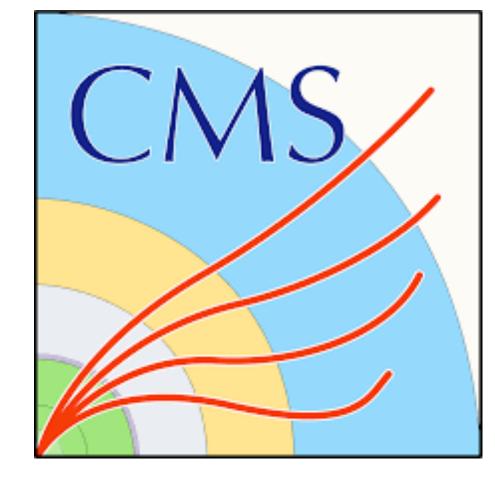


**TeV Particle Astrophysics 2021** 

Deepak Kumar, On behalf of CMS Collaboration

29<sup>th</sup> October 2021



# Dark Matter Searches at CMS

## Indian Institute of Science, Bangalore, India

## Dark Matter and detection channels

Existence of dark matter (DM) know from astrophysics and cosmos. Not observed directly.

It rarely interact with ordinary matter via electro-magnetic interaction(WIMPs)

From Cosmology, Universe contains 25% dark matter

### **1. Direct detection**

DM interacts with ordinary matters such as nucleons.

### **2. Indirect detection.**

- DM self-annihilate or decay in outer space.
- **3. Particle colliders.** 
  - Produce DM particles in a laboratory.

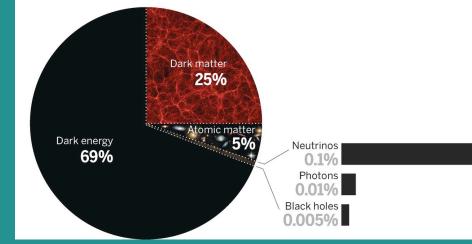
If DM can be produced at hadron colliders, then Large Hadron Collider (LHC) is the best laboratory to detect.

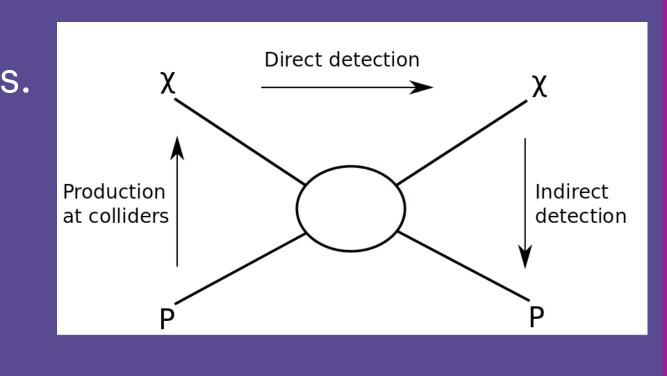
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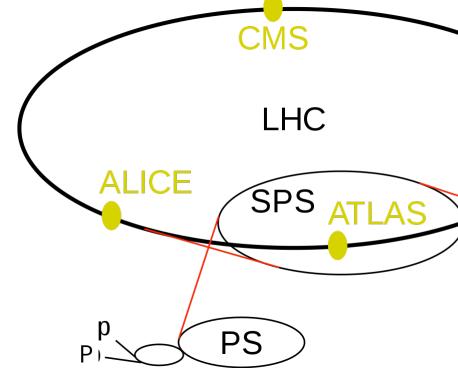
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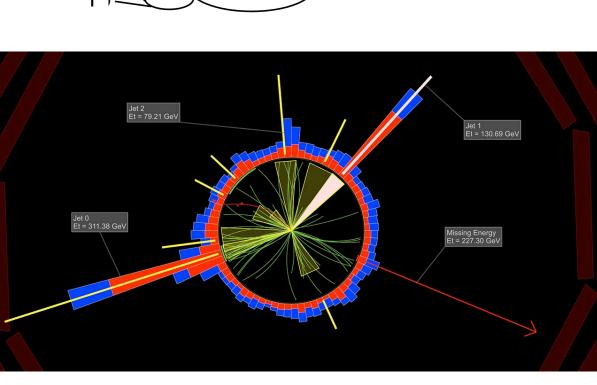


he multiple components that compose our universe









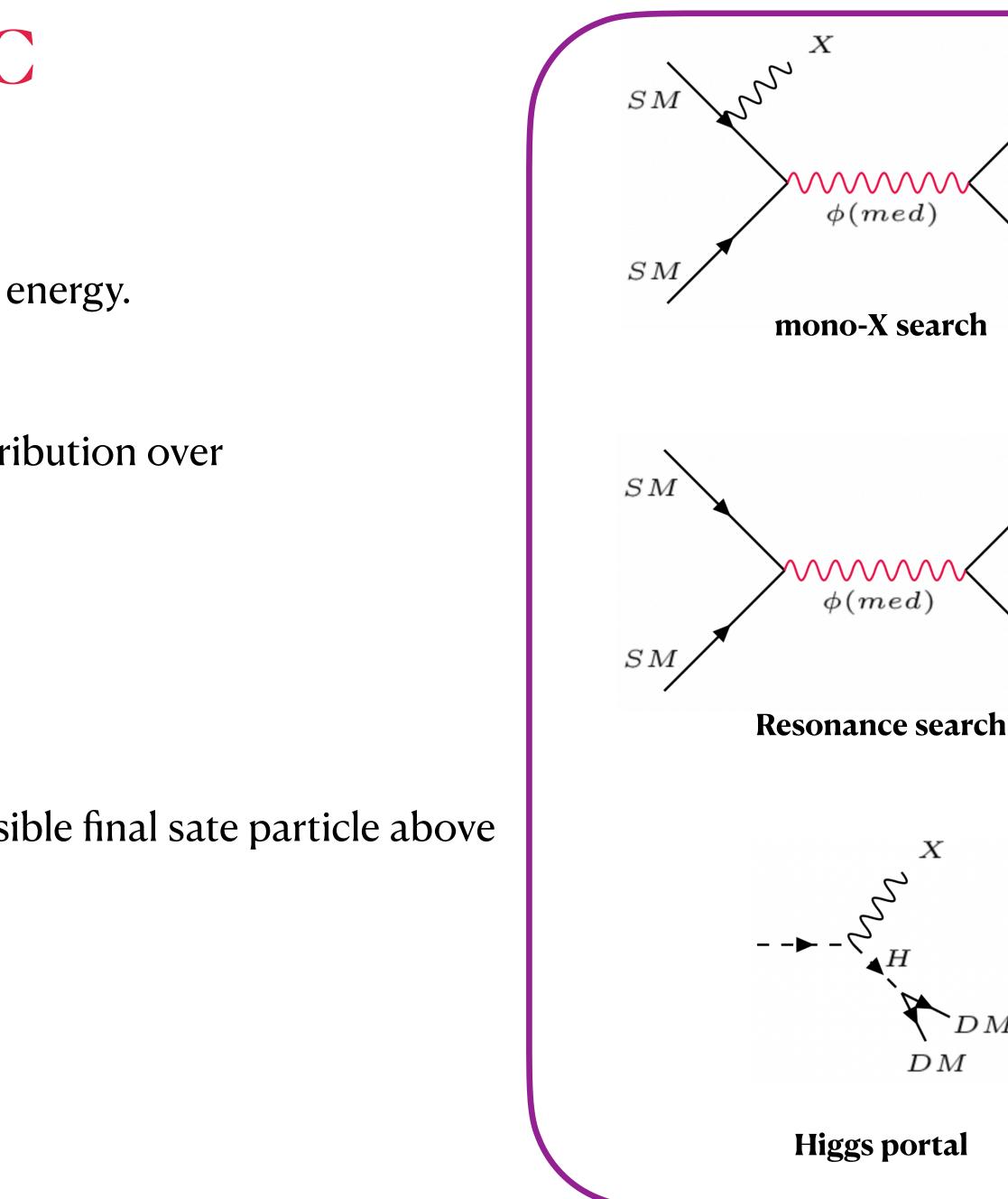


## Dark matter detection at LHC

## • mono-X search:

- Standard model particle recoil against missing energy.
- Tag from radiation or associated production
- Expect signal in the tail of missing energy distribution over the standard model background.
- Resonance search:
  - DM decays to standard model particle.
  - Expect signal peak in invariant mass of two visible final sate particle above the standard model background.
- Higgs portal:
  - Higgs decays to DM candidates.

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## Outline

## • Analyses for main talk:

- mono-J/V search : <u>EXO-</u>
- mono-Z search : <u>EXO-</u>
- mono-Higgs : <u>EXO-</u>
- Dilepton resonance search
- Dijet resonance search :
- Full list of analyses are available here:
  - Public results

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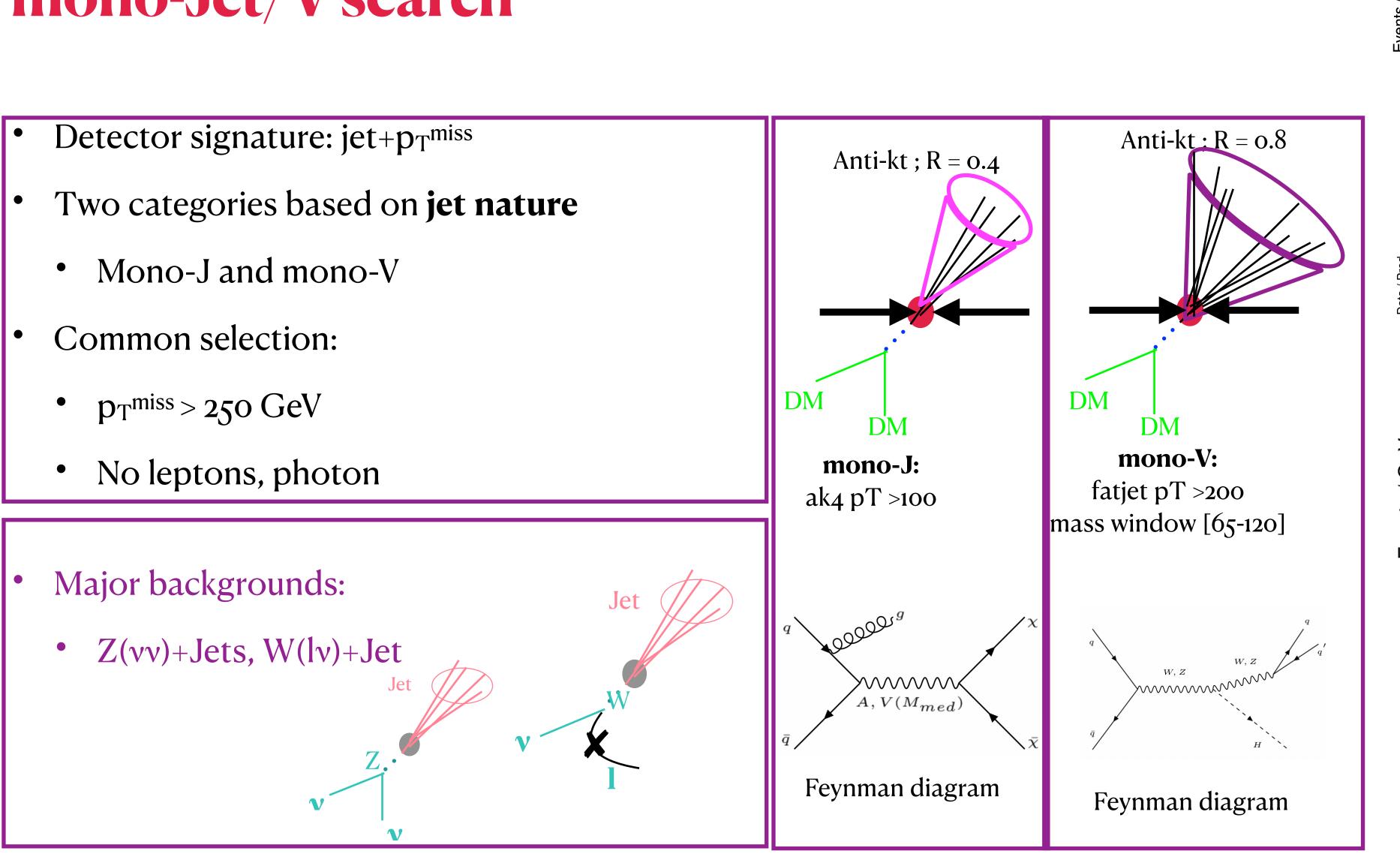
Dark Matter@CMS, 29 Oct 2021

: <u>EXO-20-004(arXiv:2107.13021)</u> : <u>EXO-19-003 (Eur. Phys. J. C 81 (2021) 13)</u> : <u>EXO-18-011(JHEP 03 (2020) 025)</u> : <u>EXO-19-019(JHEP 07 (2021) 208)</u> : <u>EXO-19-012(JHEP 05 (2020) 033)</u>



## mono-Jet/V search

- - Mono-J and mono-V

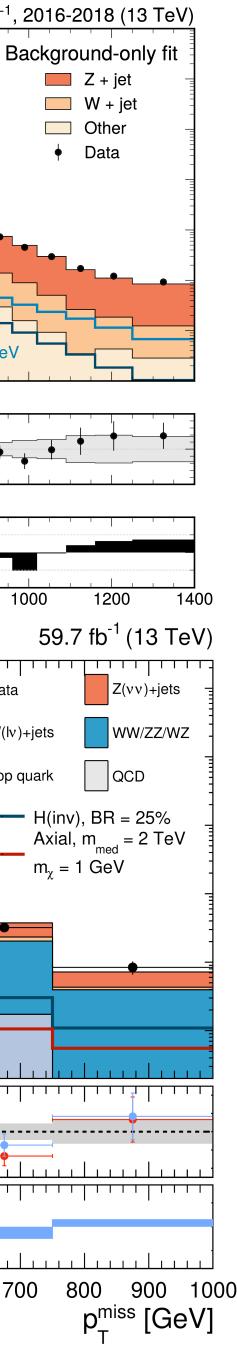


No excess of event observed with respect to SM background expectations.

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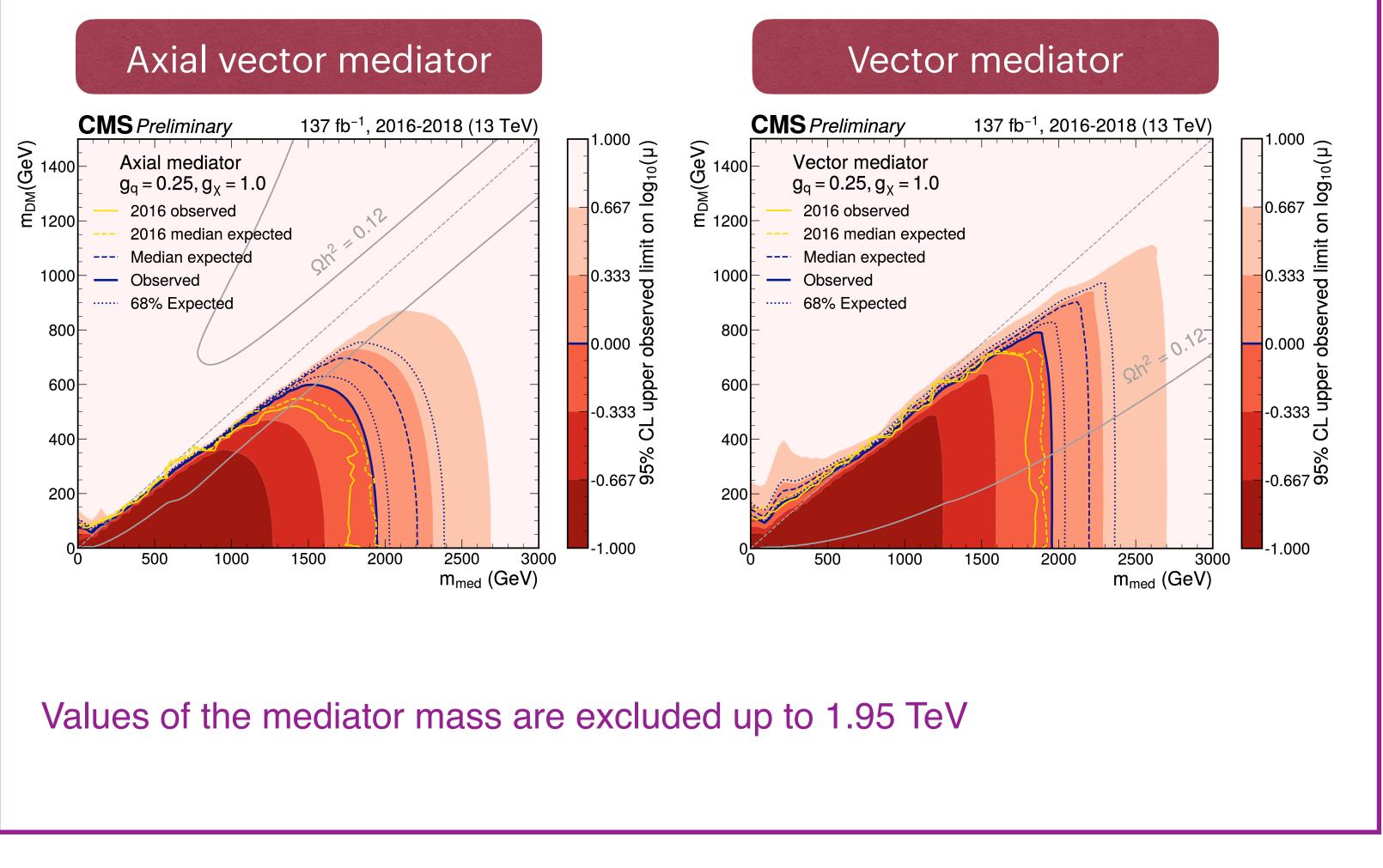
## Dark Matter@CMS, 29 Oct 2021

137 fb<sup>-1</sup>, 2016-2018 (13 TeV) Monojet GeV **CMS** *Preliminary* Events H(inv), BR = 25% $10^{-1}$ Axial,  $m_{med} = 2 \text{ TeV}$ ,  $m_y = 1 \text{ GeV}$ 400 600 1000 800 **Л**ар 2910<sup>3</sup> - Data **CMS** Preliminary Mono-V (high-purity) W(lv)+jets Events / 2018 Top guark  $m_{\gamma} = 1 \text{ GeV}$ 10<sup>-1</sup>  $10^{-2}$ Pred. 🔶 🔶 🔶 🕂 Post-fit 0.6 <del>.</del> <u>(Data-Prec</u> σ 600 700 300 400 500

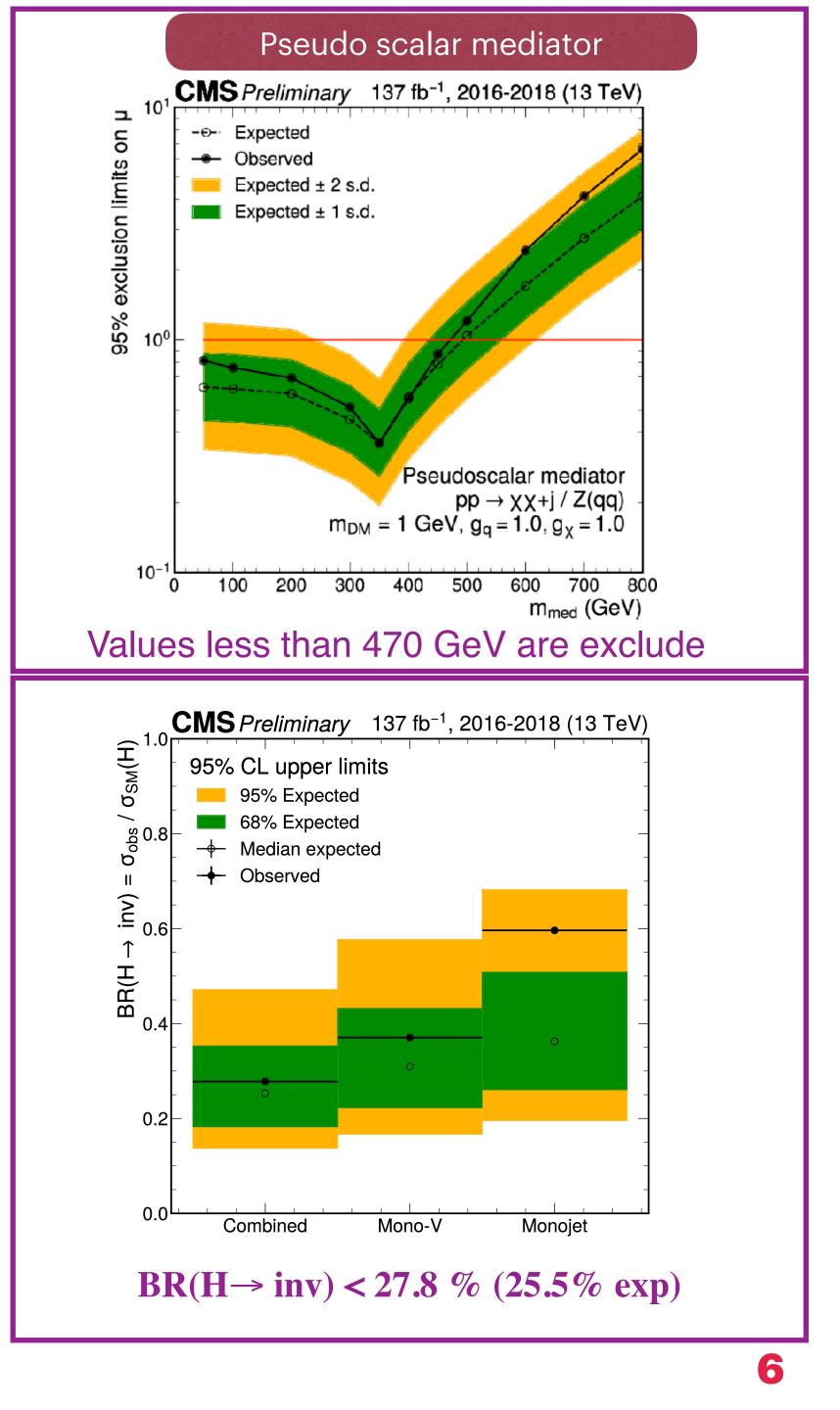




## mono-Jet/V search



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## mono-Z(II) Search

Signature: Z(II)+MET

**Model for interpretations:** Simplified model

2HDM+a

 $p_T^{miss}$  for simplified,  $m_T$  for 2HDM+a model

$$m_{\mathrm{T}} = \sqrt{2p_{\mathrm{T}}^{Z}p_{\mathrm{T}}^{\mathrm{miss}}(1-\cos(\Delta\phi_{\ell\ell-\vec{p}_{\mathrm{T}}^{\mathrm{miss}}}))},$$

### **Basic selection:**

- p<sub>T(l1)</sub> > 25 GeV, p<sub>T(l2)</sub> > 20 GeV
- |m<sub>II</sub> m<sub>z</sub>| < 15 GeV
- MET > 100 GeV

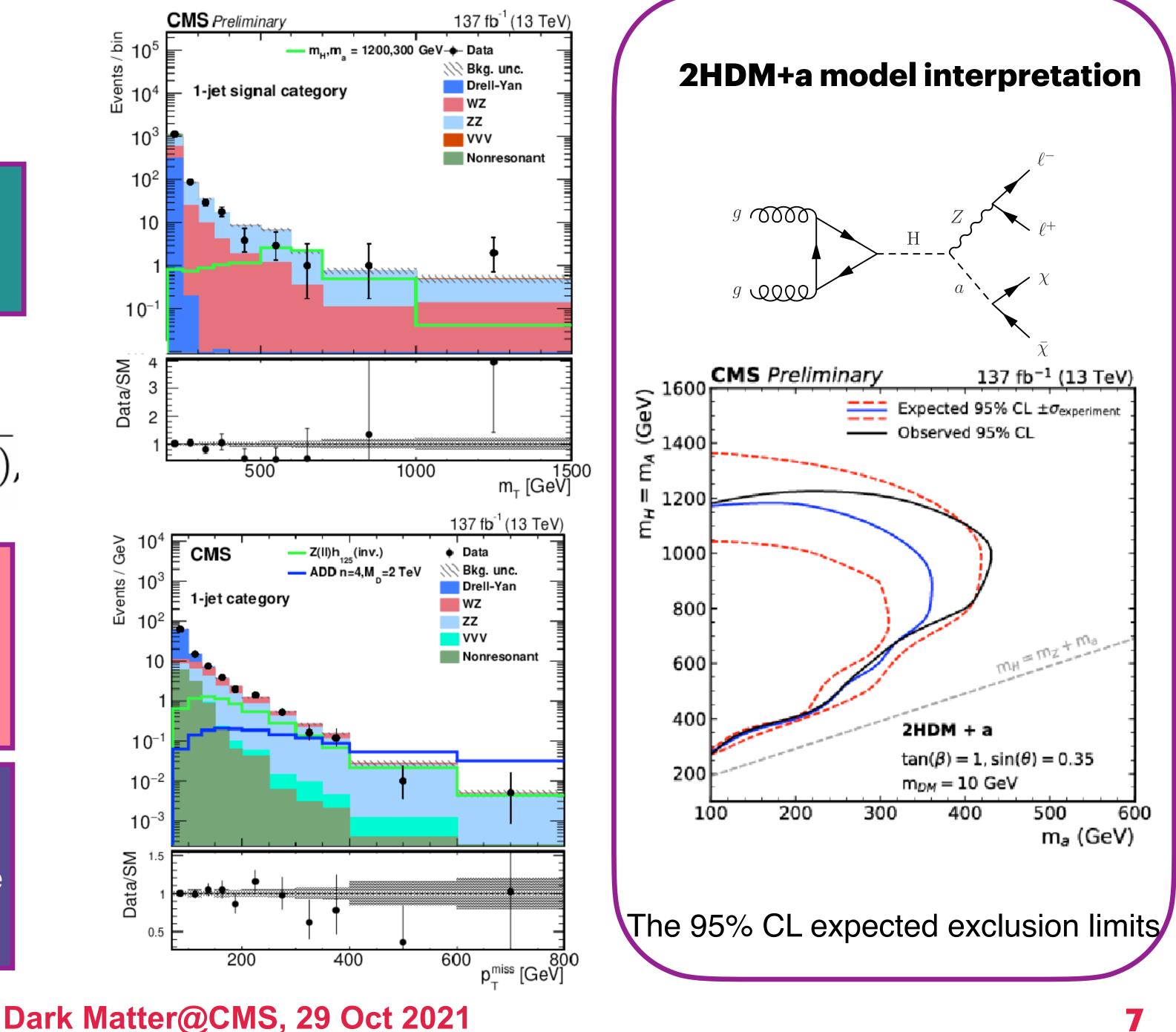
### **Backgrounds:**

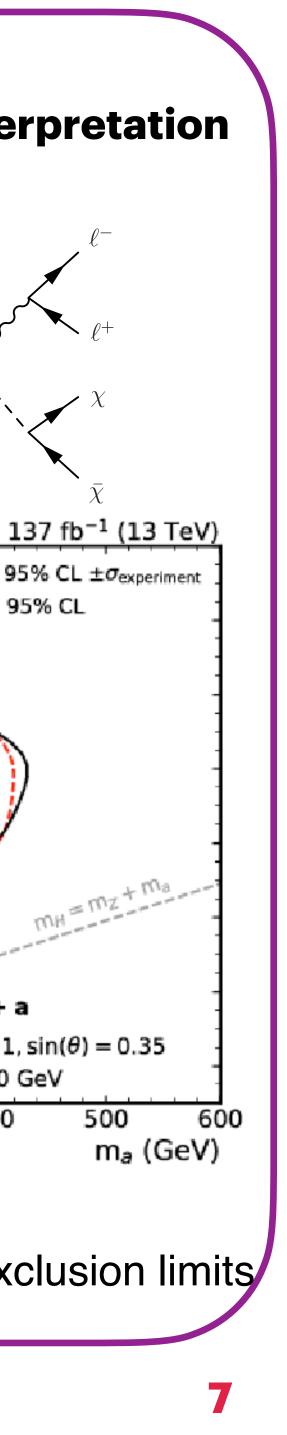
- Drell-Yan, WZ, ZZ, VVV
- Dedicated Control regions to model the background

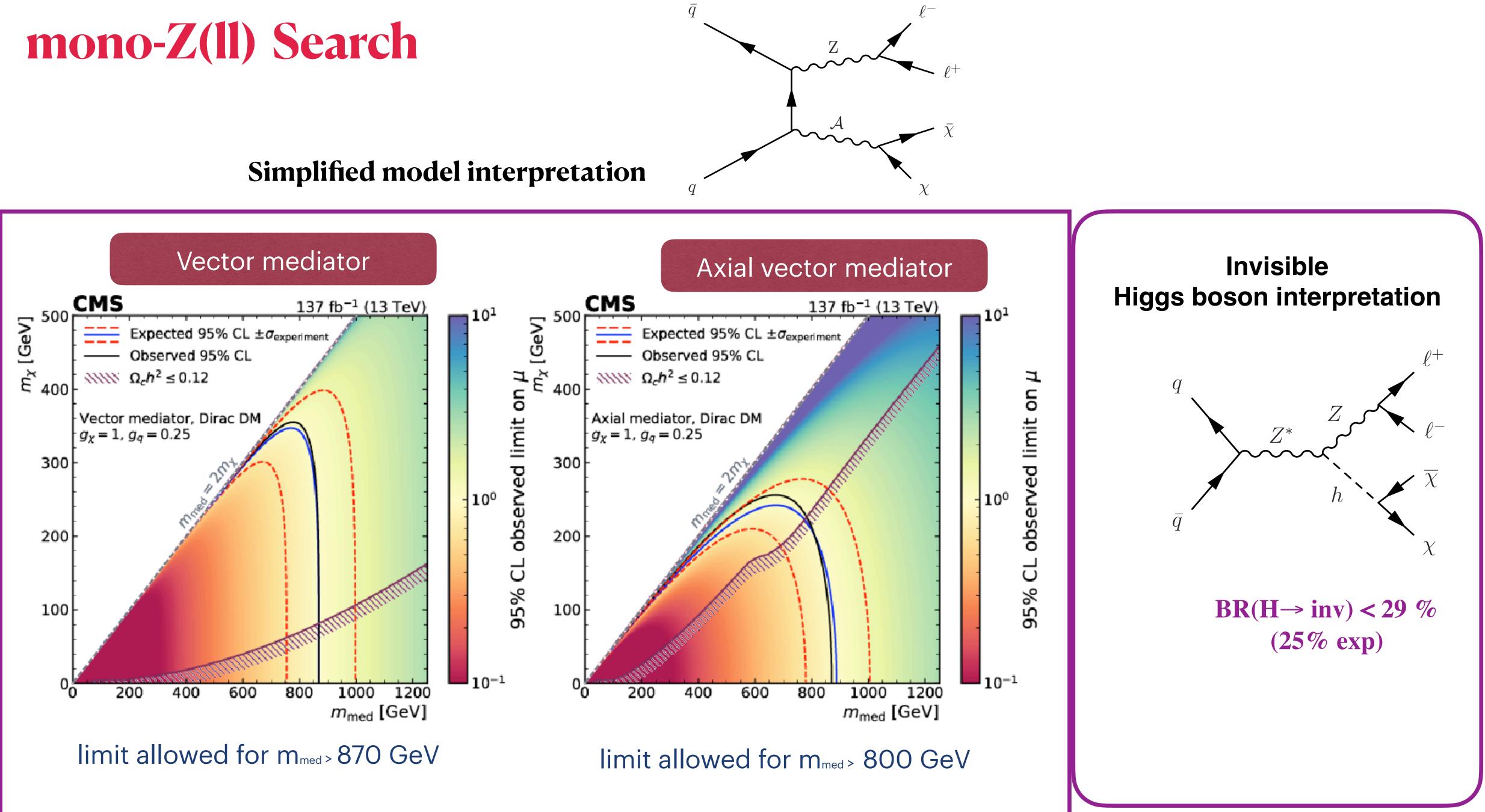
Events / bin 10<sup>5</sup> 10<sup>3</sup> 10<sup>2</sup> 10 10<sup>-1</sup> Data/SM 3 2

Events / 10<sup>3</sup> 10<sup>2</sup> 10 10<sup>-1</sup> 10<sup>-2</sup> 10<sup>-3</sup> Data/SM 1 0.5

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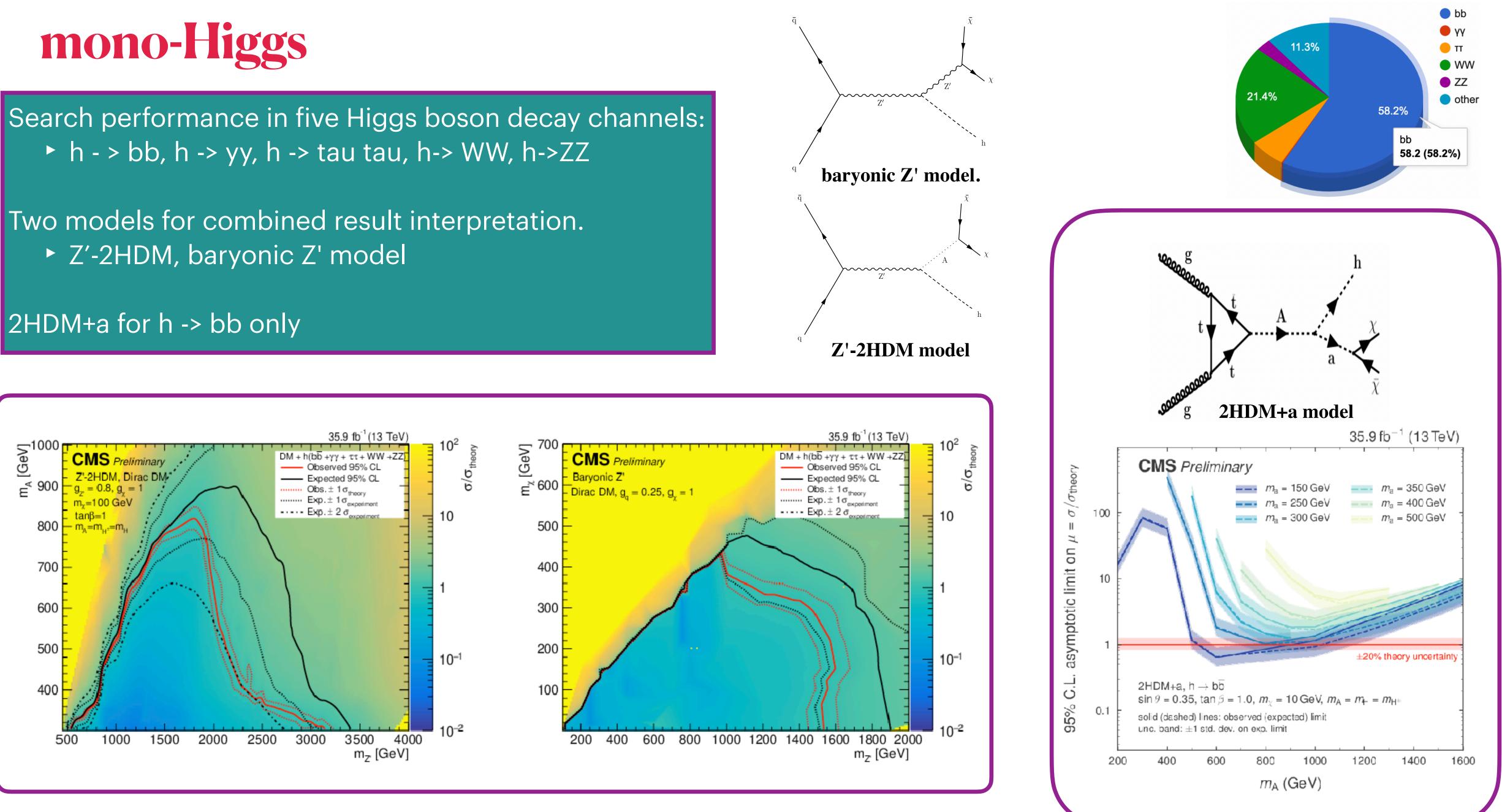
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h -> bb, h -> yy, h -> tau tau, h-> WW, h->ZZ

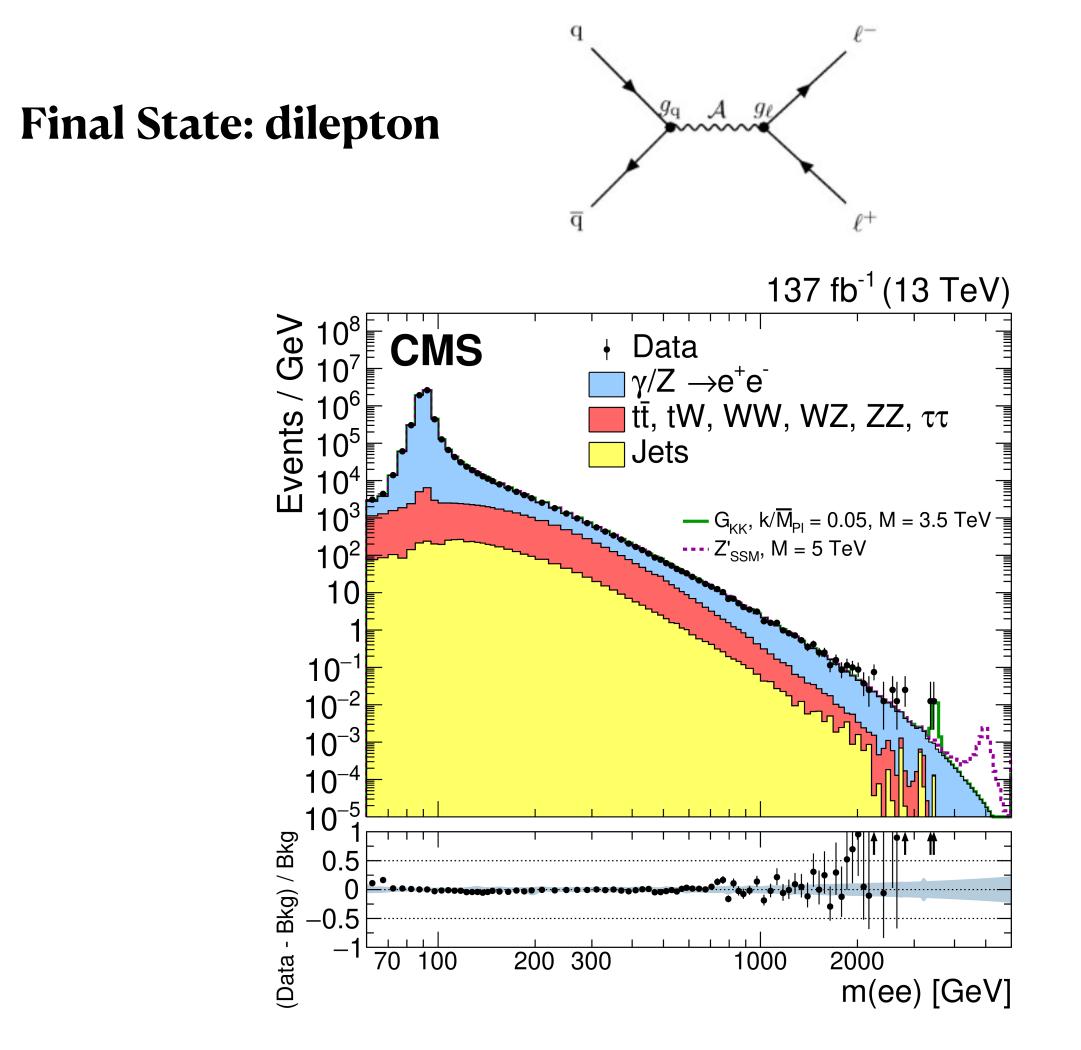
Z'-2HDM, baryonic Z' model



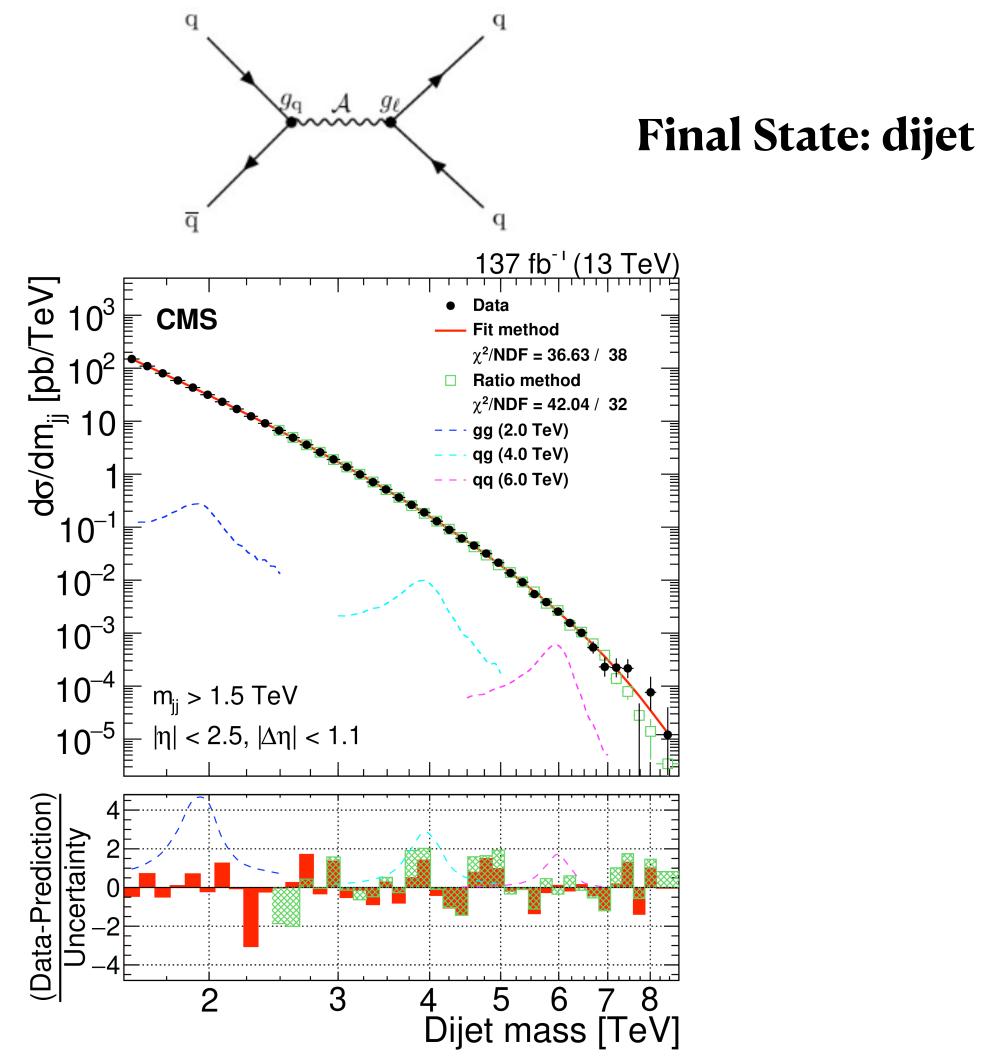
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## **Resonance search**



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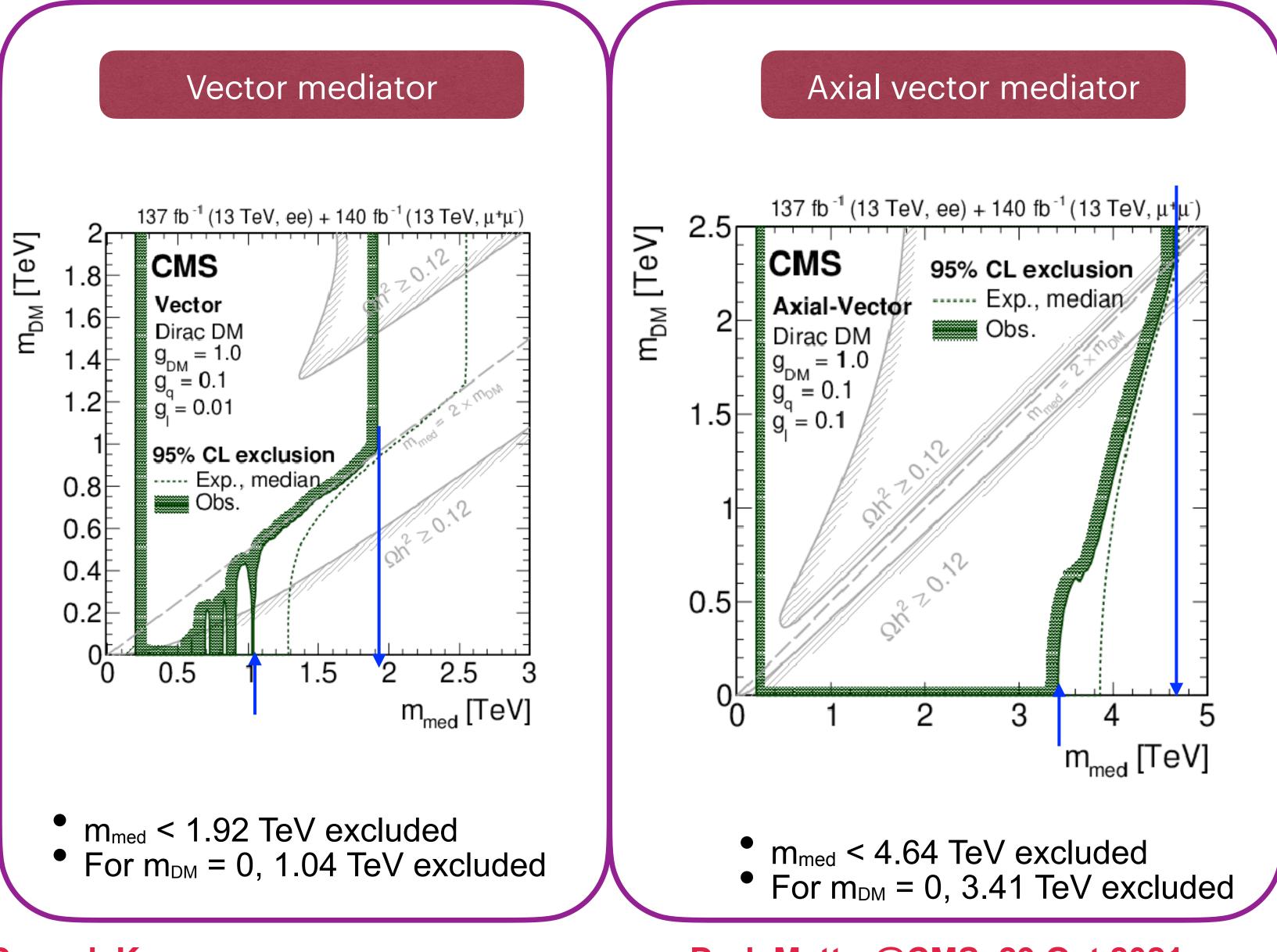


No peak is observed with respect to the SM background expectations.

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## **Resonance search**



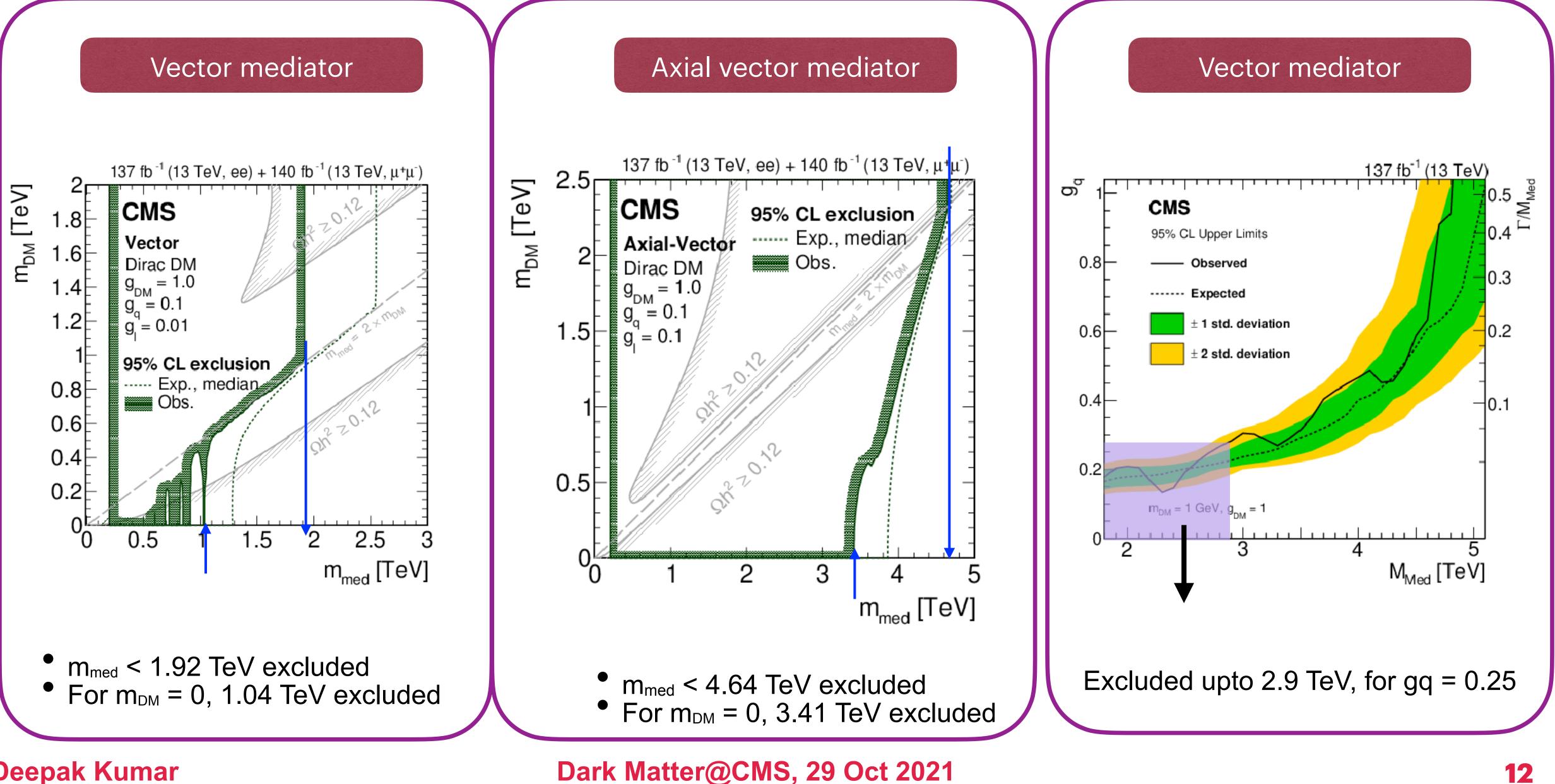
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## **Dilepton final sate**



## **Resonance search**



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## **Dilepton final sate**

## Dijet final sate



## Summary

- A brief summary of very exciting dark matter searches at CMS is presented.
- Full list of results are available here:

## • Public results

- No excess observed in any of the analysis.
- More results are coming soon with full run2

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## Dark Higgs

## • Dark Higgs:

- Higgs boson in Dark sector, responsible for generating the mass of DM.
- Signature :
  - Dark Higgs + missing energy
  - Dark Higgs (s) ->WW), WW (dileptonic decay

## • Basic selections:

- $160 \text{ GeV} < m_s < 400 \text{ GeV}$
- 195 GeV <  $m_z$  < 2500 GeV

## • Backgrounds:

• WW, Top, Drell–Yan

**Representative Feynman diagrams for the dark Higgs simplified model** 

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## Deepak Kumar

 $m_{\mathrm{T}}^{\ell\,\mathrm{r}}$ 

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# • Fir

q

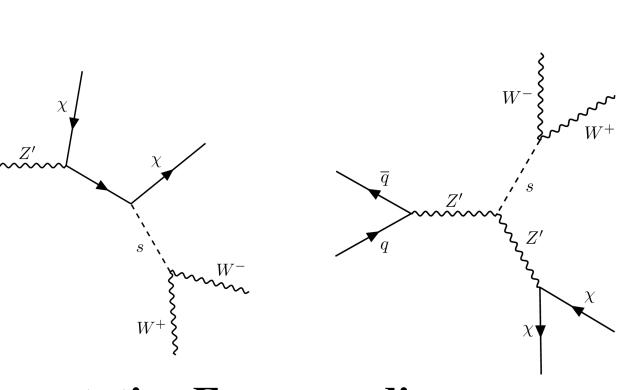
### **Observable:**

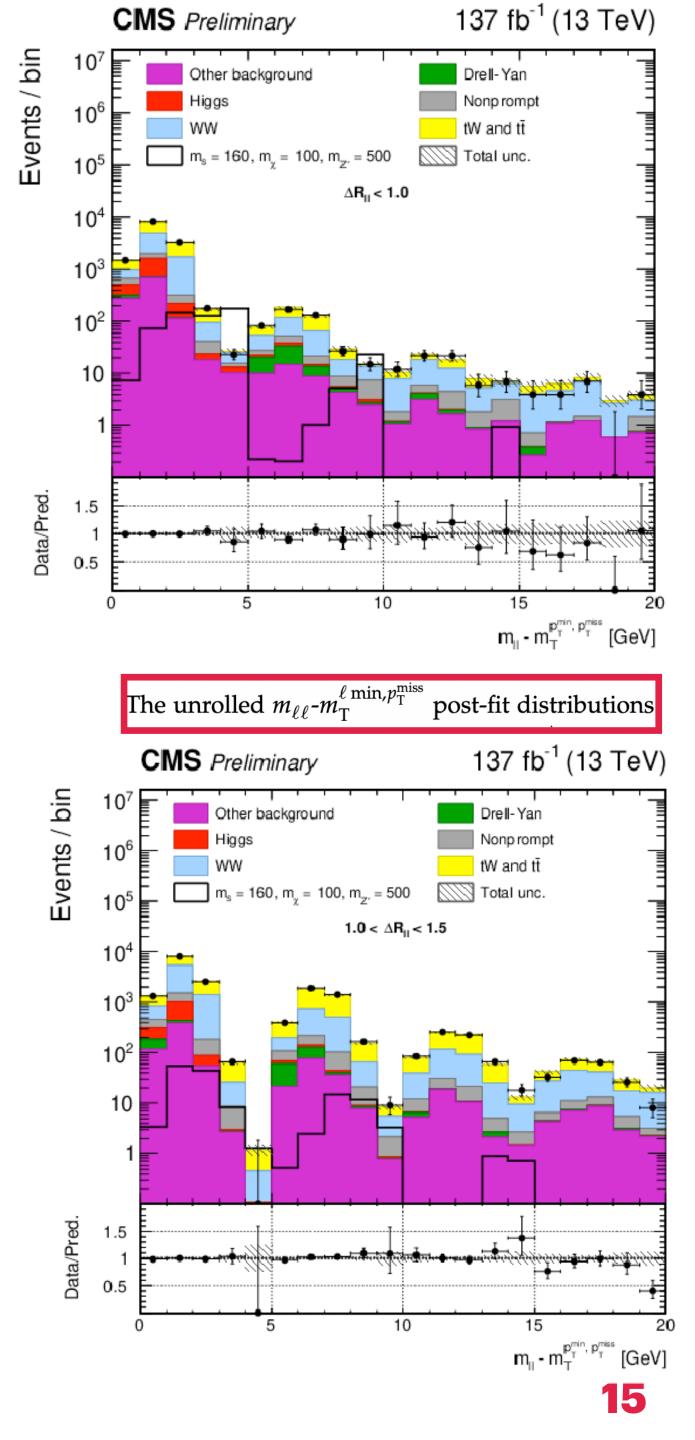
Transverse mass of the trailing lepton plus  $p_T^{miss}$  system

$$\min_{p_{\mathrm{T}}^{\mathrm{miss}}} = \sqrt{2p_{\mathrm{T}}^{\ell\,\mathrm{min}}p_{\mathrm{T}}^{\mathrm{miss}}\left[1 - \cos\Delta\phi(\vec{p}_{\mathrm{T}}^{\ell\,\mathrm{min}}, \vec{p}_{\mathrm{T}}^{\mathrm{miss}})\right]}$$

### • Final fit:

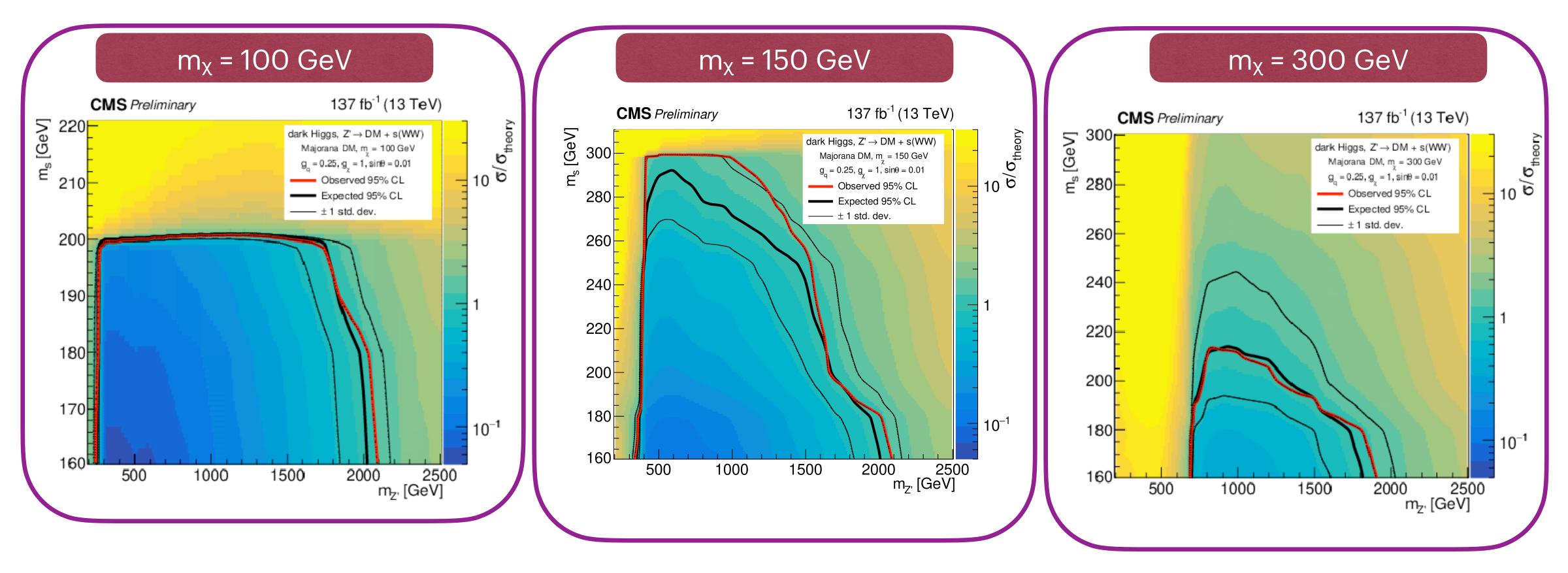
 2D profiled fit to the invariant mass of the dilepton system, m<sub>II</sub> and the m<sub>T</sub><sup>Imin, pTmiss</sup>





## Dark Higgs

- No significant deviation from the standard model prediction is observed.
- parameters



The most stringent limit is set for a  $m_x = 150$  GeV,  $m_s < 300$  GeV and  $480 < m_z < 1200$  GeV excluded

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Upper limits at 95% confidence level on the production cross section of dark matter are set in the dark Higgs model

