# Energy reconstruction of cosmic ray for LHAASO-KM2A

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## Purpose

- The all-particle energy spectrum exhibits two interesting structures from 10<sup>14</sup> to 10<sup>18</sup> eV, the "knee" and the "second knee". An explanation of these features is thought to be an important step in understanding the origin of the high-energy particles.
- A finite energy attained during the acceleration process leakage from the Galaxy
- Cosmic-ray compositions and hadronic interaction models
- ◆ A combination of electron and muon sizes is insensitive to the type of primary particle.

Quarter array simulation data :

QGSII\_Gheisha: Proton He CNO MgAlSi Fe, Theta:0-11° Slope: -2

normalizing to Gaisser energy spectrum





log(E<sub>0</sub>/GeV)





## **Energy reconstruction method 1**

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 $N_{\mu}$ 

$$N_{\mu} = \left(\frac{E_0}{E_c^{\pi}}\right)^{\beta} A^{1-\beta} \approx 1.69 \times 10^4 \cdot A^{0.10} \left(\frac{E_0}{1 \text{ PeV}}\right)^{0.90}$$

$$N_e \approx 5.95 \times 10^5 \cdot A^{-0.046} \left(\frac{E_0}{1 \text{ PeV}}\right)^{1.046}$$

$$= \frac{\ln E_0 / E_c^{\pi}}{\ln \frac{3}{2} N_{ch}} = 0.85 \lg \left(\frac{E_0}{E_c^{\pi}}\right)$$

$$N_{\mu} = n_c \ln N_{ch} = \beta \ln \left(\frac{E_0}{E_c^{\pi}}\right), \quad \text{with} \quad \beta = \frac{\ln N_{ch}}{\ln \frac{3}{2} N_{ch}} \approx 0.85$$

$$= A(E_0 / (AE_c^{\pi}))^{\beta}$$

### **Energy reconstruction method 1**

 $N_{em} = N_e^{0.68} * N_{\mu}^{0.32}$ 

 $\log 10(E) = a + b \cdot \log 10(N_{em})$ 





 $(E_{rec} - E_{true})/E_{true}$ 





#### **Energy reconstruction method 2**





 $N_{em} = \rho_{50}^{0.68} * N_{\mu}^{0.32}$  $\log 10(E) = a + b * \log 10(N_{em})$ 







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## **Energy reconstruction method 3**



A new combination of electron and muon sizes (method 3) which will be shown to be insensitive to the type of primary particle.

 ✓ The cosmic ray all-particle energy spectrum between 10<sup>14</sup> to 10<sup>17</sup> eV using LHAASO-KM2A full array.







