



The KM3NeT infrastructure: status, perspectives and preliminary results

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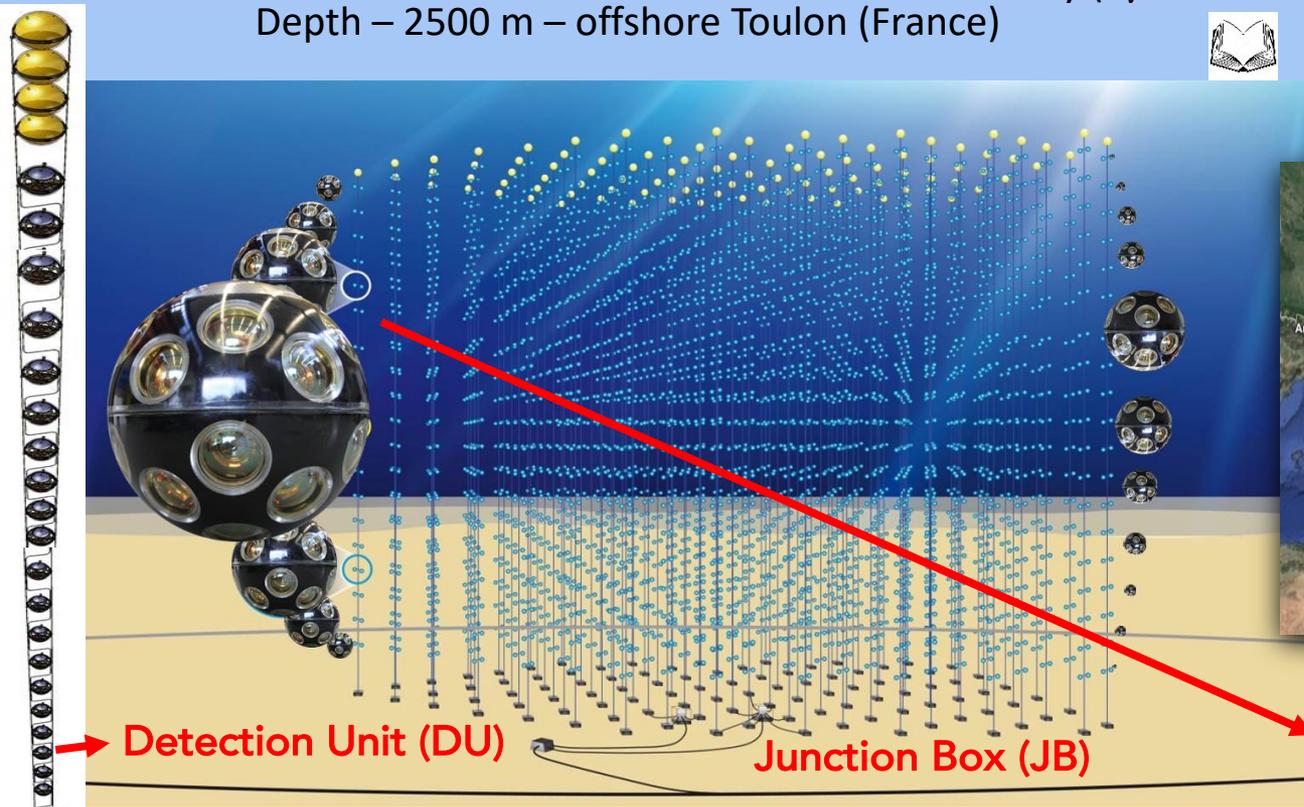
The KM3NeT detectors

KM3NeT is a research infrastructure underwater in the Mediterranean Sea that will host a network of neutrino telescopes:

- KM3NeT/ARCA** (Astroparticle Research with Cosmics in the Abyss)
 - discovery and observation of high energy cosmic neutrino sources ($E_\nu \sim \text{GeV-PeV}$)
 - Depth – 3500 m – offshore Sicily (Italy)
- KM3NeT/ORCA** (Oscillation Research with Cosmics in the Abyss)
 - determination of the neutrino mass hierarchy ($E_\nu \sim \text{MeV - GeV}$)
 - Depth – 2500 m – offshore Toulon (France)



**KM3NeT 2.0 Letter of Intent:
J.Phys. G43 (2016) 084001**



Detection Unit (DU)

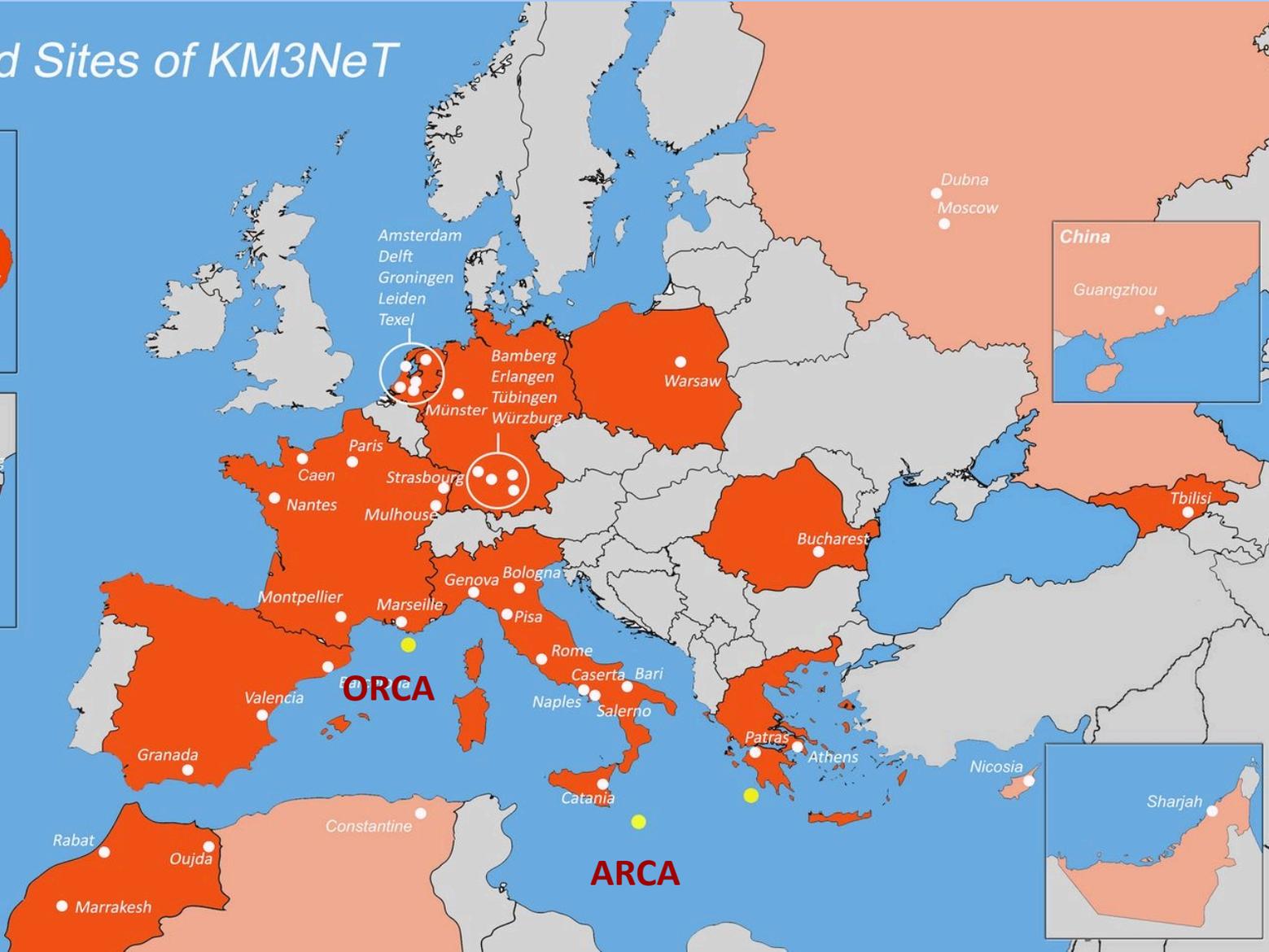
Junction Box (JB)

Optical sensor (DOM)
31 PMTs of 3 inches

The collaboration

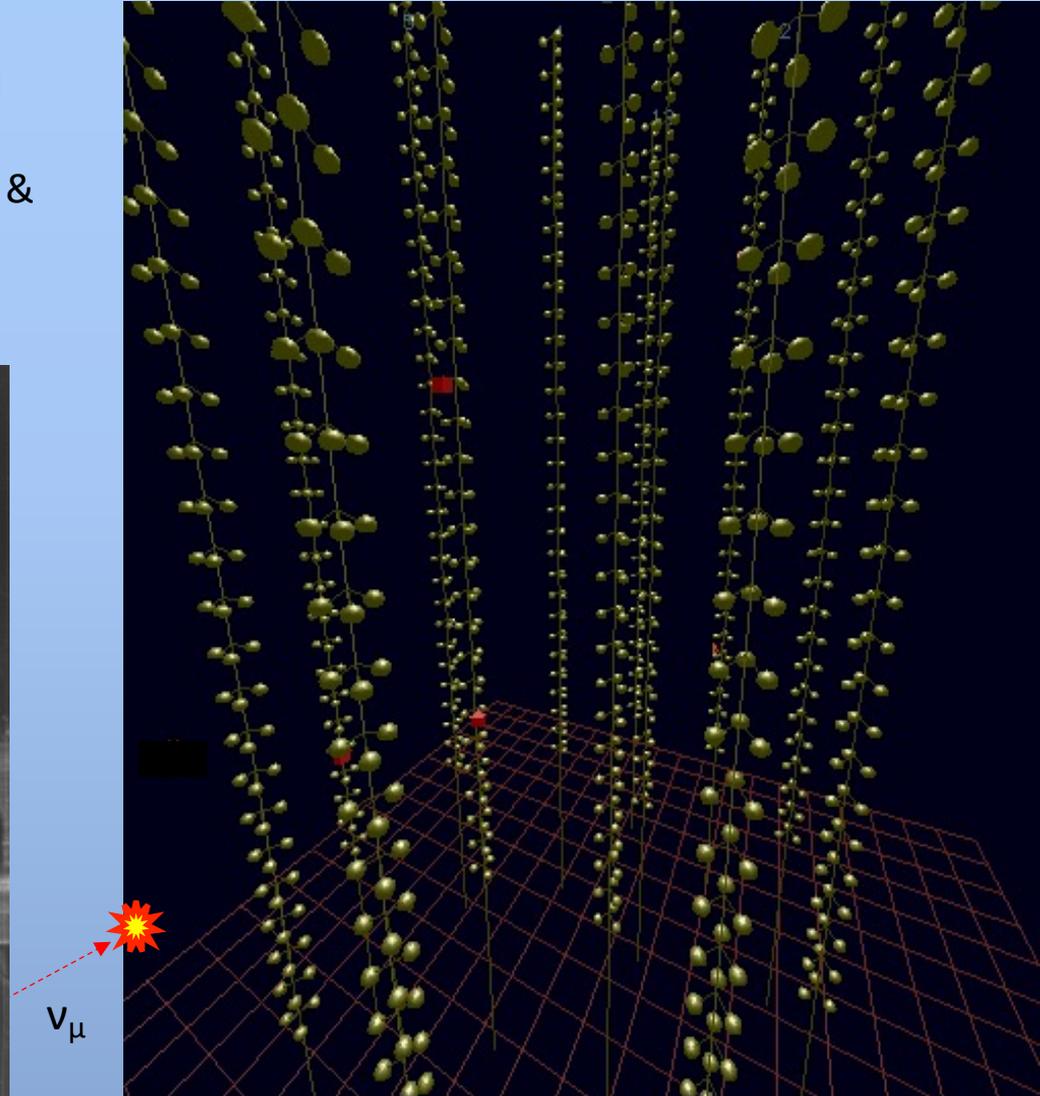


Cities and Sites of KM3NeT



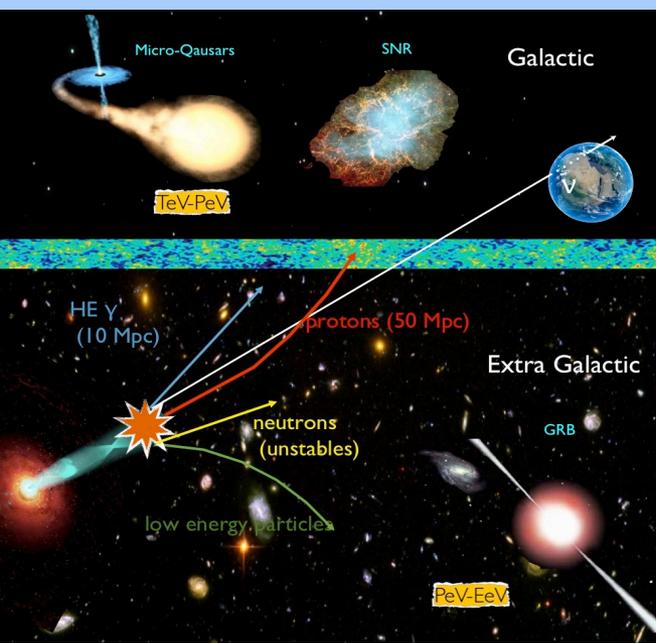
The concept of Cherenkov neutrino telescopes

- 3D lattice of photomultipliers collecting Cherenkov photons due to relativistic charged particles from ν interactions
- ν direction reconstructed using photon time & position





The scientific case

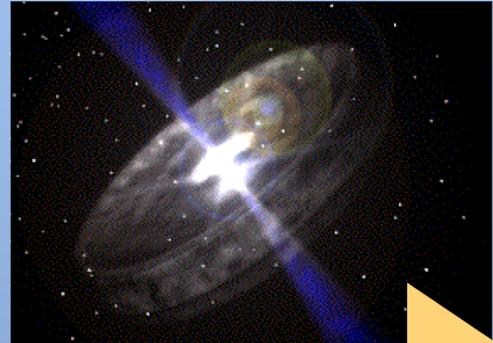
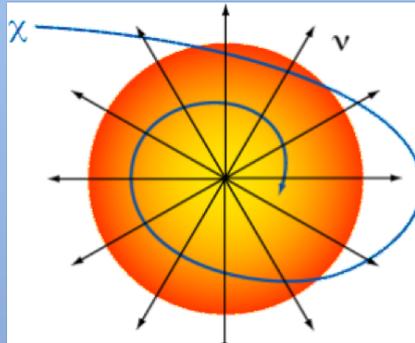
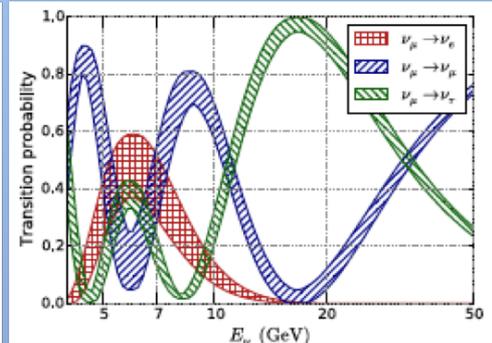
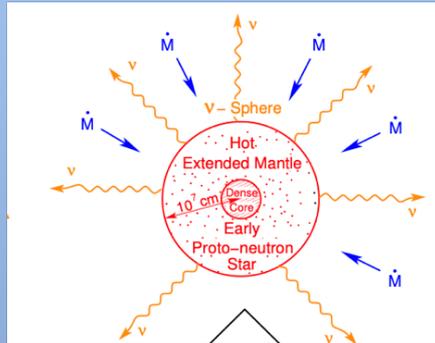


ASTROPHYSICS
Perfect probes : undeflected and unabsorbed
Multimessenger approach

PARTICLE PHYSICS
ATMOSPHERIC NEUTRINOS

- ν oscillations
- sterile neutrinos
- mass hierarchy
- tau appearance
- non-standard interactions

Super Novae explosion Neutrino oscillation Dark Matter HE neutrinos Multi-messenger program



Neutrino Energy from MeV to PeV

ARCA + ORCA

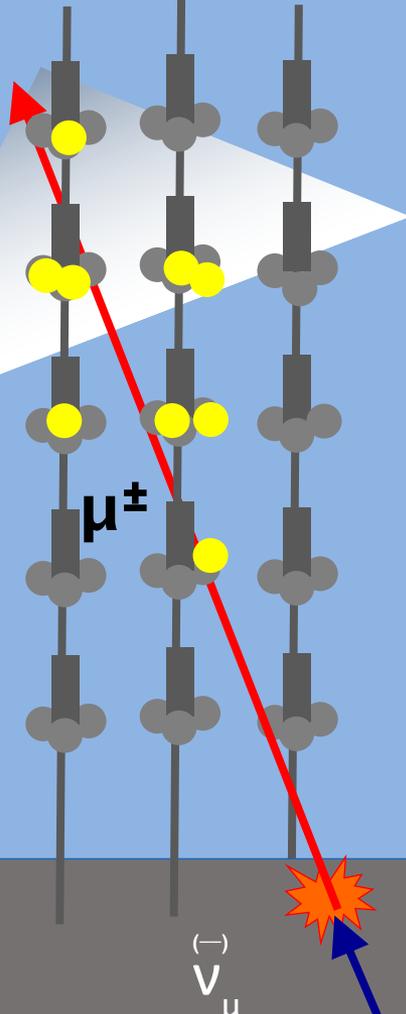
ORCA

ARCA

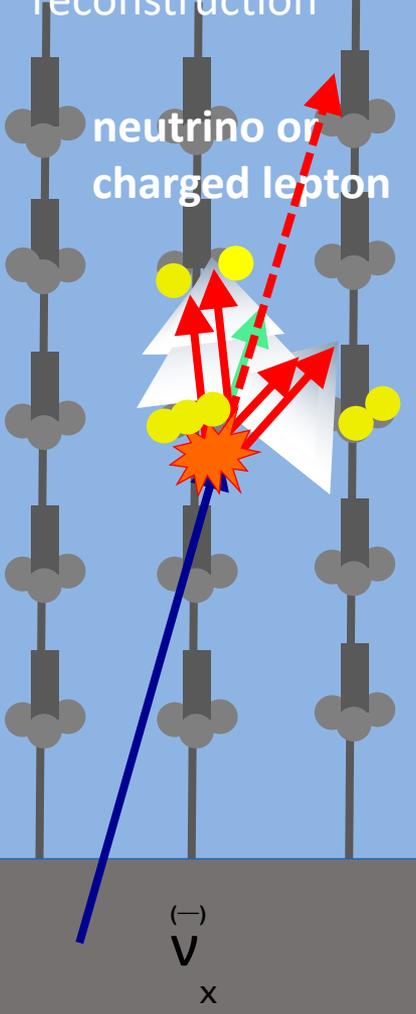


Events in a neutrino telescope

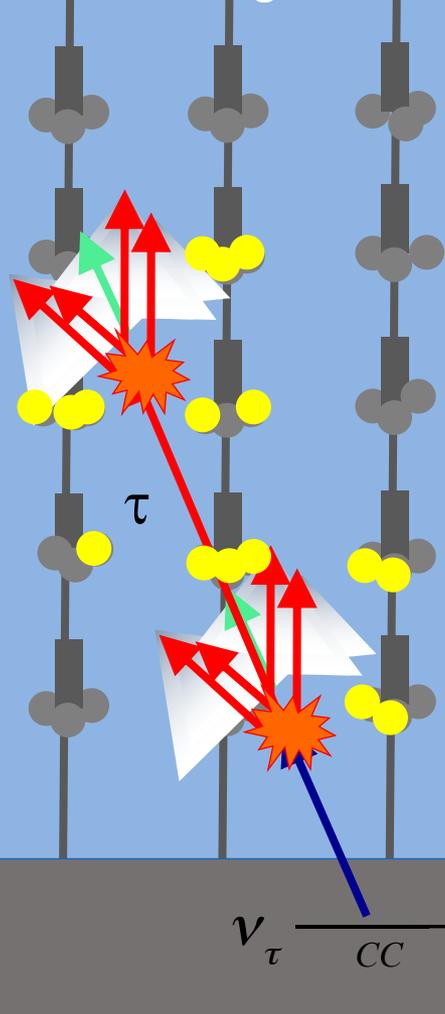
CC ν_μ
track like events
good pointing



CC ν_e + all flavours NC
shower like events –
good energy
reconstruction

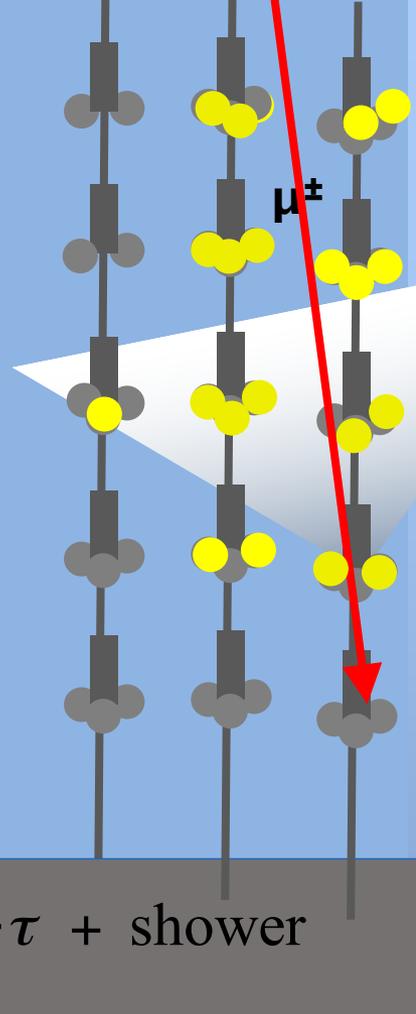


CC ν_τ
“double bang”



BACKGROUND !!

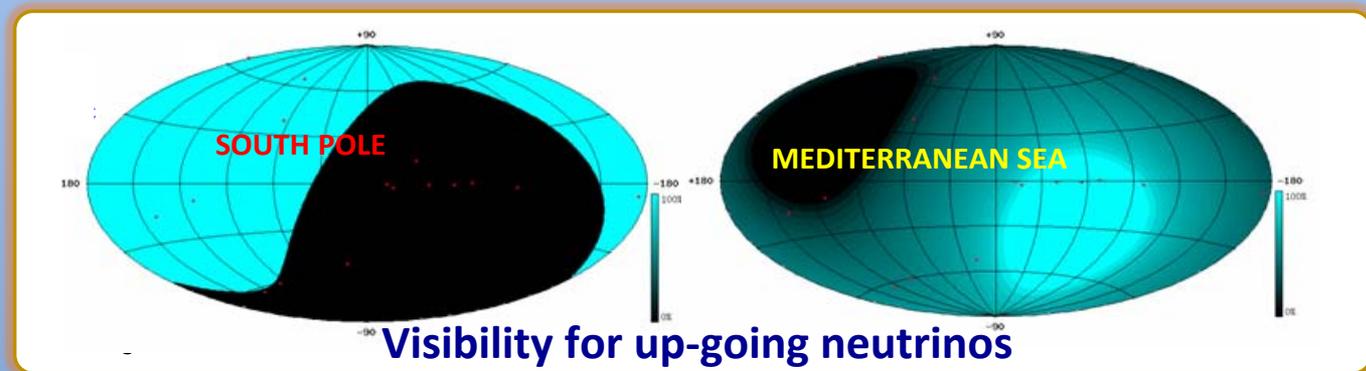
atmospheric muon



$$\nu_\tau \xrightarrow{CC} \tau + \text{shower}$$

Why the Mediterranean Sea

- Optical properties of water → Mapping the **Southern sky** with unprecedented angular resolution
 - tracks: $< 0.2^\circ$ @ 10TeV - ν_μ CC (neutrino astronomy)
 - showers: $< 2^\circ$ $E_\nu > 50$ TeV ν_e CC + ν_x NC (\sim calorimetric measurement)
- Visibility of the Galactic region → $\sim 70\%$ for the Galactic Centre



The KM3NeT detectors - Design

31 PMTs x 3" PMTs

The optical sensor:
the **D**igital **O**ptical **M**odule (DOM)

a 3D array of optical sensors

Seafloor Network

~ 40 cm

The **D**etection **U**nit (DU)

String:

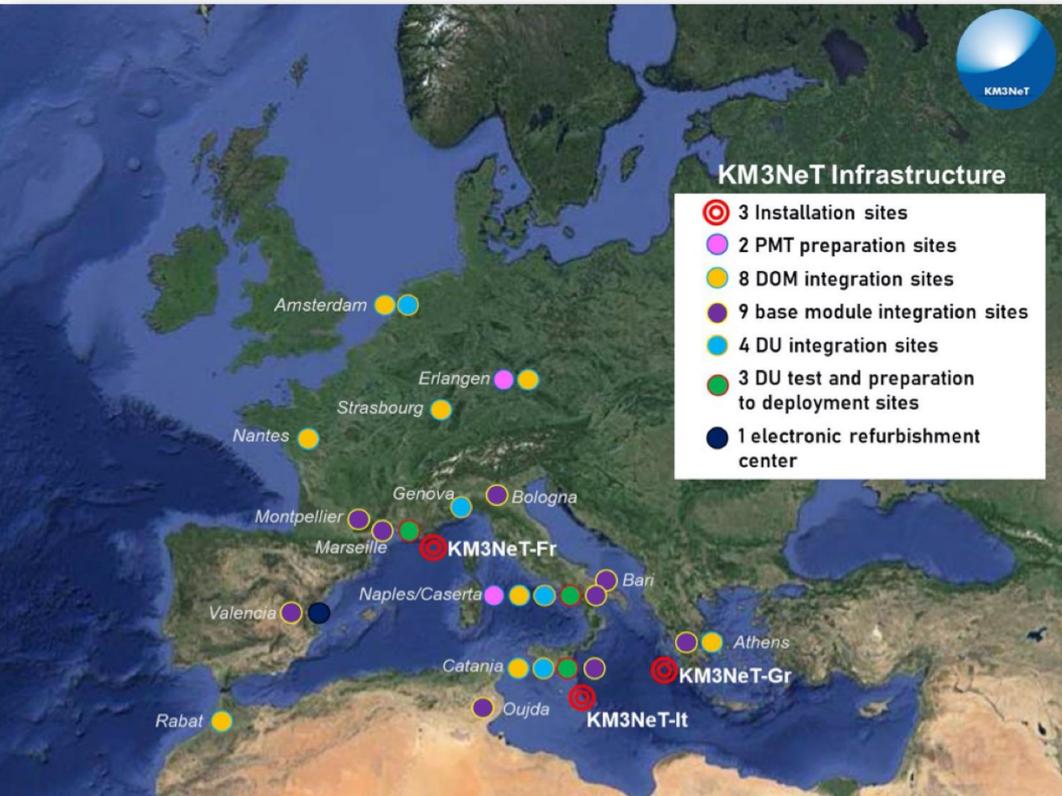
- 1 Buoy
- 2 Dyneema ropes
- 18 DOMs
- 1 Anchor
- Electro-optical backbone:
 - Flexible hose 7mm
 - Oil-filled
 - 18 fibres
 - 2 copper wires



Improved background rejection
Compact and cost effective design:
photocathode area $\approx 3 \times 10''$ PMTs

1 Building Block (BB) = 115 DUs
ARCA= 2 BB = 230 DUs
ORCA= 1 BB

Several integration sites



DOMs:

- 8 integration sites
- 640 DOM produced (400 ARCA, 240 ORCA)
- 100 currently in progress

BMs:

- 9 integration sites
- 27 BMs produced
- 6 currently in progress

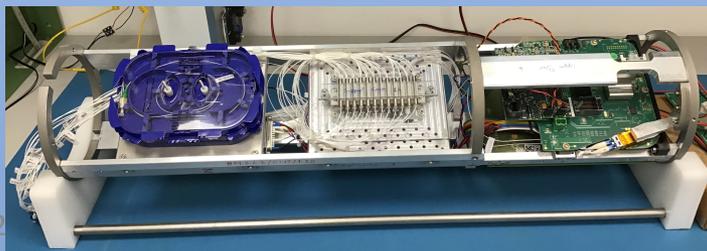
DUs:

- 5 integration sites
- 13 DUs produced
- 8 currently in progress

Total: 22 integration sites!
(last year: 15)



Base Modules



Digital Optical Modules

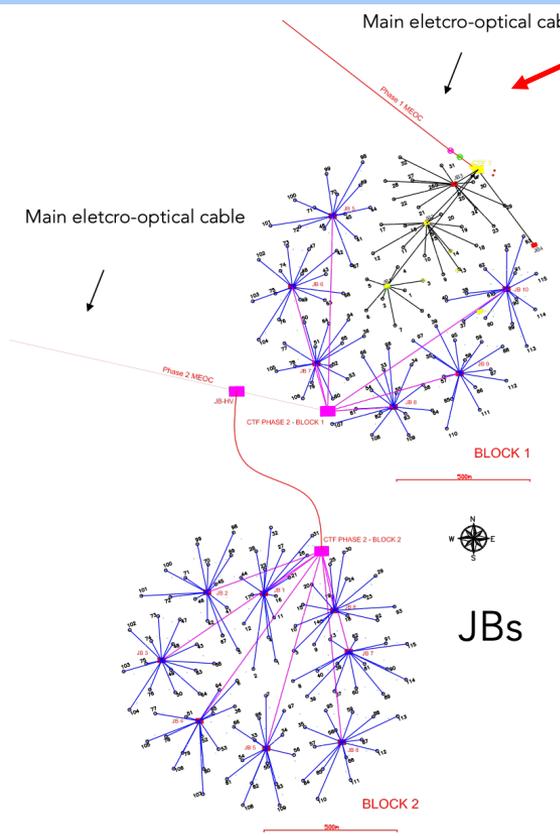


Detection Units

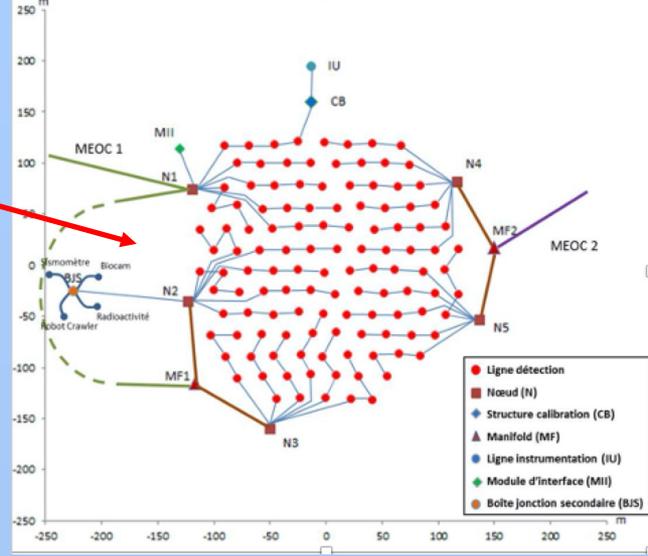


KM3NeT/ARCA (KM3NeT-It site)

Footprint of the two telescopes



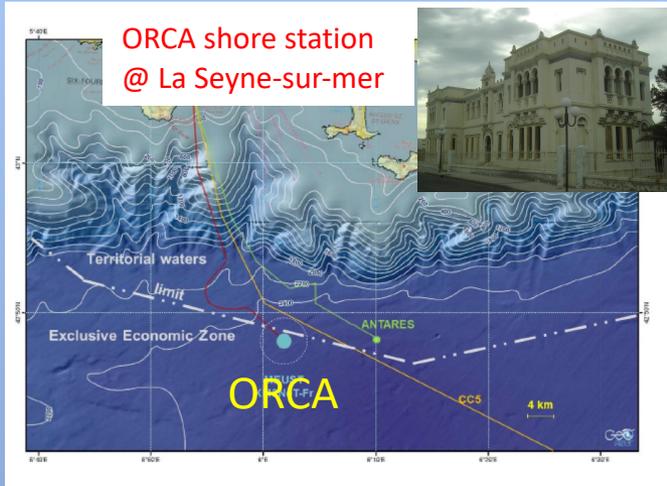
- same technology
- same design
- different density of active sensors
- modular design
- power and data distribution with a single backbone cable
- breakouts at DOMs
- network of submarine cables + Junction Boxes
- connection to shore via a main e/o cable
- all-data-to-shore



KM3NeT/ORCA (KM3NeT-Fr site)

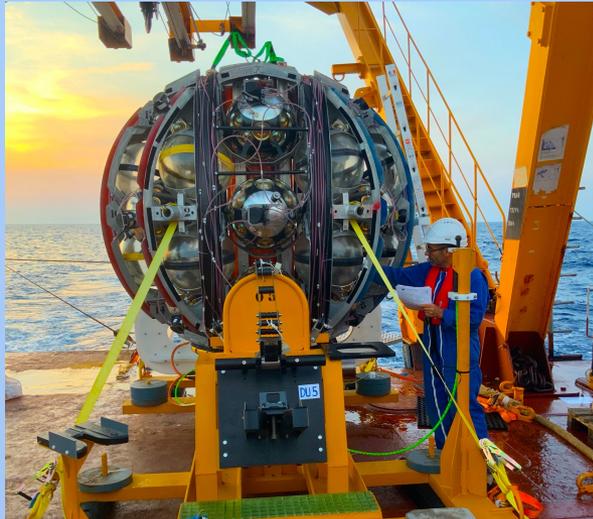


ARCA shore station
@ Portopalo di Capo Passero

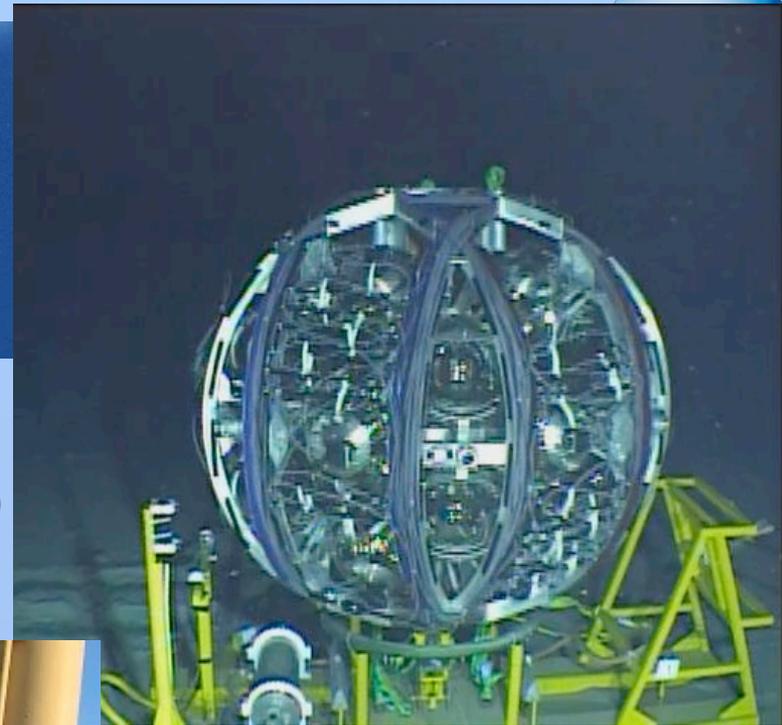


	ARCA	ORCA
Location	Italy	France
DU distance	90 m	20 m
DOM spacing	36 m	9 m
DU height	~ 800 m	~ 200 m
Instrumented mass	2*500 Mton	7 Mton
Depth	3500 m	2500 m

KM3NeT Deployment



The LOM
(Launcher of Optical Modules)

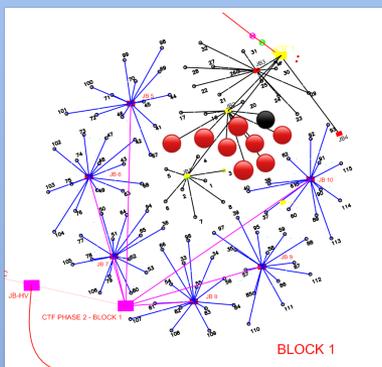


- Rapid deployment
- Multiple strings/sea campaign
- Autonomous/ROV unfurling
- Reusable

Status of ARCA

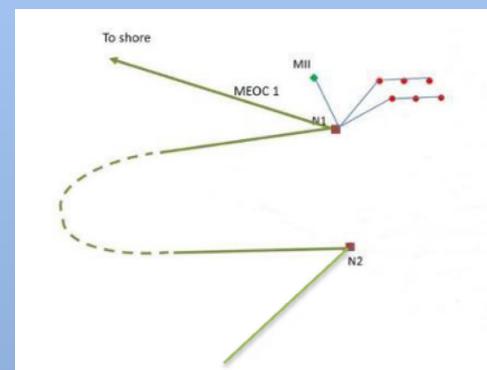
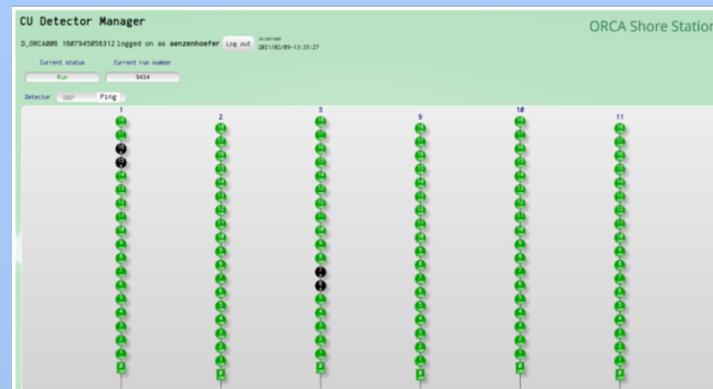
December 2015 → First DU
 April 2021 → + 5 DU & 1 JB
 September 2021 → + 3 DU

9 DU: **8 DU in data taking**



Status of ORCA

6 DU in data taking
 since February 2020



Next sea campaign : Spring 2022
12-14 Detection Units + 2 Junction Boxes

Next sea campaign : November 2021
+ 7 Detection Units



Some events in ARCA





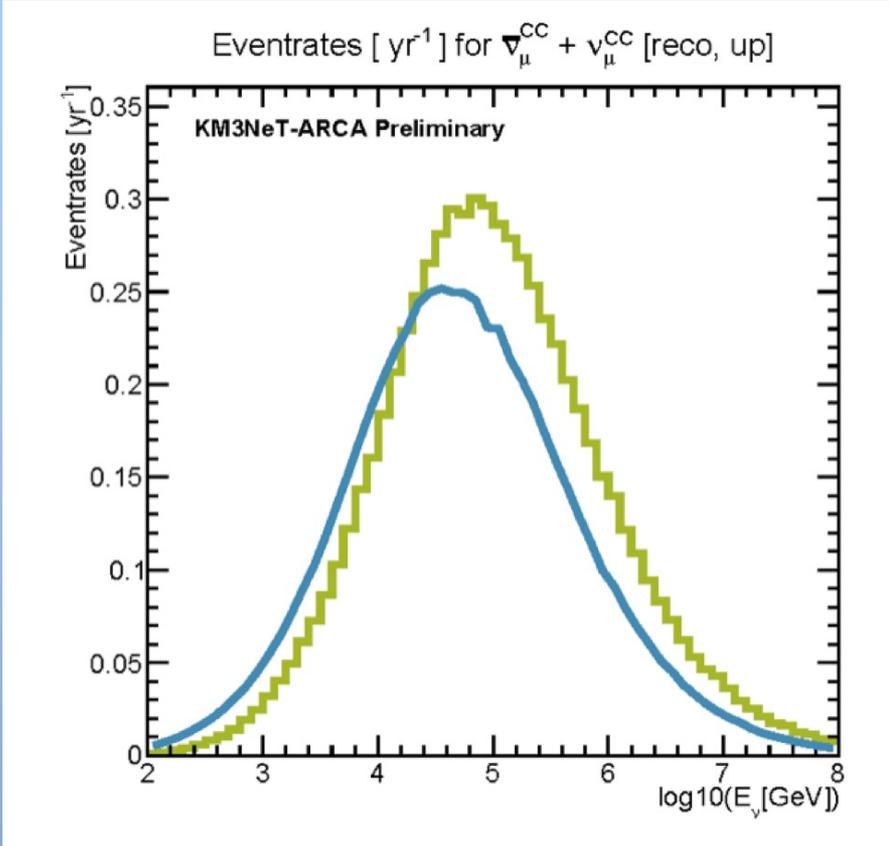
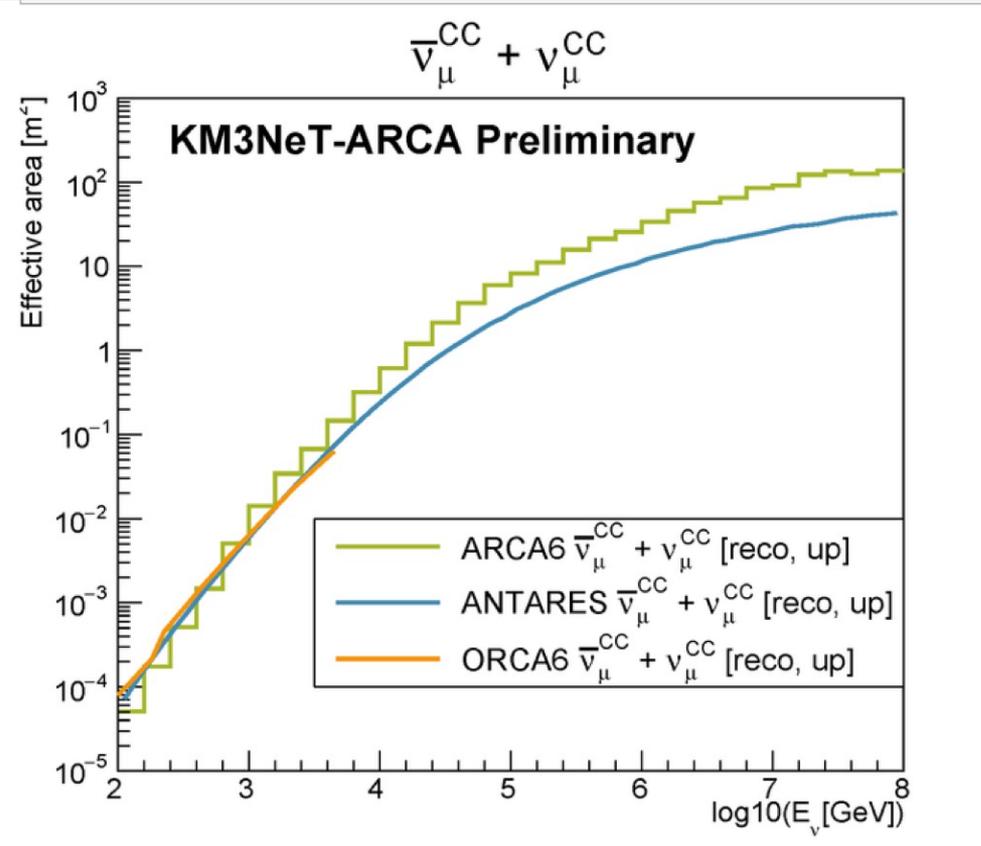
KM3NeT effective area

Upgoing tracks selected

ARCA6 + ORCA6 area > ANTARES

Number of events per year for a cosmic diffuse flux

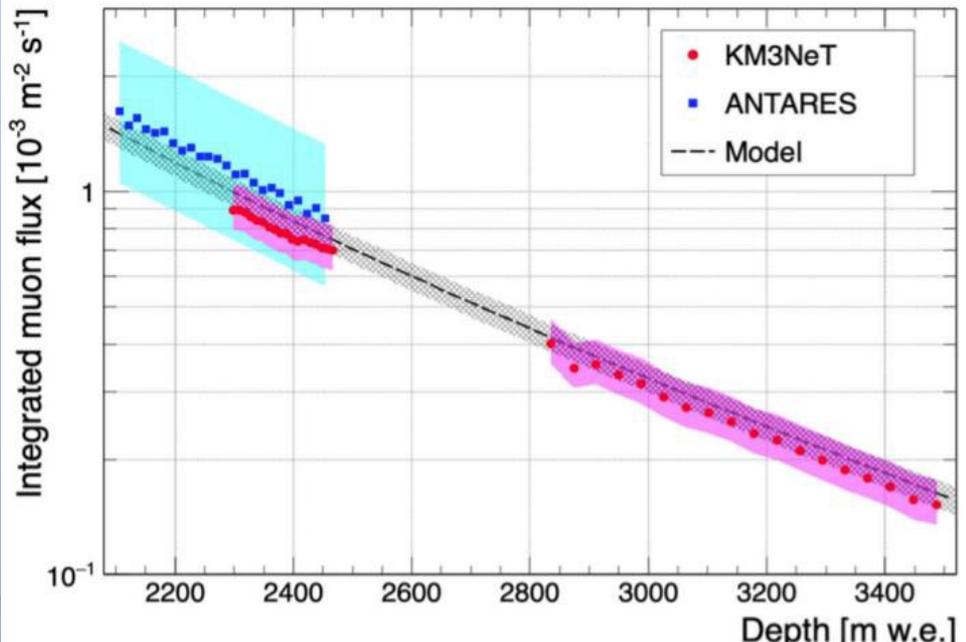
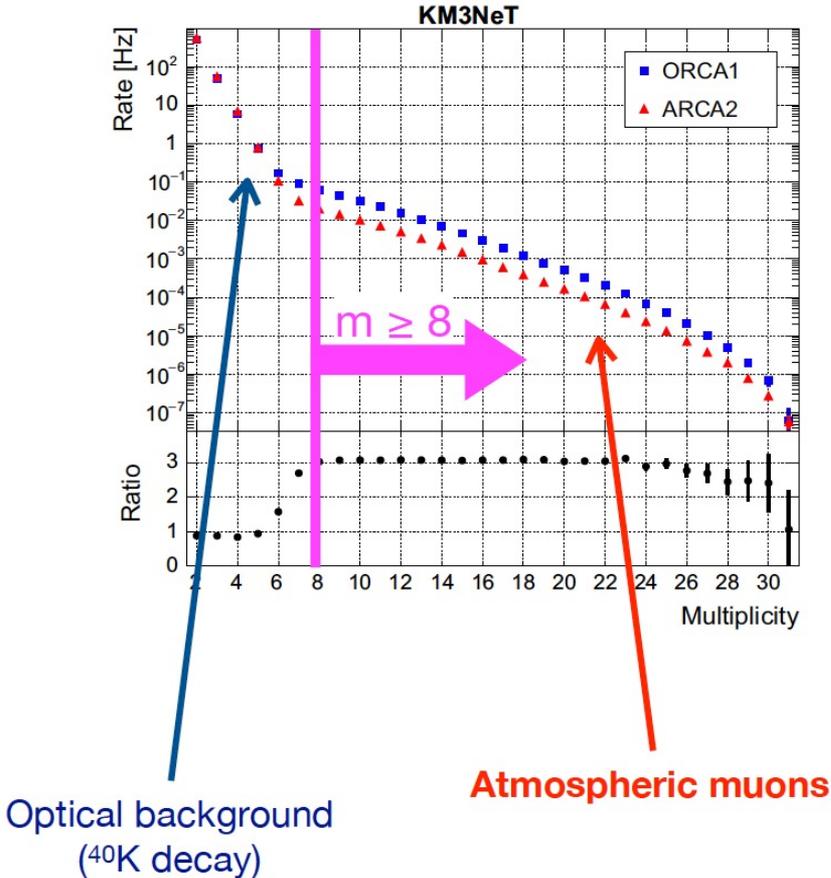
$$\Phi = 10^{-8} E^{-2} \text{GeV}^{-1} \text{cm}^{-2} \text{s}^{-1} \text{sr}^{-1}$$



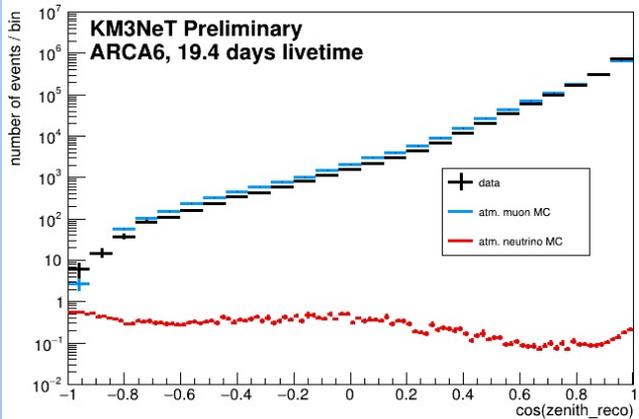
Atmospheric muon flux vs depth

- Single-DOM measurement
- Useful calibration tool
- Good comparison with ANTARES and theoretical model

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Selection of atmospheric neutrinos



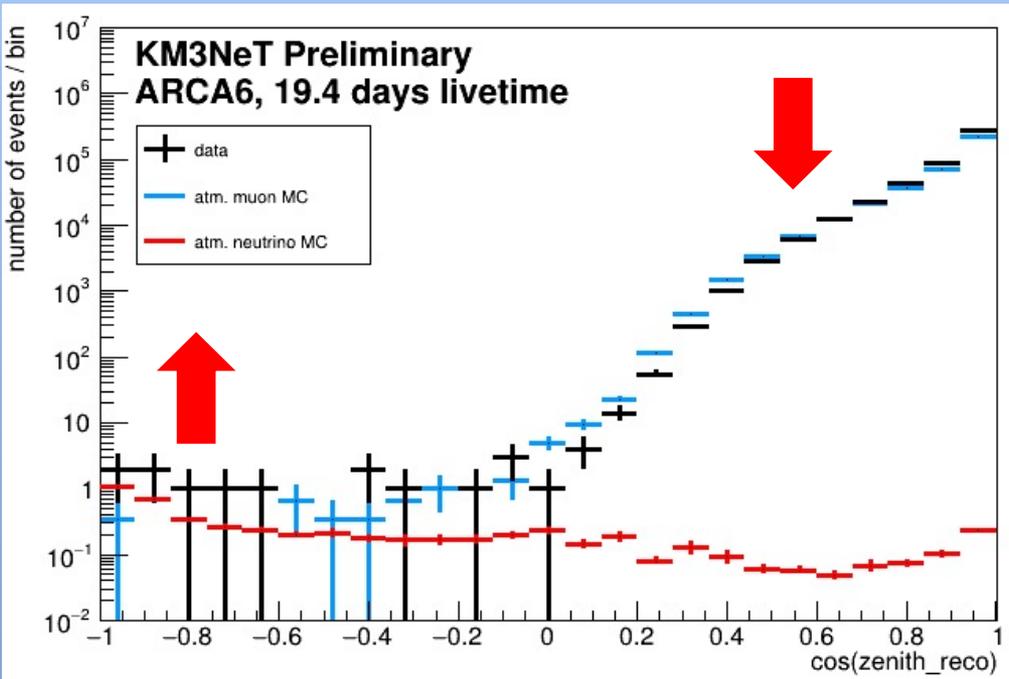
Only neutrinos can cross the Earth

Upgoing tracks = neutrino induced events
 Downgoing tracks = dominated by atm muons

After reconstruction quality cuts

15 up-going tracks

MC expectations: 4 atm ν + 7 atm μ



Core Collapse SuperNovae (CCSN)

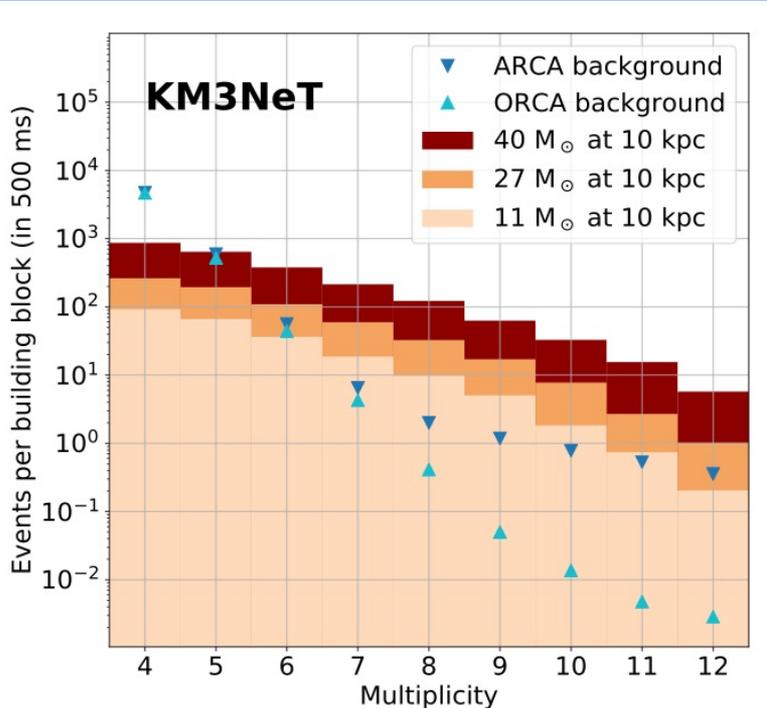


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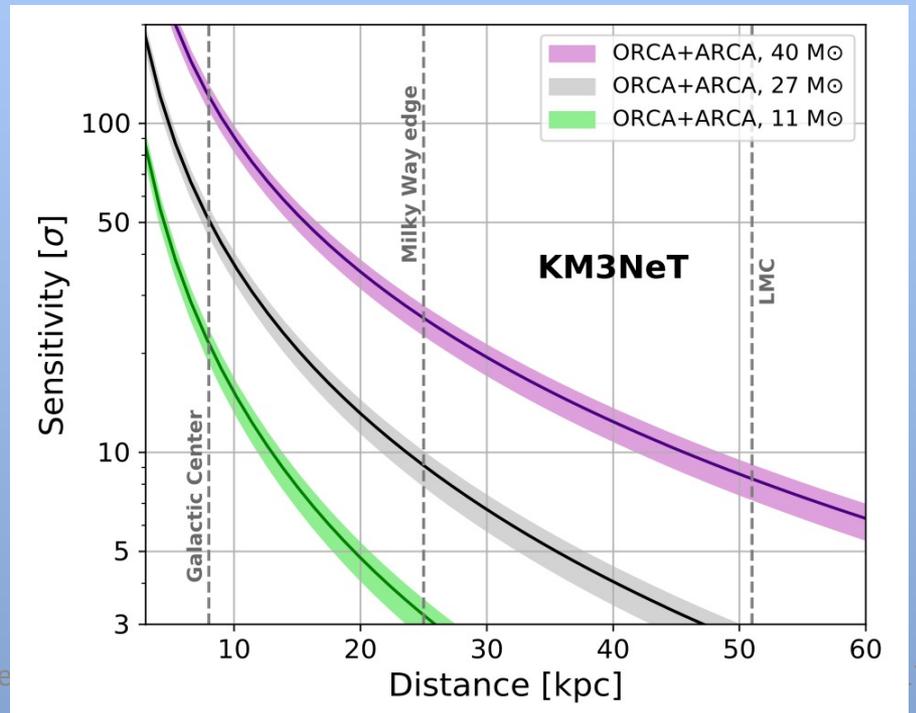
ARCA6+ORCA6 already sensitive to 60% of Galactic CCSNe (<11 kpc)
real time trigger active in SNEWS since early 2019

> 5 σ for ARCA+ORCA for 27M \odot at a distance ~ 36 kpc

ORCA 1 BB + ARCA 1BB



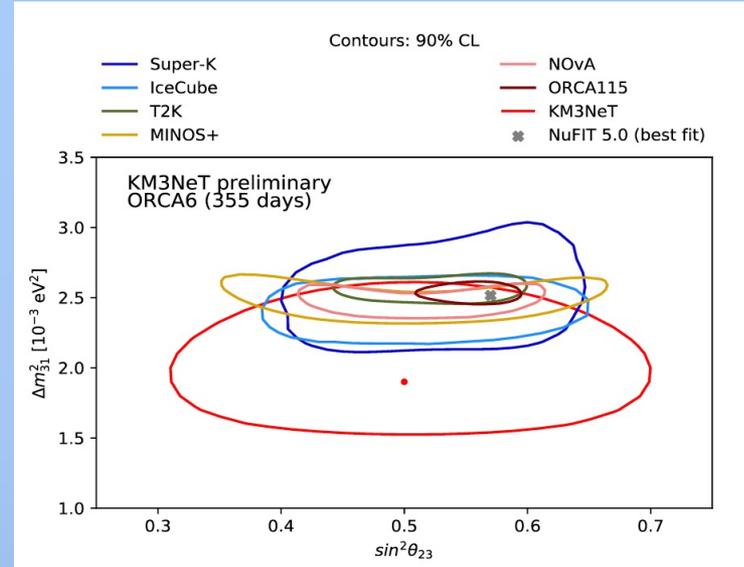
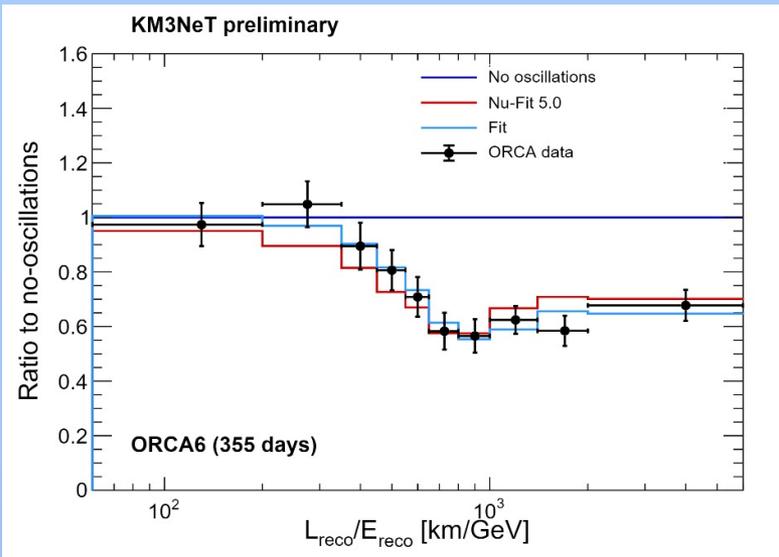
The KM3NeT



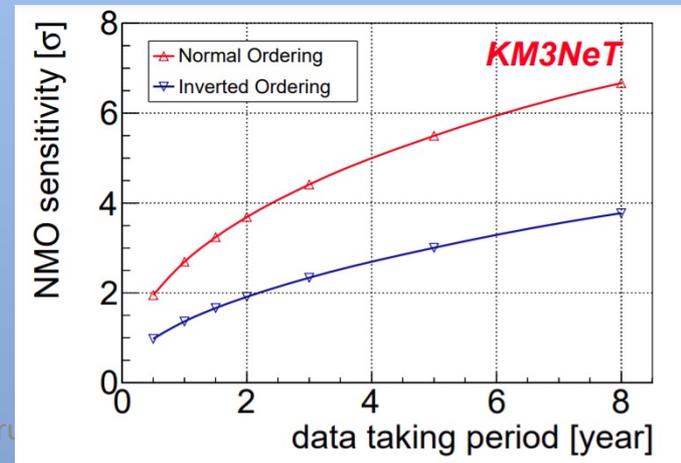
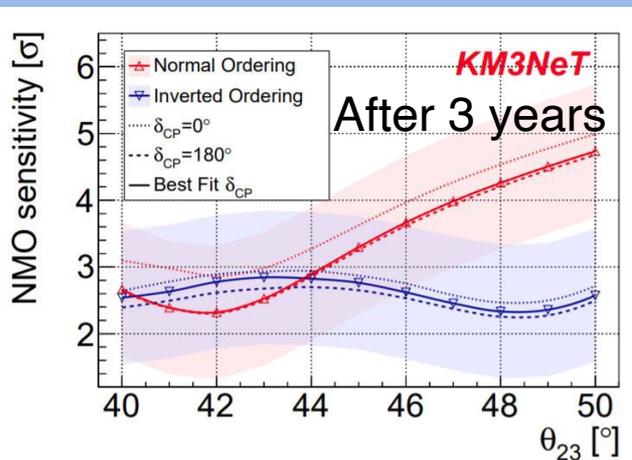
Neutrino studies with KM3NeT/ORCA



Oscillations – sensitivity after 3 y and Preliminary results



Neutrino Mass Ordering



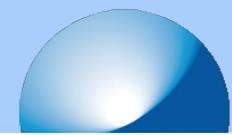


Conclusions and perspectives

- KM3NeT is taking over the responsibility of neutrino astronomy in the Mediterranean Sea from ANTARES
- Technique fully validated
- First preliminary results confirm expectations
- Detectors are in construction
- 8 ARCA + 6 ORCA Detection Units are taking data
- Next sea campaigns:
 - ORCA : November 2021 → + 7 DUs
February 2022 → + 3 DUs & Calibration Base & Instrumentation Unit
 - ARCA : Spring 2022 → 12 DUs & 1JB + & Calibration Base & Instrumentation Unit

Backup slides





Background

Atmospheric muons

direction cut \rightarrow upward going events

useful for detector calibration

Atmospheric neutrinos

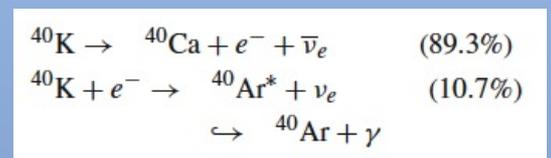
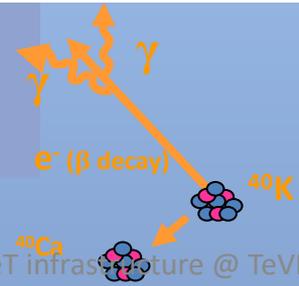
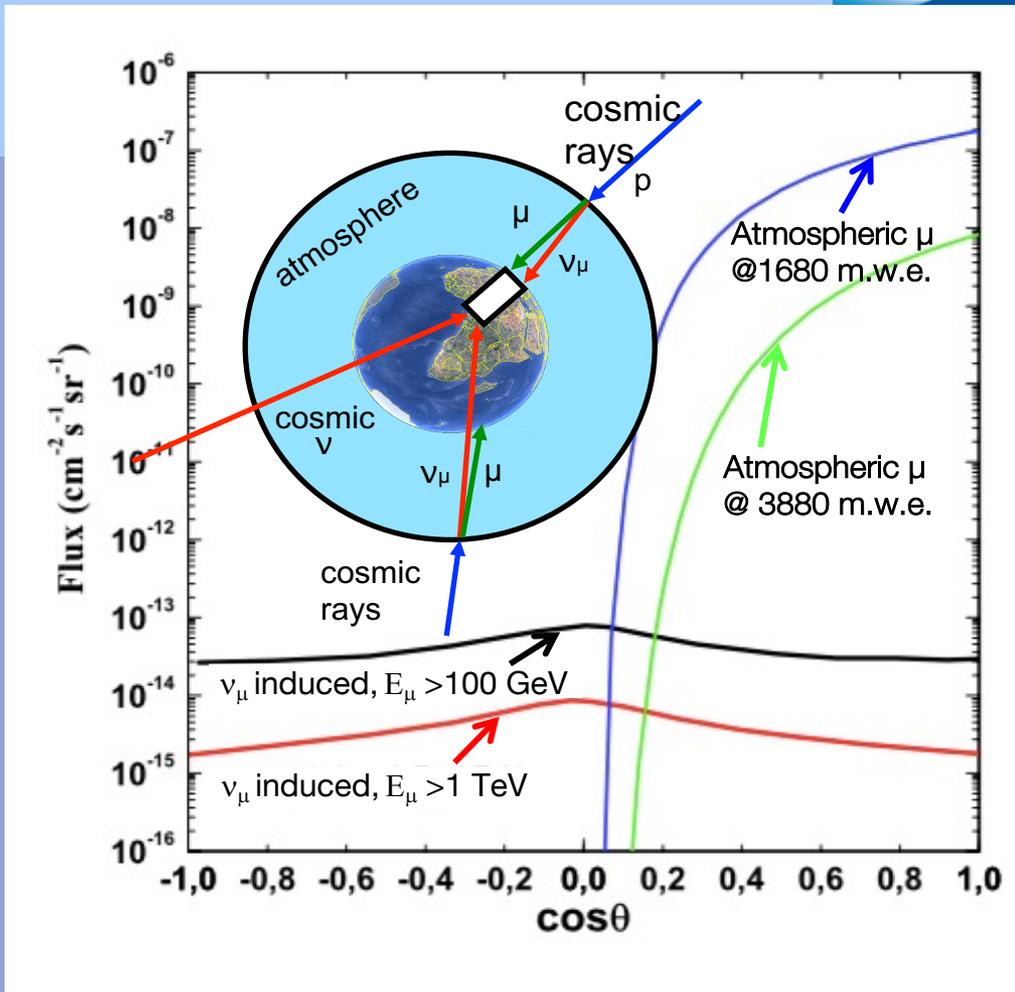
irreducible background

neutrino oscillations \rightarrow mass hierarchy

Environmental background

- ^{40}K decay
- bioluminescence
-

rejection : causal correlation of the signals



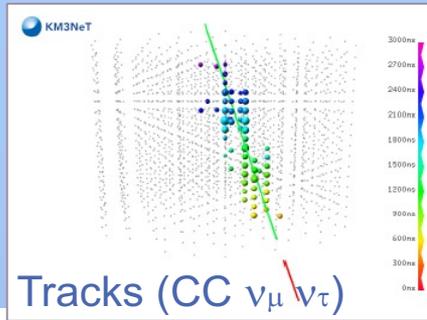
Reconstruction performances

Excellent optical properties of sea water → good angular resolution

Tracks → ν_μ CC (neutrino astronomy)

Angular resolution $< 0.2^\circ$ @10 TeV

Energy resolution $\sim 30\%$



Showers → ν_e CC + ν_x NC

Angular resolution $< 2^\circ$ $E_\nu > 50$ TeV

Energy resolution $< 5\%$

Contained events, calorimetric meas.

