

# Searching for fractionally charged particles with DAMPE

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On behalf of the DAMPE collaboration



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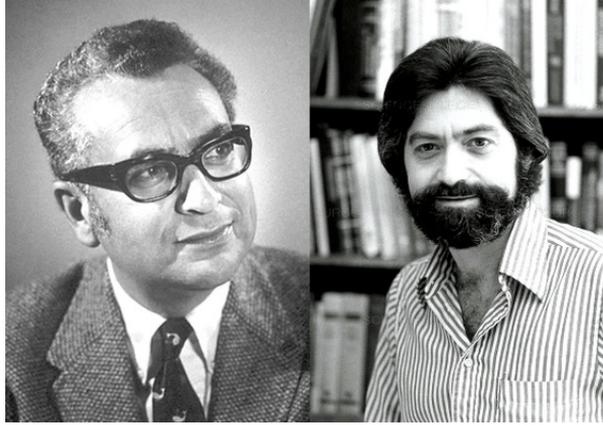
# Outline

- Motivation
- Previous searches of FCP
- DAMPE experiment
- Search for FCP with DAMPE
- Summary

# Motivation



Oil Drop Experiment



Quark Model

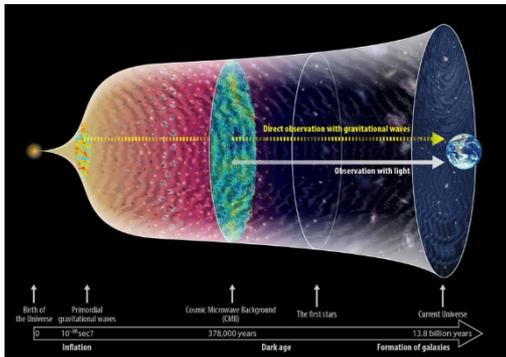


QCD Theory

- Since the oil drop experiment performed in 1909, all particles are measured as having charges of **multiples of electron charge**.
- In 1964, **quark model** for hadrons was proposed by Gell-man and Zweig.
- Due to the **QCD theory**, the quarks will not exist freely.
- Fractionally Charged Particle (FCP) is supposed to carry **any non-integer** charge.

# The possible origins of FCP

The basic assumption: **FCP is a kind of heavy lepton**



Early universe



Supernova explosion



Extensive air shower

*There are three possible sources of FCP in cosmic rays:*

- **First**, it may be produced at the early Universe after the Big Bang.
- **Second**, it may be produced through high-energy astrophysical processes.
- **Third**, it may be produced in the extensive air shower of cosmic-rays.

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# The previous experiments of large volume

For searches of **underground**, the target FCP should have energy above **hundreds GeV** to penetrate the rocks

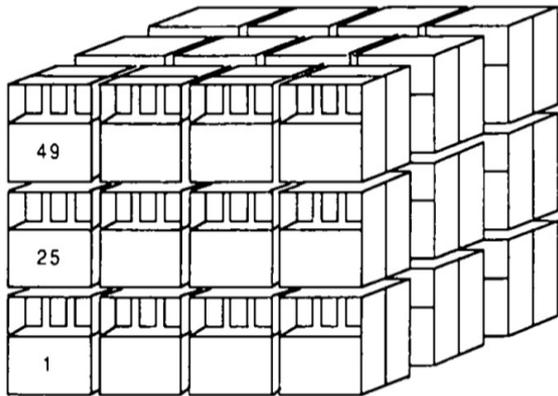


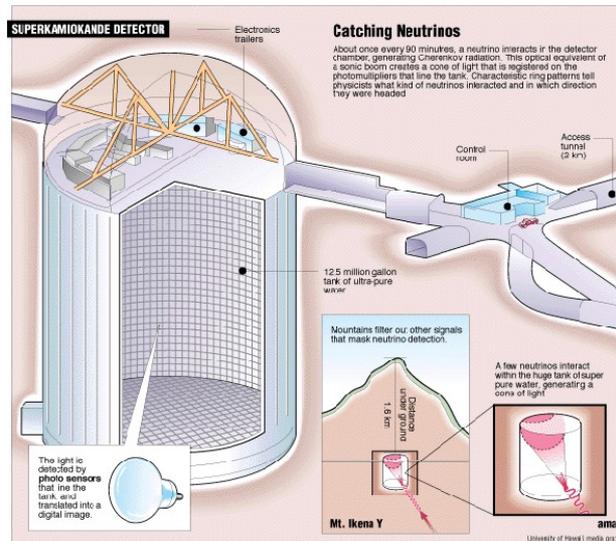
Fig. 1. The LSD experimental detector. The 72 tanks are considered as divided into 24 vertical columns (e.g. tanks 1–25–49 form the first telescope).

LSD 1800 m

$$\Phi\left(\frac{1}{3}\right) = 2.3 \times 10^{-13} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

$$\Phi\left(\frac{2}{3}\right) = 2.7 \times 10^{-13} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

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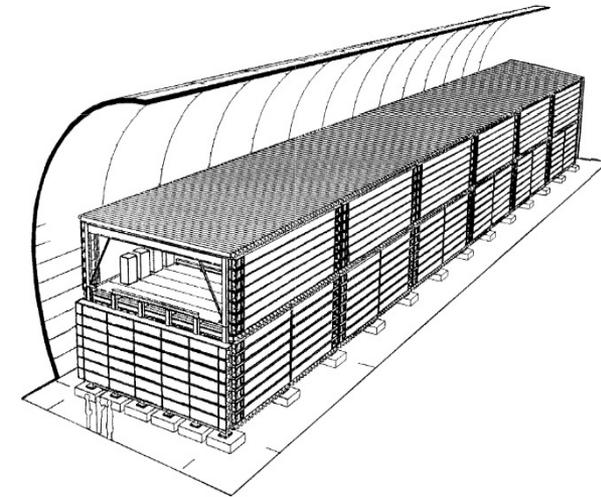


Kamiokande II 1000 m

$$\Phi\left(\frac{1}{3}\right) = 2.1 \times 10^{-15} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

$$\Phi\left(\frac{2}{3}\right) = 2.3 \times 10^{-15} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

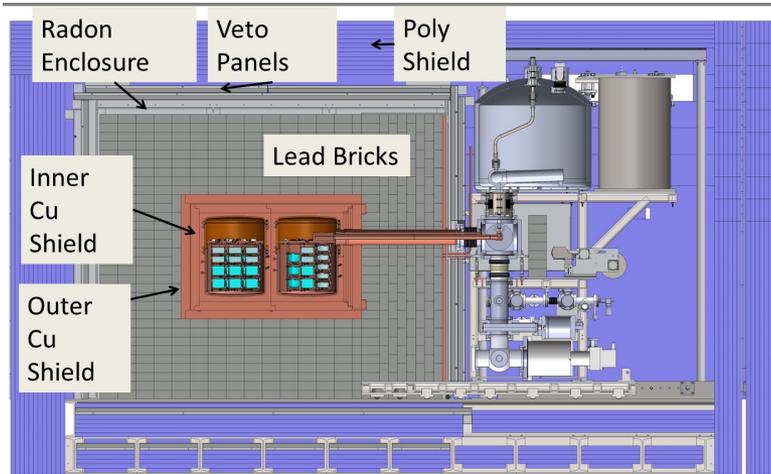
TeVPA2021



MARCO 1400 m

$$\Phi\left(\frac{1}{4} \sim \frac{2}{3}\right) = 6.1 \times 10^{-16} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

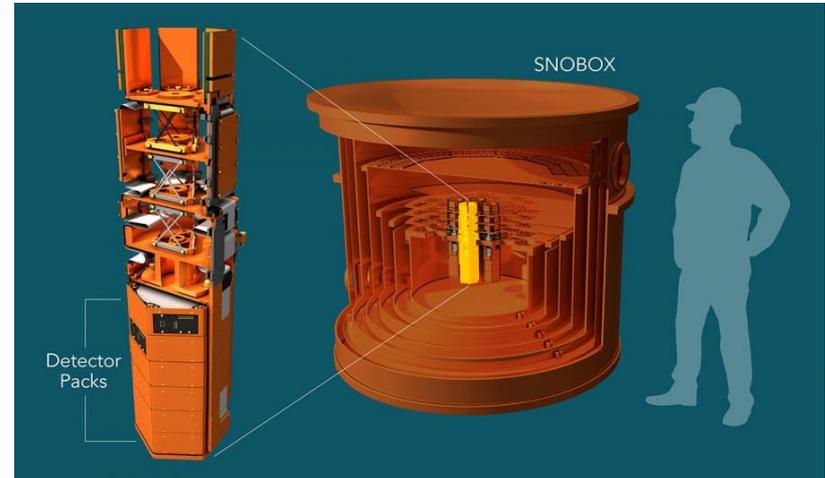
# The previous experiments of charge sensitivity



MAJORANA 1600 m

$$\frac{1}{1000} < Q < \frac{1}{6}$$

$$\Phi\left(\frac{1}{6} \sim \frac{1}{30}\right) = 2 \times 10^{-9} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$



CDMS II 700 m

$$\frac{1}{200} < Q < \frac{1}{6}$$

$$\Phi\left(\frac{1}{160}\right) = 1.36 \times 10^{-7} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$

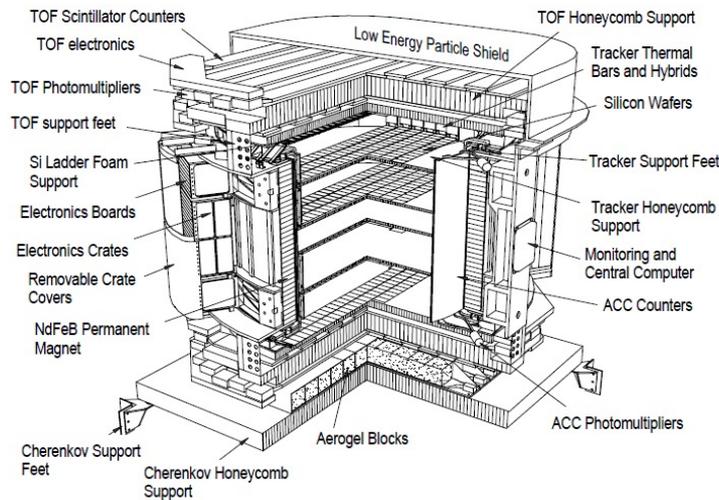
Searches for **Millicharge**, these experiments have very good capability of charge measurement

$$\text{CDMSlite: } \frac{1}{1000} < Q < \frac{1}{6}$$

$$\text{TEXONO: } 10^{-6} < Q < 10^{-3}$$

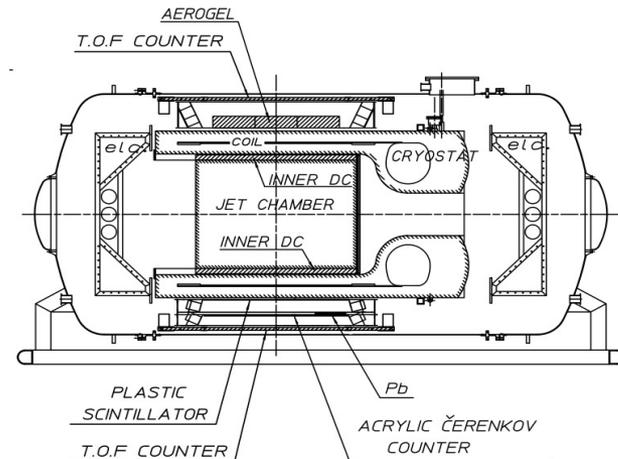
# The previous experiments in space

For searches **in space**, the target FCP need not to penetrate the rocks, it can have energy **as low as ~GeV level** (the geomagnetic cutoff)



AMS01 space shuttle

$$\Phi\left(\frac{2}{3}\right) = 3.0 \times 10^{-7} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

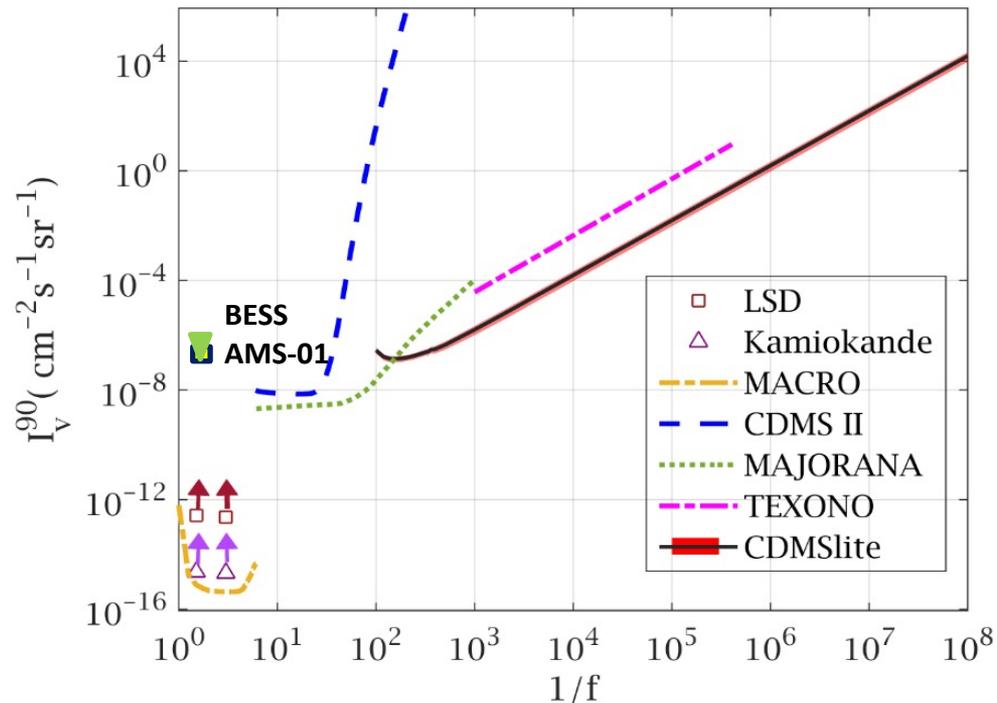


BESS balloon

$$\Phi\left(\frac{2}{3}\right) = 4.5 \times 10^{-7} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

# The previous experiments

The flux upper limit versus the inverse charge value

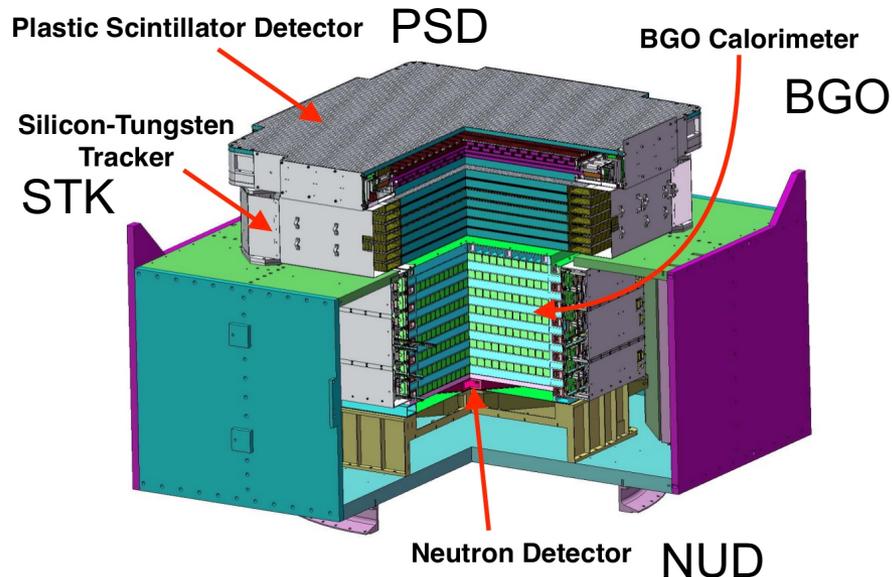


- **DAMPE** has been working stably on orbit for **nearly six years**.
- A lot of scientific data was accumulated.
- We hope to do something in searching for FCP as an on-orbit apparatus.

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# DAMPE experiment



## Dark Matter Particle Explorer(DAMPE)

### Main Scientific Goals:

- Origin and Propagation of Cosmic-Rays
  - Dark Matter Indirect Detection
  - Gamma-ray Astronomy
- 
- Orbit: sun-synchronous
  - Altitudes: 500 km
  - Period: about 90 minutes
  - 5 million events/day
  - 16 GB/day downlink
  - Launched on Dec.17<sup>th</sup> 2015
  - Life time > 5 years

Charge measurement (**PSD, STK**)  
 Precise tracking (**STK + BGO**)  
 Precise energy measurement (**BGO**)  
 Particle identification (**BGO + NUD**)

# DAMPE collaboration

## CHINA

- Purple Mountain Observatory, CAS, Nanjing
- University of Science and Technology of China, Hefei
- Institute of High Energy Physics, CAS, Beijing
- National Space Science Center, CAS, Beijing
- Institute of Modern Physics, CAS, Lanzhou



## ITALY

- INFN Perugia and University of Perugia
- INFN Bari and University of Bari
- INFN Lecce and University of Salento
- INFN LNGS and Gran Sasso Science Institute

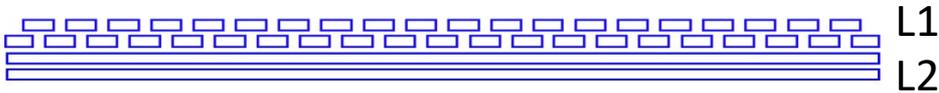


## SWITZERLAND

- University of Geneva

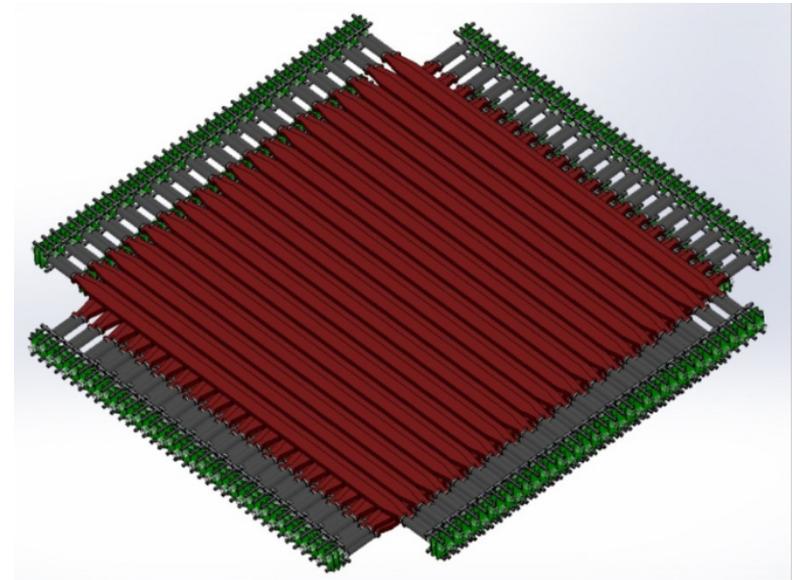
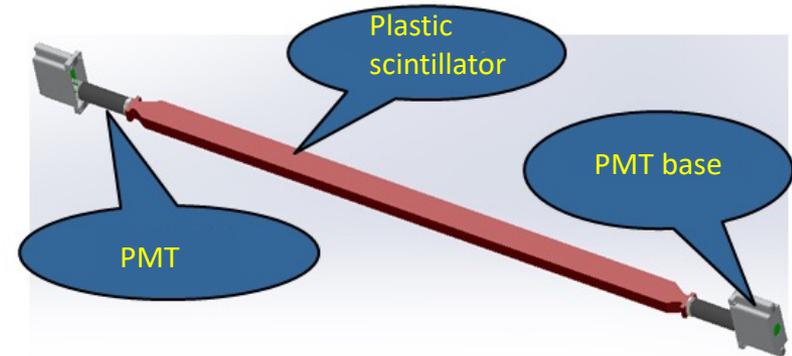


# Plastic Scintillator Detector (PSD)

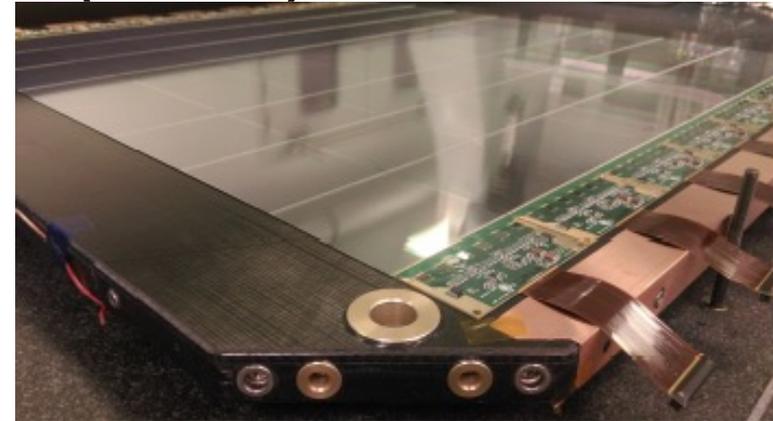
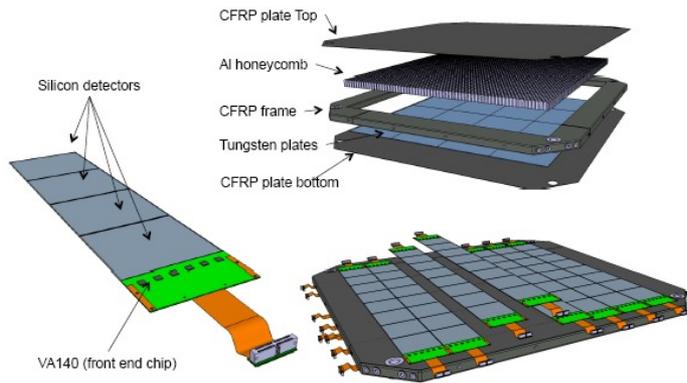


PSD is located on the top of the payload

- Active area: 82 cm × 82 cm
- Number of layers: 2
- 41 modules each layer
- A PMT at each end of strip
- Each PMT provides two signals (from Dy5 and Dy8 for large dynamic range )
- Charge resolution: 6% for  $Z = 1$

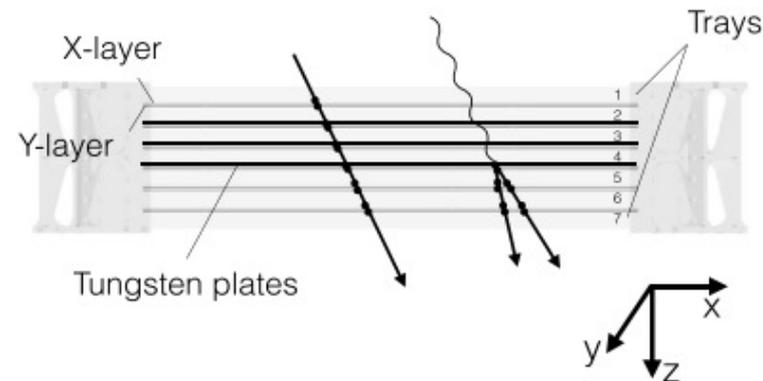


# Silicon Tungsten tracker (STK)



STK is composed by:

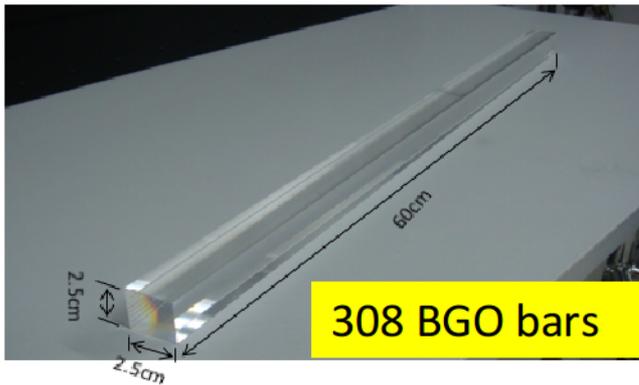
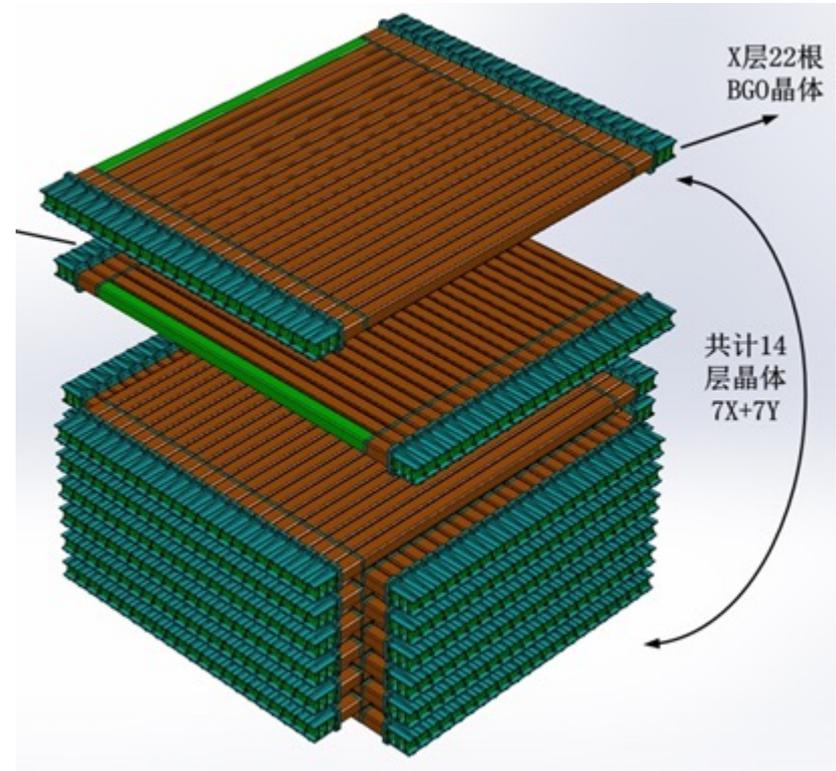
- Active area: 758 mm × 758 mm
- 7 CERP trays, 3 with tungsten plates for photon conversion
- Number of layers: 6
- 192 Si ladders. 16 on each sensitive face(12)
- each tray is orthogonal to the previous one to allow 3D tracking



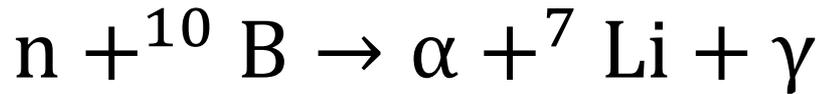
# BGO calorimeter

## BGO structure:

- 14 layers of 22 BGO crystals
- Dimension of BGO bar:  $2.5 \times 2.5 \times 60 \text{ cm}^3$
- Hodoscopic stacking alternating orthogonal layers
- r.l: 32X0, NIL:1.6
- Two PMTs coupled with each BGO crystal bar in two ends
- Electronics boards attached to each side of module

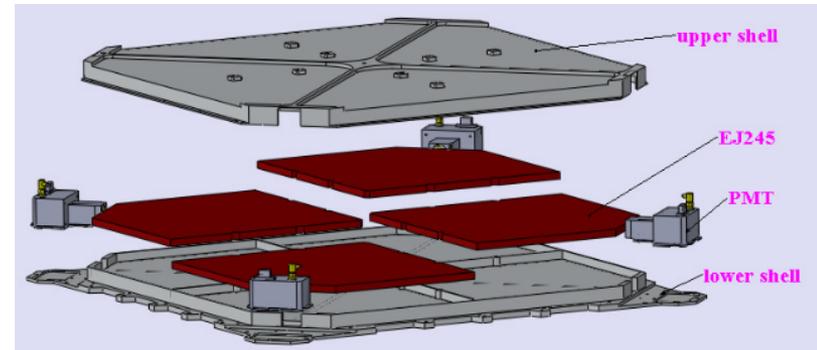


# Neutron Detector

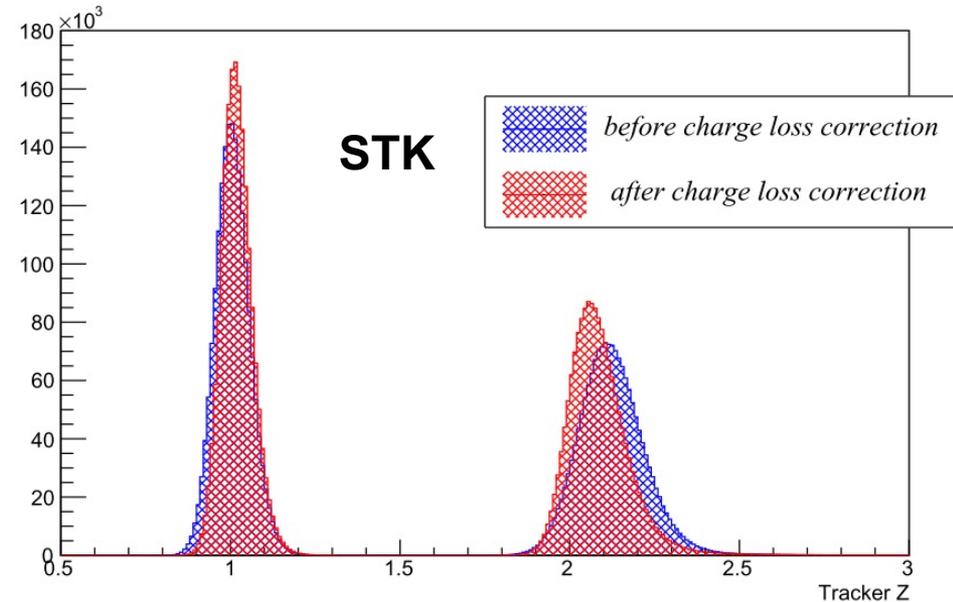
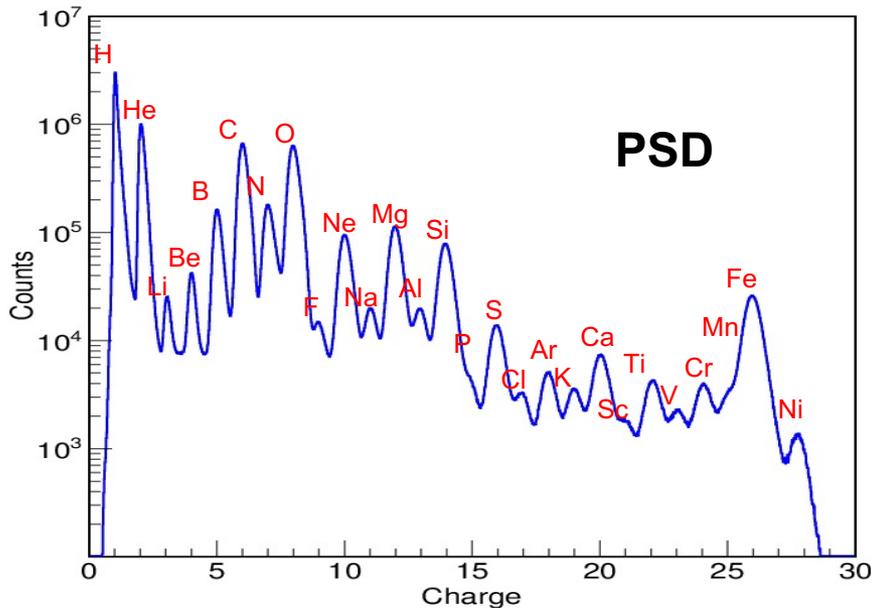


## NUD designed parameter:

- 4 boron-doped plastic scintillators ( 30 cm × 30 cm × 1 cm)
- Active area: 61 × 61 cm<sup>2</sup>
- Energy range: 2 - 60 MeV for single detector
- Energy resolution: ≤ 10% at 30 MeV



# Performance of charge measurement



**Left:** The charge discrimination of **PSD** in Z from 1 to 28

**Right:** The charge discrimination of **STK** between proton and helium

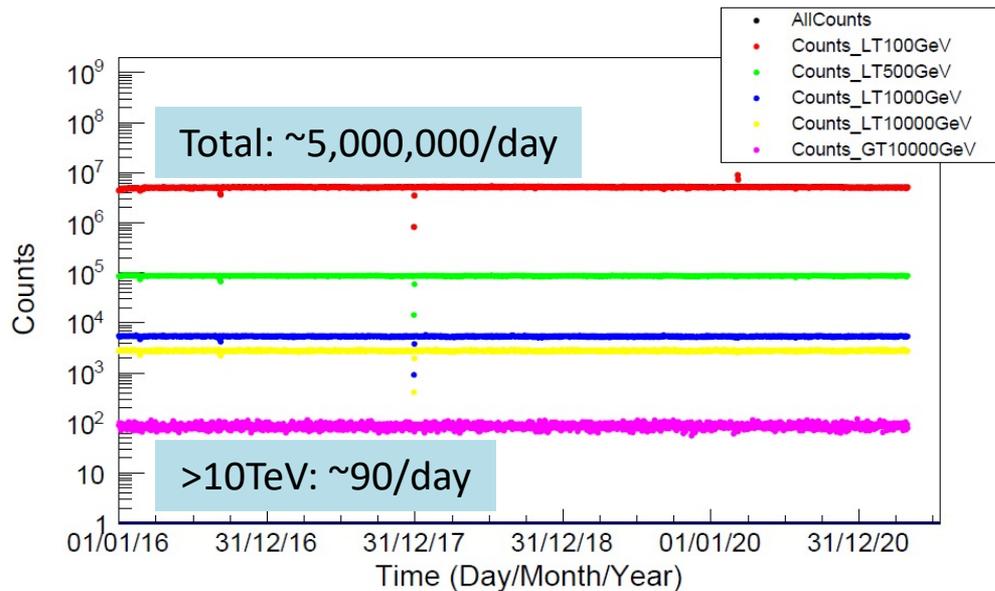
Charge resolution for singly charged particles:

PSD: **0.06e**

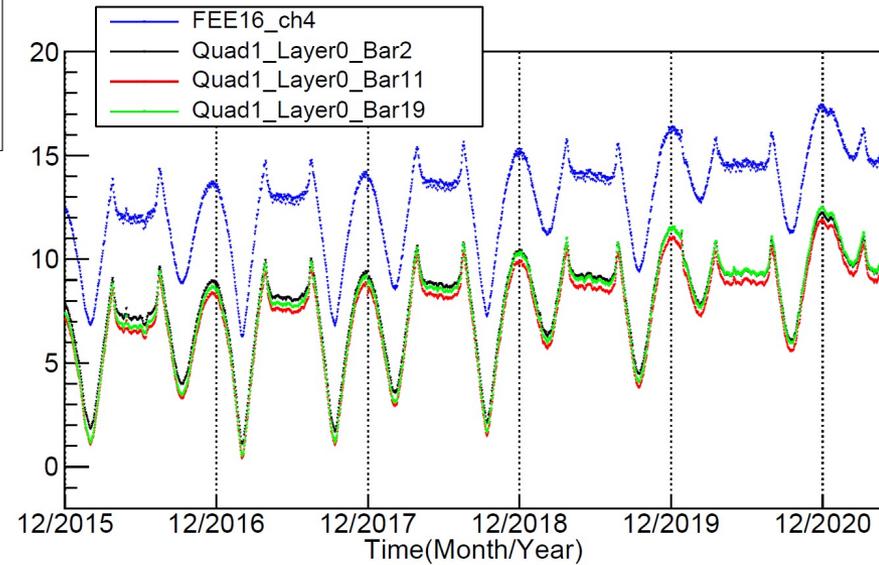
STK: **0.04e**

# DAMPE status

## Trigger rate per day



## Daily temperature variation of BGO

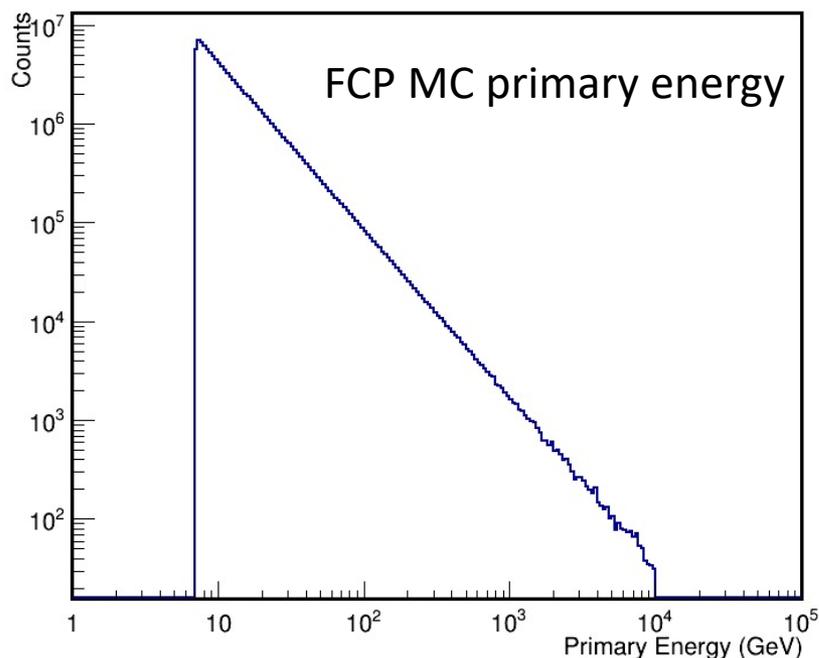


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# Data samples

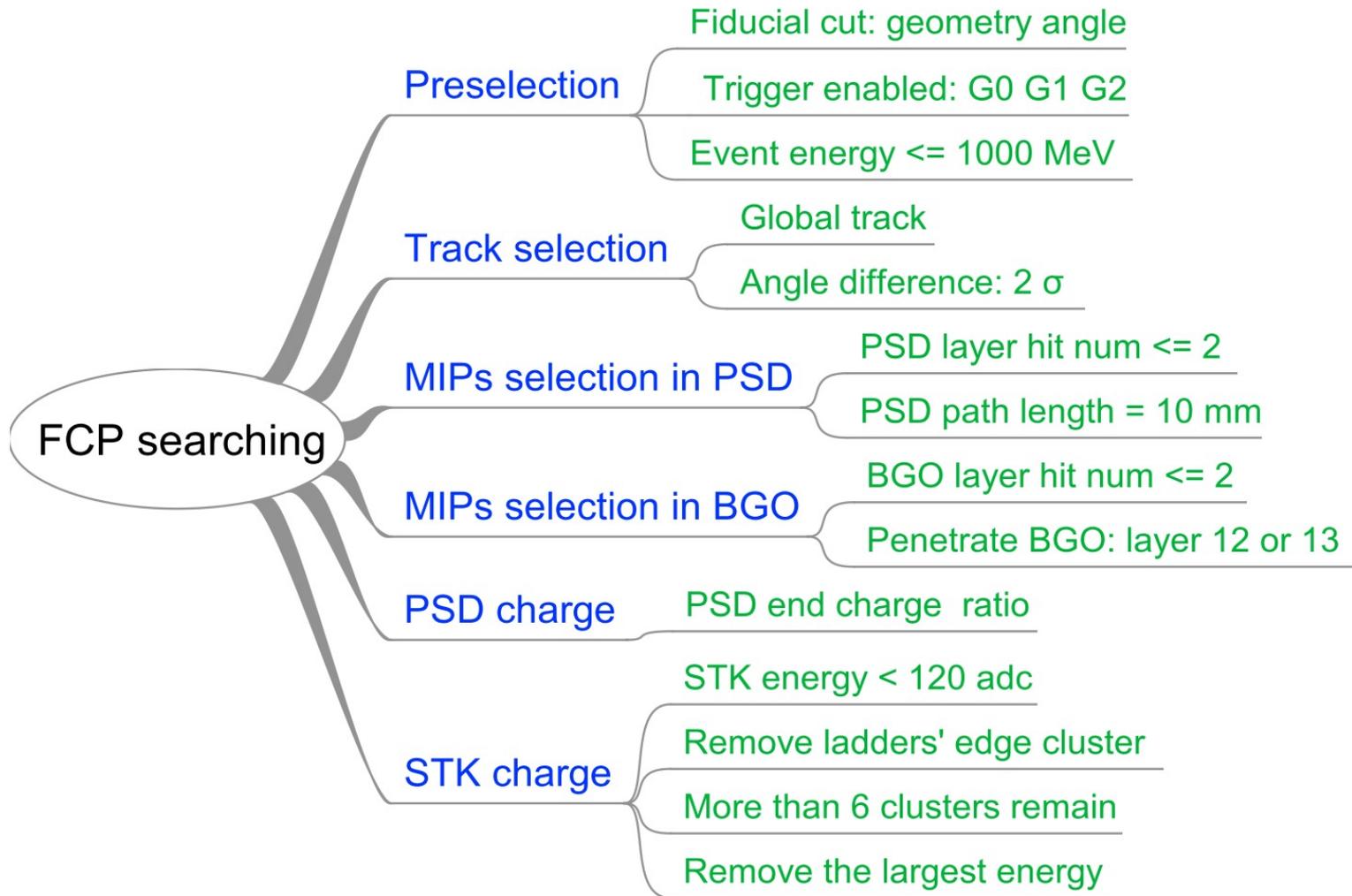
- Five years on-orbit data
- Proton simulation based on Geant4 in DAMPE software 10 GeV - 100 TeV
- FCP simulation based on Geant4 in DAMPE software 7 GeV – 10 TeV



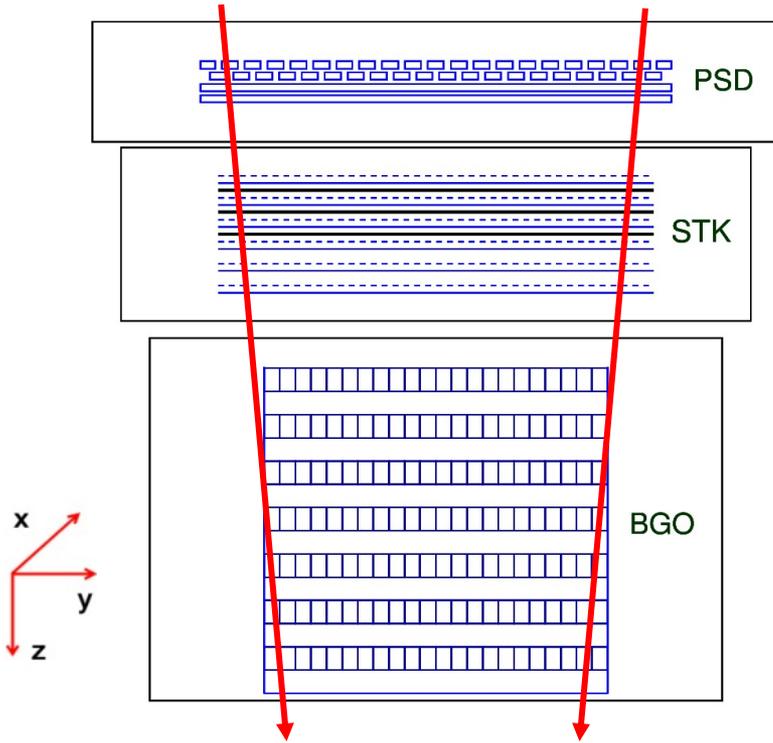
## FCP simulation:

- Created a virtual particle
- Charge with  $\frac{2}{3}e$
- Mass with 1.2 GeV
- Add ionization and multi scattering process
- Energy spectrum obey the  $E^{-3}$
- Spheric particle source

# Searching FCP with DAMPE



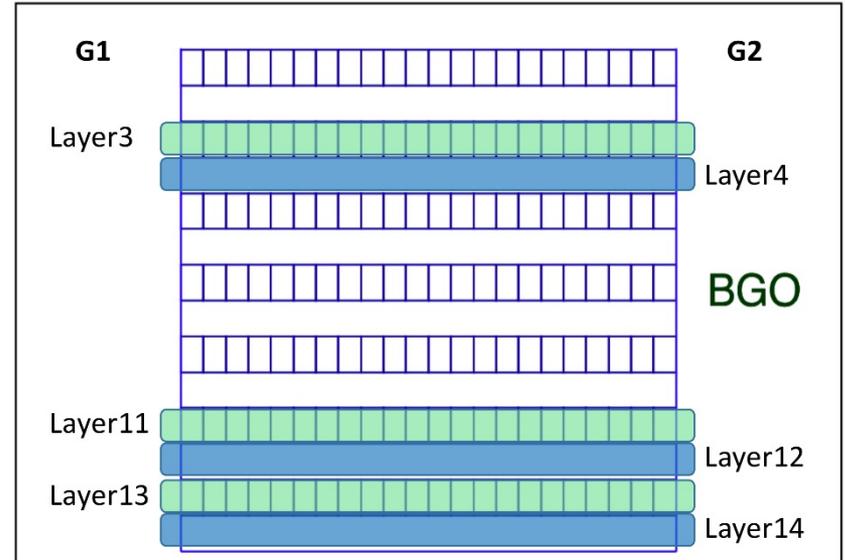
# Pre-selection



## Fiducial cut:

Constrain the positions of injection and ejection to maintain the event in the whole detector

**MIPs Trigger:** G1||G2 trigger enabled  
 $-20^\circ < \text{latitude} < 20^\circ$



Since the calibrated trigger threshold of G1 G2 is  $\sim 0.2$  MIPs, higher than the  $1/3$  charged particles ( $1/9$  MIPs), we aim to search the  $2/3$  charged particles.

## Mips energy:

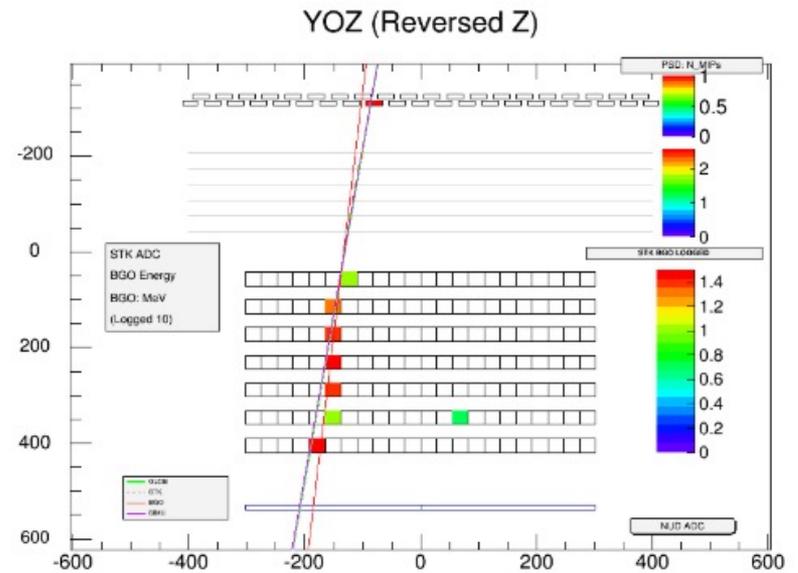
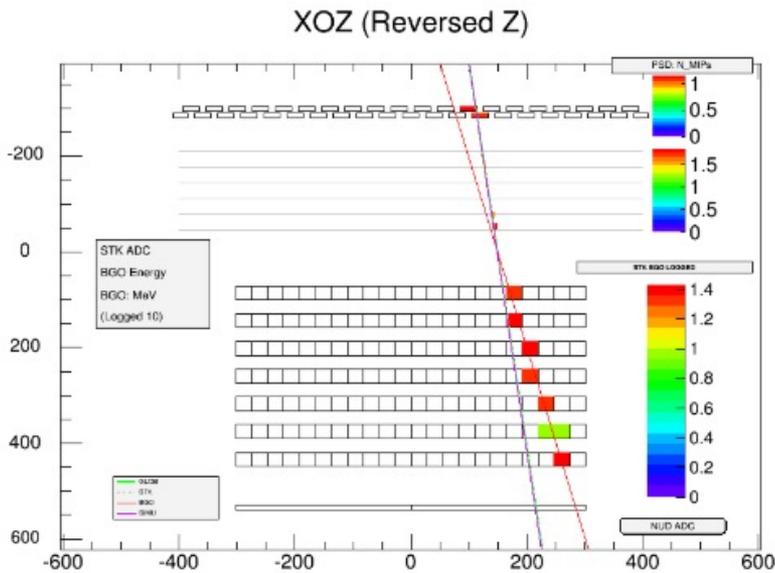
1000 MeV loose cut since singly charged particle depositing energy is about 700 MeV

# Track selection

**Global track:**  
should be reconstructed in STK.

**Angle difference between Global Track and BGO Track:**

The hits in BGO calorimeter can be used to reconstruct a BGO track. If global track deflects too much from BGO track, there may be scatters happened



**MC\_Proton event Angle > 6°**

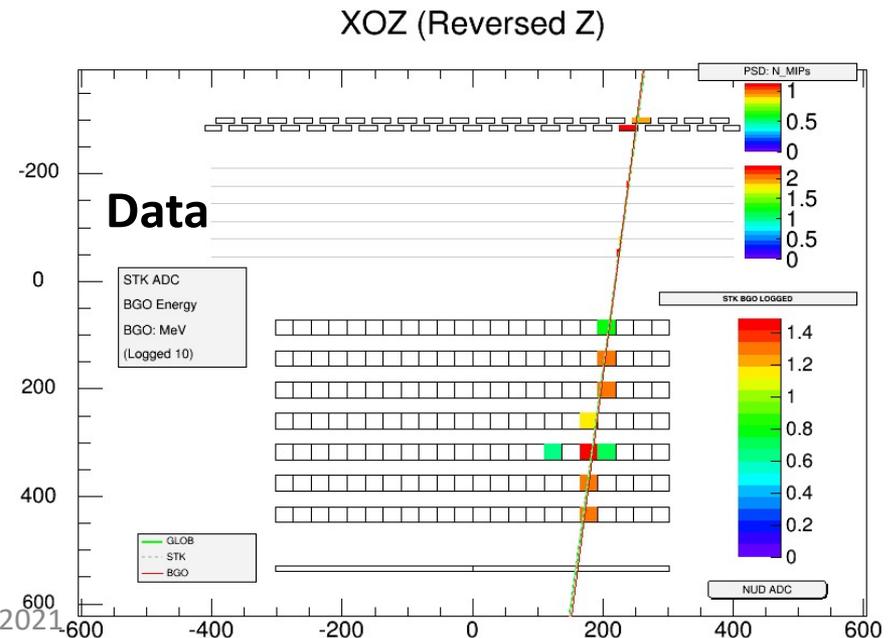
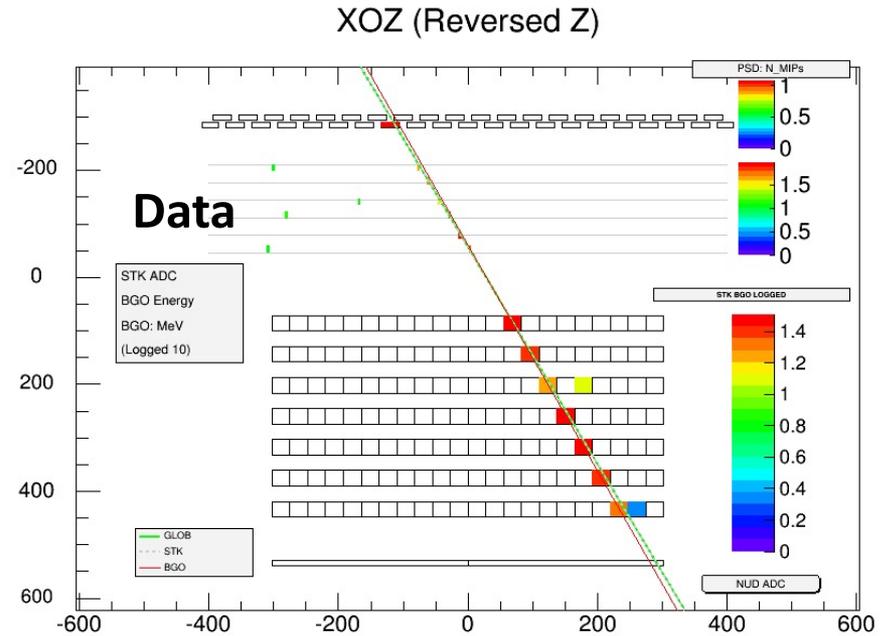
# MIPs selection

## MIP in PSD:

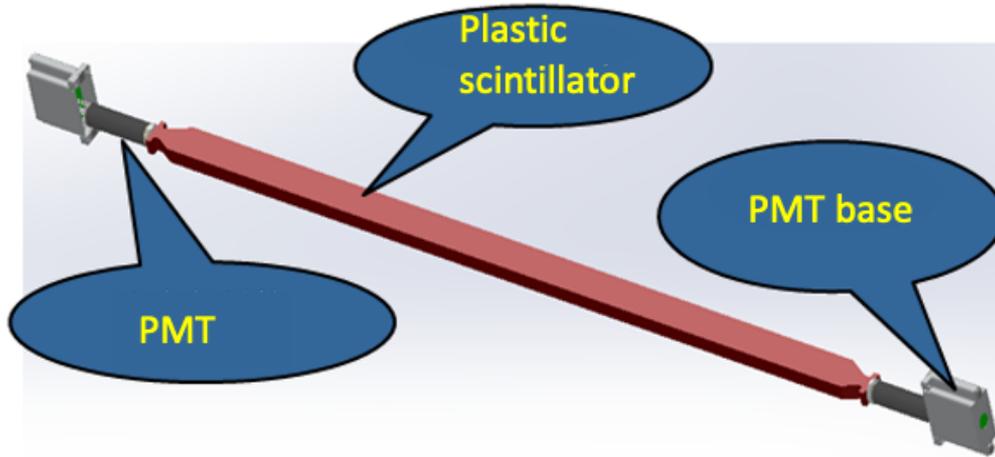
- Every Layer hit bars  $\leq 2$
- Path length = 10 mm

## MIP in BGO:

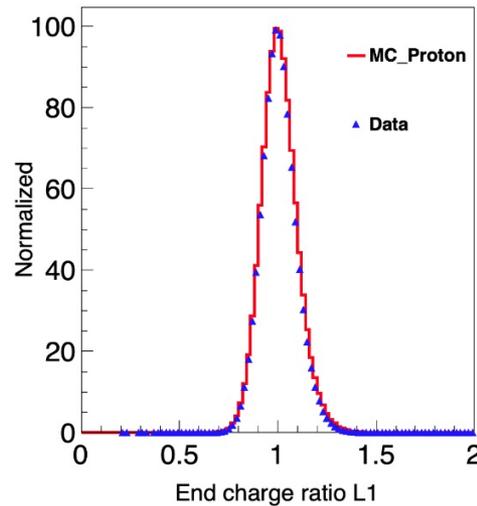
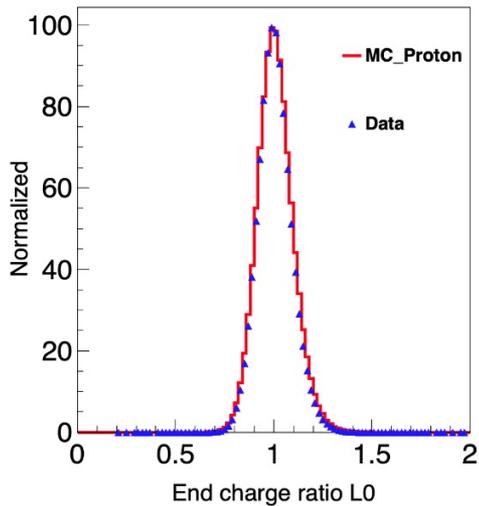
- Over-threshold(2 MeV) hits no more than 2 in one layer along the track and no more than 28 in the whole calorimeter
- Hit layers  $> 5$  in both YOZ and XOZ, **last two layers** should be fired



# PSD charge

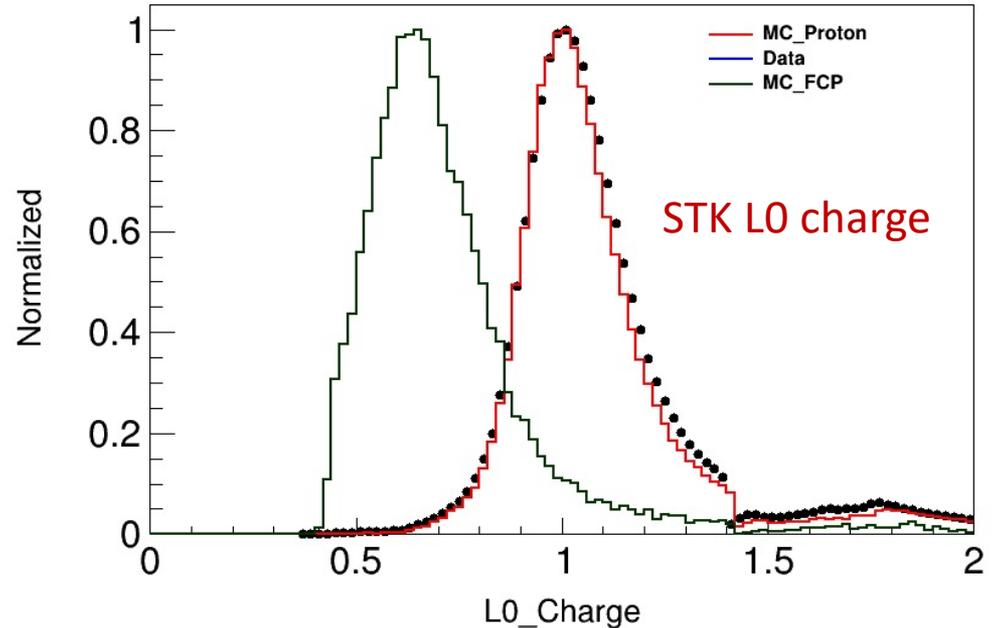
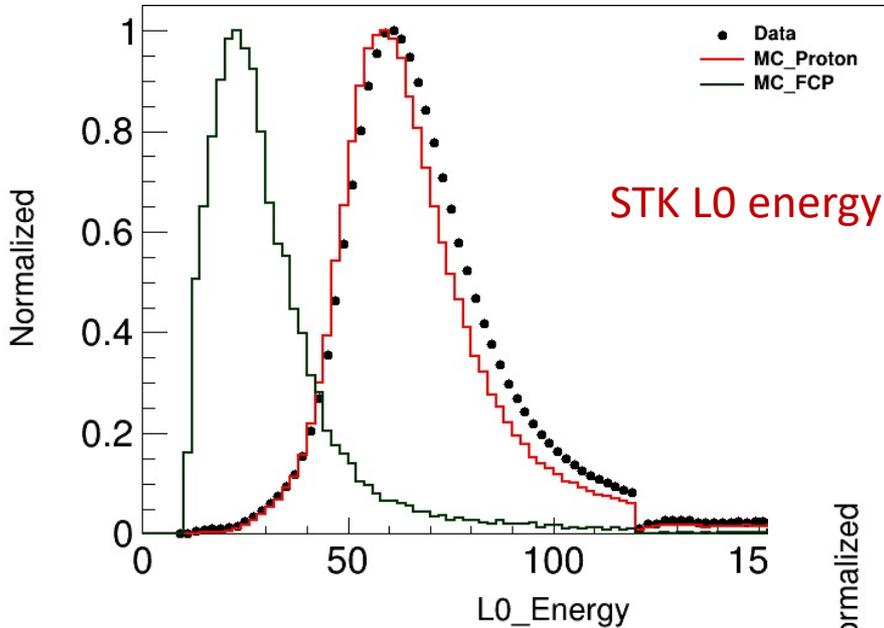


To ensure the reliability of charge reconstruction, the ratio of two ends of one strip should be consistent



# STK sub-layer energy and charge

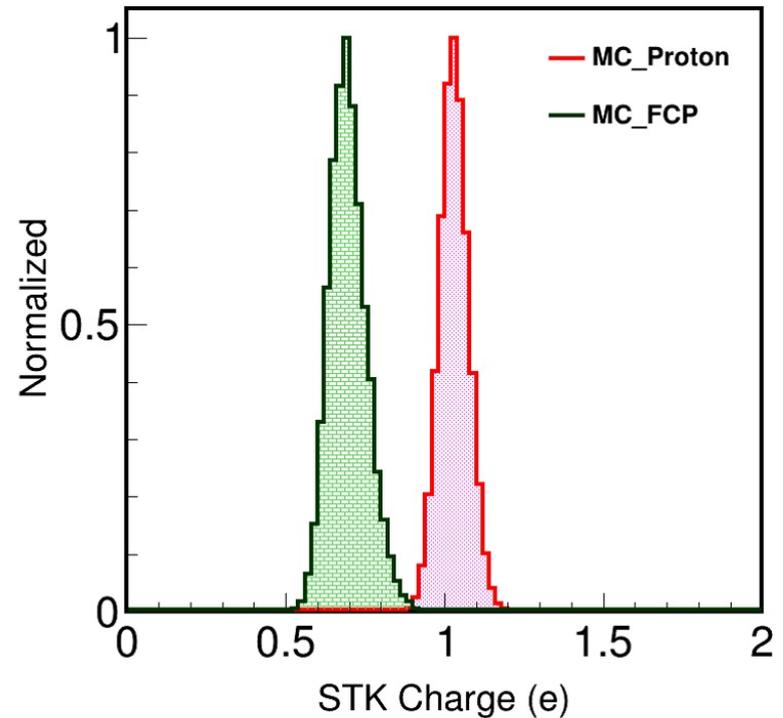
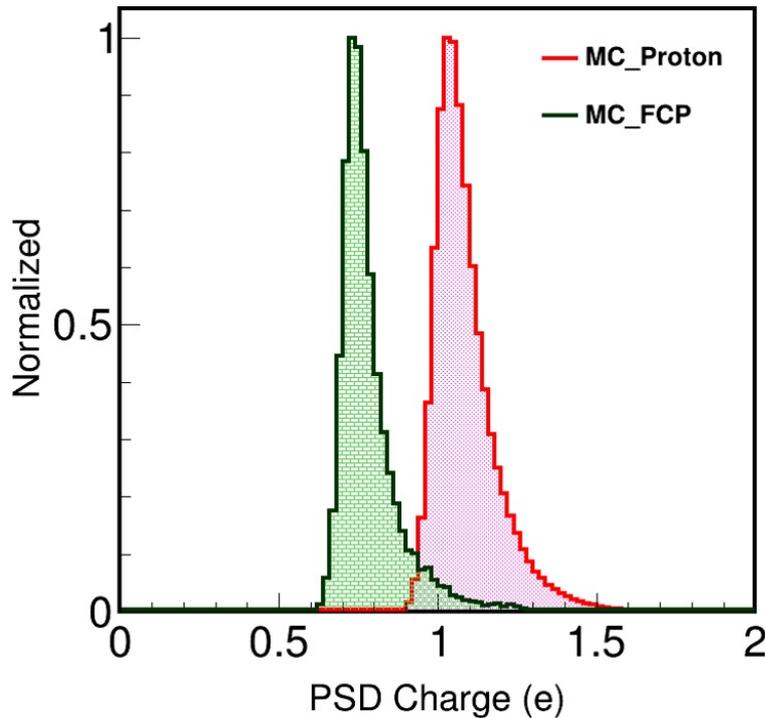
- Take layer 1 for example



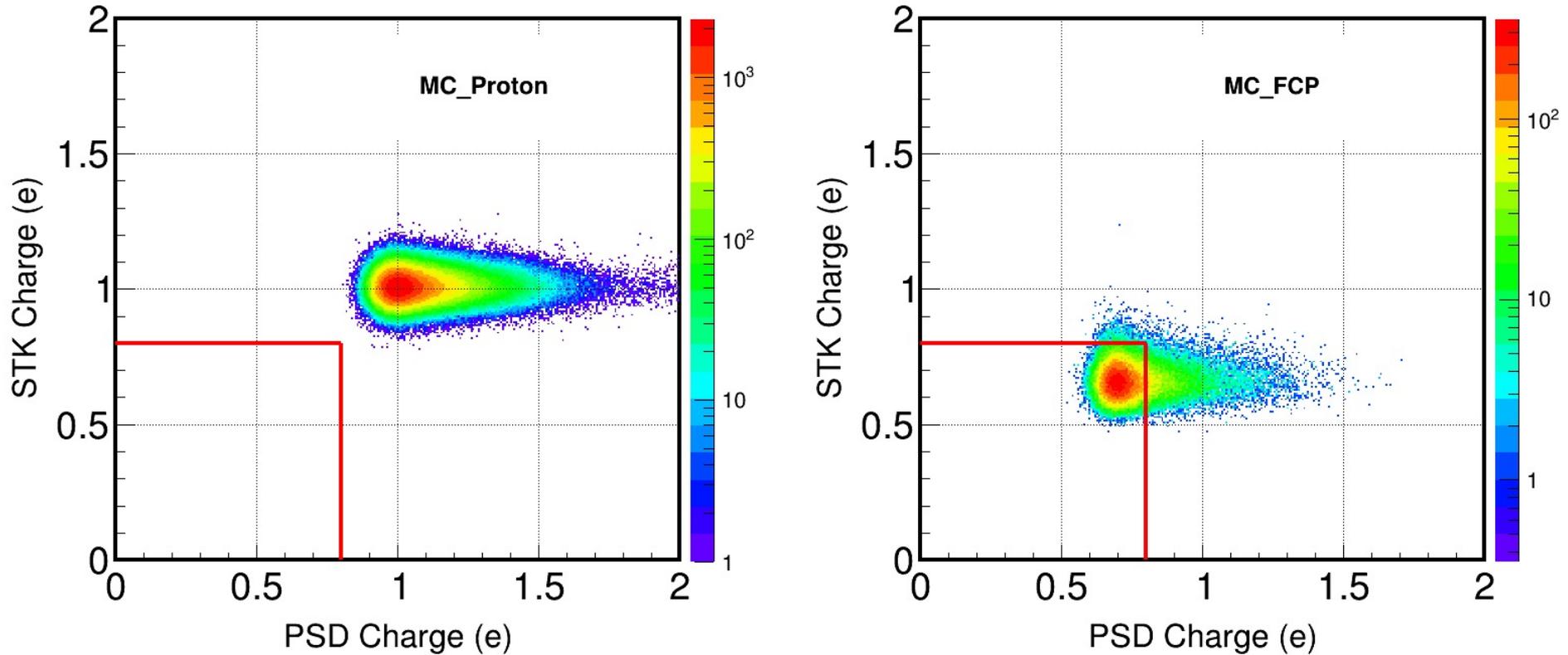
# Charge reconstruction

$$Q_{PSD} = \frac{Q_0 + Q_1}{2}$$

$$Q_{STK} = \frac{\sum_{i=1}^N Q_i}{N}$$

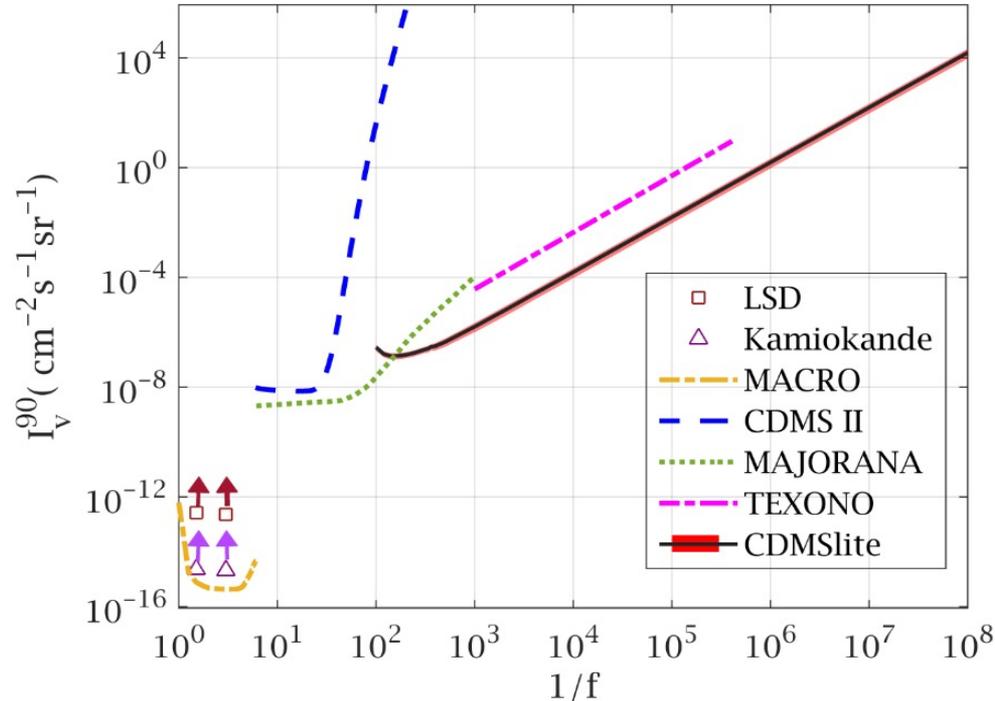


# Signal region for FCP



A signal region can be defined by PSD and STK MC as the red lines shows, the charge of two lines are set to  $0.8e$ . The signal region is effective to exclude the background and has a **80% efficiency** of covering the FCP.

# Results from underground experiments



With the **large volume and long exposure time**, The MACRO released the most stringent upper limit; with **high degree of charge sensitivity**, the CDMS and others could measure very small charge.

# Comparison with same type experiments

Experiments	Geometry acceptance ( $\text{cm}^2\text{sr}$ )	Exposure time (s)	Upper limit $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}$
AMS-01	3000	$3.6 \times 10^4$	$3.0 \times 10^{-7}$
BESS	1500	$3.2 \times 10^5$	$4.5 \times 10^{-7}$
<b>DAMPE</b>	3000	$2.4 \times 10^7$	<b>To be released</b>

- AMS-01 has a large acceptance, but short of the exposure time
- BESS integrates four times of flights to achieve a longer exposure time
- DAMPE has a relatively larger acceptance and the longest exposure time. We hope to release a lower upper limit in the near future.

# Summary

- The history of FCP has been reviewed briefly.
- The DAMPE experiment has been introduced.
- The selection criteria to search FCP with DAMPE have been studied.
- A MC simulation has been performed and an evaluation of the detection efficiency has been carried out.
- A comparison between DAMPE and other equipment has been done. DAMPE is hopeful to release a lower upper limit in space.

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Thank you!