

Stochastic Fluctuations of Low-Energy Cosmic Rays

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Sarah Recchia, and Stefano Gabici.

Phan et al., 2021, PhRvL, 127, 141101

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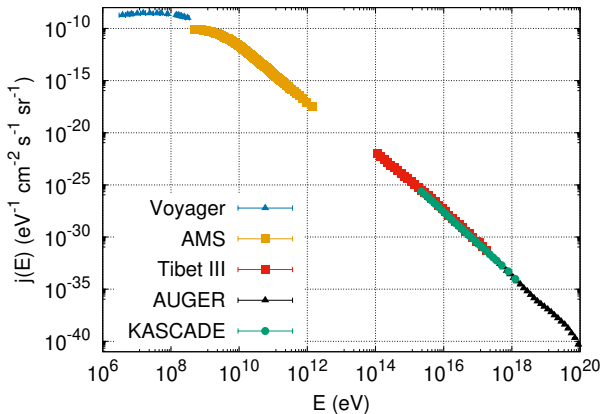


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de Paris

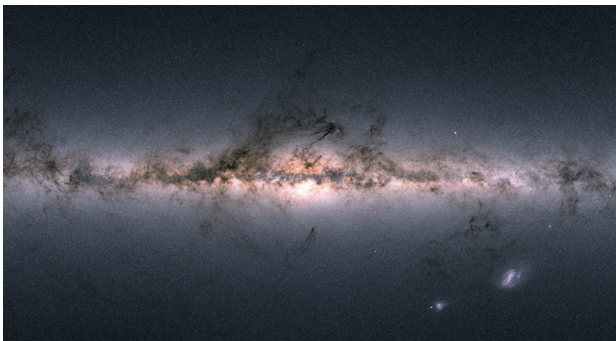


- **Physical Context**
- **Stochastic Model for Low-Energy Cosmic Rays**
- **Results and Implications for Voyager Data**
- **Conclusions and Perspectives**

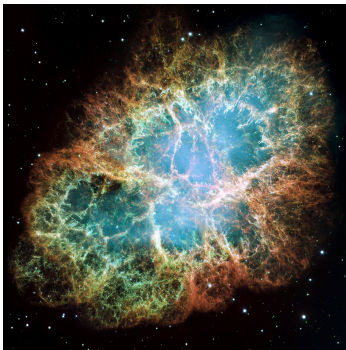
- We will focus on cosmic rays of energy below 10 GeV since these particles affect the dynamics of various processes in the Universe including the formation of new stars.

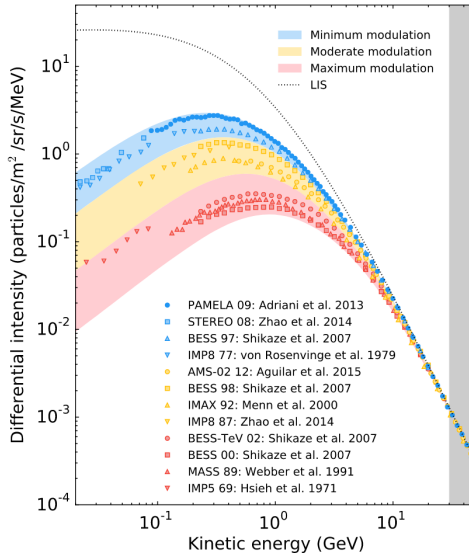


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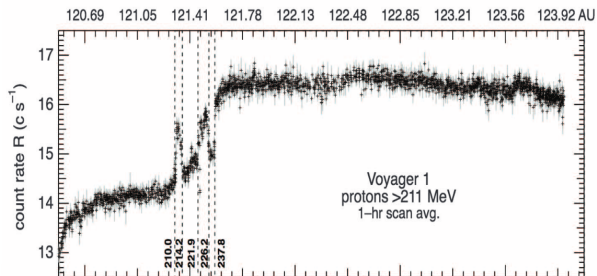


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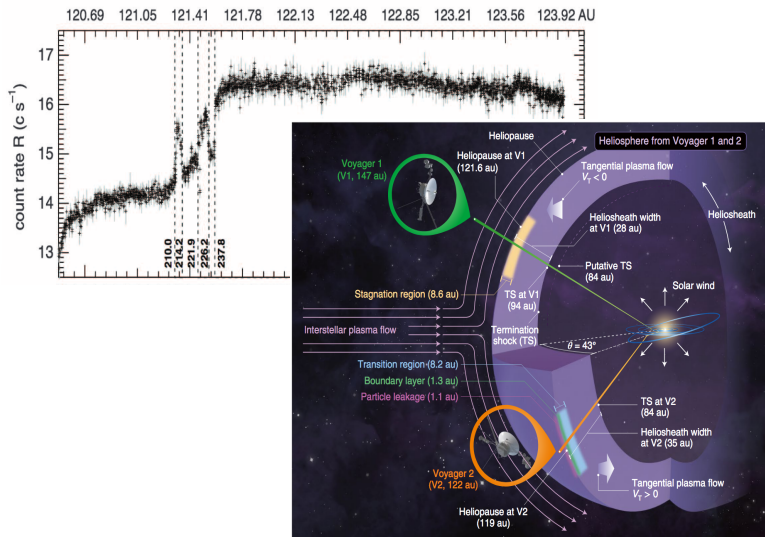


Voyager Probes



Krimigis et al., *Science*, 2013
Strauss, *Nature Astronomy*, 2019

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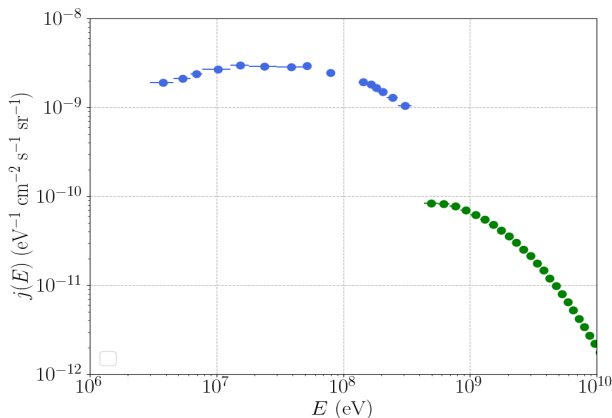
Strauss, *Nature Astronomy*, 2019

October 26, 2021

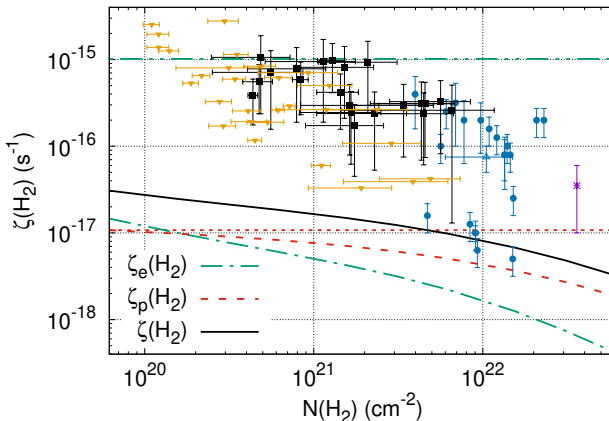
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Ionization Rate in Diffuse Clouds

- Ionization rate is **the production rate of H_2^+ ions**.
- A difference of about **1 to 2 orders of magnitude** between the ionization rate in diffuse MCs estimated from the Voyager spectra and the observed data.



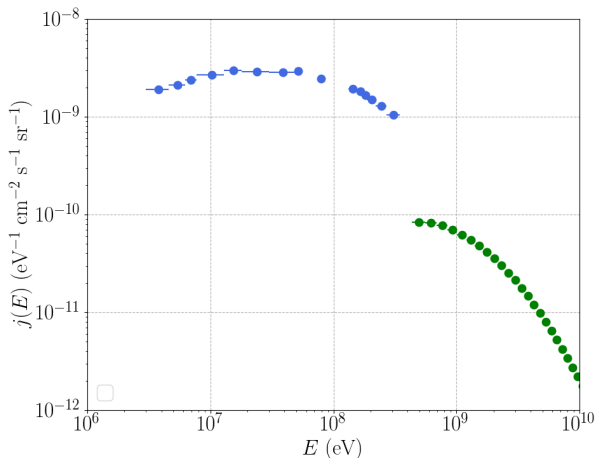
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How to Fit Voyager Data?

- The transport equation of cosmic rays:

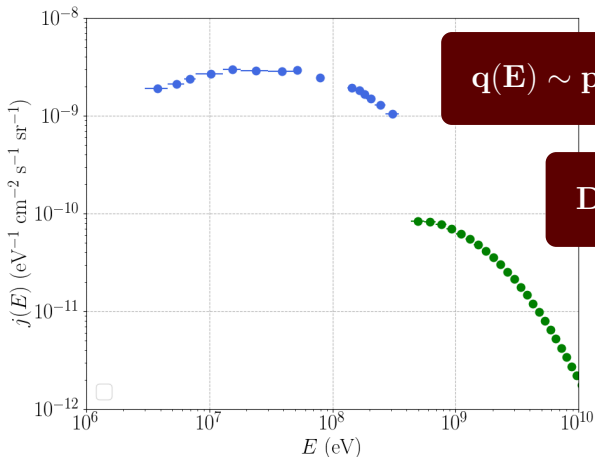
$$\frac{\partial \psi}{\partial t} + \frac{\partial}{\partial z} (u\psi) - D\nabla^2 \psi + \frac{\partial}{\partial E} (\dot{E}\psi) = q(r, z, E, t),$$



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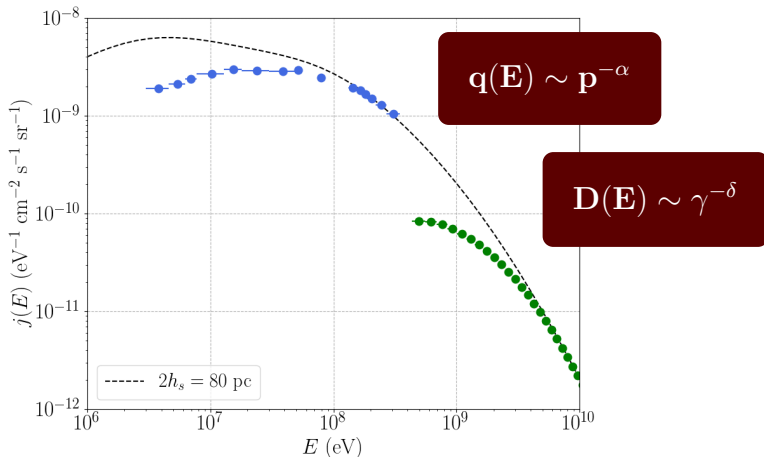
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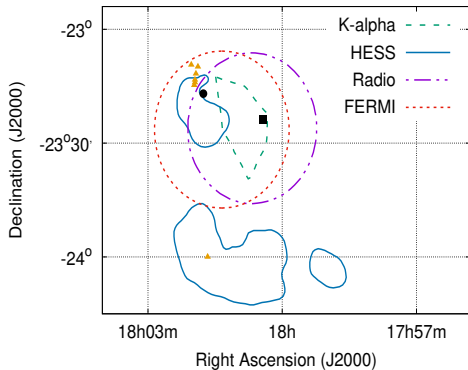
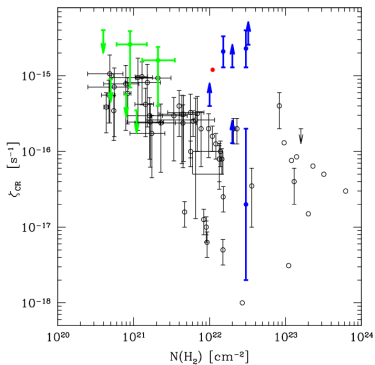
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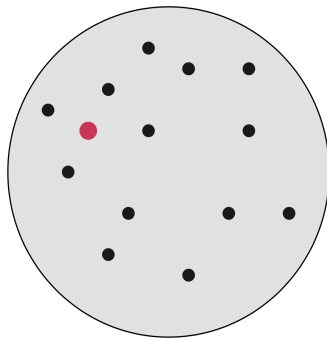
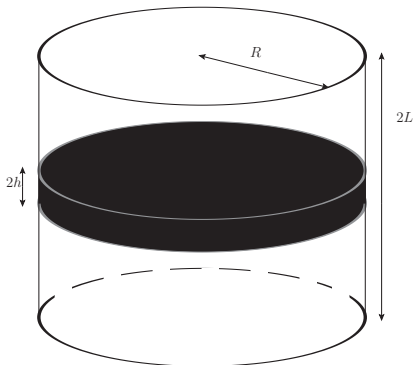
Sources of Low-Energy Cosmic Rays?



Gabici & Montmerle, *Pos ICRC*, 2015
Phan et al., *A&A*, 2020

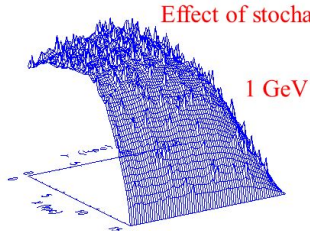
- The transport equation of cosmic rays for discrete sources in the disk:

$$\frac{\partial \psi}{\partial t} + \frac{\partial}{\partial z} (u\psi) - D\nabla^2 \psi + \frac{\partial}{\partial E} (\dot{E}\psi) = q(r, z, E, t),$$
$$q(r, z, E, t) = \sum_{i=1}^{N_s} Q(E) \frac{\delta(r - r_i)}{2\pi r_i} \delta(z - z_i) \delta(t - t_i).$$

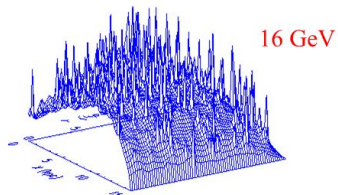


An Example for Stochasticity

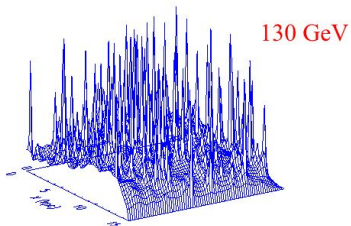
particle #0 electrons:1.02e+03 MeV



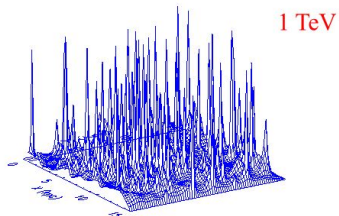
particle #0 electrons:1.64e+04 MeV



particle #0 electrons:1.31e+05 MeV



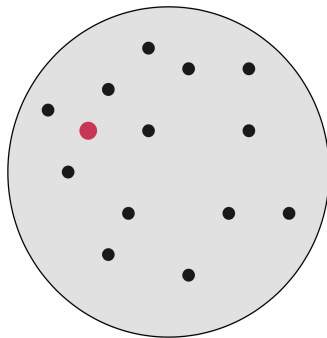
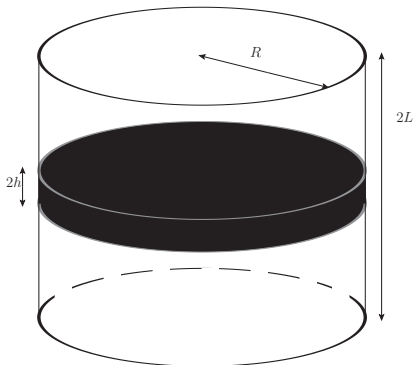
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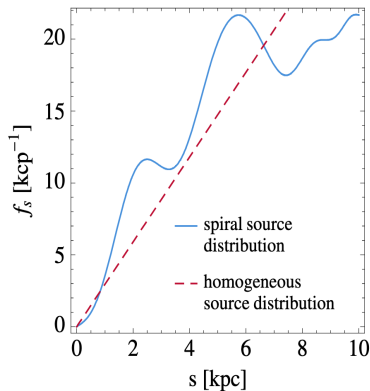
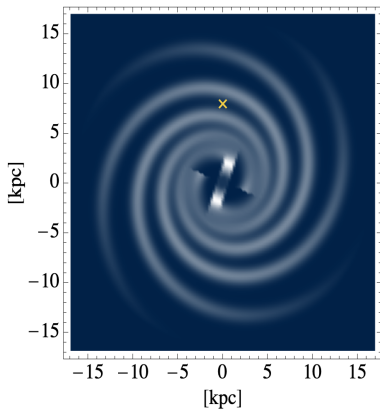
Strong & Moskalenko ICRC2001

- The transport equation of cosmic rays for discrete sources in the disk:

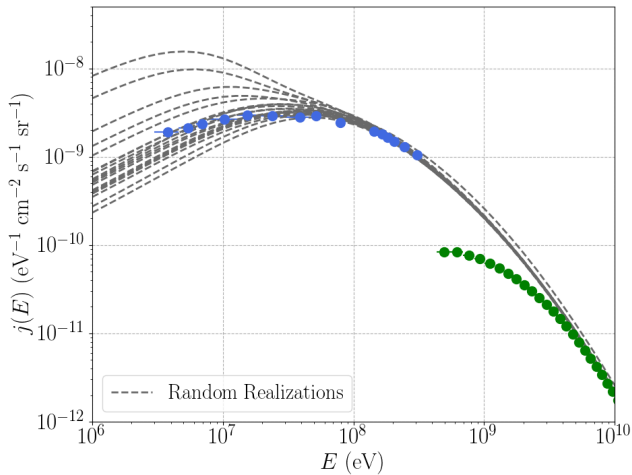
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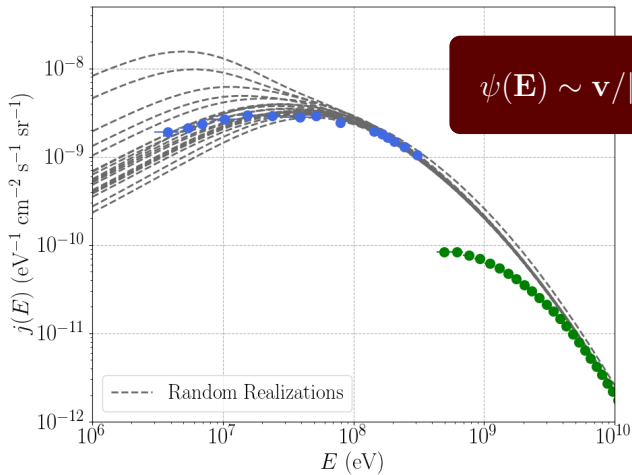


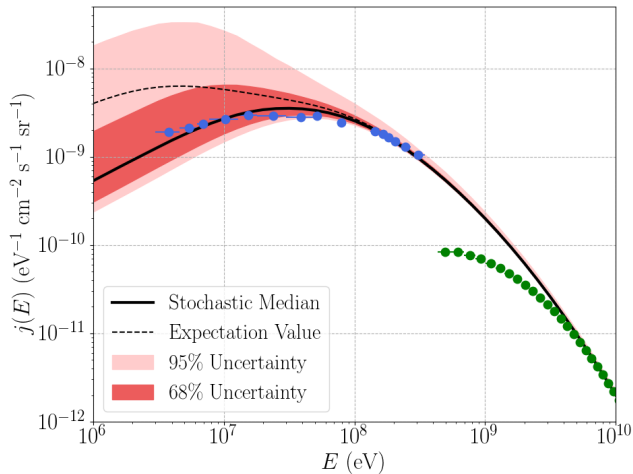
Radial Distribution of Sources



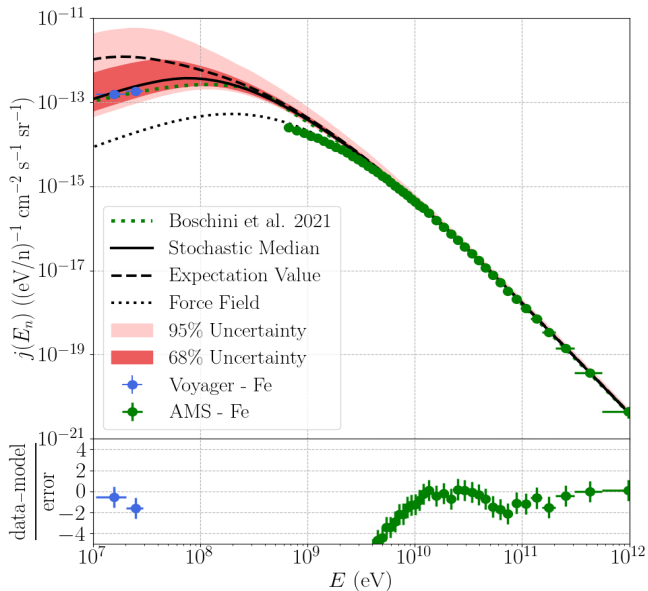
Mertsch, *JCAP*, 2011
Ahlers et al., *PhRvD*, 2009



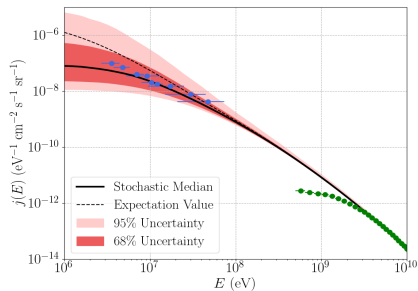
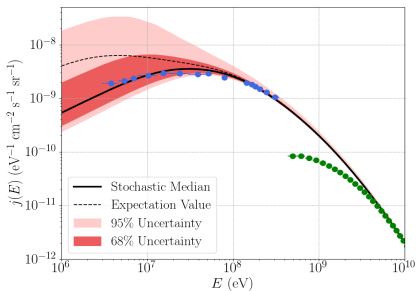




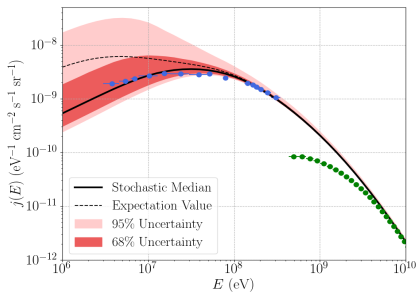
Preliminary Results on Cosmic-Ray Iron



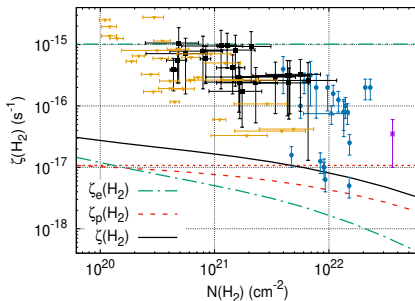
- Stochasticity is important for the modelling of low-energy cosmic rays and this effects allows us explain the Voyager data without requiring any un-physical breaks.



- The stochastic fluctuations might open up a way to explain both the **Voyager data** and the **ionization rate**.



Voyager data



Ionization rate data