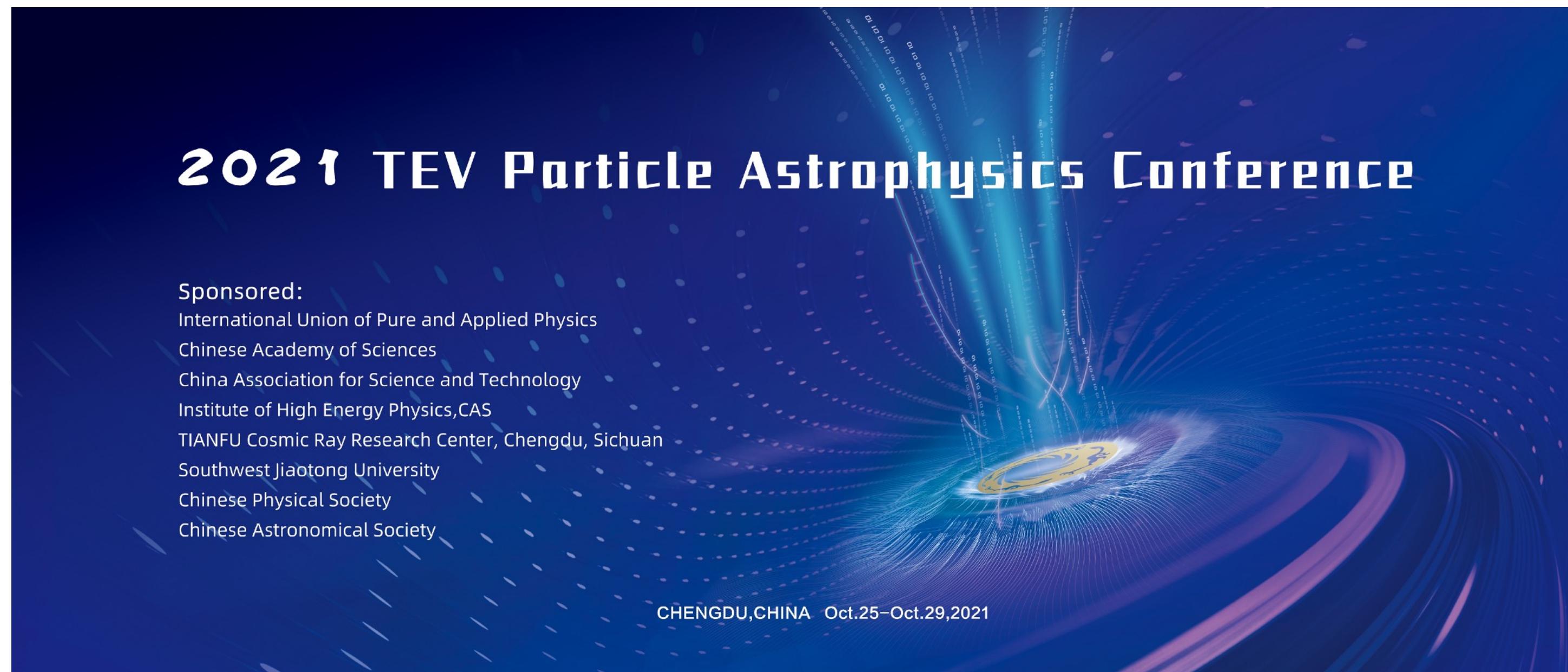


# Implications of Li to O data of AMS-02 on our understanding cosmic-ray propagation

**Michael Korsmeier**  
in collaboration with Alessandro Cuoco

Mostly based on:  
*Phys.Rev.D 103 (2021) 10, 103016*



# Outline

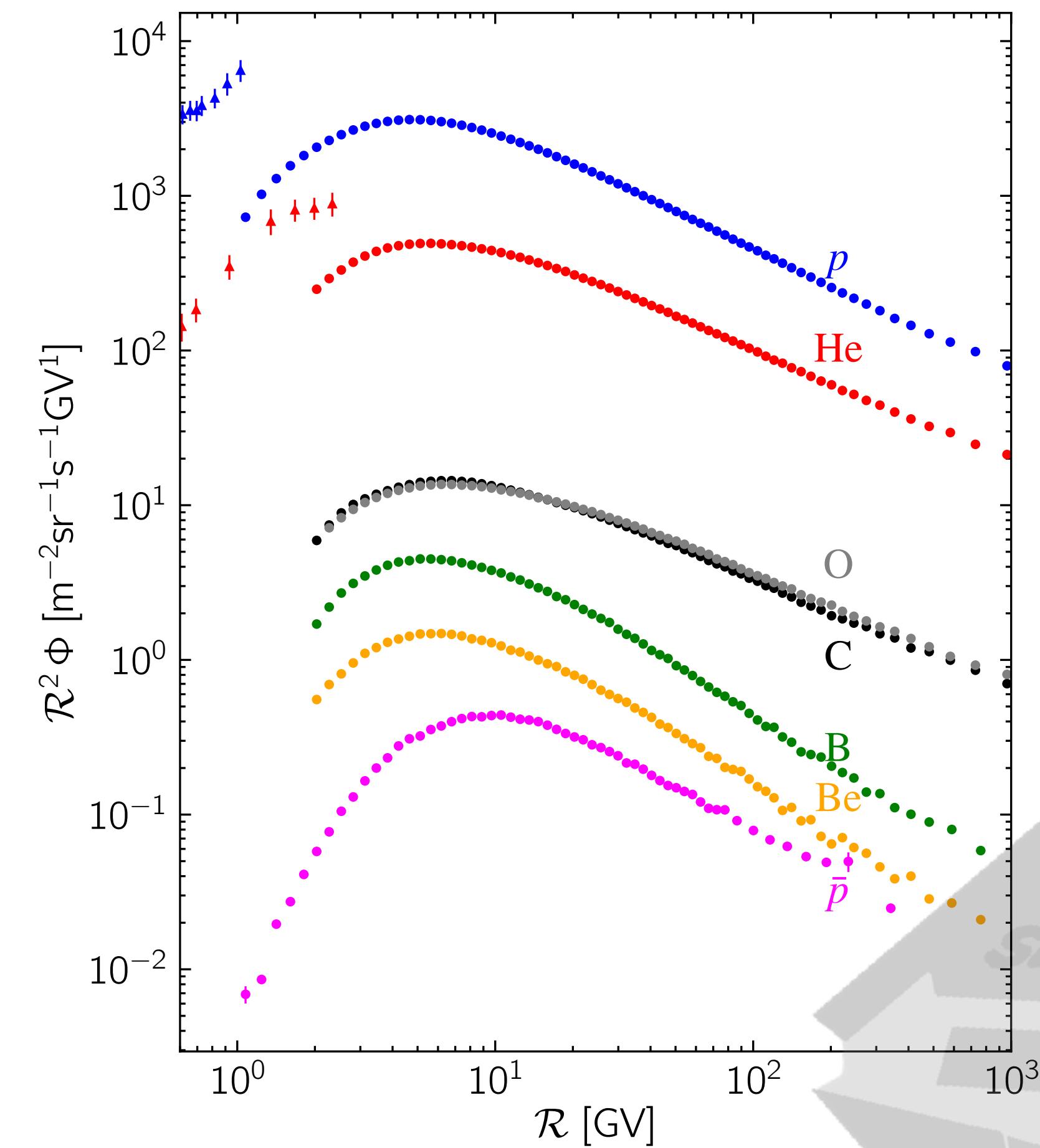
**Introduction to CR propagation**

**Global fit of CR data from Li to O**

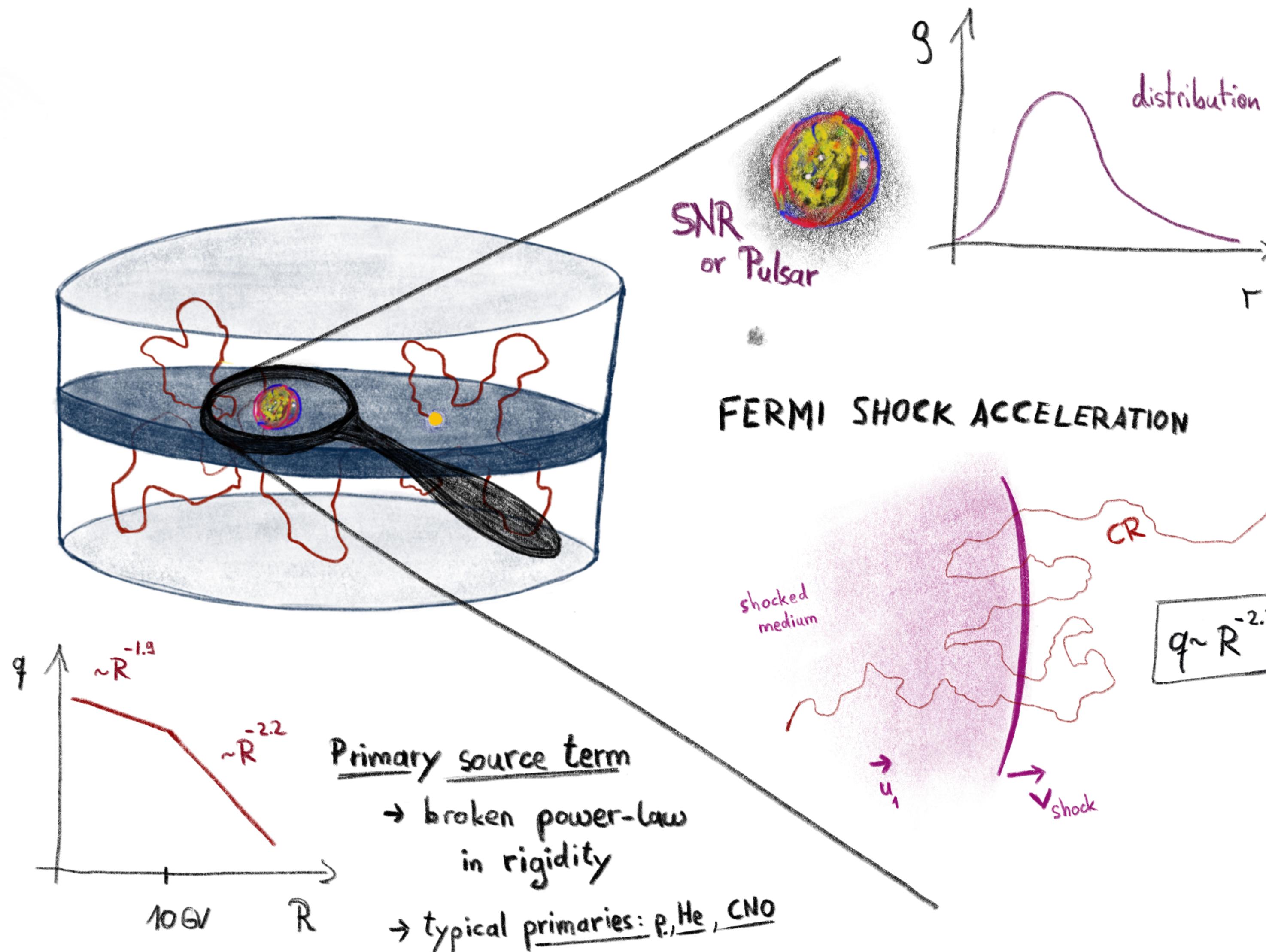
**Cross section uncertainties**

**Correlations in the AMS-02 data**

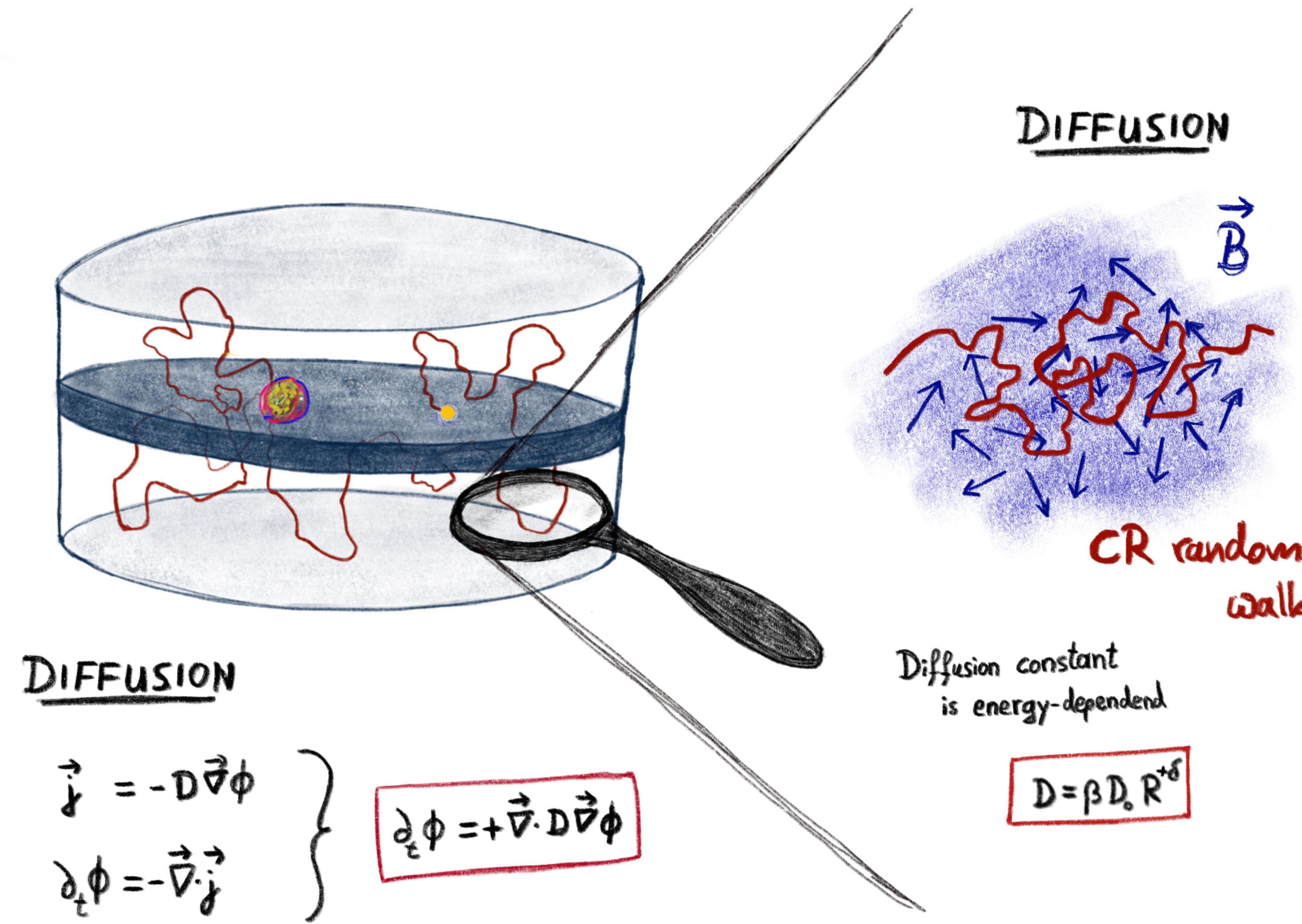
**Conclusion & outlook**



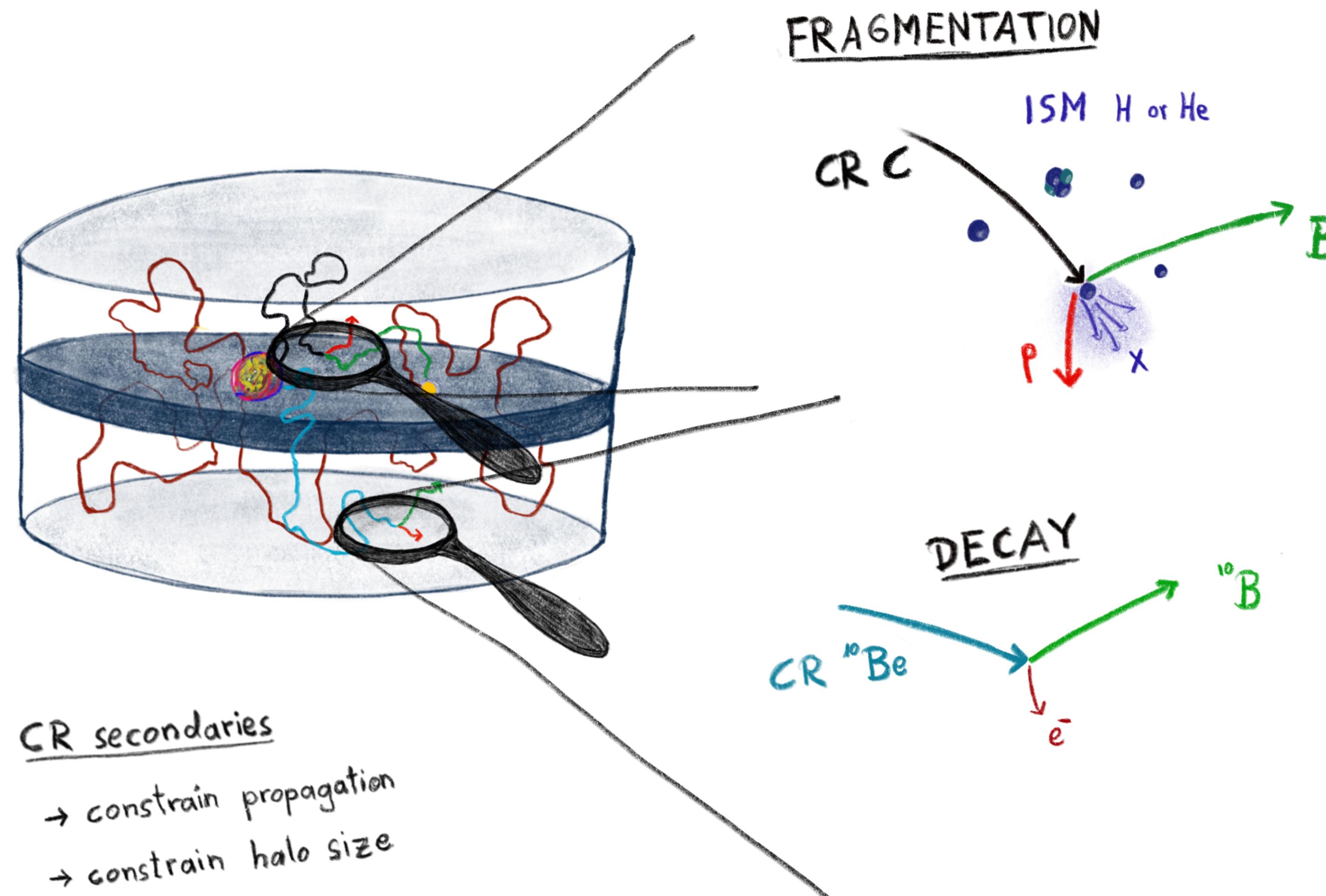
# Modeling cosmic-ray propagation



# Modeling cosmic-ray propagation



# Modeling cosmic-ray propagation

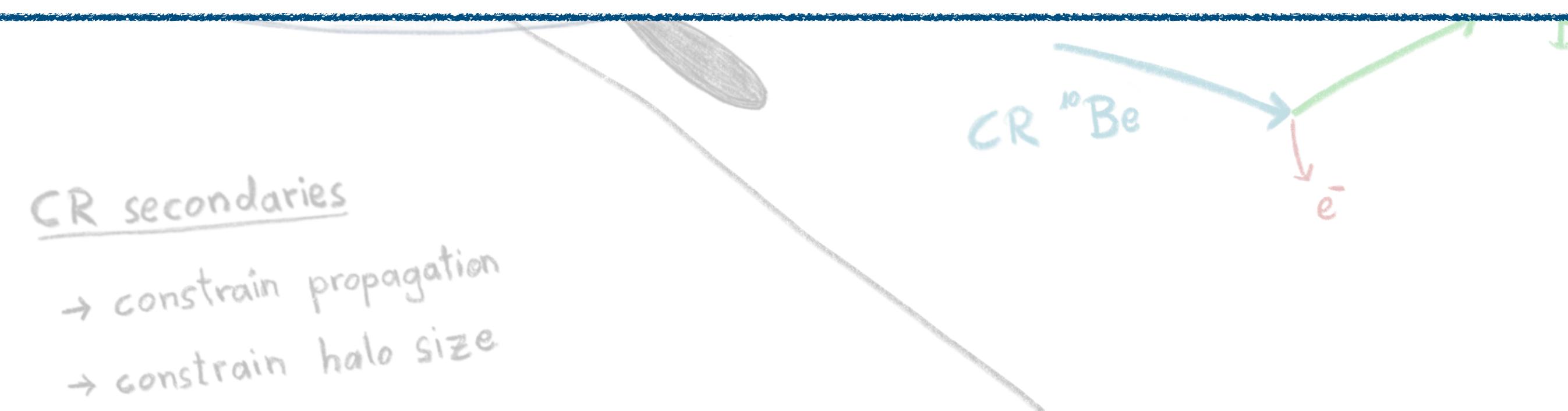


# Modeling cosmic-ray propagation



**CR propagation is described by diffusion equations.**

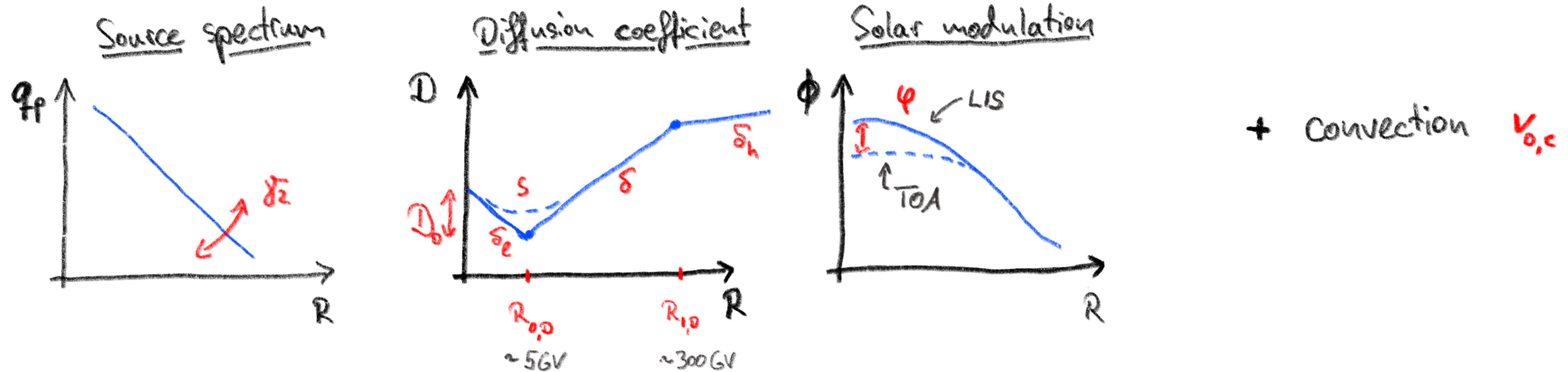
We use the **GALPROP** code to solve them.



# CR propagation models

We explore **5 different setups** for CR propagation:

**BASE    BASE+ $v_A$     BASE+inj    BASE+inj+ $v_A$     BASE+inj+ $v_A$ -diff.brk**

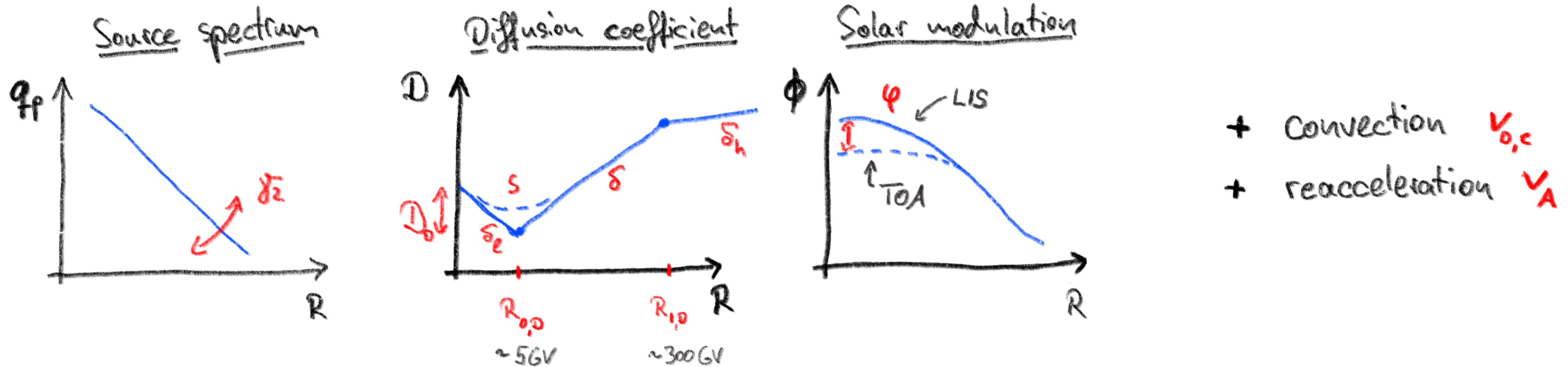


$$\frac{d\psi}{dt} = q(\mathbf{x}, p) + \nabla \cdot (D_{xx} \nabla \psi - \mathbf{V} \psi) + \frac{\partial}{\partial p} p^2 D_{pp} \frac{\partial}{\partial p} \frac{1}{p^2} \psi - \frac{\partial}{\partial p} \left( \frac{dp}{dt} \psi - \frac{p}{3} \nabla \cdot \mathbf{V} \psi \right) - \frac{1}{\tau_f} \psi - \frac{1}{\tau_r} \psi$$

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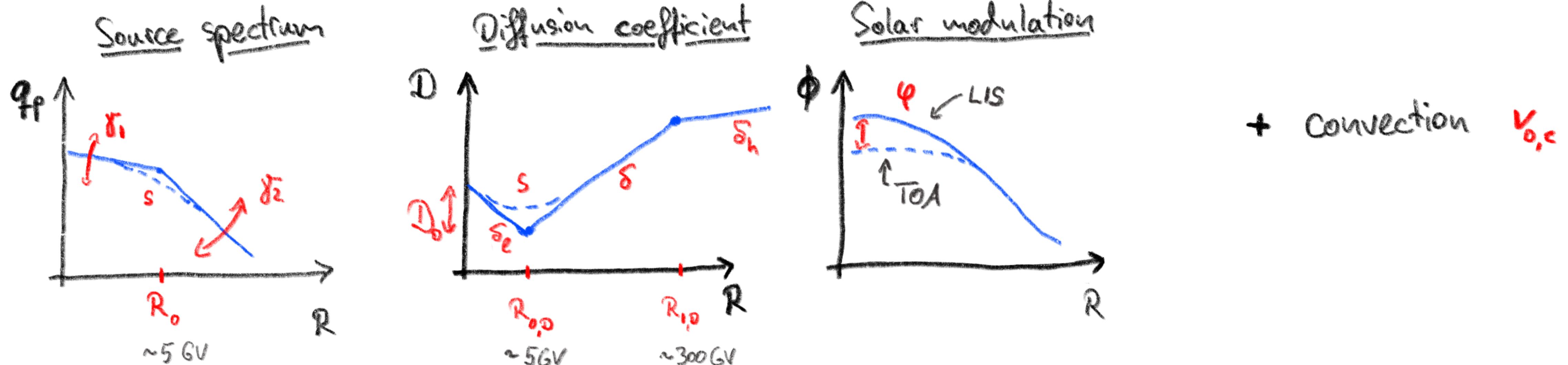


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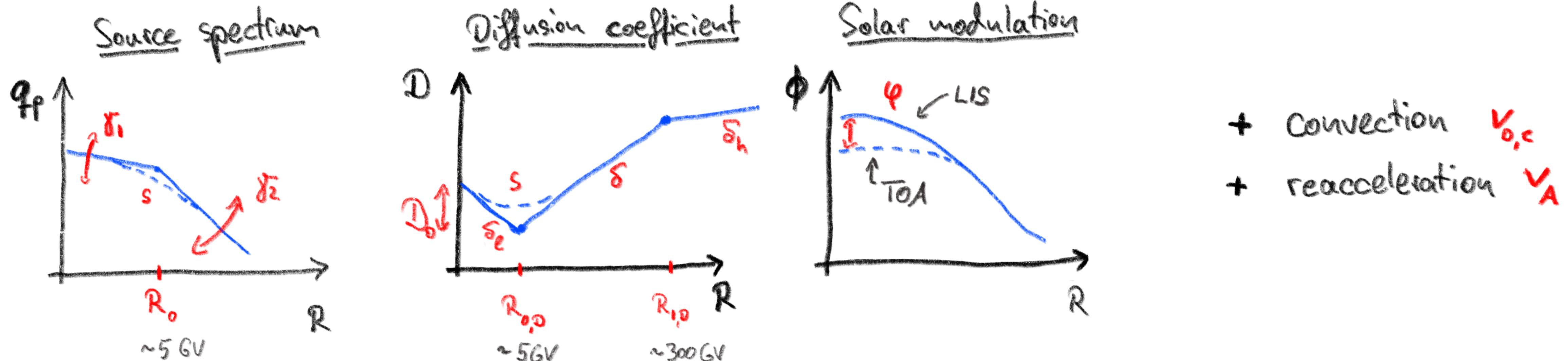


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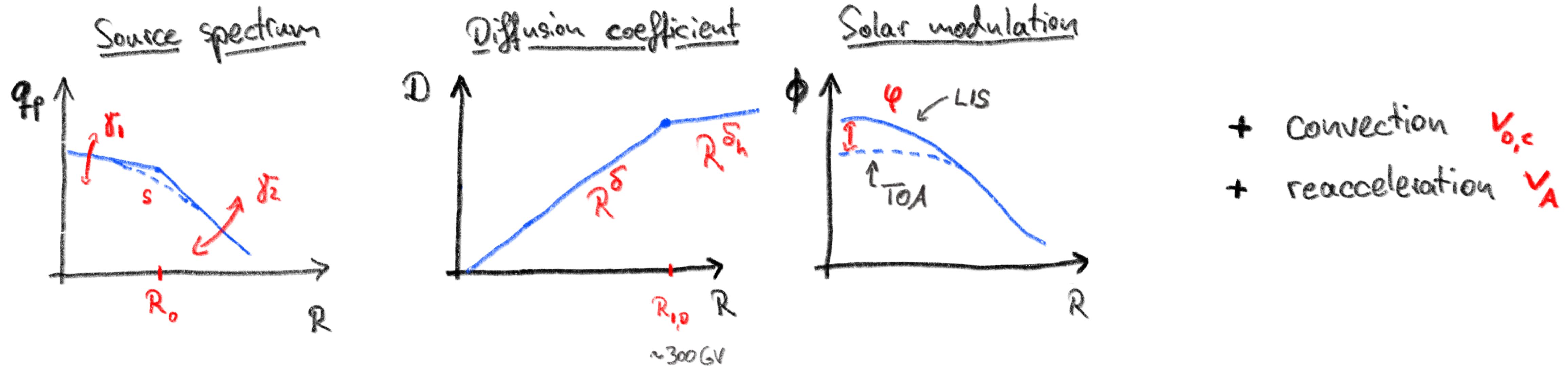


$$\frac{d\psi}{dt} = q(\mathbf{x}, p) + \nabla \cdot (D_{xx} \nabla \psi - \mathbf{V} \psi) + \frac{\partial}{\partial p} p^2 D_{pp} \frac{\partial}{\partial p} \frac{1}{p^2} \psi - \frac{\partial}{\partial p} \left( \frac{dp}{dt} \psi - \frac{p}{3} \nabla \cdot \mathbf{V} \psi \right) - \frac{1}{\tau_f} \psi - \frac{1}{\tau_r} \psi$$

# CR propagation models

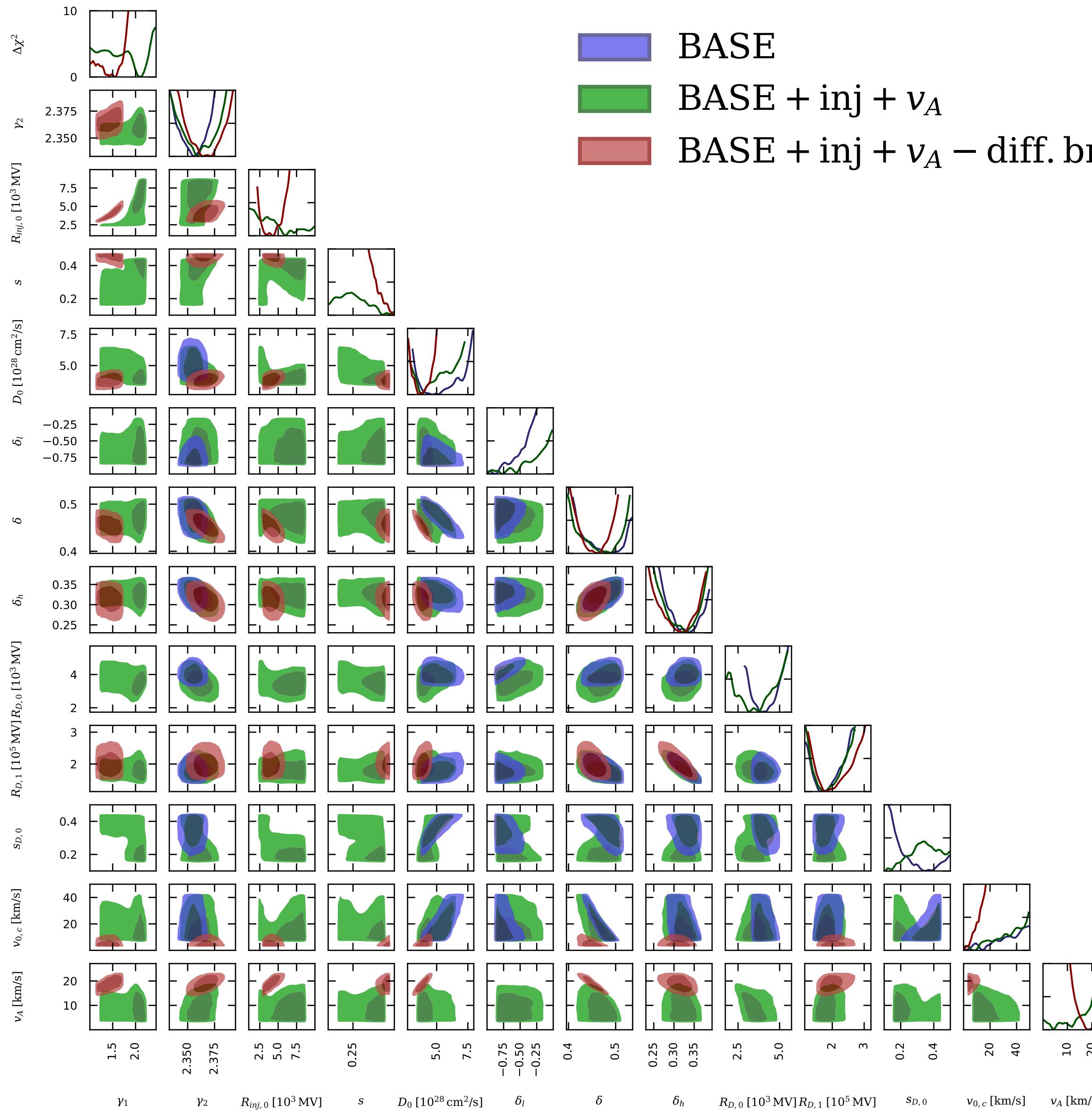
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# Global fit

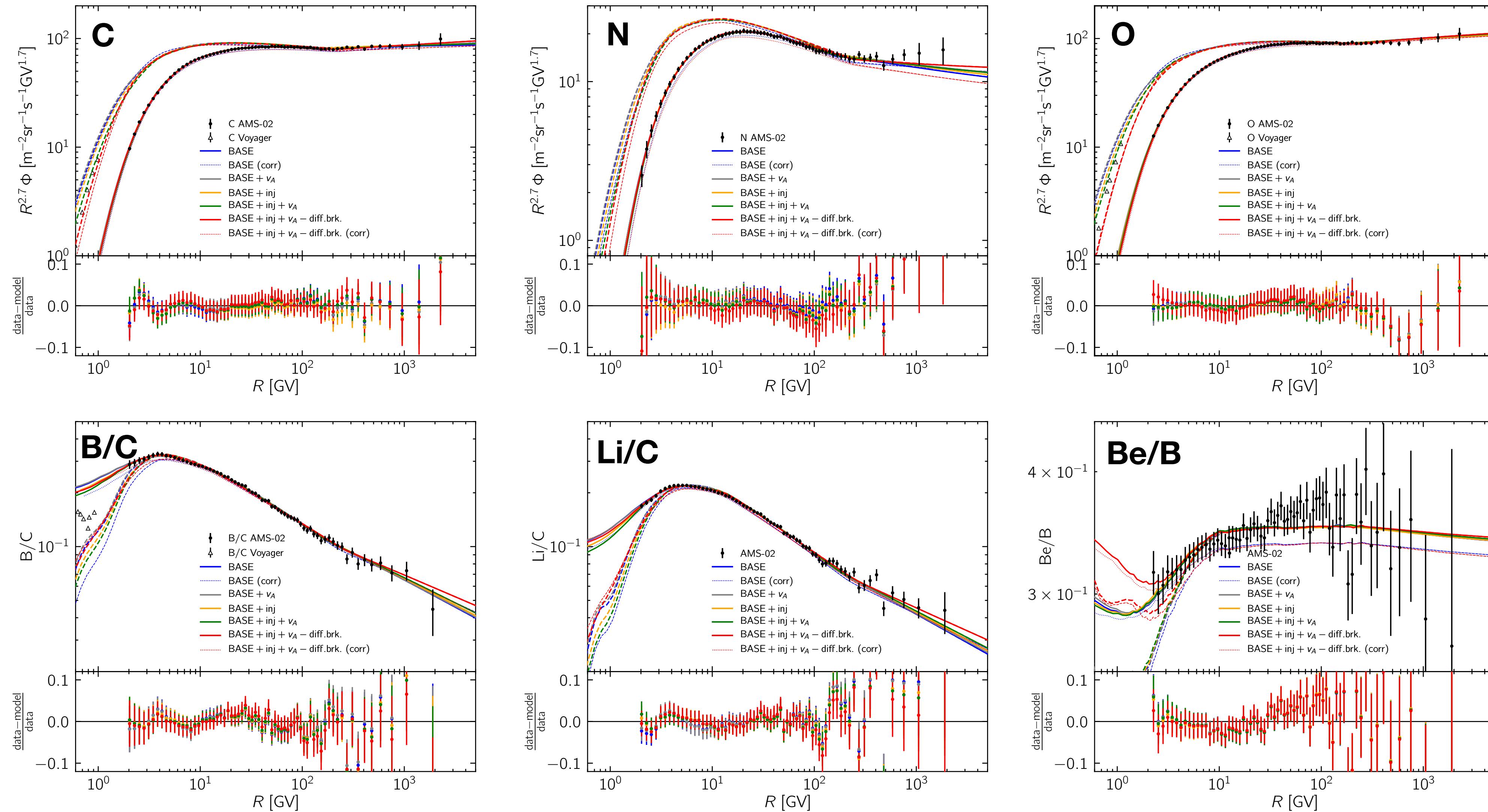


We investigate five propagation setups and perform several consistency checks

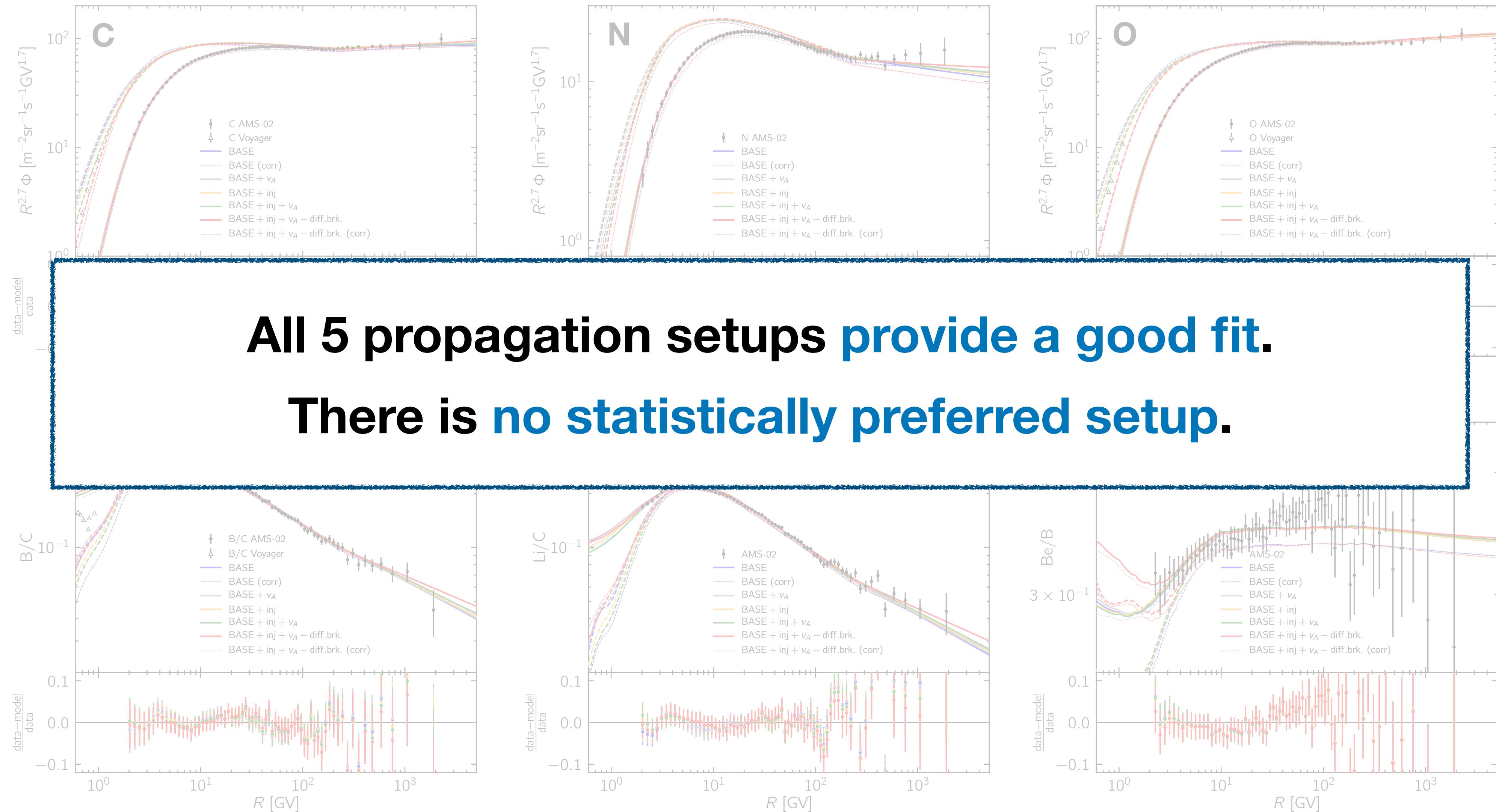
We use MultiNest to sample the large parameter space of up to 27 parameters

Parameters for CR propagation and cross section nuisance parameters are sampled at the same time

# Results of the global fits

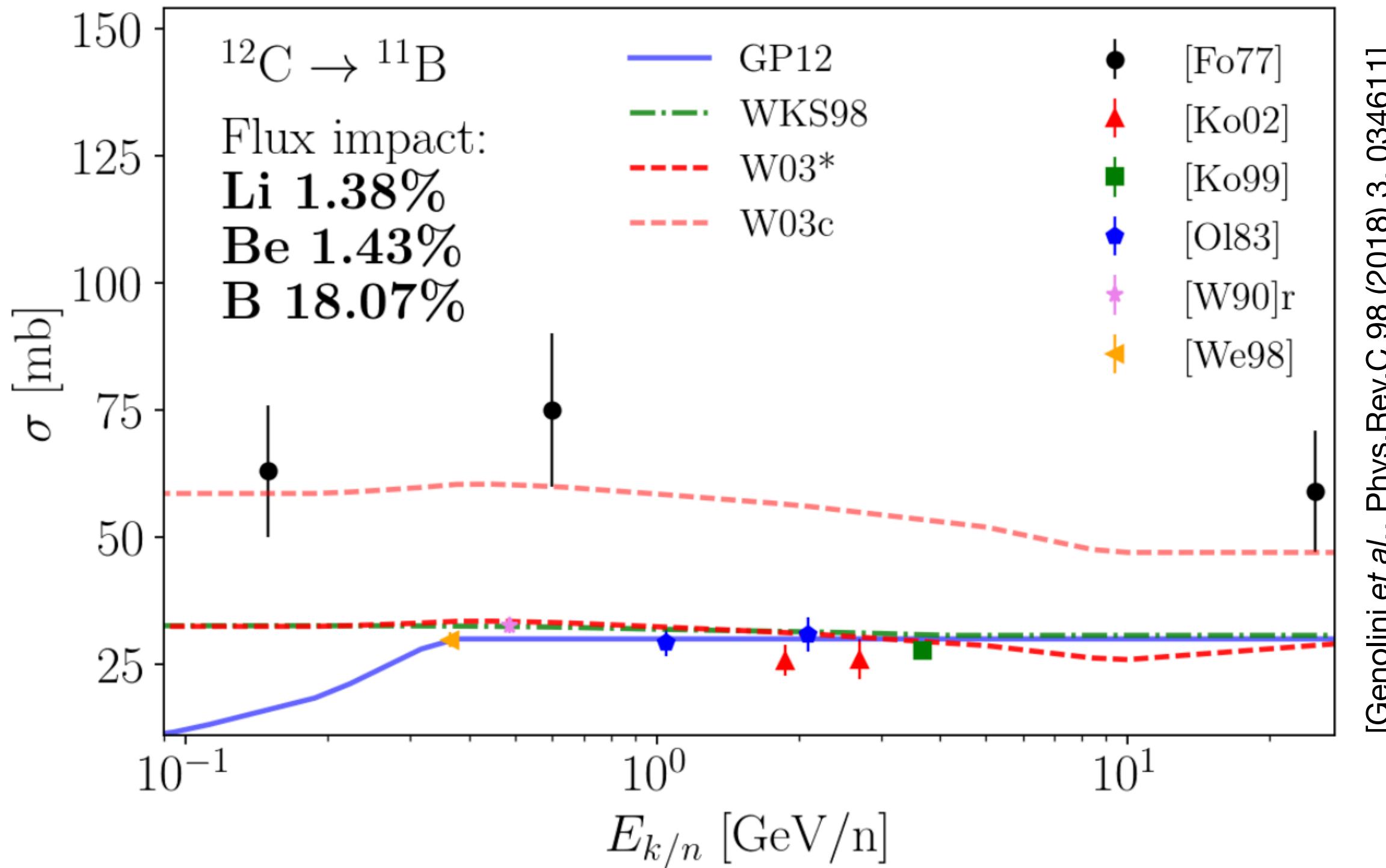


# Results of the global fits



# Systematic uncertainty: fragmentation cross sections

## Example: Fragmentation of $^{12}\text{C}$ to $^{11}\text{B}$



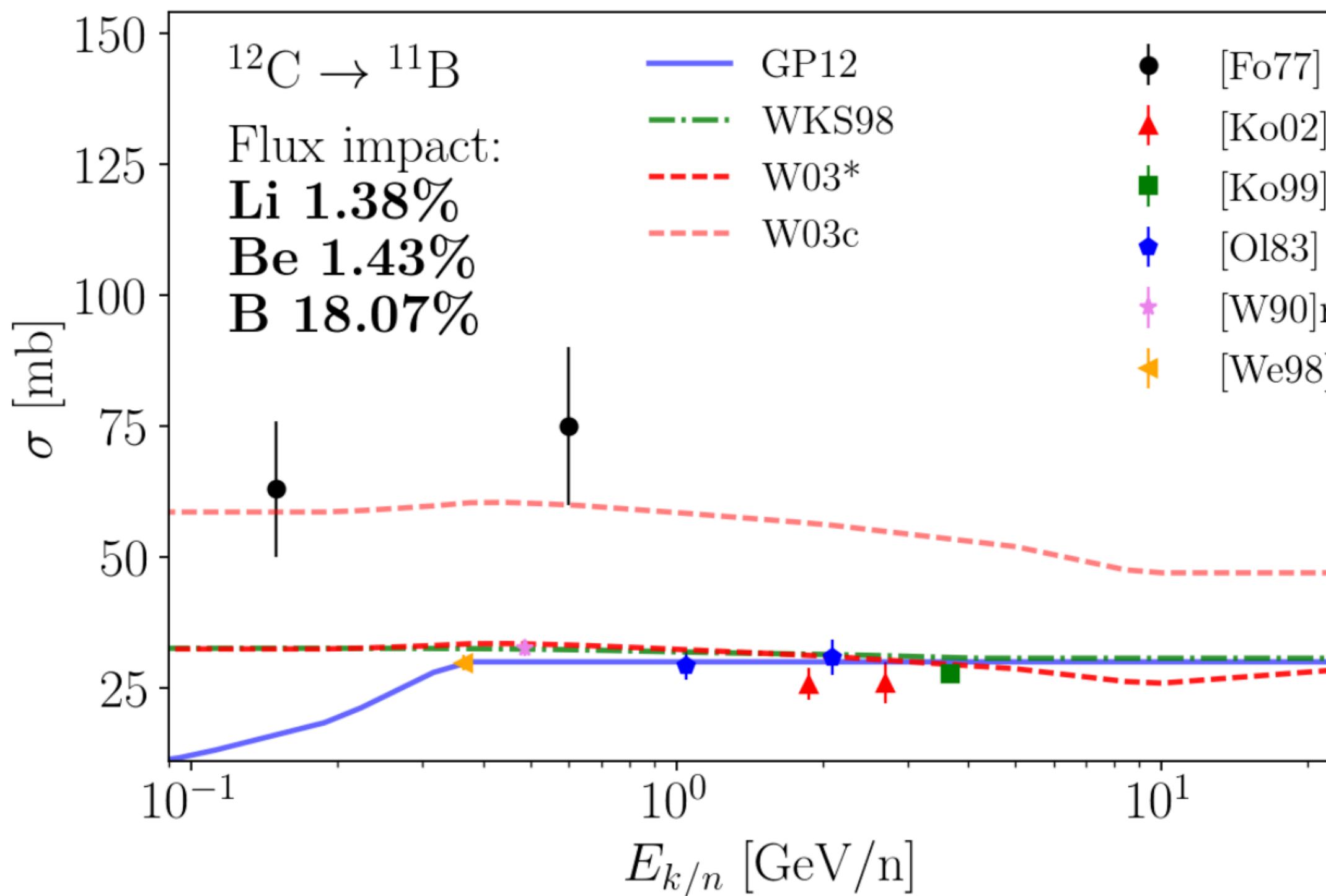
[Genolini et al., Phys. Rev. C 98 (2018) 3, 034611]

**Systematic uncertainties in the fragmentation cross sections are larger than those in the measured CR spectra!**

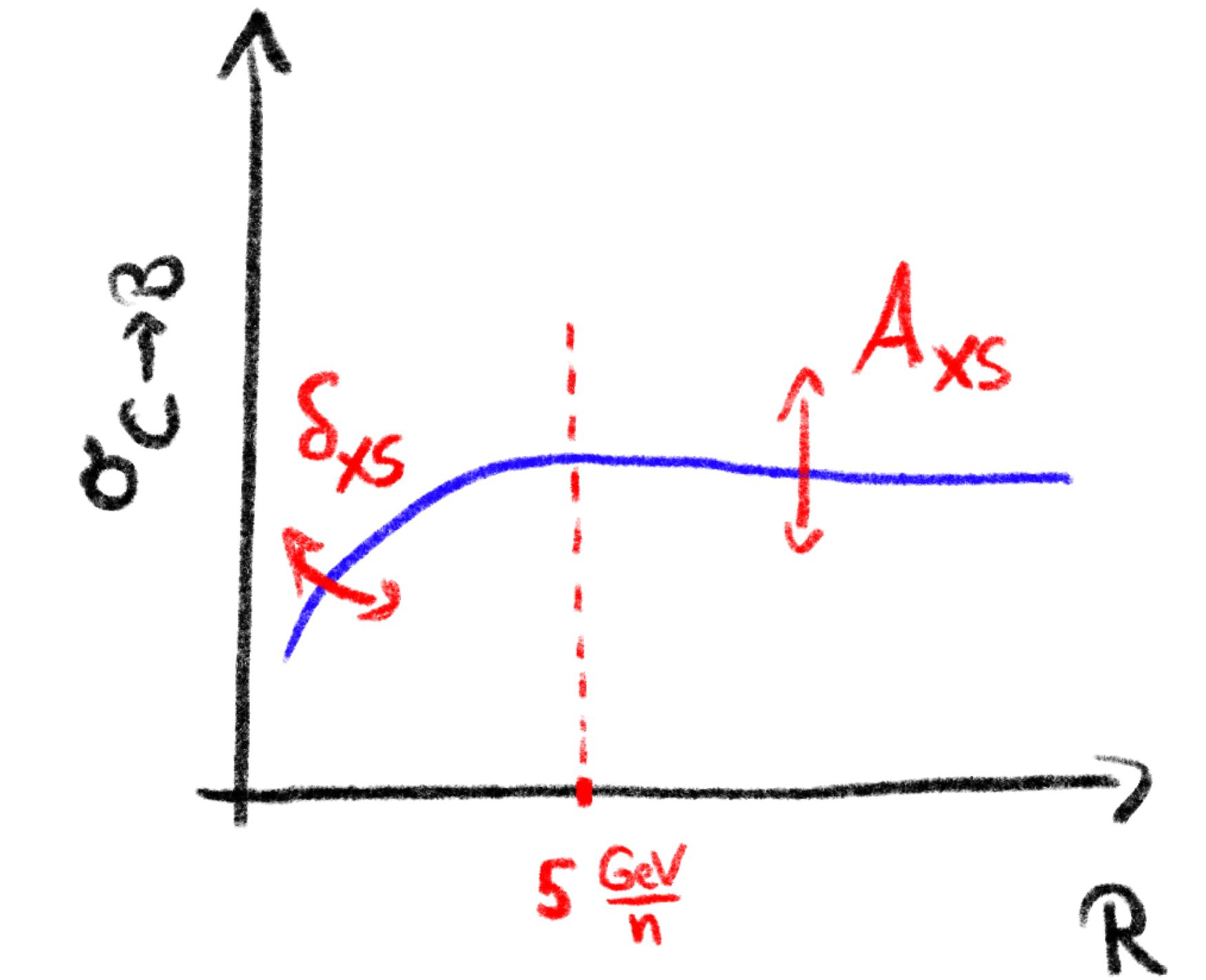
→ See also: Talk by P. De la Torre Luque

# Systematic uncertainty: fragmentation cross sections

## Example: Fragmentation of $^{12}\text{C}$ to $^{11}\text{B}$



[Genolini et al., Phys. Rev. C 98 (2018) 3, 034611]



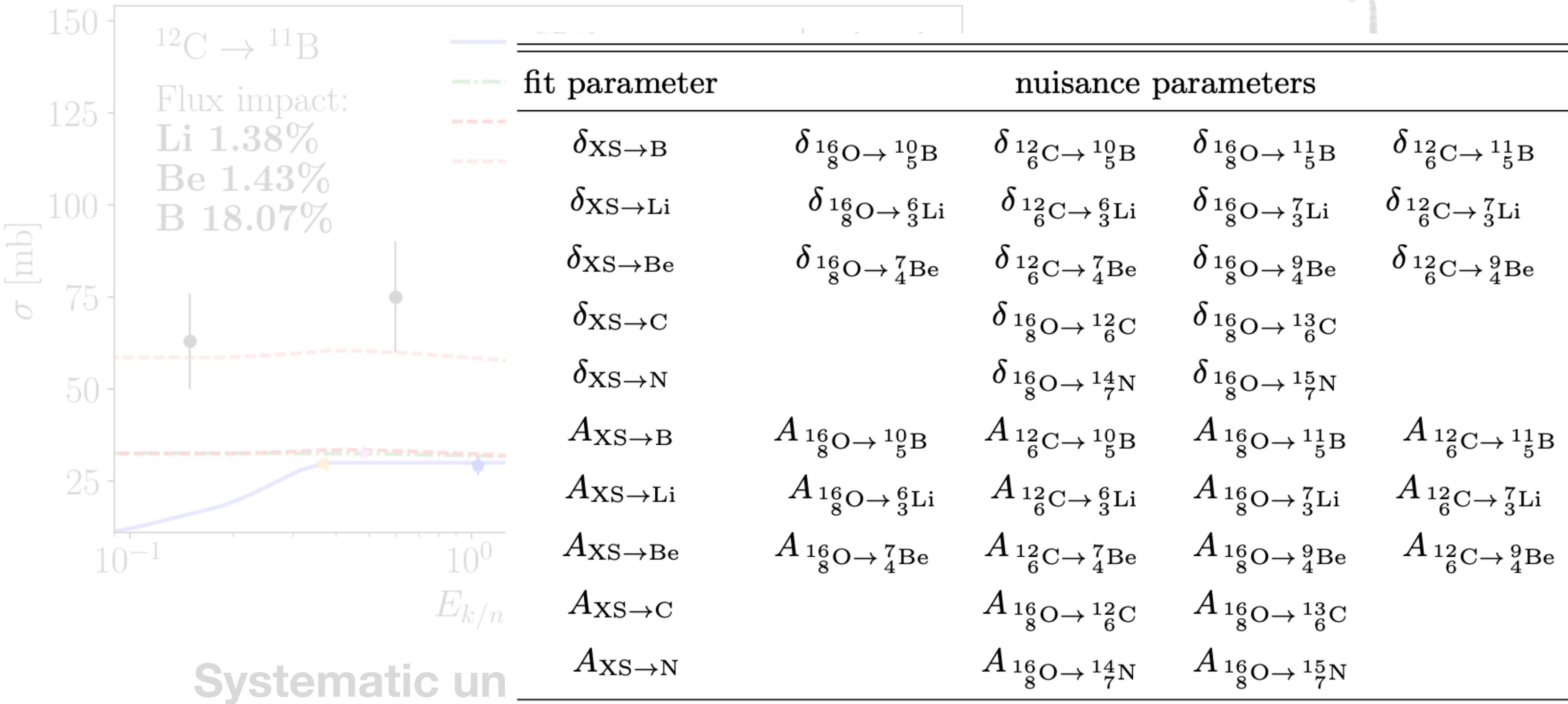
**Systematic uncertainties in the fragmentation cross sections are larger than those in the measured CR spectra!**

We perform a **global fit and profile over nuisance parameters** in the most relevant fragmentation cross sections.

→ See also: Talk by P. De la Torre Luque

# Systematic uncertainty: fragmentation cross sections

## Example: Fragmentation of $^{12}\text{C}$ to $^{11}\text{B}$



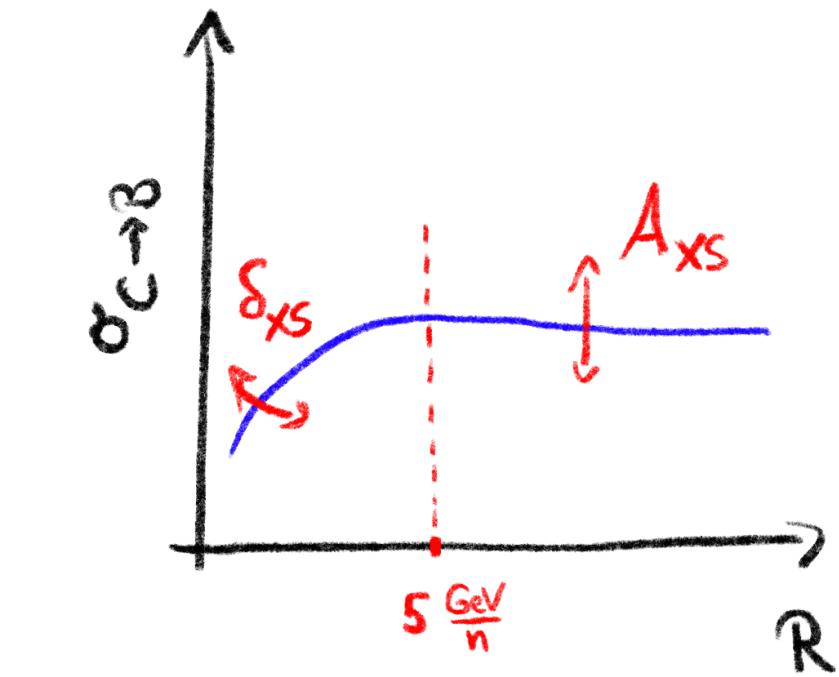
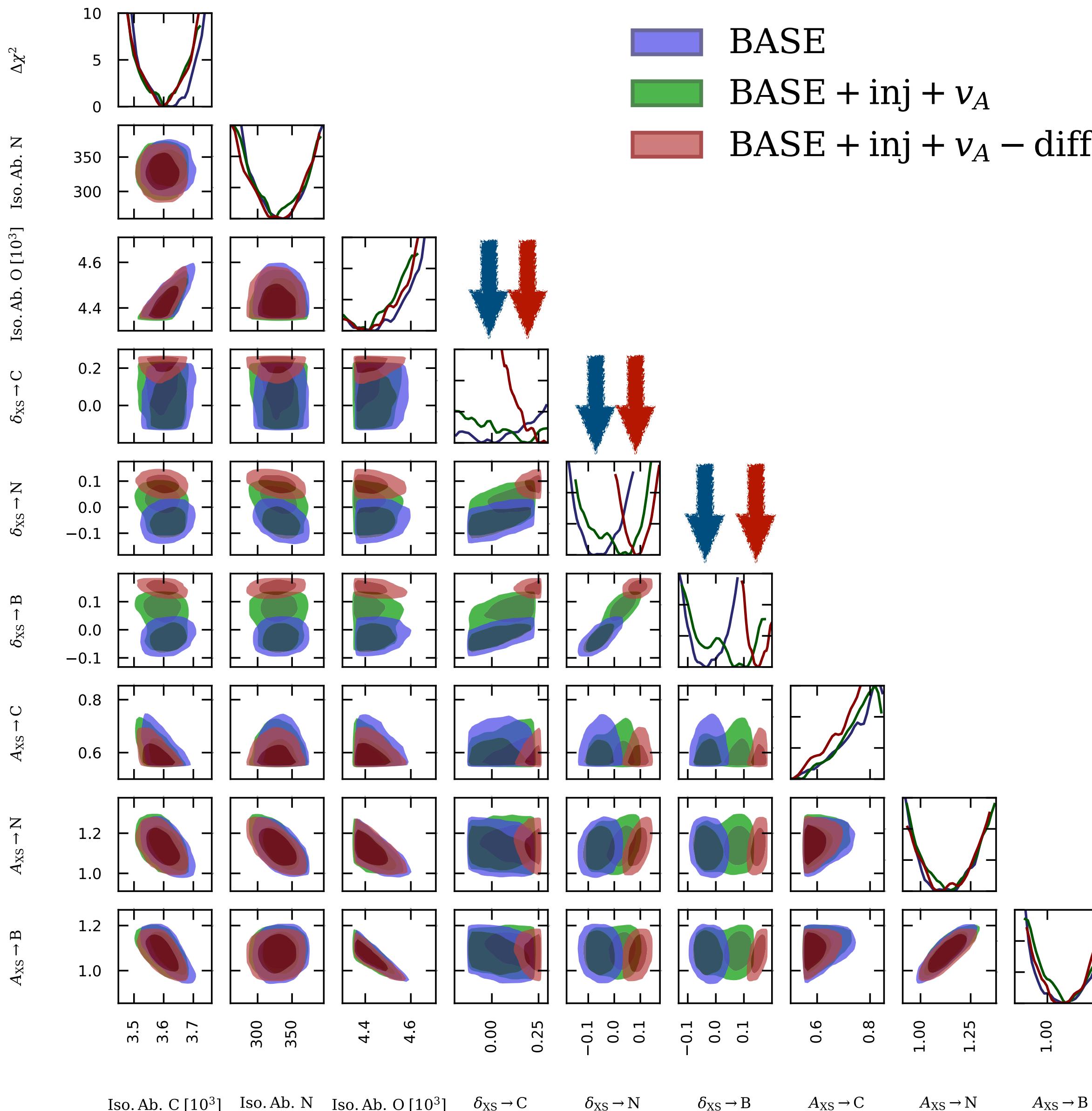
Systematic un

fragmentation cross sections are larger  
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nuisance parameters in the most  
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→ See also: Talk by P. De la Torre Luque

# Cross section nuisance parameters



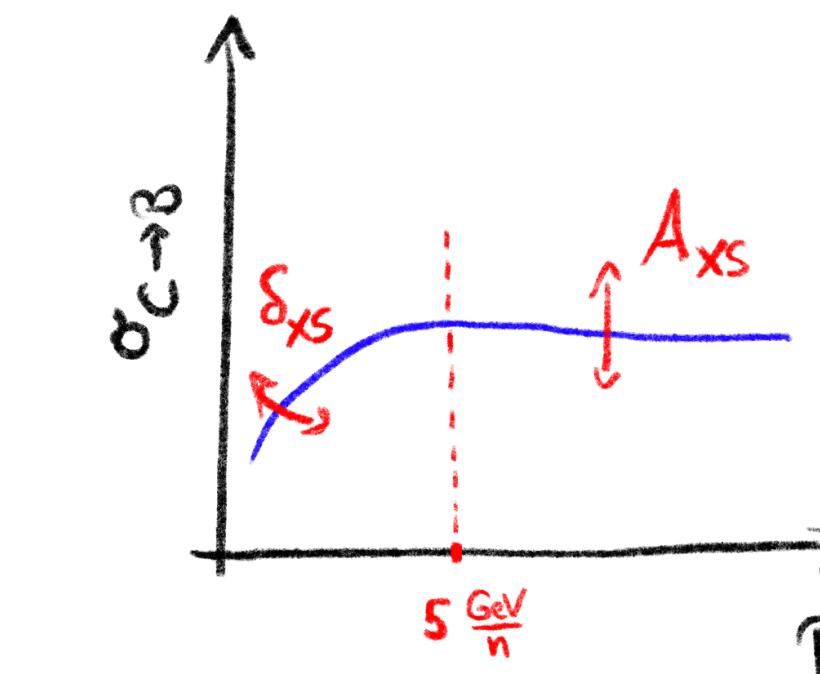
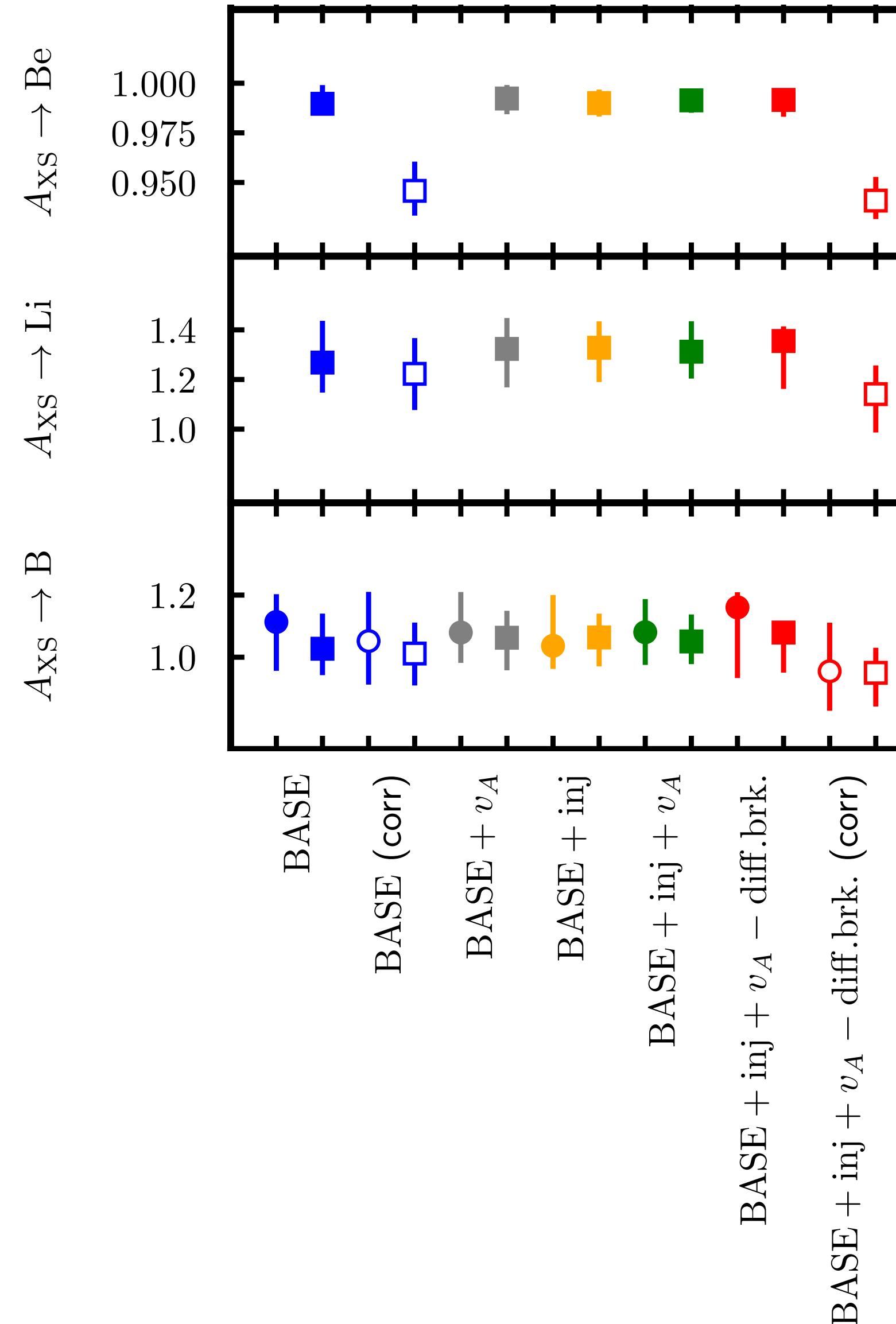
The default cross section parametrization is "[GALPROP 12](#)"

**BASE** is compatible with the default cross section

**BASE+inj+v<sub>A</sub>-diff.brk** converges at  $\delta_{XS} \sim 0.2$

**Li cross section are increased by ~25%**

# Cross section nuisance parameters



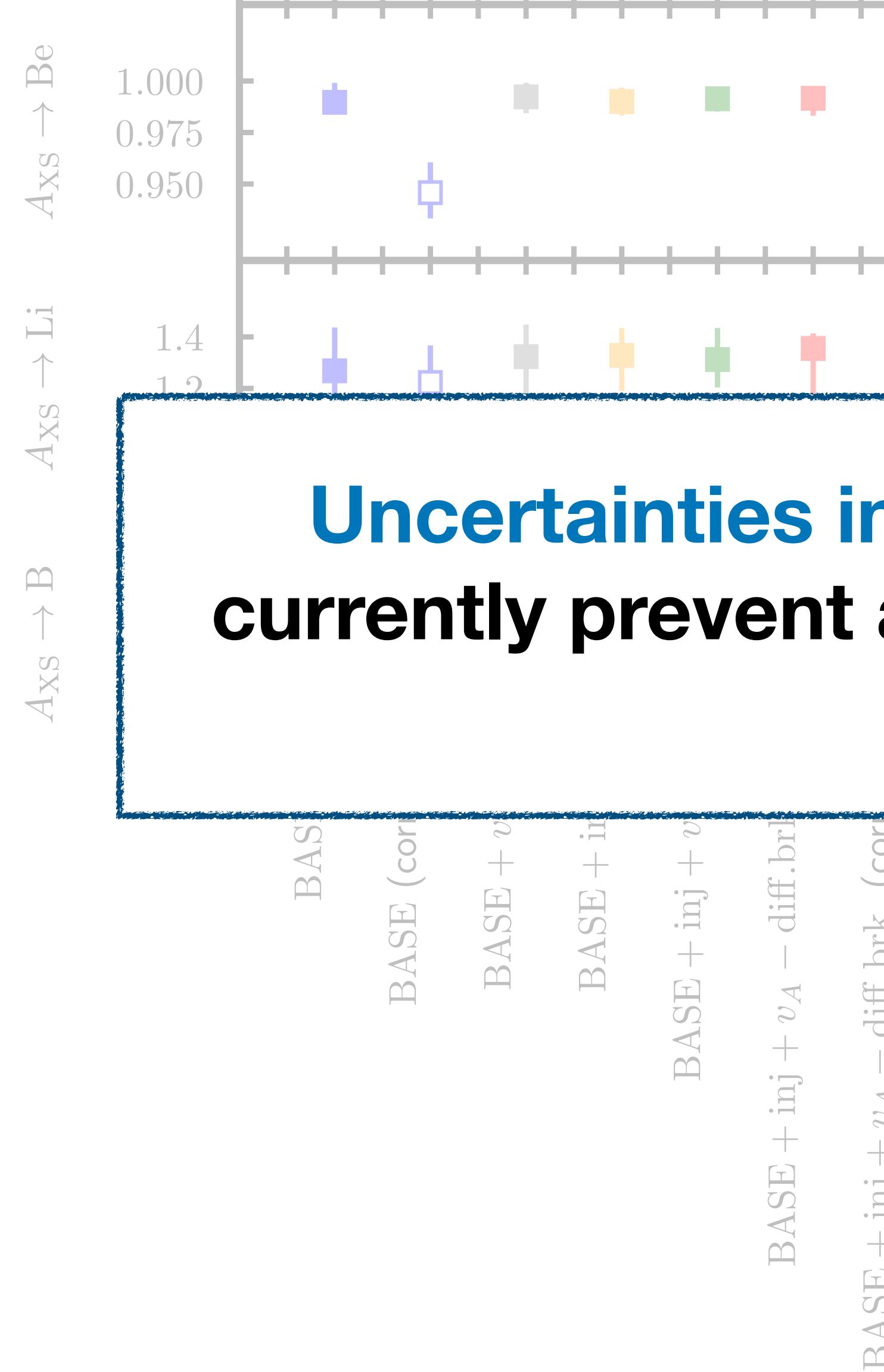
The default cross section  
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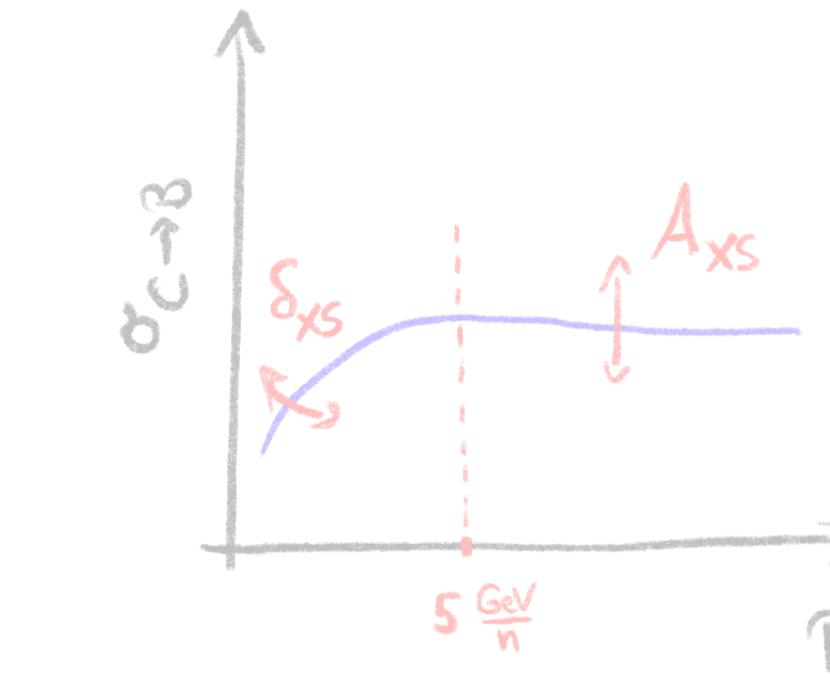
**BASE+inj+v<sub>A</sub>-diff.brk**  
converges at  $\delta_{\text{XS}} \sim 0.2$

**Li cross section are  
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# Cross section nuisance parameters



**Uncertainties in the fragmentation cross sections currently prevent a better understanding of cosmic ray propagation.**

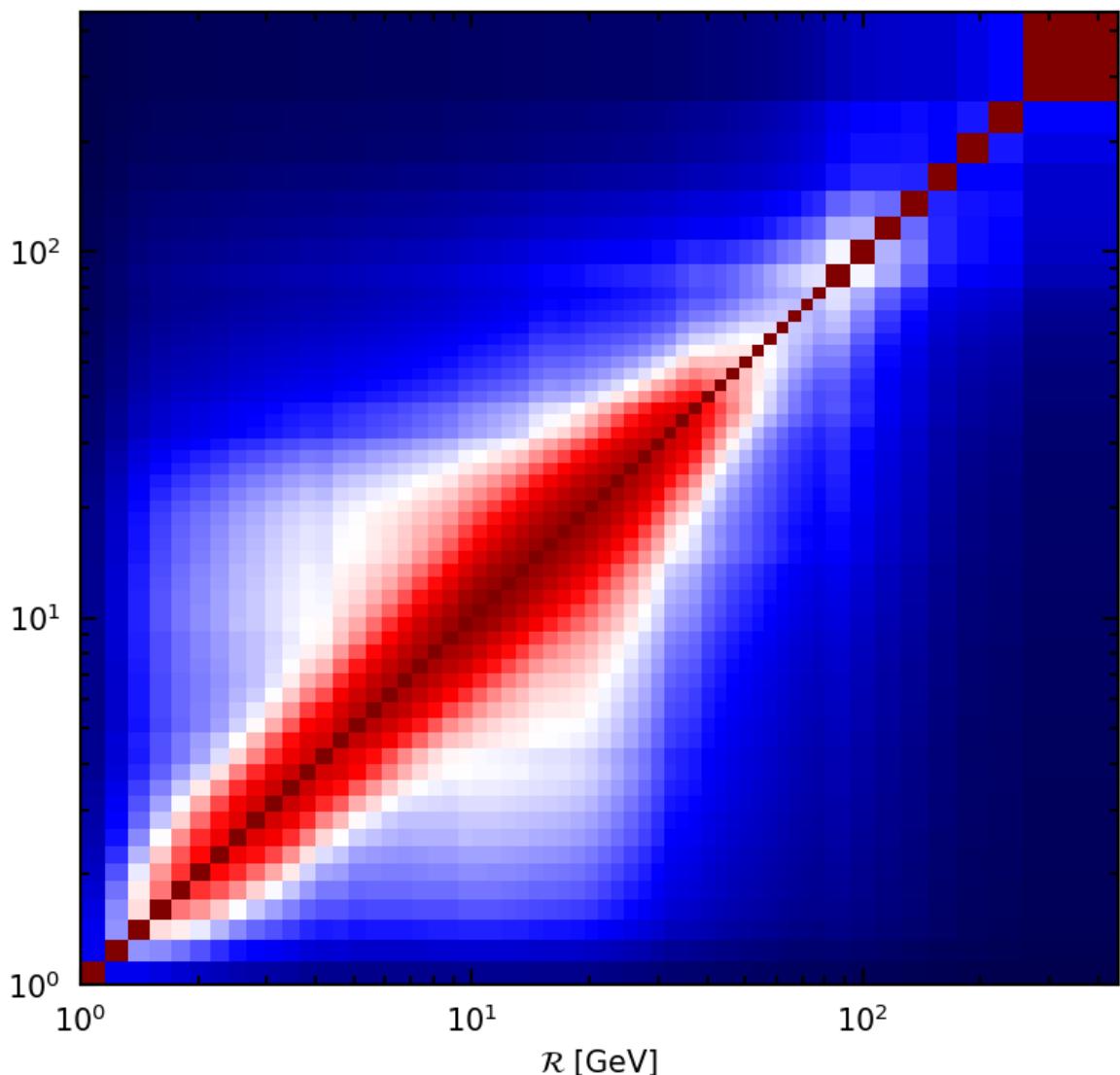
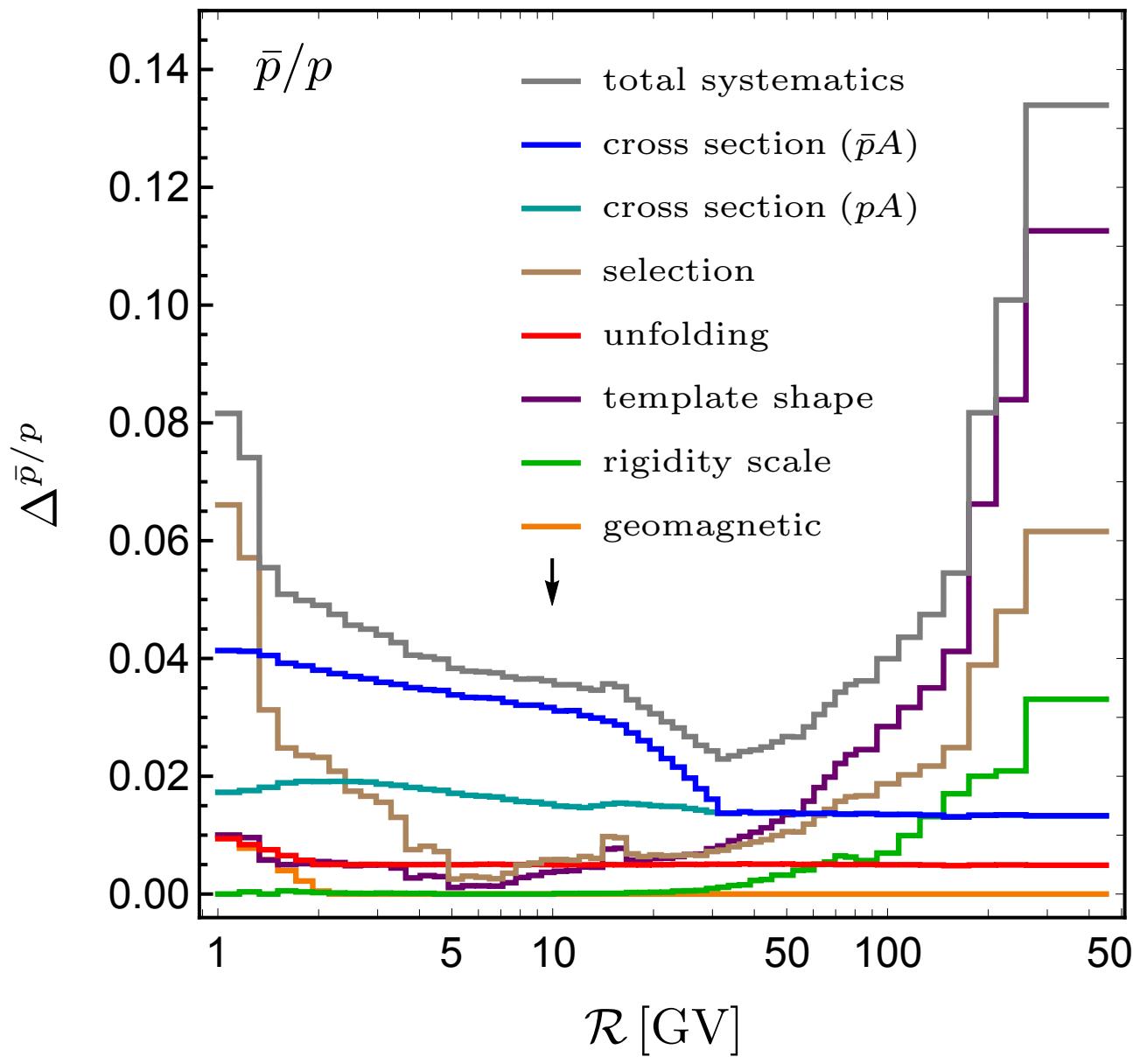


**BASE+inj+v<sub>A</sub>-diff.brk converges at  $\delta_{XS} \sim 0.2$**

**Li cross section are increased by  $\sim 25\%$**

# Correlation in the cosmic-ray data of AMS-02

[Heisig, MK, Winkler; PRR; 2020]



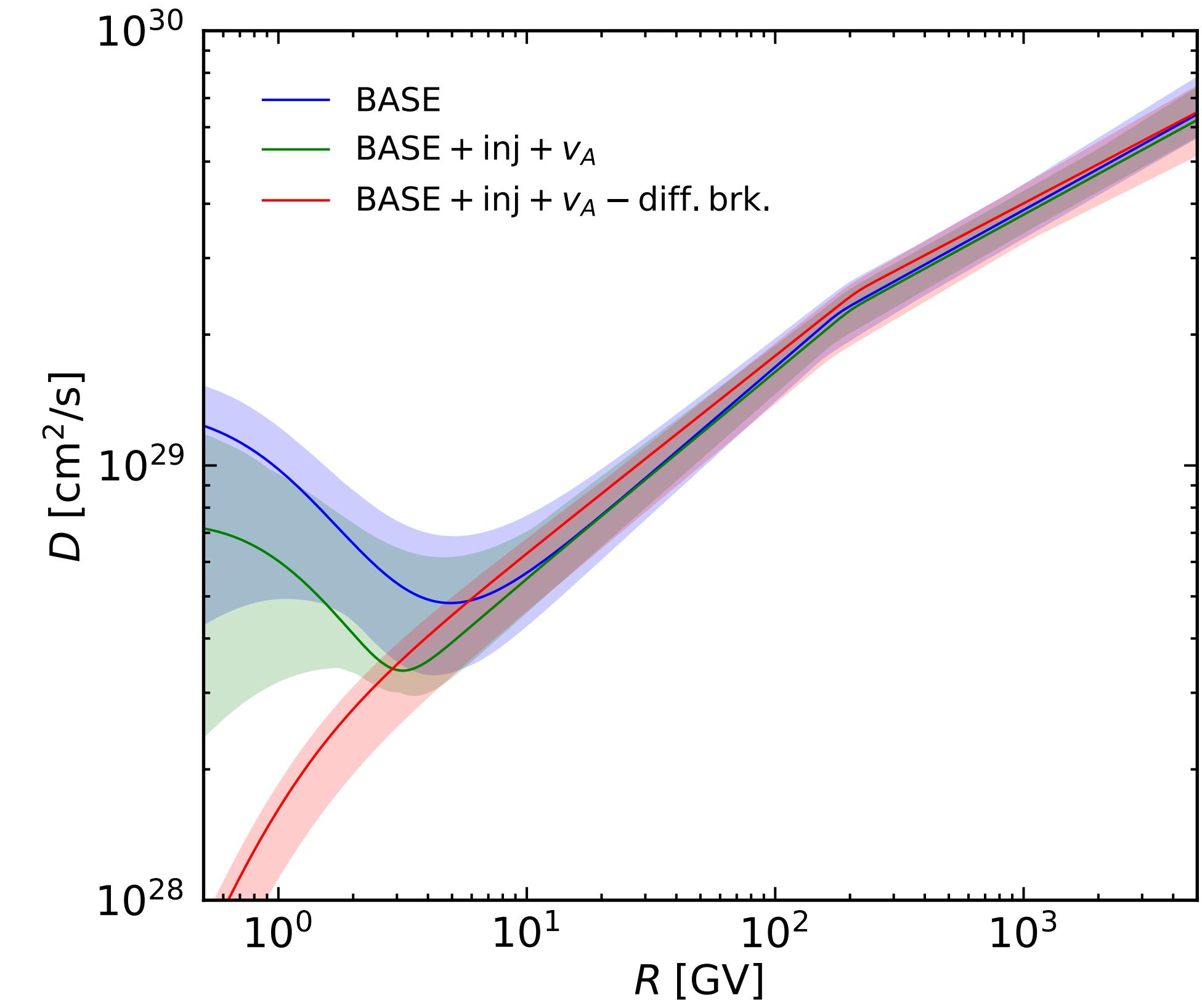
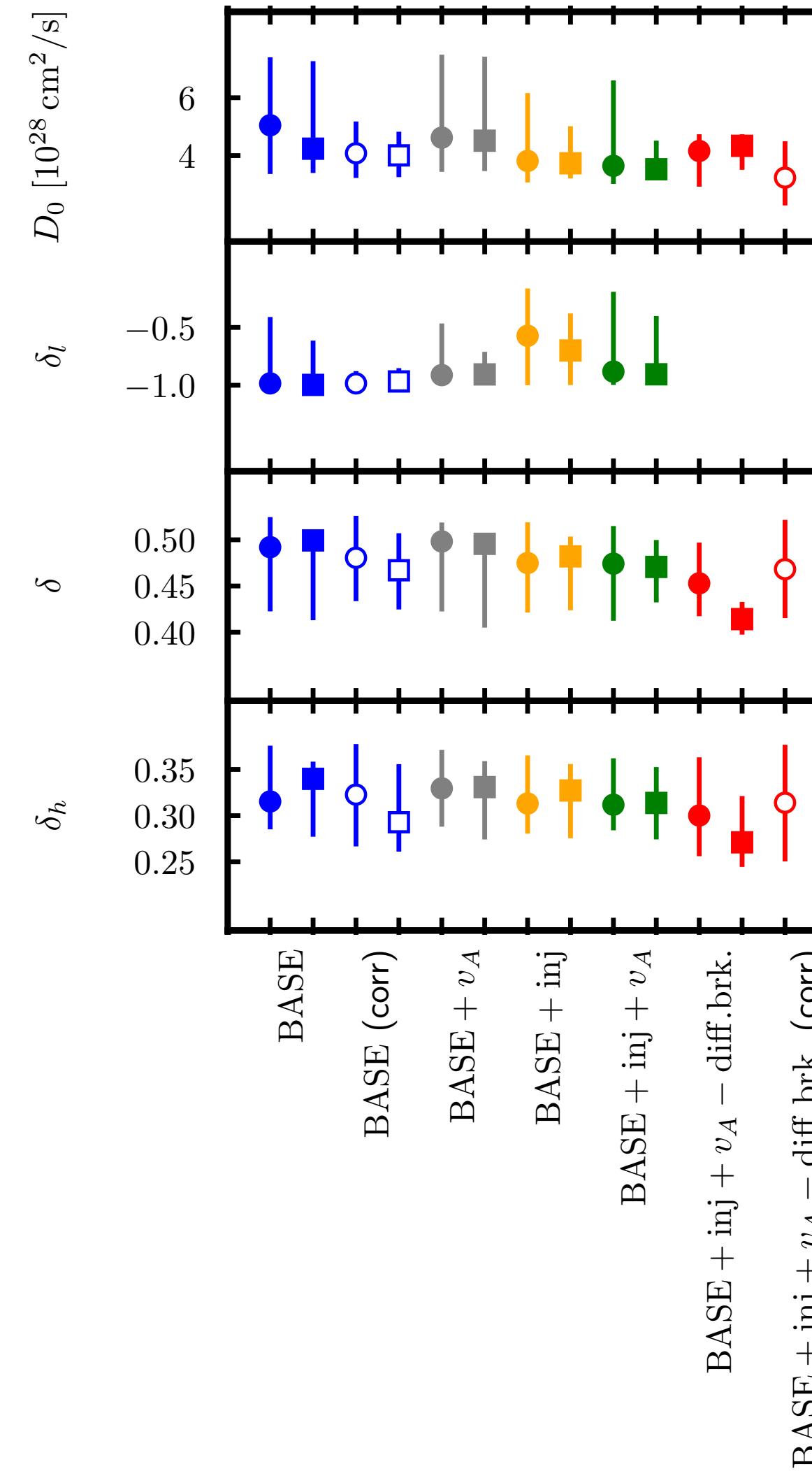
The **AMS-02 collaboration does not provide the correlation of the flux data points**

We **model the covariance matrix by splitting the systematic uncertainties into separate contributions and attributing a correlation length to each contribution**

The **inclusion of correlation does not change our conclusions!**

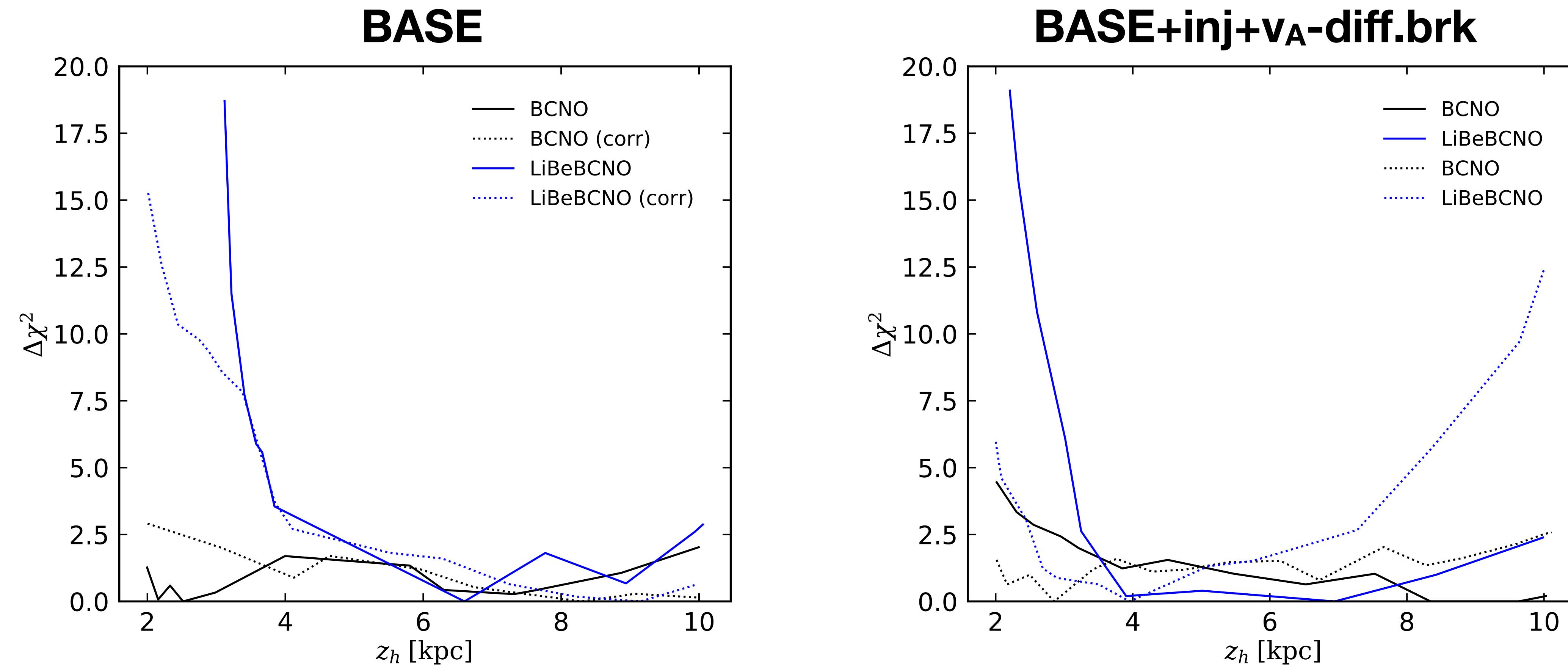
$$\mathcal{V}_{ij} = \sigma_i \sigma_j \exp \left( -\frac{1}{2} \left( \frac{R_i - R_j}{\ell_{\text{corr}}} \right)^2 \right)$$

# Parameter constraints



**The diffusion coefficient is well constrained above 10 GV**

# Parameter constraints



The combination of B and Be data allows to constrain  $z_h$

# Conclusions

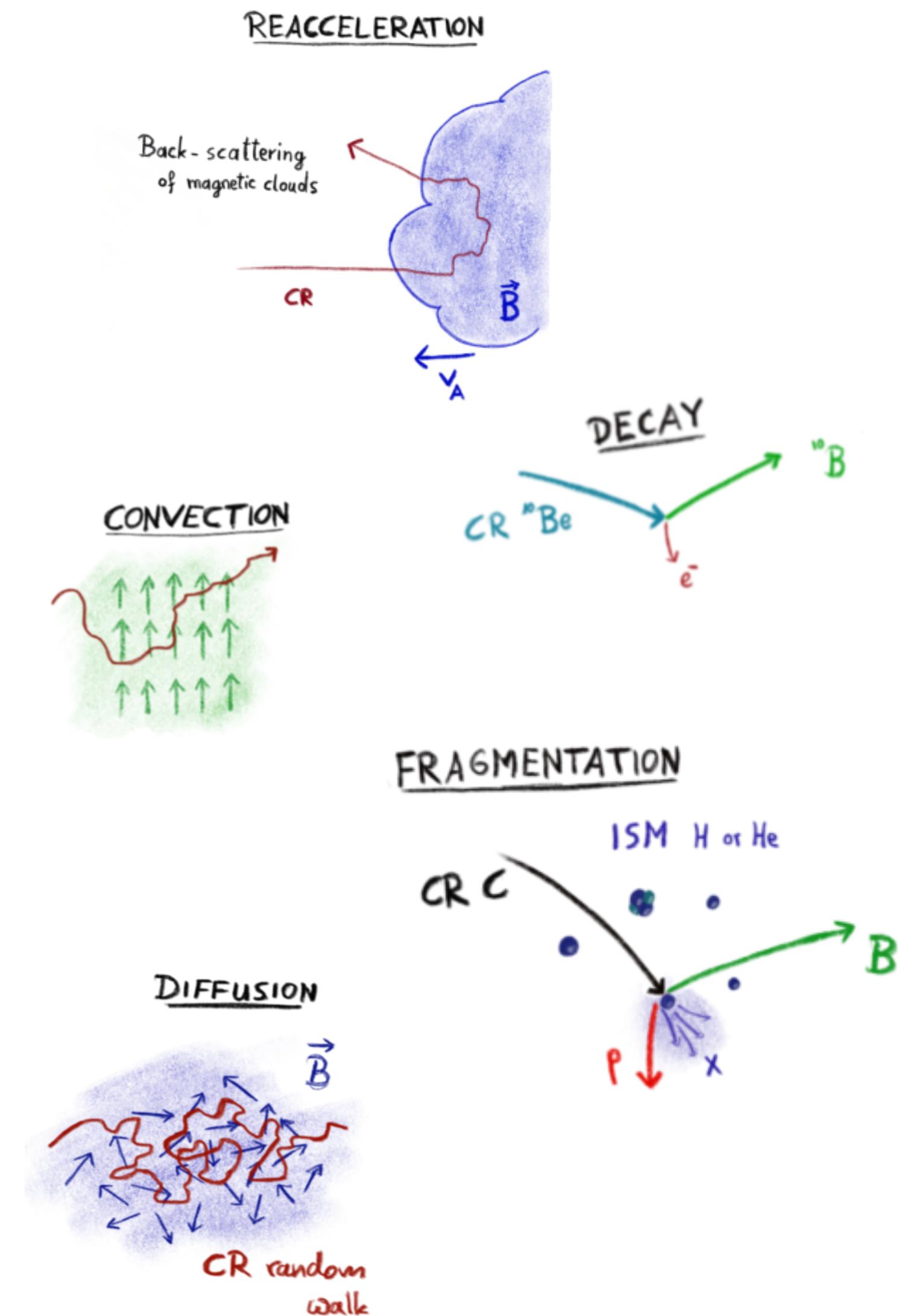
Precision data requires **precision analysis** and an honest treatment of systematic uncertainties

CR nuclei from **Li to O** are consistent with the traditional CR diffusion models

There is **no clear preference** for one CR propagation model because of uncertainties in the secondary fragmentation cross sections

Small halo heights of  $z_h < 3$  kpc are excluded

The diffusion coefficient is well constrained above 10 GeV



# **Backup**

# Free parameters

TABLE III. Summary of free CR parameters in the five different frameworks adopted to describe CR propagation.

	BASE	BASE+ $v_A$	BASE+inj	BASE+inj+ $v_A$	BASE+inj+ $v_A$ -diff.brk.	prior
$\gamma_1$	$\gamma_1 = \gamma_2$	$\gamma_1 = \gamma_2$	free	free	free	[0.0, 2.0]
$\gamma_2$	free	free	free	free	free	[2.1, 2.5]
$R_{\text{inj},0}$	-	-	free	free	free	[1, 10] GV
$s$	-	-	free	free	free	[0.1, 0.7]
$D_0$	free	free	free	free	free	[1e28, 1e29] $\text{cm}^2\text{s}^{-1}$
$\delta_l$	free	free	free	free	$\delta_l = \delta$	[-1, 0]
$\delta$	free	free	free	free	free	[0.2, 0.7]
$\delta_h$	free	free	free	free	free	[0.2, 0.7]
$R_{D,0}$	free	free	free	free	-	[1, 10] GV
$R_{D,1}$	free	free	free	free	free	[1e5, 5e5] GV
$s_{D,0}$	free	free	free	free	-	[0.1, 0.7]
$v_{0,c}$	free	free	free	free	free	[0, 50] km/s
$v_A$	-	free	-	free	free	[0, 50] km/s
Iso. Ab. $^{12}_6\text{C}$	free	free	free	free	free	[3300, 4000]
Iso. Ab. $^{14}_7\text{N}$	free	free	free	free	free	[200, 500]
Iso. Ab. $^{16}_8\text{O}$	free	free	free	free	free	[4200, 5000]
$\varphi_{\text{AMS-02}}$	free	free	free	free	free	$600 \pm 30$ MV
#par	13	14	16	17	14	

# Best fit table

TABLE IV. Fit results. For all 14 fits we report the total  $\chi^2$ , the contribution to the  $\chi^2$  from each single species, the number of degrees of freedom, and the best-fit value and  $1\sigma$  error for each parameter.

	BASE		BASE(corr)		BASE + $v_A$		BASE + inj		BASE + inj + $v_A$		BASE + inj + $v_A$ - diff.brk.		BASE + inj + $v_A$ - diff.brk. (corr)	
data set	BCNO	LiBeBCNO	BCNO	LiBeBCNO	BCNO	LiBeBCNO	BCNO	LiBeBCNO	BCNO	LiBeBCNO	BCNO	LiBeBCNO	BCNO	LiBeBCNO
#dof	252	383	252	383	251	382	249	380	248	379	251	382	251	382
$\chi^2$	72.4	170.0	423.2	593.6	72.8	169.0	67.9	160.7	67.3	158.9	74.2	168.9	415.4	590.2
$\chi_N^2$	15.9	18.7	148.1	146.9	19.3	17.2	14.8	15.2	19.3	16.5	17.2	20.0	151.5	149.1
$\chi_O^2$	14.0	13.7	61.8	63.0	11.3	14.8	12.8	13.8	11.2	12.4	14.8	12.1	62.1	62.2
$\chi_C^2$	13.1	15.2	127.8	124.9	12.7	12.7	16.0	16.9	11.7	15.6	13.7	15.4	122.4	122.5
$\chi_{Be/B}^2$	-	42.2	-	82.9	-	42.0	-	43.3	-	42.4	-	40.6	-	83.3
$\chi_{Li/C}^2$	-	46.5	-	82.5	-	47.5	-	39.1	-	39.2	-	41.2	-	85.3
$\chi_{B/C}^2$	27.8	30.2	78.4	80.5	28.7	29.5	24.1	26.9	24.0	28.1	25.8	33.4	75.8	82.7
$\gamma_1$	-	-	-	-	-	-	$2.18^{+0.04}_{-0.51}$	$2.21^{+0.04}_{-0.07}$	$2.08^{+0.10}_{-0.30}$	$2.10^{+0.10}_{-0.06}$	$1.20^{+0.42}_{-0.16}$	$1.64^{+0.04}_{-0.07}$	$1.15^{+0.08}_{-0.12}$	$1.14^{+0.24}_{-0.11}$
$\gamma_2$	$2.357^{+0.003}_{-0.005}$	$2.365^{+0.005}_{-0.002}$	$2.34^{+0.01}_{-0.01}$	$2.360^{+0.014}_{-0.009}$	$2.353^{+0.006}_{-0.004}$	$2.361^{+0.009}_{-0.002}$	$2.368^{+0.002}_{-0.017}$	$2.371^{+0.004}_{-0.005}$	$2.360^{+0.008}_{-0.004}$	$2.378^{+0.003}_{-0.005}$	$2.362^{+0.016}_{-0.004}$	$2.389^{+0.005}_{-0.004}$	$2.365^{+0.008}_{-0.020}$	$2.373^{+0.013}_{-0.005}$
$R_{inj,0}$ [10 <sup>3</sup> MV]	-	-	-	-	-	-	$8.85^{+1.15}_{-4.33}$	$8.31^{+0.91}_{-1.05}$	$6.20^{+1.87}_{-1.84}$	$6.98^{+2.10}_{-0.33}$	$3.28^{+1.82}_{-0.59}$	$5.18^{+0.65}_{-0.30}$	$2.93^{+0.20}_{-0.37}$	$2.61^{+1.01}_{-0.10}$
$s$	-	-	-	-	-	-	$0.48^{+0.02}_{-0.27}$	$0.45^{+0.04}_{-0.06}$	$0.45^{+0.03}_{-0.20}$	$0.487^{+0.006}_{-0.043}$	$0.490^{+0.009}_{-0.052}$	$0.493^{+0.007}_{-0.044}$	$0.39^{+0.10}_{-0.06}$	$0.494^{+0.005}_{-0.038}$
$D_0$ [10 <sup>28</sup> cm <sup>2</sup> /s]	$5.05^{+0.99}_{-1.34}$	$4.24^{+0.96}_{-0.44}$	$4.08^{+0.33}_{-0.55}$	$4.01^{+0.32}_{-0.45}$	$4.62^{+1.28}_{-0.46}$	$4.52^{+1.94}_{-0.39}$	$3.82^{+1.03}_{-0.30}$	$3.73^{+0.69}_{-0.20}$	$3.65^{+0.87}_{-0.18}$	$3.53^{+0.25}_{-0.10}$	$4.16^{+0.33}_{-0.88}$	$4.34^{+0.14}_{-0.66}$	$3.24^{+0.94}_{-0.36}$	$3.60^{+0.19}_{-0.33}$
$\delta_t$	$-0.98^{+0.22}_{-0.01}$	$-0.997^{+0.11}_{-0.001}$	$-0.98^{+0.03}_{-0.02}$	$-0.97^{+0.03}_{-0.02}$	$-0.91^{+0.11}_{-0.07}$	$-0.91^{+0.08}_{-0.09}$	$-0.57^{+0.22}_{-0.37}$	$-0.70^{+0.09}_{-0.28}$	$-0.88^{+0.36}_{-0.06}$	$-0.91^{+0.13}_{-0.04}$	-	-	-	-
$\delta$	$0.49^{+0.03}_{-0.04}$	$0.499^{+0.002}_{-0.033}$	$0.48^{+0.03}_{-0.01}$	$0.47^{+0.02}_{-0.01}$	$0.498^{+0.007}_{-0.045}$	$0.496^{+0.004}_{-0.056}$	$0.47^{+0.02}_{-0.03}$	$0.48^{+0.01}_{-0.03}$	$0.47^{+0.02}_{-0.03}$	$0.471^{+0.009}_{-0.014}$	$0.45^{+0.02}_{-0.02}$	$0.414^{+0.013}_{-0.005}$	$0.47^{+0.02}_{-0.04}$	$0.43^{+0.02}_{-0.01}$
$\delta_h$	$0.315^{+0.045}_{-0.008}$	$0.340^{+0.007}_{-0.033}$	$0.32^{+0.03}_{-0.02}$	$0.293^{+0.032}_{-0.009}$	$0.33^{+0.02}_{-0.02}$	$0.331^{+0.008}_{-0.027}$	$0.31^{+0.03}_{-0.02}$	$0.33^{+0.02}_{-0.03}$	$0.31^{+0.03}_{-0.01}$	$0.31^{+0.01}_{-0.01}$	$0.30^{+0.04}_{-0.02}$	$0.271^{+0.026}_{-0.007}$	$0.31^{+0.02}_{-0.03}$	$0.311^{+0.007}_{-0.044}$
$R_{D,0}$ [10 <sup>3</sup> MV]	$3.94^{+0.52}_{-0.35}$	$4.05^{+0.43}_{-0.14}$	$3.87^{+0.14}_{-0.12}$	$3.85^{+0.16}_{-0.05}$	$3.97^{+0.21}_{-0.36}$	$4.25^{+0.10}_{-0.35}$	$4.07^{+0.20}_{-0.53}$	$4.01^{+0.14}_{-0.37}$	$3.02^{+0.81}_{-0.23}$	$3.37^{+0.43}_{-0.41}$	-	-	-	-
$R_{D,1}$ [10 <sup>5</sup> MV]	$1.80^{+0.13}_{-0.30}$	$1.52^{+0.48}_{-0.08}$	$2.00^{+0.25}_{-0.22}$	$2.09^{+0.14}_{-0.42}$	$1.88^{+0.12}_{-0.35}$	$1.63^{+0.19}_{-0.07}$	$1.65^{+0.35}_{-0.13}$	$1.49^{+0.36}_{-0.06}$	$2.02^{+0.09}_{-0.46}$	$1.68^{+0.12}_{-0.08}$	$2.14^{+0.16}_{-0.40}$	$2.33^{+0.16}_{-0.46}$	$1.96^{+0.62}_{-0.11}$	$2.12^{+0.25}_{-0.29}$
$s_{D,0}$	$0.38^{+0.06}_{-0.11}$	$0.32^{+0.06}_{-0.07}$	$0.15^{+0.03}_{-0.02}$	$0.16^{+0.03}_{-0.01}$	$0.36^{+0.06}_{-0.07}$	$0.31^{+0.13}_{-0.05}$	$0.12^{+0.19}_{-0.02}$	$0.13^{+0.06}_{-0.02}$	$0.13^{+0.20}_{-0.02}$	$0.109^{+0.033}_{-0.004}$	-	-	-	-
$v_{0,c}$ [km/s]	$3.34^{+21.76}_{-2.49}$	$1.81^{+17.74}_{-0.70}$	$9.09^{+7.89}_{-8.68}$	$12.11^{+5.83}_{-6.91}$	$0.27^{+23.83}_{-0.06}$	$0.84^{+27.41}_{-0.22}$	$13.18^{+14.33}_{-12.26}$	$4.92^{+10.66}_{-4.85}$	$5.02^{+18.32}_{-2.27}$	$2.30^{+6.45}_{-1.31}$	$0.34^{+3.88}_{-0.23}$	$0.004^{+1.515}_{-0.000}$	$0.89^{+5.05}_{-0.75}$	$1.81^{+2.30}_{-1.63}$
$v_A$ [km/s]	-	-	-	-	$8.65^{+3.51}_{-7.81}$	$0.54^{+6.04}_{-0.24}$	-	-	$10.68^{+2.94}_{-9.29}$	$10.85^{+3.55}_{-4.79}$	$19.23^{+3.65}_{-3.77}$	$24.04^{+0.91}_{-2.90}$	$16.24^{+5.30}_{-1.35}$	$20.14^{+1.44}_{-1.49}$
Iso.Ab. C [10 <sup>3</sup> ]	$3.59^{+0.08}_{-0.02}$	$3.59^{+0.04}_{-0.02}$	$3.48^{+0.03}_{-0.14}$	$3.37^{+0.11}_{-0.06}$	$3.63^{+0.02}_{-0.04}$	$3.60^{+0.03}_{-0.02}$	$3.58^{+0.05}_{-0.03}$	$3.59^{+0.03}_{-0.04}$	$3.640^{+0.009}_{-0.068}$	$3.57^{+0.03}_{-0.02}$	$3.58^{+0.06}_{-0.04}$	$3.54^{+0.05}_{-0.01}$	$3.47^{+0.08}_{-0.12}$	$3.36^{+0.16}_{-0.02}$
Iso.Ab. N	$325.38^{+17.75}_{-6.27}$	$306.87^{+17.12}_{-7.38}$	$276.35^{+44.56}_{-20.91}$	$280.12^{+23.61}_{-35.03}$	$348.86^{+7.40}_{-25.27}$	$323.27^{+9.25}_{-17.15}$	$333.27^{+23.85}_{-21.66}$	$307.74^{+18.80}_{-8.56}$	$327.91^{+14.14}_{-8.92}$	$313.82^{+8.68}_{-16.42}$	$337.18^{+26.21}_{-38.86}$	$300.77^{+14.11}_{-14.55}$	$308.24^{+16.04}_{-49.55}$	$228.85^{+61.43}_{-7.77}$
Iso.Ab. O [10 <sup>3</sup> ]	$4.35^{+0.18}_{-0.02}$	$4.41^{+0.05}_{-0.04}$	$4.40^{+0.05}_{-0.10}$	$4.40^{+0.05}_{-0.08}$	$4.41^{+0.05}_{-0.05}$	$4.41^{+0.03}_{-0.09}$	$4.38^{+0.05}_{-0.04}$	$4.37^{+0.04}_{-0.07}$	$4.42^{+0.03}_{-0.09}$	$4.34^{+0.07}_{-0.01}$	$4.313^{+0.181}_{-0.004}$	$4.34^{+0.11}_{-0.02}$	$4.32^{+0.23}_{-0.01}$	$4.41^{+0.12}_{-0.05}$
$\delta_{XS \rightarrow C}$	$-0.08^{+0.23}_{-0.08}$	$0.03^{+0.14}_{-0.13}$	$0.17^{+0.09}_{-0.13}$	$0.13^{+0.14}_{-0.07}$	$-0.05^{+0.08}_{-0.05}$	$-0.12^{+0.18}_{-0.06}$	$0.17^{+0.06}_{-0.21}$	$0.11^{+0.12}_{-0.07}$	$0.23^{+0.03}_{-0.24}$	$0.15^{+0.05}_{-0.07}$	$0.28^{+0.02}_{-0.09}$	$0.25^{+0.04}_{-0.03}$	$0.26^{+0.03}_{-0.06}</math$	