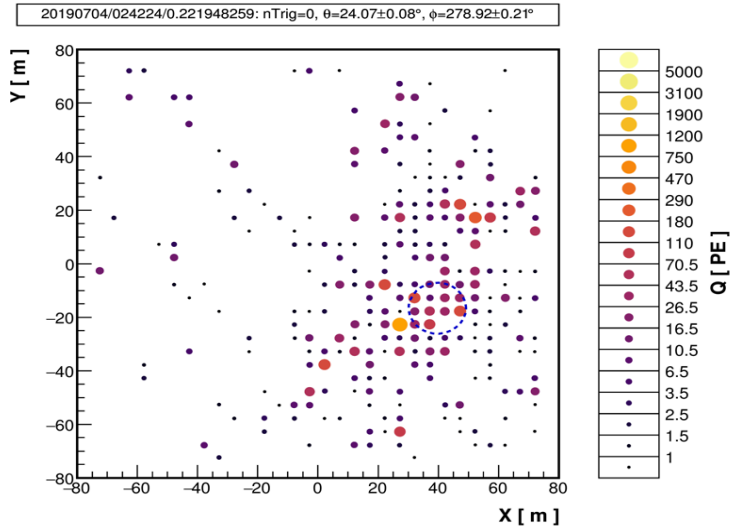
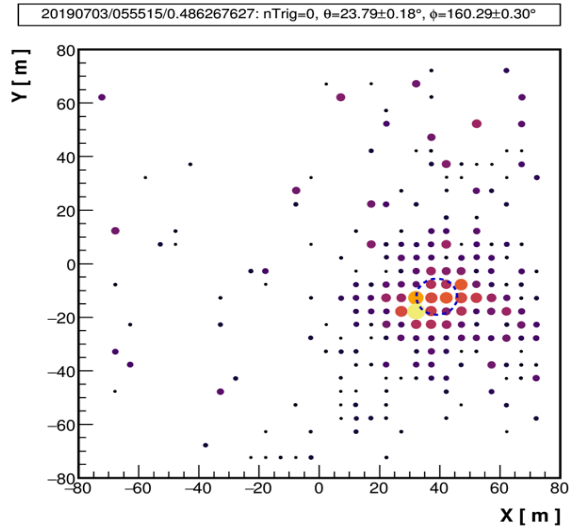
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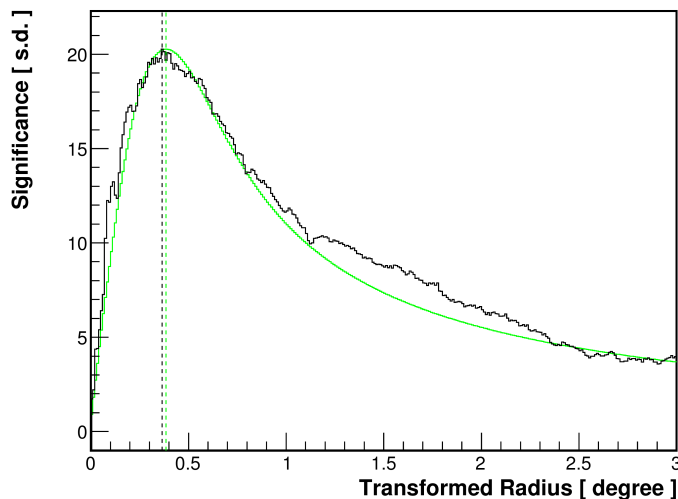
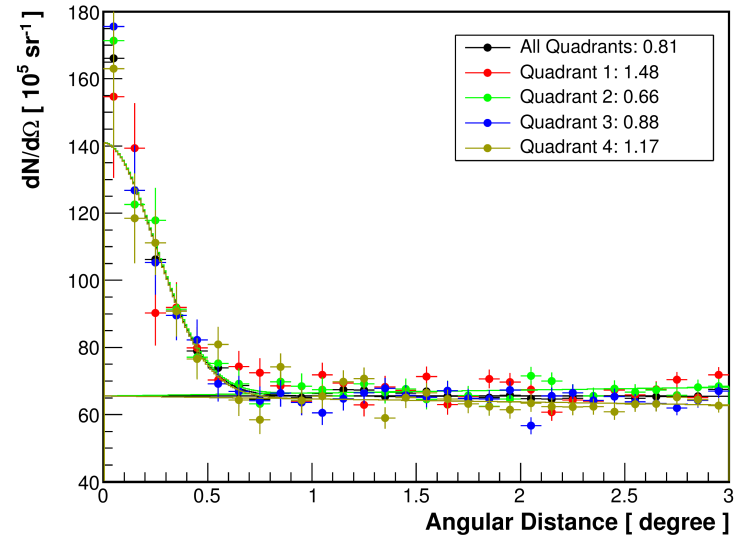
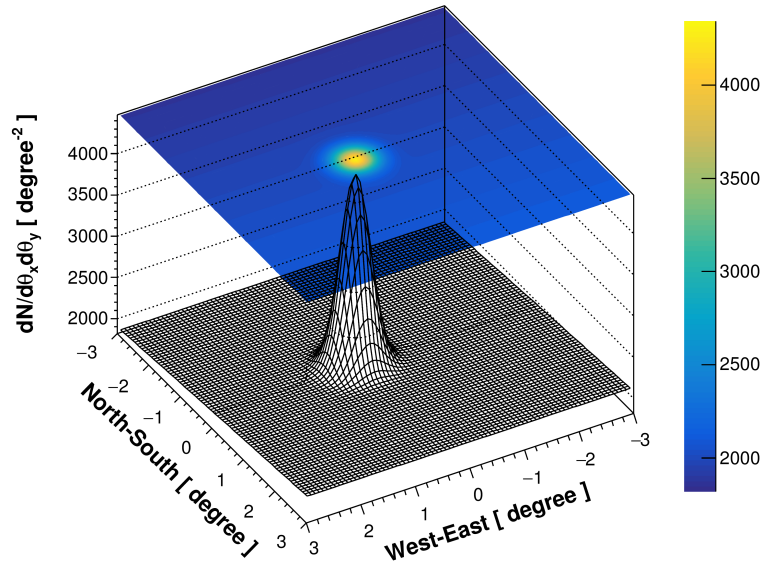
Background rejection @ compactness

Top3: $N_{hit}=236$ $C=39$

Ptop3: $N_{hit}=261$



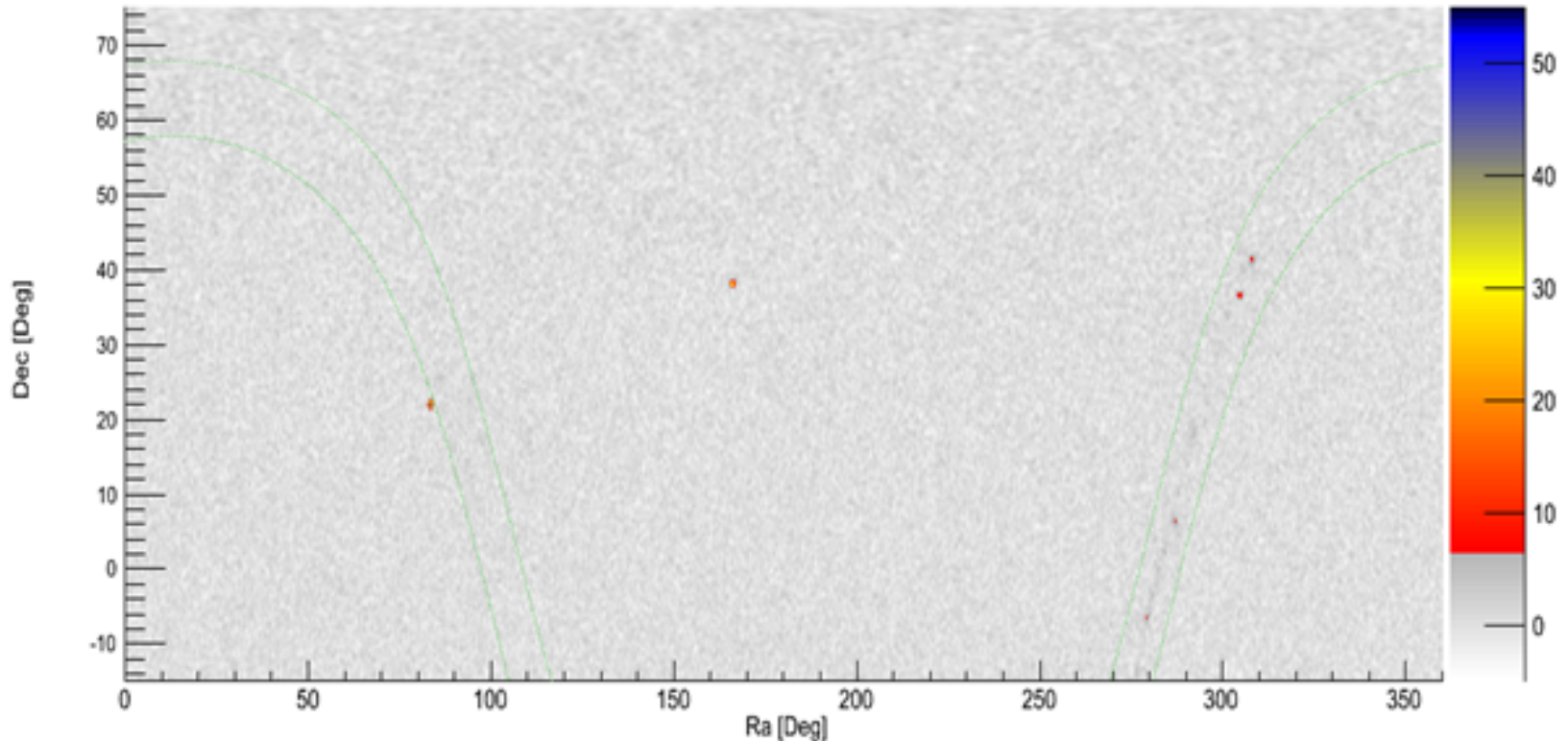
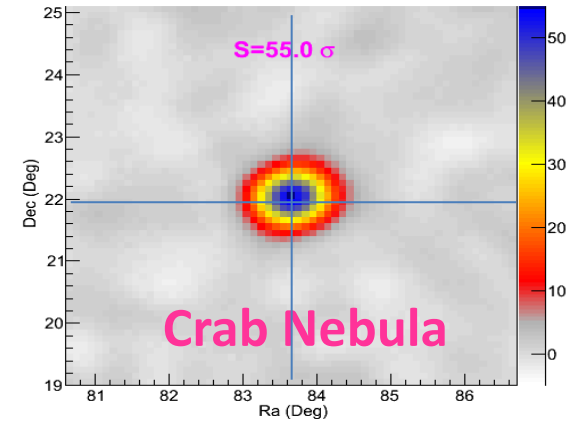
Crab Nebular observation per month



- 20 sigma/month;
- Sigma is 0.25 deg;
- Pointing is less than 0.1

200 days Sky map

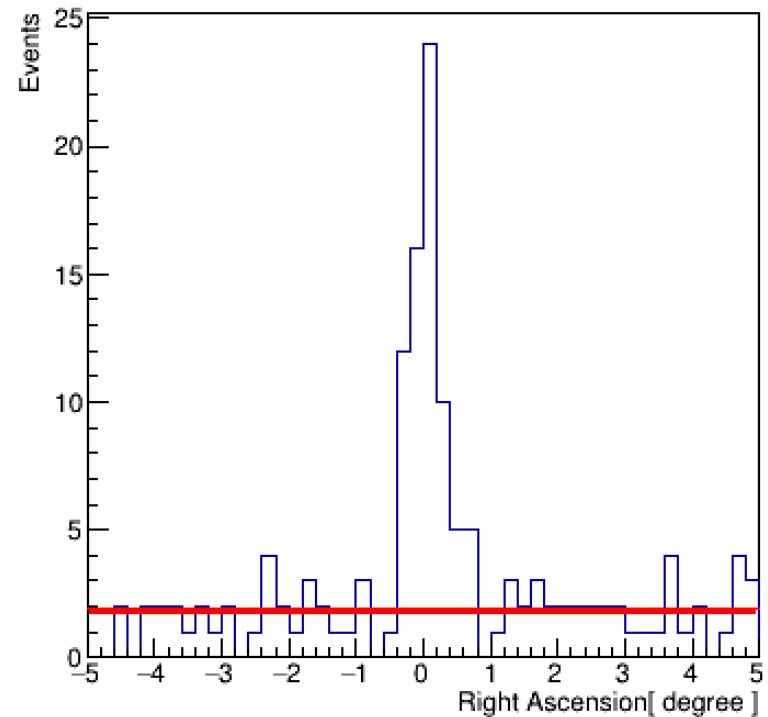
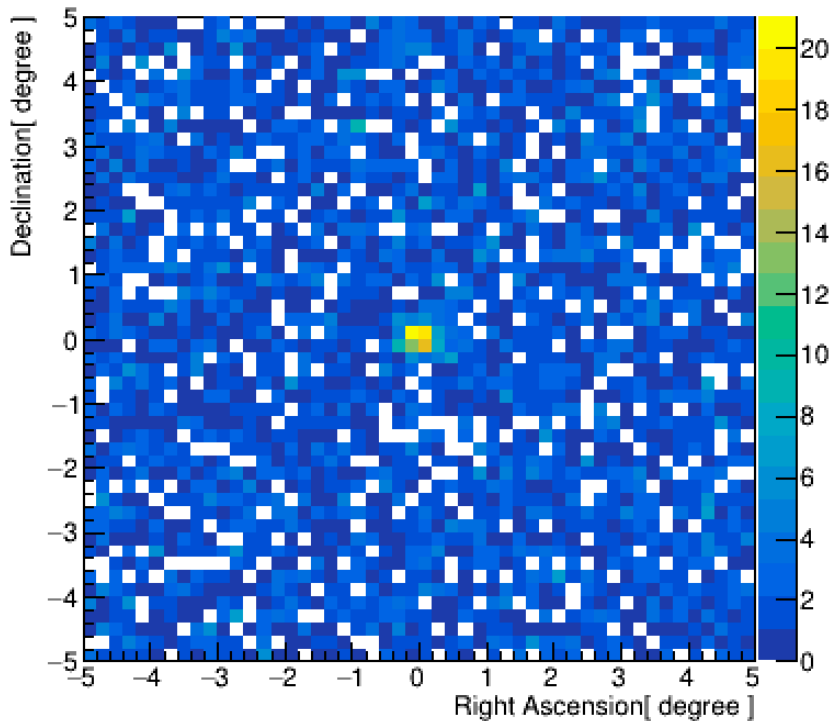
- Data samples
 - Nfit>100;
 - 60 mCU
 - PSF: 0.26
 - 6 hot points with significance $>7 \sigma$

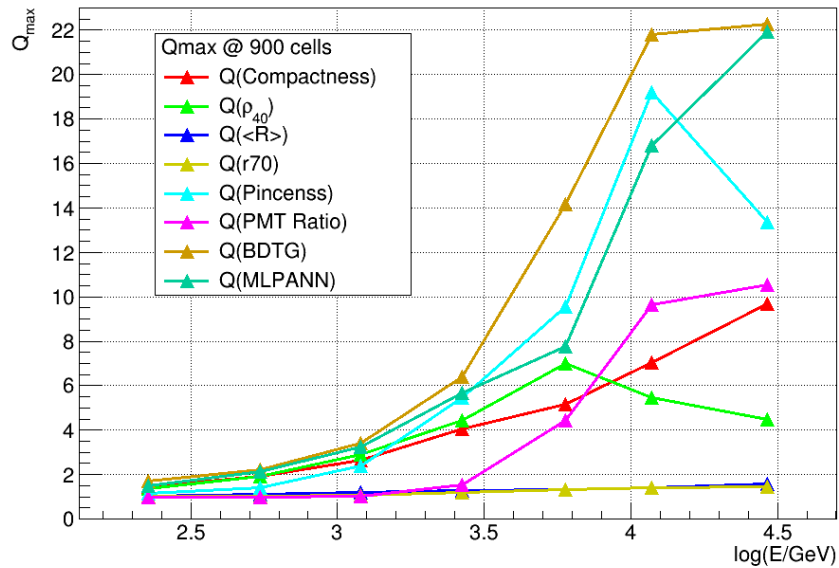


Gamma-rich sample from the Crab

- Hard cuts on high energy sample: $S/N \sim 12$

Crab



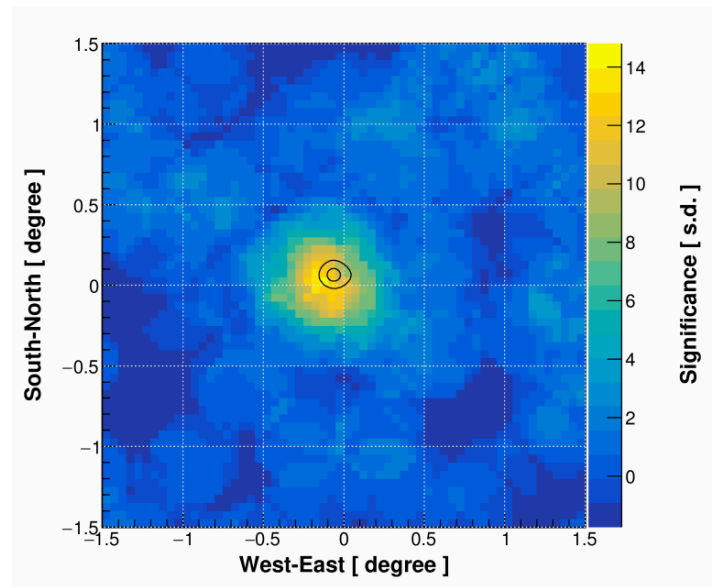
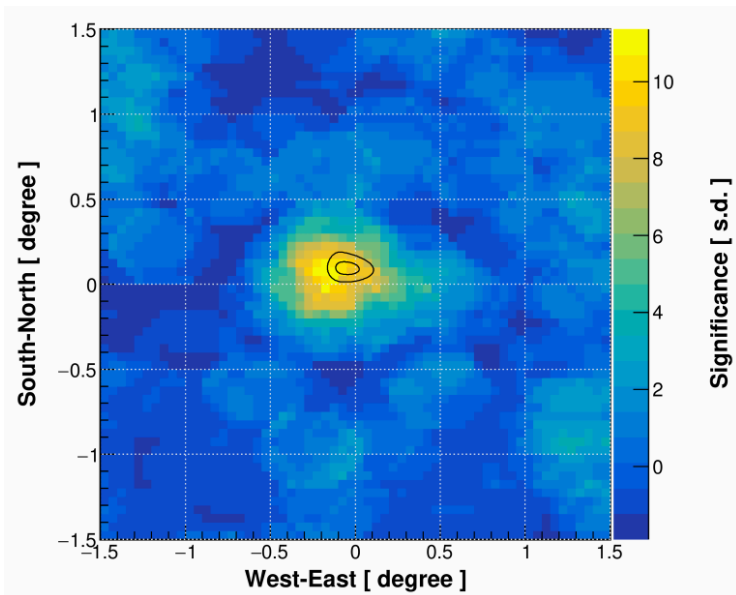


Compactness cut

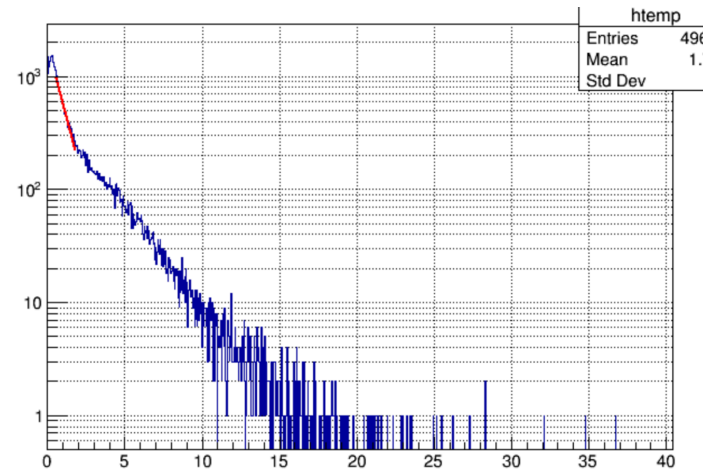
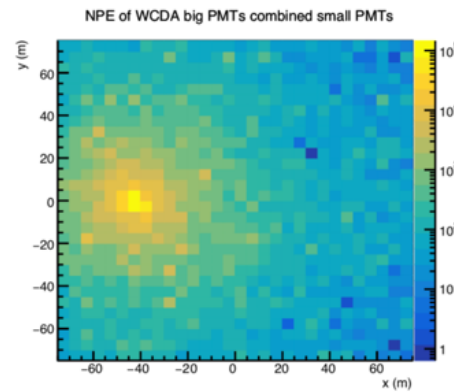
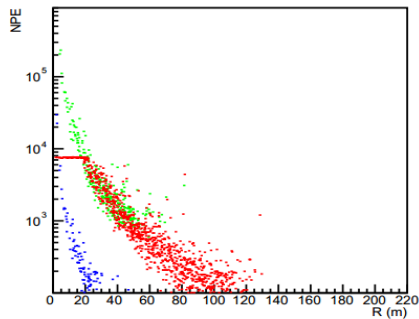
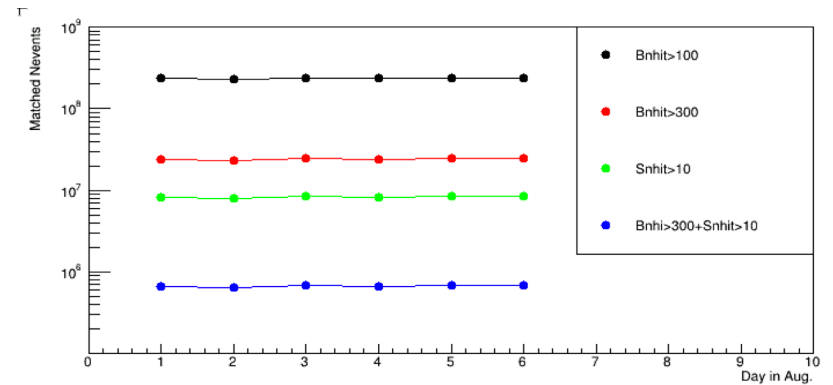
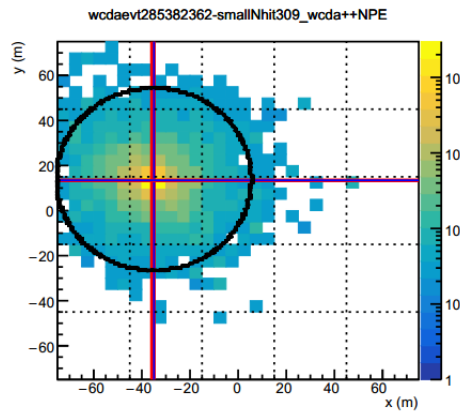
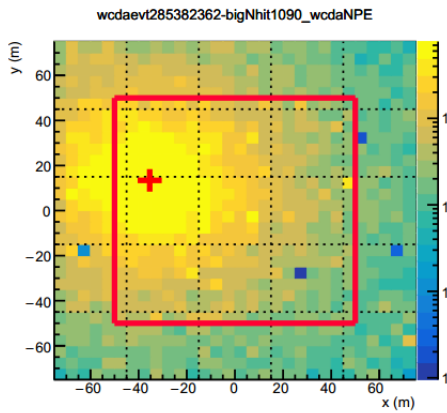
nfitc:300-500

BDTG

Improvement: 32%

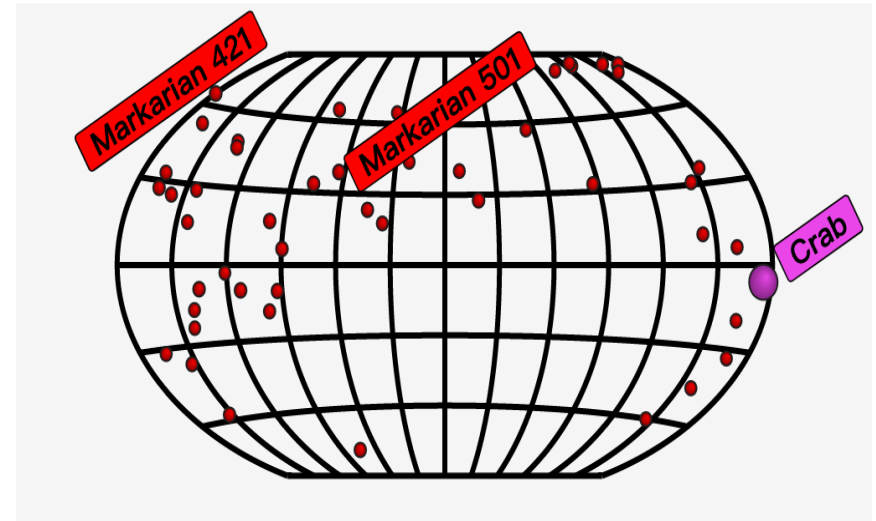
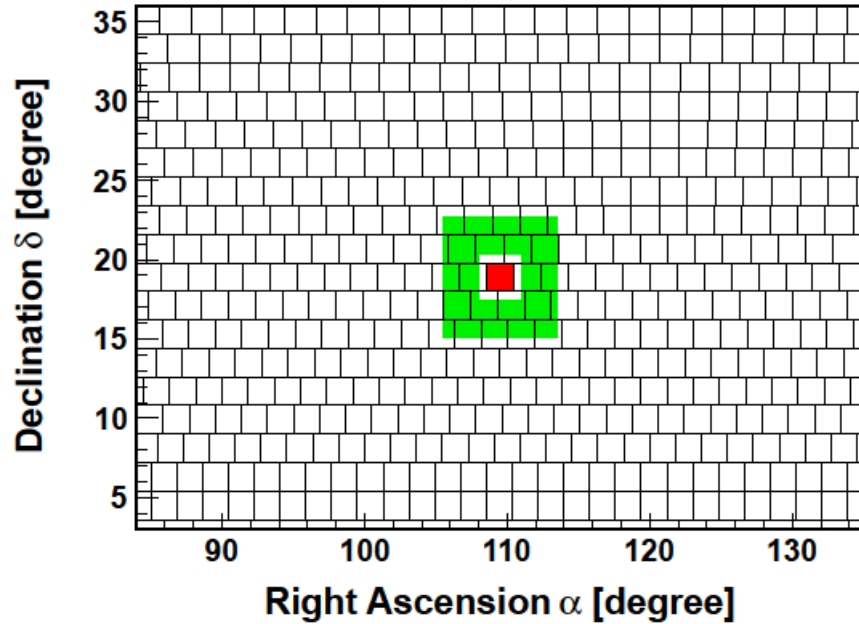


Joint B-PMT + S-PMT event analysis

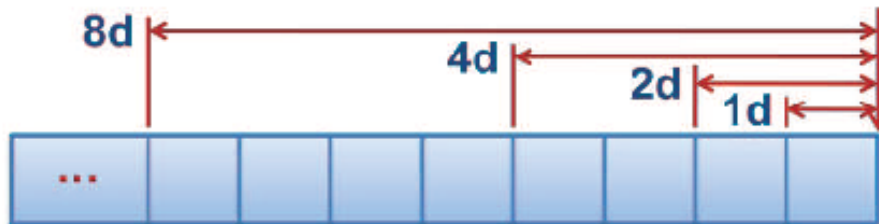


- New G/P separation parameter ?

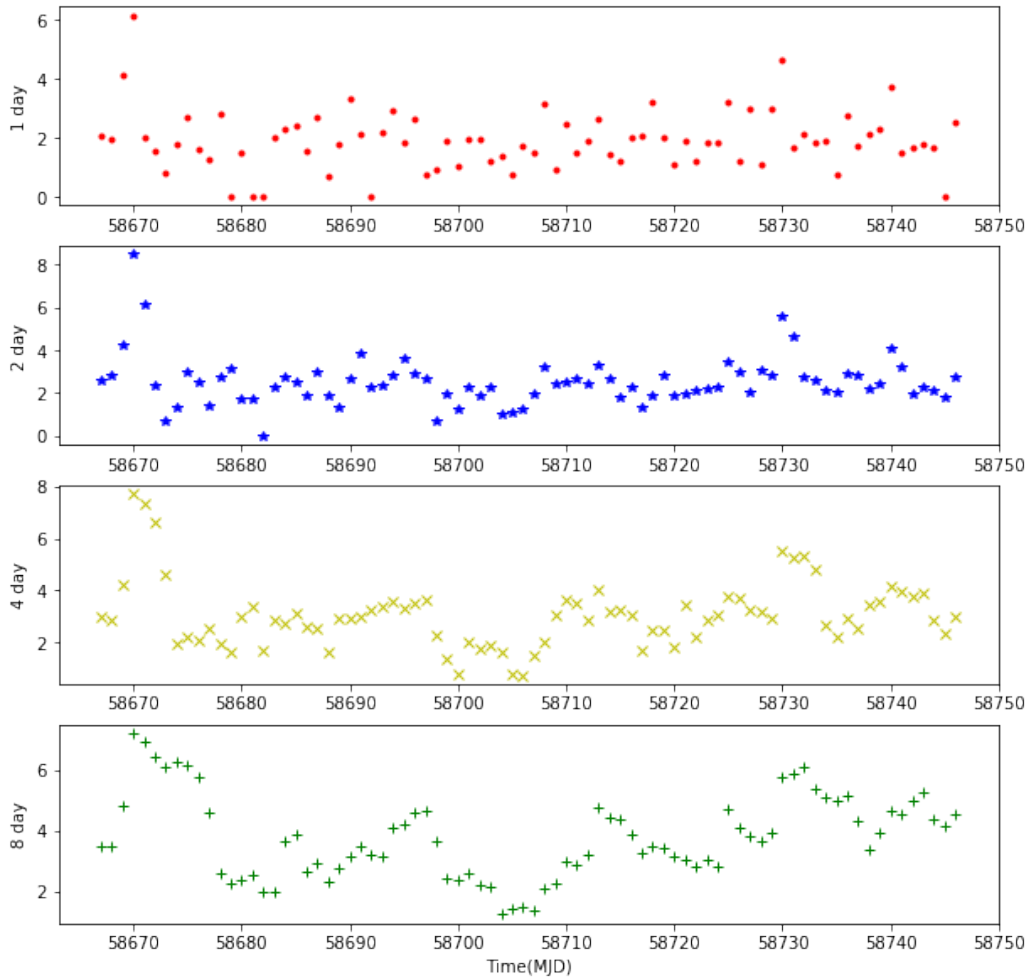
"real time" VHE gamma-ray flares



- 62 VHE blazars from TeVCat
- Crab: a direct check for the monitoring scheme



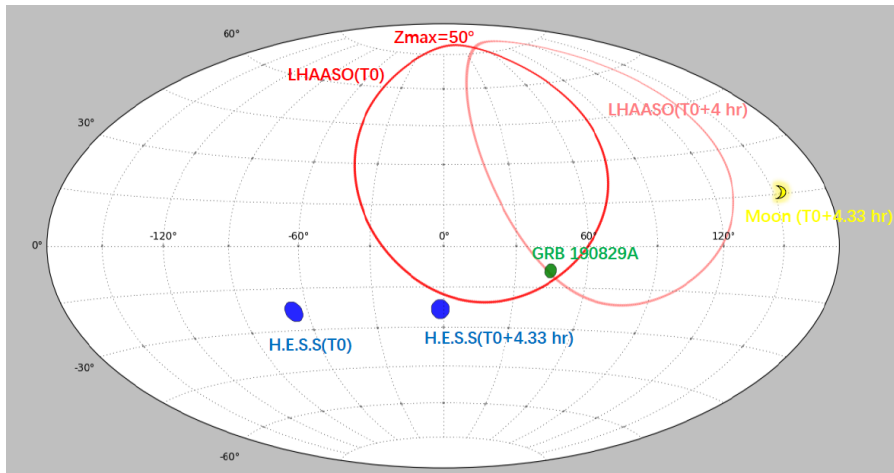
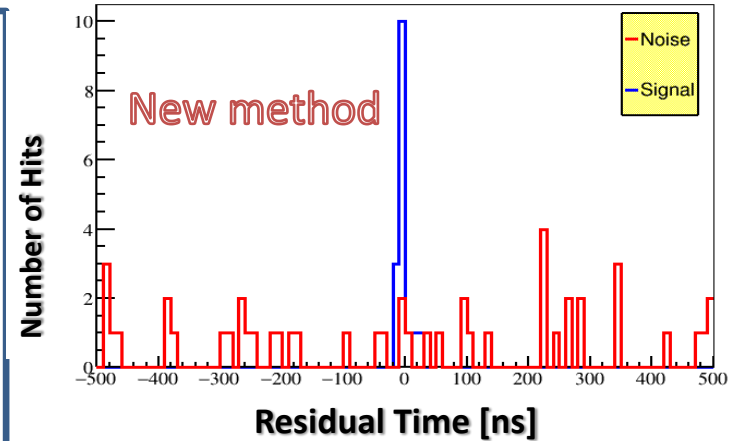
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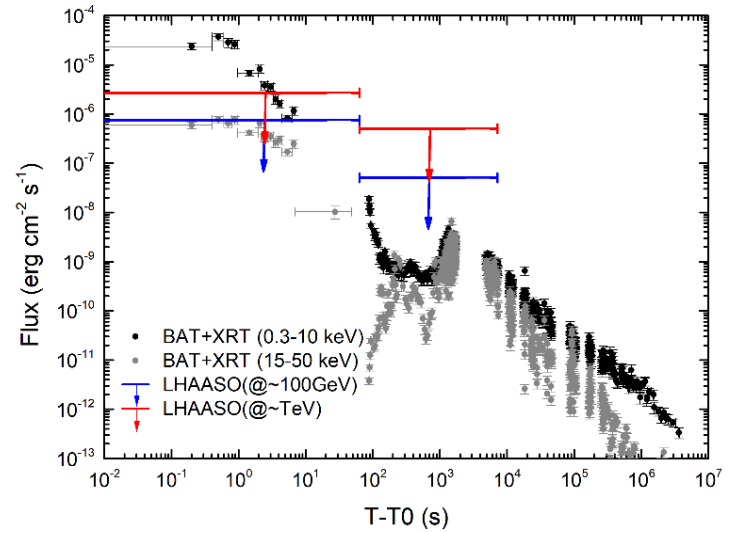
- Preliminary result
- Mrk421 flaring behavior
- Event selection:
 - Without LED signal;
 - $N_{fitc} > 100$
 - $Zenith < 60^\circ$
 - $Compactness > 15$

GRB follow-up

- Develop a new method to search for GRBs
- Set up GRBs Monitor software
- Carry out analysis of GRB190829A
- 36 GRB follow-ups (Since June 2019)
the statistical analysis is on the way.

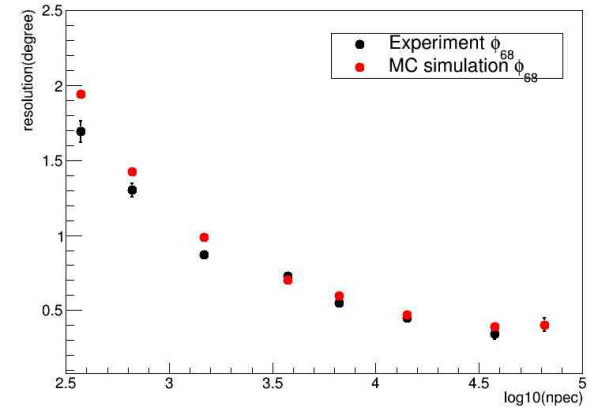
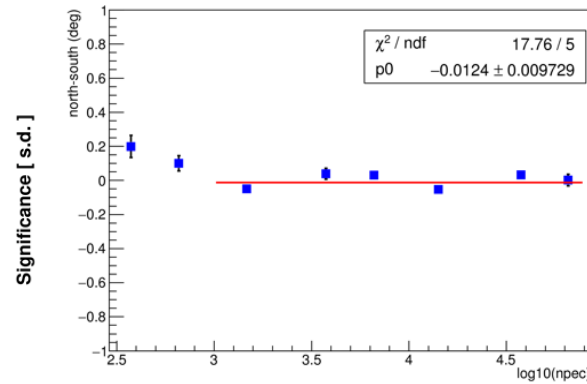
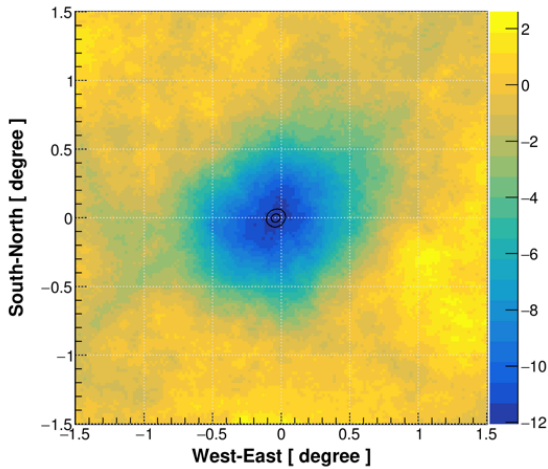


GRB190829A @ LHAASO-WCDA1

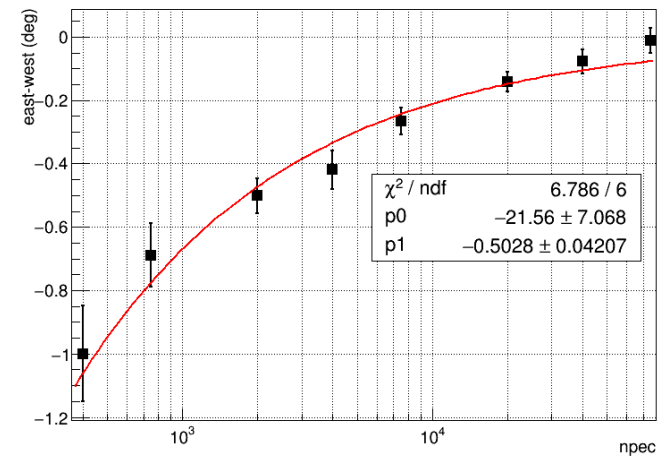
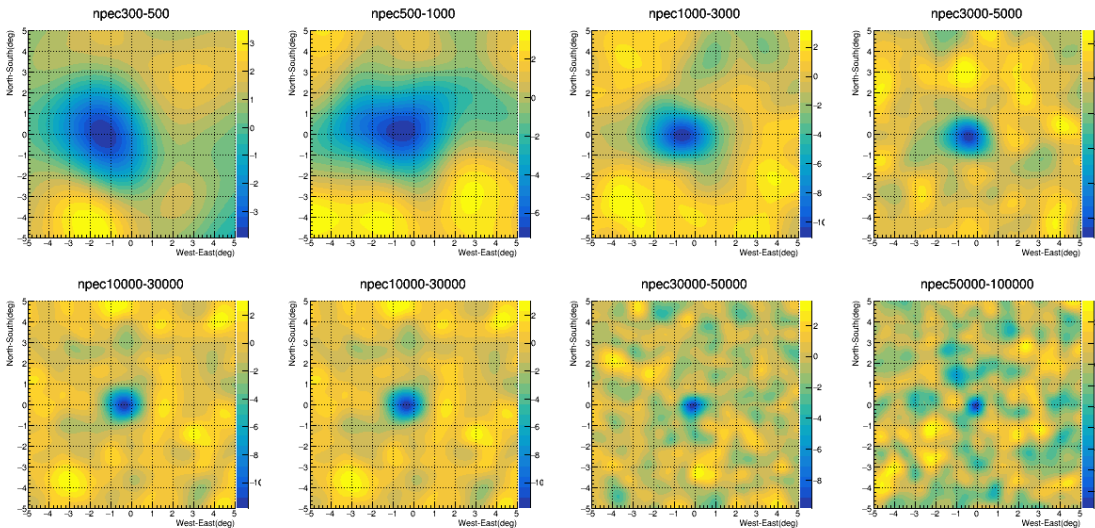


Moon shadow: Geomagnetic spectrometer

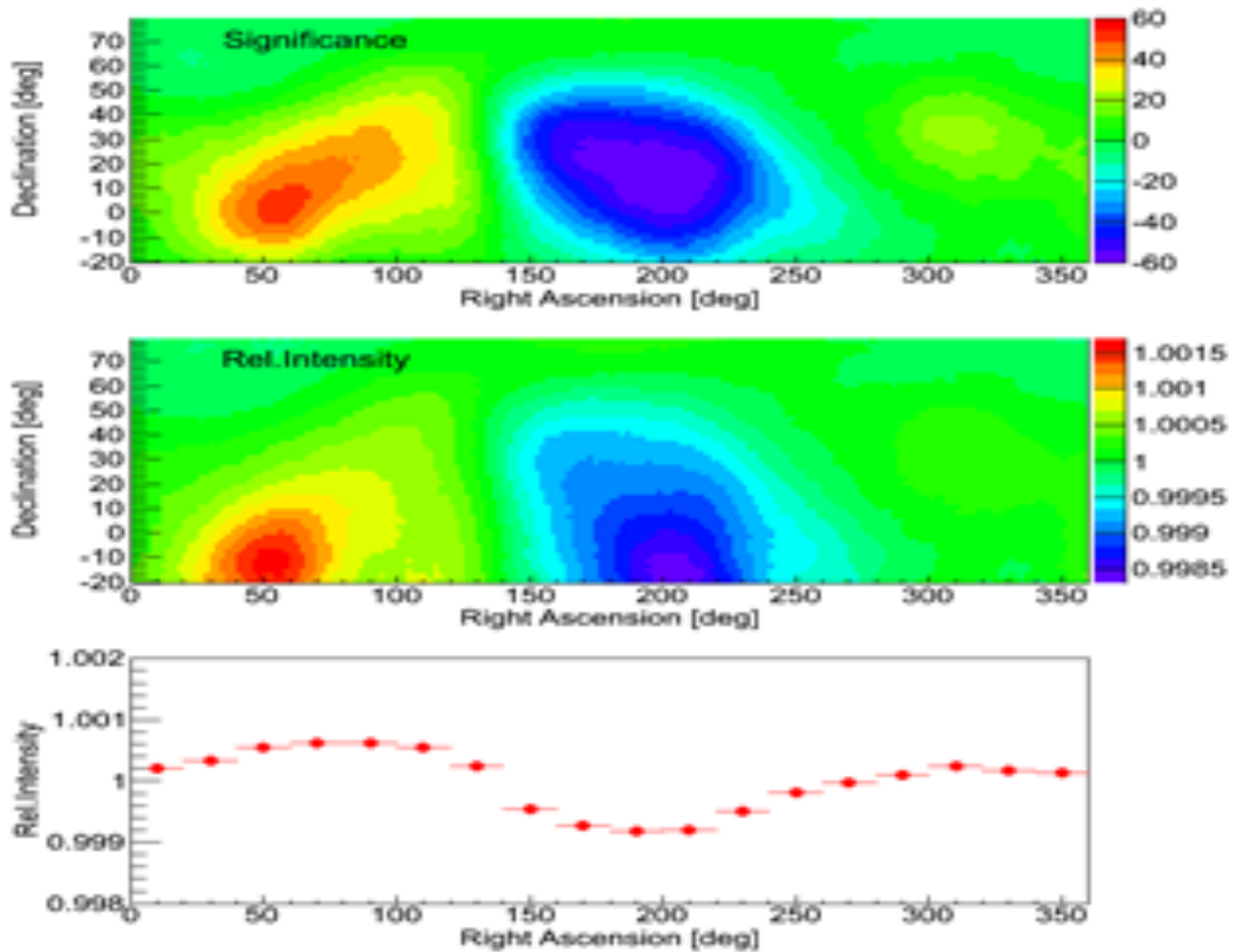
$N_{\text{fit}} > 300$



Pointing error $< 0.05^\circ$

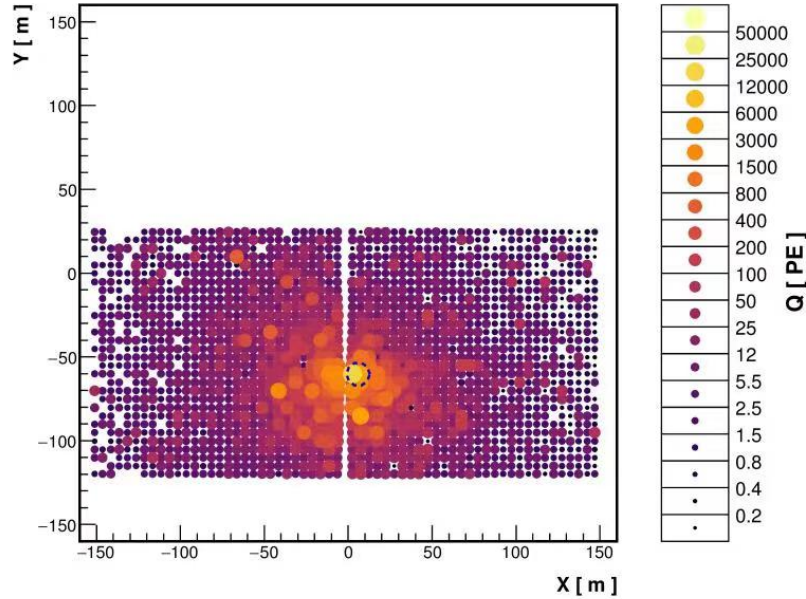


CR-anisotropy @ 120 days data

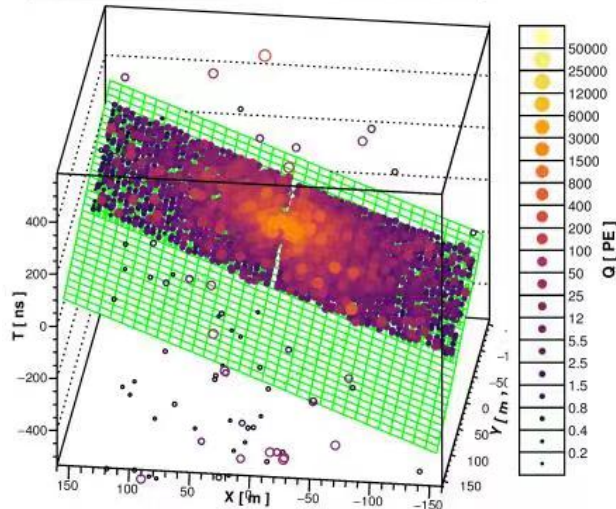


WCDA-1 + WCDA-2 = a full beautiful event

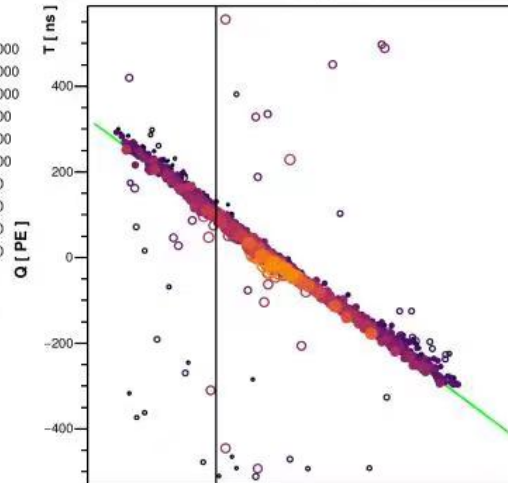
20191214/203329/0.155216835: nTrig=0, $\theta=32.60\pm 0.01^\circ$, $\phi=157.11\pm 0.03^\circ$



20191214/203329/0.155216835: nTrig=0, $\theta=32.60\pm 0.01^\circ$, $\phi=157.11\pm 0.03^\circ$



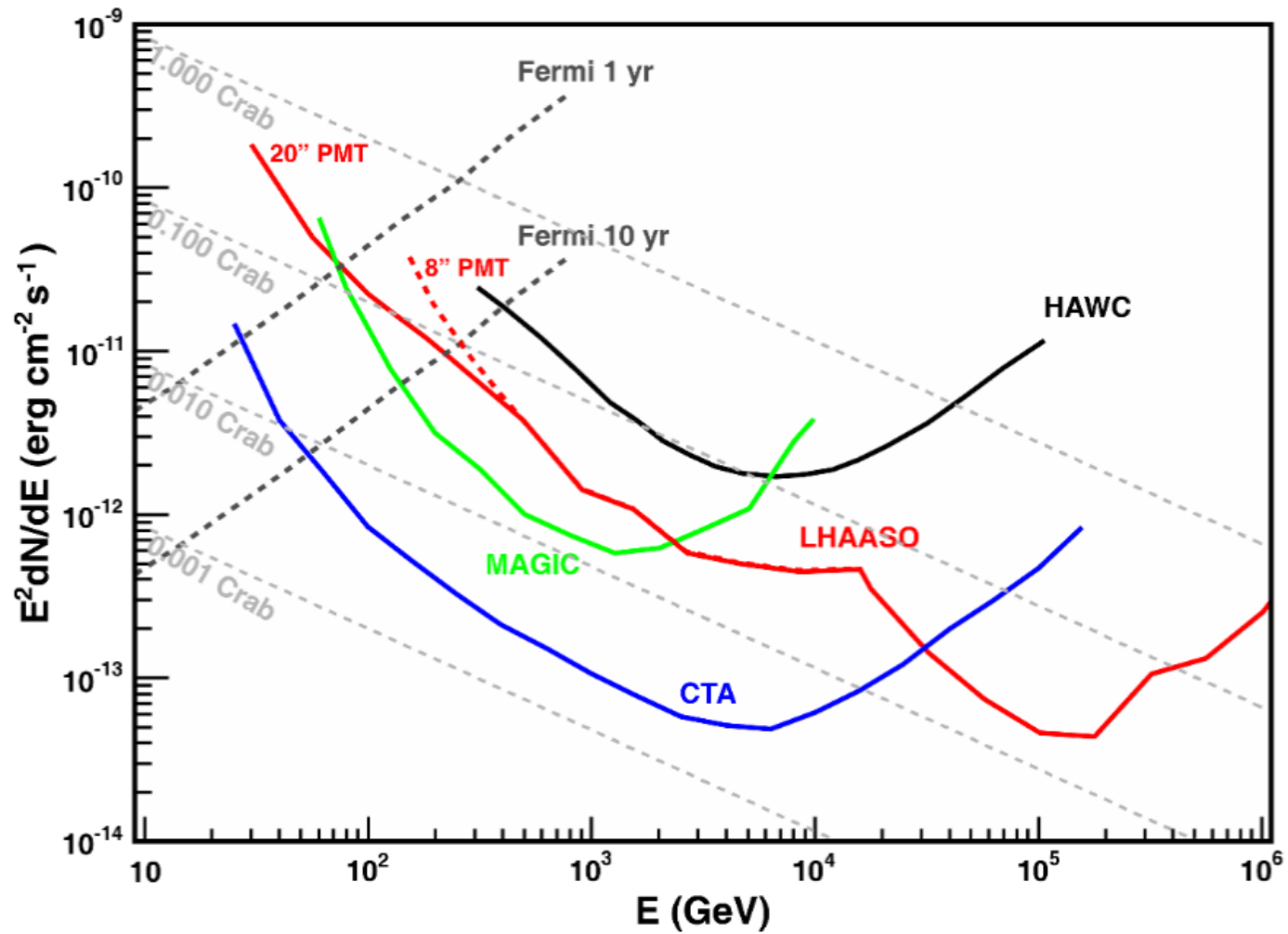
0/0 #33058: nHit=1800, nFit=1653, $\Delta\alpha=0.02^\circ$, $\chi^2=53354.3 / 1650$



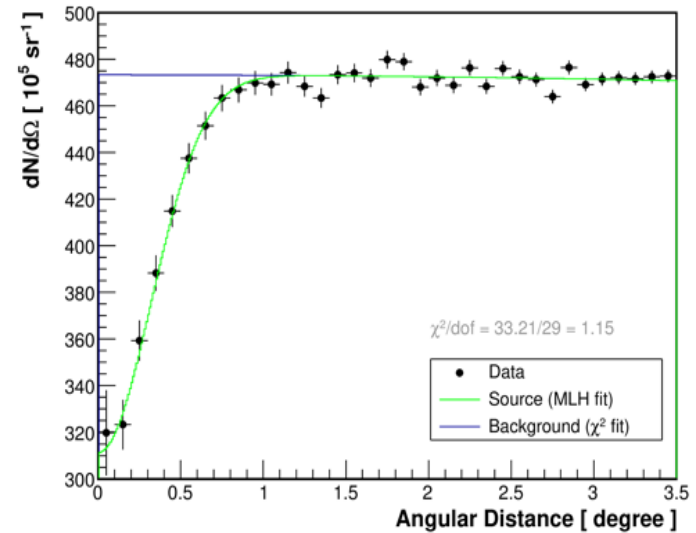
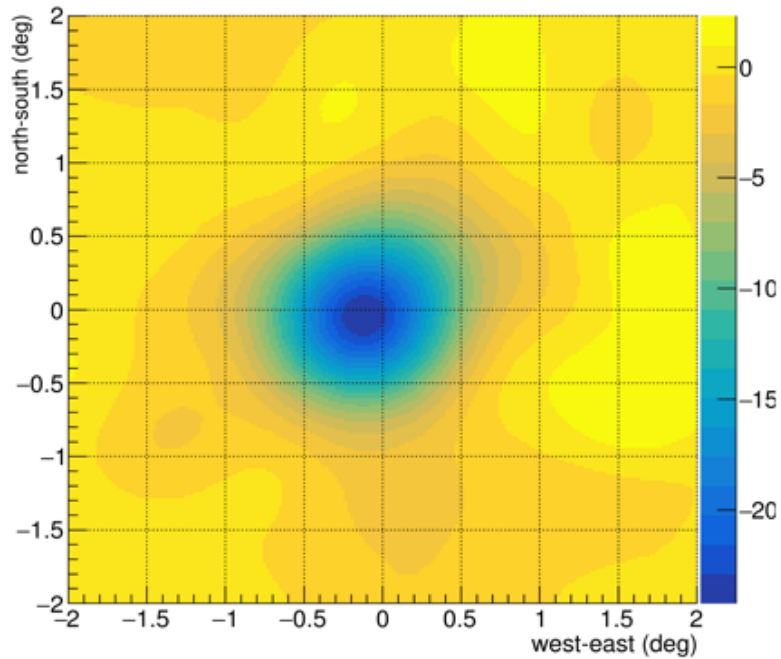
Summary and outlook

- WCDA-1 works smoothly in scientific operation mode;
 - Crab significance is around 20 sigma per month;
 - Moon shadow significance is around 20 sigma per month;
 - Pointing error is less than 0.1 deg;
 - Sky map, Moon/Sun shadow and CR anisotropy have been observed;
 - To understand the systematic error of the analysis;
- WCDA-2 begin to take physics data in Dec. 24th 2019;
 - Looking forward to news from 20-inch PMT
- At the end of 2020, WCDA-3 will be ready.

backup



Pointing error + angular resolution



1->2、3号水池的PMTs

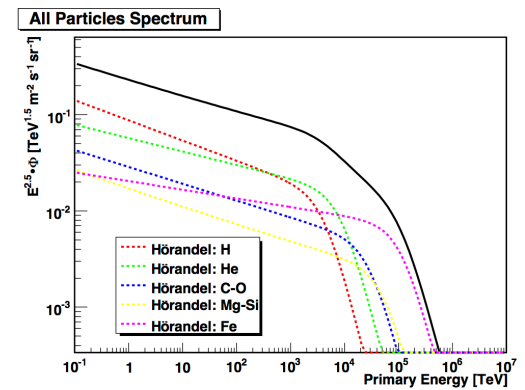
- 主要变化:



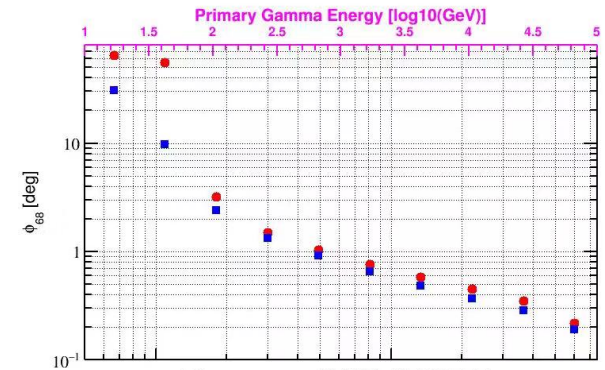
100GeV的 γ : 8inch eff=1.7% ; 20inch eff=10.5%

MC simulation details

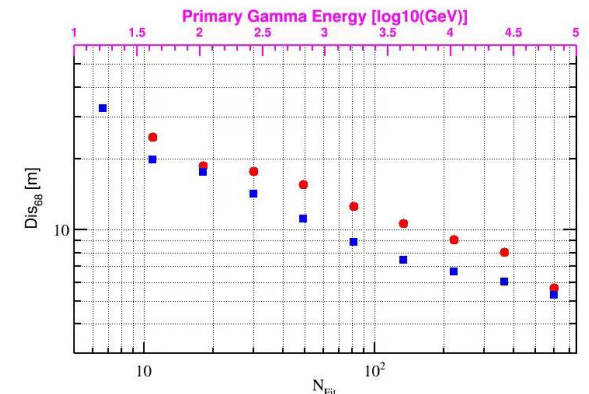
- ◆ CORSIKA v74000; total size 2 T
- ◆ Hadronic models:
 - ◆ FLUKA + EPOS-LHC
- ◆ PrimaryID:
 - ◆ Gamma ;
 - ◆ 5 groups (H/He/CNO/MgAlSi/Fe)
 - ◆ Flux from Horandel model;
- ◆ primary energy 10 GeV to 100 TeV / 1 PeV;
 - ◆ 5 decades: 1.e10-1.e11 / 1.e11-1.e12
 - ◆ 1.e12-1.e13 / 1.e13-1.e14 / 1.e14_1.e15
- ◆ Zenith angles: $0^\circ - 60^\circ$;
- ◆ Azimuth angle $0^\circ - 360^\circ$;
- ◆ Shower core projection area: 2000 X 2000 m2
- ◆ G4WCDA 4.1 + 4.2;
 - ◆ 11.e12-1.e13 eV: 3e7
 - ◆ Total size: 230 T
- ◆ Core resolution around 10m @100N_{fit} and angular resolution 0.5° @100 N_{fit}



Δ angle gamma@10GeV-100TeV



Δ Core gamma@10GeV-100TeV



WCDA-1 data taking

