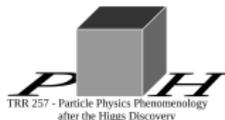


Introduction to Monte Carlo Event Generators

Stefan Gieseke

*Institut für Theoretische Physik
KIT*

Lectures at MCnet Beijing School 2021
University of Chinese Academy of Sciences, Beijing, China
28 June-2 July 2021

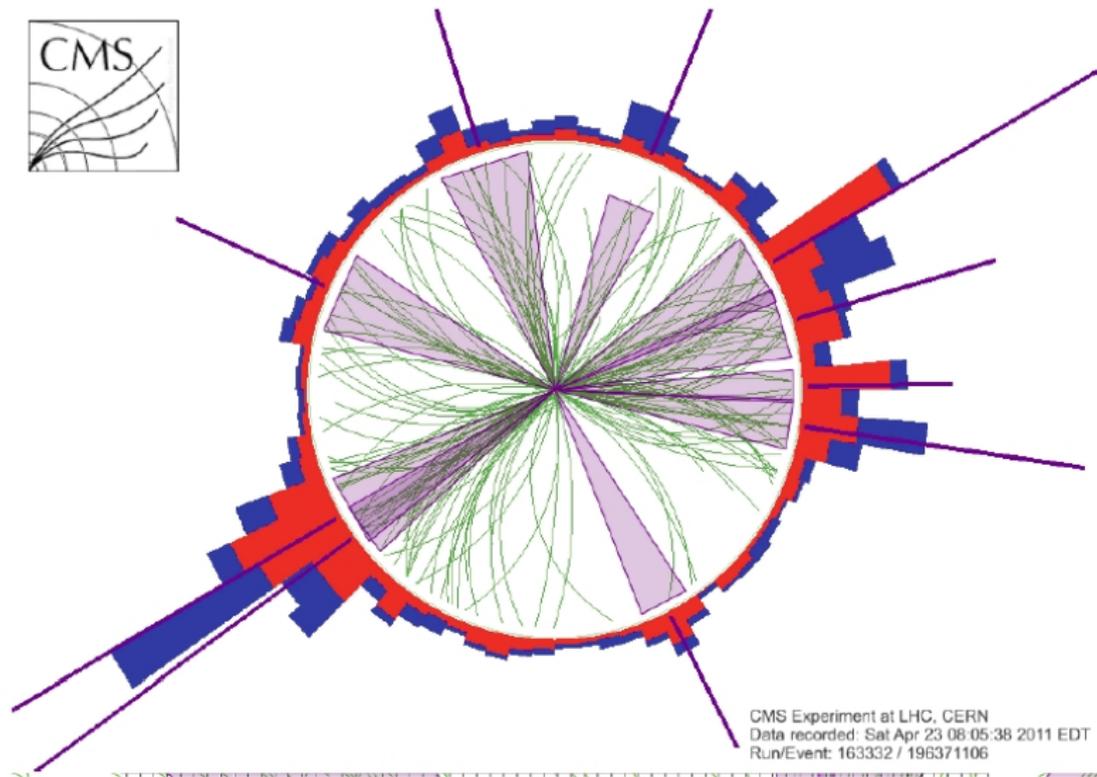


Motivation: jets



[Google Images]

Motivation: jets (at LHC of course)



Why Monte Carlos?

We want to understand

$$\mathcal{L}_{\text{int}} \longleftrightarrow \text{Final states} .$$

Why Monte Carlos?

LHC experiments require
sound understanding of signals and *backgrounds*.



Full detector simulation.



Fully exclusive hadronic final state.

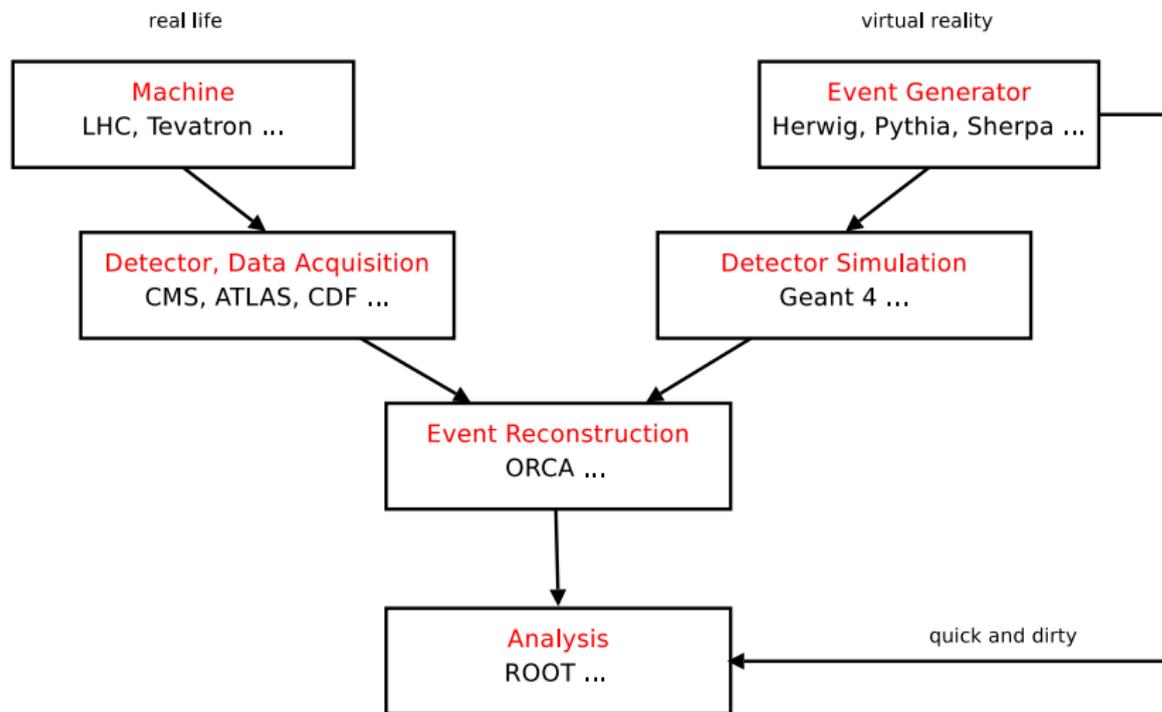


Monte Carlo event generator with
parton shower, hadronization model, decays of unstable
particles.



Parton level computations.

Experiment and Simulation



Monte Carlo Event Generators

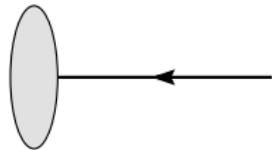
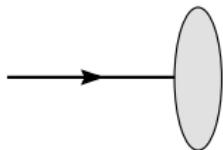
- Complex final states in full detail (jets).
- Arbitrary observables and cuts from final states.
- Studies of new physics models.

- Rates and topologies of final states.
- Background studies.
- Detector Design.
- Detector Performance Studies (Acceptance).

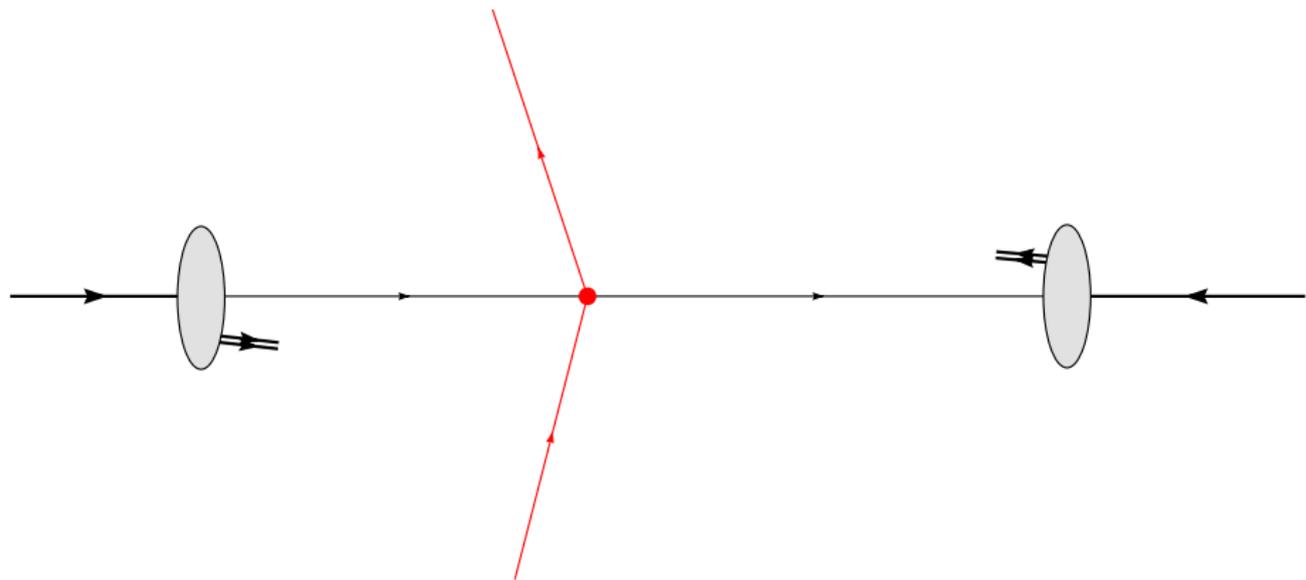
- *Obvious* for calculation of observables on the quantum level

$$|A|^2 \longrightarrow \text{Probability.}$$

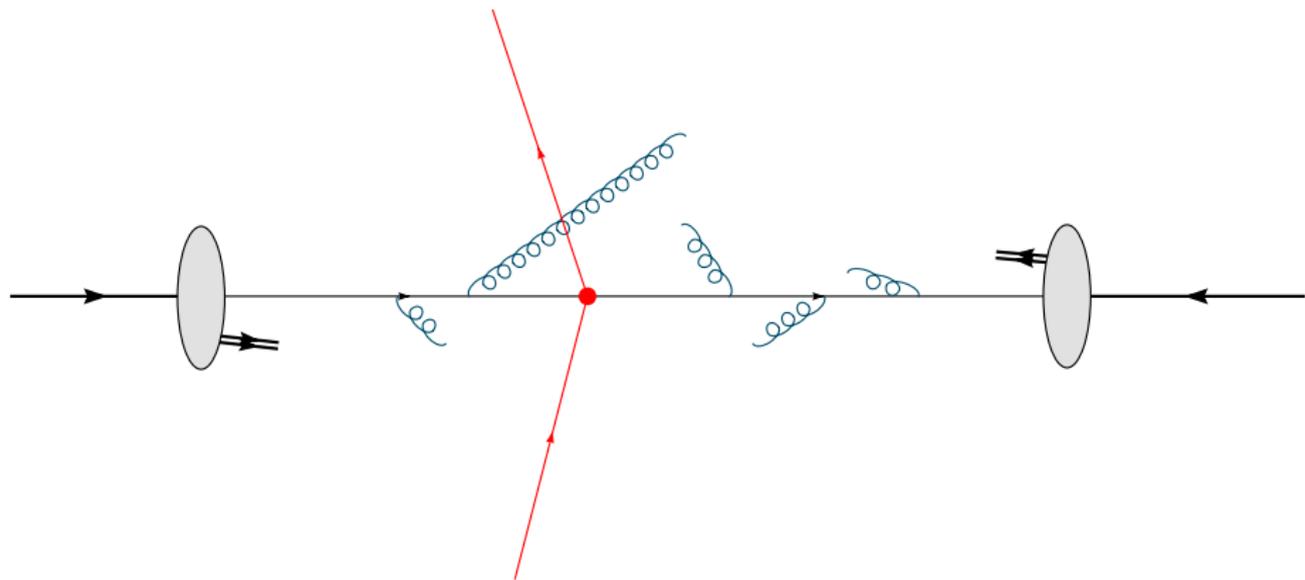
pp Event Generator



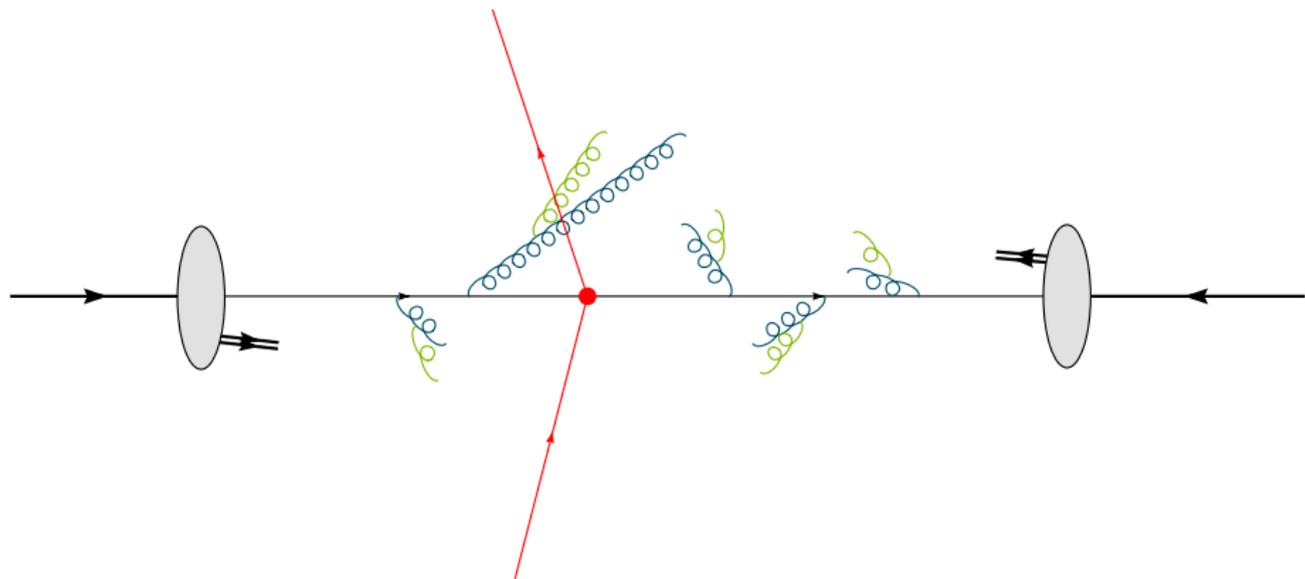
pp Event Generator



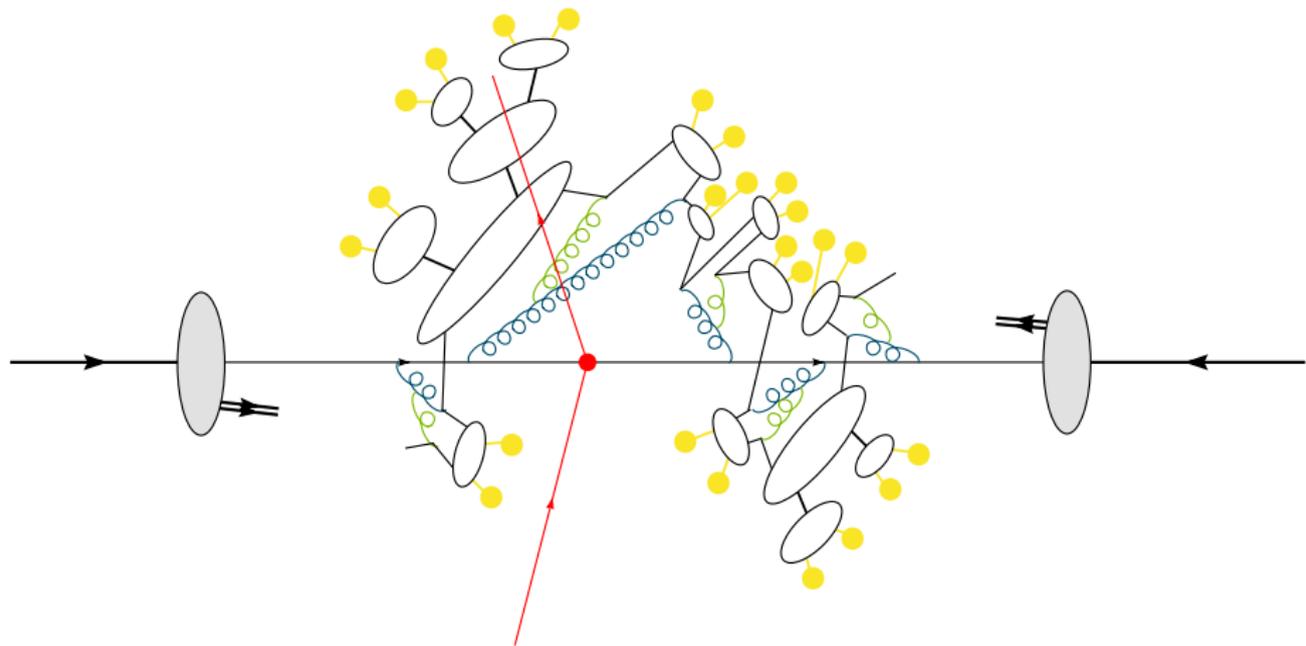
pp Event Generator



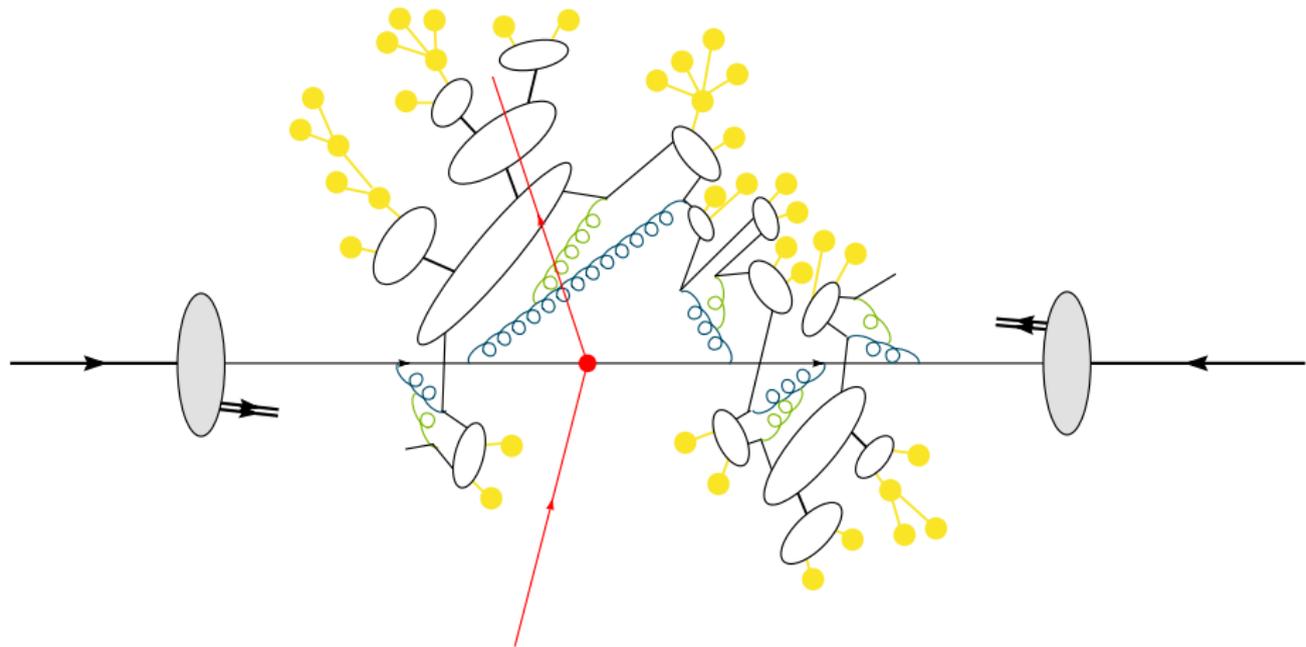
pp Event Generator



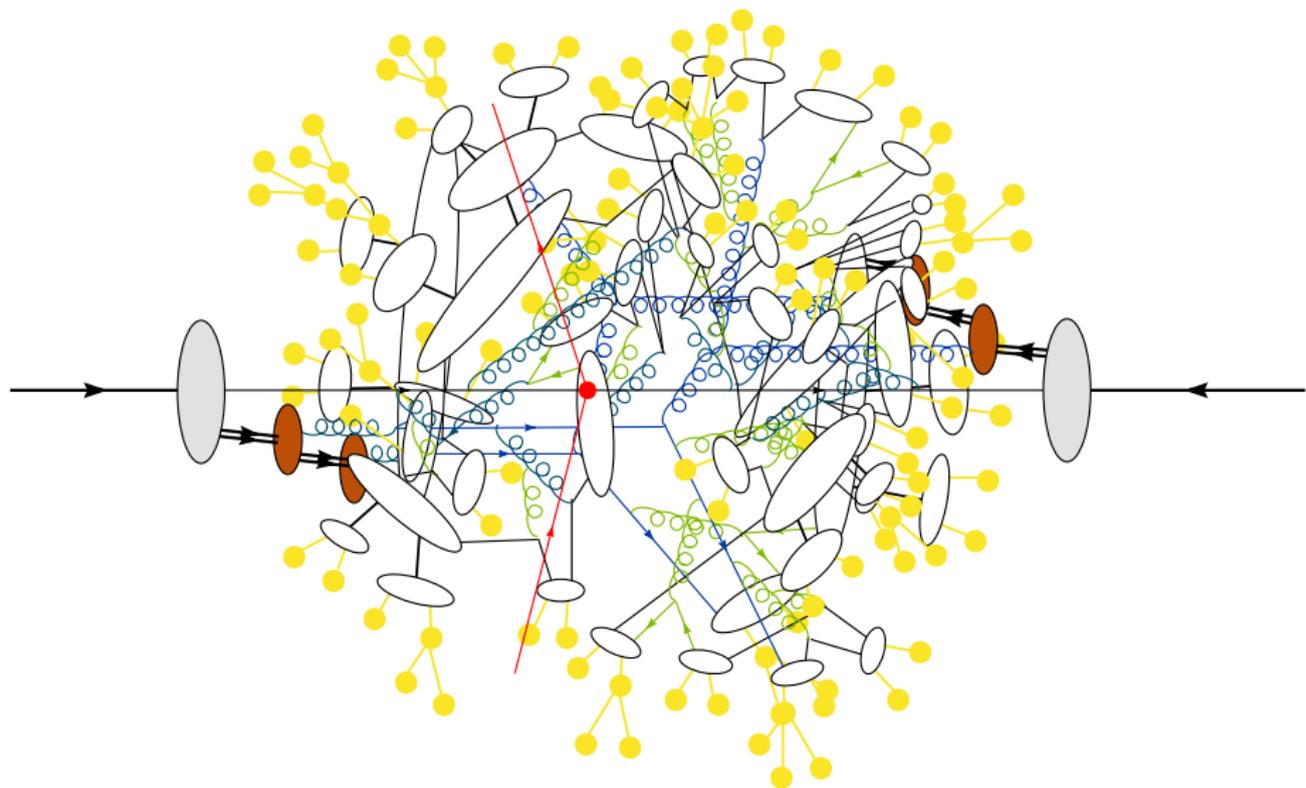
pp Event Generator



pp Event Generator



pp Event Generator



Divide and conquer

Partonic cross section from Feynman diagrams

$$d\sigma = d\sigma_{\text{hard}} dP(\text{partons} \rightarrow \text{hadrons})$$

$$\begin{aligned} dP(\text{partons} \rightarrow \text{hadrons}) = & dP(\text{resonance decays}) && [\Gamma > Q_0] \\ & \times dP(\text{parton shower}) && [\text{TeV} \rightarrow Q_0] \\ & \times dP(\text{hadronisation}) && [\sim Q_0] \\ & \times dP(\text{hadronic decays}) && [O(\text{MeV})] \end{aligned}$$

Underlying event from multiple partonic interactions

$$d\sigma \longleftarrow d\sigma(\text{QCD } 2 \rightarrow 2)$$

Plan for these lectures

- Monte Carlo Methods
- Hard Scattering
- Parton Showers

Monte Carlo Methods

Monte Carlo Methods

Introduction to the most important MC sampling (= integration) techniques.

- ① Hit and miss.
- ② Simple MC integration.
- ③ (Some) methods of variance reduction.
- ④ Adaptive MC, VEGAS.
- ⑤ Multichannel.
- ⑥ Mini event generator in particle physics.

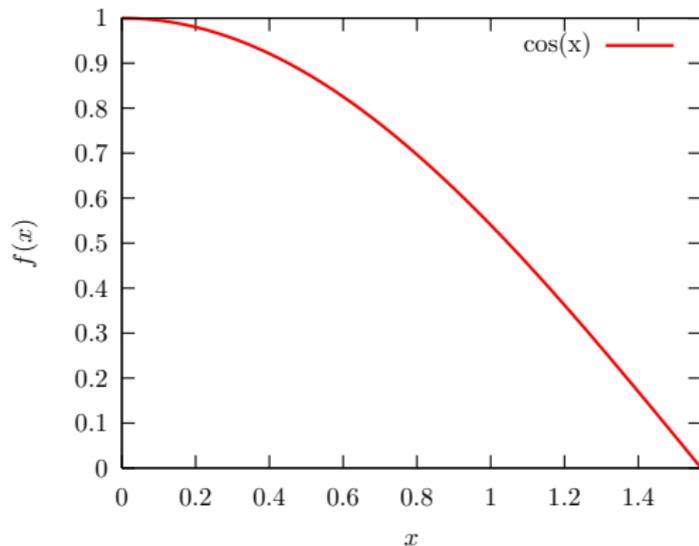
Probability

Probability density:

$$dP = f(x) dx$$

is probability to find value x .

Example: $f(x) = \cos(x)$.



Probability

Probability density:

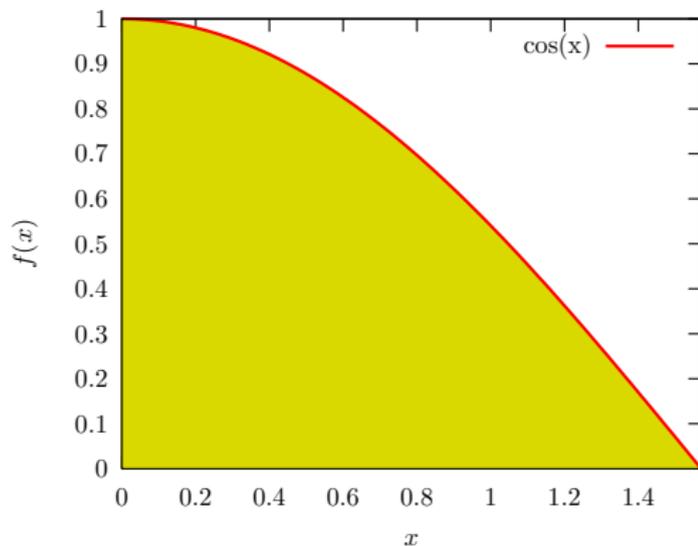
$$dP = f(x) dx$$

is probability to find value x .

$$F(x) = \int_{x_0}^x f(x) dx$$

is called *probability distribution*.

Example: $f(x) = \cos(x)$.



Probability

Probability density:

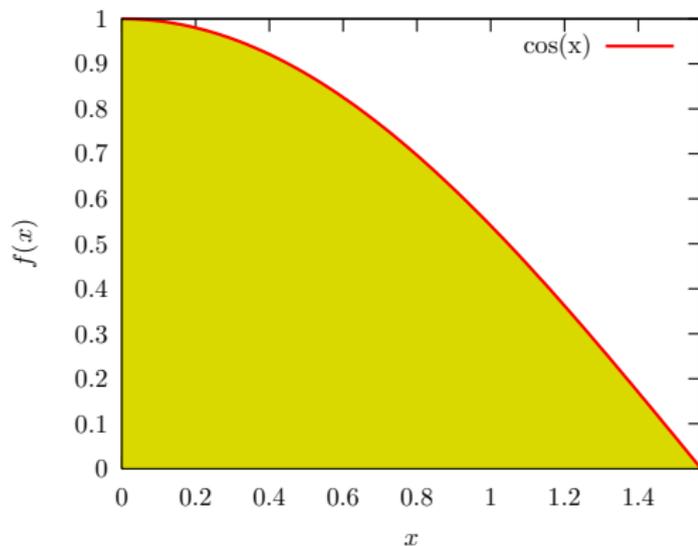
$$dP = f(x) dx$$

is probability to find value x .

$$F(x) = \int_{x_0}^x f(x) dx$$

is called *probability distribution*.

Example: $f(x) = \cos(x)$.



Probability \sim Area

Hit and Miss

Hit and miss method:

- throw N random points (x, y) into region.
- Count hits N_{hit} ,
i.e. whenever $y < f(x)$.

Then

$$I \approx V \frac{N_{\text{hit}}}{N}.$$

approaches 1 again in our example.

Hit and Miss

Hit and miss method:

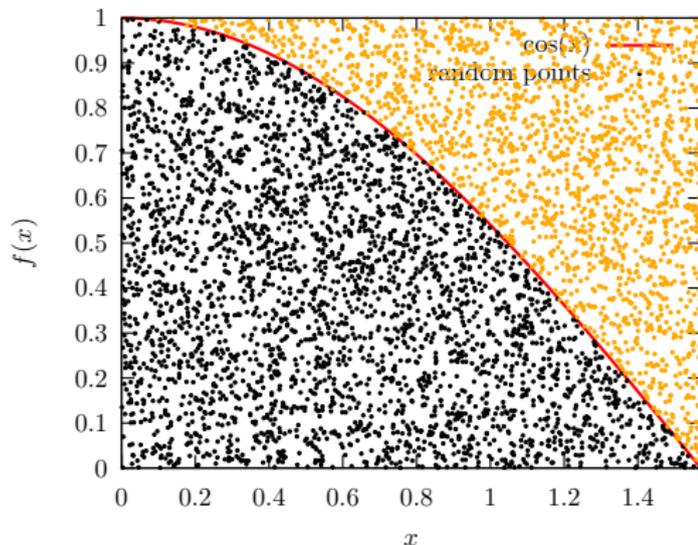
- throw N random points (x, y) into region.
- Count hits N_{hit} ,
i.e. whenever $y < f(x)$.

Then

$$I \approx V \frac{N_{\text{hit}}}{N}.$$

approaches 1 again in our
example.

Example: $f(x) = \cos(x)$.



Hit and Miss

Hit and miss method:

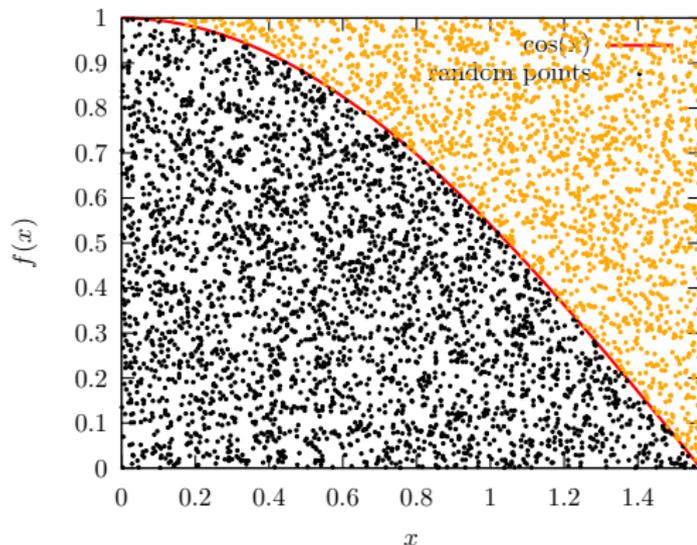
- throw N random points (x, y) into region.
- Count hits N_{hit} ,
i.e. whenever $y < f(x)$.

Then

$$I \approx V \frac{N_{\text{hit}}}{N}.$$

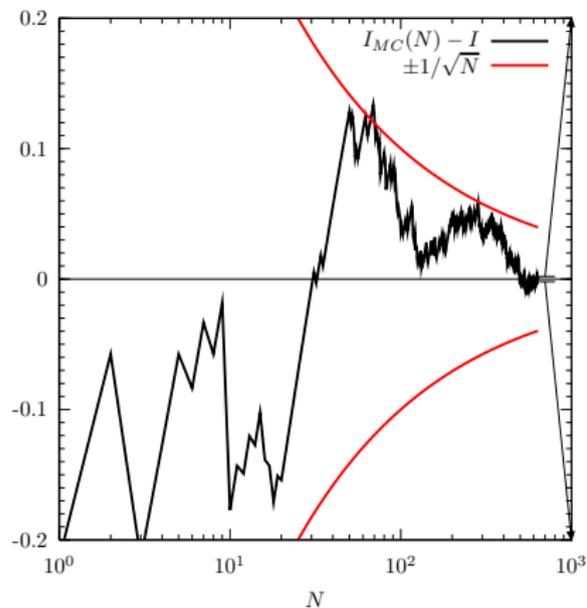
approaches 1 again in our
example.

Example: $f(x) = \cos(x)$.



Every **accepted** value of x can be considered an **event** in this picture. As $f(x)$ is the 'histogram' of x , it seems obvious that the x values are distributed as $f(x)$ from this picture.

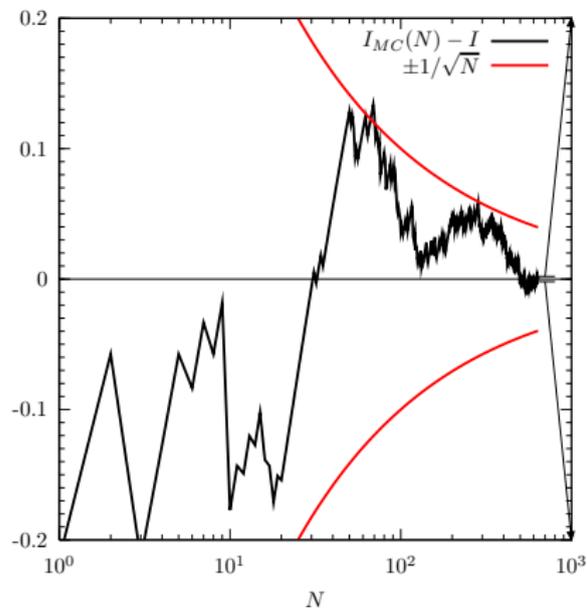
Hit and Miss



How well does it converge?

Error $1/\sqrt{N}$.

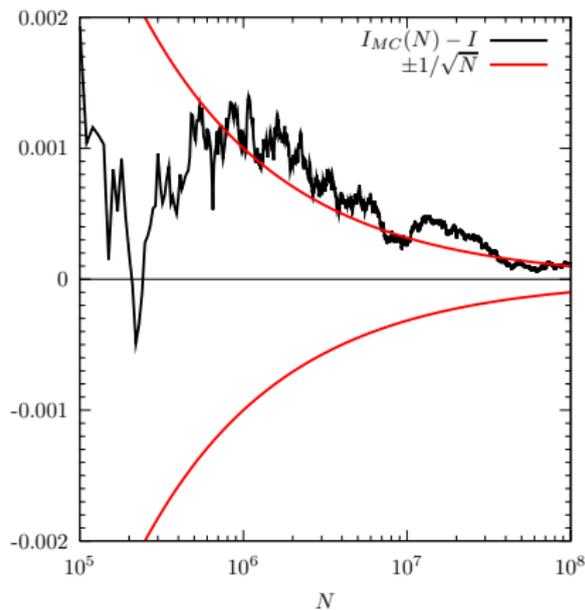
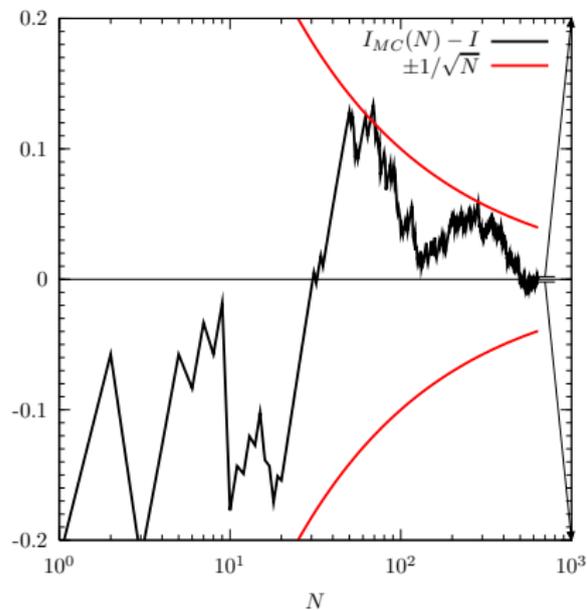
Hit and Miss



More points, zoom in...

Error $1/\sqrt{N}$.

Hit and Miss



Error $1/\sqrt{N}$.

Hit and Miss

This method is used in many event generators. However, it is not sufficient as such.

- Can handle any density $f(x)$, however wild and unknown it is.
- $f(x)$ should be bounded from above.
- Sampling will be very *inefficient* whenever $\text{Var}(f)$ is large.

Improvements go under the name **variance reduction** as they improve the error of the crude MC at the same time.

Simple MC integration

Mean value theorem of integration:

$$\begin{aligned} I &= \int_{x_0}^{x_1} f(x) dx \\ &= (x_1 - x_0) \langle f(x) \rangle \end{aligned}$$

(Riemann integral).

Simple MC integration

Mean value theorem of integration:

$$\begin{aligned} I &= \int_{x_0}^{x_1} f(x) dx \\ &= (x_1 - x_0) \langle f(x) \rangle \\ &\approx (x_1 - x_0) \frac{1}{N} \sum_{i=1}^N f(x_i) \end{aligned}$$

(Riemann integral).

Simple MC integration

Mean value theorem of integration:

$$\begin{aligned} I &= \int_{x_0}^{x_1} f(x) dx \\ &= (x_1 - x_0) \langle f(x) \rangle \\ &\approx (x_1 - x_0) \frac{1}{N} \sum_{i=1}^N f(x_i) \end{aligned}$$

(Riemann integral).

Sum doesn't depend on ordering

→ randomize x_i .

Simple MC integration

Mean value theorem of integration:

$$\begin{aligned} I &= \int_{x_0}^{x_1} f(x) dx \\ &= (x_1 - x_0) \langle f(x) \rangle \\ &\approx (x_1 - x_0) \frac{1}{N} \sum_{i=1}^N f(x_i) \end{aligned}$$

(Riemann integral).

Sum doesn't depend on ordering

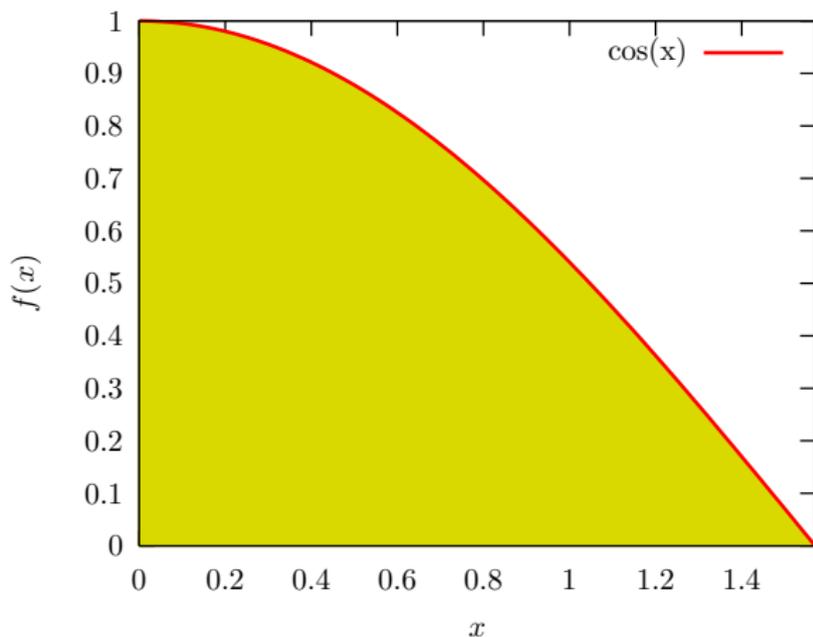
→ randomize x_i .

Yields a flat distribution of events x_i ,
but weighted with *weight* $f(x_i)$ (→ unweighting).

Simple MC integration

Pictorially:

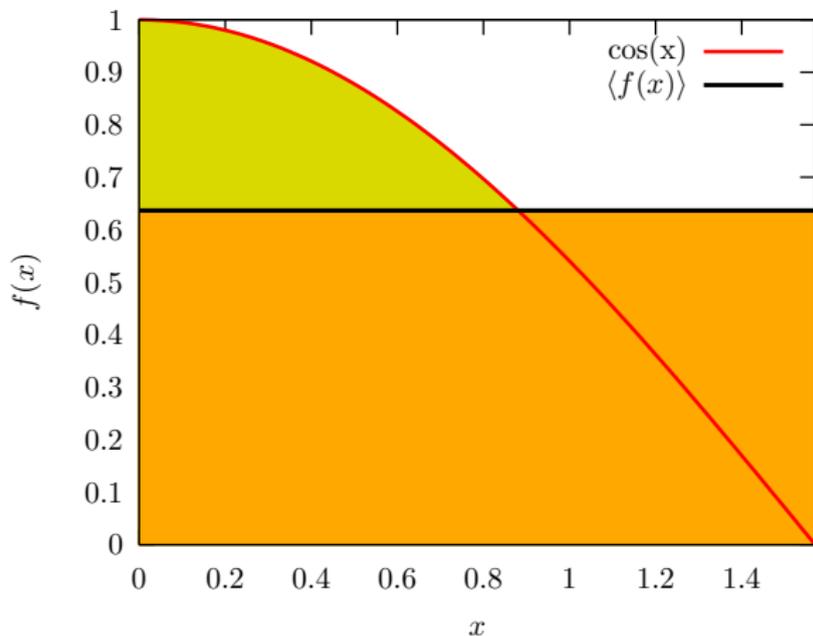
$$\begin{aligned} I &= \int_{x_0}^{x_1} f(x) dx \\ &= (x_1 - x_0) \langle f(x) \rangle \end{aligned}$$



Simple MC integration

Pictorially:

$$I = \int_{x_0}^{x_1} f(x) dx$$
$$= (x_1 - x_0) \langle f(x) \rangle$$

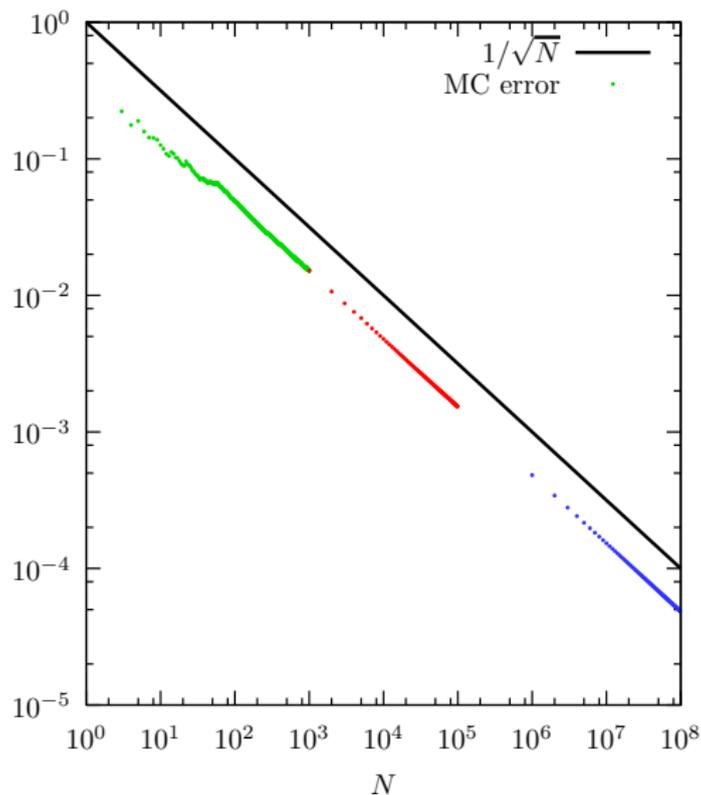


Simple MC integration

What's the error?

Again, looks like

$$\sigma \sim \frac{1}{\sqrt{N}}$$



Simple MC integration

What's the error?

We can calculate it (central limit theorem for the average):

In general: *Crude MC*

$$\begin{aligned} I &= \int f dV \\ &\approx V \langle f \rangle \pm V \sqrt{\frac{\langle f \rangle^2 - \langle f^2 \rangle}{N}} \\ &\approx V \langle f \rangle \pm V \frac{\sigma}{\sqrt{N}} \end{aligned}$$

Simple MC integration

What's the error?

We can calculate it (central limit theorem for the average):

Our example: $\cos(x)$, $0 \leq x \leq \pi/2$,
compute σ_{MC} from

$$\langle f \rangle = \frac{1}{N} \sum_{i=1}^N f(x_i)$$

$$\langle f^2 \rangle = \frac{1}{N} \sum_{i=1}^N f^2(x_i).$$

Simple MC integration

What's the error?

We can calculate it (central limit theorem for the average):

Compute σ directly ($V = \pi/2$):

$$\langle f \rangle = \int_0^{\pi/2} \cos(x) dx = 1$$

$$\langle f^2 \rangle = \int_0^{\pi/2} \cos^2(x) dx = \frac{\pi}{4}$$

then

$$\sigma = \sqrt{1^2 - \frac{\pi}{4}} \approx 0.4633.$$

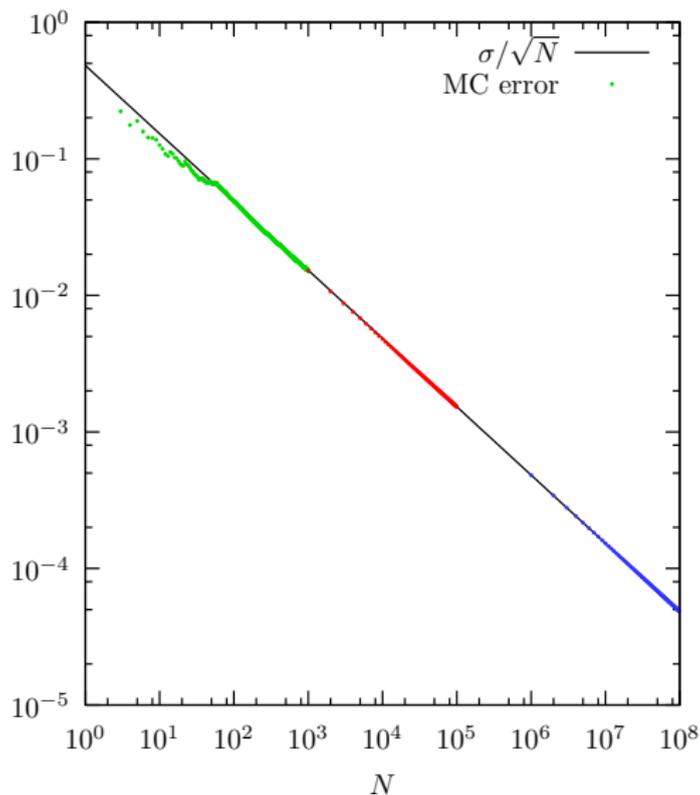
Simple MC integration

What's the error?

Now, compare

$$\sigma_{MC} = \frac{0.4633}{\sqrt{N}}$$

with error estimate
from MC.



Simple MC integration

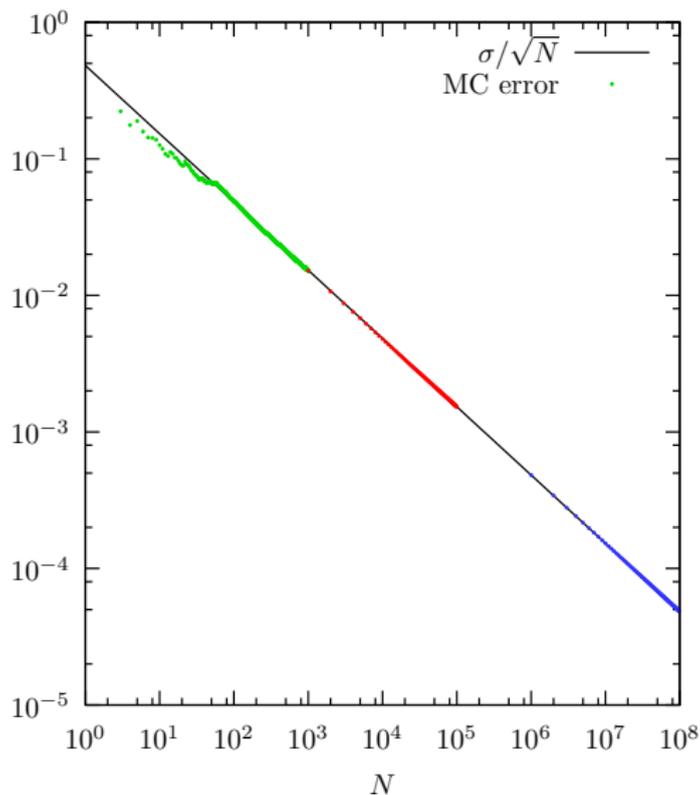
What's the error?

Now, compare

$$\sigma_{MC} = \frac{0.4633}{\sqrt{N}}$$

with error estimate
from MC.

Spot on.



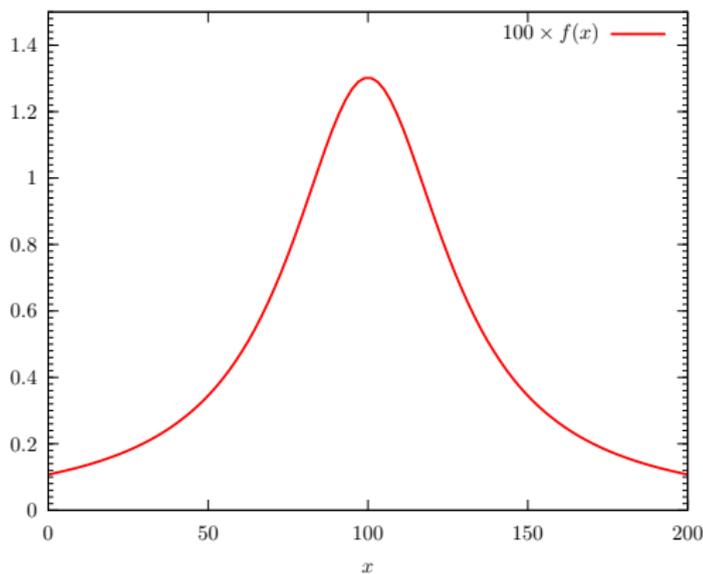
Inverting the Integral

Another basic MC method, based on the observation that

$$\textit{Probability} \sim \textit{Area}$$

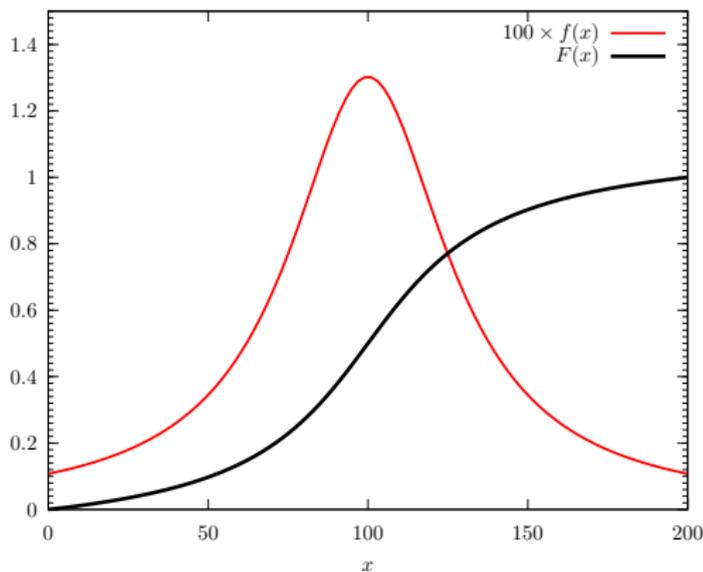
Inverting the Integral

- Probability density $f(x)$. Not necessarily normalized.



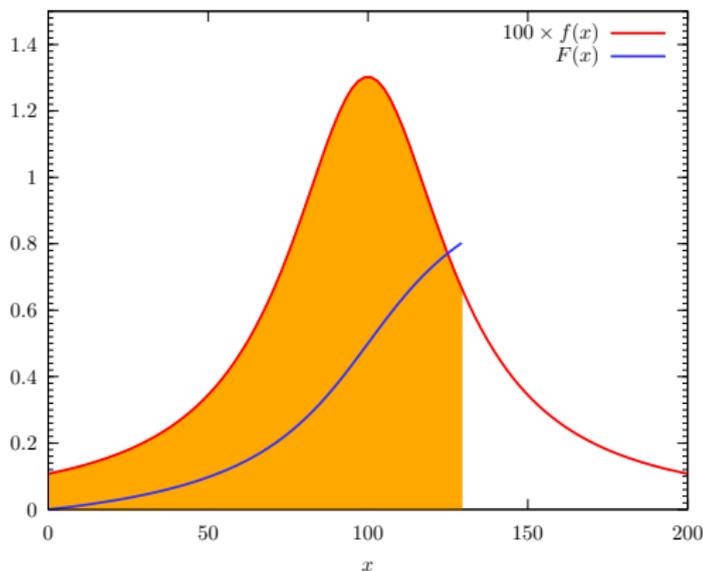
Inverting the Integral

- Probability density $f(x)$. Not necessarily normalized.
- Integral $F(x)$ known,



Inverting the Integral

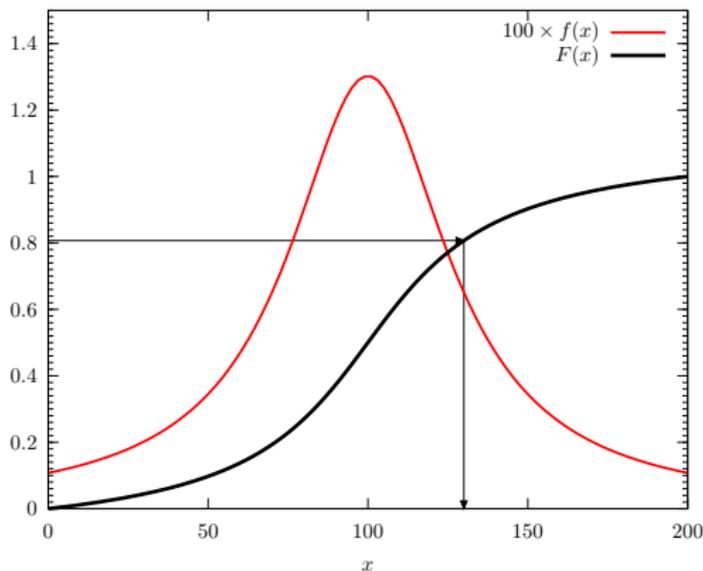
- Probability density $f(x)$. Not necessarily normalized.
- Integral $F(x)$ known,
- $P(x < x_s) = F(x_s)$.



Inverting the Integral

- Probability density $f(x)$. Not necessarily normalized.
- Integral $F(x)$ known,
- $P(x < x_s) = F(x_s)$.
- Probability = 'area', distributed evenly,

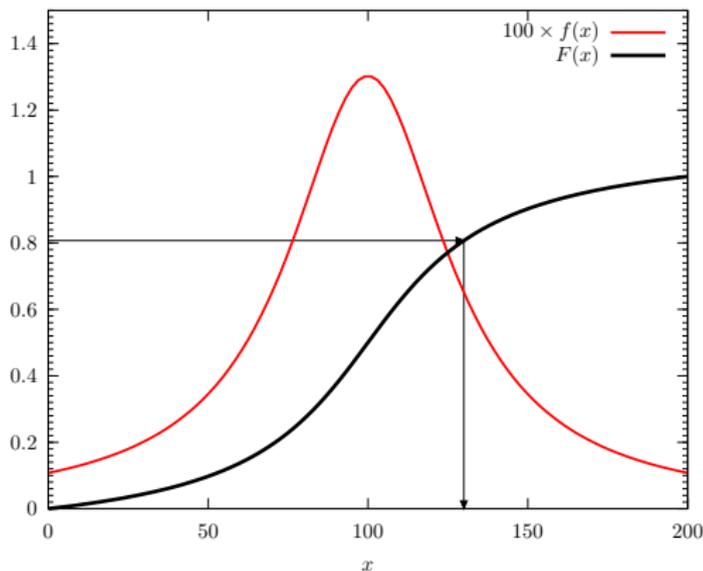
$$\int_{x_0}^x dP = r$$



Inverting the Integral

- Probability density $f(x)$. Not necessarily normalized.
- Integral $F(x)$ known,
- $P(x < x_s) = F(x_s)$.
- Probability = 'area', distributed evenly,

$$\int_{x_0}^x dP = r$$



Sample x according to $f(x)$ with

$$x = F^{-1} \left[F(x_0) + r(F(x_1) - F(x_0)) \right].$$

Inverting the Integral

Another basic MC method, based on the observation that

$$\textit{Probability} \sim \textit{Area}$$

Sample x according to $f(x)$ with

$$x = F^{-1} \left[F(x_0) + r(F(x_1) - F(x_0)) \right].$$

Optimal method, but we need to know

- The integral $F(x) = \int f(x) dx$,
- It's inverse $F^{-1}(y)$.

That's rarely the case for real problems.

But very powerful in combination with other techniques.

Importance sampling

Error on Crude MC $\sigma_{MC} = \sigma/\sqrt{N}$.

\implies Reduce error by reducing variance of integrand.

Importance sampling

Error on Crude MC $\sigma_{MC} = \sigma/\sqrt{N}$.

\implies Reduce error by reducing variance of integrand.

Idea: *Divide out the singular structure.*

$$I = \int f dV = \int \frac{f}{p} p dV \approx \left\langle \frac{f}{p} \right\rangle \pm \sqrt{\frac{\langle f^2/p^2 \rangle - \langle f/p \rangle^2}{N}}.$$

where we have chosen $\int p dV = 1$ for convenience.

Note: need to sample flat in $p dV$, so we better know $\int p dV$ and it's inverse.

Importance sampling

Consider error term:

$$\begin{aligned} E &= \left\langle \frac{f^2}{p^2} \right\rangle - \left\langle \frac{f}{p} \right\rangle^2 = \int \frac{f^2}{p^2} p dV - \left[\int \frac{f}{p} p dV \right]^2 \\ &= \int \frac{f^2}{p} dV - \left[\int f dV \right]^2. \end{aligned}$$

Importance sampling

Consider error term:

$$E = \int \frac{f^2}{p} dV - \left[\int f dV \right]^2 .$$

Best choice of p ? Minimises $E \rightarrow$ functional variation of error term with (normalized) p :

$$\begin{aligned} 0 = \delta E &= \delta \left(\int \frac{f^2}{p} dV - \left[\int f dV \right]^2 + \lambda \int p dV \right) \\ &= \int \left(-\frac{f^2}{p^2} + \lambda \right) dV \delta p , \end{aligned}$$

Importance sampling

Consider error term:

$$E = \int \frac{f^2}{p} dV - \left[\int f dV \right]^2 .$$

Best choice of p ? Minimises $E \rightarrow$ functional variation of error term with (normalized) p :

$$0 = \delta E = \int \left(-\frac{f^2}{p^2} + \lambda \right) dV \delta p ,$$

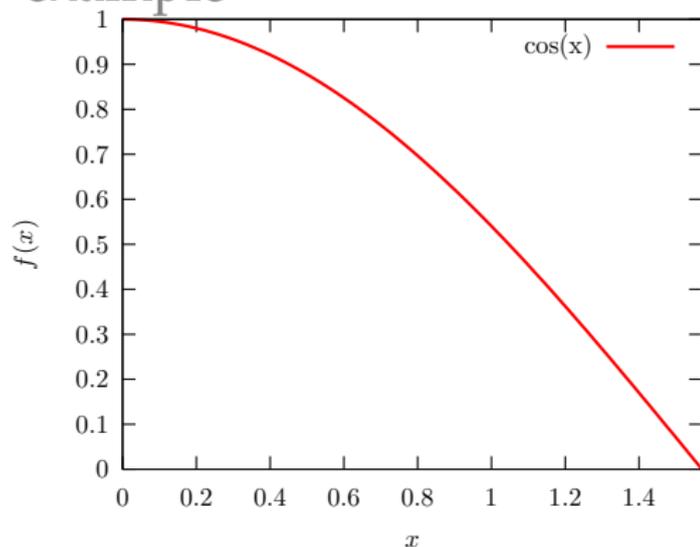
hence

$$p = \frac{|f|}{\sqrt{\lambda}} = \frac{|f|}{\int |f| dV} .$$

Choose p as close to f as possible.

Importance sampling — example

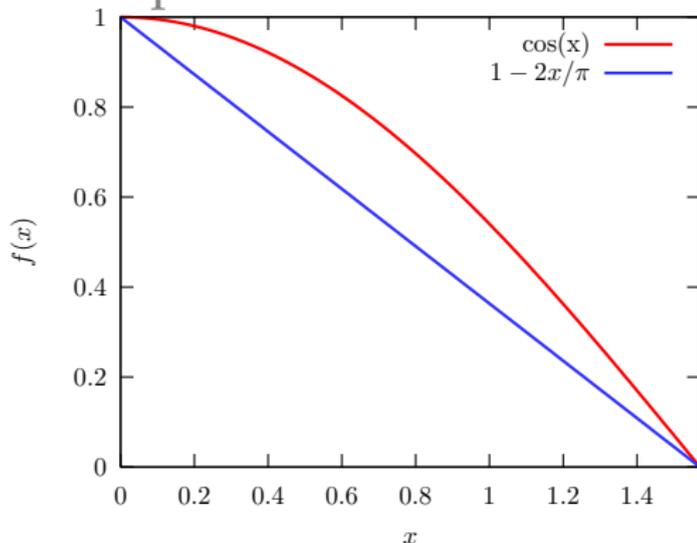
Improving $\cos(x)$
sampling,



Importance sampling — example

Improving $\cos(x)$
sampling,

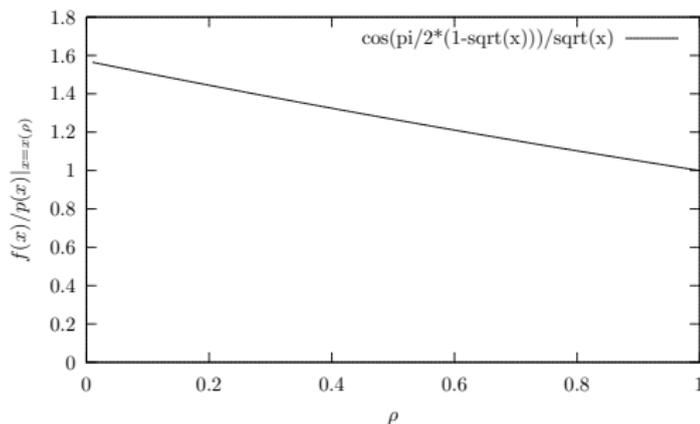
$$\begin{aligned} I &= \int_0^{\pi/2} \cos(x) dx \\ &= \int_0^{\pi/2} \frac{\cos(x)}{1 - \frac{2}{\pi}x} \left(1 - \frac{2}{\pi}x\right) dx \\ &= \int_0^1 \frac{\cos(x)}{1 - \frac{2}{\pi}x} \Bigg|_{x=x(\rho)} d\rho . \end{aligned}$$



Importance sampling — example

Improving $\cos(x)$
sampling,

$$\begin{aligned} I &= \int_0^{\pi/2} \cos(x) dx \\ &= \int_0^{\pi/2} \frac{\cos(x)}{1 - \frac{2}{\pi}x} \left(1 - \frac{2}{\pi}x\right) dx \\ &= \int_0^1 \frac{\cos(x)}{1 - \frac{2}{\pi}x} \Bigg|_{x=x(\rho)} d\rho. \end{aligned}$$



Sample x with *inverting the integral* technique (flat random number ρ),

$$x = \frac{\pi}{2} \left(1 - \sqrt{1 - \rho}\right) \hat{=} \frac{\pi}{2} (1 - \sqrt{\rho}) \quad \left(I = \int_0^1 \frac{\cos\left(\frac{\pi}{2} (1 - \sqrt{\rho})\right)}{\sqrt{\rho}} d\rho. \right)$$

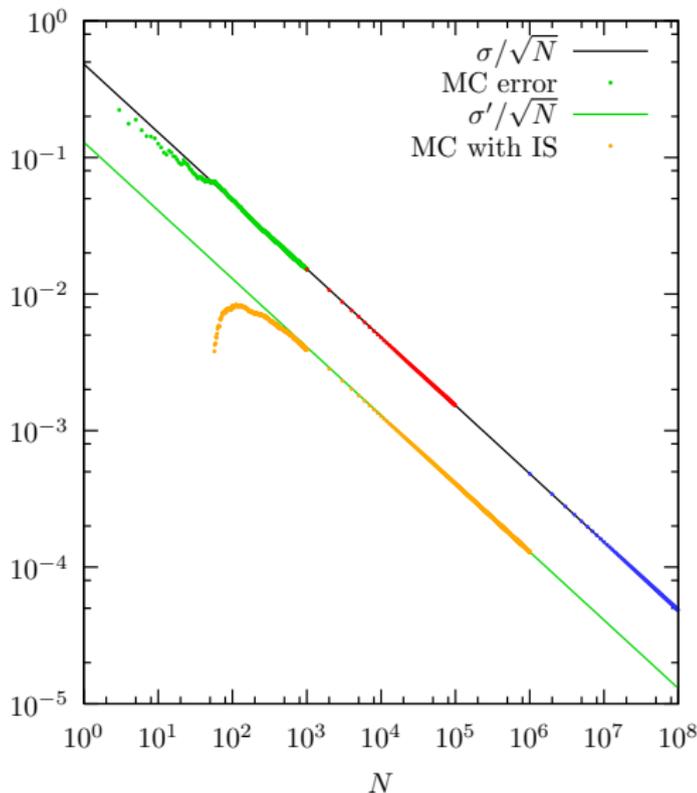
Importance sampling — example

Improving $\cos(x)$
sampling,

much better
convergence,

about 80% “accepted
events”.

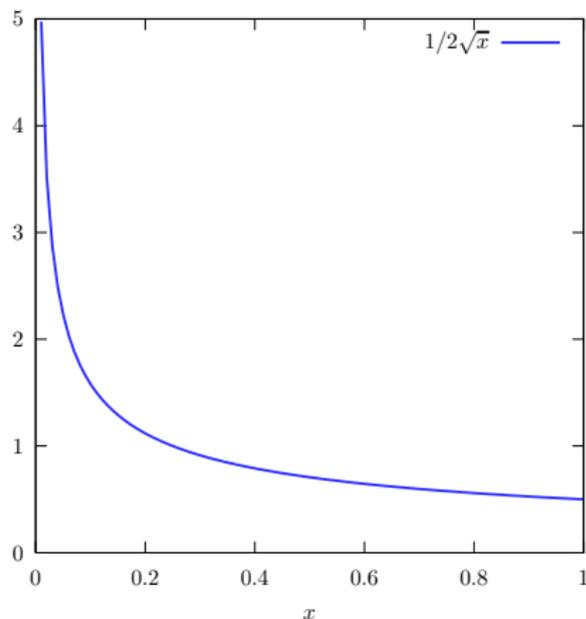
Reduced variance
($\sigma' = 0.027$)
⇒ better efficiency.



Importance sampling — better example

More interesting for **divergent integrands**, eg

$$\frac{1}{2\sqrt{x}},$$



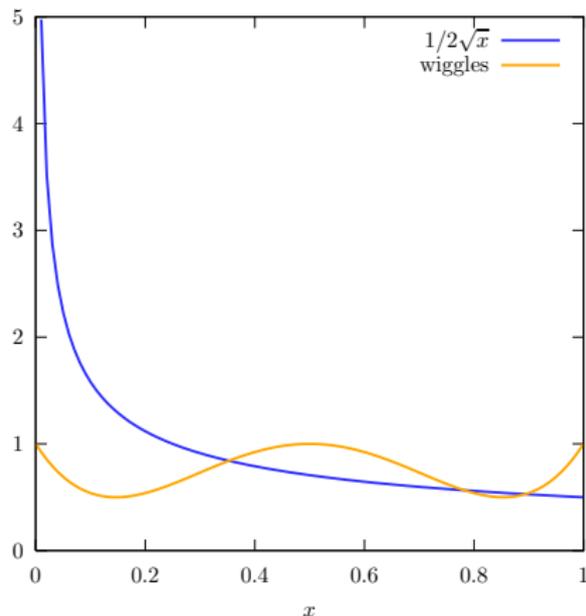
Importance sampling — better example

More interesting for **divergent integrands**, eg

$$\frac{1}{2\sqrt{x}},$$

with some wiggles,

$$p(x) = 1 - 8x + 40x^2 - 64x^3 + 32x^4.$$



Importance sampling — better example

More interesting for **divergent integrands**, eg

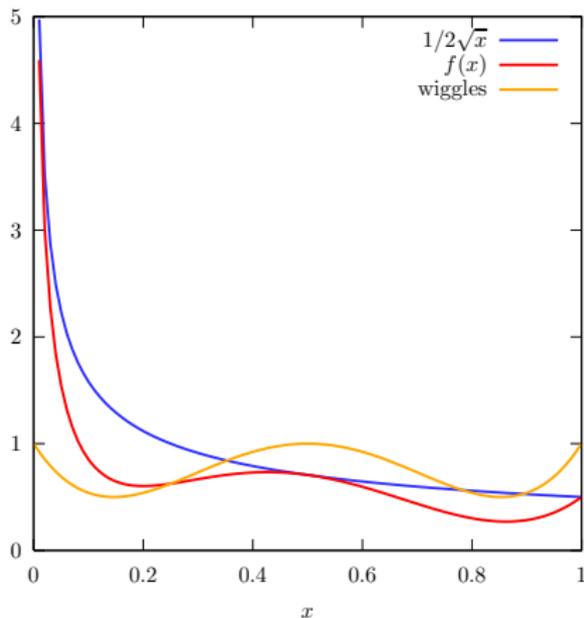
$$\frac{1}{2\sqrt{x}},$$

with some wiggles,

$$p(x) = 1 - 8x + 40x^2 - 64x^3 + 32x^4.$$

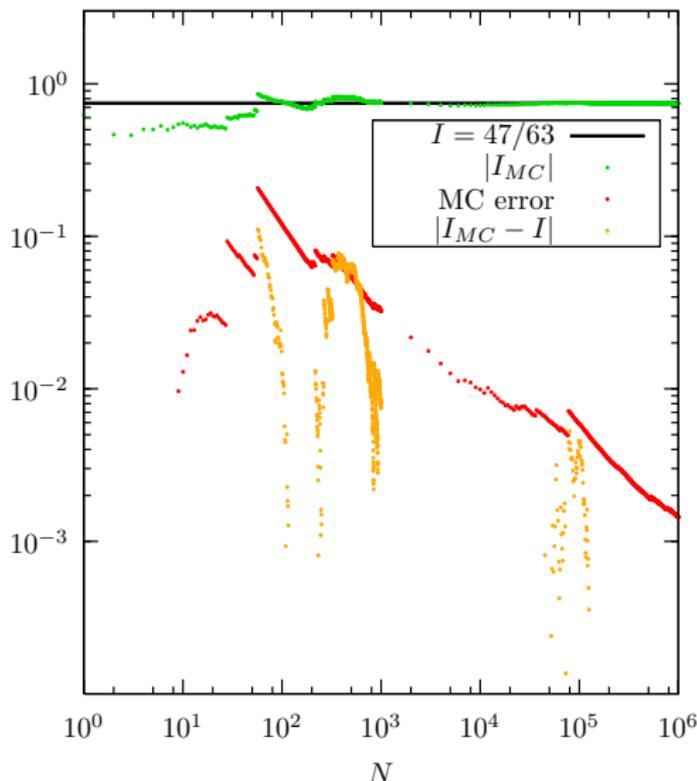
i.e. we want to integrate

$$f(x) = \frac{p(x)}{2\sqrt{x}}.$$



Importance sampling — better example

- Crude MC gives result in reasonable 'time'.
- Error a bit unstable.
- Event generation with maximum weight $w_{\max} = 20$. (that's arbitrary.)
- hit/miss/events with $(w > w_{\max}) = 36566/963434/617$ with 1M generated events.

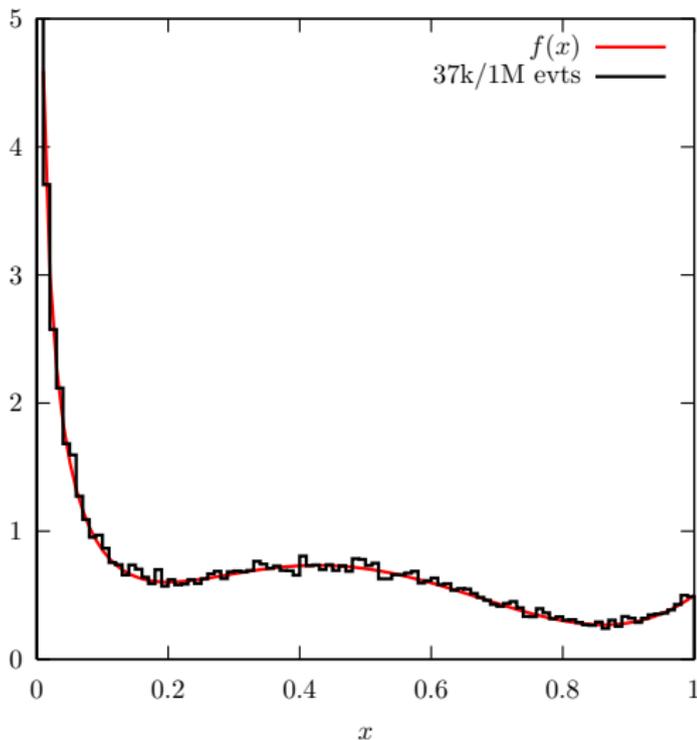


Importance sampling — better example

Want events:

use hit+mass variant
here:

- Choose new random number r
- $w = f(x)$ in this case.
- if $r < w/w_{\max}$ then “hit”.
- MC efficiency = hit/ N .

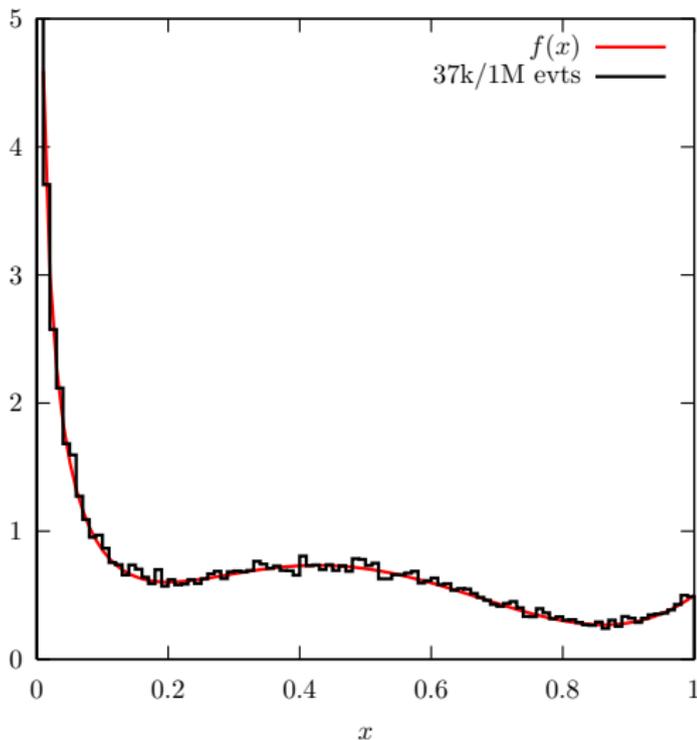


Importance sampling — better example

Want events:

use hit+mass variant
here:

- Choose new random number r
- $w = f(x)$ in this case.
- if $r < w/w_{\max}$ then “hit”.
- MC efficiency = hit/ N .
- Efficiency for MC events only 3.7%.
- Note the wiggly histogram.



Importance sampling — better example

Now importance sampling, i.e. divide out $1/2\sqrt{x}$.

$$\begin{aligned}\int_0^1 \frac{p(x)}{2\sqrt{x}} dx &= \int_0^1 \left(\frac{p(x)}{2\sqrt{x}} / \frac{1}{2\sqrt{x}} \right) \frac{dx}{2\sqrt{x}} \\ &= \int_0^1 p(x) d\sqrt{x} \\ &= \int_0^1 p(x(\rho)) d\rho \\ &= \int_0^1 1 - 8\rho^2 + 40\rho^4 - 64\rho^6 + 32\rho^8 d\rho\end{aligned}$$

so,

$$\rho = \sqrt{x}, \quad d\rho = \frac{dx}{2\sqrt{x}}$$

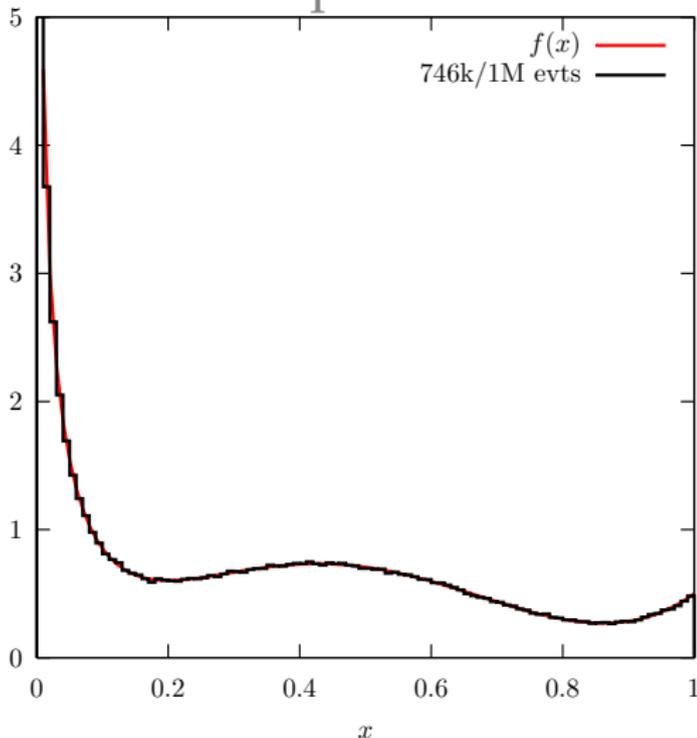
x sampled with *inverting the integral* from flat random numbers ρ , $x = \rho^2$.

Importance sampling — better example

$$\int_0^1 \frac{p(x)}{2\sqrt{x}} dx = \int_0^1 p(x(\rho)) d\rho$$

with

$$\rho = \sqrt{x}, \quad d\rho = \frac{dx}{2\sqrt{x}}$$



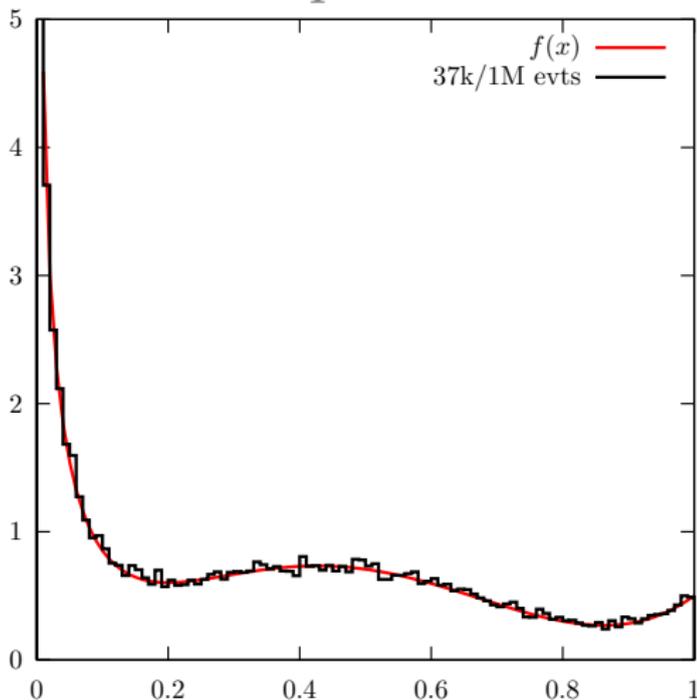
Events generated with $w_{\max} = 1$, as $p(x) \leq 1$, no guesswork needed here! Now, we get **74.6%** MC efficiency.

Importance sampling — better example

$$\int_0^1 \frac{p(x)}{2\sqrt{x}} dx = \int_0^1 p(x(\rho)) d\rho$$

with

$$\rho = \sqrt{x}, \quad d\rho = \frac{dx}{2\sqrt{x}}$$

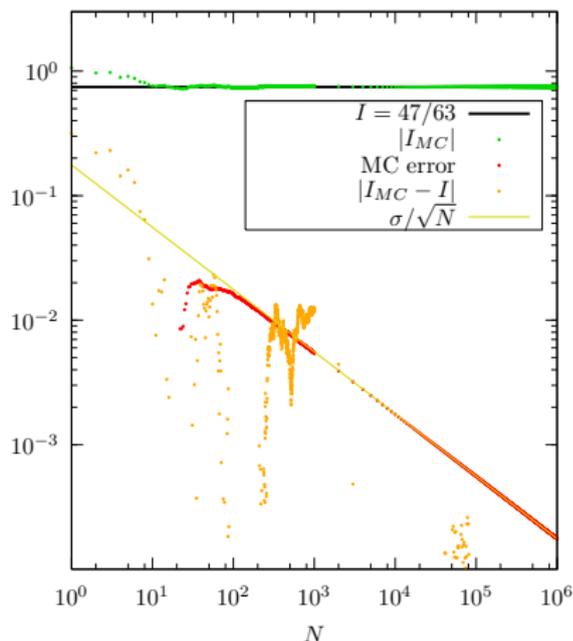
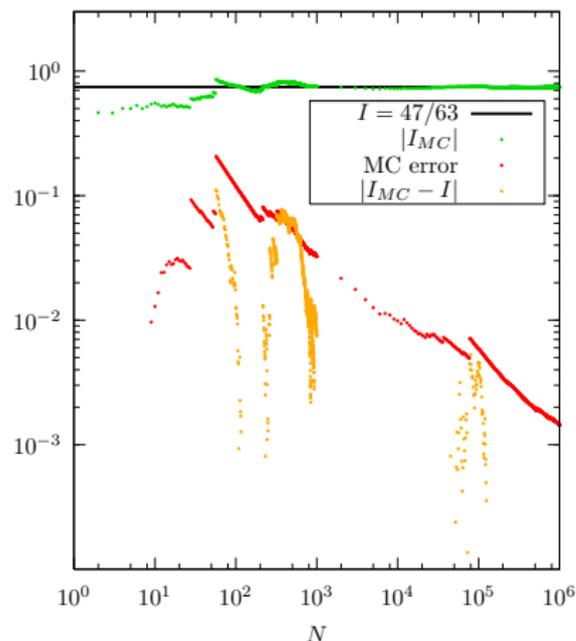


Events generated with $w_{\max} = 1$, as $p(x) \leq 1$, no guesswork needed here! Now, we get 74.6% MC efficiency.

... as opposed to 3.7%.

Importance sampling — better example

Crude MC vs Importance sampling.



100× more events needed to reach same accuracy.

Importance sampling — another useful example

Breit–Wigner peaks appear in many realistic MEs for cross sections and decays.

$$I = \int_{s_0}^{s_1} \frac{ds}{(s - m^2)^2 + m^2\Gamma^2}$$

Importance sampling — another useful example

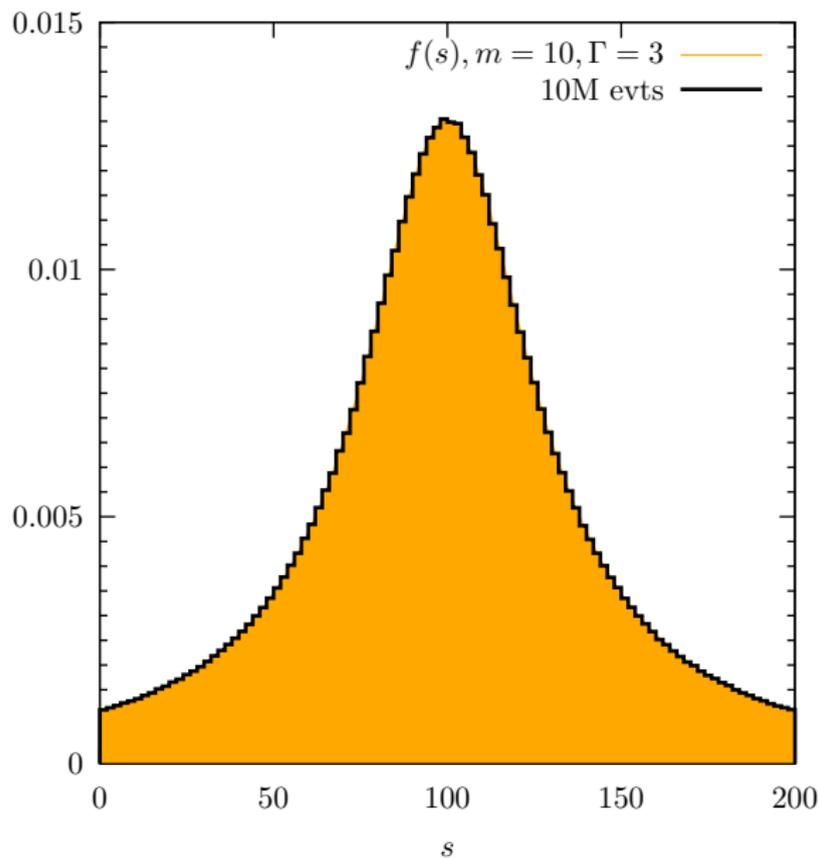
Breit–Wigner peaks appear in many realistic MEs for cross sections and decays.

$$\begin{aligned} I &= \int_{s_0}^{s_1} \frac{ds}{(s - m^2)^2 + m^2\Gamma^2} = \frac{1}{m\Gamma} \int_{y_0}^{y_1} \frac{dy}{y^2 + 1} \quad \left(y = \frac{s - m^2}{m\Gamma}\right) \\ &= \frac{1}{m\Gamma} \arctan \frac{s - m^2}{m\Gamma} \Big|_{s_0}^{s_1} \end{aligned}$$

Inverting the integral gives (“tan mapping”).

$$\begin{aligned} f(s) &= \frac{m\Gamma}{(s - m^2)^2 + m^2\Gamma^2} , \\ F(s) &= \arctan \frac{s - m^2}{m\Gamma} = \rho , \\ F^{-1}(\rho) &= m^2 + m\Gamma \tan \rho . \end{aligned}$$

Importance sampling — another useful example



VEGAS

- Classic algorithm.
- Automatic importance sampling.
- Adopt grid size.
- Often used for multidimensional integration.
- Very robust.

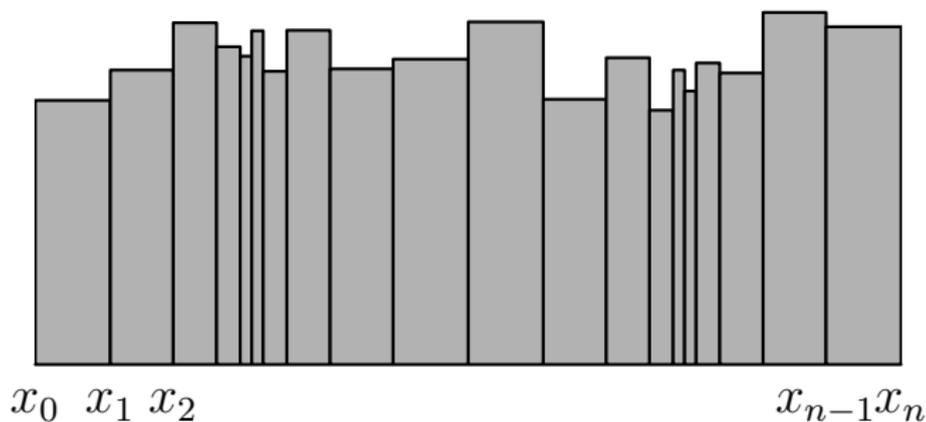
VEGAS

- start with equidistant grid x_0, x_1, \dots, x_N .
- Sample a number of points $(x_{s,i}, f(x_{s,i}))$, compute first estimate of integral as $\langle f \rangle$.
- Resize grid:
choose x'_i such that contribution from partial areas inside $x_i < x < x_{i+1}$ to integral is $\langle f \rangle / N$.
- Remember, optimal $p(x) \sim |f(x)|$.
- Sample again with same number of points into every bin $x_i < x < x_{i+1}$. Results in step weight function with steps

$$p_i = \frac{1}{N(x_i - x_{i-1})}, \quad x_i < x < x_{i+1} .$$

- \Rightarrow Sample often where density is high.

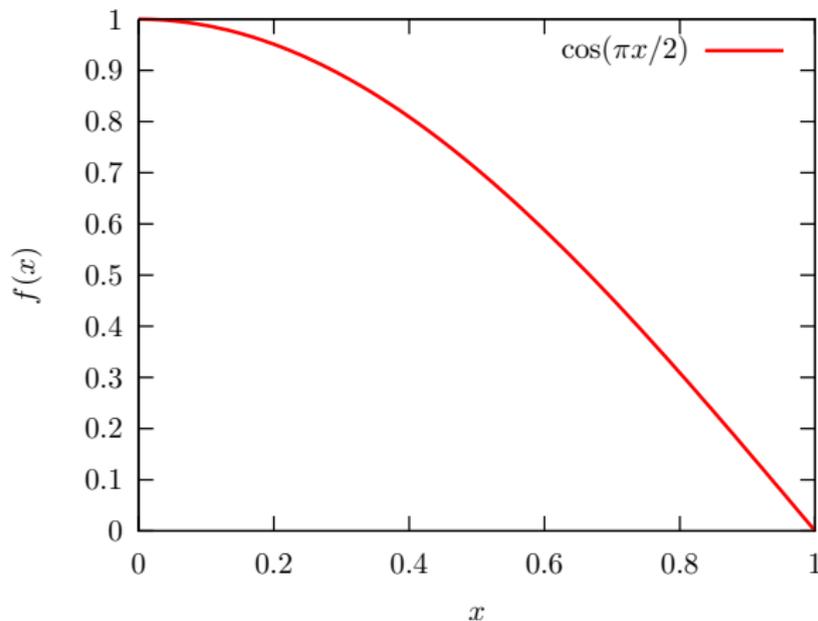
Rebinning:



[from T. Ohl, VAMP]

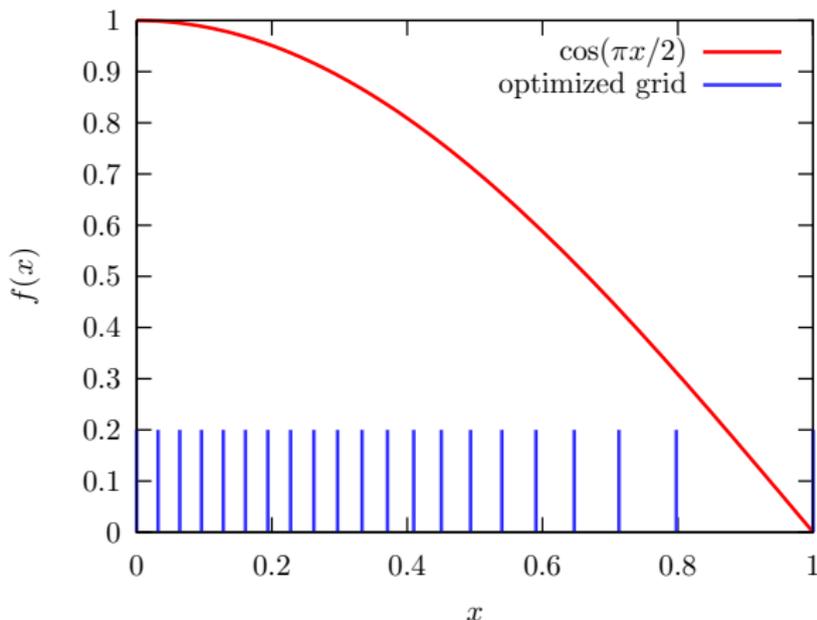
VEGAS

Example: $\cos(\frac{\pi x}{2})$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.



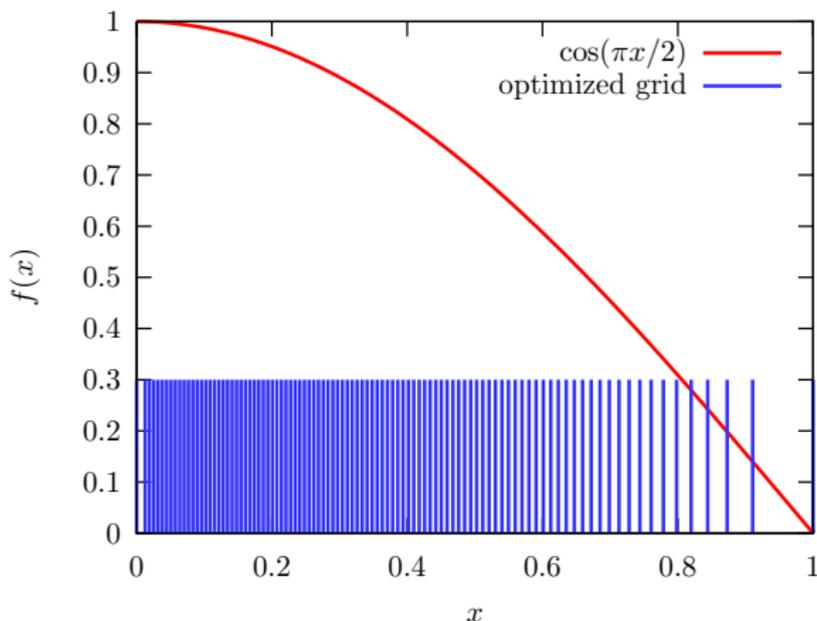
VEGAS

Example: $\cos(\frac{\pi x}{2})$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.



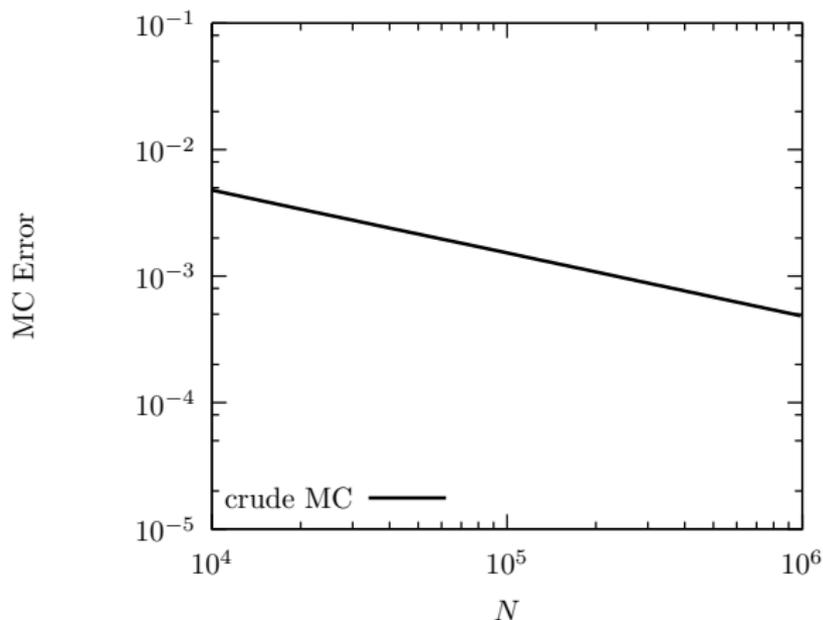
VEGAS

Example: $\cos(\frac{\pi x}{2})$
 $N_{\text{grid}} = 20,100$
Convergence
improved.



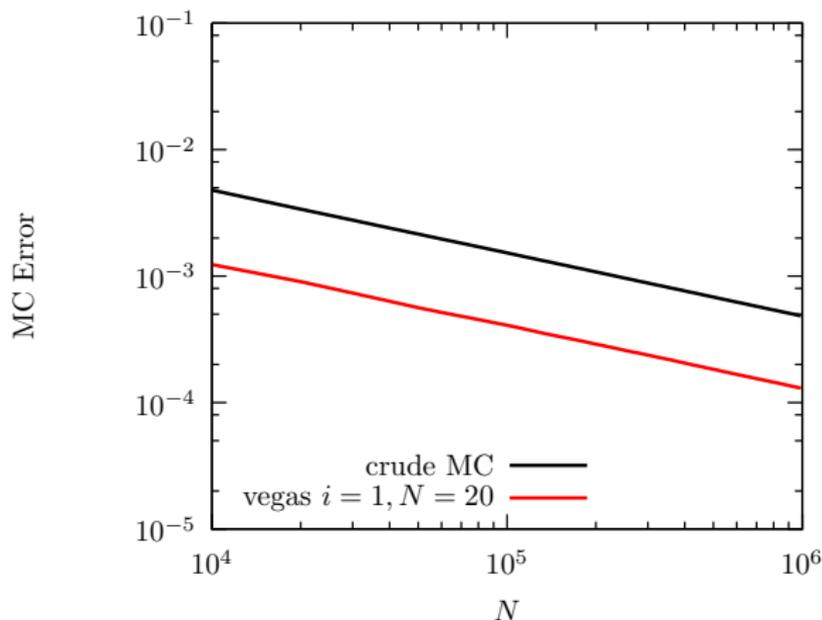
VEGAS

Example: $\cos\left(\frac{\pi x}{2}\right)$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.



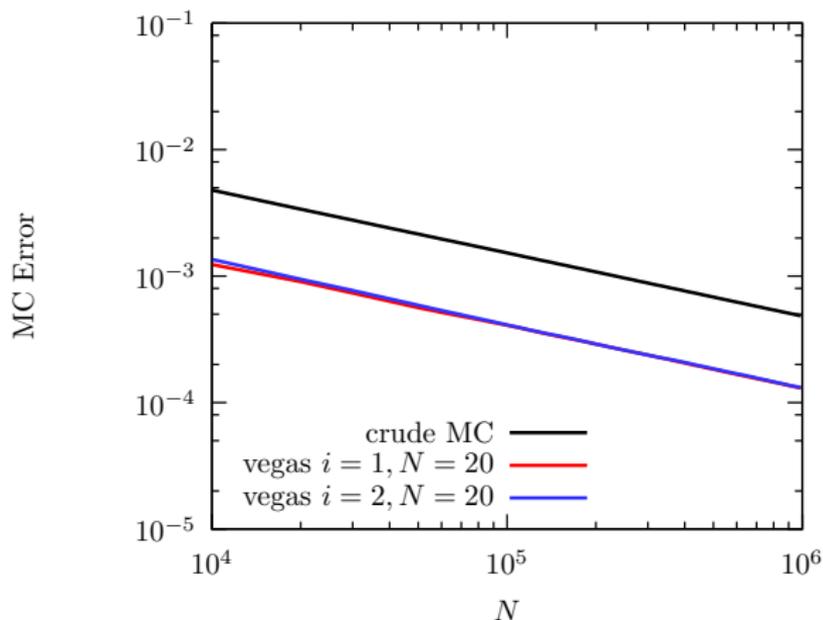
VEGAS

Example: $\cos\left(\frac{\pi x}{2}\right)$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.



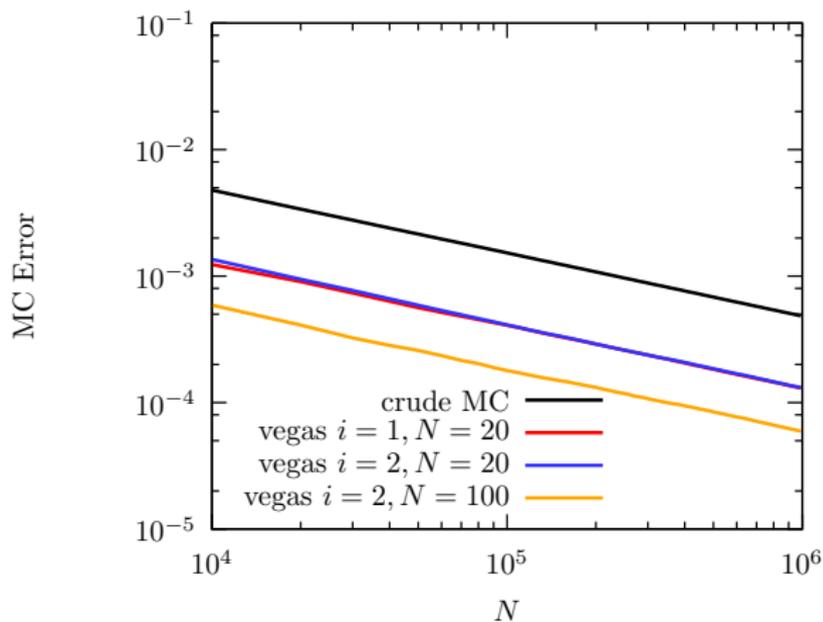
VEGAS

Example: $\cos\left(\frac{\pi x}{2}\right)$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.



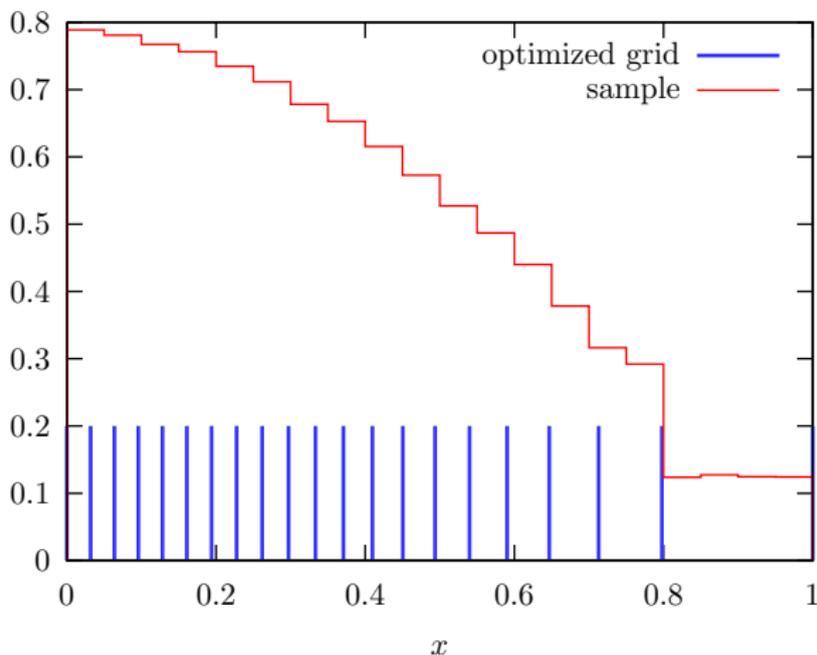
VEGAS

Example: $\cos\left(\frac{\pi x}{2}\right)$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.



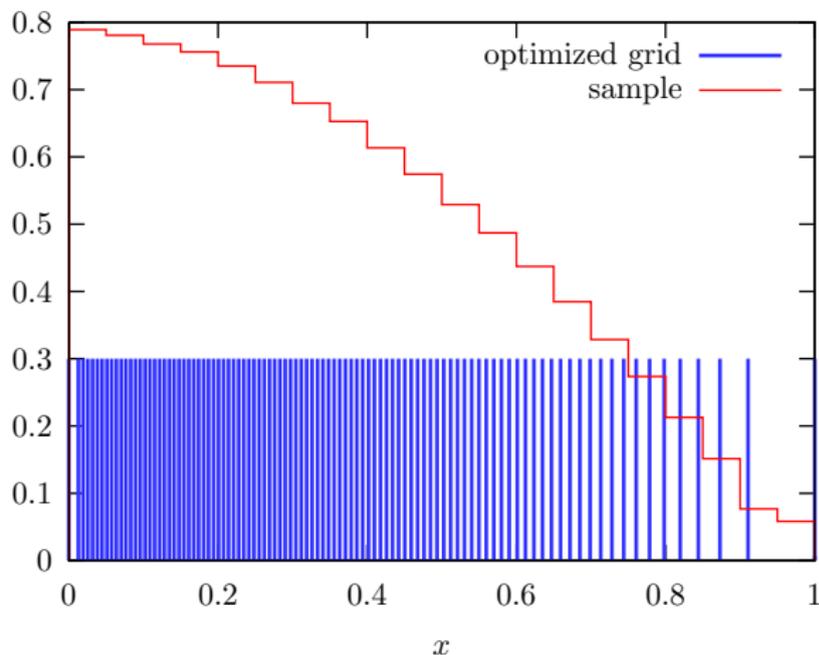
VEGAS

Example: $\cos(\frac{\pi x}{2})$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.



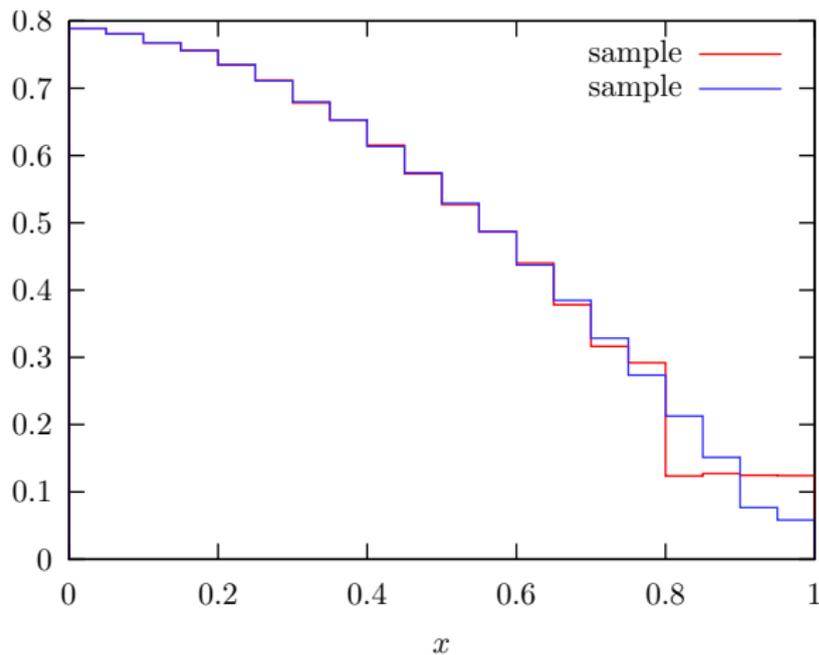
VEGAS

Example: $\cos(\frac{\pi x}{2})$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.

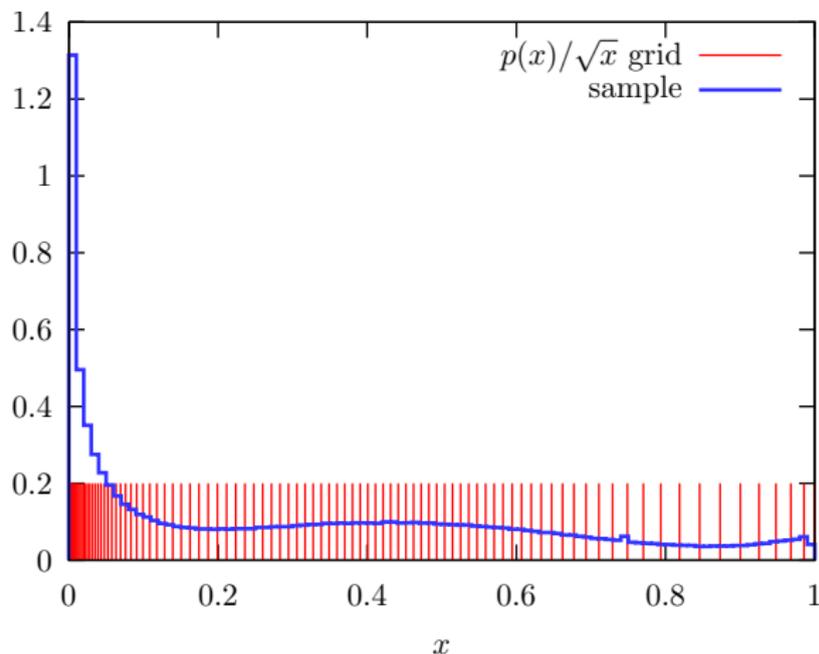


VEGAS

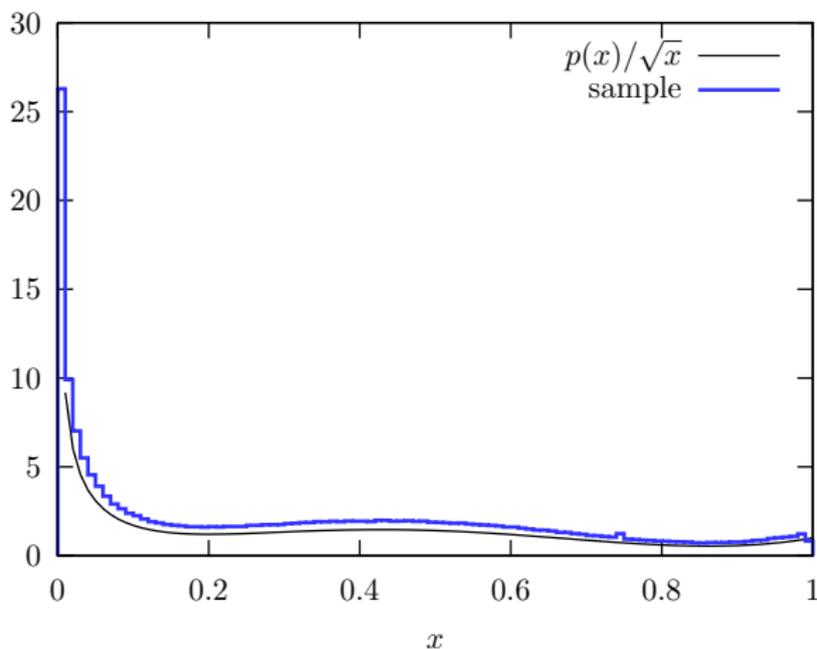
Example: $\cos\left(\frac{\pi x}{2}\right)$
 $N_{\text{grid}} = 20, 100$
Convergence
improved.



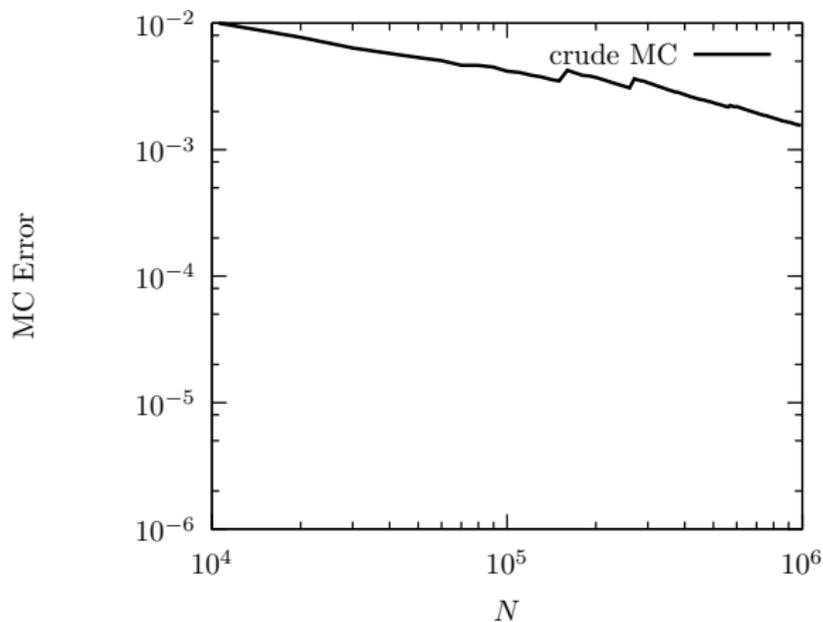
Second example:
 $p(x)/\sqrt{x}$
 (divergence with
 wiggles)



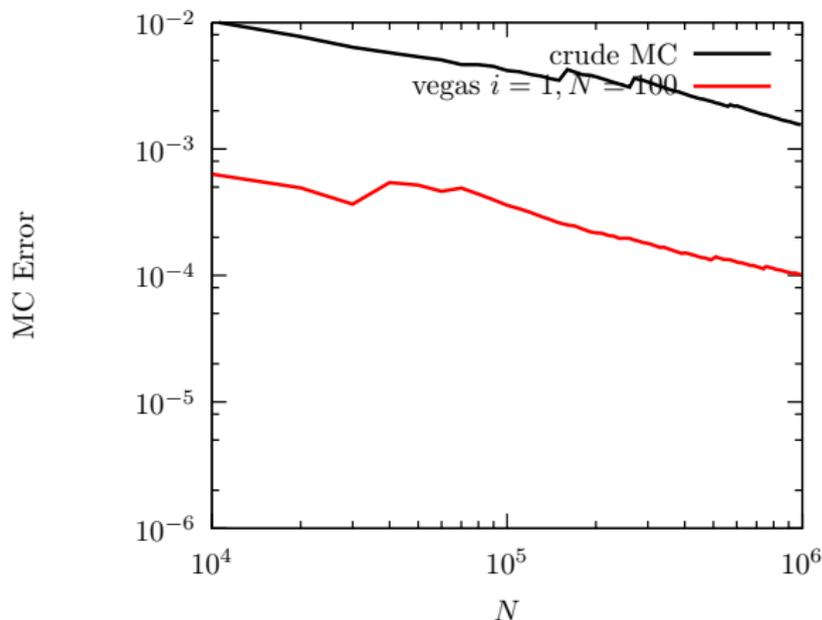
Second example:
 $p(x)/\sqrt{x}$
(divergence with
wiggles)



Second example:
 $p(x)/\sqrt{x}$
(divergence with
wiggles)

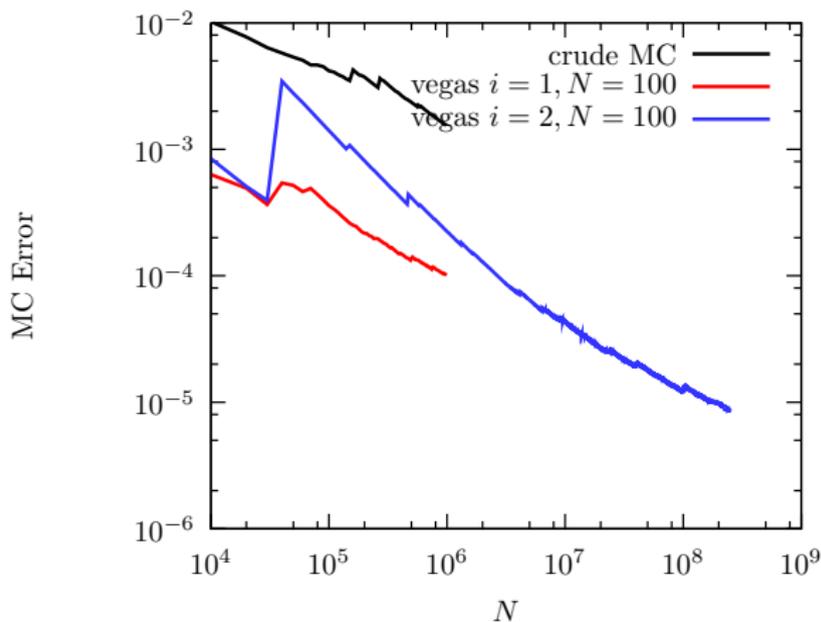


Second example:
 $p(x)/\sqrt{x}$
(divergence with
wiggles)



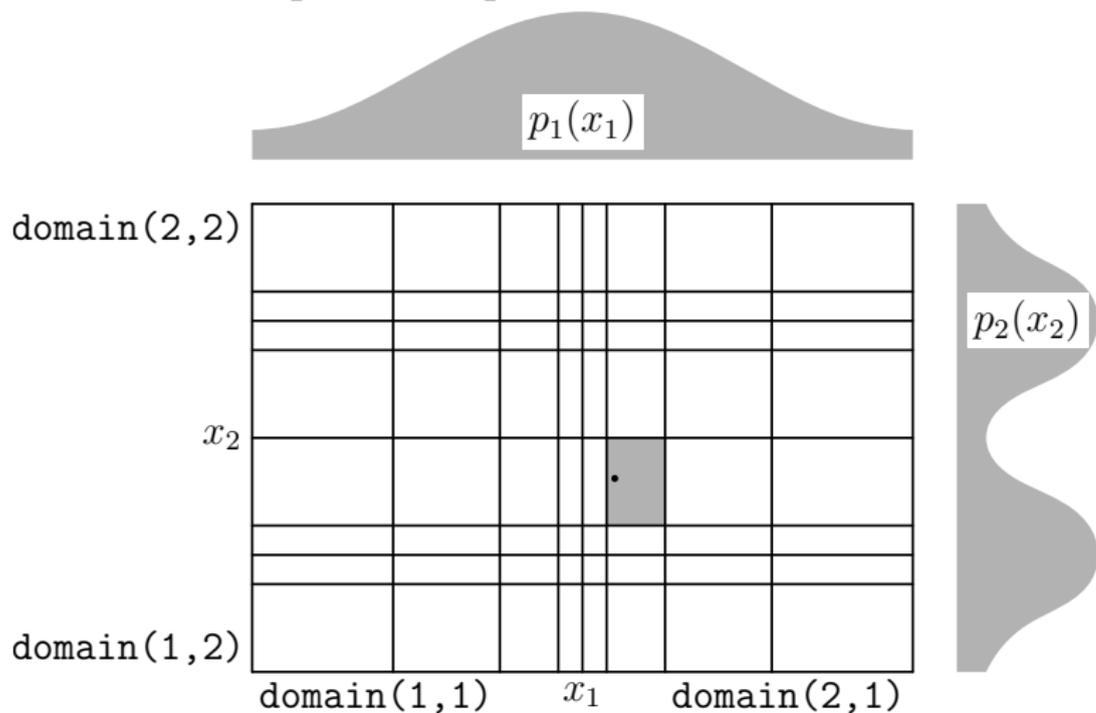
Acc 10^{-4} after $N = 10^6$ comparable with 'inverting the integral'.

Second example:
 $p(x)/\sqrt{x}$
(divergence with
wiggles)



VEGAS

Problem to adapt in multiple dimensions:

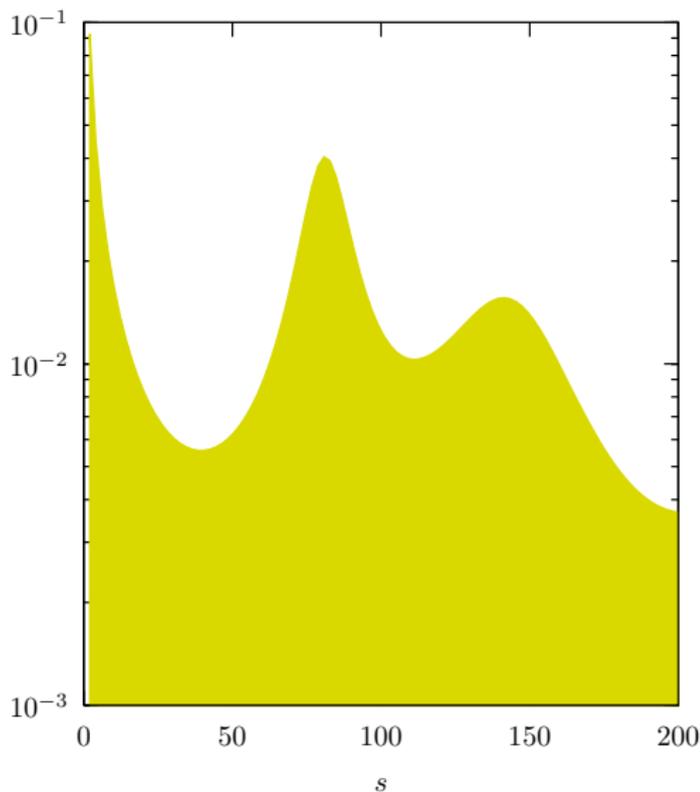


[from T. Ohl, VAMP]

Multichannel MC

Typical problem:

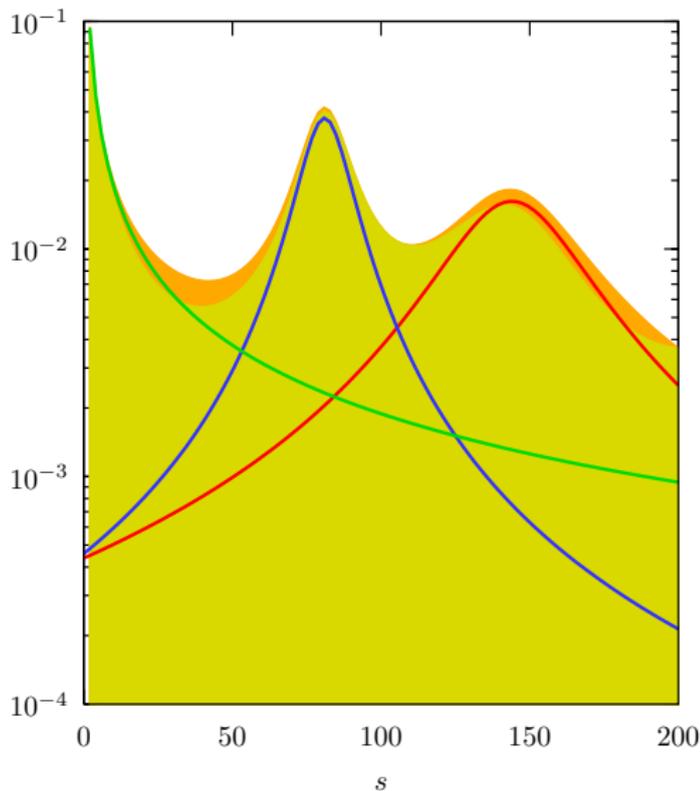
- $f(s)$ has multiple peaks (\times wiggles from ME).



Multichannel MC

Typical problem:

- $f(s)$ has multiple peaks (\times wiggles from ME).
- Usually have some idea of the peak structure.

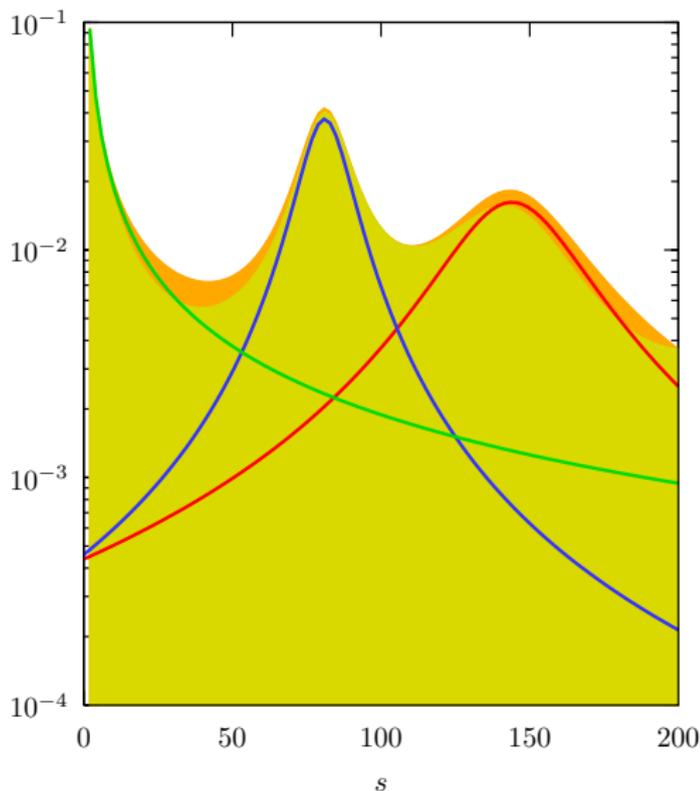


Multichannel MC

Typical problem:

- $f(s)$ has multiple peaks (\times wiggles from ME).
- Usually have some idea of the peak structure.
- Encode this in sum of sample functions $g_i(s)$ with weights $\alpha_i, \sum_i \alpha_i = 1$.

$$g(s) = \sum_i \alpha_i g_i(s) .$$



Multichannel MC

Now rewrite

$$\begin{aligned}\int_{s_0}^{s_1} f(s) ds &= \int_{s_0}^{s_1} \frac{f(s)}{g(s)} g(s) ds \\ &= \int_{s_0}^{s_1} \frac{f(s)}{g(s)} \sum_i \alpha_i g_i(s) ds \\ &= \sum_i \alpha_i \int_{s_0}^{s_1} \frac{f(s)}{g(s)} g_i(s) ds\end{aligned}$$

Now $g_i(s) ds = d\rho_i$ (inverting the integral).

Multichannel MC

Now rewrite

$$\begin{aligned}\int_{s_0}^{s_1} f(s) ds &= \int_{s_0}^{s_1} \frac{f(s)}{g(s)} g(s) ds \\ &= \int_{s_0}^{s_1} \frac{f(s)}{g(s)} \sum_i \alpha_i g_i(s) ds \\ &= \sum_i \alpha_i \int_{s_0}^{s_1} \frac{f(s)}{g(s)} g_i(s) ds\end{aligned}$$

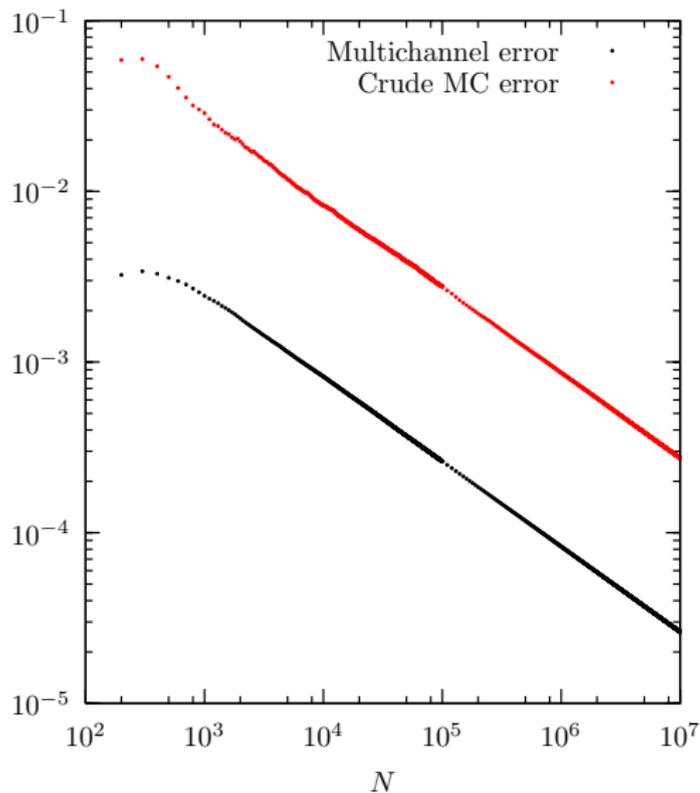
Now $g_i(s) ds = d\rho_i$ (inverting the integral).

Select the distribution $g_i(s)$ you'd like to sample next event from acc to weights α_i .

α_i can be optimized after a number of trials.

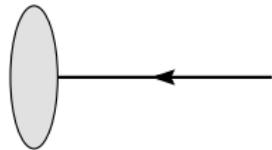
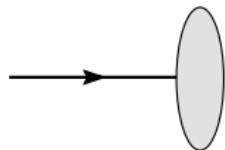
Multichannel MC

Works quite well:

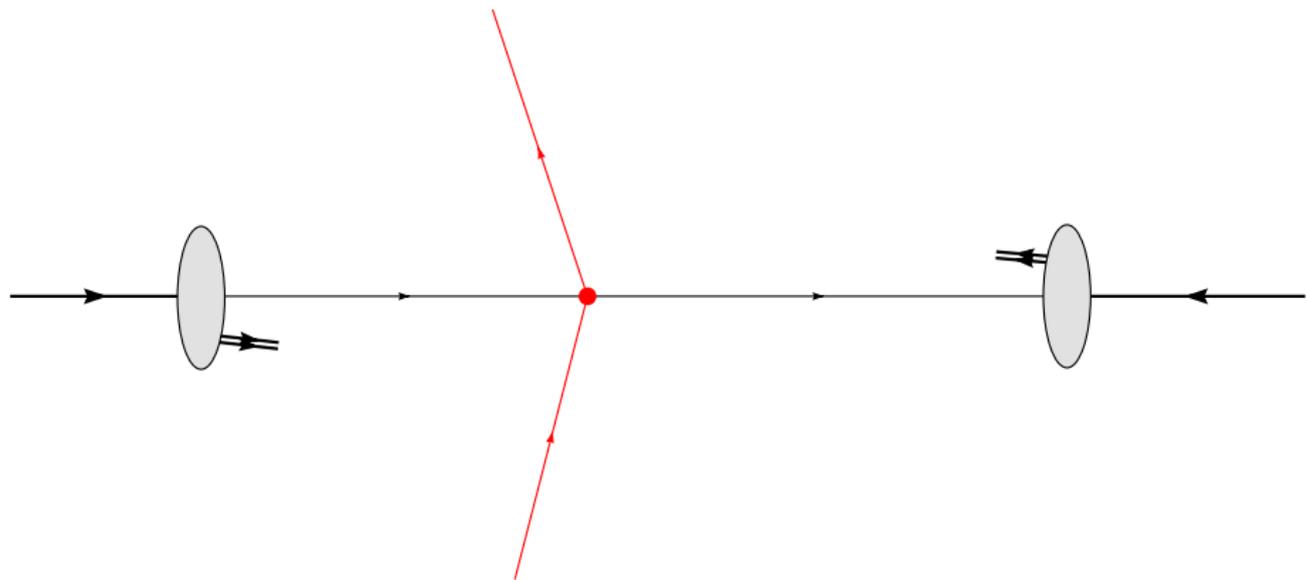


Hard Scattering

Hard scattering

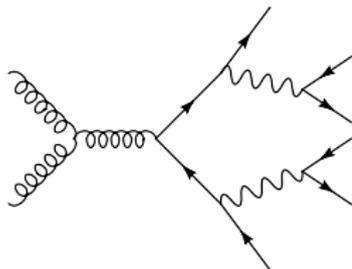


Hard scattering



Matrix elements

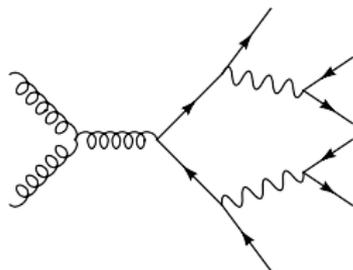
- Perturbation theory/Feynman diagrams give us (fairly accurate) final states for a few number of legs ($O(1)$).



- OK for very inclusive observables.

Matrix elements

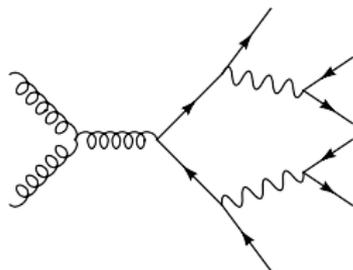
- Perturbation theory/Feynman diagrams give us (fairly accurate) final states for a few number of legs ($O(1)$).



- OK for very inclusive observables.
- Starting point for further simulation.
- Want exclusive final state at the LHC ($O(100)$).

Matrix elements

- Perturbation theory/Feynman diagrams give us (fairly accurate) final states for a few number of legs ($O(1)$).



- OK for very inclusive observables.
- Starting point for further simulation.
- Want exclusive final state at the LHC ($O(100)$).
- Want arbitrary cuts.
- \rightarrow use Monte Carlo methods.

Matrix elements

Where do we get (LO) $|M|^2$ from?

- Most/important simple processes (SM and BSM) are ‘built in’.
- Calculate yourself (≤ 3 particles in final state).
- Matrix element generators:
 - MadGraph/MadEvent.
 - Comix/AMEGIC (part of Sherpa).
 - HELAC/PHEGAS.
 - Whizard.
 - CalcHEP/CompHEP.

generate code or event files that can be further processed.

- \rightarrow FeynRules interface to ME generators.

Also NLO mostly automatically available.

See “Matching and Merging”.

Cross section formula

From Matrix element, we calculate

$$\sigma = \int f_i(x_1, \mu^2) f_j(x_2, \mu^2) \frac{1}{F} \sum |M|^2 dx_1 dx_2 d\Phi_n ,$$

Cross section formula

From Matrix element, we calculate

$$\sigma = \int f_i(x_1, \mu^2) f_j(x_2, \mu^2) \frac{1}{F} \sum |M|^2 \Theta(\text{cuts}) dx_1 dx_2 d\Phi_n ,$$

Cross section formula

From Matrix element, we calculate

$$\sigma = \int f_i(x_1, \mu^2) f_j(x_2, \mu^2) \frac{1}{F} \overline{\sum} |M|^2 \Theta(\text{cuts}) dx_1 dx_2 d\Phi_n ,$$

now,

$$\frac{1}{F} dx_1 dx_2 d\Phi_n = J(\vec{x}) \prod_{i=1}^{3n-2} dx_i \quad \left(d\Phi_n = (2\pi)^4 \delta^{(4)}(\dots) \prod_{i=1}^n \frac{d^3\vec{p}}{(2\pi)^3 2E_i} \right)$$

such that

$$\begin{aligned} \sigma &= \int g(\vec{x}) d^{3n-2}\vec{x} , & \left(g(\vec{x}) = J(\vec{x}) f_i f_j \overline{\sum} |M|^2 \Theta(\text{cuts}) \right) \\ &= \frac{1}{N} \sum_{i=1}^N \frac{g(\vec{x}_i)}{p(\vec{x}_i)} = \frac{1}{N} \sum_{i=1}^N w_i . \end{aligned}$$

Cross section formula

From Matrix element, we calculate

$$\sigma = \int f_i(x_1, \mu^2) f_j(x_2, \mu^2) \frac{1}{F} \overline{\sum} |M|^2 \Theta(\text{cuts}) dx_1 dx_2 d\Phi_n ,$$

now,

$$\frac{1}{F} dx_1 dx_2 d\Phi_n = J(\vec{x}) \prod_{i=1}^{3n-2} dx_i \quad \left(d\Phi_n = (2\pi)^4 \delta^{(4)}(\dots) \prod_{i=1}^n \frac{d^3\vec{p}}{(2\pi)^3 2E_i} \right)$$

such that

$$\begin{aligned} \sigma &= \int g(\vec{x}) d^{3n-2}\vec{x} , & \left(g(\vec{x}) = J(\vec{x}) f_i f_j \overline{\sum} |M|^2 \Theta(\text{cuts}) \right) \\ &= \frac{1}{N} \sum_{i=1}^N \frac{g(\vec{x}_i)}{p(\vec{x}_i)} = \frac{1}{N} \sum_{i=1}^N w_i . \end{aligned}$$

We generate **events** \vec{x}_i with **weights** w_i .

Mini event generator

- We generate pairs (\vec{x}_i, w_i) .

Mini event generator

- We generate pairs (\vec{x}_i, w_i) .
- Use immediately to book weighted histogram of arbitrary observable (possibly with additional cuts!)

Mini event generator

- We generate pairs (\vec{x}_i, w_i) .
- Use immediately to book weighted histogram of arbitrary observable (possibly with additional cuts!)
- Keep event \vec{x}_i with probability

$$P_i = \frac{w_i}{w_{\max}} .$$

Generate events with same frequency as in nature!

Mini event generator

- We generate pairs (\vec{x}_i, w_i) .
- Use immediately to book weighted histogram of arbitrary observable (possibly with additional cuts!)
- Keep event \vec{x}_i with probability

$$P_i = \frac{w_i}{w_{\max}},$$

where w_{\max} has to be chosen sensibly.

→ reweighting, when $\max(w_i) = \bar{w}_{\max} > w_{\max}$, as

$$P_i = \frac{w_i}{\bar{w}_{\max}} = \frac{w_i}{w_{\max}} \cdot \frac{w_{\max}}{\bar{w}_{\max}},$$

i.e. reject events with probability $(w_{\max}/\bar{w}_{\max})$ afterwards.

Mini event generator

- We generate pairs (\vec{x}_i, w_i) .
- Use immediately to book weighted histogram of arbitrary observable (possibly with additional cuts!)
- Keep event \vec{x}_i with probability

$$P_i = \frac{w_i}{w_{\max}} .$$

Generate events with same frequency as in nature!

Matrix elements

Some comments:

- Use common Monte Carlo techniques to generate events efficiently. Goal: small variance in w_i distribution!

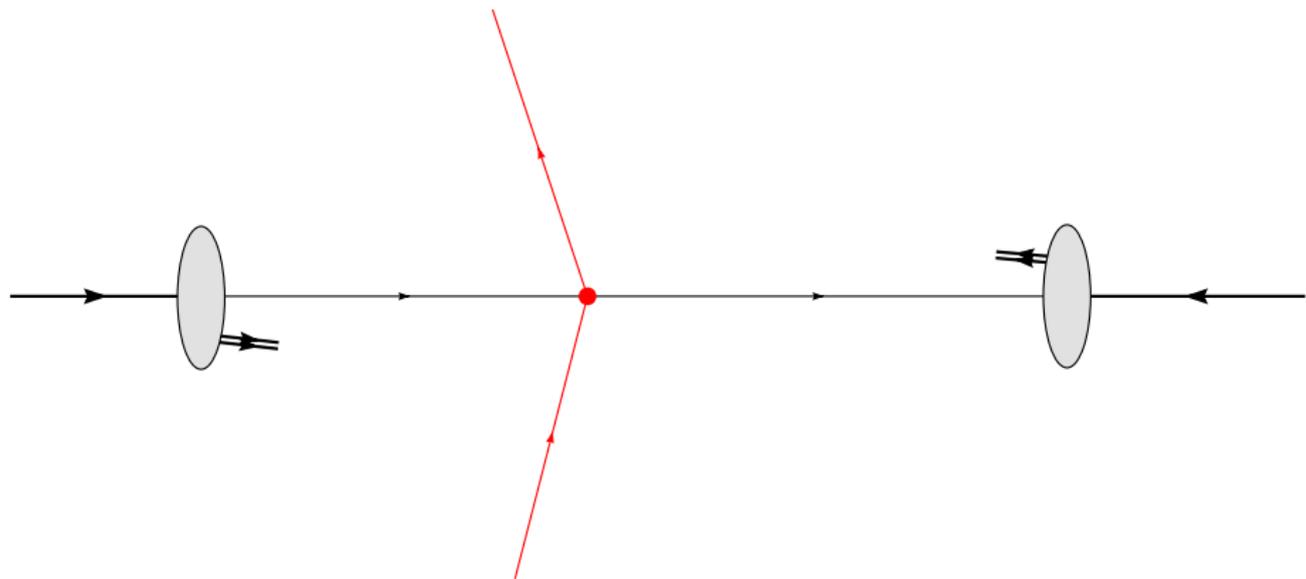
Matrix elements

Some comments:

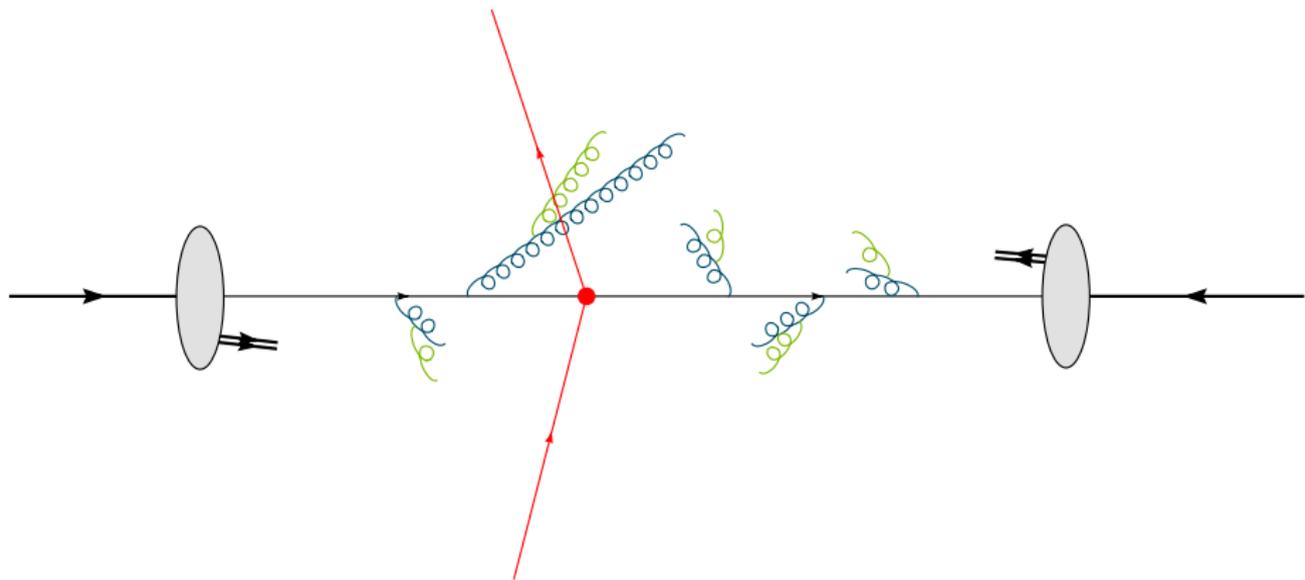
- Use common Monte Carlo techniques to generate events efficiently. Goal: small variance in w_i distribution!
- Efficient generation closely tied to knowledge of $f(\vec{x}_i)$, *i.e.* the matrix element's propagator structure.
→ build phase space generator already while generating ME's automatically.

Parton Showers

Hard matrix element



Hard matrix element \rightarrow parton showers



Parton showers

Quarks and gluons in final state, pointlike.

Parton showers

Quarks and gluons in final state, pointlike.

- Know short distance (short time) fluctuations from matrix element/Feynman diagrams: $Q \sim \text{few GeV to } O(\text{TeV})$.

- Measure hadronic final states, long distance effects, $Q_0 \sim 1 \text{ GeV}$.

Parton showers

Quarks and gluons in final state, pointlike.

- Know short distance (short time) fluctuations from matrix element/Feynman diagrams: $Q \sim \text{few GeV to } O(\text{TeV})$.
- Parton shower evolution, multiple gluon emissions become resolvable at smaller scales. $\text{TeV} \rightarrow 1 \text{ GeV}$.
- Measure hadronic final states, long distance effects, $Q_0 \sim 1 \text{ GeV}$.

Parton showers

Quarks and gluons in final state, pointlike.

- Know short distance (short time) fluctuations from matrix element/Feynman diagrams: $Q \sim \text{few GeV to } O(\text{TeV})$.
- Parton shower evolution, multiple gluon emissions become resolvable at smaller scales. $\text{TeV} \rightarrow 1 \text{ GeV}$.
- Measure hadronic final states, long distance effects, $Q_0 \sim 1 \text{ GeV}$.

Dominated by large logs, terms

$$\alpha_s^n \log^{2n} \frac{Q}{Q_0} \sim 1 .$$

Generated from emissions *ordered* in Q .

Parton showers

Quarks and gluons in final state, pointlike.

- Know short distance (short time) fluctuations from matrix element/Feynman diagrams: $Q \sim \text{few GeV to } O(\text{TeV})$.
- Parton shower evolution, multiple gluon emissions become resolvable at smaller scales. $\text{TeV} \rightarrow 1 \text{ GeV}$.
- Measure hadronic final states, long distance effects, $Q_0 \sim 1 \text{ GeV}$.

Dominated by large logs, terms

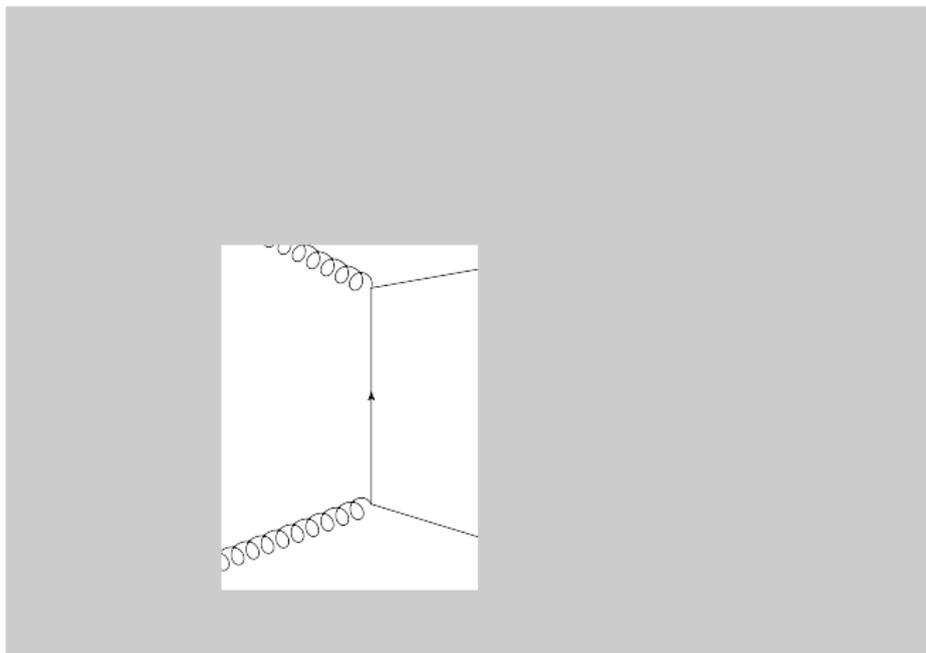
$$\alpha_s^n \log^{2n} \frac{Q}{Q_0} \sim 1 .$$

Generated from emissions *ordered* in Q .

Soft and/or collinear emissions.

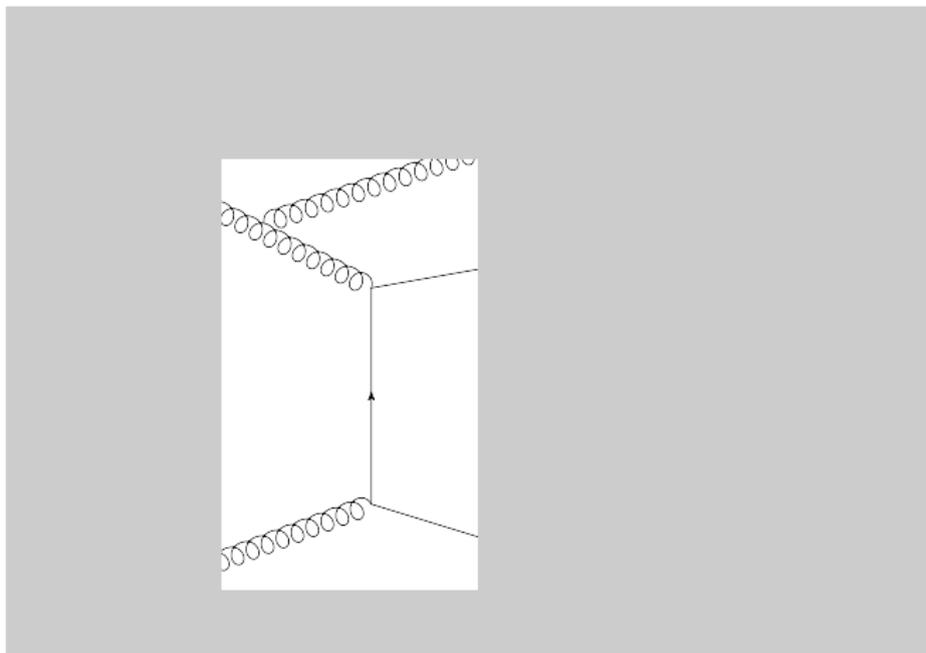
ME approximated by parton cascade

Evolution in scale, typically $Q \sim 1 \text{ TeV}$ down to $Q \sim 1 \text{ GeV}$.



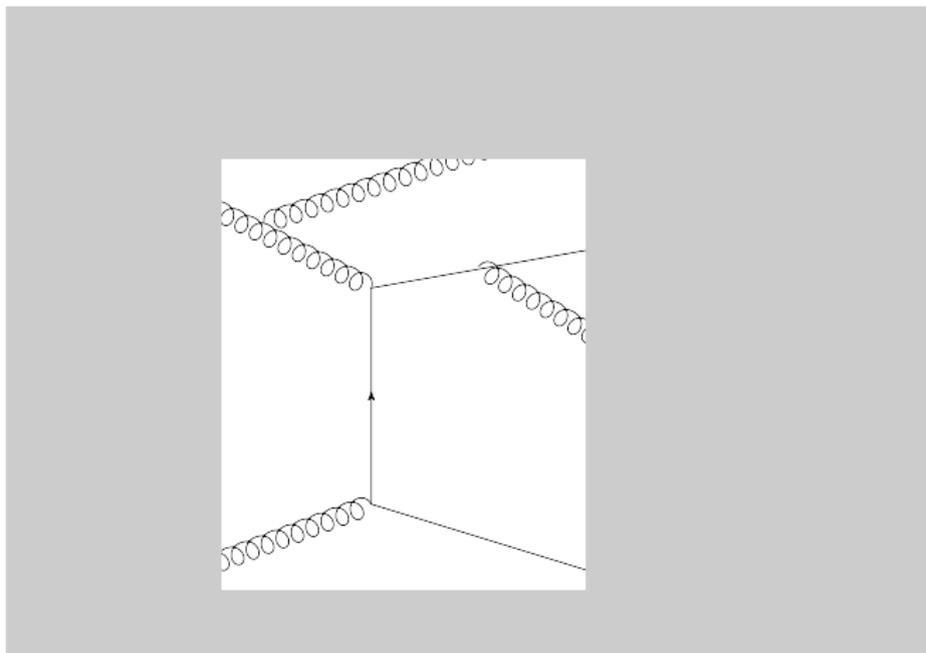
ME approximated by parton cascade

Evolution in scale, typically $Q \sim 1 \text{ TeV}$ down to $Q \sim 1 \text{ GeV}$.



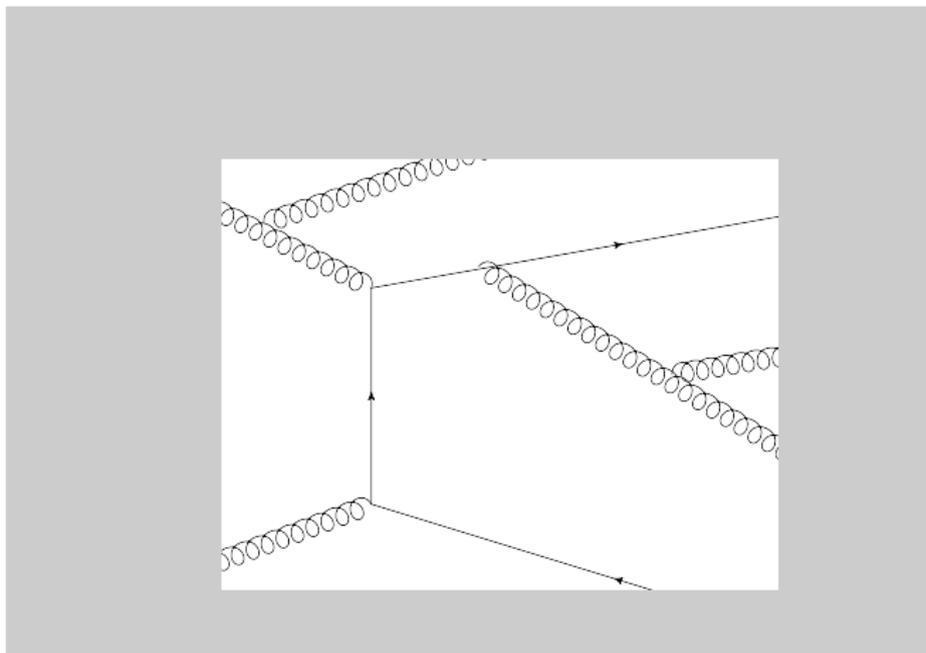
ME approximated by parton cascade

Evolution in scale, typically $Q \sim 1 \text{ TeV}$ down to $Q \sim 1 \text{ GeV}$.



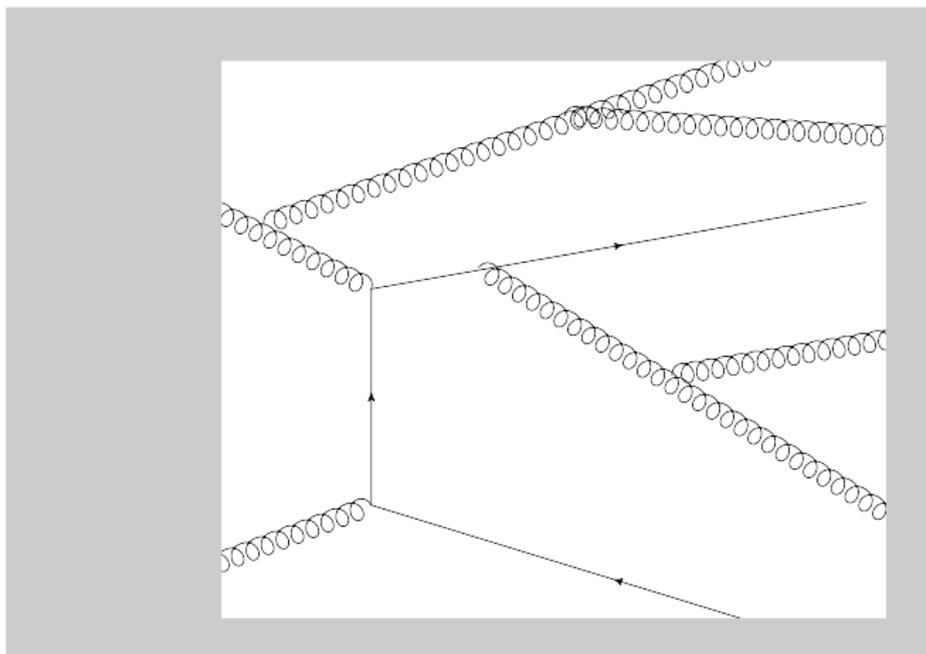
ME approximated by parton cascade

Evolution in scale, typically $Q \sim 1 \text{ TeV}$ down to $Q \sim 1 \text{ GeV}$.



ME approximated by parton cascade

Evolution in scale, typically $Q \sim 1 \text{ TeV}$ down to $Q \sim 1 \text{ GeV}$.



e^+e^- annihilation

Good starting point: $e^+e^- \rightarrow q\bar{q}g$:

Final state momenta in one plane (orientation usually averaged).

Write momenta in terms of

$$x_i = \frac{2p_i \cdot q}{Q^2} \quad (i = 1, 2, 3),$$

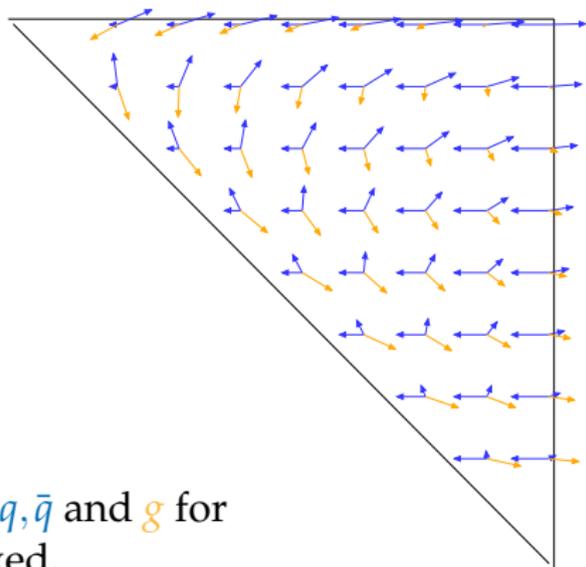
$$0 \leq x_i \leq 1, x_1 + x_2 + x_3 = 2,$$

$$q = (Q, 0, 0, 0),$$

$$Q \equiv E_{cm}.$$

Fig: momentum configuration of q, \bar{q} and g for given point (x_1, x_2) , \bar{q} direction fixed.

$(x_1, x_2) = (x_q, x_{\bar{q}})$ -plane:

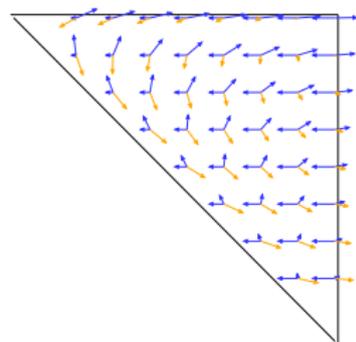
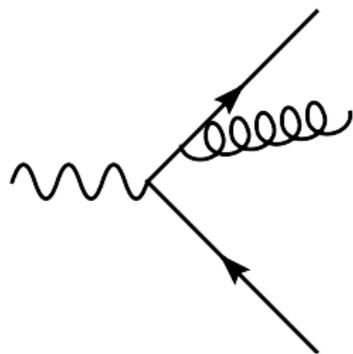


e^+e^- annihilation

Differential cross section:

$$\frac{d\sigma}{dx_1 dx_2} = \sigma_0 \frac{C_F \alpha_S}{2\pi} \frac{x_1^2 + x_2^2}{(1-x_1)(1-x_2)}$$

Collinear singularities: $x_1 \rightarrow 1$ or $x_2 \rightarrow 1$. Soft singularity: $x_1, x_2 \rightarrow 1$.



e^+e^- annihilation

Differential cross section:

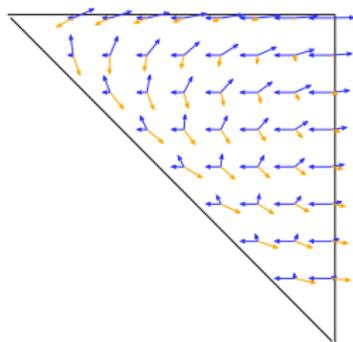
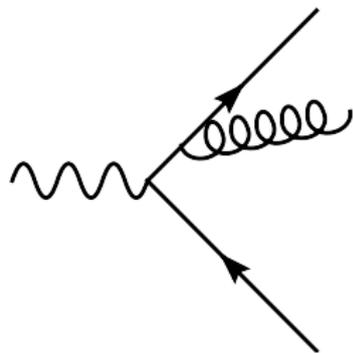
$$\frac{d\sigma}{dx_1 dx_2} = \sigma_0 \frac{C_F \alpha_S}{2\pi} \frac{x_1^2 + x_2^2}{(1-x_1)(1-x_2)}$$

Collinear singularities: $x_1 \rightarrow 1$ or $x_2 \rightarrow 1$. Soft singularity: $x_1, x_2 \rightarrow 1$.

Rewrite in terms of x_3 and $\theta = \angle(q, g)$:

$$\frac{d\sigma}{d\cos\theta dx_3} = \sigma_0 \frac{C_F \alpha_S}{2\pi} \left[\frac{2}{\sin^2\theta} \frac{1 + (1-x_3)^2}{x_3} - x_3 \right]$$

Singular as $\theta \rightarrow 0$ and $x_3 \rightarrow 0$.



e^+e^- annihilation

Can separate into two jets as

$$\begin{aligned}\frac{2d\cos\theta}{\sin^2\theta} &= \frac{d\cos\theta}{1-\cos\theta} + \frac{d\cos\theta}{1+\cos\theta} \\ &= \frac{d\cos\theta}{1-\cos\theta} + \frac{d\cos\bar{\theta}}{1-\cos\bar{\theta}} \\ &\approx \frac{d\theta^2}{\theta^2} + \frac{d\bar{\theta}^2}{\bar{\theta}^2}\end{aligned}$$

e^+e^- annihilation

Can separate into two jets as

$$\begin{aligned}\frac{2d\cos\theta}{\sin^2\theta} &= \frac{d\cos\theta}{1-\cos\theta} + \frac{d\cos\theta}{1+\cos\theta} \\ &= \frac{d\cos\theta}{1-\cos\theta} + \frac{d\cos\bar{\theta}}{1-\cos\bar{\theta}} \\ &\approx \frac{d\theta^2}{\theta^2} + \frac{d\bar{\theta}^2}{\bar{\theta}^2}\end{aligned}$$

So, we rewrite $d\sigma$ in collinear limit as

$$d\sigma = \sigma_0 \sum_{\text{jets}} \frac{d\theta^2}{\theta^2} \frac{\alpha_S}{2\pi} C_F \frac{1+(1-z)^2}{z} dz$$

e^+e^- annihilation

Can separate into two jets as

$$\begin{aligned}\frac{2d\cos\theta}{\sin^2\theta} &= \frac{d\cos\theta}{1-\cos\theta} + \frac{d\cos\theta}{1+\cos\theta} \\ &= \frac{d\cos\theta}{1-\cos\theta} + \frac{d\cos\bar{\theta}}{1-\cos\bar{\theta}} \\ &\approx \frac{d\theta^2}{\theta^2} + \frac{d\bar{\theta}^2}{\bar{\theta}^2}\end{aligned}$$

So, we rewrite $d\sigma$ in collinear limit as

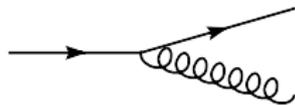
$$\begin{aligned}d\sigma &= \sigma_0 \sum_{\text{jets}} \frac{d\theta^2}{\theta^2} \frac{\alpha_S}{2\pi} C_F \frac{1+(1-z)^2}{z} dz \\ &= \sigma_0 \sum_{\text{jets}} \frac{d\theta^2}{\theta^2} \frac{\alpha_S}{2\pi} P(z) dz\end{aligned}$$

with DGLAP splitting function $P(z)$.

Collinear limit

Universal DGLAP splitting kernels for collinear limit:

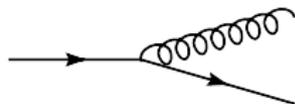
$$d\sigma = \sigma_0 \sum_{\text{jets}} \frac{d\theta^2}{\theta^2} \frac{\alpha_S}{2\pi} P(z) dz$$



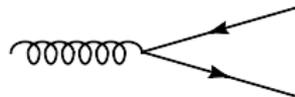
$$P_{q \rightarrow qg}(z) = C_F \frac{1+z^2}{1-z}$$



$$P_{g \rightarrow gg}(z) = C_A \frac{(1-z(1-z))^2}{z(1-z)}$$



$$P_{q \rightarrow gq}(z) = C_F \frac{1+(1-z)^2}{z}$$



$$P_{g \rightarrow qq}(z) = T_R(1-2z(1-z))$$

Collinear limit

Universal DGLAP splitting kernels for collinear limit:

$$d\sigma = \sigma_0 \sum_{\text{jets}} \frac{d\theta^2}{\theta^2} \frac{\alpha_S}{2\pi} P(z) dz$$

Note: Other variables may equally well characterize the collinear limit:

$$\frac{d\theta^2}{\theta^2} \sim \frac{dQ^2}{Q^2} \sim \frac{dp_{\perp}^2}{p_{\perp}^2} \sim \frac{d\tilde{q}^2}{\tilde{q}^2} \sim \frac{dt}{t}$$

whenever $Q^2, p_{\perp}^2, t \rightarrow 0$ means “collinear”.

Collinear limit

Universal DGLAP splitting kernels for collinear limit:

$$d\sigma = \sigma_0 \sum_{\text{jets}} \frac{d\theta^2}{\theta^2} \frac{\alpha_S}{2\pi} P(z) dz$$

Note: Other variables may equally well characterize the collinear limit:

$$\frac{d\theta^2}{\theta^2} \sim \frac{dQ^2}{Q^2} \sim \frac{dp_{\perp}^2}{p_{\perp}^2} \sim \frac{d\tilde{q}^2}{\tilde{q}^2} \sim \frac{dt}{t}$$

whenever $Q^2, p_{\perp}^2, t \rightarrow 0$ means “collinear”.

- θ : HERWIG
- Q^2 : PYTHIA ≤ 6.3 , SHERPA.
- p_{\perp} : PYTHIA ≥ 6.4 , ARIADNE, Catani–Seymour showers.
- \tilde{q} : Herwig++.

Resolution

Need to introduce **resolution** t_0 , e.g. a cutoff in p_{\perp} . Prevent us from the singularity at $\theta \rightarrow 0$.

Emissions below t_0 are **unresolvable**.

Finite result due to virtual corrections:



The diagram shows two Feynman diagrams separated by a plus sign, followed by an equals sign and the word "finite". The first diagram is a horizontal line with a red wavy line (representing a gluon emission) attached to it. The second diagram is a horizontal line with a red loop (representing a virtual correction) attached to it.

unresolvable + virtual emissions are included in Sudakov form factor via unitarity (see below!).

Towards multiple emissions

Starting point: factorisation in collinear limit, single emission.

$$\sigma_{2+1}(t_0) = \sigma_2(t_0) \int_{t_0}^t \frac{dt'}{t'} \int_{z_-}^{z_+} dz \frac{\alpha_S}{2\pi} \hat{P}(z) = \sigma_2(t_0) \int_{t_0}^t dt W(t) .$$

Towards multiple emissions

Starting point: factorisation in collinear limit, single emission.

$$\sigma_{2+1}(t_0) = \sigma_2(t_0) \int_{t_0}^t \frac{dt'}{t'} \int_{z_-}^{z_+} dz \frac{\alpha_S}{2\pi} \hat{P}(z) = \sigma_2(t_0) \int_{t_0}^t dt W(t).$$

Simple example:

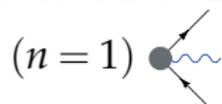
Multiple photon emissions, strongly ordered in t .

We want

$$W_{\text{sum}} = \sum_{n=1} W_{2+n} = \frac{\int \left| \begin{array}{c} \nearrow \\ \bullet \\ \searrow \\ \text{---} \\ \nearrow \end{array} \right|^2 d\Phi_1 + \int \left| \begin{array}{c} \nearrow \\ \bullet \\ \searrow \\ \text{---} \\ \nearrow \\ \text{---} \\ \nearrow \end{array} \right|^2 d\Phi_2 + \int \left| \begin{array}{c} \nearrow \\ \bullet \\ \searrow \\ \text{---} \\ \nearrow \\ \text{---} \\ \nearrow \\ \text{---} \\ \nearrow \end{array} \right|^2 d\Phi_3 + \dots}{\left| \begin{array}{c} \nearrow \\ \bullet \\ \searrow \end{array} \right|^2}$$

for any number of emissions.

Towards multiple emissions



$$W_{2+1} = \left(\int \left| \text{Diagram 1} \right|^2 + \left| \text{Diagram 2} \right|^2 d\Phi_1 \right) / \left| \text{Diagram 3} \right|^2 = \frac{2}{1!} \int_{t_0}^t dt W(t) .$$

The equation uses Feynman diagrams to represent terms in a series. The numerator consists of two terms: an integral over phase space $d\Phi_1$ of the squared magnitude of a diagram with two outgoing lines and one incoming wavy line, plus the squared magnitude of a similar diagram where the wavy line is on the other side. The denominator is the squared magnitude of the vertex diagram shown in the $(n=1)$ diagram above. The result is equal to $\frac{2}{1!} \int_{t_0}^t dt W(t)$.

Towards multiple emissions

$(n = 1)$ 

$$W_{2+1} = \left(\int \left| \begin{array}{c} \nearrow \\ \leftarrow \\ \leftarrow \\ \searrow \end{array} \right|^2 + \left| \begin{array}{c} \nearrow \\ \leftarrow \\ \leftarrow \\ \searrow \\ \leftarrow \end{array} \right|^2 d\Phi_1 \right) / \left| \begin{array}{c} \nearrow \\ \leftarrow \\ \leftarrow \\ \searrow \end{array} \right|^2 = \frac{2}{1!} \int_{t_0}^t dt W(t).$$

$(n = 2)$ 

$$W_{2+2} = \left(\int \left| \begin{array}{c} \nearrow \\ \leftarrow \\ \leftarrow \\ \searrow \\ \leftarrow \end{array} \right|^2 + \left| \begin{array}{c} \nearrow \\ \leftarrow \\ \leftarrow \\ \searrow \\ \leftarrow \\ \leftarrow \end{array} \right|^2 + \left| \begin{array}{c} \nearrow \\ \leftarrow \\ \leftarrow \\ \searrow \\ \leftarrow \\ \leftarrow \end{array} \right|^2 + \left| \begin{array}{c} \nearrow \\ \leftarrow \\ \leftarrow \\ \searrow \\ \leftarrow \\ \leftarrow \end{array} \right|^2 d\Phi_2 \right) / \left| \begin{array}{c} \nearrow \\ \leftarrow \\ \leftarrow \\ \searrow \end{array} \right|^2$$

$$= 2^2 \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' W(t') W(t'') = \frac{2^2}{2!} \left(\int_{t_0}^t dt W(t) \right)^2.$$

We used

$$\int_{t_0}^t dt_1 \dots \int_{t_0}^{t_{n-1}} dt_n W(t_1) \dots W(t_n) = \frac{1}{n!} \left(\int_{t_0}^t dt W(t) \right)^n.$$

Towards multiple emissions

Easily generalized to n emissions  by induction. *i.e.*

$$W_{2+n} = \frac{2^n}{n!} \left(\int_{t_0}^t dt W(t) \right)^n$$

Towards multiple emissions

Easily generalized to n emissions  by induction. *i.e.*

$$W_{2+n} = \frac{2^n}{n!} \left(\int_{t_0}^t dt W(t) \right)^n$$

So, in total we get

$$\sigma_{>2}(t_0) = \sigma_2(t_0) \sum_{k=1}^{\infty} \frac{2^k}{k!} \left(\int_{t_0}^t dt W(t) \right)^k = \sigma_2(t_0) \left(e^{2 \int_{t_0}^t dt W(t)} - 1 \right)$$

Towards multiple emissions

Easily generalized to n emissions  by induction. *i.e.*

$$W_{2+n} = \frac{2^n}{n!} \left(\int_{t_0}^t dt W(t) \right)^n$$

So, in total we get

$$\begin{aligned} \sigma_{>2}(t_0) &= \sigma_2(t_0) \sum_{k=1}^{\infty} \frac{2^k}{k!} \left(\int_{t_0}^t dt W(t) \right)^k = \sigma_2(t_0) \left(e^{2 \int_{t_0}^t dt W(t)} - 1 \right) \\ &= \sigma_2(t_0) \left(\frac{1}{\Delta^2(t_0, t)} - 1 \right) \end{aligned}$$

Sudakov Form Factor

$$\Delta(t_0, t) = \exp \left[- \int_{t_0}^t dt W(t) \right]$$

Towards multiple emissions

Easily generalized to n emissions  by induction. *i.e.*

$$W_{2+n} = \frac{2^n}{n!} \left(\int_{t_0}^t dt W(t) \right)^n$$

So, in total we get

$$\begin{aligned} \sigma_{>2}(t_0) &= \sigma_2(t_0) \sum_{k=1}^{\infty} \frac{2^k}{k!} \left(\int_{t_0}^t dt W(t) \right)^k = \sigma_2(t_0) \left(e^{2 \int_{t_0}^t dt W(t)} - 1 \right) \\ &= \sigma_2(t_0) \left(\frac{1}{\Delta^2(t_0, t)} - 1 \right) \end{aligned}$$

Sudakov Form Factor in QCD

$$\Delta(t_0, t) = \exp \left[- \int_{t_0}^t dt W(t) \right] = \exp \left[- \int_{t_0}^t \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \right]$$

Sudakov form factor

Note that

$$\begin{aligned}\sigma_{\text{all}} &= \sigma_2 + \sigma_{>2} = \sigma_2 + \sigma_2 \left(\frac{1}{\Delta^2(t_0, t)} - 1 \right), \\ \Rightarrow \Delta^2(t_0, t) &= \frac{\sigma_2}{\sigma_{\text{all}}}.\end{aligned}$$

Two jet rate = $\Delta^2 = P^2$ (No emission in the range $t \rightarrow t_0$).

Sudakov form factor = No emission probability .

Often $\Delta(t_0, t) \equiv \Delta(t)$.

- Hard scale t , typically CM energy or p_{\perp} of hard process.
- Resolution t_0 , two partons are resolved as two entities if inv mass or relative p_{\perp} above t_0 .
- P^2 (not P), as we have two legs that evolve independently.

Sudakov form factor from Markov property

Unitarity

$$\begin{aligned} P(\text{"some emission"}) + P(\text{"no emission"}) \\ = P(0 < t \leq T) + \bar{P}(0 < t \leq T) = 1. \end{aligned}$$

Multiplication law (no memory)

$$\bar{P}(0 < t \leq T) = \bar{P}(0 < t \leq t_1) \bar{P}(t_1 < t \leq T)$$

Sudakov form factor from Markov property

Unitarity

$$\begin{aligned} P(\text{"some emission"}) + P(\text{"no emission"}) \\ = P(0 < t \leq T) + \bar{P}(0 < t \leq T) = 1. \end{aligned}$$

Multiplication law (no memory)

$$\bar{P}(0 < t \leq T) = \bar{P}(0 < t \leq t_1) \bar{P}(t_1 < t \leq T)$$

Then subdivide into n pieces: $t_i = \frac{i}{n}T, 0 \leq i \leq n$.

$$\begin{aligned} \bar{P}(0 < t \leq T) &= \lim_{n \rightarrow \infty} \prod_{i=0}^{n-1} \bar{P}(t_i < t \leq t_{i+1}) = \lim_{n \rightarrow \infty} \prod_{i=0}^{n-1} (1 - P(t_i < t \leq t_{i+1})) \\ &= \exp \left(- \lim_{n \rightarrow \infty} \sum_{i=0}^{n-1} P(t_i < t \leq t_{i+1}) \right) = \exp \left(- \int_0^T \frac{dP(t)}{dt} dt \right). \end{aligned}$$

Sudakov form factor

Again, no-emission probability!

$$\bar{P}(0 < t \leq T) = \exp\left(-\int_0^T \frac{dP(t)}{dt} dt\right)$$

So,

$$\begin{aligned} dP(\text{first emission at } T) &= dP(T)\bar{P}(0 < t \leq T) \\ &= dP(T)\exp\left(-\int_0^T \frac{dP(t)}{dt} dt\right) \end{aligned}$$

That's what we need for our parton shower! Probability density for next emission at t :

$$dP(\text{next emission at } t) = \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \exp\left[-\int_{t_0}^t \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz\right]$$

Parton shower Monte Carlo

Probability density:

$dP(\text{next emission at } t) =$

$$\frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \exp \left[- \int_{t_0}^t \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \right]$$

Conveniently, the probability distribution is $\Delta(t)$ itself.

Parton shower Monte Carlo

Probability density:

$$dP(\text{next emission at } t) = \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \exp \left[- \int_{t_0}^t \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \right]$$

Conveniently, the probability distribution is $\Delta(t)$ itself.

Hence, parton shower very roughly from (HERWIG):

- 1 Choose flat random number $0 \leq \rho \leq 1$.
- 2 If $\rho < \Delta(t_{\max})$: no resolvable emission, stop this branch.
- 3 Else solve $\rho = \Delta(t_{\max})/\Delta(t)$
(= no emission between t_{\max} and t) for t .
Reset $t_{\max} = t$ and goto 1.

Determine z essentially according to integrand in front of exp.

Parton shower Monte Carlo

Probability density:

$$dP(\text{next emission at } t) = \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \exp \left[- \int_{t_0}^t \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \right]$$

Conveniently, the probability distribution is $\Delta(t)$ itself.

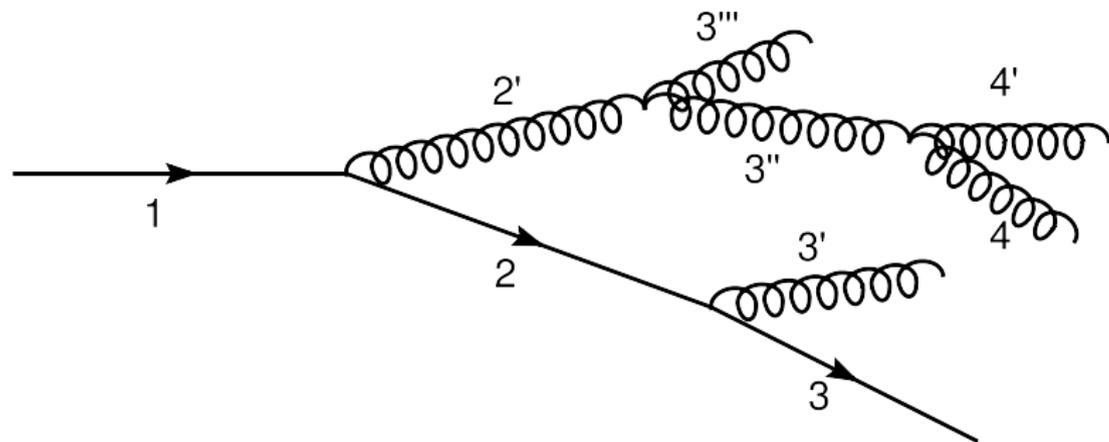
- That was old HERWIG variant. Relies on (numerical) integration/tabulation for $\Delta(t)$.
- Pythia, now also Herwig++, use the **Veto Algorithm**.
- Method to sample x from distribution of the type

$$dP = F(x) \exp \left[- \int^x dx' F(x') \right] dx .$$

Simpler, more flexible, but slightly slower.

Parton cascade

Get tree structure, ordered in evolution variable t :



Here: $t_1 > t_2 > t_3; t_2 > t_{3'}$ etc.

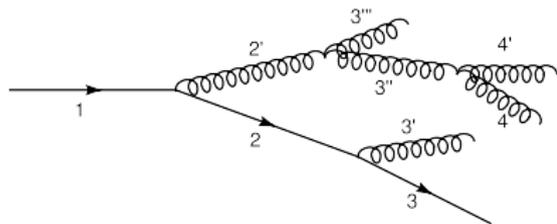
Construct four momenta from (t_i, z_i) and (random) azimuth ϕ .

Not at all unique!

Many (more or less clever) choices still to be made.

Parton cascade

Get tree structure, ordered in evolution variable t :



- t can be $\theta, Q^2, p_{\perp}, \dots$
- Choice of hard scale t_{\max} not fixed. “Some hard scale”.
- z can be light cone momentum fraction, energy fraction, ...
- Available parton shower phase space.
- Integration limits.
- Regularisation of soft singularities.
- ...

Good choices needed here to describe wealth of data!

Soft emissions

- Only *collinear* emissions so far.
- Including *collinear+soft*.
- *Large angle+soft* also important.

Soft emissions

- Only *collinear* emissions so far.
- Including *collinear+soft*.
- *Large angle+soft* also important.

Soft emission: consider *eikonal factors*,
here for $q(p+q) \rightarrow q(p)g(q)$, soft g :

$$u(p) \not{\epsilon} \frac{\not{p} + \not{q} + m}{(p+q)^2 - m^2} \longrightarrow u(p) \frac{p \cdot \epsilon}{p \cdot q}$$

soft factorisation. Universal, *i.e.* independent of emitter.
In general:

$$d\sigma_{n+1} = d\sigma_n \frac{d\omega}{\omega} \frac{d\Omega}{2\pi} \frac{\alpha_S}{2\pi} \sum_{ij} C_{ij} W_{ij} \quad (\text{"QCD-Antenna"})$$

with

$$W_{ij} = \frac{1 - \cos \theta_{ij}}{(1 - \cos \theta_{iq})(1 - \cos \theta_{jq})} .$$

Soft emissions

We define

$$W_{ij} = \frac{1 - \cos \theta_{ij}}{(1 - \cos \theta_{iq})(1 - \cos \theta_{qj})} \equiv W_{ij}^{(i)} + W_{ij}^{(j)}$$

with

$$W_{ij}^{(i)} = \frac{1}{2} \left(W_{ij} + \frac{1}{1 - \cos \theta_{iq}} - \frac{1}{1 - \cos \theta_{qj}} \right) .$$

$W_{ij}^{(i)}$ is only collinear divergent if $q \parallel i$ etc .

Soft emissions

We define

$$W_{ij} = \frac{1 - \cos \theta_{ij}}{(1 - \cos \theta_{iq})(1 - \cos \theta_{qj})} \equiv W_{ij}^{(i)} + W_{ij}^{(j)}$$

with

$$W_{ij}^{(i)} = \frac{1}{2} \left(W_{ij} + \frac{1}{1 - \cos \theta_{iq}} - \frac{1}{1 - \cos \theta_{qj}} \right).$$

$W_{ij}^{(i)}$ is only collinear divergent if $q \parallel i$ etc .

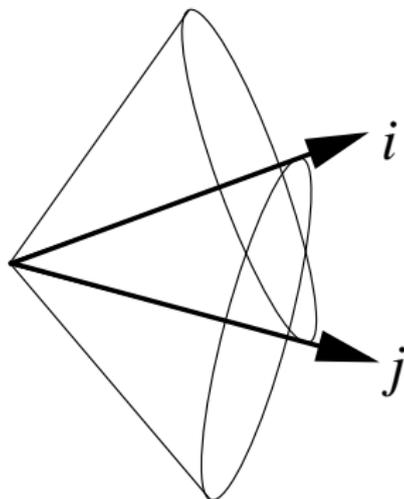
After integrating out the azimuthal angles, we find

$$\int \frac{d\phi_{iq}}{2\pi} W_{ij}^{(i)} = \begin{cases} \frac{1}{1 - \cos \theta_{iq}} & (\theta_{iq} < \theta_{ij}) \\ 0 & \text{otherwise} \end{cases}$$

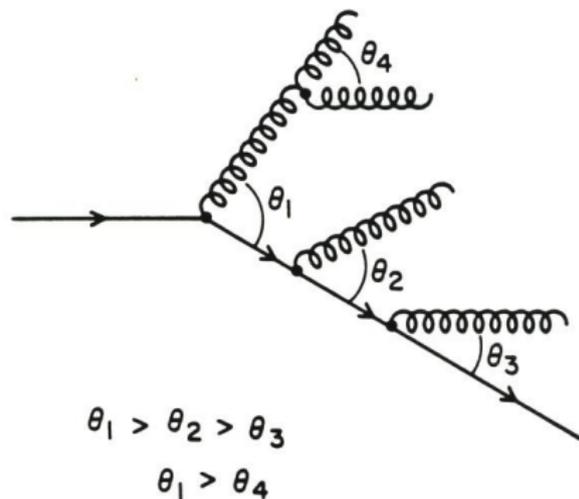
That's angular ordering.

Angular ordering

Radiation from parton i is bound to a cone, given by the colour partner parton j .



Results in angular ordered parton shower and suppresses soft gluons viz. hadrons in a jet.



Colour coherence from CDF

Events with 2 hard (> 100 GeV) jets and a soft 3rd jet (~ 10 GeV)

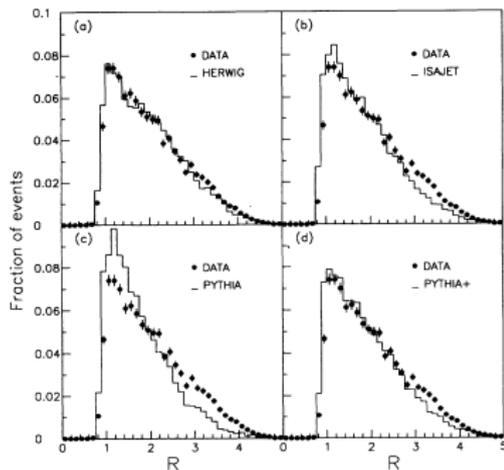


FIG. 14. Observed R distribution compared to the predictions of (a) HERWIG; (b) ISAJET; (c) PYTHIA; (d) PYTHIA+.

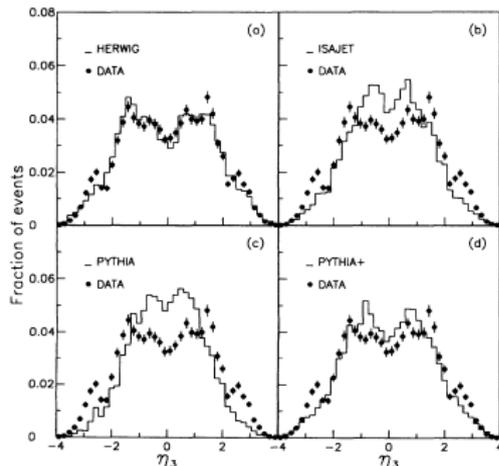


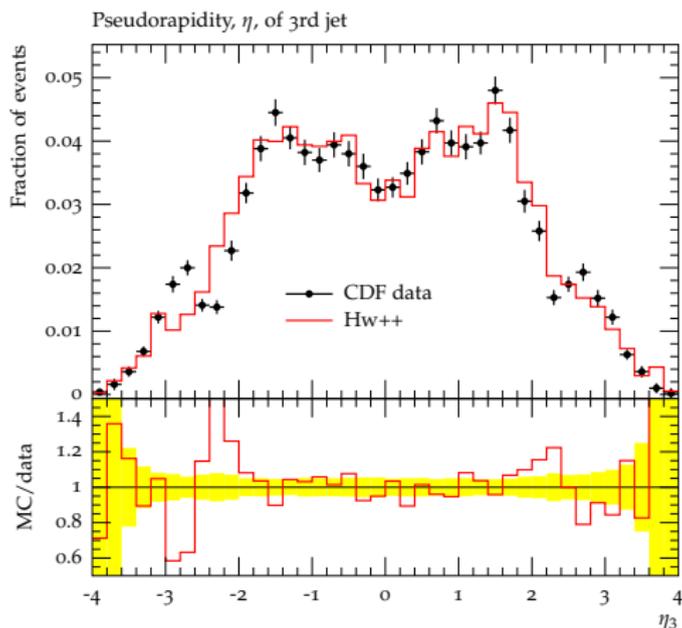
FIG. 13. Observed η_3 distribution compared to the predictions of (a) HERWIG; (b) ISAJET; (c) PYTHIA; (d) PYTHIA+.

F. Abe *et al.* [CDF Collaboration], *Phys. Rev. D* **50** (1994) 5562.

Best description with angular ordering.

Colour coherence from CDF

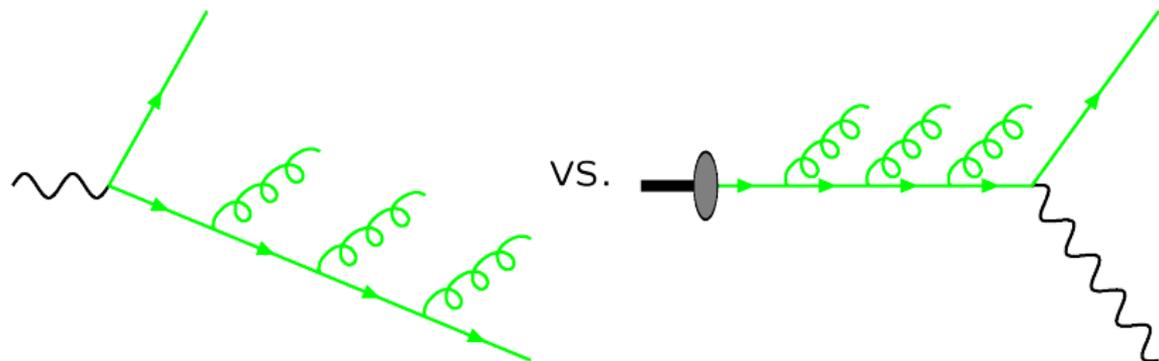
Events with 2 hard (> 100 GeV) jets and a soft 3rd jet (~ 10 GeV)



F. Abe *et al.* [CDF Collaboration], Phys. Rev. D **50** (1994) 5562.

Best description with angular ordering.

Initial state radiation



Similar to final state radiation. Sudakov form factor ($x' = x/z$)

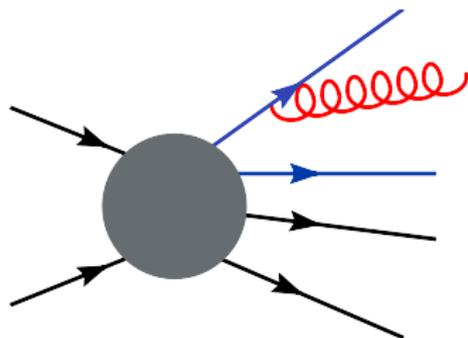
$$\Delta(t, t_{\max}) = \exp \left[- \sum_b \int_t^{t_{\max}} \frac{dt}{t} \int_{z_-}^{z_+} dz \frac{\alpha_S(z, t)}{2\pi} \frac{x' f_b(x', t)}{x f_a(x, t)} \hat{P}_{ba}(z, t) \right]$$

Have to **divide out the pdfs.**

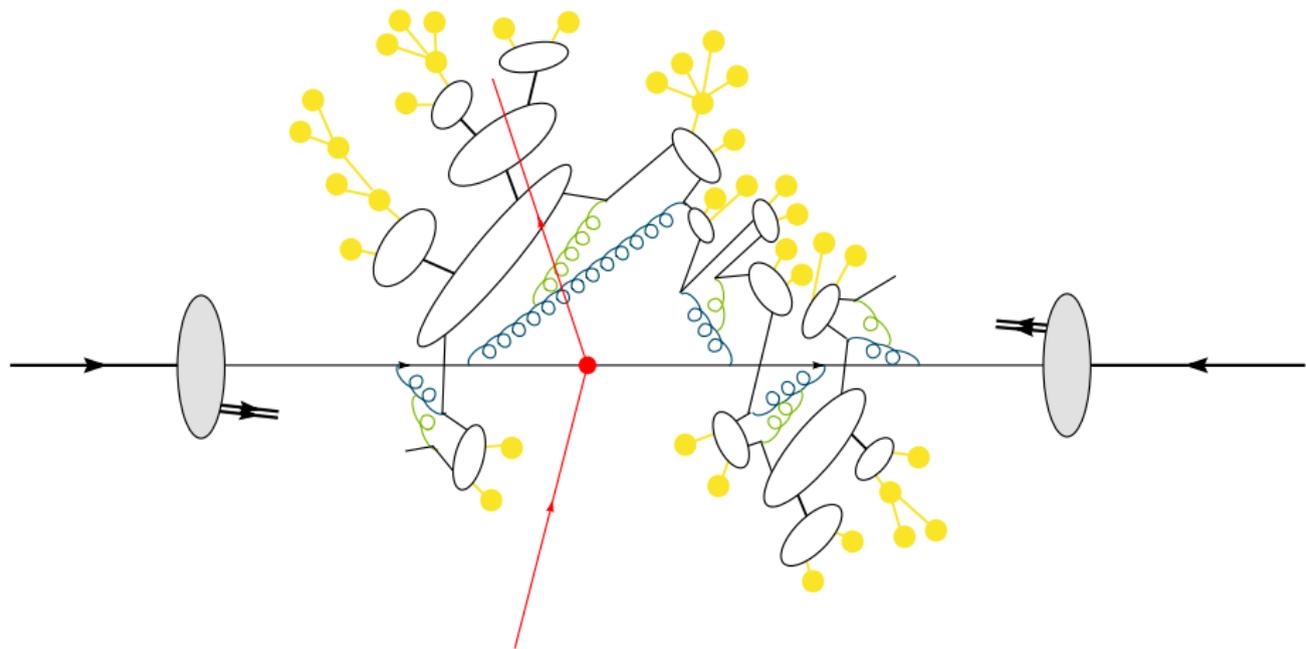
Dipoles

Exact kinematics when recoil is taken by *spectator(s)*.

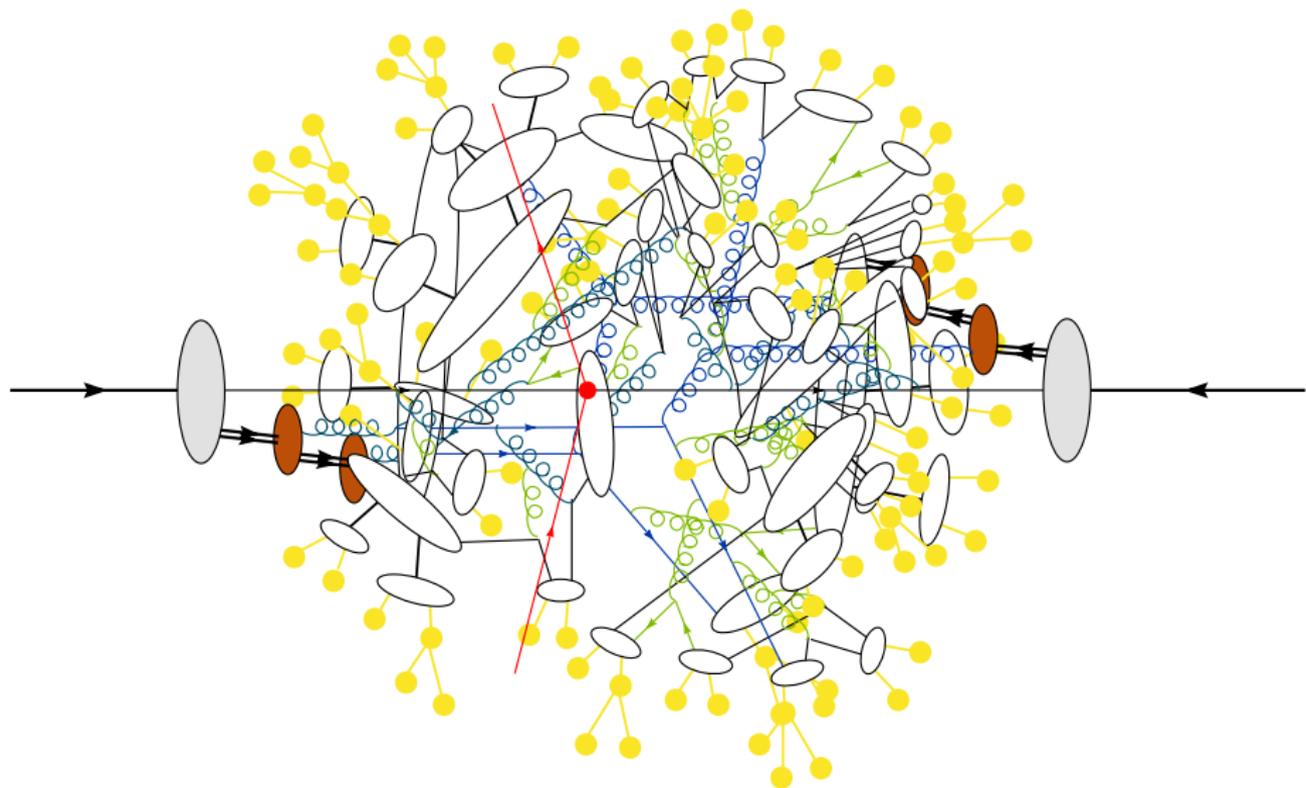
- Dipole showers.
- Ariadne.
- Recoils in Pythia.
- New dipole showers, based on
 - Catani Seymour dipoles.
 - QCD Antennae.
 - Herwig, Sherpa, Vincia, Dire, ...
 - Goal: matching with NLO.
- Generalized to IS–IS, IS–FS.



Brief graphical summary



Brief graphical summary



A few plots

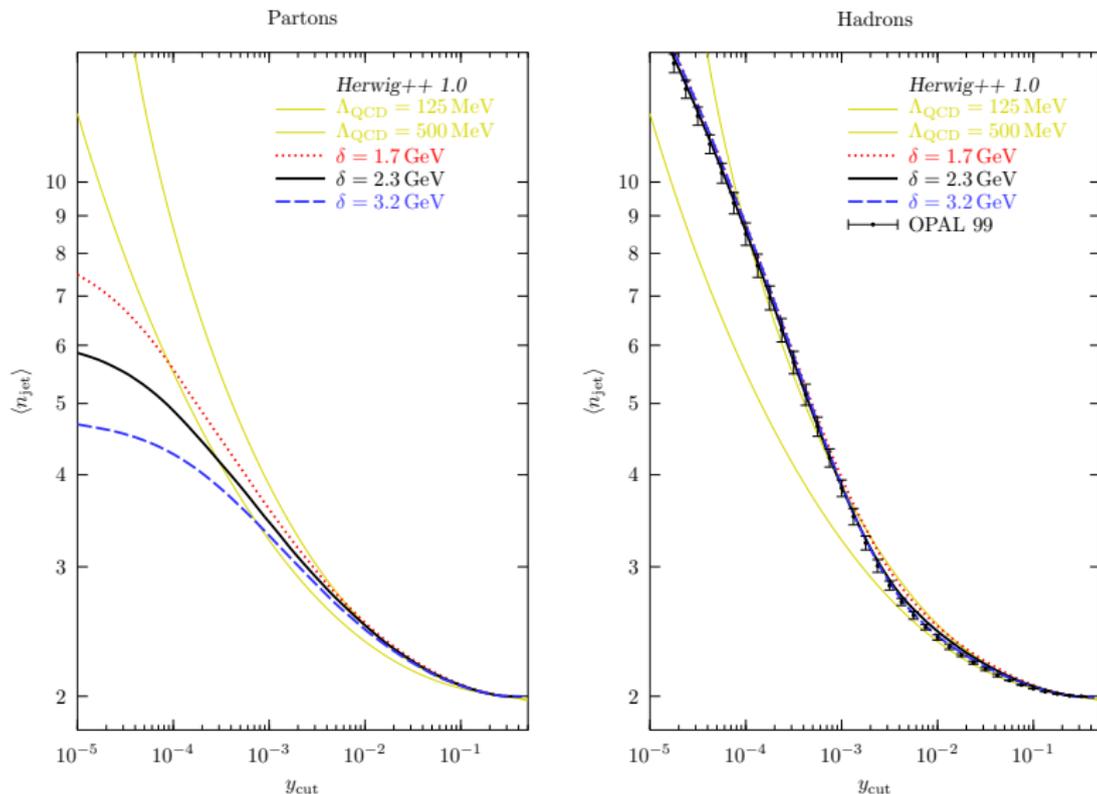
How well does it work?

- $e^+e^- \rightarrow$ hadrons, mostly at LEP.
- Jet shapes, jet rates, event shapes, identified particles...
- 'Tuning' of parameters.
- Use *all* analyses available in Rivet.
- Want to get *everything* right with *one* parameter set.
- Compare to literally ≈ 20000 plots.

- Check out <http://herwig.hepforge.org>
(\rightarrow Plots) for many more and comparisons with the latest release.

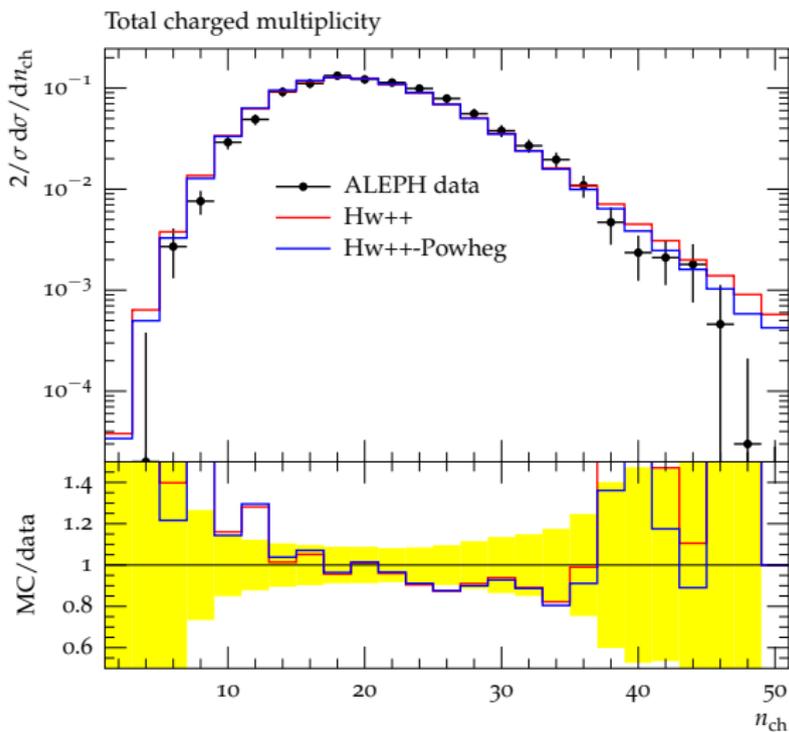
How well does it work?

Smooth interplay between shower and hadronization.



How well does it work?

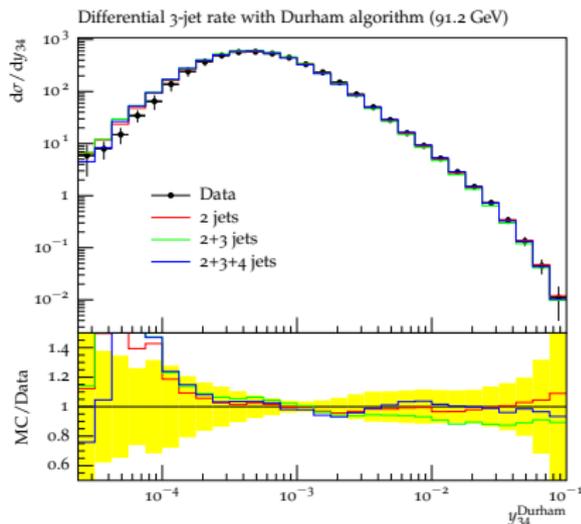
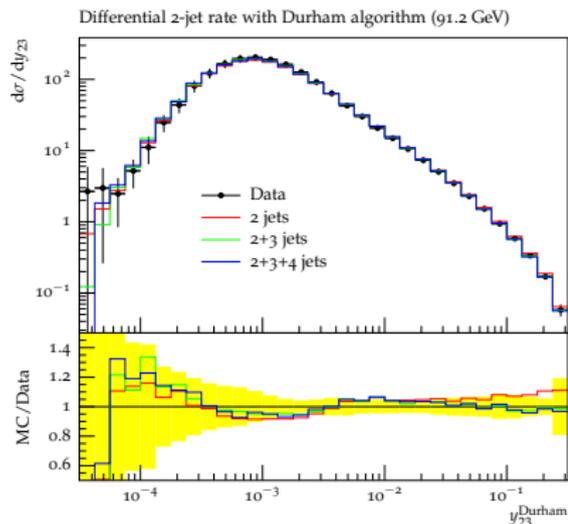
N_{ch} at LEP. Crucial for t_0 (Herwig++ 2.5.2)



How well does it work?

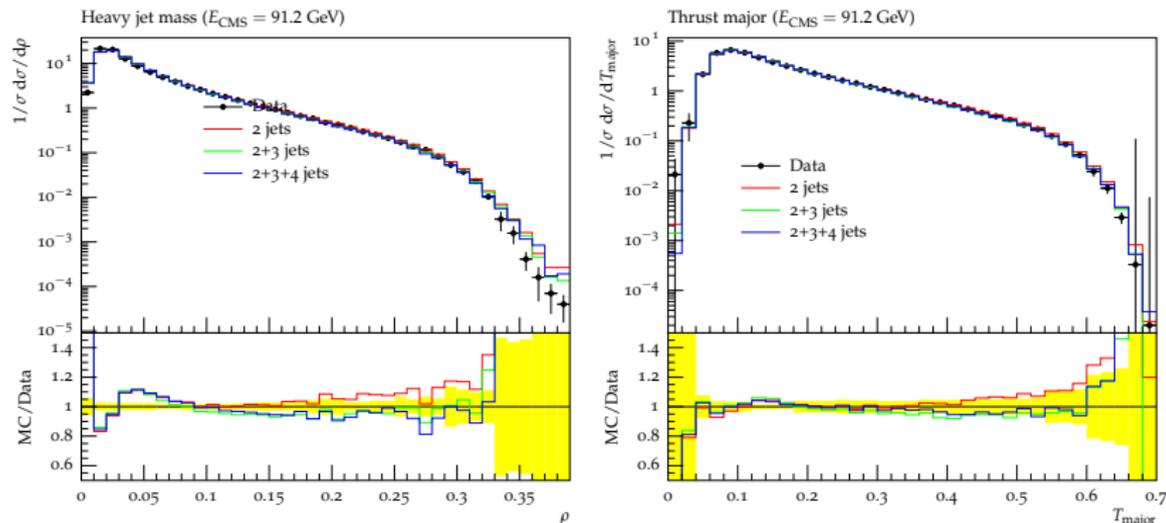
How well does it work?

Differential Jet Rates at LEP (Herwig++ pre-3.0). Dipole shower + some merging



How well does it work?

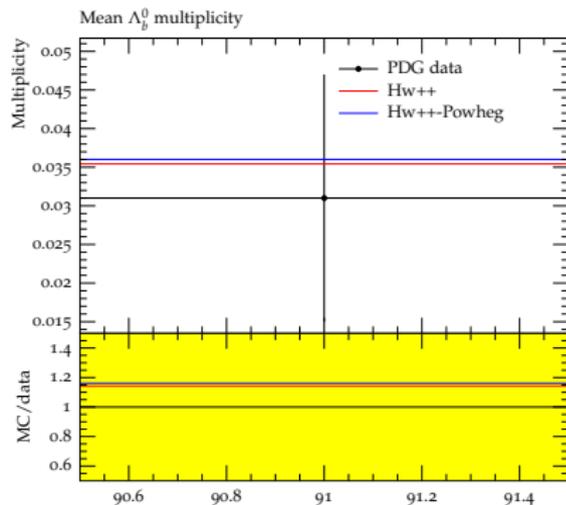
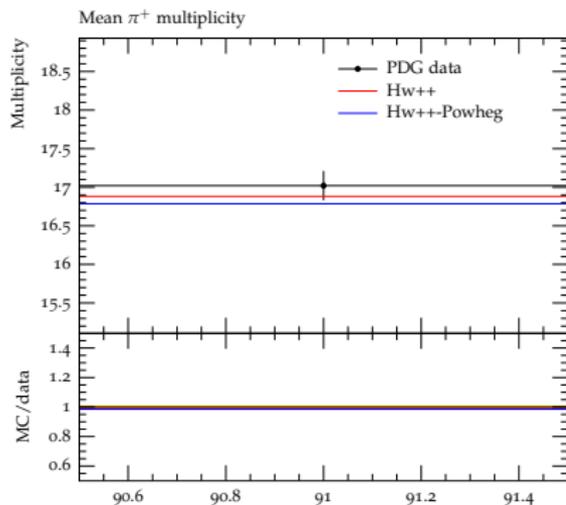
Event Shapes at LEP (Herwig++ pre-3.0).
Dipole shower + some merging



Parton showers do very well, today!

How well does it work?

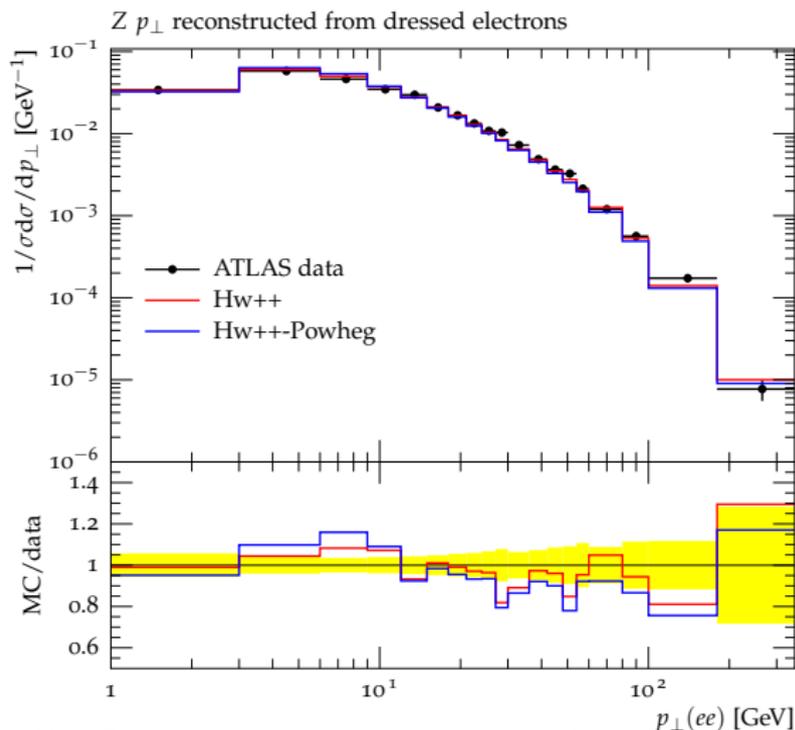
Hadron Multiplicities at LEP (e.g. π^+ , Λ_b^0).



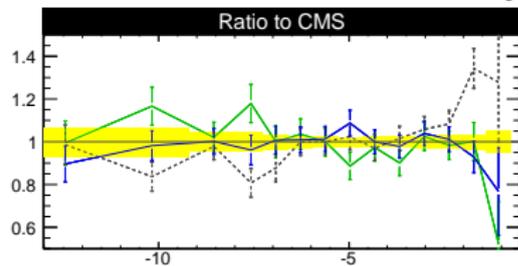
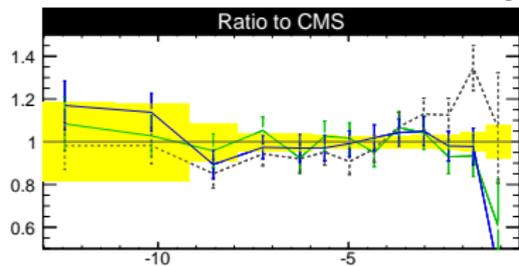
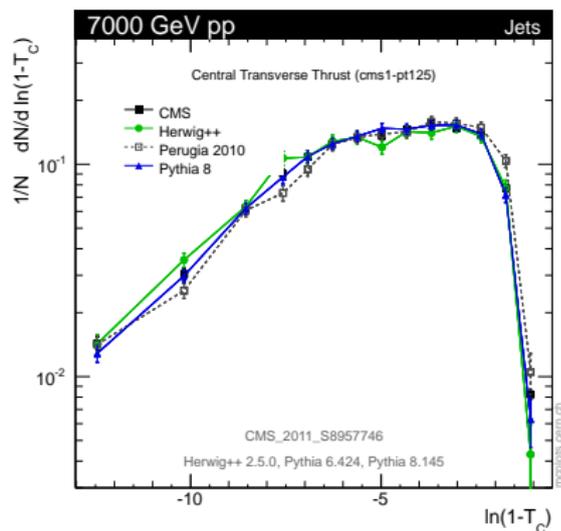
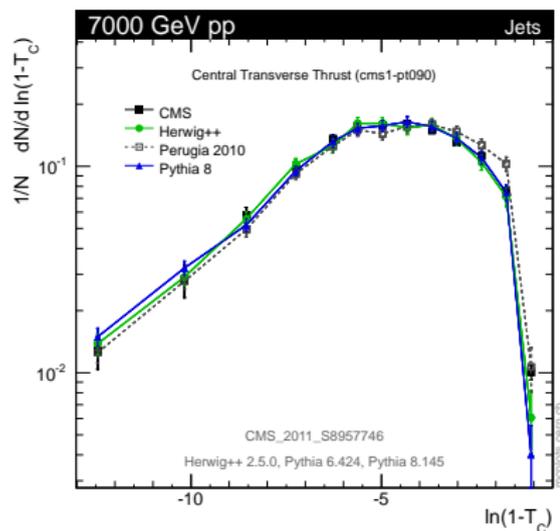
How well does it work?

$p_{\perp}(Z^0) \rightarrow$ intrinsic k_{\perp} (LHC 7 TeV).

See also in context of matching/marging.

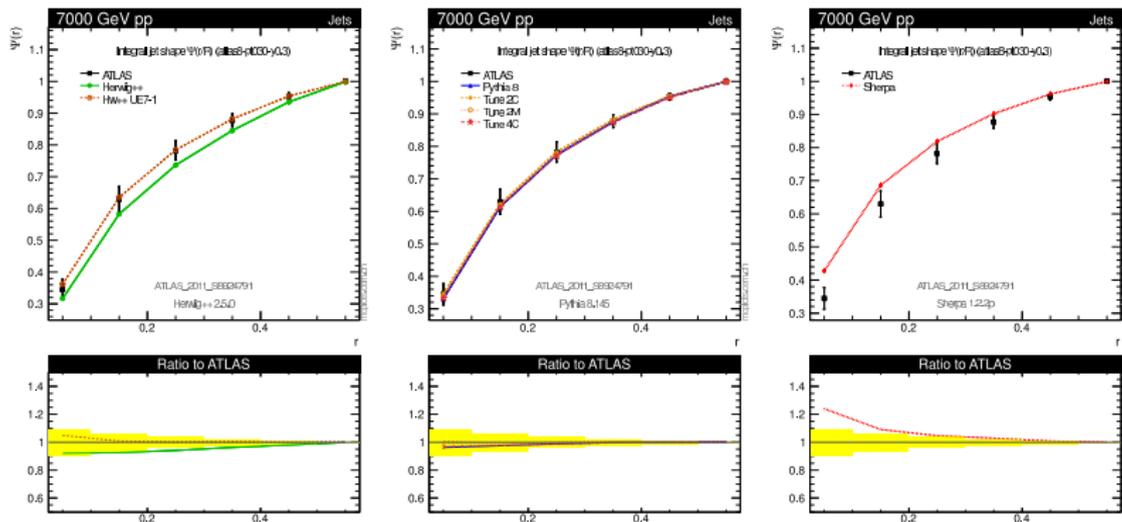


Transverse thrust



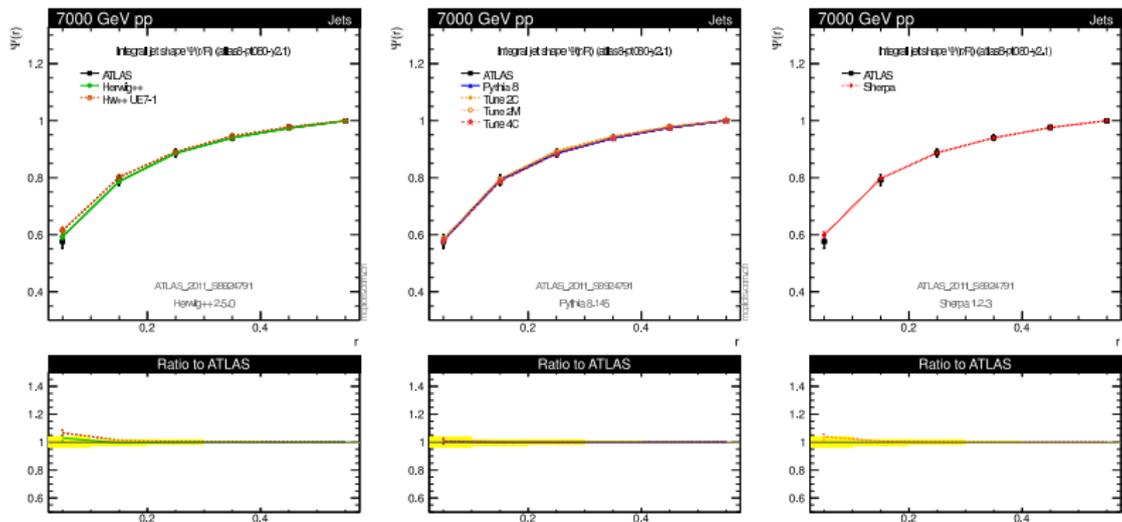
Integral jet shapes

not too hard, central ($30 < p_T/\text{GeV} < 40; 0 < |y| < 0.3$)



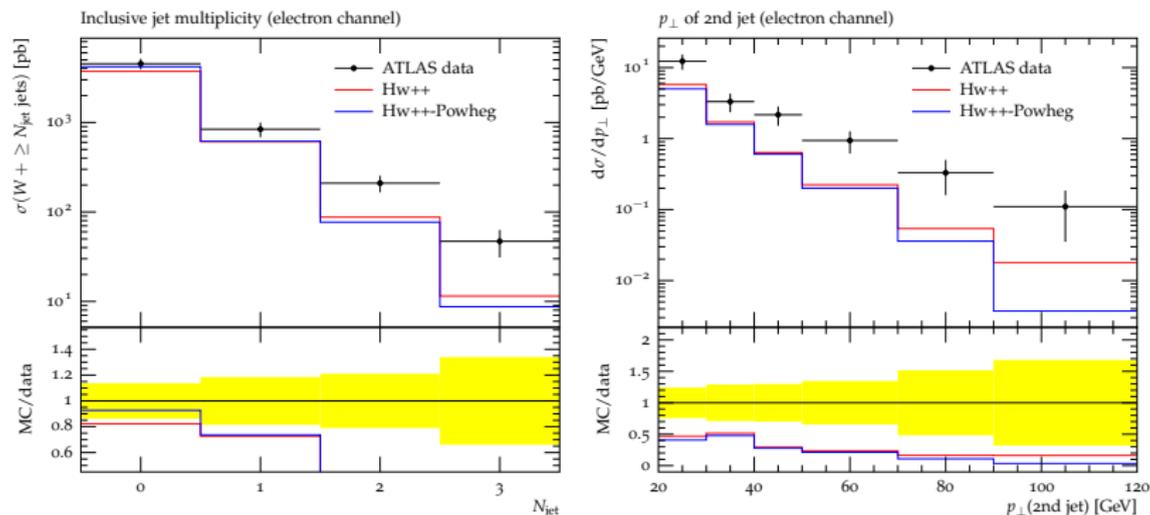
Integral jet shapes

harder, more forward ($80 < p_T/\text{GeV} < 110; 1.2 < |y| < 2.1$)



Limits of parton showers

$W + \text{jets}$, LHC 7 TeV.



Higher jets not covered by parton shower only \rightarrow merging.