

# RHIC上电磁探针研究进展与实验计划

杨驰

山东大学





# Outline

- Introduction
- Search for Breit-Wheeler pair production
  - Summary of the results on EM probes
    - Recent detector upgrades
      - Future plan
        - Summary

## Dileptons and real photons —— ideal electromagnetic probes

- Suffer no strong interaction
- Produced at all stages of the system evolution
- Bring production information to final state
- Can probe temperature of the system during its hottest phases
- Sensitive to electromagnetic fields

# **Fundamental Interactions: Light & Matter**



Plots from O. Pike

2021/8/17

# Why is the Breit-Wheeler Process So Elusive?

DECEMBER 15, 1934

<sup>,</sup> Breit

heeler

PHYSICAL REVIEW

VOLUME 46

### Collision of Two Light Quanta

G. BREIT\* AND JOHN A. WHEELER,\*\* Department of Physics, New York University

(Received October 23, 1934)

As has been reported at the Washington meeting, pair production due to collisions of cosmic rays with the temperature radiation of interstellar space is much too small to be of any interest. We do not give the explicit calculations, since the result is due to the orders of magnitude rather than exact relations. It is also hopeless to try to observe the pair formation in laboratory experiments with two beams of x-rays or  $\gamma$ -rays meeting each other on account of the smallness of  $\sigma$  and the insufficiently large available densities of quanta. In the considerations of Williams, however, the large nuclear electric fields lead to large densities of quanta in moving frames of reference. This, together with the large number of nucleii available in unit volume of ordinary materials, increases the effect to observable amounts. Analyzing the field of the nucleus into quanta by a procedure similar to that of v. Weizsäcker,<sup>4</sup> he finds that if one quantum  $h\nu$ 

# Breit and Wheeler mentioned it was hard, and they even introduced field generated by heavy ion.

Second Series

November 15, 1930

Vol. 36, No. 10

# PHYSICAL REVIEW

SCATTERING OF HARD  $\gamma\text{-}\mathrm{RAYS}$ 

BY C. Y. CHAO\* Norman Bridge Laboratory of Physics, California Institute of Technology (Received October 13, 1930)



赵忠尧先生1930年实际首次观 察到正负电子对的湮灭辐射

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# **Search for Breit-Wheeler Pair Production**

| Beam<br>+<br>Laser   | Gamma-ray (100 GeV)                     | Optical laser photon (eV)   | As numerous 100+ pe<br>planned or being built   | tawatt lasers facilities are<br>world-wide, a detailed |
|--|---|-----------------------------|---|--|
| Laser  |   |                             | Electrodynamics has b   | pecome more important                                  |
|  |   |                             | than ever.  |  |
| Beam<br>+  | ~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~ |                             | Recently proposed experiments   | for achieving the Breit-Wheeler process:               |
| Beam   | Gamma-ray (MeV)                         | Gamma-ray (MeV)             | <u>Nature Photonics</u> <b>8</b> , 434–436 (20<br><u>Nature Communications</u> <b>7</b> , Articl<br><u>Communications Physics</u> <b>1</b> , Articl | 14)<br>e number: 13686 (2016)<br>e number: 93 (2018)   |
| Beam<br>+  | ~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~ |                             |   |  |
| Target   | Gamma-ray (GeV)                         | Thermal x-ray field (100 eV | )   |  |
|  |   | National Ignition           | on Facility (USA)   | aser Megajoule (France)                                |
|  |   |                             |   |  |
|  |   |                             |   | Alter Alter  |
|  |   |                             |   |  |
|  |   |                             | VI in 192 hours   | g pulse: 1.4 ML in 176 hears                           |
|  |   |                             | P. Laberuerie   |  |
| SLAC E   | -144 experiment:                        | _                           | Orion Laser (UK)  | omega ep (USA)   |
| first sia  | n of positron production in             | n liaht-by-liaht            |   |  |
| scattering   |   |                             |   |  |
| Burke et al., Phys.Rev. Lett. 79, 1626 (1997)<br>Hu & Müller, Phys.Rev. Lett. 107, 090402 (2010) |   |                             |   |  |
|  |   |                             | Long-pulse: 5 kJ in 10 beams  | Long-pulse: 40 kJ in 60 beams                          |
|  |   | ©                           | British Crown Owned Copyright / AWE   | University of Rochester                                |

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# **Ultra-Peripheral Collisions Provide Opportunities**

## Main challenges in detection:

- Peak of the same order of that in Compton scattering and Dirac annihilation.
- Center-of-mass energy > 2m<sub>e</sub>
- High photon density need.





## UPC of heavy ions solve the central challenges:

- Highly Lorentz contracted EM field, quantized as a flux of quasi-real photons
- |B|~10<sup>14-16</sup>T
- High photon density:  $Z\alpha \sim 1$
- Z<sup>4</sup>~40,000,000 times higher yield than collisions of single charged particle pairs
- Clean background, no hadronic collisions

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# **Particle Identification**



## Chinese STAR group contribute to the eID/muID:

- TPC: FDU+SDU+SINAP
- TOF: FDU+SINAP+THU+USTC



## Electron identification: TPC+TOF

Phys. Rev. Lett. 94, 062301 (2005)

## Muon identification: TPC or TPC+MTD

J.D.Brandenburg, Nucl. Phys. A 982, 192 (2019) T.C.Huang, et.al., NIM.A 833, 88-93 (2016)

# **Particle Identification**



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# Search for Breit-Wheeler Process at RHIC-STAR

Phys. Rev. Lett. 127, 052302 (2021)

**1.** Observe 6085 exclusive e<sup>+</sup>e<sup>-</sup> pairs from data collected in 2010 at STAR.



2. No vector meson contribution visible

## 3. Energy spectrum

# 4. Photon transverse polarization & spatial distribution



https://www.bnl.gov/newsroom/news.php?a=119023

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# **Cross-Section in STAR Acceptance**

**STAR**: Au+Au at  $\sqrt{s_{NN}}$  = 200 GeV,  $|y^{ee}| < 1$ ,  $P_{\perp} < 0.1$  GeV,  $P_T^e > 0.2$  GeV,  $|\eta^e| < 1$ .



 $\sigma(\gamma\gamma \rightarrow e^+e^-)$  in STAR Acceptance: 0.261  $\pm$  0.004(stat.)  $\pm$  0.013(sys.)  $\pm$  0.034(scale) mb STARLight: 0.22mb Generalized EPA: 0.26mb QED: 0.26mb

## Measured cross-section consist with various theory calculations within $\pm$ 1 $\sigma$

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# Polar Angle of the e<sup>+</sup> Momentum and the Beam Axis



 $\theta'$ : angle of the e<sup>+</sup> momentum w.r.t the beam axis in the pair rest frame

- Breit-Wheeler process exhibits an enhancement toward a small polar angle
- Highly virtual photon interactions should have an isotropic distribution

 $\gamma\gamma 
ightarrow e^+e^-$  : Individual e<sup>+</sup>/e<sup>-</sup> preferentially aligned along beam direction

S. Brodsky, T. Kinoshita and H. Terazawa, Phys. Rev. D 4, 1532 (1971)

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# e<sup>+</sup>e<sup>-</sup> Pair Transverse Momentum Distribution



60%-80%: Phys. Rev. Lett. 121, 132301 (2018)

|  | Ultraperipheral |      |      | Peripheral |              |      |
|--|-----------------|------|------|------------|--------------|------|
|  | Measured        | QED  | SC   | SL         | Measured     | QED  |
| $\sqrt{\langle P_{\perp}^2 \rangle}$ (MeV) | $38.1\pm0.9$    | 37.6 | 35.4 | 35.9       | $50.9\pm2.5$ | 48.5 |

- Peak at very low p<sub>T</sub>
- Impact parameter dependence

# **Polarized Photon-Photon Collisions**

# Gluons are highly linearly polarized.

Metz & 周剑, 2011

# **CGC** is highly linearly polarized state.

| <u>QCD:</u>                              | <u>QED:</u>                        | <ul> <li>By direct analogy, in extreme QED fields, the linear</li> </ul> |
|--|------------------------------------|--|
| $gg  ightarrow q \overline{q}$           | $\gamma\gamma  ightarrow l^+l^-$   | photon polarization distribution can be measured                         |
| $\Delta \phi = \phi^{q\bar{q}} - \phi^q$ | $\Delta \phi = \phi^{ll} - \phi^l$ | via $\cos(2\Delta\phi)$ and $\cos(4\Delta\phi)$ modulations              |

## The same analysis applies to the QED case.

# **Azimuthal Modulations Predicted for Real Photon**



|                         | Ultr         | Ultraperipheral |    |    | Peripheral |      |
|-------------------------|--------------|-----------------|----|----|------------|------|
|                         | Measured     | QED             | SC | SL | Measured   | QED  |
| $ A_{4\Delta\phi} $ (%) | $16.8\pm2.5$ | 16.5            | 19 | 0  | $27\pm 6$  | 34.5 |
| $ A_{2\Delta\phi} $ (%) | $2.0\pm2.4$  | 0               | 5  | 5  | $6\pm 6$   | 0    |

## Results at STAR are well consistent with the theory prediction.

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# Link to Vacuum Birefringence



Maybe more closely related to the vacuum dichroism? *K. Hattori, H. Taya, and S. Yoshida, JHEP 01,93 (2021)* 

# Coherent Low $p_T e^+e^-$ in Au+Au and U+U

Phys. Rev. Lett. 121, 132301 (2018)



Dimuon channel: Jian Zhou's talk at 11:30 Tue.

- Significant enhancement
- Excesses concentrate below p<sub>T</sub> ≈ 0.15 GeV/c
- Data are consistent with hadronic expectation when  $p_T > 0.15$  GeV/c
- Coherent photon-photon and photon-nucleon interactions in HHIC

# We Proceed as We Planned

# In UPC

ChiYang, ISMD2019 talk, Santa Fe

- Provide baseline
- Exclusive and clean e<sup>+</sup>e<sup>-</sup> detection
- Individual e should be preferentially aligned longitudinally and the azimuthal angle between the e<sup>+</sup>e<sup>-</sup> pair momentum and the individual e momenta should display modulations

# In HHIC

- Study extreme magnetic field and potential medium effect
- A+A centralities and different species (A+B, STAR isobaric data)
- Kinematics WRT B-field (reaction plane), mass, p<sub>T</sub>, longitudinal alignment

The energy spectrum and angular distribution provide the information needed to map the spatial extent of the intense electromagnetic fields produced by ultrarelativistic heavy nuclei for the first time.

Can be compared to different models (*Q.Y.Shou,Y.G.Ma, P.Sorensen, A. H.Tang, F. Videbæk, and H. Wang, Phys.Lett. B* 749, 215 (2015)) which incorporate the charge distribution with a Lorentz boost.

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# **Dielectron in Beam Energy Scan Phase I**



• Low mass excesses are consistent with  $\rho$  broadening scenario from RHIC top energy down to 19.6GeV



- Emission rate is dominant in the T<sub>c</sub> region
- More clear pictures of the excess versus lifetime and total baryon density in BES-II

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# **Direct Virtual Photon at RHIC Top Energy**



## **Model calculations:**

- Consistent with the yield within uncertainties except some bins in 6o-8o%
- Simultaneously describe both dielectron and direct virtual photon yields

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# **Summary of Dielectron and Direct Photons at STAR**

## One decade passed, on dileptons and direct photons as EM probes:

3 PRL(200GeV A+A Dielectron, Low pT Peripherial Dielectron, Breit-Wheeler Process) 2 PLB(19.6GeV A+A Dielectron, 200GeV A+A Direct Photon)

 $_{3}$  PRC(200GeV p+p Dielectron, 200GeV A+A, Dielectron v<sub>2</sub>)

## Achieved:

- ✓ Rho broadening observed
- ✓ Possible IMR excess observed
- ✓ Thermal photon observed
- ✓ Photon-photon interaction in HHICs
- ✓ Breit-Wheeler pair production in UPCs

## Target/Focus:

chiral symmetry restoration

**QGP** thermal radiation

*medium temperature* 

QGP in extreme EM field

QED under extreme EM field

Main challenges: Low S/B, Hadron spectrum, Statistics, PID

Our TOF and MTD enables analysis at STAR with the lepton channels.

**Excellent STAR detectors push us moving forward!** 

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## **Detector Upgrades for Beam Energy Scan Phase II at RHIC**



| iTPC upgrade  | EPD upgrade  | eTOF upgrade  |  |
|---|--|---|--|
| Continuous pad rows<br>Replace all inner TPC sectors  | Replace Beam Beam Counter  | Add CBM TOF modules and electronics<br>(FAIR Phase o) |  |
| η <1.5 (was 1.0)                                      | 2.1< η <5.1  | -1.6<η<-1.1   |  |
| p <sub>T</sub> >60 MeV/c (was 150MeV/c)               | Better trigger & b/g reduction   | Extend forward PID capability                         |  |
| Better dE/dx resolution<br>Better momentum resolution | Greatly improved Event Plane info (esp. 1 <sup>st</sup> -<br>order EP) | Allows higher energy range of Fixed Target program    |  |
| Fully operational in 2019                             | Fully operational in 2018  | Fully operational in 2019                             |  |

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# **New Detectors at STAR Forward Rapidity**



## FST, 3 Silicon disks: at 146, 160, and 173 cm from IP

Built on successful experience with STAR IST

- Single-sided double-metal mini-strip sensors
  - ✓ Granularity: fine in f and coarse in R
  - ✓ Si from Hamamatsu
- Frontend chips: APV25-S1 → IST all in hand
- Reuse IST DAQ system and cooling system

| Detector | pp and pA                               | AA  |
|----------|---|---|
| ECal     | ~10%/√E                                 | ~20%/√E   |
| HCal     | ~50%/√E+10%                             |   |
| Tracking | charge separation<br>photon suppression | 0.2 <p<sub>T&lt;2 GeV/c<br/>with 20-30% 1/p<sub>T</sub></p<sub> |

## FCS: 7 m from the IP

**ECal:** reuse PHENIX SHASHLYK 1496 Ch.

• Lateral tower Size  $5.5 \times 5.5 \times 33 \text{ cm}^3$  (18X<sub>o</sub>)

HCal: Fe/Sc (20mm/3 mm) sandwich 520 Ch.

- Lateral tower size 10 x 10 cm<sup>2</sup>, ~ 4.5 $\lambda$
- ✓ in close collaboration with EIC R&D

#### **Preshower:**

• Existing EPD, with additional splitter

## FTT, 4 sTGC disks: at 307, 325, 343 and 361 cm from IP

- location inside Magnet pole tip opening
  - ✓ inhomogeneous magnetic field
- 4 quadrants double sided sTGC → 1 disk
  - ✓ sTGC technique developed by ATLAS
- Position resolution: ~200 um
- Readout: based on VMM-chips

STAR note o648: <u>https://drupal.star.bnl.gov/STAR/starnotes/public/sno648</u>

# Forward sTGC Tracker, fSTAR

From 2018 till now:

- ✓ Three versions of module prototypes
- $\checkmark$  Three versions of electronics prototype
- ✓ 25 modules in mass production

30 x 30 cm<sup>2</sup> prototype



60 x 60 cm<sup>2</sup> prototype



Module Production: SDU



Integrations & DAQ: BNL

Commissioning & software: BNL, SDU

| Detector                  | Produced        | Shipped         | Installed       |
|---------------------------|-----------------|-----------------|-----------------|
| 1 <sup>st</sup> prototype | Oct.2018        | Jan.2019        | Jun.2019        |
| 2 <sup>nd</sup> prototype | Jan.2019        | Jul.2020        | May 2021        |
| 3 <sup>rd</sup> prototype | Oct.2020        | N/A             | N/A             |
| Final modules             | <u>May 2021</u> | <u>Jun.2021</u> | <u>Sep.2021</u> |

#### 55 x 55 cm<sup>2</sup> pentagon







Electronics: USTC



Joint test at SDU in Sep.2020.

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# **Organizational Structure in STAR Forward Upgrade**



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# The Physics at RHIC beyond 2021+

## **Opportunities in RHIC 2021+ on dileptons and photons**

#### Mid-rapidity -1.5<η<1.5

#### A+A

#### Beam:

Full Energy (200 GeV) Au+Au

#### Physics Topics:

a deep look into the properties of the QGP:

- γ & e+e- pairs
- Chiral symmetry restoration
- Temperature and lifetime of hot, dense medium
- Lower momentum  $\pi$ , K, p spectra
- Hypertriton Lifetime
   Measurement
- Precision measurements of direct photon yields and vn

#### p+A, p+p

#### Beam:

500 GeV: p+p 200 GeV: p+p and p+A

#### Physics Topics:

Improve statistical precision:

- TMD measurements, i.e. Collins, Sivers, ...
- Access s & Ds through Kaons in jets
- Measurement of GPD  $E_g$  through UPC J/ $\!\Psi$
- First access to Wigner functions through di-jets in UPC
- Gluon and quark vacuum
   fragmentation
- Gluon and quark fragmentation in nuclear medium
- Nuclear dependence of Collins FF

2019 Event Display : Au+Au 19.6 GeV Full tracking with all iTPC sectors

#### Forward-rapidity 2.8< η <4.2

#### A+A

#### Beam:

#### Full Energy AuAu (2023/25)

#### Physics Topics:

- Temperature dependence of viscosity through flow harmonics up to η ~4
- Longitudinal decorrelation up to η ~4
- Strong rapidity dependence in Global Lambda Polarization

#### p+A, p+p

Beam: 500 GeV: p+p 200 GeV: p+p and p+A

#### Physics Topics:

- TMD measurements at high x

   o transversity → tensor charge
   o Sivers through DY, direct γ and tagged jets
- Gluon PDFs for nuclei
- R<sub>pA</sub> for direct photons & DY, and hadrons
- Test of Saturation predictions through di-hadrons,  $\gamma$  -Jets, di-jets



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## **Current and Future Opportunities on EM Probes at RHIC**

Au+Au 54 GeV and Au+Au 27 GeV data collected in Run 17 -- Smaller heavy flavor semi-leptonic decay in IMR

Isobaric collision data collected in Run 18

-- Same A, different Z – similar hadronic interaction, different initial magnetic field

BES-II in Run 19, Run 20 and Run 21

-- Different total baryon densities, scan in high total baryon density region. Connect with NICA-MPD EM probes research

## Beyond 2021+

-- Mid-rapidity: e<sup>+</sup>e<sup>-</sup> measurement at μ<sub>B</sub>~o Link to chiral symmetry restoration

-- Precise p spectra

Thermal radiation from QGP

-- IMR slope

Direct virtual photon at RHIC top energy

- -- Thermal photon slope
- Photon-photon collisions

-- EM field

 $\frac{1}{M_{ee}} \left( \frac{1}{\text{GeV/c}^2} \right)^{\frac{1}{4}}$ Uncertainty prediction with iTPC + 4 billion events

杨驰,第十三届全国粒子物理学术会议,山东大学,青岛

# $here (GeV/c^2)$

BNL, RICE, SDU, USTC

Need man power

# From RHIC to EIC

## Current and future plans at RHIC provide opportunities in:

### Detector R&D with techniques potentially used in EIC

- HCal+SiPM readout same as EIC-fHCal (joint STAR EIC R&D)
- Silicon technique for EIC tracker
- sTGC technique for EIC trigger/tracker
- ✓ Help to realize the scientific promise of the EIC
- Inform the physics program
- Quantify experimental requirements
- Especially on photon-nucleus interaction
- Train the young talents especially on detector R&Ds for EIC
- Several tens of the graduate/undergraduate students working on fSTAR





# Summary I

- Plenty of important results have been achieved from RHIC-STAR based on electromagnetic probes from the passed decade.
- Recent research extend the study on EM probes to fundamental QED process. The photoproduction study is crucial for the understanding of various topics:
  - ✓ Initial EM field in HICs
  - ✓ QED under extreme EM field
  - ✓ QGP under extreme EM field
  - ✓ Interdisciplinary with astronomy, laser physics, quantum magnetohydrodinamics, ...

# Summary II

- Chinese STAR group have made and are making crucial and significant contributions to most of the key sub-systems at STAR.
- These sub-systems support our physics program precisely!



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