

Observation of the large-scale anisotropy of cosmic rays with LHAASO

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Introduction

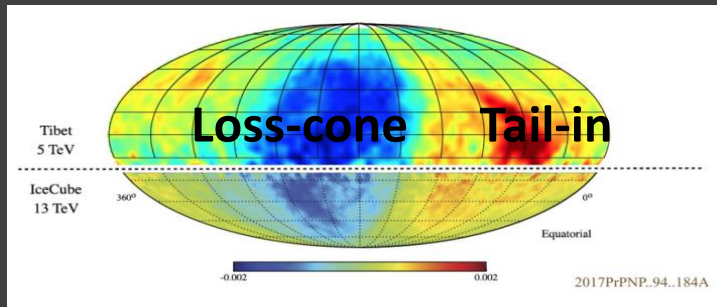
Large-scale anisotropy with KM2A

Conclusion

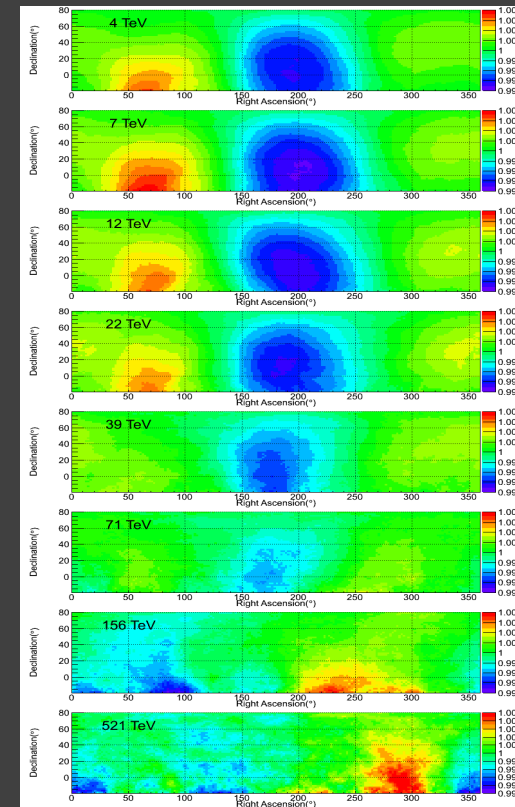
Introduction

□ Anisotropy

- a probe to study the origin and propagation of cosmic rays
- the intensity and phase are obviously in energy dependence



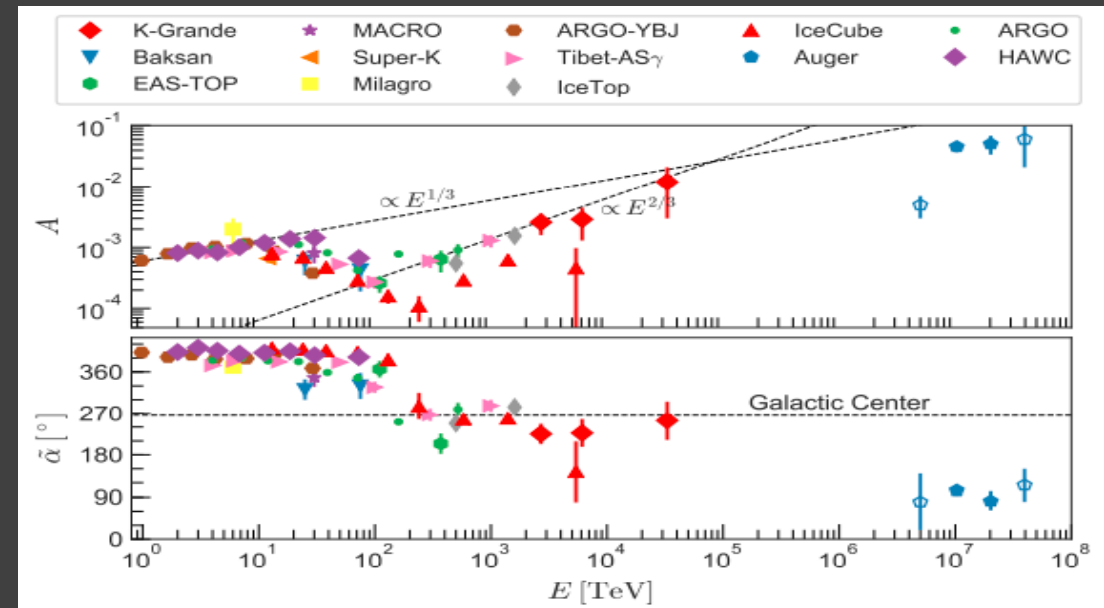
ARGO



Bartoli et al. 2018

□ Puzzles

The irregularity of energy dependence of amplitude and phase —Why ?



JuliaBecker Tjus et al., Physics Reports, 872

Large-scale anisotropy with KM2A

□ simulation

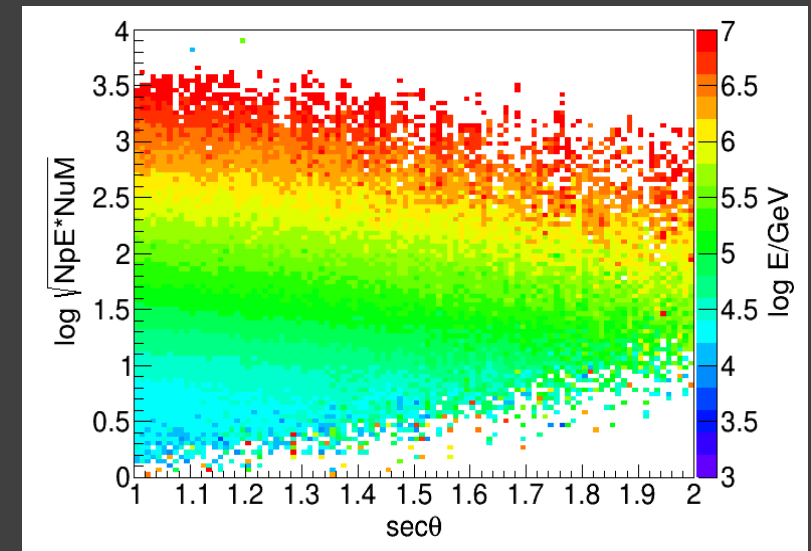
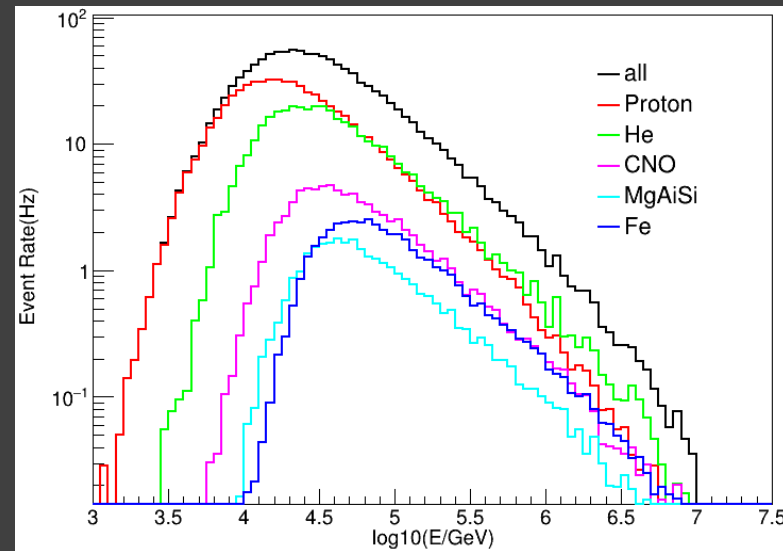
QGSJETII-04 & Gheisha

1TeV-10PeV

Zenith 0-70°

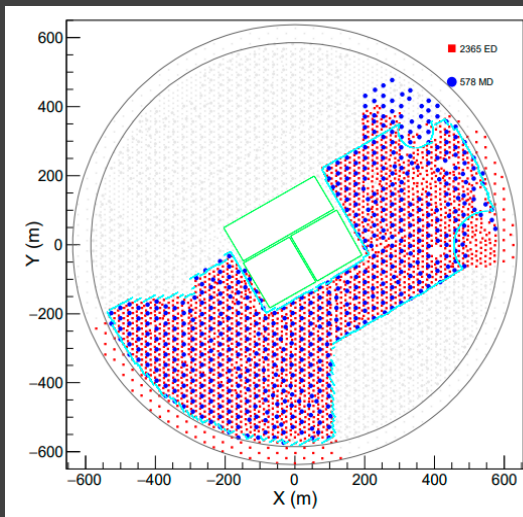
$\sim 7.8 \times 10^7$ events

Gaisser model



$$\log E = f(\log \sqrt{N_p E * NuM}, \theta)$$

- full year of 2021
- core locate in the half array of KM2A

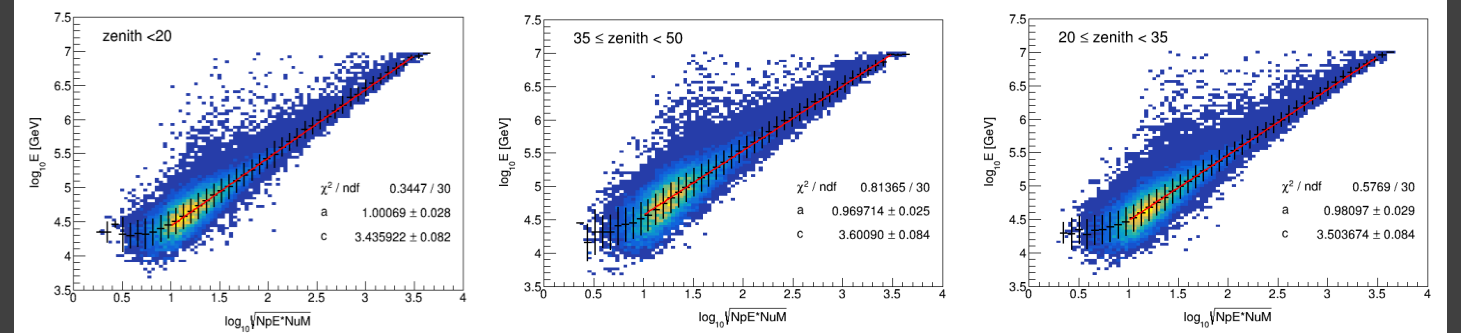
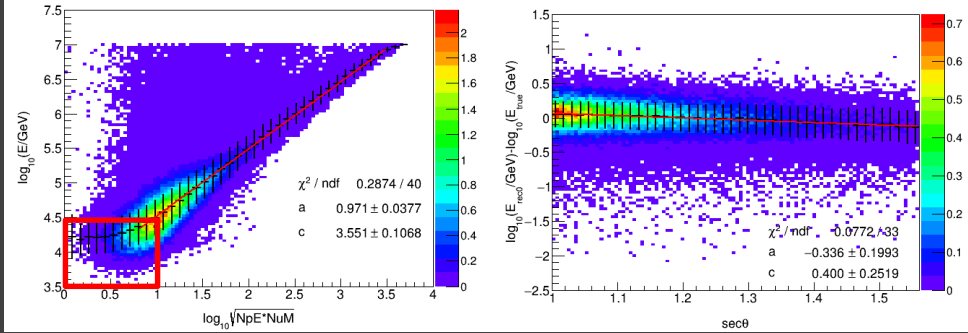


parameter	sample1	Sample2
zenith	<50	<50
NtrigE	>=20	>=20
NfiltE	>=20	>=20
NpE2 (r=40-100, dt:-50,100)	>0	>15
NuM3 (r=40-200, dt:-50,100)	>0	>0
dr (芯位距离边界距离)	>0	>0
NpE1 (r<rwind+50, dt:-50,100)	-----	NpE1/NpE2>2

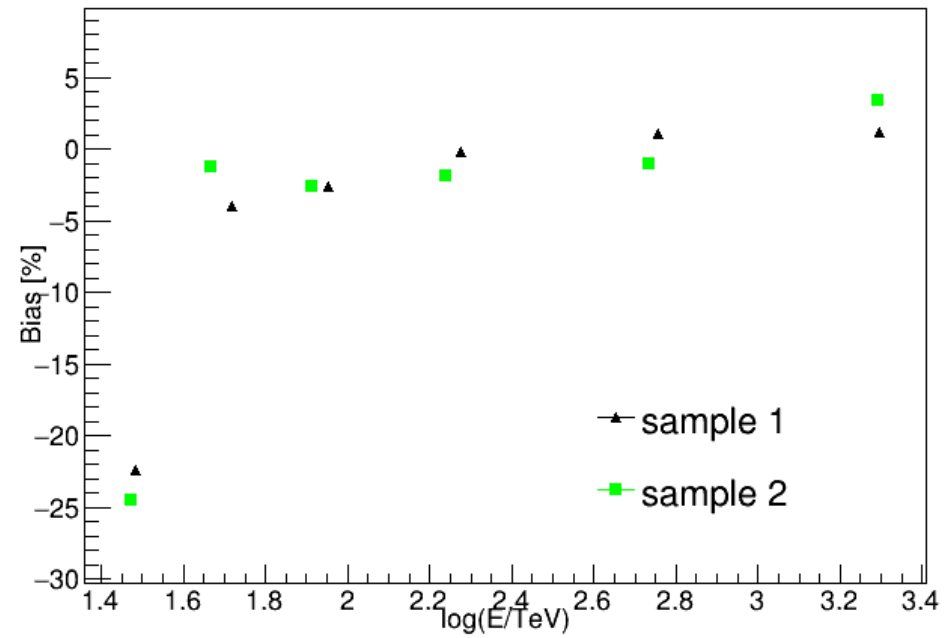
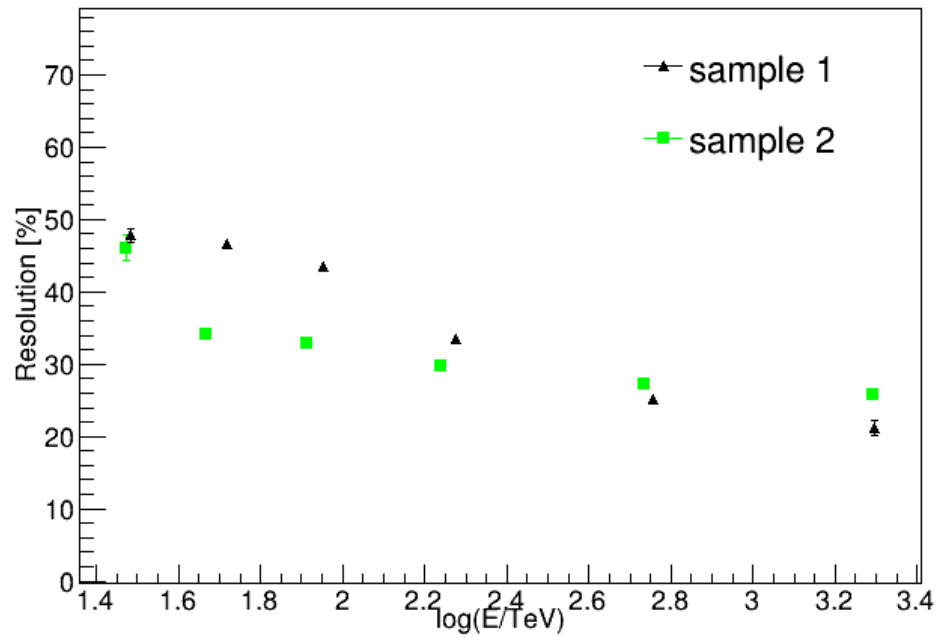
Energy estimation

Sample 1 $\log E = f(\log \sqrt{NpE * NuM}, \theta)$

Sample 2 $\log E = f(\log \sqrt{NpE * NuM})$

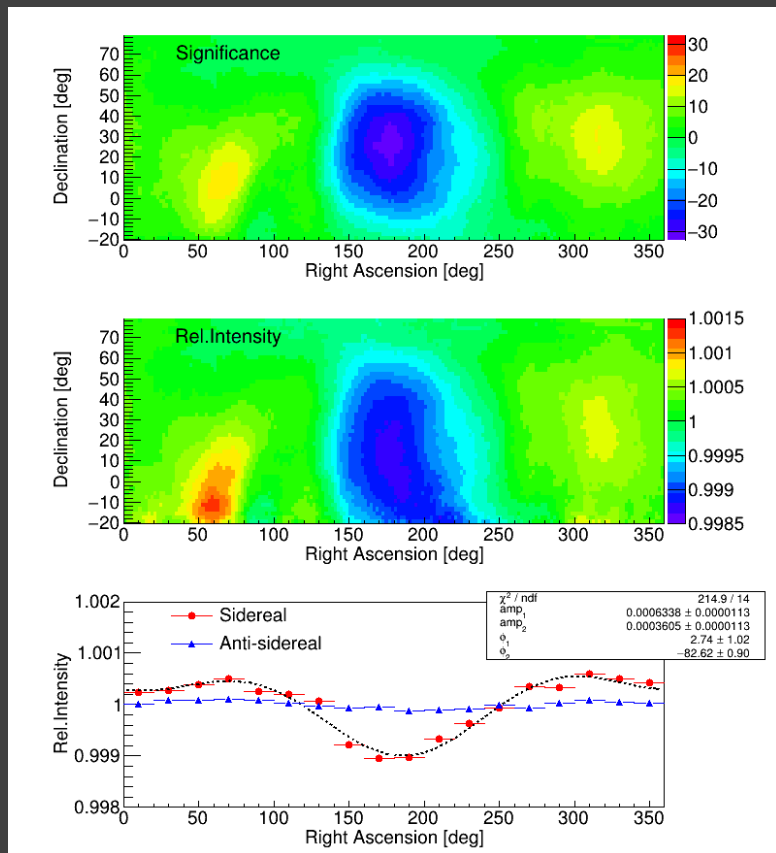


logErec	Sample 1 77.5			Sample 2 (3 zenith bins) 96.7		
	mean (TeV)	Resolution (%)	Bias (%)	mean (TeV)	Resolution (%)	Bias (%)
3.5~4.5	30.43	47.8±1.0	-22.4	29.60	46.1±1.8	-24.5
4.5~4.75	52.24	46.6±0.8	-4	46.54	34.2±0.3	-1.2
4.75~5.0	89.89	43.6±0.8	-2.6	81.56	32.8±0.3	-2.6
5.0~5.5	188.92	33.4±0.5	-0.2	173.23	29.8±0.2	-1.9
5.5~6	570.26	25.1±0.6	1.1	542.61	27.3±0.3	-1
>=6	1983.24	21.2±1.1	1.2	1967.07	25.87±0.6	3.4

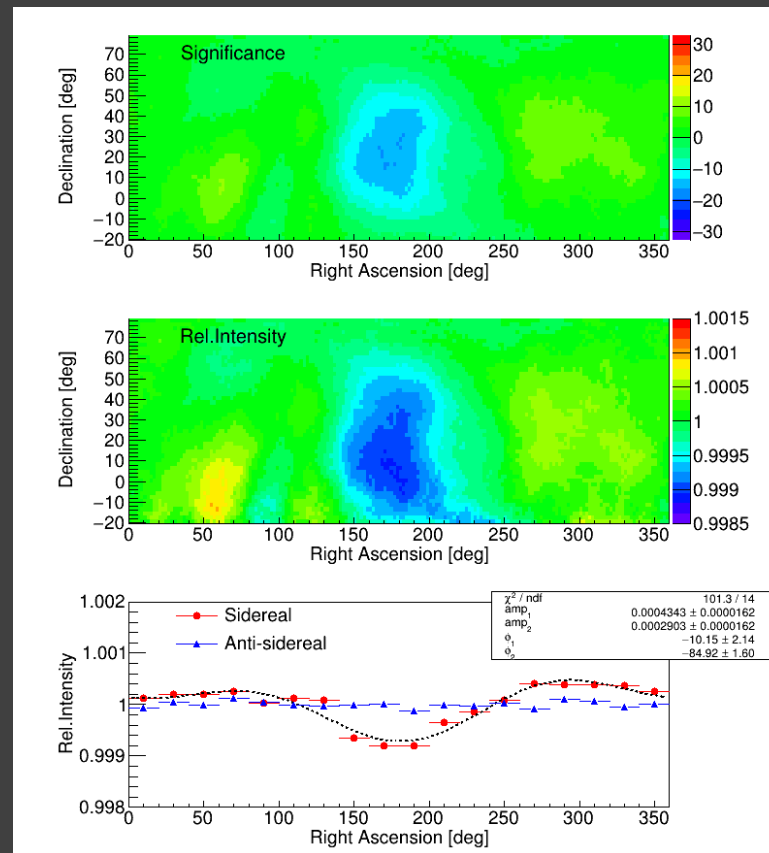


Anisotropy in sidereal time

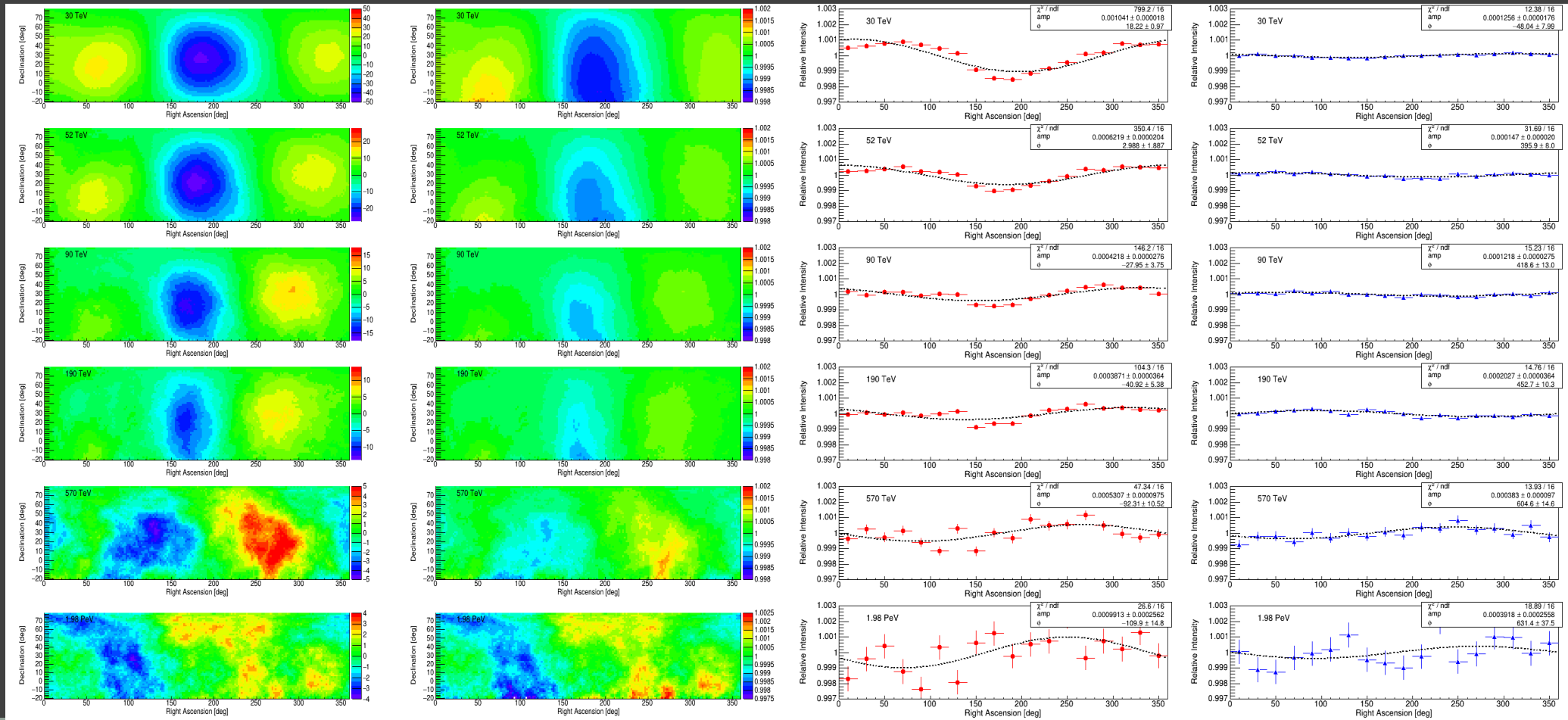
sample 1 $\sim 1.6 \times 10^{10}$ events



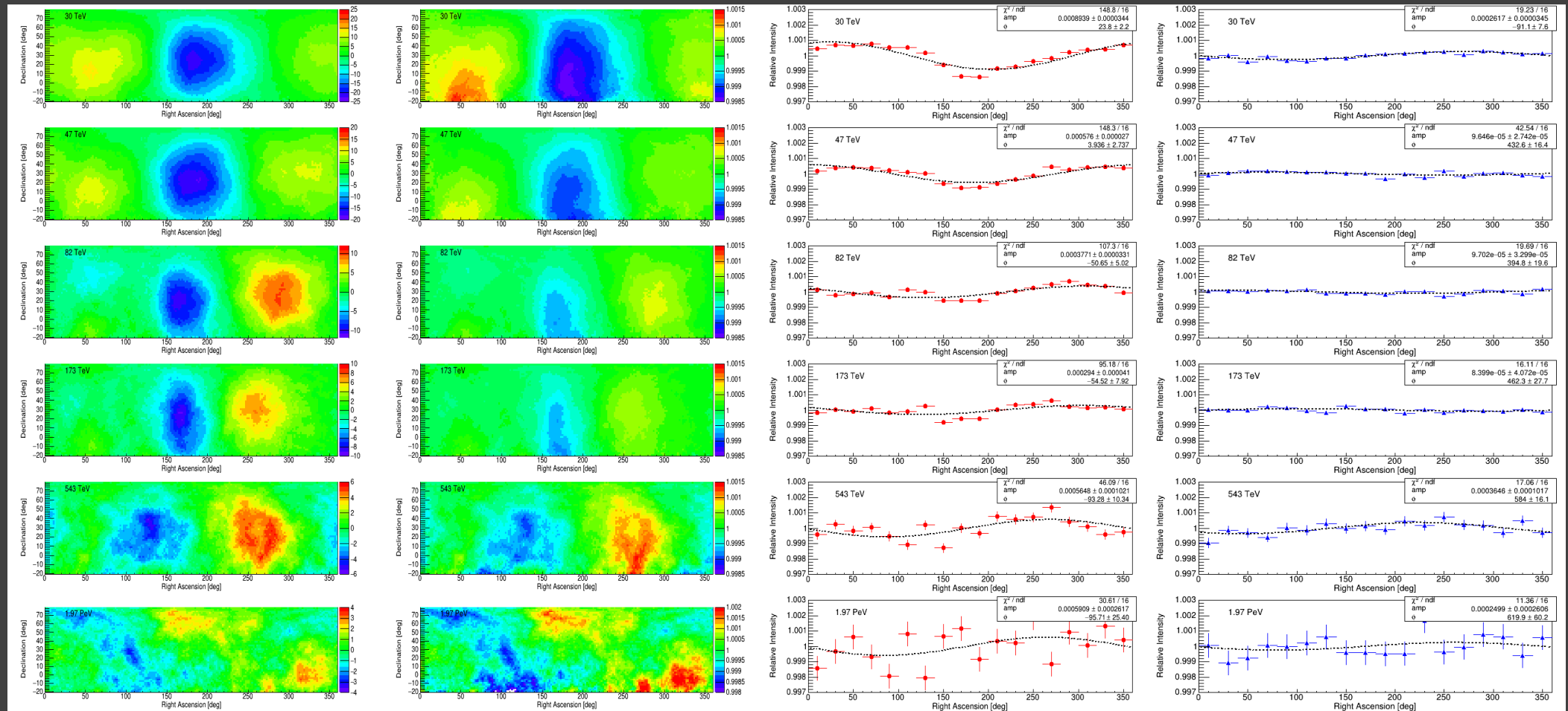
sample 2 $\sim 7.6 \times 10^9$ events



Evolution with energy-sample 1



Evolution with energy-sample 2



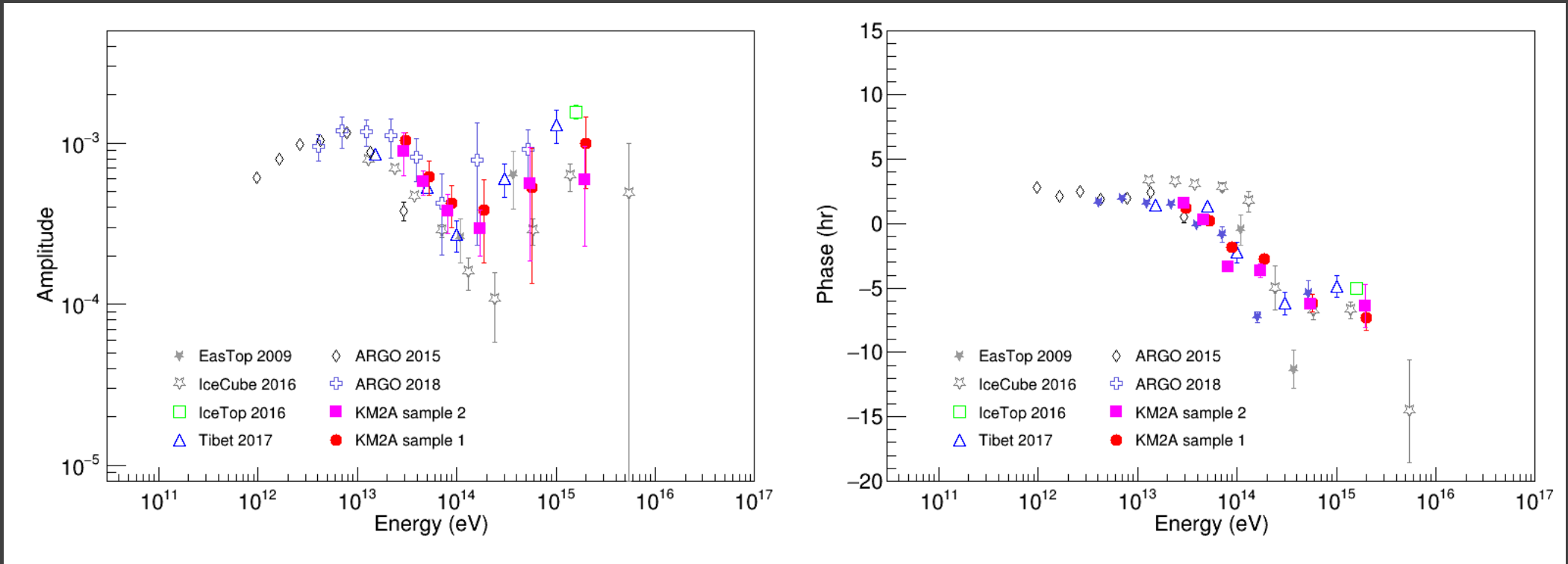
Sample 1

Bin	log(Erec) (GeV)	Mean Energy (TeV)	Events (10 ⁸)	Amp (10 ⁻⁴)	Φ (deg)	Amp_Anti (10 ⁻⁴)
1	3.5~4.5	30.43	64.66	10.41±0.18	18.22±0.97	1.256
2	4.5~4.75	52.24	47.75	6.22±0.20	2.99±1.89	1.47
3	4.75~5.0	89.89	26.24	4.22±0.28	-27.95±3.75	1.218
4	5.0~5.5	189.92	15.10	3.87±0.36	-40.92±5.38	2.027
5	5.5~6	570.26	2.10	5.31±0.98	-92.31±10.52	3.83
6	>=6	1983.24	0.30	9.91±2.56	-109.9±14.8	3.918

Sample 2

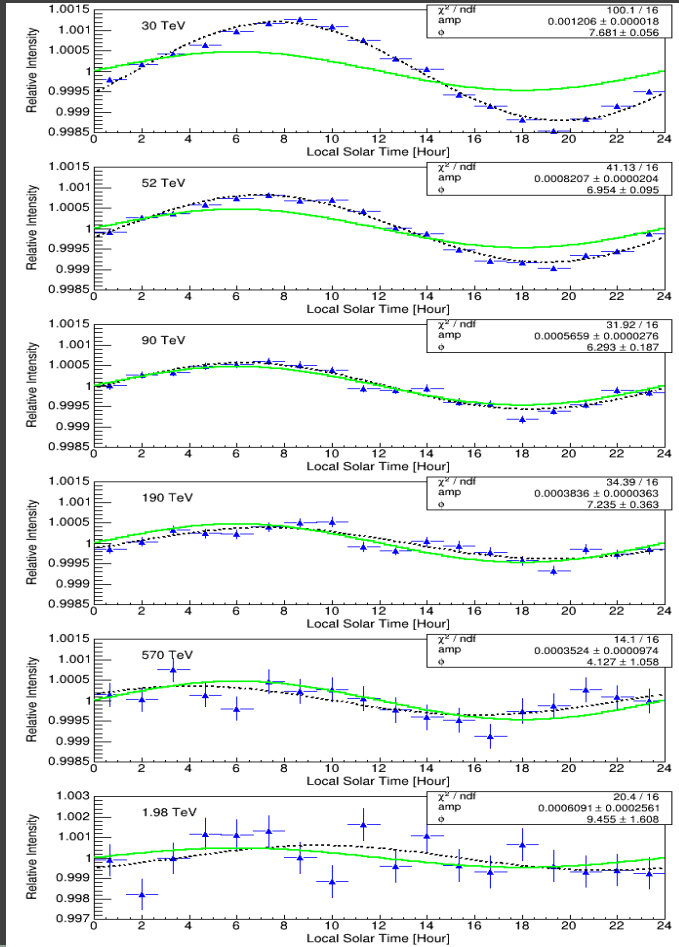
Bin	log(Erec) (GeV)	Mean Energy (TeV)	Events (10 ⁸)	Amp (10 ⁻⁴)	Φ (deg)	Amp_Anti (10 ⁻⁴)
1	3.5~4.5	29.60	16.81	8.94±0.34	23.8±2.2	2.62
2	4.5~4.75	46.54	26.48	5.76±0.27	3.94±2.74	0.96
3	4.75~5.0	81.56	18.27	3.77±0.33	-50.65±5.02	0.97
4	5.0~5.5	173.23	12.07	2.94±0.41	-54.52±7.92	0.84
5	5.5~6	542.61	1.92	5.65±1.02	-93.28±10.34	3.65
6	>=6	1967.07	0.29	5.91±2.62	-95.71±25.40	2.50

Evolution with energy - amplitude & phase

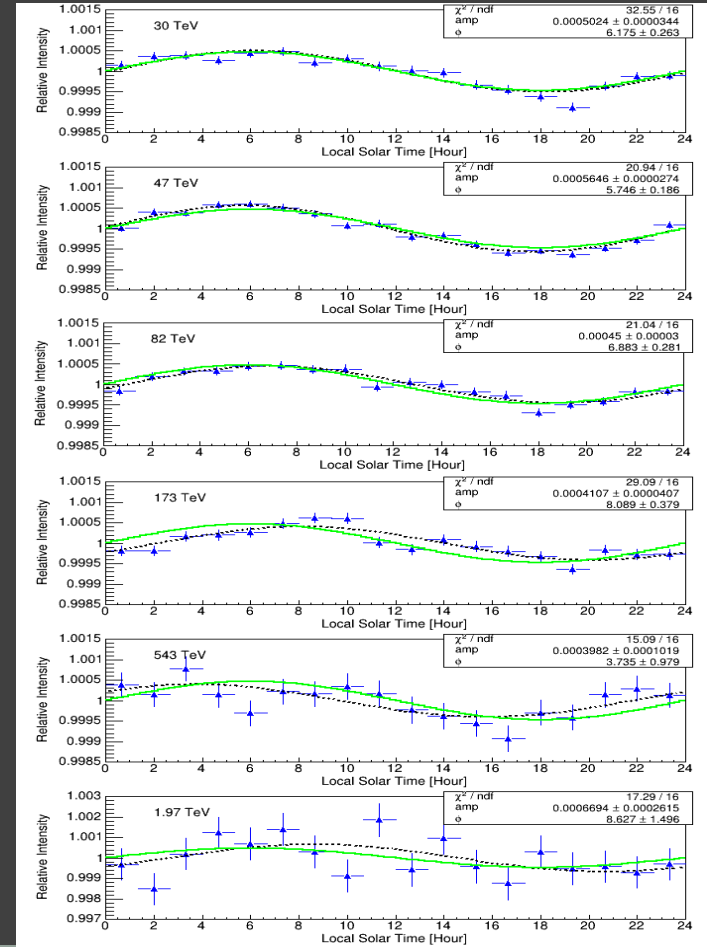


Solar

Sample 1



Sample 2



Conclusion

- Consistent results of large-scale anisotropy in sidereal with the others were observed.
- Relative intensity from stricter-cut sample more approach the expected CG anisotropy.
- Project :
 - more works for the CG effect
 - extend the high energy range
 - anisotropy of heavy and light component of cosmic rays

**Thanks for your
attention !**

Supplement

