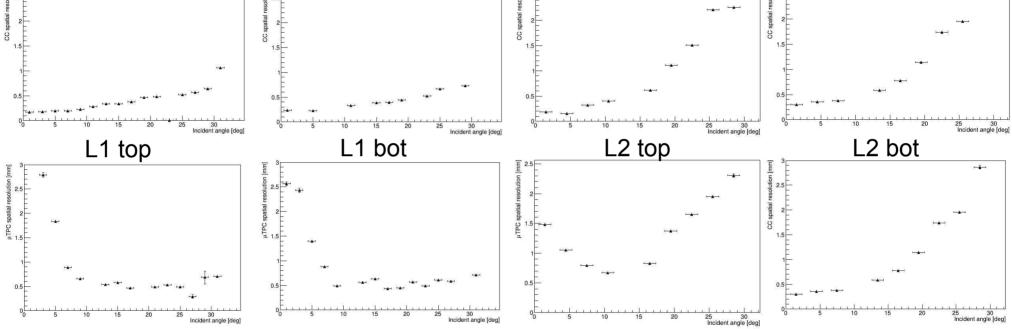
ANALYSIS UPDATES

Cosmic data, March. 2020 - Ferrara

R.Farinelli

CC & µTPC: incident angle





- •L1 performs systematically better than L2, expecially if the beam incident angle is larger than 15°.
- μ TPC behavior in L1 is flat between 10° and 30° while in L2 doesn't. Might it be due to the tracking system?
- •No significant difference are evaluated between the top and the bottom part of L1 and L2.

What are the systematic to include in these results and why L2 is always worsen than L1?





TOY MC - COSMIC SETUP

This is a simulation of the cosmic setup for L1 and L2. The purpose of this toy is to evaluate the contribution of the tracking system as a function of the incident angle.

Cosmic particle with a path orthogonal to the ground plane are shooted on L1 and L2. A shift is used to hit the setup with tracks having an incident angle w.r.t. the cylinder surface between 0 and 90° on L1 and between 0 and 40° on L2.

Up to now, only positive X positions are considered.

Position is evaluated with CC alogrithm only.

Cosmic data, March. 2020 - Ferrara



TOY MC - Cosmic setup

1. Generate a random number from 0° to 40° as the L2 incident angle

2. Evaluate the L1 incident angle from geometry: L1_angle = asin (R2/R1 * sin (L2_angle))

3. Smear the position on the L1up, L1down, L2up, L2down as a function of the expected performance from experimental data with planar triple-GEM

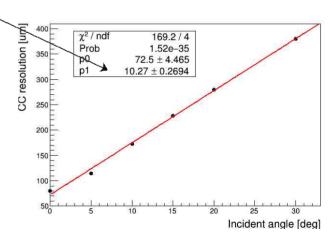
4. Choose 3 points and fit them with a line, then measure the residual of the forth point

5. Repeat action 1-4 for a large sample

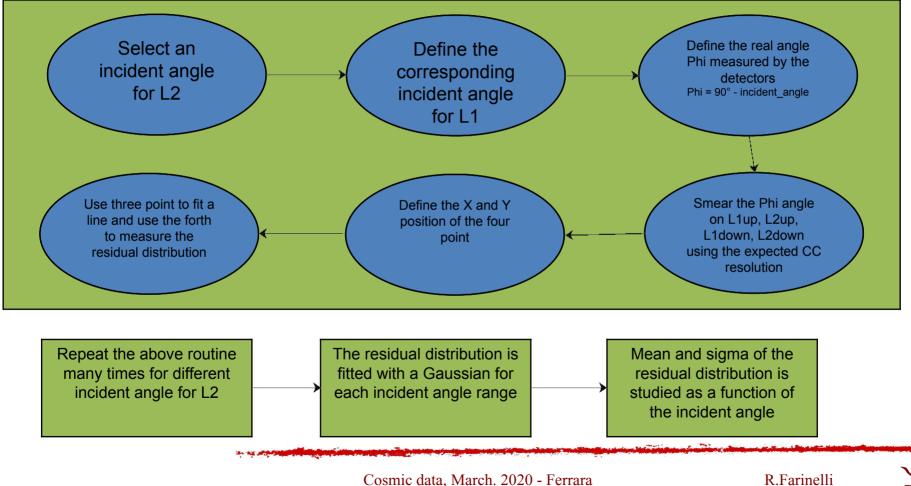
X

y

- 6. Analyze the residual distrubution with a Gaussian fit
- 7. Plot the mean and the sigma as a function of the incident angle on the test detector
- 8. Evaluate the tracking contribution with a squared substraction of the expected resolution for each incident angle



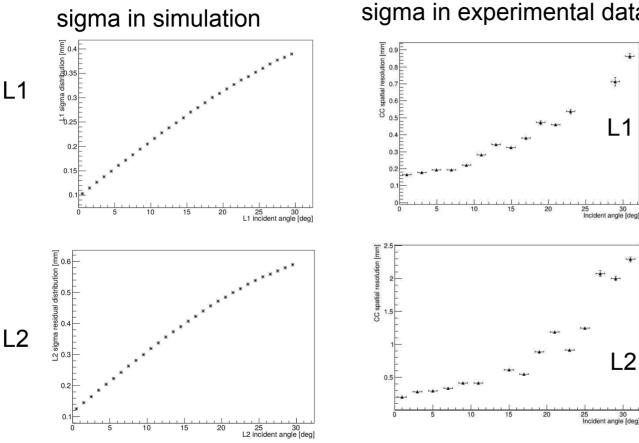
TOY MC - Cosmic setup



R.Farinelli



TOY MC - Sigma simulation vs experimental



sigma in experimental data

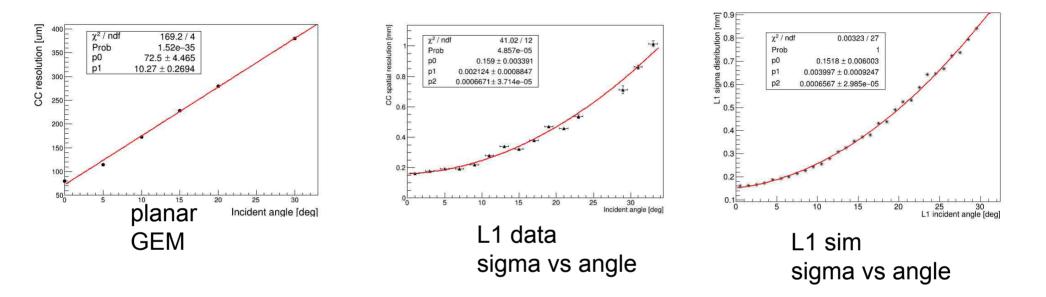
1. The sigma measured contains the contribution of the tracking system and the detector itself

Simulation underestimate the 2. experimantal data, expecially on L2 3. shape of the simulated The and experimental data is different too

The resolution of the planar GEM is not 4.

the best one to simulate the CGEM

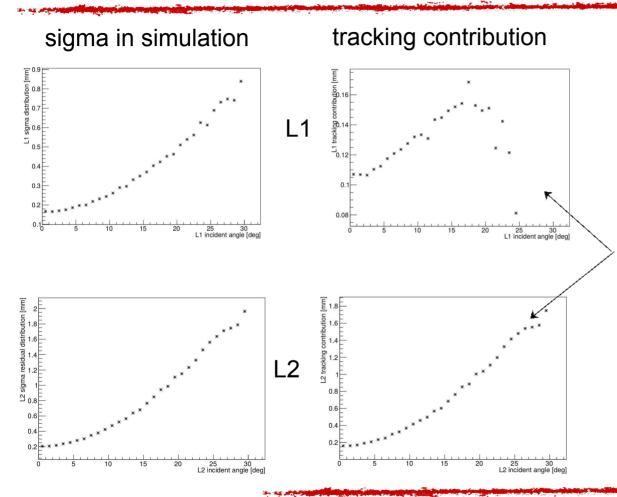
new TOY MC - CC resolution of CGEM



- 1. The performance of the CGEM is measured by real data is very different from the one measured by the cosmic setup
- 2. I try a different function to reproduce the resolution of the CC: a parabola instead of a line
- 3. The parabola for the simulation of the L1 resolution has been properly tuned to match the data of L1
- 4. The same values are used also for L2 and the simulated result (next slide) are in agreement with the experiement (previous slide)

Cosmic data, March. 2020 - Ferrara

new TOY MC - Tracking system contribution



1. Using the sigma value as a function of the incident angle measured in the previous slide, it is possible to measure the contribution of the tracking system.

2. The sigma of the Guassian is determined by the intrinsic resolution of the detector and the contribution of the tracking system, then the effect of the tracking system can be extracted with a squared subtraction:

sqrt((residual sigma)² - (CC planar resolution)²)

3. The result has to be understood. The contribution of the tracking system seems to be very small on L1. It is important to understand what is the role of the mean value of the residual and how it could worsen the tracking system.

TOY MC - Conclusion

- 1. A contribution of the tracking system of about 100-150 µm for L1 and 200-1600 µm for L2 has been evaluated as a function of the incident angle
- 2. The CC resolution from the planar GEM is not enough to describe the data
- 3. A parabolic shape of the CC resolution vs incident angle seems to describe better the data
- 4. The resolution on L2 is systematically worsen thant L1

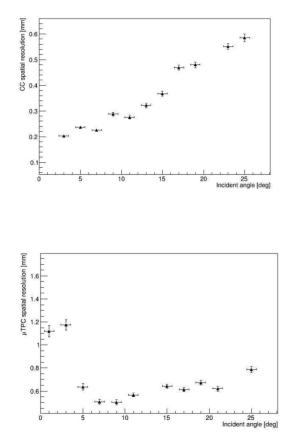


Ok, we can go back to the L1 performance

BES-ITA 1-2 Apr. 2020 - Vidyo



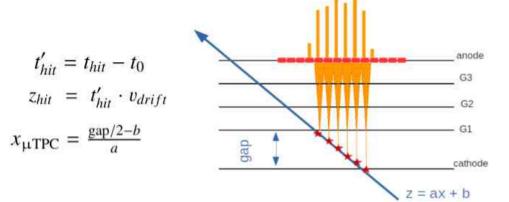
CC and TPC vs angle (L1top)

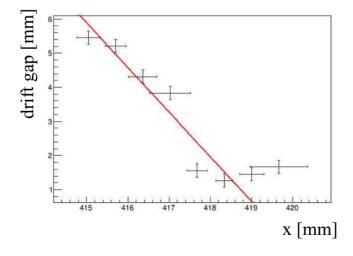


Until μ TPC will to be better than CC, every merge studies is useless and they will not be implemented in CGEMBOSS

--> Let's concentrate to the μ TPC

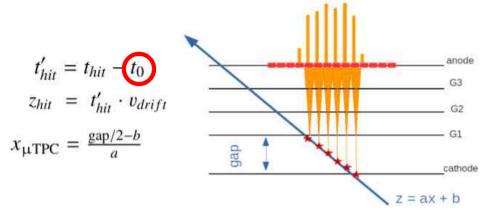
µTPC present status in CGEM-IT



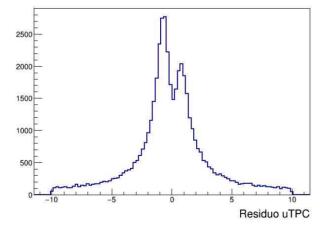


 μ TPC events are recostructed with error bars. Two methods are used in CGEMBOSS but in the present situation no differences show up.

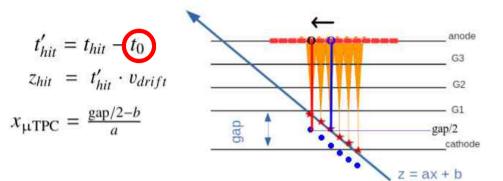
µTPC present status: time reference

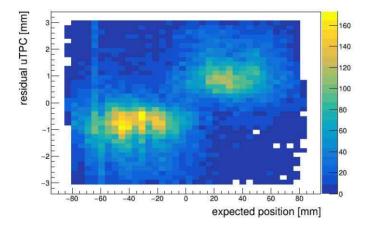


Double peak is observed in the μ TPC residual distribution. This problem is due to the **time reference**.



µTPC present status: time reference





Double peak is observed in the μ TPC residual distribution. This problem is due to the **time reference**.

Positive position shift on one side and negative position to the opposite side.

A wrong time reference can shift the μ TPC points (from red stards to blue dots) then a shift in the μ TPC positions.

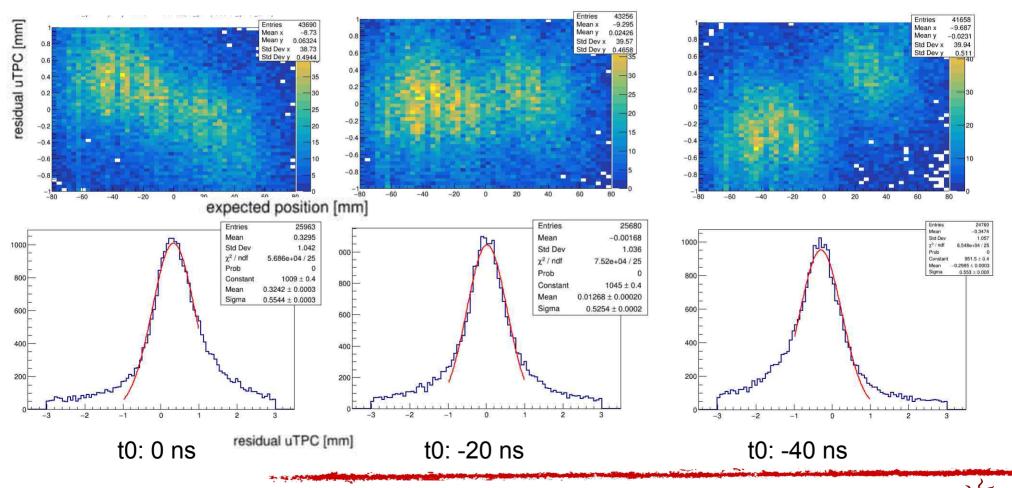
Positive angle shift on one side, negative angle to the other one.

Let's select only negative expected position

BES-ITA 1-2 Apr. 2020 - Vidyo



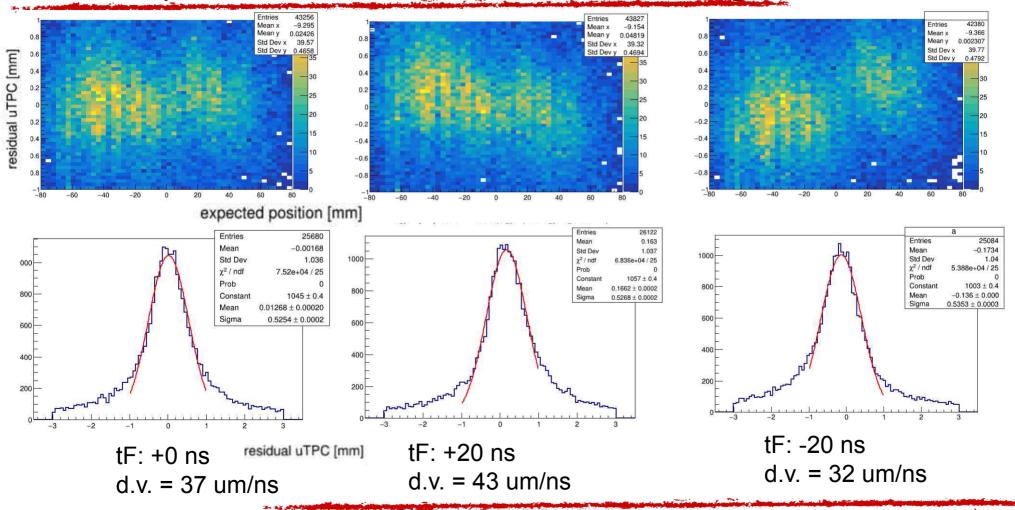
µTPC: time reference studies - L1top



BES-ITA 1-2 Apr. 2020 - Vidyo

R.Farinelli

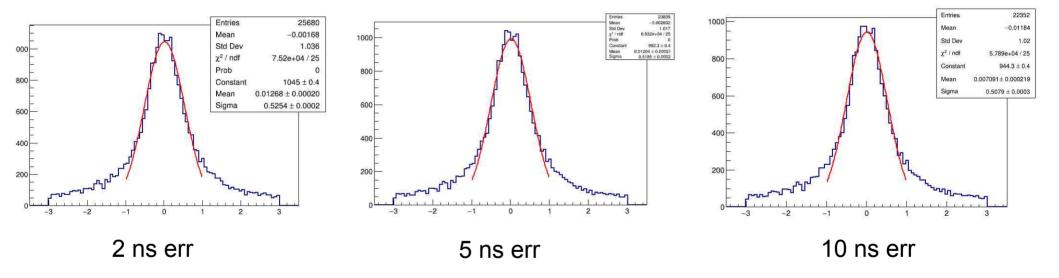
µTPC: time width studies - L1top



BES-ITA 1-2 Apr. 2020 - Vidyo

R.Farinelli 29 17

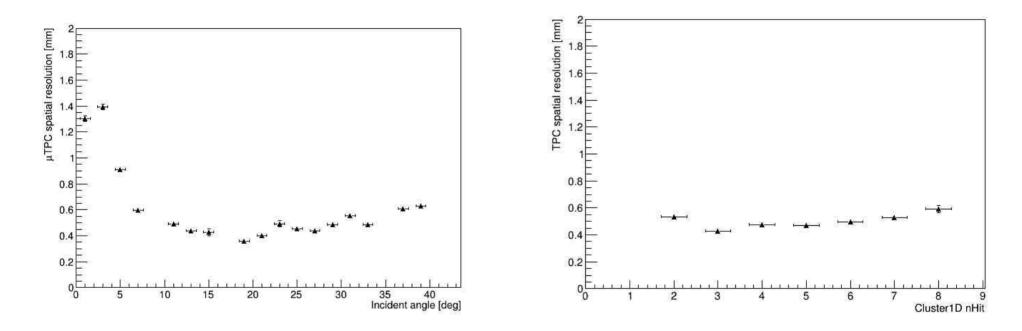
µTPC: time error studies - L1top





R.Farinelli

µTPC studies



This is the behavior as a function of the incident angle and cluster size

BES-ITA 1-2 Apr. 2020 - Vidyo

R.Farinelli 39 19



- 1. Optimization of the μ TPC variables are needed
- 2. Preliminary studies on the time reference shows important effects
- 3. Minor effects are observed in the drift velocity scan and in the error size scan
- 4. More studies will be needed because the actual results could not improve the µTPC below 500µm



LUT results

Cosmic data, March. 2020 - Ferrara



LUT: results

The variables list is:

GENERAL SETTING runs high voltage values energy mode

GEOMETRY ROC id, FEB id, TIGER id strip, side, layer

CALIBRATION qdc slope, constant and saturation value

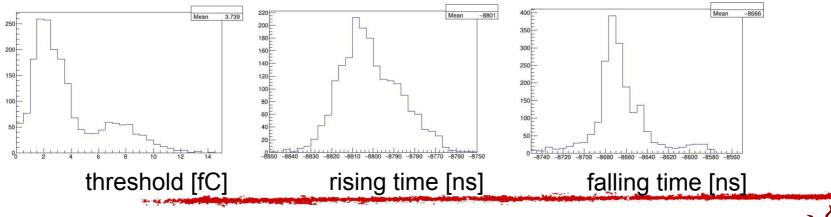
THRESHOLD

voltage thr. E and T branches voltage baseline E and T branches charge cut (fC) E and T branches effective charge cut (fC) on the channel

NOISE (out of time) rate (Hz) , mean charge (fC)

SIGNAL (in-time) rate (Hz), mean chare (fC), max charge (fC) leading and falling time (ns) quality

--> New package to extract the variables: ReadCosmicRayData-00-00-21

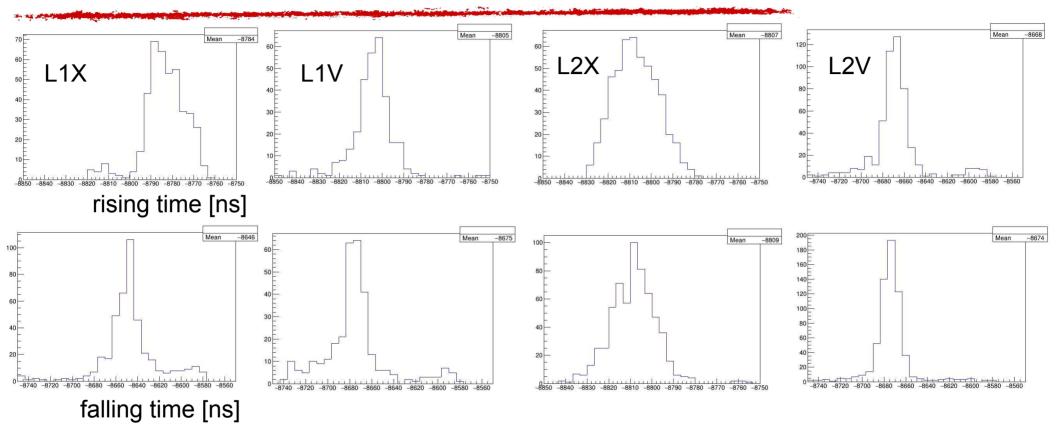


Cosmic data, March. 2020 - Ferrara

R.Farinelli



LUT: Time reference



Differences of several tens of ns are present between on L1 and L2, but also between channels of the same layer



LUT - Conclusion

- 1. Information channel by channel can be used to improve the resolution of the detector such us the time-walk effect
- 2. Moreover other information such us time reference could be used for futher studies
- 3. Is the calibration team available to develope the corrections channel by channel?

All - Conclusion

- 1. L1 detector is the one with the smaller contribution of the tracking system
- 2. uTPC optimization is feasible but now we need to implement it chip by chip or channel by channel
- 3. LUT can provide threshold value and it could be used to perform the time-walk correction

