

Spin Hydrodynamics and local spin polarization in hydrodynamic approaches

Shi Pu

**Department of Modern Physics,
University of Science and Technology of China**

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Outline

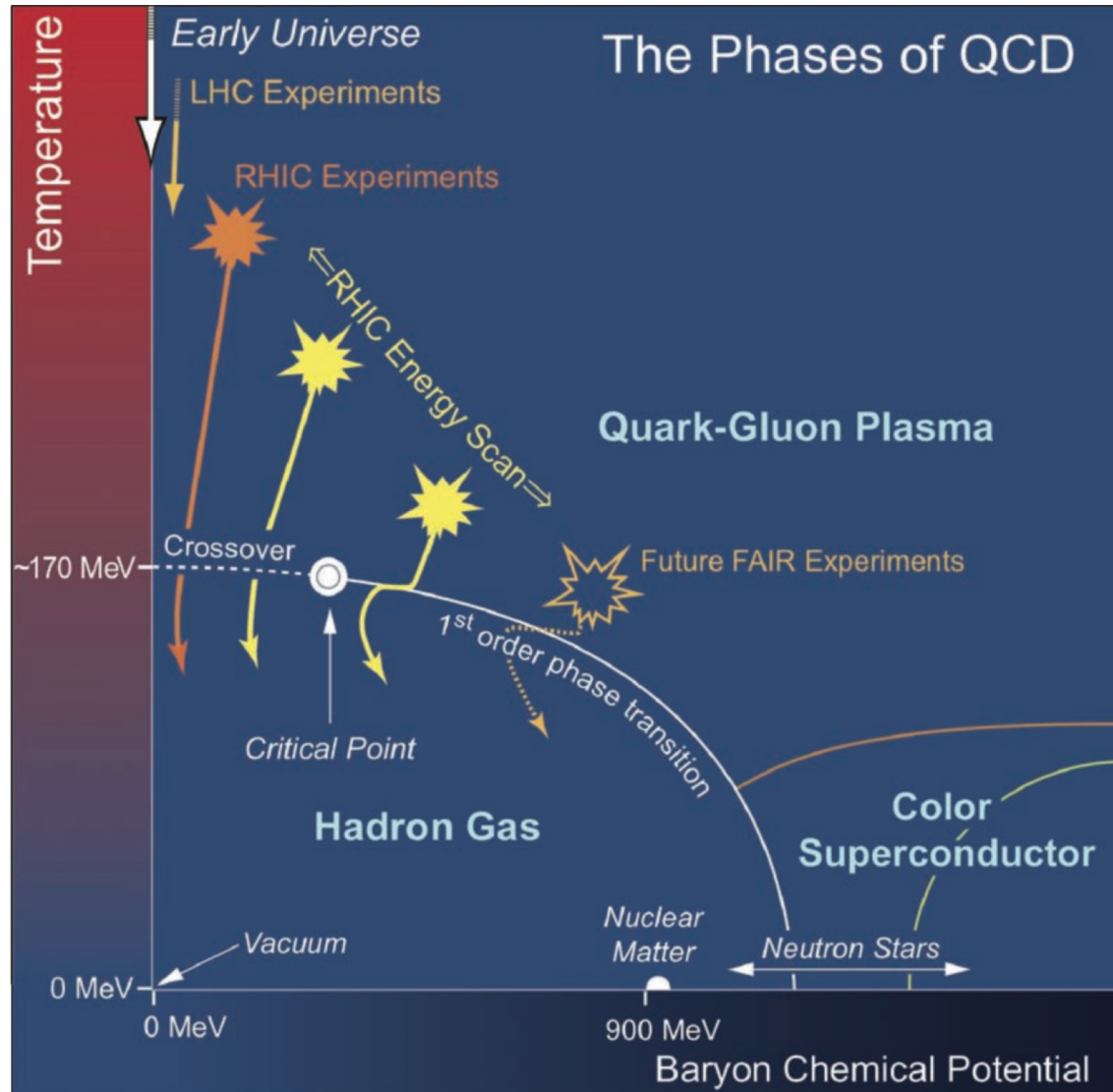
- **Introduction**
- **Belinfante form of spin hydrodynamics**
- **Local spin polarization in hydrodynamical approach**
- **Summary**

Introduction

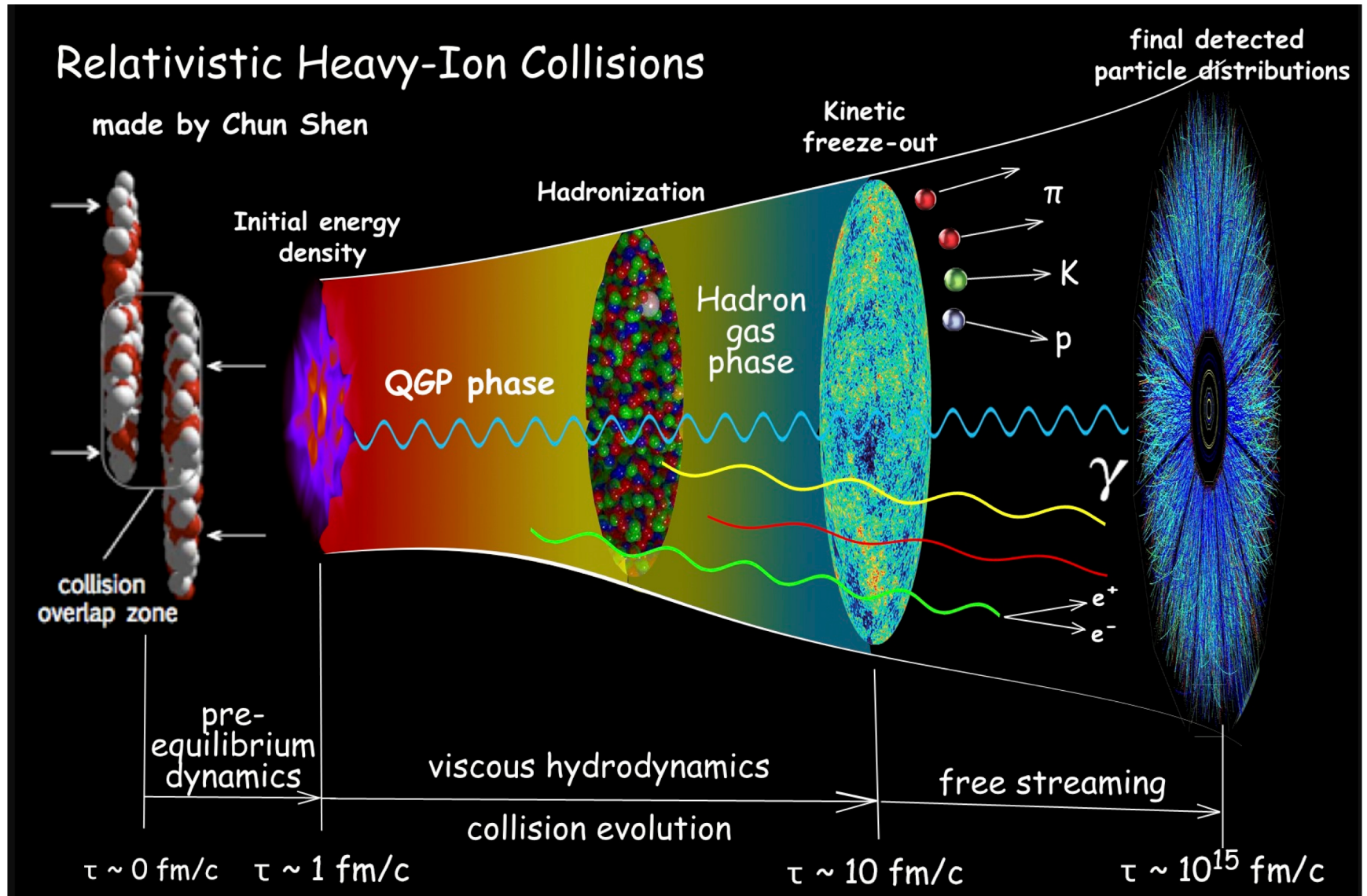
Phases of QCD



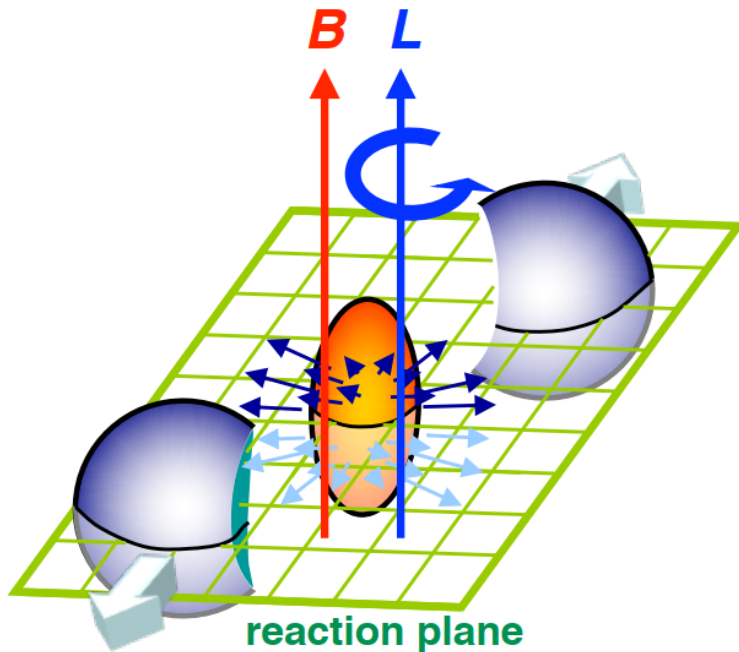
**T.D. Lee (1974) and Collins (1975):
Heavy ion collision to create a new
form of matter!**



Relativistic heavy ion collisions



Huge angular momentum

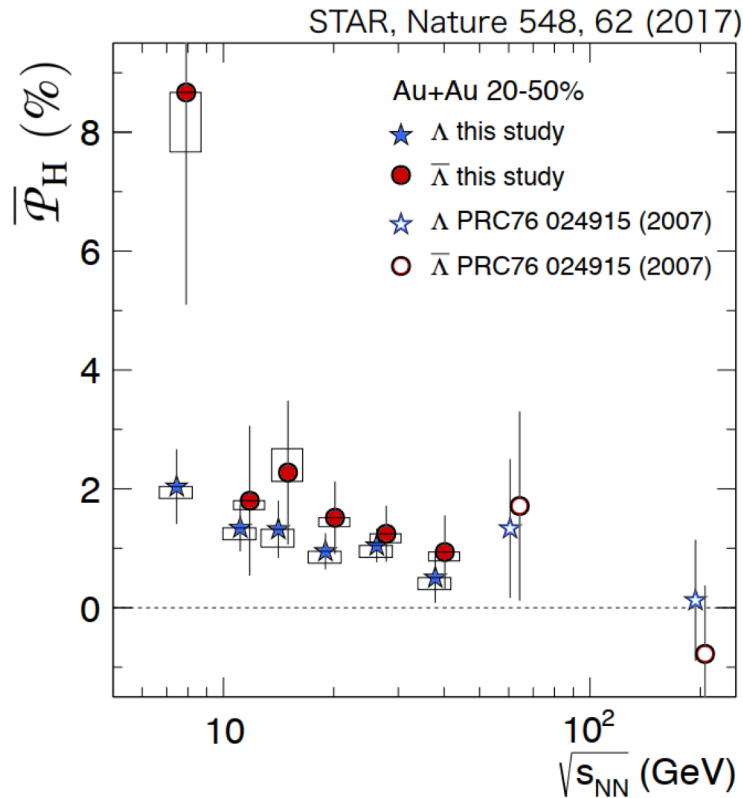


- Huge global orbital angular momenta are produced

$$L \sim 10^5 \hbar$$

- How do orbital angular momenta be transferred to the matter created?

Global Polarization of Λ and $\bar{\Lambda}$



- $\sqrt{s_{NN}} < 62.4\text{GeV}$, we observe the signal for polarization of Λ and $\bar{\Lambda}$
- The lower energy, the stronger polarization effects
- $P_{\bar{\Lambda}} > P_{\Lambda}$

$$P_{\Lambda} \approx \frac{1}{2} \frac{\omega}{T} + \frac{\mu_{\Lambda} B}{T}$$

$$P_{\bar{\Lambda}} \approx \frac{1}{2} \frac{\omega}{T} - \frac{\mu_{\Lambda} B}{T}$$

$\omega = (9 \pm 1) \times 10^{21}/\text{s}$,
greater than previously
observed in any system.

Liang, Wang, PRL (2005)

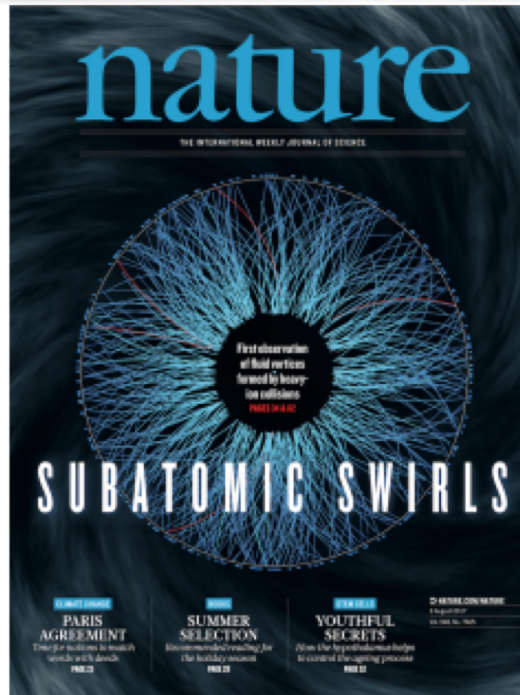
Betz, Gyulassy, Torrieri, PRC (2007)

Becattini, Piccinini, Rizzo, PRC (2008)

Becattini, Karpenko, Lisa, Upsal, Voloshin, PRC (2017)

Fang, Pang, Q. Wang, X. Wang, PRC (2016)

The fastest fluid

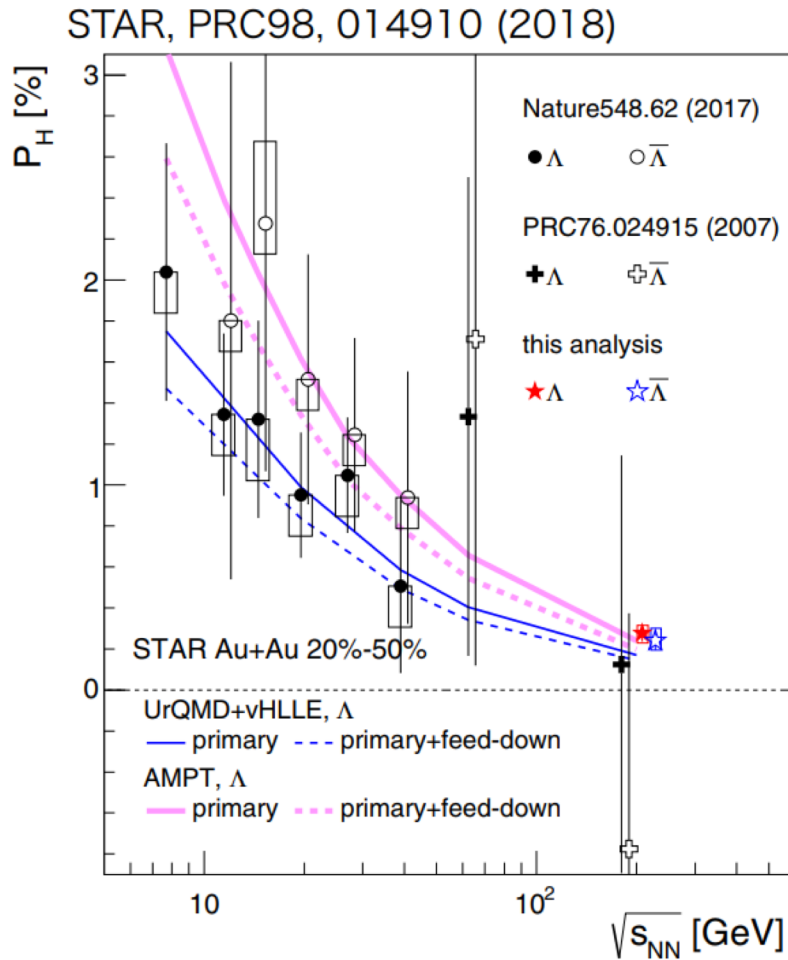


The Fastest Fluid

by Sylvia Morrow

Superhot material spins at an incredible rate.

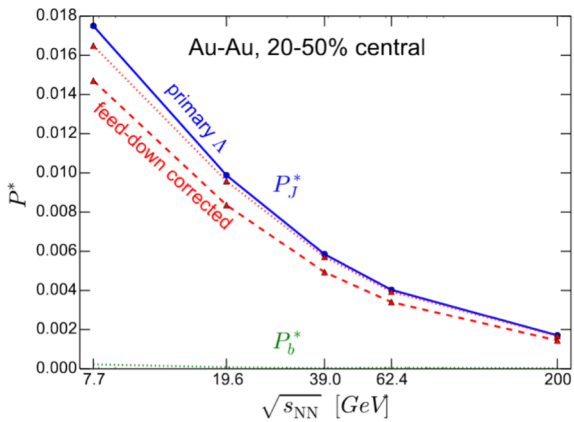
Global Polarization



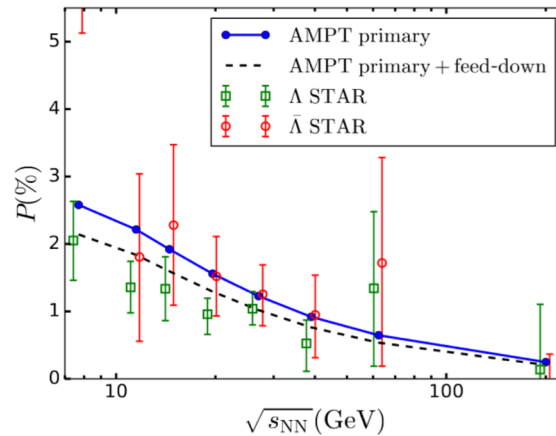
The results from both UrQMD+hydro and AMPT are consistent with the experimental data.

- **UrQMD+vHLLC: Karpenko, Becattini, EPJC(2017)**
- **AMPT: Li, Pang, Wang, Xia, PRC (2017)**

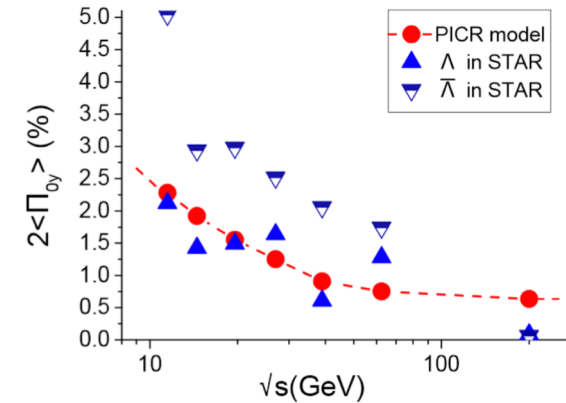
Global Polarization from different models



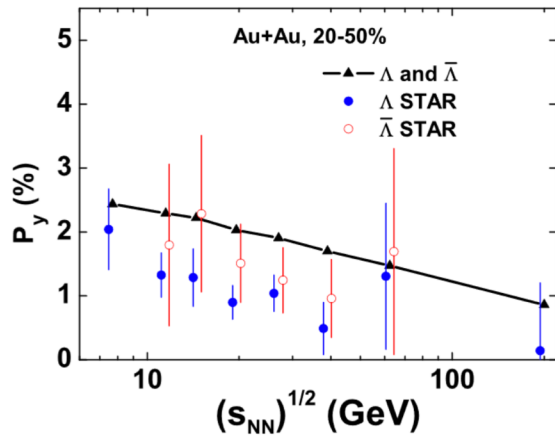
Karpenko, Becattini, EPJC(2017)



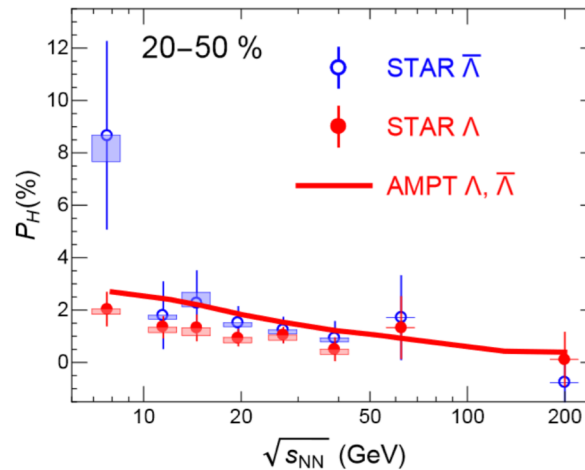
Li, Pang, Wang, Xia PRC(2017)



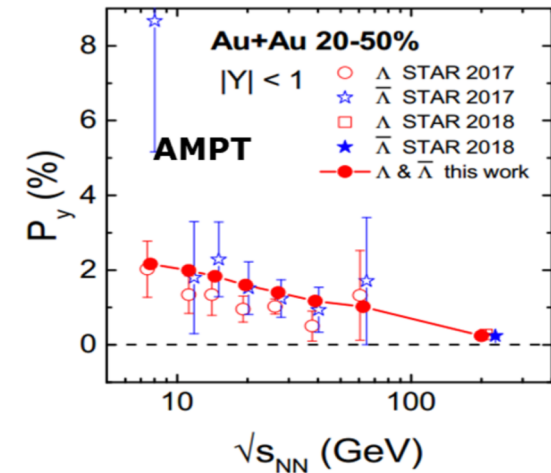
Xie, Wang, Csernai, PRC(2017)



Sun, Ko, PRC(2017)

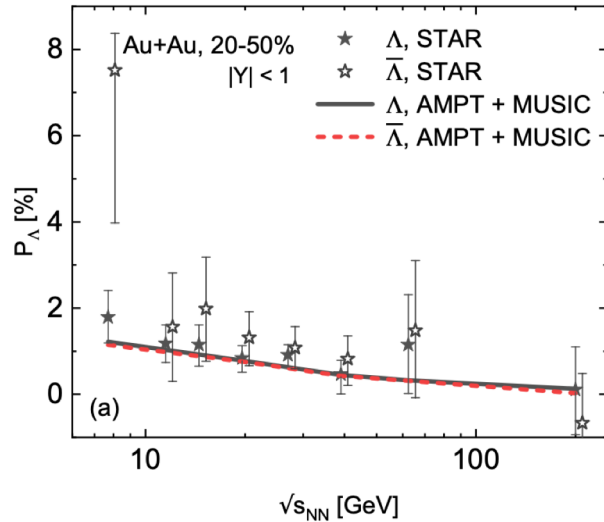


Shi, Li, Liao, PLB(2018)

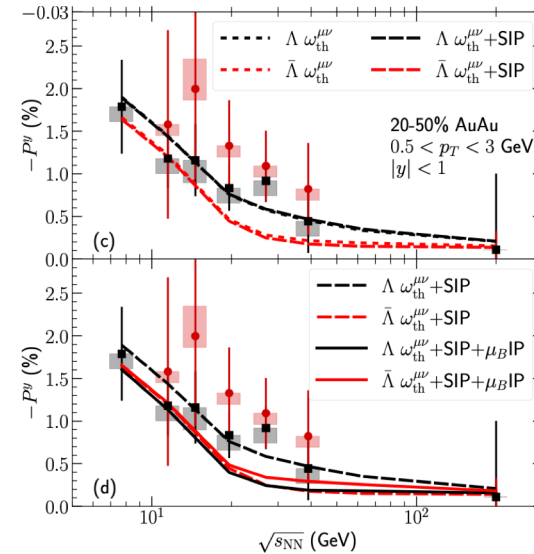


Wei, Deng, Huang, PRC(2019)

Global Polarization from different models

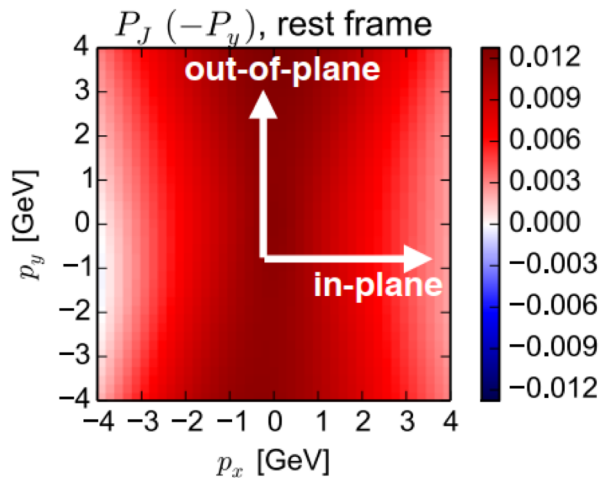
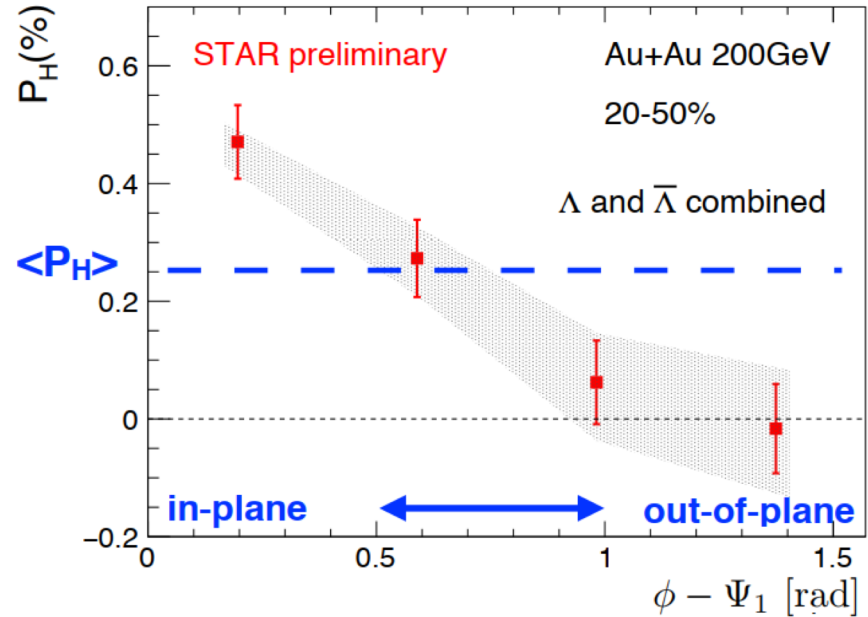
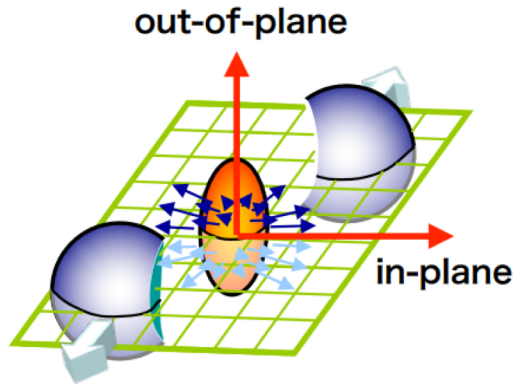


B.C. Fu, K. Xu, X.G. Huang, H.C. Song,
Phys. Rev. C 103, 024903 (2021)



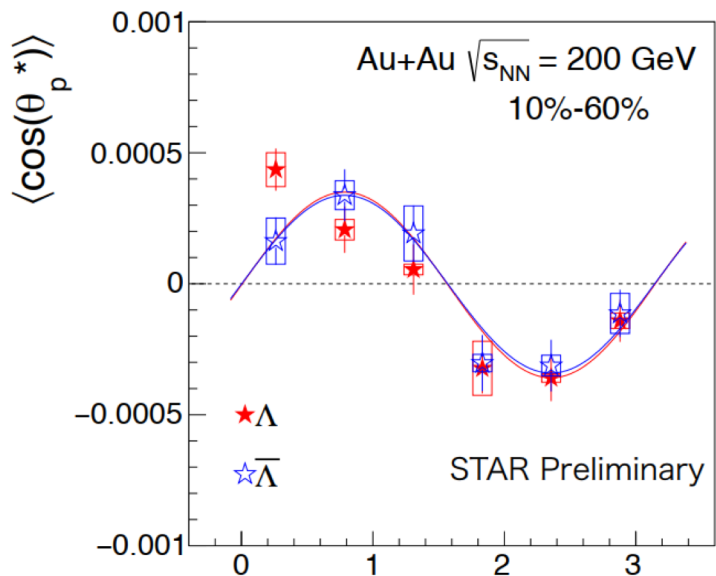
S. Ryu, V. Jupic, C. Shen,
arXiv:2106.08125

Local Polarization

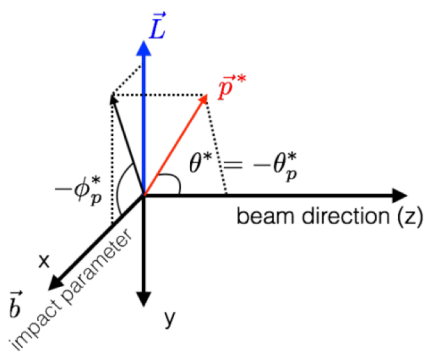
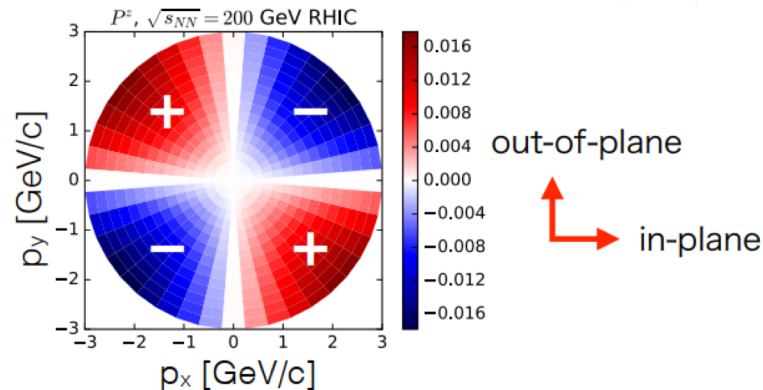


- **Exp data:**
 P_H in-plane $>$ P_H out-of-plane
- **Simulations:**
Sign is opposite of expected!

Local Polarization along beam direction



Again, sign is opposite of expected!



$$\frac{dN}{d\Omega^*} = \frac{1}{4\pi} (1 + \alpha_H \mathbf{P}_H \cdot \mathbf{p}_p^*)$$

$$\langle \cos \theta_p^* \rangle = \int \frac{dN}{d\Omega^*} \cos \theta_p^* d\Omega^*$$

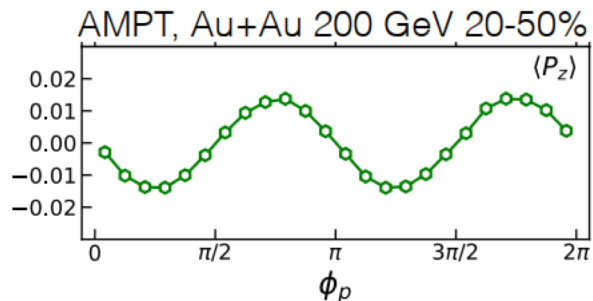
$$= \alpha_H P_z \langle (\cos \theta_p^*)^2 \rangle$$

$$\therefore P_z = \frac{\langle \cos \theta_p^* \rangle}{\alpha_H \langle (\cos \theta_p^*)^2 \rangle}$$

$$= \frac{3 \langle \cos \theta_p^* \rangle}{\alpha_H} \quad (\text{if perfect detector})$$

α_H : hyperon decay parameter
 θ_p^* : θ of daughter proton in Λ rest frame

UrQMD : *Becattini, Karpenko, PRL (2018)*

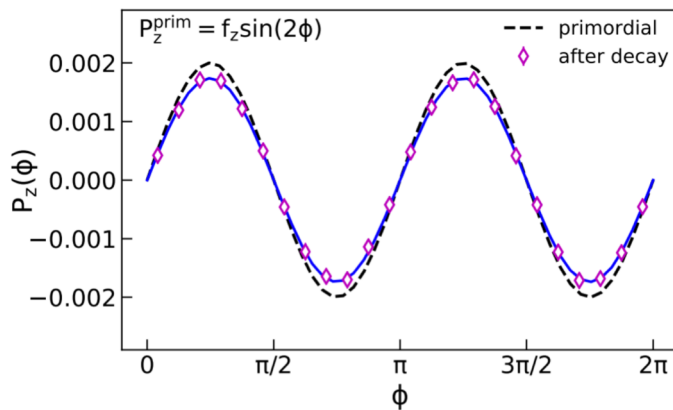


AMPT: *Xia, Li, Tang, Wang, PRC (2018)*

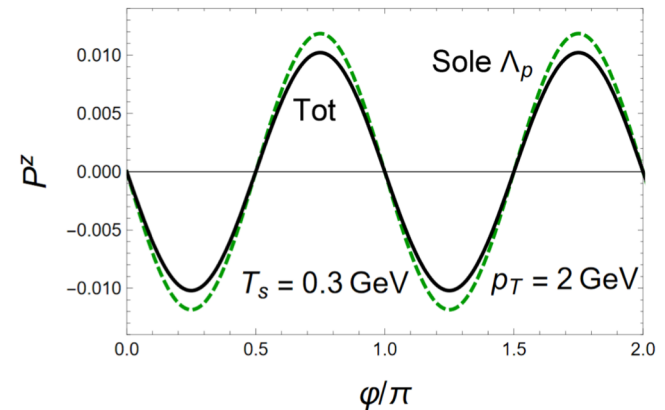
Feed-down effects: NO!

- Feed-down effects

Lambda may come from decays of heavier particles



Xia, Li, Huang, Huang PRC(2019)



Becattini, Cao, Speranza, EPJC(2019)

Different approaches

- **Spin hydrodynamics**

Florkowski, Friman, Jaiswal, Ryblewski, Speranza (2017-2018);

Montenegro, Tinti, Torrieri (2017-2019);

Hattori, Hongo, Huang, Matsuo, Taya PLB(2019)

Fukushima, SP, Lecture Note (2020); PLB(2021)

S.Y. Li, M.A Stephanov, H.U Yee, arXiv:2011.12318

D. She, A. Huang, D.F. Hou, J.F Liao, arXiv: 2105.04060

....

- **Quantum kinetic theory for massive fermions**

- **Other approaches:**

Different approaches

- **Spin hydrodynamics**
- **Quantum kinetic theory for massive fermions**
- **Other approaches:**

Quantum kinetic theory (massless fermion)

- Hamiltonian formulism, effective theory

Son, Yamamoto, PRL, (2012); PRD (2013)

- Path integration

Stephanov, Yin, PRL (2012);

Chen, Son, Stephanov, Yee, Yin, PRL, (2014);

J.W. Chen, J.Y. Pang, SP, Q. Wang, PRD (2014)

- Wigner function (Quantum field theory)

- hydrodynamics, equilibrium

J.W. Chen, SP, Q. Wang, X.N. Wang, PRL (2013);

- out-of-equilibrium, quantum field theory

Y. Hidaka, SP, D.L. Yang, PRD(RC) (2017)

- Other studies

A.P. Huang, S.Z. Su, Y. Jiang, J.F. Liao, P.F. Zhuang, PRD (2018)

- World-line formulism

N. Muller, R, Venugopalan PRD 2017

Also see Talks at Chirality Workshop in 2018, 2019 and Quark Matter Chirality section 2019

Quantum kinetic theory (massive fermions)

- Quantum kinetic theory (for massive fermions)
- Collision term with quantum corrections

Weickgenannt, Sheng, Speranza, Wang, Rischke, PRD 100, 056018 (2019)

Hattori, Hidaka, Yang, PRD100, 096011 (2019); Yang, Hattori, Hidaka, arXiv: 2002.02612.

Liu, Mameda, Huang, arXiv:2002.03753.

Wang, Guo, Shi, Zhuang, PRD100, 014015 (2019),

Li, Yee, PRD100, 056022 (2019)

Recent reviews:

Gao, Ma, SP, Wang, NST 31 (2020) 9, 90

Gao, Liang, Wang, Int.J.Mod.Phys A 36 (2021), 2130001

Hidaka, SP, Yang, Wang, in preparation

Different approaches

- Spin hydrodynamics
- Quantum kinetic theory for massive fermions
- Other approaches:
 - Side-jump effect *Liu, Sun, Ko PRL(2020)*
 - Mesonic mean-field *Csernai, Kapusta, Welle, PRC(2019)*
 - Using different vorticity *Wu, Pang, Huang, Wang, PRR (2019)*
- Also see recent review
J.H. Gao, G.L. Ma, SP, Q. Wang, NST 31(2020)9, 90

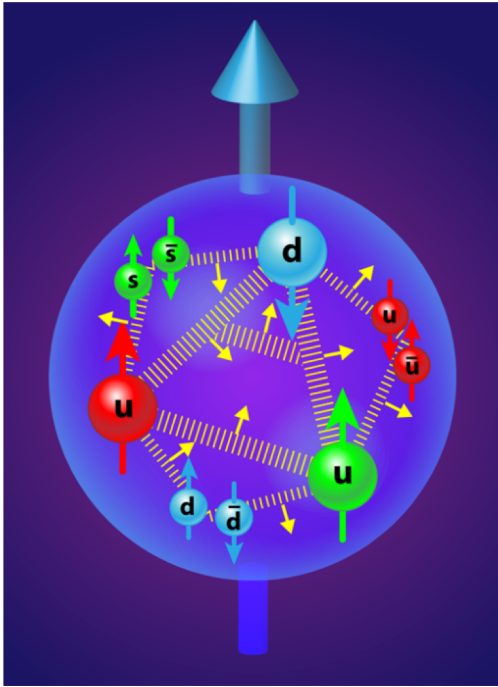
Spin hydrodynamics

- **Relativistic hydrodynamics + spin degree of freedom**
- **Problem: how to introduce spin of a massive fermionic fluid in a relativistic theory?**

Connection to “spin physics” (QCD)

- Proton spin problem:

(slides from Hatta-son's talk)



$$\frac{1}{2} = \frac{1}{2} \Delta \Sigma + \Delta G + L^q + L^g$$

↑
Quarks'
helicity

↑
Gluons'
helicity

↑
Orbital angular
Momentum (OAM)

Total angular momentum conservation

- **Nöther's theorem :**

$$x^\mu \rightarrow x'^\mu = \Lambda^\mu_\nu x^\nu = (\delta^\mu_\nu + \epsilon^\mu_\nu) x^\nu, \quad A^\mu(x) \rightarrow A'^\mu(x) = \Lambda^\mu_\nu A^\nu(\Lambda^{-1}x),$$
$$\psi(x) \rightarrow \psi'(x) = \Lambda_{\frac{1}{2}} \psi(\Lambda^{-1}x),$$



$$\partial_\lambda (J_A^{\lambda\mu\nu} + J_\psi^{\lambda\mu\nu}) = 0$$

- **Nöther current**

Gauge
part

$$J_A^{\lambda\mu\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_\lambda A^\alpha)} \Delta A^{\mu\nu\alpha} = -F^\lambda_\alpha (x^\mu \partial^\nu - x^\nu \partial^\mu) A^\alpha - F^{\lambda\mu} A^\nu + F^{\lambda\nu} A^\mu.$$

Fermionic
part

$$J_\psi^{\lambda\mu\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_\lambda \psi)} \Delta \psi^{\mu\nu} = \bar{\psi} i \gamma^\lambda (x^\mu \partial^\nu - x^\nu \partial^\mu - i \Sigma^{\mu\nu}) \psi$$

- How to define the orbital and spin parts?

Jaffe-Manohar (canonical) decomposition

- Canonical energy momentum tensor Non-symmetric

Gauge part $T_{A,\text{can}}^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_\mu A^\alpha)} \partial^\nu A^\alpha - g^{\mu\nu} \mathcal{L}_A = -F_\alpha^\mu \partial^\nu A^\alpha + \frac{1}{4} g^{\mu\nu} F^{\alpha\beta} F_{\alpha\beta}$

Fermionic part $T_{\psi,\text{can}}^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \psi)} \partial^\nu \psi - g^{\mu\nu} \mathcal{L}_\psi = \bar{\psi} i \gamma^\mu \partial^\nu \psi - g^{\mu\nu} \bar{\psi} (i \gamma^\alpha D_\alpha - m) \psi$

- Canonical decomposition

$$J^{\mu\nu\lambda} = \underbrace{x^\nu T_{can}^{\mu\lambda} - x^\lambda T_{can}^{\mu\nu}}_{\text{Orbital angular momentum}} - \underbrace{\frac{1}{2} \epsilon^{\mu\nu\lambda\rho} \bar{\psi} \gamma_5 \gamma_\rho \psi}_{\text{Quark helicity (spin)}} + \underbrace{F^{\mu\lambda} A^\nu - F^{\mu\nu} A^\lambda}_{\text{Gluon helicity (Spin)}}$$

- Jaffe-Manohar decomposition (1990)

$$\frac{1}{2} = \frac{1}{2} \Delta \Sigma + \Delta G + L_{can}^q + L_{can}^g \quad \underline{\text{Operators are not gauge invariant.}}$$

Pseudo-gauge transformation

- The variation of (canonical) spin is the anti-symmetric part of energy momentum tensor

$$0 = \partial_\lambda J^{\lambda\mu\nu} = \partial_\lambda (x^\mu T_{\text{can}}^{\lambda\nu} - x^\nu T_{\text{can}}^{\lambda\mu} + S_{\text{can}}^{\lambda\mu\nu}) \Rightarrow T_{\text{can}}^{\mu\nu} - T_{\text{can}}^{\nu\mu} = -\partial_\lambda S_{\text{can}}^{\lambda\mu\nu},$$

- Belinfante energy momentum tensor

$$T_{\text{Bel}}^{\mu\nu} \equiv T_{\text{can}}^{\mu\nu} + \partial_\lambda K_{\text{Bel}}^{\lambda\mu\nu}$$

$$K^{\lambda\mu\nu} = -K^{\mu\lambda\nu}$$

$$K_{\text{Bel}}^{\lambda\mu\nu} = \frac{1}{2} (S_{\text{can}}^{\lambda\mu\nu} - S_{\text{can}}^{\mu\lambda\nu} + S_{\text{can}}^{\nu\mu\lambda})$$

Conserved

$$\partial_\mu T_{\text{Bel}}^{\mu\nu} = 0$$

$$T_{\text{Bel}}^{\mu\nu} - T_{\text{Bel}}^{\nu\mu} = 0$$

Symmetric

$$T_{A,\text{Bel}}^{\mu\nu} \equiv -F_\alpha^\mu F^{\nu\alpha} + \frac{1}{4} g^{\mu\nu} F^{\alpha\beta} F_{\alpha\beta},$$

$$T_{\psi,\text{Bel}}^{\mu\nu} \equiv \bar{\psi} i \gamma^\mu \overleftrightarrow{D}^\nu \psi + \frac{1}{4} \varepsilon^{\mu\nu\lambda\rho} \partial_\lambda (\bar{\psi} \gamma_5 \gamma_\rho \psi).$$

Gauge invariant

Ji (Belinfante) decomposition

- Belinfante improved total angular momentum

$$J_{\text{Bel}}^{\lambda\mu\nu} \equiv J^{\lambda\mu\nu} + \partial_\rho (x^\mu K_{\text{Bel}}^{\rho\lambda\nu} - x^\nu K_{\text{Bel}}^{\rho\lambda\mu})$$

$$\partial_\lambda J^{\lambda\mu\nu} = 0 \quad \longleftrightarrow \quad \partial_\lambda J_{\text{Bel}}^{\lambda\mu\nu} = 0$$

$$J_{A/\psi, \text{Bel}}^{\lambda\mu\nu} = x^\mu \tilde{T}_{A/\psi, \text{Bel}}^{\lambda\nu} - x^\nu \tilde{T}_{A/\psi, \text{Bel}}^{\lambda\mu}$$

- Ji decomposition (1997)

$$\frac{1}{2} = J_q + J_g$$

Table of two different forms

- Canonical (Jaffe-Manohar) decomposition

$$\mathbf{J} = \underbrace{-\frac{1}{2}\bar{\psi}\gamma_5\boldsymbol{\gamma}\psi}_{\frac{1}{2}\Delta\Sigma} + \underbrace{\mathbf{E} \times \mathbf{A}}_{\Delta G} + \underbrace{-i\psi^\dagger(\mathbf{x} \times \boldsymbol{\nabla})\psi}_{L_{\text{can}}^q} + \underbrace{\mathbf{E}(\mathbf{x} \times \boldsymbol{\nabla})\mathbf{A}}_{L_{\text{can}}^g}$$

$$T_{\psi,\text{can}}^{\mu\nu} = \bar{\psi}i\gamma^\mu \partial^\nu \psi - g^{\mu\nu}\bar{\psi}(i\gamma^\alpha D_\alpha - m)\psi$$

non-symmetric
Not gauge invariant

- Belinfante (Ji) decomposition

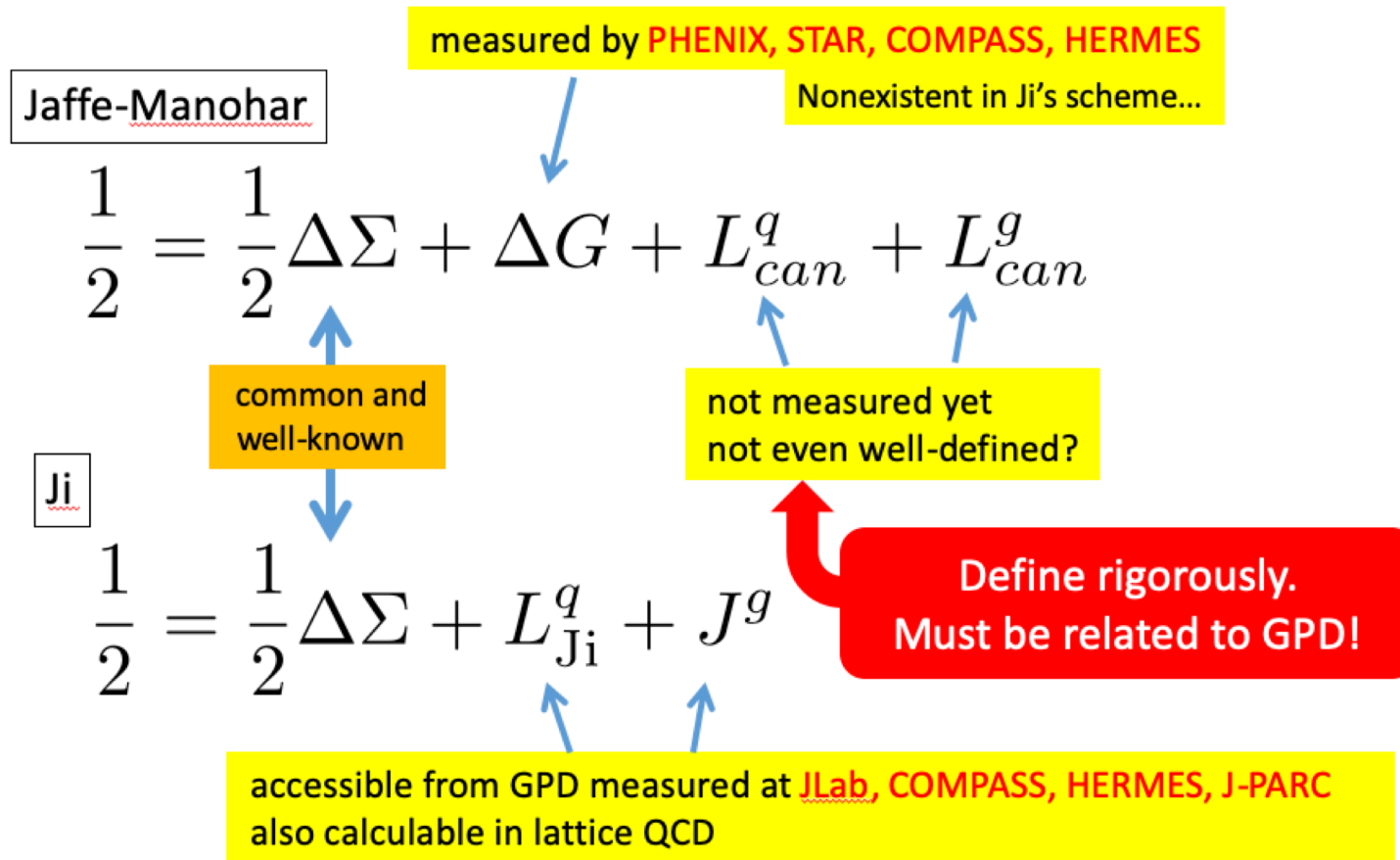
$$\mathbf{J} = \underbrace{-\frac{1}{2}\bar{\psi}\gamma_5\boldsymbol{\gamma}\psi}_{\frac{1}{2}\Delta\Sigma} + \underbrace{-i\psi^\dagger(\mathbf{x} \times \mathbf{D})\psi}_{L_{\text{Ji}}^q} + \underbrace{\mathbf{x} \times (\mathbf{E} \times \mathbf{B})}_{J_{\text{Ji}}^g}$$

$$T_{\psi,\text{Bel}}^{\mu\nu} \equiv \bar{\psi}i\gamma^\mu \overleftrightarrow{D}^\nu \psi + \frac{1}{4}\varepsilon^{\mu\nu\lambda\rho}\partial_\lambda(\bar{\psi}\gamma_5\gamma_\rho\psi)$$

Symmetric
Gauge invariant

Connected by
pseudo gauge
transformation

Two spin communities divided



(slides from Hatta-san's talk)

E. Leader, C. Lorce, Phys. Rept. 541 (2014) 163-248

GLW decomposition

- **Another Pseudo gauge transformation**

$$T_{\text{can}}^{\mu\nu} = T_{\text{GLW}}^{\mu\nu} + \frac{1}{2} \partial_\lambda \left(\Phi_{\text{can}}^{\lambda, \mu\nu} + \Phi_{\text{can}}^{\mu, \nu\lambda} + \Phi_{\text{can}}^{\nu, \mu\lambda} \right)$$

$$S_{\text{can}}^{\lambda, \mu\nu} = S_{\text{GLW}}^{\lambda, \mu\nu} - \Phi_{\text{can}}^{\lambda, \mu\nu} \quad \Phi_{\text{can}}^{\lambda, \mu\nu} \equiv S_{\text{GLW}}^{\mu, \lambda\nu} - S_{\text{GLW}}^{\nu, \lambda\mu}$$

$$\partial_\lambda S_{\text{can}}^{\lambda, \mu\nu}(x) = T_{\text{can}}^{\nu\mu} - T_{\text{can}}^{\mu\nu} = -\partial_\lambda S_{\text{GLW}}^{\mu, \lambda\nu}(x) + \partial_\lambda S_{\text{GLW}}^{\nu, \lambda\mu}(x).$$

$$T_{\text{Bel}}^{\mu\nu} = T_{\text{GLW}}^{\mu\nu} - \frac{1}{2} \partial_\lambda \left(S_{\text{GLW}}^{\nu, \lambda\mu} + S_{\text{GLW}}^{\mu, \lambda\nu} \right)$$

Textbook written by de Groot, van Leeuwen, and van Weert

Review: W. Florkowski, R. Ryblewski and Avdhesh Kumar, Prog. Part. Nucl. Phys. 108 (2019) 103709

- **Microscopic kinetic theory: GLW is the classical one.**

$$T_{\text{GLW}}^{\mu\nu}(x) = \frac{1}{m} \text{tr}_4 \int d^4k k^\mu k^\nu \mathcal{W}(x, k) = \frac{1}{m} \int d^4k k^\mu k^\nu \mathcal{F}(x, k).$$

Question

- **Which kinds of energy momentum tensor are measured or preferred by the experiments?**
- **Hints:**
 - ✓ Ordinary relativistic hydrodynamics formulism are symmetric.
(Relatively easy to be extended to spin hydro ?)
 - ✓ Anomalous (magneto-) hydrodynamics (including the spin current for massless fermions) are symmetric.
(Relatively easy to be checked in massless limit)
 - ✓ Maybe, we eventually need to add the gluons' contributions.
(A gauge invariant macroscopic theory may be more acceptable.)

Canonical form of spin hydrodynamics

Ref:

***K. Hattori, M. Hongo, X.-G.Huang, M. Matsuo, H. Taya,
“Fate of spin polarization in a relativistic fluid: An entropy-current analysis,”
Phys. Lett. B795 (2019) 100–106, arXiv:1901.06615 [hep-th].***

Also see recent work:

***S.Y. Li, M.A Stephanov, H.U Yee, “Non-dissipative second-order transport,
spin, and pseudo-gauge transformations in hydrodynamics”,
arXiv:2011.12318***

***D. She, A. Huang, D.F. Hou, J.F Liao, “Relativistic Viscous Hydrodynamics
with Angular Momentum”, arXiv: 2105.04060***

Basic conservation equations

- Total angular momentum conservation

$$\partial_\alpha J_{can}^{\alpha\mu\nu} = 0 \quad J_{can}^{\alpha\mu\nu} = \underbrace{x^\mu T_{can}^{\alpha\nu} - x^\nu T_{can}^{\alpha\mu}}_{\text{Orbital part}} + \underbrace{\Sigma^{\alpha\mu\nu}}_{\text{Spin tensor}}$$



$$\partial_\alpha \Sigma^{\alpha\mu\nu} = T_{can}^{\nu\mu} - T_{can}^{\mu\nu}$$



- Energy-momentum conservation

$$\partial_\mu T_{can}^{\mu\nu} = 0$$



- Currents conservation

$$\partial_\mu j^\mu = 0$$



Common strategy for derivation of fluid equations

- **Tensor decomposition:**

- Parallel or perpendicular to fluid velocity u^μ
- Traceless part and other part

- **Gradient (∂) expansion: $\partial X \ll X$**

- **Entropy principle:**

to derive the general expression for all components of tensors

An example: charge currents

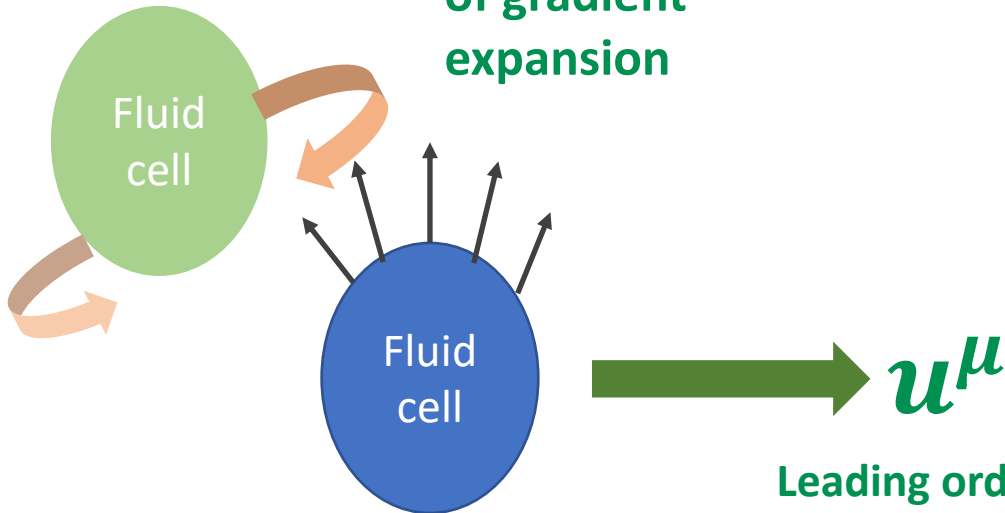
- Charge currents:

$$j^\mu = \underbrace{nu^\mu}_{\substack{\text{Parallel to fluid} \\ \text{velocity } u^\mu; \\ \text{Leading order} \\ \text{of gradient} \\ \text{expansion}}} + j_{(1)}^\mu$$

n : charge density
↓

$j_{(1)}^\mu$
Perpendicular to fluid velocity u^μ ;
Higher orders of gradient expansion

Higher orders:
exchange the heat
and particles with
other cells.



Leading order:
moving along the u^μ in average

Spin tensor decomposition

- Analogy to the decomposition for currents:

$$j^\mu = nu^\mu + j_{(1)}^\mu$$

Parallel to fluid velocity u^μ ;
Leading order

Perpendicular to fluid velocity u^μ ;
Higher order

- One can assume that

$$\Sigma^{\alpha\mu\nu} = u^\alpha S^{\mu\nu} + \Sigma_{(1)}^{\alpha\mu\nu}$$

spin tensor

Parallel to fluid velocity u^μ ;
Leading order

Perpendicular to fluid velocity u^μ ;
Higher order

Spin density

Modified thermodynamic relations

- ## Density vs Chemical potential

Charge density: n

Spin density: $S^{\mu\nu}$



Charge chemical potential: μ

Spin chemical potential: $\omega^{\mu\nu}$

Physical interpretation:
Variation of total energy when we add one particle with spin $S^{\mu\nu}$

- ## Thermodynamic relations

$$e + p = Ts + \mu n + \omega_{\mu\nu} S^{\mu\nu}$$

energy density
pressure
temperature X entropy density
spin chemical potential
X
spin density

- ## Gibbs relations

$$de = Tds + \mu dn + \omega_{\mu\nu} dS^{\mu\nu} \quad dp = sdT + nd\mu + S^{\mu\nu} d\omega_{\mu\nu}$$

Orders of $S^{\mu\nu}$ and $\omega^{\mu\nu}$

- In Ref. *K. Hattori, M. Hongo, X.-G.Huang, M. Matsuo, H. Taya, Phys. Lett. B795 (2019) 100–106.*

$$S^{\mu\nu}, \omega^{\mu\nu} \sim O(\partial^1); \omega_{\mu\nu} S^{\mu\nu} \sim O(\partial^2)$$

- In our recent work, *K. Fukushima, SP, PLB 2010.01608*

$$S^{\mu\nu} \sim O(\mathbf{1}), \omega^{\mu\nu} \sim O(\partial^1); \omega_{\mu\nu} S^{\mu\nu} \sim O(\partial^{\mathbf{1}})$$

Density is classic $O(1)$, but the variation of energy is quantum $O(\partial^1)$!

- We only consider the spin hydro up to the **first order** in gradient expansion.

Entropy production rate

- Two ways to derive the entropy flow:

- Directly using

$$u_\nu \partial_\mu T_{can}^{\mu\nu} + \mu \partial_\mu j^\mu = 0 \quad + \text{Gibbs relation}$$

➡ $\partial_\mu \mathcal{S}_{can}^\mu \geq 0$

K. Hattori, M. Hongo, X.-G.Huang, M. Matsuo, H. Taya, PLB795 (2019) 100–106.

- Using the extended entropy flow

W. Israel, J. Stewart, Annals Phys. 118, 341 (1979)

Relativistic fluid
generation
of Gibbs relation

$$\begin{aligned} \mathcal{S}_{can}^\mu &= \frac{u_\nu}{T} \Theta^{\mu\nu} + \frac{p}{T} u^\mu - \frac{\mu}{T} j^\mu - \frac{1}{T} \omega_{\rho\sigma} S^{\rho\sigma} u^\mu + \mathcal{O}(\partial^2) \\ &= s u^\mu + \frac{u_\nu}{T} \Theta_{(1)}^{\mu\nu} - \frac{\mu}{T} j_{(1)}^\mu + \mathcal{O}(\partial^2), \end{aligned}$$

➡ $\partial_\mu \mathcal{S}_{can}^\mu \geq 0$

K. Fukushima, SP, PLB 2010.01608

Constraints from entropy principle

$$T_{can}^{\mu\nu} = T_{(0)}^{\mu\nu} + T_{(1)}^{\mu\nu}$$

Leading order
Next-to-Leading order

$$T_{(1s)}^{\mu\nu} = T_{(1)}^{\mu\nu} + T_{(1)}^{\nu\mu}$$

$$T_{(1a)}^{\mu\nu} = T_{(1)}^{\mu\nu} - T_{(1)}^{\nu\mu}$$

symmetric
anti-symmetric

$$\partial_\mu \mathcal{S}_{can}^\mu = T_{(1s)}^{\mu\nu} \partial_\mu \frac{u_\nu}{T} - j_{(1)}^\mu \partial_\mu \frac{\mu}{T} + \frac{2T_{(1a)}^{\mu\nu}}{T} \left(\omega_{\mu\nu} + \frac{1}{2} T \omega_{\mu\nu}^{th} \right) \geq 0$$

Ordinary terms
not related to spin

Thermal vorticity

$$\omega_{th}^{\mu\nu} = (g^{\mu\alpha} - u^\mu u^\alpha)(g^{\nu\beta} - u^\nu u^\beta) \times [\partial_\alpha(u_\beta/T) - \partial_\beta(u_\alpha/T)]$$

Non-relativistic
limit



$$\epsilon^{ijk} \omega_{th}^{ij} \sim \left(\nabla \times \frac{\mathbf{V}}{T} \right)^k$$

Global equilibrium

Ordinary terms
not related to spin

$$\omega_{th}^{\mu\nu} = (g^{\mu\alpha} - u^\mu u^\alpha)(g^{\nu\beta} - u^\nu u^\beta) \times [\partial_\alpha(u_\beta/T) - \partial_\beta(u_\alpha/T)]$$

Thermal vorticity

$$\partial_\mu \mathcal{S}_{can}^\mu = T_{(1s)}^{\mu\nu} \partial_\mu \frac{u_\nu}{T} - j_{(1)}^\mu \partial_\mu \frac{\mu}{T} + \frac{2T_{(1a)}^{\mu\nu}}{T} \left(\omega_{\mu\nu} + \frac{1}{2} T \omega_{\mu\nu}^{th} \right) = 0$$

Must vanish!

Spin chemical potential $\omega^{\mu\nu}$ must be related to thermal vorticity as $-T\omega_{\mu\nu}^{th}/2$ in global equilibrium!

Widely proved by many approaches:

F. Becattini, L. Bucciattini, E. Grossi, and L. Tinti, Eur. Phys. J. C 75, 191 (2015)

F. Becattini, W. Florkowski, and E. Speranza, Physics Letters B 789, 419 (2019)

K. Hattori, M. Hongo, X.-G. Huang, M. Matsuo, H. Taya, Phys. Lett. B795 (2019) 100–106.

...

Also see recent reviews:

Y. C. Liu and X. G. Huang, Nucl. Sci. Tech. 31,56 (2020)

J.H. Gao, G.L. Ma, SP, Q. Wang, Nucl. Sci. Tech 31 (2020) 9, 90

Local equilibrium

$$\partial_\mu \mathcal{S}_{can}^\mu = T_{(1s)}^{\mu\nu} \partial_\mu \frac{u_\nu}{T} - j_{(1)}^\mu \partial_\mu \frac{\mu}{T} + \frac{2T_{(1a)}^{\mu\nu}}{T} \left(\omega_{\mu\nu} + \frac{1}{2} T \omega_{\mu\nu}^{th} \right) \geq 0$$

- **Tensor decomposition of energy momentum tensor**

symmetric $T_{(1s)}^{\mu\nu} = \underset{\substack{\uparrow \\ \text{heat flow}}}{h^\mu u^\nu} + h^\nu u^\mu + \underset{\substack{\uparrow \\ \text{viscous tensor}}}{\pi^{\mu\nu}}$

anti-symmetric $T_{(1a)}^{\mu\nu} = q^\mu u^\nu - q^\nu u^\mu + \phi^{\mu\nu}$

$$\begin{aligned} \partial_\mu \mathcal{S}_{can}^\mu = & \left(h^\mu - \frac{e+p}{n} j_{(1)}^\mu \right) \frac{n}{e+p} (g_{\nu\alpha} - u_\nu u_\alpha) \partial^\nu \frac{\mu}{T} \\ & + \frac{\pi^{\mu\nu}}{T} \partial_{\langle\mu} u_{\nu\rangle} + \frac{1}{3} \frac{1}{T} \pi_\mu^\mu (\partial \cdot u) \\ & + q^\mu \left[-\frac{1}{T} (u \cdot \partial) u_\mu + \partial_\mu \frac{1}{T} + 4 \frac{\omega_{\mu\nu} u^\nu}{T} \right] \\ & + \phi^{\mu\nu} [\omega_{\mu\nu}^{th} + 2\beta \omega_{\mu\nu}] \geq 0 \end{aligned}$$

} dissipative terms in ordinary fluids

} new terms related to spin

Entropy principle

$$\begin{aligned}
 \partial_\mu \mathcal{S}_{can}^\mu &= \left(h^\mu - \frac{e+p}{n} j_{(1)}^\mu \right) \frac{n}{e+p} (g_{\nu\alpha} - u_\nu u_\alpha) \partial^\nu \frac{\mu}{T} \\
 &+ \frac{\pi^{\mu\nu}}{T} \partial_{\langle\mu} u_{\nu\rangle} + \frac{1}{3} \frac{1}{T} \pi_\mu^\mu (\partial \cdot u) \\
 &+ q^\mu \left[-\frac{1}{T} (u \cdot \partial) u_\mu + \partial_\mu \frac{1}{T} + 4 \frac{\omega_{\mu\nu} u^\nu}{T} \right] \\
 &+ \phi^{\mu\nu} [\omega_{\mu\nu}^{th} + 2\beta \omega_{\mu\nu}] \geq 0
 \end{aligned}$$

} **dissipative terms
in ordinary fluids**

} **new terms related
to spin**

- **To ensure the entropy production rate be always positive, the only possible way is**

$$\begin{aligned}
 q^\mu &= \lambda [(u \cdot \partial) u^\mu + \frac{1}{T} \Delta^{\mu\nu} \partial_\nu T - 4 \omega^{\mu\nu} u_\nu], \\
 \phi^{\mu\nu} &= 2\gamma [T \omega_{th}^{\mu\nu} + 2(g^{\mu\alpha} - u^\mu u^\alpha)(g^{\nu\beta} - u^\nu u^\beta) \omega_{\alpha\beta}] / T.
 \end{aligned}$$

$\lambda, \gamma \geq 0$ are new transport coefficients

Brief summary of canonical form

- Energy momentum tensor has anti-symmetric part

$$T_{can}^{\mu\nu} = T_{(0)}^{\mu\nu} + T_{(1)}^{\mu\nu} \quad T_{(1s)}^{\mu\nu} = T_{(1)}^{\mu\nu} + T_{(1)}^{\nu\mu} \quad \text{symmetric}$$

Leading order
Next-to-Leading order
 $T_{(1a)}^{\mu\nu} = T_{(1)}^{\mu\nu} - T_{(1)}^{\nu\mu}$
anti-symmetric

- In global equilibrium, spin chemical potential is related to thermal vorticity

$$\omega^{\mu\nu} = -\frac{1}{2}T\omega_{th}^{\mu\nu}$$

Power counting: $\omega_{\mu\nu}^{th} \sim \mathcal{O}(\partial^1)$
 $\rightarrow \omega_{\mu\nu} \sim \mathcal{O}(\partial^1)$

- Symmetric part is as the same as the ordinary fluid. The expression for anti-symmetric part can be derived by entropy principle.

$$T_{(1a)}^{\mu\nu} = q^\mu u^\nu - q^\nu u^\mu + \phi^{\mu\nu}$$

$$q^\mu = \lambda[(u \cdot \partial)u^\mu + \frac{1}{T}\Delta^{\mu\nu}\partial_\nu T - 4\omega^{\mu\nu}u_\nu],$$

$$\phi^{\mu\nu} = 2\gamma[T\omega_{th}^{\mu\nu} + 2(g^{\mu\alpha} - u^\mu u^\alpha)(g^{\nu\beta} - u^\nu u^\beta)\omega_{\alpha\beta}]/T.$$

Power counting: $T_{(1a)}^{\mu\nu} \sim \mathcal{O}(\partial^1)$
 $\partial_\rho(u^\rho S^{\mu\nu}) = -2T_{(1a)}^{\mu\nu} \sim \mathcal{O}(\partial^1)$
 $\rightarrow S^{\mu\nu} \sim \mathcal{O}(1)$

$\lambda, \gamma \geq 0$ are new transport coefficients

Belinfante form of spin hydrodynamics

Ref.

K. Fukushima, SP,

***"Spin Hydrodynamics and Symmetric Energy - Momentum Tensors
– A current induced by the spin vorticity", PLB 817 (2021) 136346***

***"Relativistic decomposition of the orbital and the spin angular
momentum in chiral physics and Feynman's angular momentum
paradox", invited lecture, 2001.00359, Lecture Notes in Physics volume
on "Strongly Interacting Matter under Rotation"***

Basic conservation equations

- **Total angular momentum conservation**

$$\begin{aligned} J_{\text{Bel}}^{\mu\nu\alpha} &= J^{\mu\nu\alpha} + \partial_\rho (x^\nu K_{\text{Bel}}^{\rho\mu\alpha} - x^\alpha K_{\text{Bel}}^{\rho\mu\nu}), \\ &= x^\nu T_{\text{Bel}}^{\mu\alpha} - x^\alpha T_{\text{Bel}}^{\mu\nu}. \end{aligned}$$

$$\partial_\mu J_{\text{Bel}}^{\mu\nu\alpha} = 0,$$

- **Energy momentum conservation**

✓ $\partial_\mu T_{\text{Bel}}^{\mu\nu} = 0.$

These two equations are equivalent!

- **Current conservation**

✓ $\partial_\mu j^\mu = 0$

No spin in Belinfante form?

$$\boxed{\partial_{\mu} J_{\text{Bel}}^{\mu\nu\alpha} = 0,} \quad \xleftrightarrow{\text{equivalent}} \quad \boxed{\partial_{\mu} T_{\text{Bel}}^{\mu\nu} = 0.}$$

The common argument:

- There is no equation for spin.
- There is no degree of freedom for spin.
- We cannot observe spin effect in Belinfante form.

Recalling what we discussed in introduction.

- Belinfante energy momentum tensor is connected to canonical one by pseudo gauge transformation.

If a physical (spin) effect disappears after a physical transformation, then, did that mean this “physical effect” is unphysical?

Or, should it be that this physical effect will appear somewhere?

Belinfante energy momentum tensor

- We take the pseudo gauge transformation

$$T_{Bel}^{\mu\nu} = T_{can}^{\mu\nu} + \partial_\lambda K_{Bel}^{\lambda\mu\nu} = T_0^{\mu\nu} + T_{(1)}^{\mu\nu}$$

Leading
order

Next-to-
Leading order

$$T_{(1)}^{\mu\nu} = \underset{\substack{\uparrow \\ \text{heat flow}}}{h^\mu u^\nu} + \underset{\substack{\uparrow \\ \text{viscous tensor}}}{h^\nu u^\mu} + \underset{\substack{\uparrow \\ \text{spin density tensor}}}{\pi^{\mu\nu}} + \frac{1}{2} \partial_\lambda (u^\mu S^{\nu\lambda} + u^\nu S^{\mu\lambda})$$

spin corrections to the
energy momentum tensor

Spin corrections

$$T_{(1)}^{\mu\nu} = \underset{\substack{\uparrow \\ \text{heat flow}}}{h^\mu} u^\nu + h^\nu \underset{\substack{\uparrow \\ \text{viscous tensor}}}{u^\mu} + \pi^{\mu\nu} + \frac{1}{2} \partial_\lambda (u^\mu \underset{\substack{\uparrow \\ \text{spin density tensor}}}{S^{\nu\lambda}} + u^\nu S^{\mu\lambda})$$

spin corrections to the energy momentum tensor

Using standard tensor decomposition, we have

$$T_{(1)}^{\mu\nu} = (\delta e_s) u^\mu u^\nu + (h^\mu + \delta h_s^\mu) u^\nu + (h^\nu + \delta h_s^\nu) u^\mu + \pi^{\mu\nu} + \delta \pi_s^{\mu\nu}$$

$$\delta e_s = u_\rho \partial_\sigma S^{\rho\sigma}, \quad \longleftrightarrow \text{Spin correction to energy density}$$

$$\delta h_s^\mu = \frac{1}{2} \Delta_\sigma^\mu \partial_\lambda S^{\sigma\lambda} + \frac{1}{2} u_\rho S^{\rho\lambda} \partial_\lambda u^\mu \quad \longleftrightarrow \text{Spin correction to heat flow}$$

$$\delta \pi_s^{\mu\nu} = \partial_\lambda (u^{<\mu} S^{\nu>\lambda}) + \frac{1}{3} \partial_\lambda (u^\sigma S^{\rho\lambda}) \Delta_{\rho\sigma}. \quad \longleftrightarrow \text{Spin correction to viscous tensor}$$

Spin will appear as corrections to the ordinary dissipative terms!

Frame dependence

- In ordinary relativistic fluid, we have Landau (energy) frame and Particle frame.
- The heat flow depends on frame. In Landau frame, there is no heat flow,

$$u_L^\mu = u^\mu + \frac{1}{e+p}(h^\mu + \delta h^\mu),$$

but the dissipative current will be modified by spin correction!

$$j_{L(1)}^\mu = \left(j_{(1)}^\mu - \frac{n}{e+p} h^\mu \right) + \delta j_{(1)}^\mu \quad \delta j_{(1)}^\mu = -\frac{n}{e+p} \delta h^\mu .$$

Non-relativistic limit



$$\delta \mathbf{j}_{(1)} = -\frac{n}{2(e+p)} \left[\nabla \times \mathbf{S} + \dot{\mathbf{v}} \times \mathbf{S} + (\nabla \cdot \mathbf{v}) \mathbf{s} - 2(\mathbf{s} \cdot \nabla) \mathbf{v} + \dot{\mathbf{s}} \right] .$$

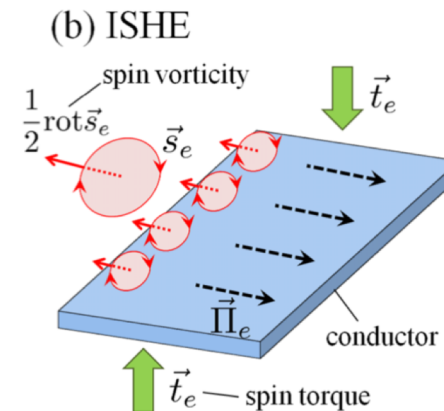
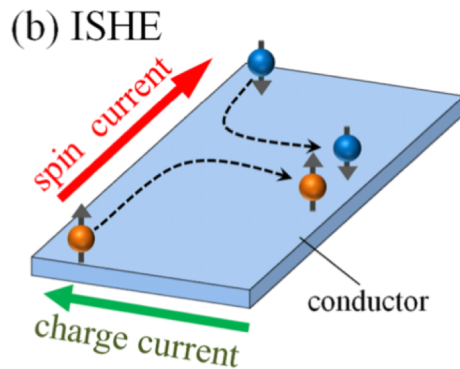
Quantum spin vorticity

- We have derived

$$\delta \mathbf{j}_{(1)} \propto -(\nabla \times \mathbf{S})$$

$\mathcal{S}^i = \epsilon^{ijk} \mathcal{S}^{ij}$
 spin density along
 i-th direction

Curl of spin will induce a current



Standard Inverse Spin Hall Effect (ISHE)

Inverse Spin Hall Effect (ISHE) understood
 by quantum spin vorticity

M. Fukuda, K. Ichikawa, M. Senami, and A. Tachibana, AIP Advances 6, 025108 (2016).

Entropy principle (1)

- Using the same method, we get the entropy production rate

$$\partial_{\mu} \mathcal{S}^{\mu} = \left(\frac{n}{e+p} h^{\mu} - j_{(1)}^{\mu} \right) \Delta_{\mu\nu} \partial^{\nu} \frac{\mu}{T} + \frac{1}{T} \pi^{\mu\nu} \partial_{\mu} u_{\nu} + \Delta$$

Spin
corrections

$$\Delta \equiv \frac{1}{2} \left[\partial_{\lambda} (u^{\mu} S^{\nu\lambda} + u^{\nu} S^{\mu\lambda}) \right] \partial_{\mu} \frac{u_{\nu}}{T} - \frac{\omega_{\rho\sigma}}{T} \partial_{\lambda} (u^{\lambda} S^{\rho\sigma}).$$

It is not in a squared form at all!

- The simplest way to ensure entropy principle is to let

$$S^{\mu\nu} = \mathbf{0}, \Delta = \mathbf{0} \quad ???$$

That is the way to “get” the common argument “No degree of freedom for spin in Belinfante form”.

Of course, it is a (trivial) solution. But, is it the **only** solution?

Entropy principle (2)

- Using the same method, we get the entropy production rate

$$\partial_\mu \mathcal{S}^\mu = \left(\frac{n}{e+p} h^\mu - j_{(1)}^\mu \right) \Delta_{\mu\nu} \partial^\nu \frac{\mu}{T} + \frac{1}{T} \pi^{\mu\nu} \partial_\mu u_\nu + \Delta$$

Spin
corrections

$$\Delta \equiv \frac{1}{2} \left[\partial_\lambda (u^\mu S^{\nu\lambda} + u^\nu S^{\mu\lambda}) \right] \partial_\mu \frac{u_\nu}{T} - \frac{\omega_{\rho\sigma}}{T} \partial_\lambda (u^\lambda S^{\rho\sigma}).$$

It is not in a squared form at all!

$$\begin{aligned} \longrightarrow \Delta = & \frac{1}{2} \partial_\mu \left[\partial_\lambda (u^\lambda S^{\mu\nu} + u^\mu S^{\nu\lambda} + u^\nu S^{\mu\lambda}) \frac{u_\nu}{T} \right] \longrightarrow \partial_\mu \delta \mathcal{S}^\mu \\ & - \frac{1}{2} \left[\partial_\lambda (u^\lambda S^{\mu\nu}) \right] \partial_\mu \frac{u_\nu}{T} - \frac{\omega_{\rho\sigma}}{T} \partial_\lambda (u^\lambda S^{\rho\sigma}). \end{aligned}$$

- We move the total derivatives to the entropy flow (redefine the entropy flow)

$$\partial_\mu (\mathcal{S}^\mu + \delta \mathcal{S}^\mu) = \dots + \Delta' \geq 0$$

Similar to the anomalous fluid *D.T. Son, P. Surowka, PRL. 103, 191601 (2009).*

Also see *S.Y. Li, M.A. Stephanov, H.U. Yee, appear in PRL, 2011.12318*

Entropy principle (3)

$$\partial_\mu(\mathcal{S}^\mu + \delta\mathcal{S}^\mu) = \dots + \Delta' \geq 0 \quad \Delta' = -\partial_\lambda(u^\lambda S^{\mu\nu}) \left(\frac{1}{2} \partial_\mu \frac{u_\nu}{T} + \frac{\omega_{\mu\nu}}{T} \right)$$

It reproduces the results in canonical form!

- In global equilibrium, $\omega^{\mu\nu} = -T\omega_{\mu\nu}^{th}/2$.
- In local equilibrium, by the tensor decomposition,

$$\partial_\lambda(u^\lambda S^{\rho\sigma}) = q^\rho u^\sigma - q^\sigma u^\rho + \phi^{\rho\sigma}$$

we can get the same results as in canonical form.

$$q^\mu = \lambda[(u \cdot \partial)u^\mu + \frac{1}{T}\Delta^{\mu\nu}\partial_\nu T - 4\omega^{\mu\nu}u_\nu],$$
$$\phi^{\mu\nu} = 2\gamma[T\omega_{th}^{\mu\nu} + 2(g^{\mu\alpha} - u^\mu u^\alpha)(g^{\nu\beta} - u^\nu u^\beta)\omega_{\alpha\beta}]/T.$$

We have re-discovered the equation of motion for spin by entropy principle!

Main equations for Belinfante form

- Energy momentum conservation

$$\partial_\mu T_{\text{Bel}}^{\mu\nu} = 0.$$

equivalent



$$\partial_\mu J_{\text{Bel}}^{\mu\nu\alpha} = 0,$$

- Current conservation

$$\partial_\mu j^\mu = 0$$

- Equations from entropy principle

$$\partial_\lambda (u^\lambda S^{\rho\sigma}) = q^\rho u^\sigma - q^\sigma u^\rho + \phi^{\rho\sigma}$$

$$q^\mu = \lambda[(u \cdot \partial)u^\mu + \frac{1}{T}\Delta^{\mu\nu}\partial_\nu T - 4\omega^{\mu\nu}u_\nu],$$

$$\phi^{\mu\nu} = 2\gamma[T\omega_{th}^{\mu\nu} + 2(g^{\mu\alpha} - u^\mu u^\alpha)(g^{\nu\beta} - u^\nu u^\beta)\omega_{\alpha\beta}]/T.$$

- Number of equations: 4+1+6= 11

- Variables: $T, \mu, S^{\mu\nu}, u^i, 1+1+6+3=11$

- Equation of state: $e = e(T, \mu, S^{\mu\nu}) + \text{Gibbs relation}$ $\omega^{\mu\nu} = \left. \frac{de}{dS^{\mu\nu}} \right|_{n,s}$

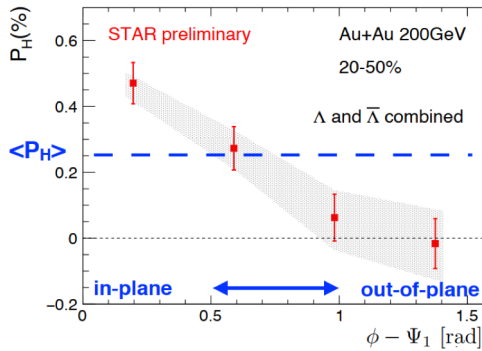
Revisit local spin polarization

Ref.

C. Yi, SP, D.L Yang

"Revisit local spin polarization beyond global equilibrium in relativistic heavy ion collisions", arXiv:2106.00238

Polarization induced by thermal vorticity



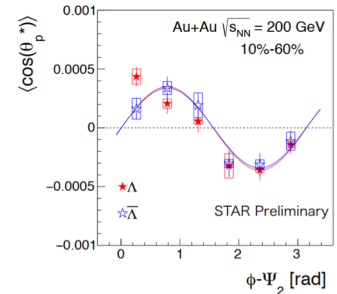
Thermal vorticity

$$\omega_{\rho\sigma}^{\text{th}} = \frac{1}{2} \left[\partial_\rho \left(\frac{u_\sigma}{T} \right) - \partial_\sigma \left(\frac{u_\rho}{T} \right) \right]$$

Distribution function: f_0

$$S^\mu(p) = \frac{1}{8m_\Lambda} \epsilon^{\mu\nu\rho\sigma} p_\nu \frac{\int d\Sigma_\lambda p^\lambda f_0 (1 - f_0) \omega_{\rho\sigma}^{\text{th}}}{\int d\Sigma_\lambda p^\lambda f_0}$$

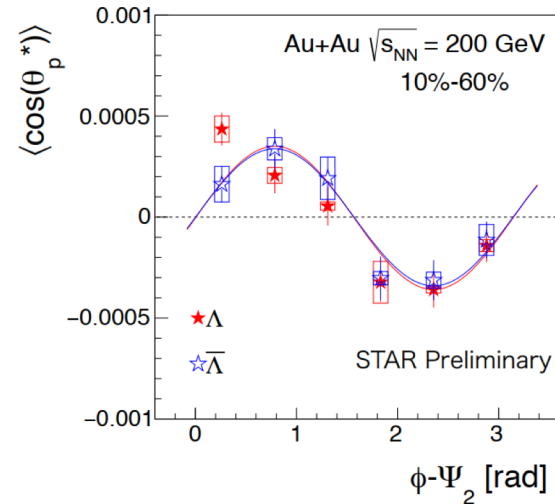
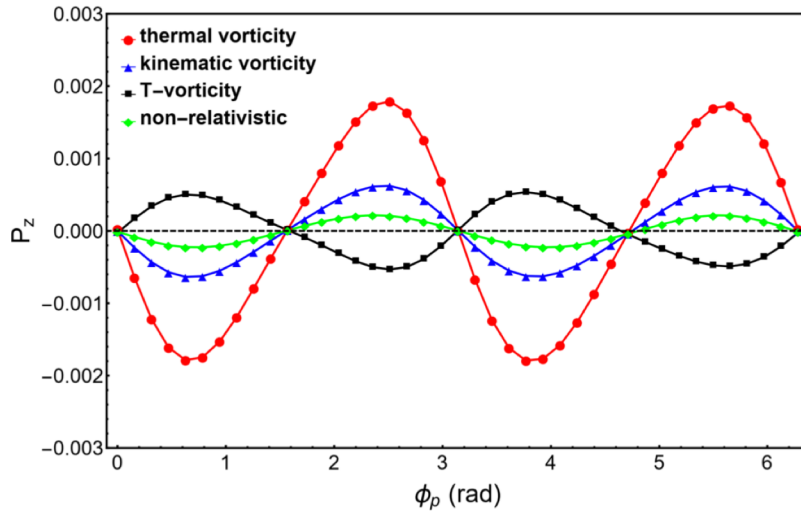
Freezeout surface



Karpenko, F. Becattini, Eur. Phys. J. C 77 (2017) 213

R.-H. Fang, L.-G. Pang, Q. Wang, X.-N. Wang, Phys. Rev. C94, 024904 (2016)

Local polarization from different vorticities



Kinematic vorticity:

$$\omega_{\mu\nu}^{(K)} = -\frac{1}{2}(\partial_\mu u_\nu - \partial_\nu u_\mu)$$

T-vorticity:

$$\omega_{\mu\nu}^{(T)} = -\frac{1}{2}[\partial_\mu(Tu_\nu) - \partial_\nu(Tu_\mu)]$$

Wu, Pang, Huang, Wang, PRR 1, 033058(2019)

- Only T-vorticity gives the right trend for both P_z and P_y
- Why T-vorticity? Out-of-equilibrium effects?

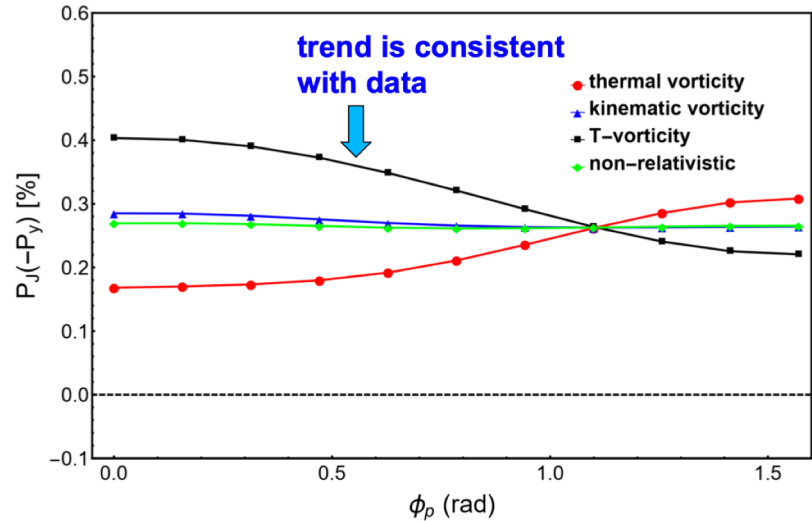
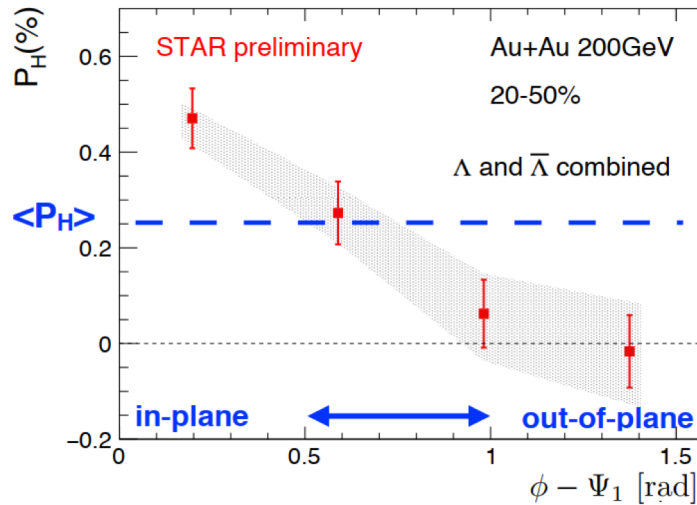
Non-Relativistic vorticity

$$\omega_{\mu\nu}^{(NR)} = \epsilon_{\nu\mu\rho\eta} u^\rho \omega^\eta$$

Thermal vorticity:

$$\omega_{\rho\sigma}^{\text{th}} = \frac{1}{2} \left[\partial_\rho \left(\frac{u_\sigma}{T} \right) - \partial_\sigma \left(\frac{u_\rho}{T} \right) \right]$$

Local polarization from different vorticities



Kinematic vorticity:

$$\omega_{\mu\nu}^{(K)} = -\frac{1}{2}(\partial_\mu u_\nu - \partial_\nu u_\mu)$$

T-vorticity:

$$\omega_{\mu\nu}^{(T)} = -\frac{1}{2}[\partial_\mu(Tu_\nu) - \partial_\nu(Tu_\mu)]$$

Wu, Pang, Huang, Wang, PRR 1, 033058(2019)

Non-Relativistic vorticity

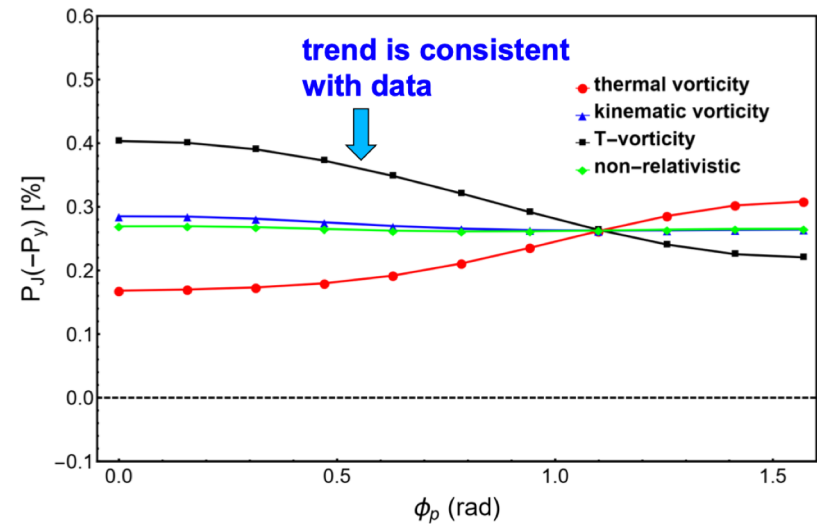
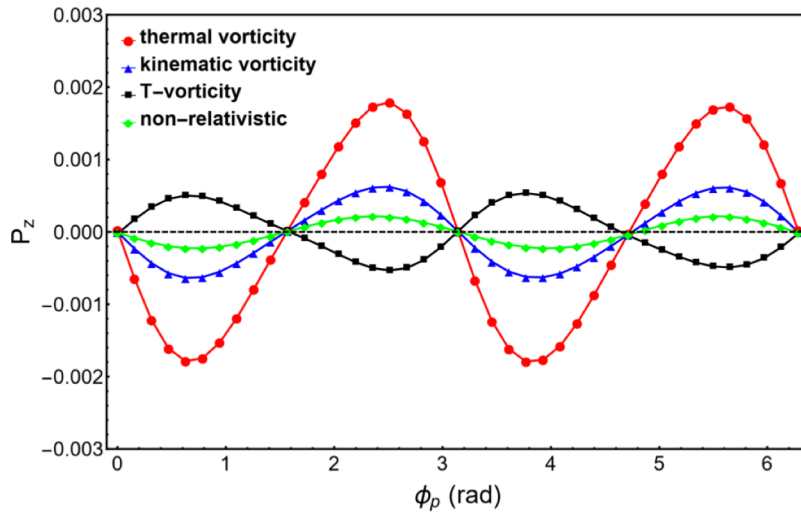
$$\omega_{\mu\nu}^{(NR)} = \epsilon_{\nu\mu\rho\eta} u^\rho \omega^\eta$$

Thermal vorticity:

$$\omega_{\rho\sigma}^{\text{th}} = \frac{1}{2} \left[\partial_\rho \left(\frac{u_\sigma}{T} \right) - \partial_\sigma \left(\frac{u_\rho}{T} \right) \right]$$

- Only T-vorticity gives the right trend for both Pz and Py
- Why T-vorticity? Out-of-equilibrium effects?

Local polarization from different vorticities



Kinematic vorticity:

$$\omega_{\mu\nu}^{(K)} = -\frac{1}{2}(\partial_\mu u_\nu - \partial_\nu u_\mu)$$

T-vorticity:

$$\omega_{\mu\nu}^{(T)} = -\frac{1}{2}[\partial_\mu(Tu_\nu) - \partial_\nu(Tu_\mu)]$$

Wu, Pang, Huang, Wang, PRR 1, 033058(2019)

Non-Relativistic vorticity

$$\omega_{\mu\nu}^{(NR)} = \epsilon_{\nu\mu\rho\eta} u^\rho \omega^\eta$$

Thermal vorticity:

$$\omega_{\rho\sigma}^{\text{th}} = \frac{1}{2} \left[\partial_\rho \left(\frac{u_\sigma}{T} \right) - \partial_\sigma \left(\frac{u_\rho}{T} \right) \right]$$

- Only T-vorticity gives the right trend for both Pz and Py
- Why T-vorticity? Out-of-equilibrium effects?

Polarization and axial current

- Recalling the original equations

$$\mathcal{S}^\mu(\mathbf{p}) = \frac{\int d\Sigma \cdot p \mathcal{J}_5^\mu(p, X)}{2m_\Lambda \int d\Sigma \cdot \mathcal{N}(p, X)},$$

- For massless fermions, the left and right handed currents read

$$\mathcal{J}_\lambda^\mu(p, X) = 2\pi \text{sign}(u \cdot p) \left\{ p^\mu + \lambda \frac{\hbar}{2} \delta(p^2) [u^\mu (p \cdot \omega) - \omega^\mu (u \cdot p) - 2S_{(u)}^{\mu\nu} \tilde{E}_\nu] \partial_{u \cdot p} + \lambda \frac{\hbar}{4} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} \partial_\nu^p \delta(p^2) \right\} f_\lambda^{(0)},$$

$$S_{(u)}^{\mu\nu} = \epsilon^{\mu\nu\alpha\beta} p_\alpha u_\beta / (2u \cdot p),$$

$$\tilde{E}_\nu = E_\nu + T \partial_\nu \frac{\mu_\lambda}{T} + \frac{(u \cdot p)}{T} \partial_\nu T - p^\sigma [\partial_{\langle \sigma} u_{\nu \rangle} + \frac{1}{3} \Delta_{\sigma\nu} (\partial \cdot u) + u_\nu D u_\sigma].$$

$$f_\lambda^{(0)} = 1 / (e^{(u \cdot p - \mu_\lambda)/T} + 1),$$

$\lambda = \pm$

+: right

-: left

Y. Hidaka, SP, and D.L. Yang, Phys. Rev. D97, 016004 (2018)

Polarization induced by different sources

- Axial currents can be decomposed as

$$\mathcal{J}_5^\mu = \mathcal{J}_{\text{thermal}}^\mu + \mathcal{J}_{\text{shear}}^\mu + \mathcal{J}_{\text{accT}}^\mu + \mathcal{J}_{\text{chemical}}^\mu + \mathcal{J}_{\text{EB}}^\mu,$$

where they are related to:

Thermal vorticity $\mathcal{J}_{\text{thermal}}^\mu = a \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} p_\nu \partial_\alpha \frac{u_\beta}{T},$

Shear viscous tensor $\mathcal{J}_{\text{shear}}^\mu = -a \frac{1}{(u \cdot p) T} \epsilon^{\mu\nu\alpha\beta} p_\alpha u_\beta p^\sigma \partial_{\langle\sigma} u_{\nu\rangle}$

Fluid acceleration $\mathcal{J}_{\text{accT}}^\mu = -a \frac{1}{2T} \epsilon^{\mu\nu\alpha\beta} p_\nu u_\alpha (D u_\beta - \frac{1}{T} \partial_\beta T).$

Gradient of chemical potential $\mathcal{J}_{\text{chemical}}^\mu = a \frac{1}{(u \cdot p)} \epsilon^{\mu\nu\alpha\beta} p_\alpha u_\beta \partial_\nu \frac{\mu}{T},$

Electromagnetic fields $\mathcal{J}_{\text{EB}}^\mu = a \frac{1}{(u \cdot p) T} \epsilon^{\mu\nu\alpha\beta} p_\alpha u_\beta E_\nu + a \frac{B^\mu}{T},$

Y. Hidaka, SP, and D.L. Yang, PRD97, 016004 (2018); C. Yi, SP, D.L. Yang, arXiv:2106.00238

Out-of-equilibrium corrections

- **Polarization vector**

$$\mathcal{P}^z(p) = \int_{-1}^{+1} dY \mathcal{S}^z(p),$$

$$\mathcal{P}^y(p) = \int_{-1}^{+1} dY \mathcal{S}^y(p),$$

- **Polarization induced by thermal vorticity, shear viscous tensor and residual part of fluid acceleration**

$$\mathcal{S}_{\text{thermal}}^\mu(\mathbf{p}) = \frac{\hbar}{8m_\Lambda N} \int d\Sigma^\sigma p_\sigma f_V^{(0)} (1 - f_V^{(0)}) \epsilon^{\mu\nu\alpha\beta} p_\nu \partial_\alpha \frac{u_\beta}{T},$$

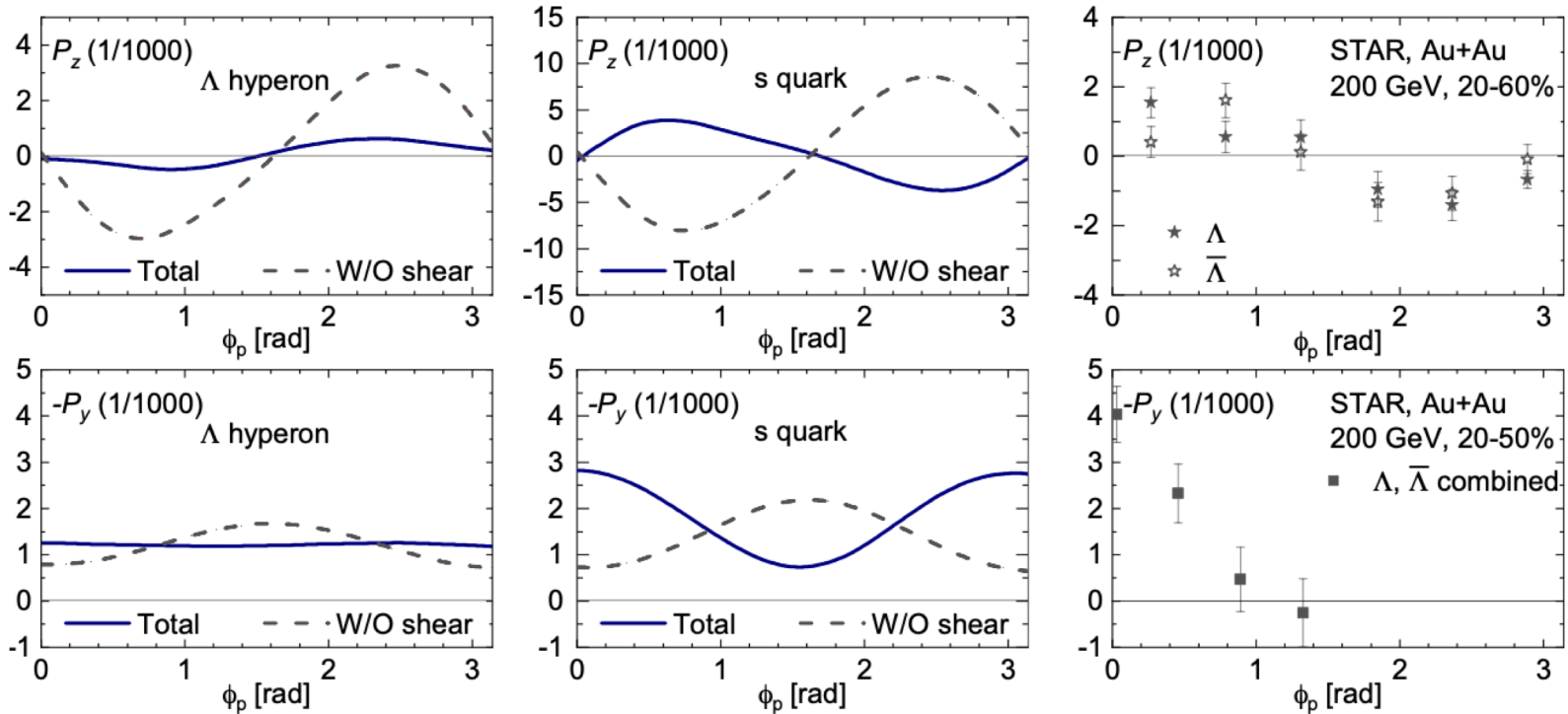
$$\mathcal{S}_{\text{shear}}^\mu(\mathbf{p}) = -\frac{\hbar}{4m_\Lambda N} \int d\Sigma \cdot p f_V^{(0)} (1 - f_V^{(0)}) \frac{\epsilon^{\mu\nu\alpha\beta} p_\alpha u_\beta}{(u \cdot p) T} \frac{1}{2} \{p^\sigma (\partial_\sigma u_\nu + \partial_\nu u_\sigma) - Du_\nu\}$$

$$\mathcal{S}_{\text{accT}}^\mu(\mathbf{p}) = -\frac{\hbar}{8m_\Lambda N} \int d\Sigma \cdot p f_V^{(0)} (1 - f_V^{(0)}) \frac{1}{T} \epsilon^{\mu\nu\alpha\beta} p_\nu u_\alpha (Du_\beta - \frac{1}{T} \partial_\beta T),$$

Recent related works

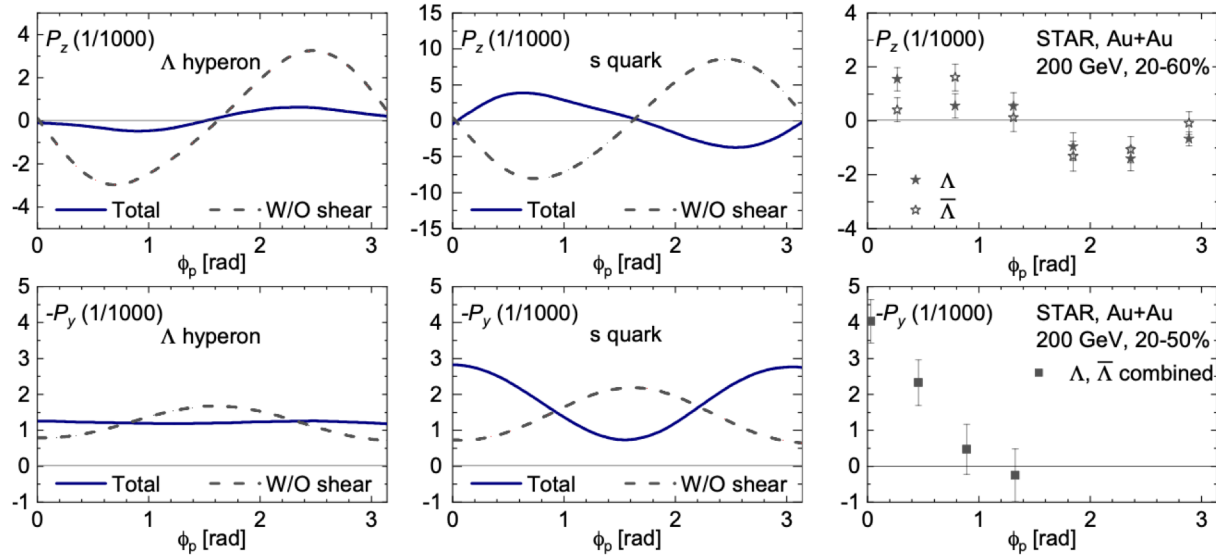
- Shear induced polarization draws some attentions.
- Shear induced Polarization from massless fermions (Theory):
Y. Hidaka, SP, and D.L. Yang, PRD97, 016004 (2018);
- Shear induced Polarization from massive fermions:
 - Theory:
S. Y. F. Liu, Y. Yin, 2103.09200
F. Becattini, M. Buzzegoli, A. Palermo, 2103.10917
 - Hydrodynamic simulations:
B. Fu, S. Y. F. Liu, L. Pang, H. Song, Y. Yin, 2103.10403
F. Becattini, M. Buzzegoli, A. Palermo, G. Inghirami, I. Karpenko, 2103.14621
C. Yi, SP, D.L. Yang, arXiv:2106.00238
- Global polarization induced by shear and gradient of chemical potential
S. Ryu, V. Jupic, C. Shen, arXiv:2106.08125

s quark scenario



B. Fu, S. Y. F. Liu, L. Pang, H. Song, Y. Yin, 2103.10403

s quark scenario: why it may works?



$$S_{\text{thermal}}^{\mu}(\mathbf{p}) = \frac{\hbar}{8m_{\Lambda}N} \int d\Sigma^{\sigma} p_{\sigma} f_V^{(0)} (1 - f_V^{(0)}) \epsilon^{\mu\nu\alpha\beta} p_{\nu} \partial_{\alpha} \frac{u_{\beta}}{T},$$

$$S_{\text{shear}}^{\mu}(\mathbf{p}) = \frac{\hbar}{4m_{\Lambda}N} \int d\Sigma \cdot p f_V^{(0)} (1 - f_V^{(0)}) \epsilon^{\mu\nu\alpha\beta} p_{\alpha} u_{\beta} \frac{1}{(u \cdot p) T} \frac{1}{2} \{p^{\sigma} (\partial_{\sigma} u_{\nu} + \partial_{\nu} u_{\sigma}) - D u_{\nu}\}$$

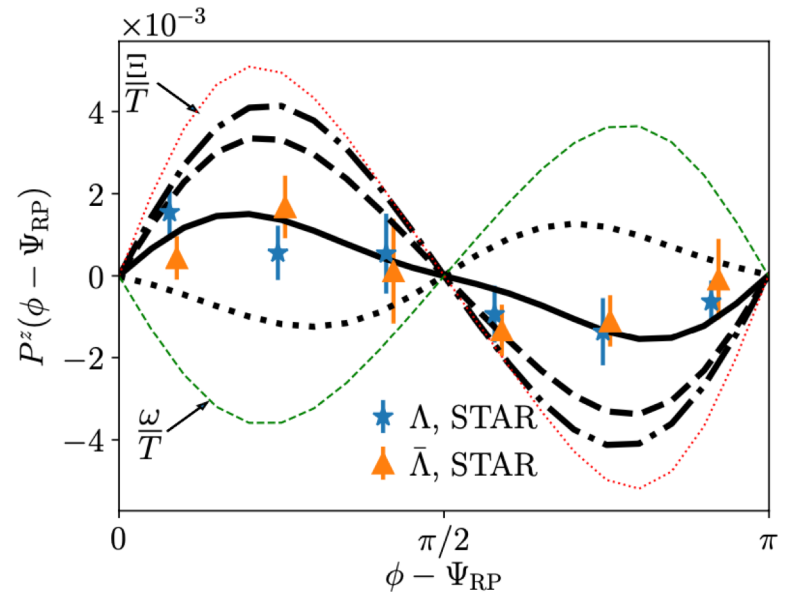
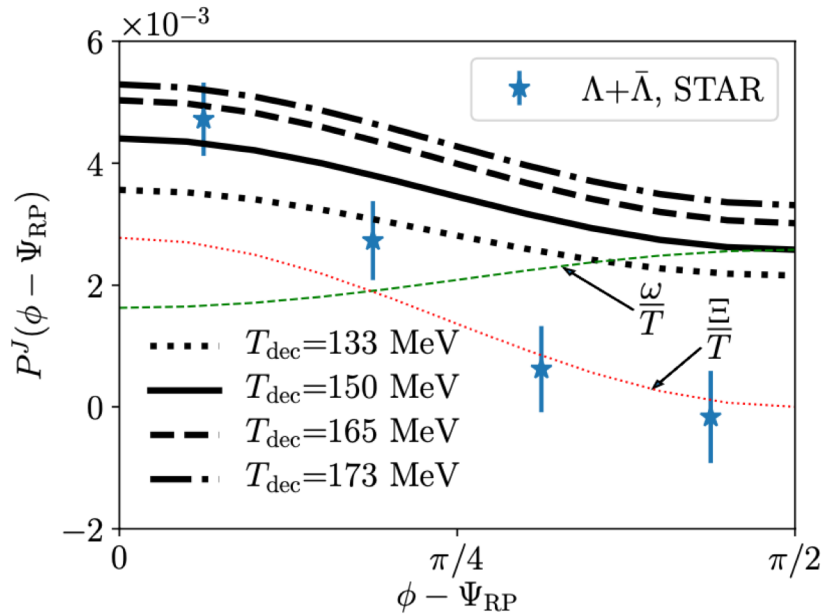
$m_{\Lambda} \rightarrow m_s$

$(u \cdot p) \sim m$

$m_s \simeq 0.3\text{GeV}$

$m_{\Lambda} \simeq 1.116\text{GeV}$

Isothermal local equilibrium



$$S_{\text{ILE}}^\mu(p) = \left| -\epsilon^{\mu\rho\sigma\tau} p_\tau \frac{\int_\Sigma d\Sigma \cdot p n_F(1-n_F) \left[\omega_{\rho\sigma} + 2\hat{t}_\rho \frac{p^\lambda}{\epsilon} \Xi_{\lambda\sigma} \right]}{8mT_{\text{dec}} \int_\Sigma d\Sigma \cdot p n_F} \right|$$

$$\omega_{\rho\sigma} = \frac{1}{2} (\partial_\sigma u_\rho - \partial_\rho u_\sigma) \quad \text{All gradient of temperature are neglected!}$$

$$\Xi_{\rho\sigma} = \frac{1}{2} (\partial_\sigma u_\rho + \partial_\rho u_\sigma) \quad \text{F. Becattini, M. Buzzegoli, A. Palermo, G. Inghirami, I. Karpenko, 2103.14621}$$

Hydrodynamic setup

- **(3+1) dimensional viscous hydrodynamic CLVisc**

L.G. Pang, H. Petersen, and X.N. Wang, Phys. Rev. C 97, 064918 (2018)

- **AMPT initial conditions**

Z.W. Lin, C. M. Ko, B.A. Li, B. Zhang, and S. Pal, Phys. Rev. C 72, 064901

- **EoS “sp95-pce”**

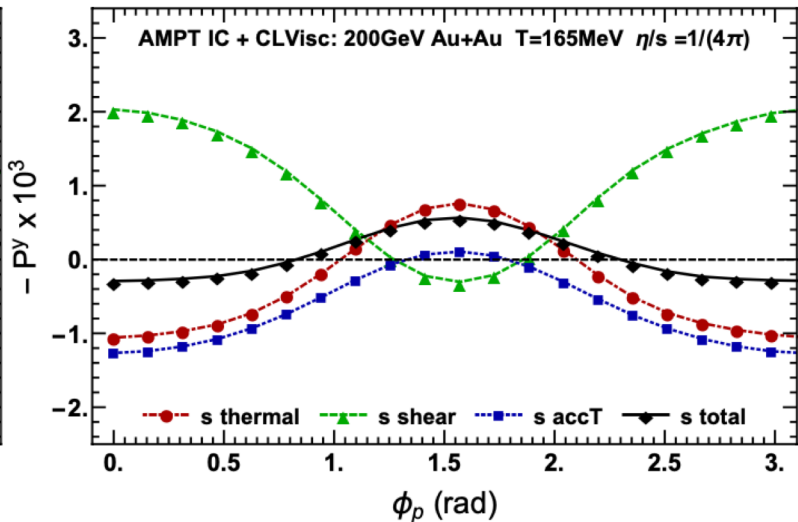
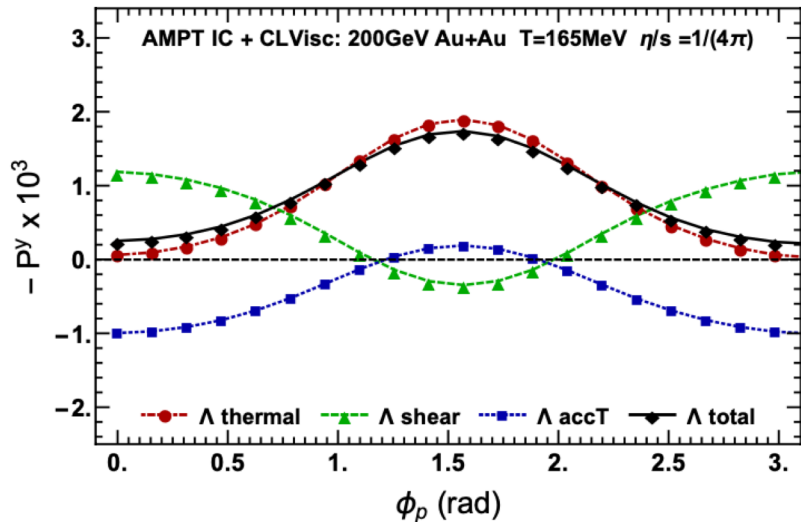
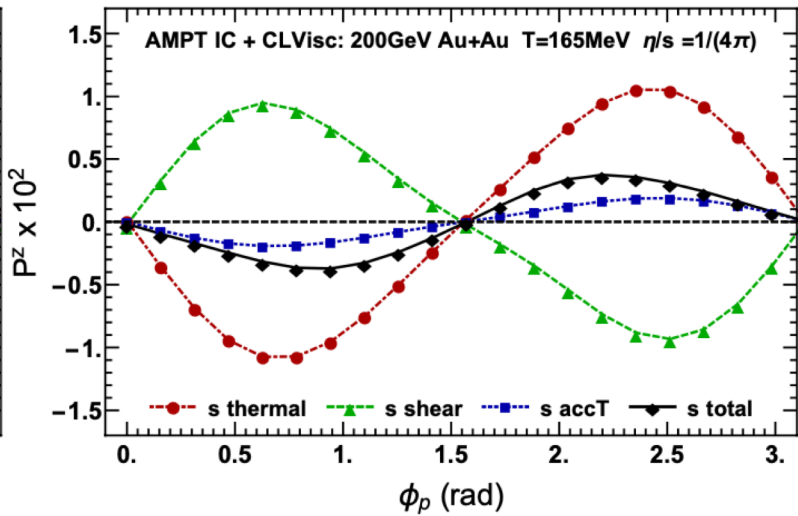
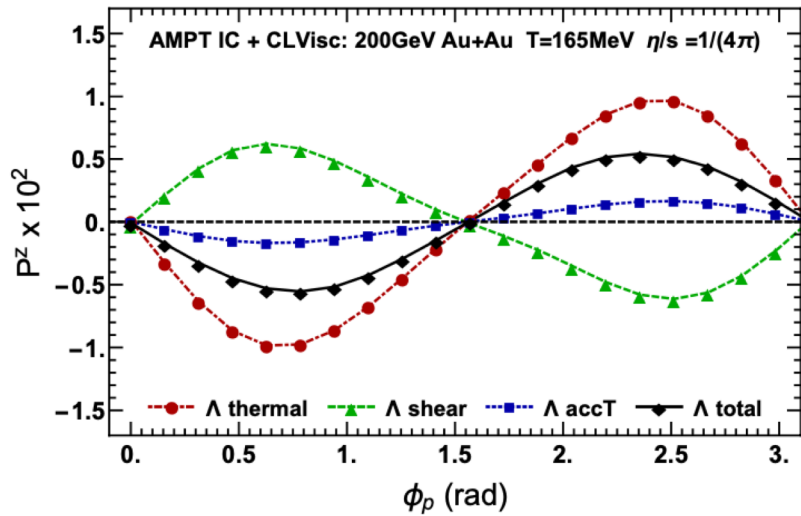
P. Huovinen and P. Petreczky, Nucl. Phys. A 837, 26 (2010)

- **Two scenarios**

B. Fu, S. Y. F. Liu, L. Pang, H. Song, and Y. Yin, (2021), 2103.10403

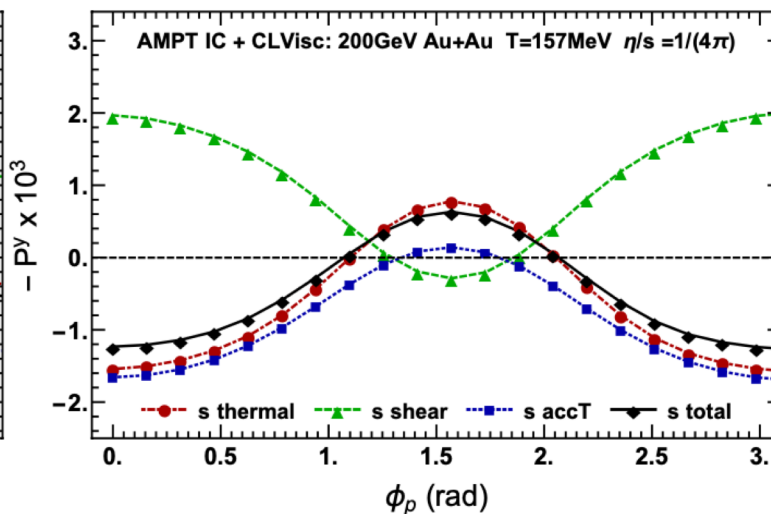
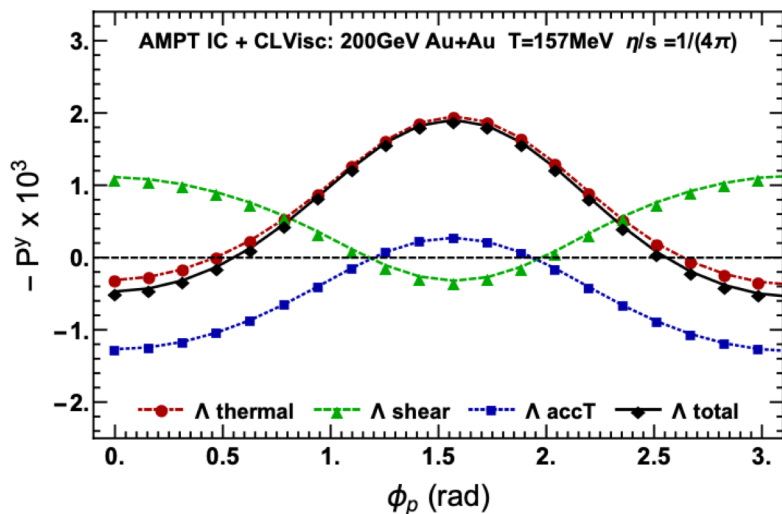
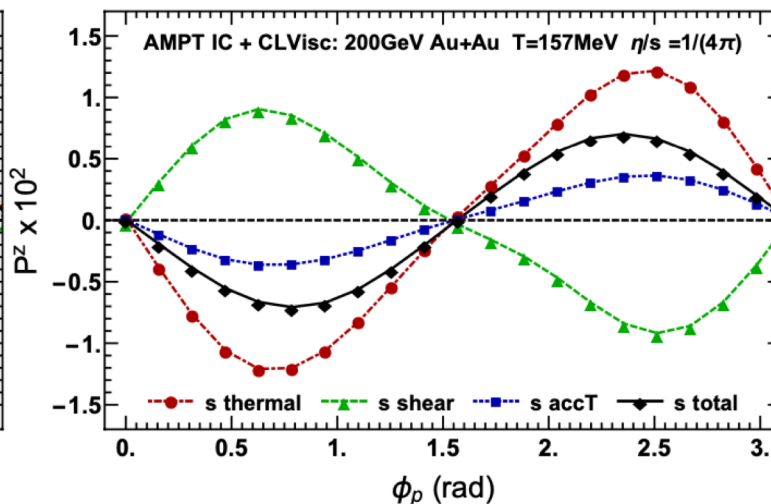
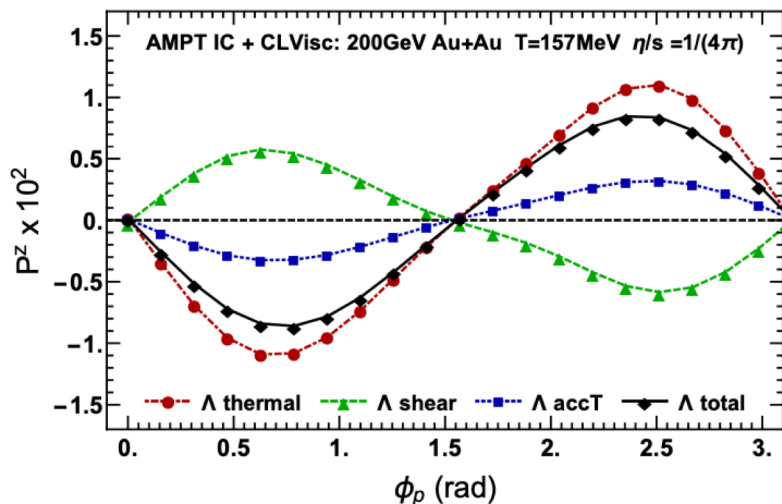
- **Lambda equilibrium scenario**
- **s quark equilibrium scenario**

Main result (I)



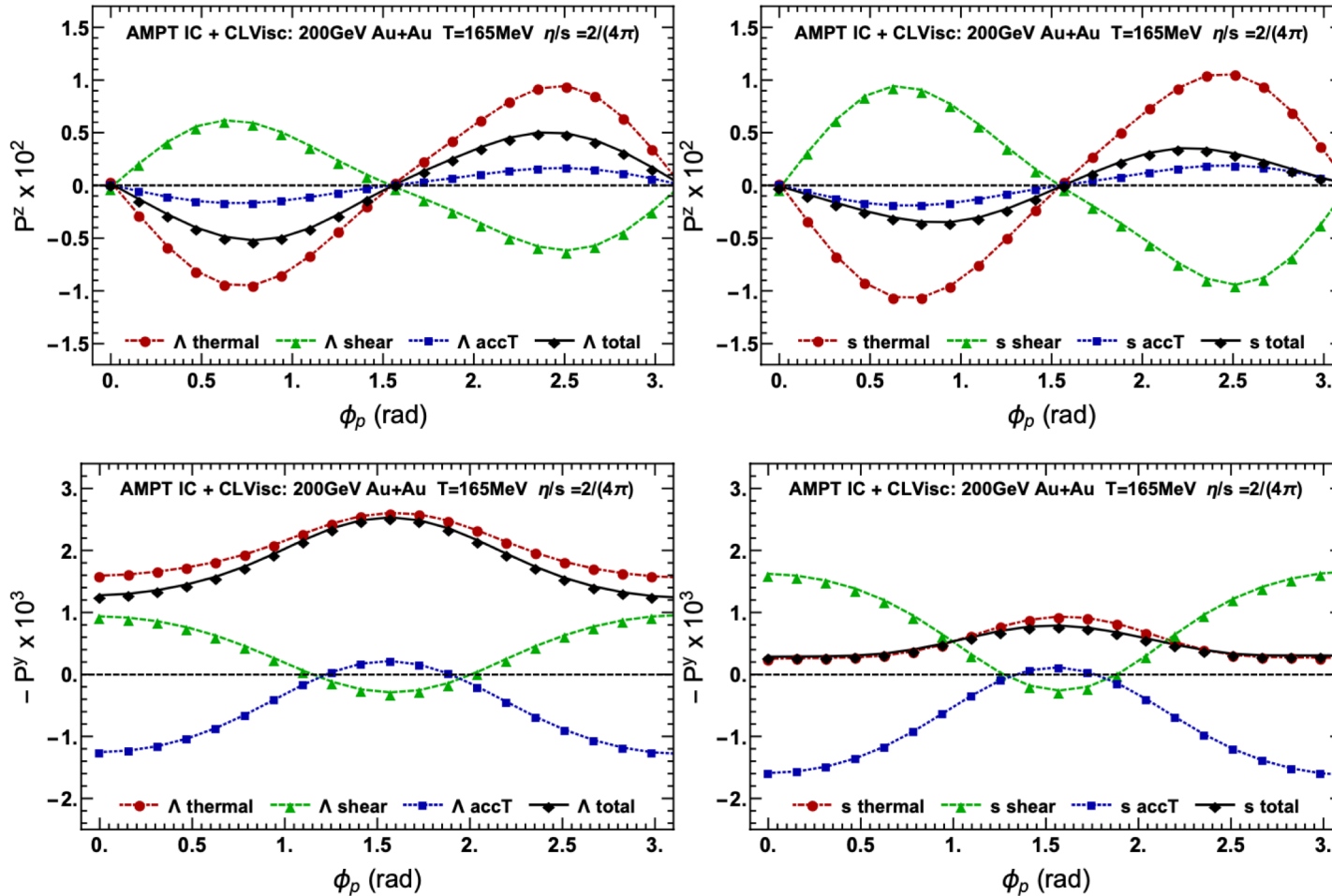
C. Yi, SP, D.L. Yang, [arXiv:2106.00238](https://arxiv.org/abs/2106.00238)

Reduce the freezeout temperature



C. Yi, SP, D.L. Yang, arXiv:2106.00238

Increase the η/s



C. Yi, SP, D.L. Yang, arXiv:2106.00238

Conclusion

- Shear induced polarization always give a “correct” sign.
- Total local polarization is very sensitive to EoS, freeze out temperature and η / s .
- The local spin polarization is still an open question. We still need to consider the out-of-equilibrium effects carefully through the spin hydrodynamics and the quantum kinetic theory with collisions.

Summary

Summary

Spin hydro in Belinfante form

- We have discussed the Belinfante energy momentum tensor, which is symmetric and gauge invariant.
- We have found the spin corrections to the dissipative terms, including quantum spin vorticity.
- By redefining the entropy flow,
 - we can reproduce the well-known results “in global equilibrium the spin chemical potential is related to thermal vorticity $\omega^{\mu\nu} = -T\omega_{\mu\nu}^{th}/2$.”
 - In Local equilibrium, we can rediscover the evolution equations for the spin effects, which is consistent with the one derived in canonical form.

Revisit local spin polarization

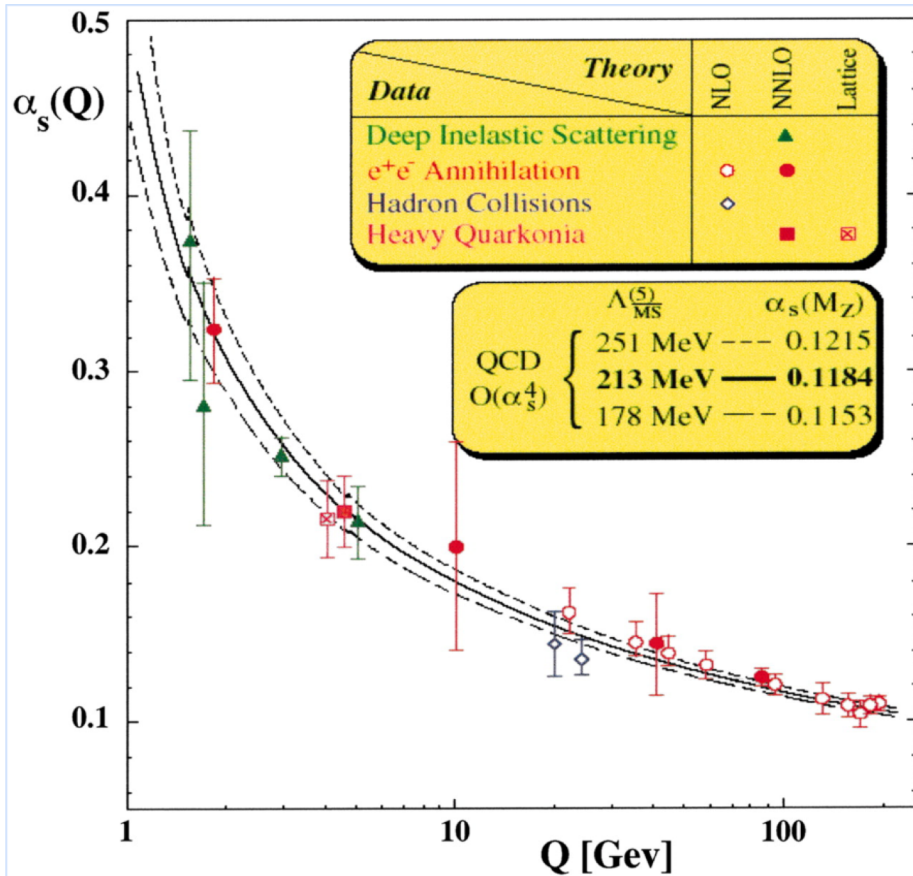
- The local spin polarization is still an open question. We still need to consider the out-of-equilibrium effects carefully through the spin hydrodynamics and the quantum kinetic theory with collisions.

Thank you for your time!

Any comments are welcome!

Backup

Asymptotic freedom of QCD



The Nobel Prize in Physics 2004



Photo from the Nobel Foundation archive.

David J. Gross

Prize share: 1/3



Photo from the Nobel Foundation archive.

H. David Politzer

Prize share: 1/3



Photo from the Nobel Foundation archive.

Frank Wilczek

Prize share: 1/3

The Nobel Prize in Physics 2004 was awarded jointly to David J. Gross, H. David Politzer and Frank Wilczek "for the discovery of asymptotic freedom in the theory of the strong interaction."

Quark Confinement

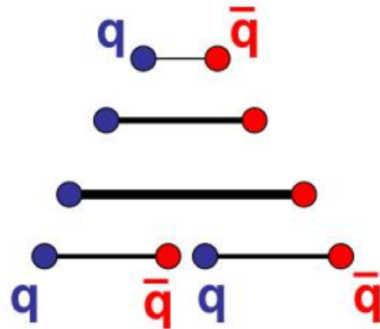
Quark Confinement:

庄子天下篇 ~ 300 B.C.

一尺之棰，日取其半，万世不竭

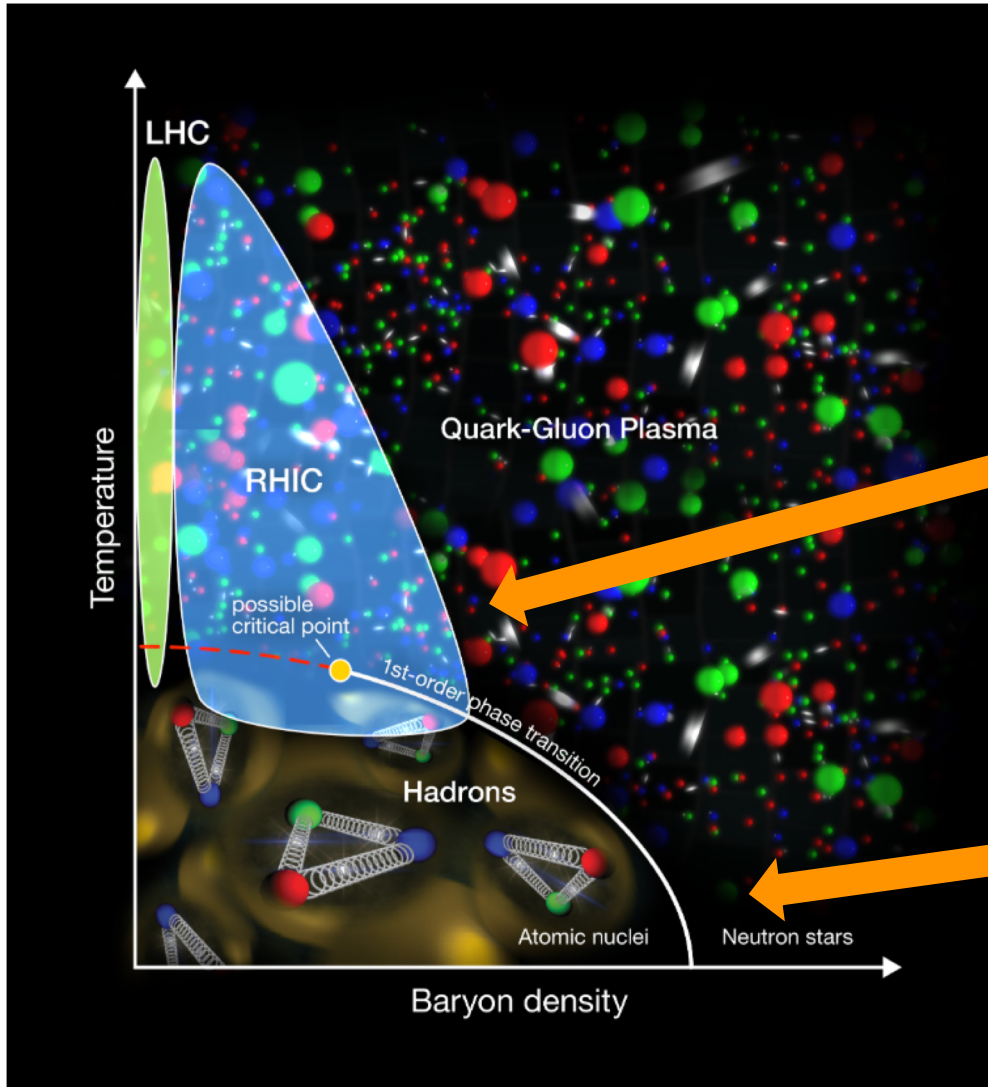
Take half from a foot long stick each day,
You will never exhaust it in million years.

QCD



Quark pairs can be produced from vacuum
No free quark can be observed

Deconfinement phase transition



High temperature

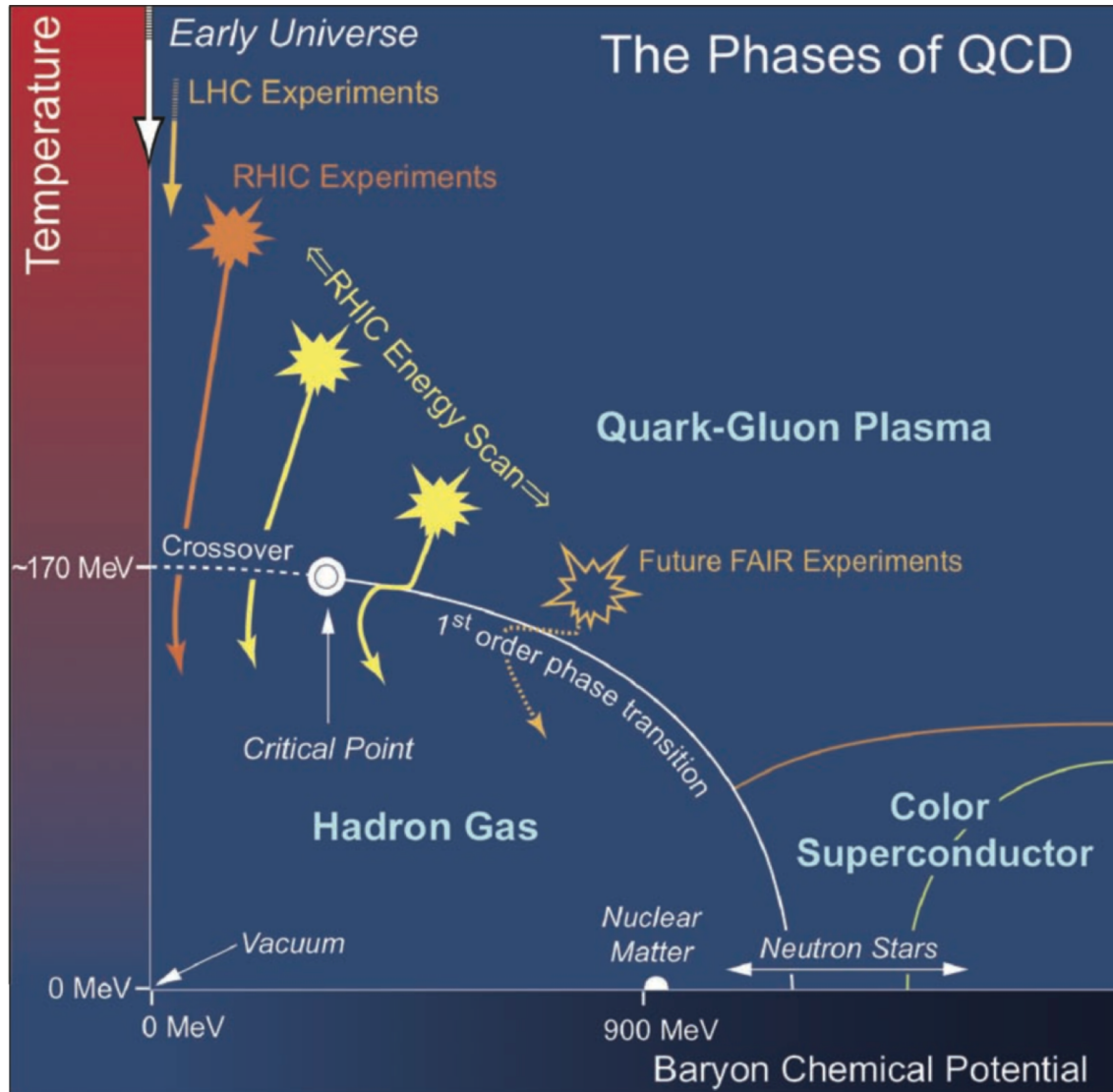


High pressure

Phases of QCD



**T.D. Lee (1974) and Collins (1975):
Heavy ion collision to create a new
form of matter!**



Symmetric energy-momentum tensor

$$T_{A,\text{Bel}}^{\mu\nu} \equiv -F^\mu_\alpha F^{\nu\alpha} + \frac{1}{4} g^{\mu\nu} F^{\alpha\beta} F_{\alpha\beta}, \quad (21)$$

$$T_{\psi,\text{Bel}}^{\mu\nu} \equiv \bar{\psi} i \gamma^\mu \overleftrightarrow{D}^\nu \psi + \frac{1}{4} \varepsilon^{\mu\nu\lambda\rho} \partial_\lambda (\bar{\psi} \gamma_5 \gamma_\rho \psi). \quad (22)$$

These are very desirable expressions and all the terms are manifestly gauge invariant, thus corresponding to physical observables in principle. At this point, one might have thought that $T_{\psi,\text{Bel}}^{\mu\nu}$ does not look symmetric with respect to μ and ν . In a quite non-trivial way one can prove that the above fermionic part is alternatively expressed as $T_{\psi,\text{Bel}}^{\mu\nu} = \bar{\psi} i \gamma^{(\mu} \overleftrightarrow{D}^{\nu)} \psi$, which is obviously symmetric.

$$\frac{1}{2} = J_q + J_g \quad (23)$$

Coming back to the angular momentum, we can introduce the Belinfante “improved” form for the angular momentum, i.e.,

$$J_{\text{Bel}}^{\lambda\mu\nu} \equiv J^{\lambda\mu\nu} + \partial_\rho (x^\mu K_{\text{Bel}}^{\rho\lambda\nu} - x^\nu K_{\text{Bel}}^{\rho\lambda\mu}). \quad (24)$$

Main equations

$$\partial_\mu T_{Bel}^{\mu\nu} = 0.$$

$$T_{Bel}^{\mu\nu} = T_{can}^{\mu\nu} + \partial_\lambda K_{Bel}^{\lambda\mu\nu} = T_0^{\mu\nu} + T_{(1)}^{\mu\nu}$$

$$T_{(1)}^{\mu\nu} = \underset{\substack{\uparrow \\ \text{heat flow}}}{h^\mu u^\nu} + \underset{\substack{\uparrow \\ \text{viscous tensor}}}{h^\nu u^\mu} + \underset{\substack{\uparrow \\ \text{spin density tensor}}}{\pi^{\mu\nu}} + \frac{1}{2} \partial_\lambda (u^\mu S^{\nu\lambda} + u^\nu S^{\mu\lambda})$$

**spin corrections to the
energy momentum tensor**

$$\partial_\lambda (u^\lambda S^{\rho\sigma}) = q^\rho u^\sigma - q^\sigma u^\rho + \phi^{\rho\sigma}$$

$$q^\mu = \lambda[(u \cdot \partial)u^\mu + \frac{1}{T}\Delta^{\mu\nu}\partial_\nu T - 4\omega^{\mu\nu}u_\nu],$$

$$\phi^{\mu\nu} = 2\gamma[T\omega_{th}^{\mu\nu} + 2(g^{\mu\alpha} - u^\mu u^\alpha)(g^{\nu\beta} - u^\nu u^\beta)\omega_{\alpha\beta}]/T.$$

Kinetic theory

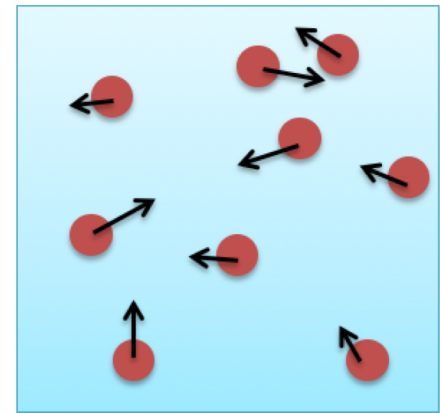
- **Assumptions:**

Mean free path \gg collision length scaling

- “distribution function” $f(x,p,t)$

how many particles in a small
volume of phase space $(x+dx, p+dp)$

e.g. Fermi-Dirac distribution function



- **Ordinary kinetic theory: Boltzmann equation**

Dynamical evolution equation for $f(x,p,t)$

Ordinary Boltzmann equation

Particle's velocity:

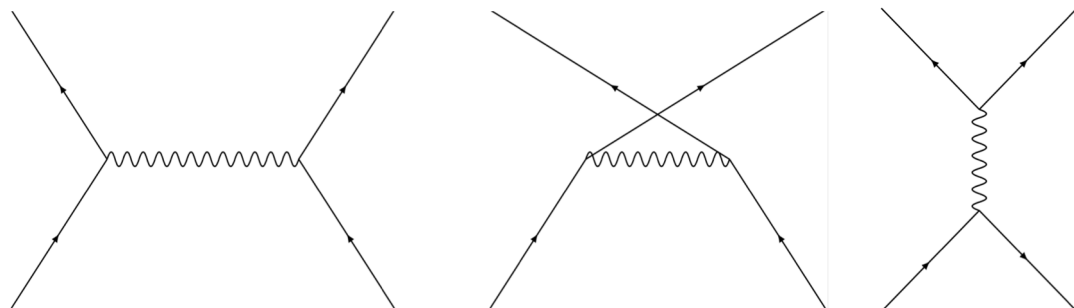
$$\dot{\mathbf{x}} = \frac{\partial \varepsilon}{\partial \mathbf{p}},$$

ε : Particle's energy

Lorentz force:

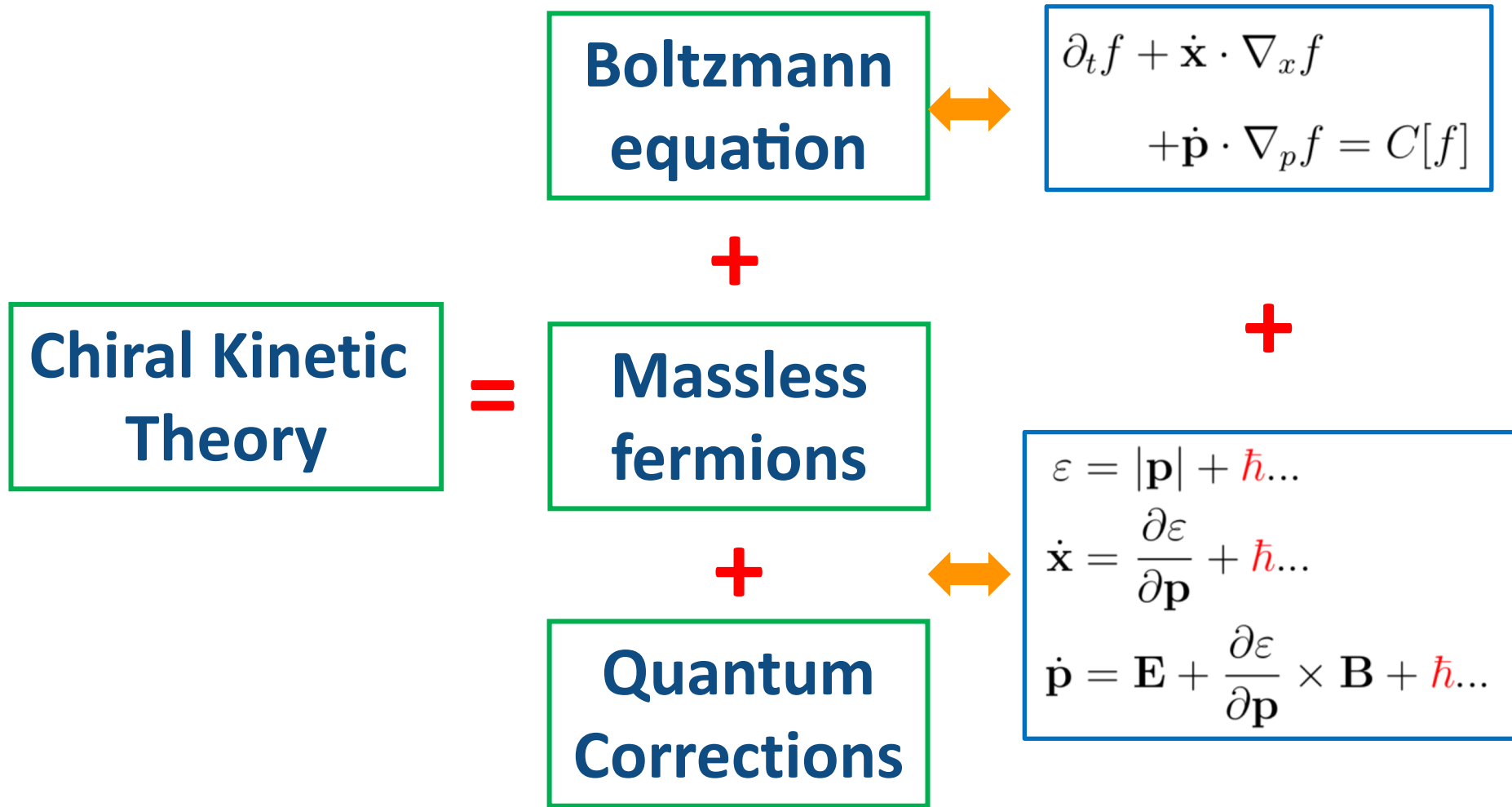
$$\dot{\mathbf{p}} = \mathbf{E} + \frac{\partial \varepsilon}{\partial \mathbf{p}} \times \mathbf{B},$$

$$\partial_t f + \dot{\mathbf{x}} \cdot \nabla_x f + \dot{\mathbf{p}} \cdot \nabla_p f = C[f],$$



Collision term:

What is Chiral kinetic theory?



Wigner function (I)

- **Wigner operator**

$$\hat{W}_{\alpha\beta} = \int \frac{d^4y}{(2\pi)^4} e^{-ip \cdot y} \bar{\psi}_{\beta}(x_+) U(x_+, x_-) \psi_{\alpha}(x_-),$$

- **Wigner function:**

Gauge link $U(x_+, x_-) \equiv e^{-iQ \int_{x_-}^{x_+} dz^{\mu} A_{\mu}(z)}$,

$$W(x, p) = \langle : \hat{W}(x, p) : \rangle$$

W operator in thermal ensemble average and normal ordering of the operators

- **Physical meaning: QFT version density matrix**

Vasak, Gyulassy, Elze, Ann. Phys. (N.Y.) 173, 462 (1987);

Elze, Heinz, Phys.Rep.183,81(1989).

Wigner function (II)

- Master equations for Wigner function:

$$\gamma_\mu \left(p^\mu + \frac{i}{2} \nabla^\mu \right) W(x, p) = 0, \quad \nabla^\mu \equiv \partial_x^\mu - Q F^\mu{}_\nu \partial_p^\nu$$

- Matrix decomposition

$$W = \frac{1}{4} \left[\mathcal{F} + i\gamma^5 \mathcal{P} + \gamma^\mu \mathcal{V}_\mu + \gamma^5 \gamma^\mu \mathcal{A}_\mu + \frac{1}{2} \sigma^{\mu\nu} \mathcal{J}_{\mu\nu} \right],$$

Charge
current

$$\mathcal{V}^\mu = \int \frac{d^4 y}{(2\pi)^4} e^{-ip \cdot y} \langle : \bar{\psi}_\beta \left(x + \frac{1}{2} y \right) \gamma^\mu U \left(x + \frac{1}{2} y, x - \frac{1}{2} y \right) \psi_\alpha \left(x - \frac{1}{2} y \right) : \rangle$$

Chiral
current

$$\mathcal{A}_\mu = \int \frac{d^4 y}{(2\pi)^4} e^{-ip \cdot y} \langle : \bar{\psi}_\beta \left(x + \frac{1}{2} y \right) \gamma^\mu \gamma^5 U \left(x + \frac{1}{2} y, x - \frac{1}{2} y \right) \psi_\alpha \left(x - \frac{1}{2} y \right) : \rangle$$

Wigner function (III)

- Left and right handed currents

$$\mathcal{J}_\mu^s(x, p) = \frac{1}{2}[\mathcal{V}_\mu(x, p) + s\mathcal{A}_\mu(x, p)], \quad s = \pm$$

- In massless limit

$$p^\mu \mathcal{J}_\mu^s(x, p) = 0,$$

$$\nabla^\mu \mathcal{J}_\mu^s(x, p) = 0,$$

$$2s(p^\lambda \mathcal{J}_s^\rho - p^\rho \mathcal{J}_s^\lambda) = -\epsilon^{\mu\nu\lambda\rho} \nabla_\mu \mathcal{J}_\nu^s.$$

hbar expansion

- **hbar (gradient) expansion**

$$\mathcal{J}_\mu^s(x, p) = \mathcal{J}_{\mu, (0)}^s(x, p) + \mathcal{J}_{\mu, (1)}^s(x, p) + \cdots,$$

- **Leading order** is the classical currents.

We introduce the initial distribution function $f(x, p)$ as input.

$$\mathcal{J}_{(0)s}^\rho(x, p) = p^\rho f_s \delta(p^2),$$

- **Next-to-leading order**

$$\mathcal{J}_{(1)s}^\rho(x, p) = -\frac{s}{2} \tilde{\Omega}^{\rho\lambda} p_\lambda \frac{df_s}{dp_0} \delta(p^2) - \frac{s}{p^2} e \tilde{F}^{\rho\lambda} p_\lambda f_s \delta(p^2).$$

$$\Omega_{\nu\sigma} = \frac{1}{2} (\partial_\nu u_\sigma - \partial_\sigma u_\nu), \text{ and } \Omega^{\mu\nu} = -\frac{1}{2} \epsilon^{\mu\nu\rho\sigma} \tilde{\Omega}_{\rho\sigma}.$$

Currents

- Integral over momentum

$$j_s^\mu = \int d^4p \mathcal{J}_s^\mu = n_s u^\mu + \xi_{B,s} B^\mu + \xi_s \omega^\mu,$$

Charge
current

$$j^\mu = \sum_{s=\pm} j_s^\mu = n u^\mu + \xi_B B^\mu + \xi \omega^\mu,$$

Chiral
current

$$j_5^\mu = \sum_{s=\pm} s j_s^\mu = n_5 u^\mu + \xi_{B5} B^\mu + \xi_5 \omega^\mu,$$

$$\xi_B = \frac{e}{2\pi^2} \mu_5,$$

$$\xi_{B5} = \frac{e}{2\pi^2} \mu,$$

$$\xi = \frac{1}{\pi^2} \mu \mu_5,$$

$$\xi_5 = \frac{1}{6} T^2 + \frac{1}{2\pi^2} (\mu^2 + \mu_5^2)$$

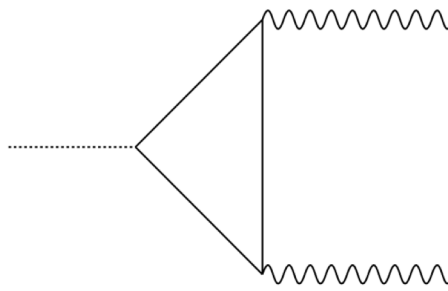
Gao, Liang, Pu, Q. Wang, X.N. Wang, PRL 109, 232301 (2012)

Chiral anomaly

- Chiral anomaly

$$\partial_{\mu} j^{\mu} = 0,$$

$$\partial_{\mu} j_5^{\mu} = -\frac{e^2}{2\pi^2} E \cdot B.$$



↑
We reproduce the
chiral anomaly
from the kinetic
theory!!!

Gao, Liang, Pu, Q. Wang, X.N. Wang, PRL 109, 232301 (2012)

Derivation of chiral kinetic theory

- **Constrain equation.**

$$\nabla^\mu \mathcal{J}_\mu^s(x, p) = 0,$$

- **We can insert our results into this equation and get the constraint equation for distribution function.**
- **Then, we need to integral over p_0 to get 3-dim form.**

J.W. Chen, SP, Q. Wang, X.N. Wang, PRL (2013);

Y. Hidaka, SP, D.L. Yang, PRD(RC) (2017)

Review: *Gao, Liang, Wang, Int.J.Mod.Phys A 36 (2021), 2130001*

Hidaka, SP, D.L. Yang, Q. Wang, invited review, in preparation

Chiral kinetic equation

$$\sqrt{G}\partial_t f + \sqrt{G}\dot{\mathbf{x}} \cdot \nabla_x f + \sqrt{G}\dot{\mathbf{p}} \cdot \nabla_p f = C[f].$$

- Particle's effective velocity:

$$\sqrt{G}\dot{\mathbf{x}} = \frac{\partial \varepsilon}{\partial \mathbf{p}} + \hbar \left(\frac{\partial \varepsilon}{\partial \mathbf{p}} \cdot \boldsymbol{\Omega} \right) \mathbf{B} + \hbar \mathbf{E} \times \boldsymbol{\Omega},$$

- Effective force:

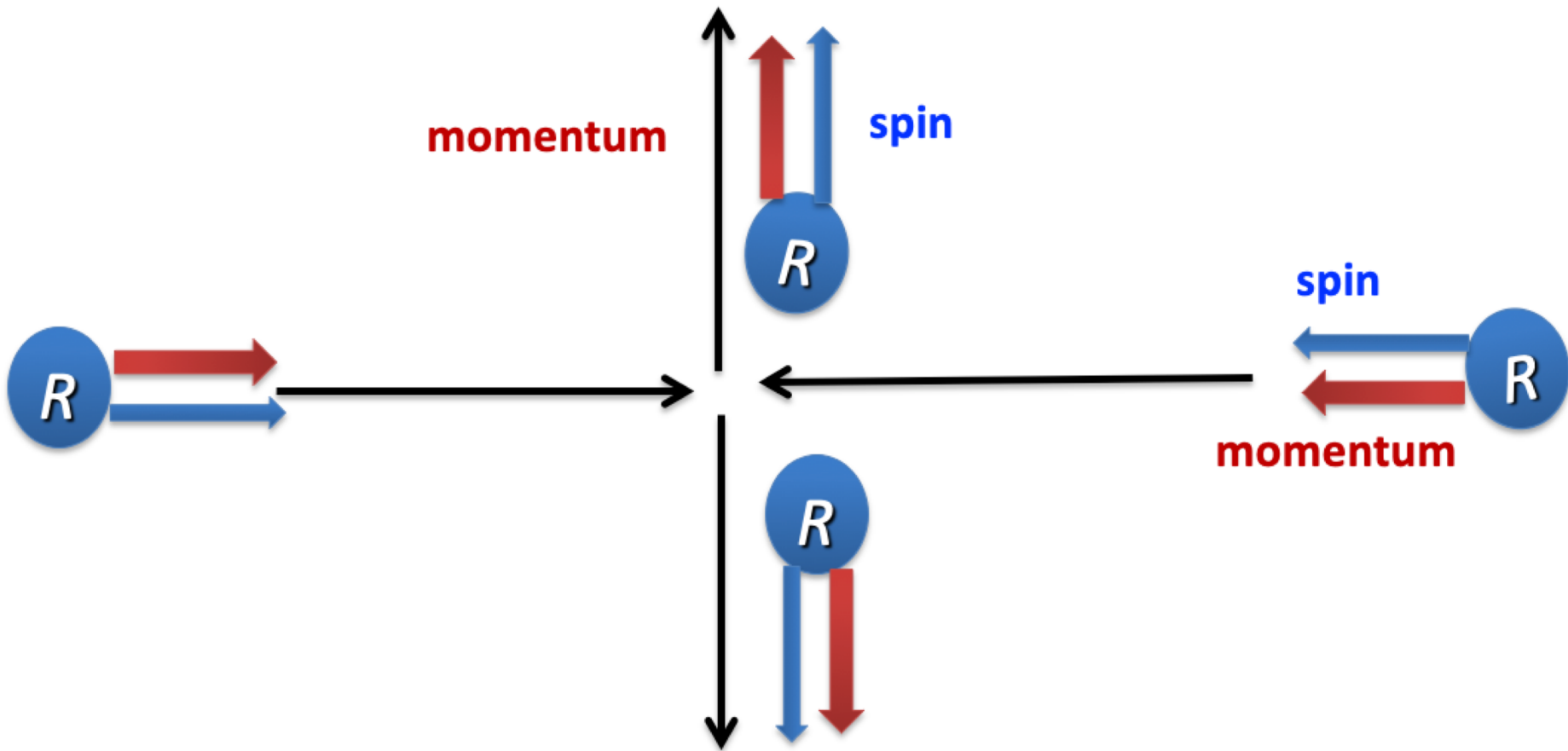
$$\sqrt{G}\dot{\mathbf{p}} = \mathbf{E} + \frac{\partial \varepsilon}{\partial \mathbf{p}} \times \mathbf{B} + \hbar (\mathbf{E} \cdot \mathbf{B}) \boldsymbol{\Omega},$$

- Berry curvature

$$\sqrt{G} = 1 + \hbar \mathbf{B} \cdot \boldsymbol{\Omega}, \quad \boldsymbol{\Omega} = \frac{\mathbf{p}}{2|\mathbf{p}|^3},$$

Side-jump (I)

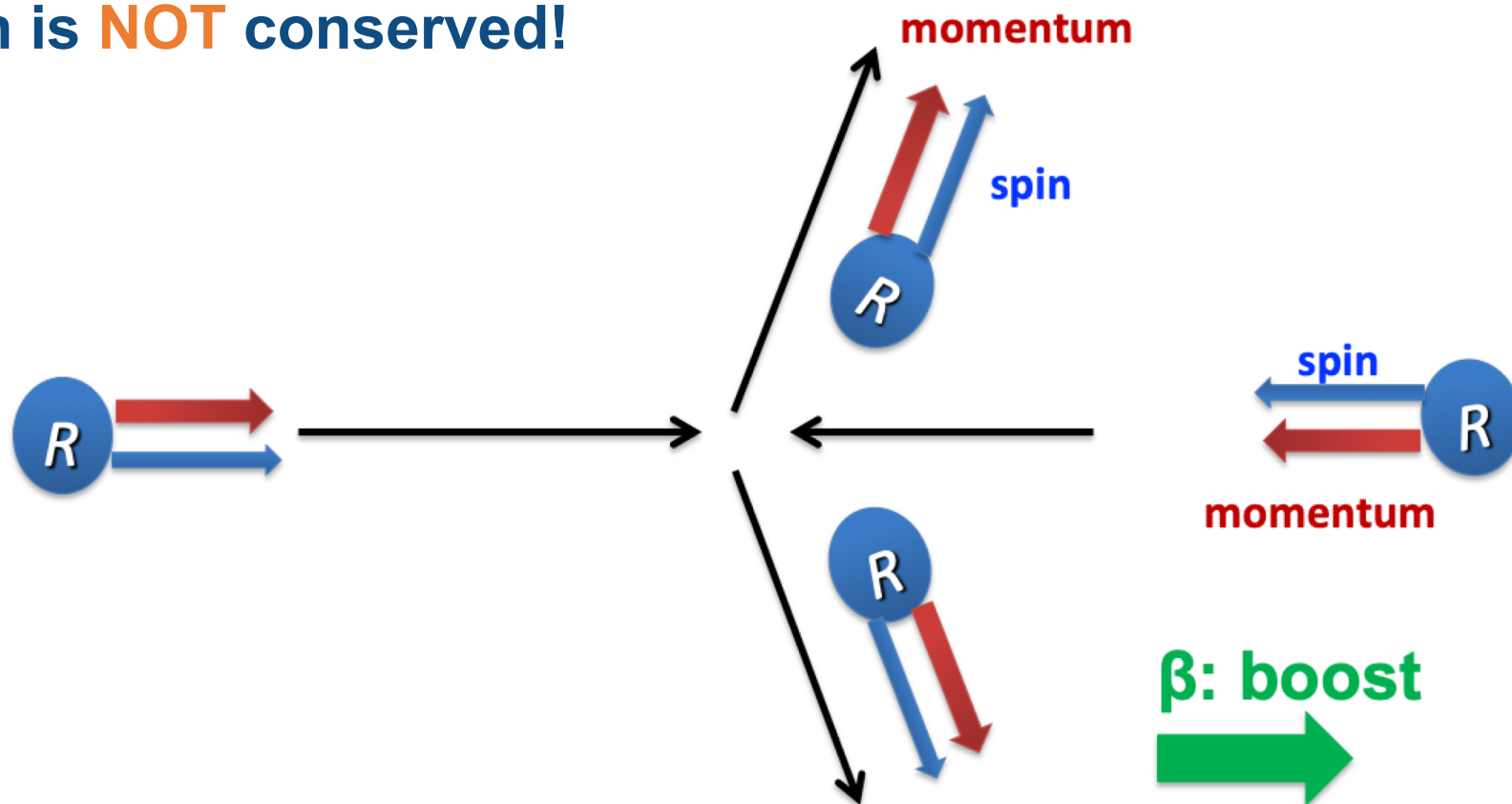
Orbital angular momentum and spin are conserved separately



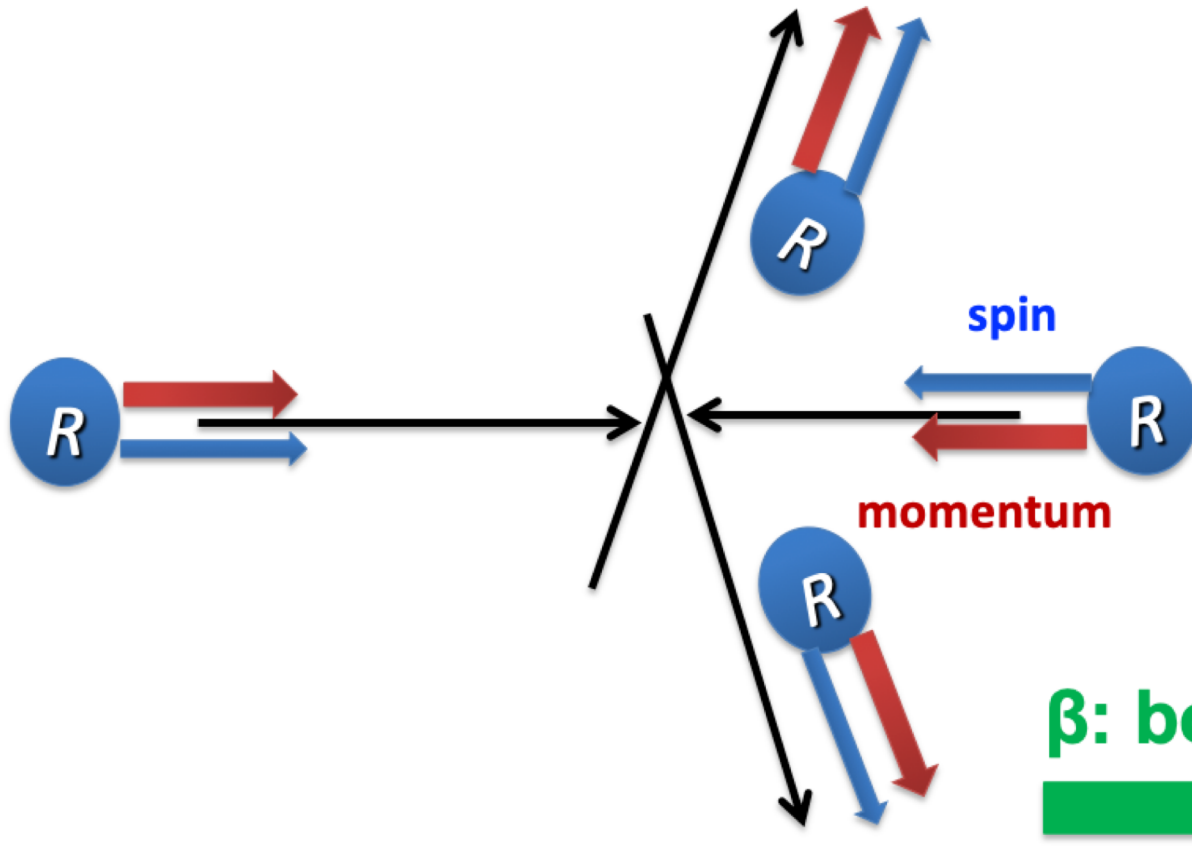
Chen, Son, Stephanov, Yee, Yin, PRL, (2014)

Side-jump (II)

Orbital angular momentum ?
Spin is **NOT** conserved!



Side-jump (III)



x has a shift!!!
“Side-jump” :

$$\mathbf{x}' = \mathbf{x} + \boldsymbol{\beta}t + \delta\mathbf{x},$$

$$\mathbf{p}' = \mathbf{p} + \boldsymbol{\beta}\varepsilon + \delta\mathbf{p},$$

$$\delta\mathbf{x} = \hbar \frac{\boldsymbol{\beta} \times \hat{\mathbf{p}}}{2|\mathbf{p}|},$$

$$\delta\mathbf{p} = \hbar \frac{\boldsymbol{\beta} \times \hat{\mathbf{p}}}{2|\mathbf{p}|} \times \mathbf{B}$$

β: boost

Chen, Son, Stephanov, PRL, (2015);
Y. Hidaka, SP, D.L. Yang, PRD (2016)

Non-trivial Lorentz symmetry

- Quantum field theory

$$j^\mu = \bar{\psi} \sigma^\mu \psi \rightarrow \Lambda_\nu^\mu j^\nu$$

- Lorentz transformation

$$x^{\mu'} = \Lambda_\nu^\mu x^\nu, \quad p^{\mu'} = \Lambda_\nu^\mu p^\nu,$$

$$f'(x', p', t') = f(x', p', t') + \hbar N^\mu (\partial_\mu^x + F_{\nu\mu} \partial_p^\nu) f,$$

Infinitesimal
Lorentz
Transform

$$\delta \mathbf{x} = \hbar \frac{\boldsymbol{\beta} \times \hat{\mathbf{p}}}{2|\mathbf{p}|},$$
$$\delta \mathbf{p} = \hbar \frac{\boldsymbol{\beta} \times \hat{\mathbf{p}}}{2|\mathbf{p}|} \times \mathbf{B}$$

*Chen, Son, Stephanov, PRL, (2015);
Y. Hidaka, SP, D.L. Yang, PRD (2016)*