# Experimental Searches for Neutrinoless Double-Beta Decays and Prospects

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## a window opened to the unknown



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neutrino oscillations

only clear failure of Standard Model neutrino masses open a new window into Physics



## Majorana v

- experimental observations  $\mathbf{v} \neq \overline{\mathbf{v}} \rightarrow \mathbf{j}$ ust helicity
- a number of good theoretical motivations for a self-conjugated massive Majorana v
- only one experimental viable way to prove it

**Neutrinoless Double Beta Decay** 

# $\beta\beta$ decay $(A, Z) \rightarrow (A, Z+2)$ + something



a number of isotopes A-Z even reach stability only through  $\beta\beta$  decay

these are the  $\beta\beta$  candidates

a small group of them have high transition energies  $Q_{\beta\beta} > 1 \text{ MeV} \dots$  and other nice features

these are the golden  $\beta\beta$  candidates

<sup>76</sup>Ge <sup>100</sup>Mo <sup>130</sup>Te <sup>136</sup>Xe

## $2\nu 2\beta - 0\nu 2\beta$

## $(\mathbf{A}, \mathbf{Z}) \rightarrow (\mathbf{A}, \mathbf{Z}+2)+2e^{-}+2\overline{v}$

- $2\nu\beta\beta$  decay observed, rare decay
- 2<sup>nd</sup> weak process allowed by SM
- measured  $T_{1/2} > 10^{18} y$

## $(A,Z) \rightarrow (A,Z+2)+2e^{-2}$

- $0\nu\beta\beta$  decay  $\Delta$ L=2 process
- forbidden in Standard Model
- light mass Majorana v exchange or other non SM diagrams





... + other diagrams

## $0\nu2\beta$ signature

 $(A,Z) \rightarrow (A,Z+2)+2e^{2}+2\overline{\nu}$ 

```
(A,Z) \rightarrow (A,Z+2)+2e^{-}
```



#### main signature $\rightarrow$ a peak at $Q_{\beta\beta}$

- typically  $Q_{\beta\beta} \sim 1-3$  MeV (the higher, the lower is the background of spurious events that mimic the signal)
- 2 e<sup>-</sup> have a short range  $\rightarrow$  detection of their sum energy

#### processes spoiling the signal and their mitigation:

- $2\nu\beta\beta \rightarrow$  energy resolution
- background  $\rightarrow$  particle identification ( $\alpha$  rejection, SS vs MS ...)

 $\rightarrow$  daughter identification ...

$$\left(T_{1/2}^{0v}\right)^{-1} = G^{0v}(Q, Z) \cdot g_A^4 \left| M^{0v} \right|^2 \cdot m_{ee}^2$$



Standard (dominant) Mechanism: exchange of a light Majorana v

## $0v2\beta$ half-life



## decay observation implies:

- → DISCOVERY
- measurement of the half-life
- $\rightarrow$  v in a Majorana particle
- → m<sub>ee</sub> is measured

$$(T_{1/2}^{0v})^{-1} = G^{0v}(Q, Z) \cdot g_A^4 |M^{0v}|^2 \cdot m_{ee}^2$$

## phase space

- is computed
- favors high  $Q_{\beta\beta}$  isotopes



**phase space** is and additional motivation for chosing high  $Q_{\beta\beta}$  isotopes (the stronger motivation is background suppression)

$$(T_{1/2}^{0v})^{-1} = G^{0v}(Q, Z) \cdot g_A^4 |M^{0v}|^2 \cdot m_{ee}^2$$

#### nuclear matrix element - is evaluated on the basis of different nuclear models

- model dependent
- "quenched  $g_A$ "? used to correct the systematic over-predictions of  $\beta$  and  $2\nu\beta\beta$  measured rates
- does it apply also to  $0\nu\beta\beta$  ? (if so predicted T<sub>1/2</sub> underestimated by a factor 2-6)



$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q, Z) \cdot g_A^4 |M^{0\nu}|^2 \cdot m_{ee}^2$$

$$m_{ee} = \sum_{k} U_{ek}^{2} m_{i} = c_{12}^{2} c_{23}^{2} m_{1} + s_{12}^{2} c_{13}^{2} e^{i\alpha} m_{2} + s_{13}^{2} e^{i\beta} m_{3}$$

#### **3** neutrinos eigenstates + mixing

- 3 mixing angles → c<sub>ii</sub> s<sub>ii</sub>
- 2 mass differences  $\rightarrow \Delta m^2_{solar} \Delta m^2_{atm}$
- 2 phases  $\rightarrow \alpha$ ,  $\beta$  (Majorana phases)
- mass ordering  $\rightarrow$  lightest neutrino mass



$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q, Z) \cdot g_A^4 |M^{0\nu}|^2 \cdot m_{ee}^2$$

$$m_{ee} = \sum_{k} U_{ek}^{2} m_{i} = c_{12}^{2} c_{23}^{2} m_{1} + s_{12}^{2} c_{13}^{2} e^{i\alpha} m_{2} + s_{13}^{2} e^{i\beta} m_{3}$$

#### **3** neutrinos eigenstates + mixing

- 3 mixing angles  $\rightarrow c_{ii} s_{ii}$
- 2 mass differences  $\rightarrow \Delta m^2_{solar} \Delta m^2$
- 2 phases  $\rightarrow \alpha$ ,  $\beta$  (Majorana phase
- mass ordering



Vissani et al. Phys Rev D 100, 073003 (2019) probability distribution for  $m_{ee}$  obtained using the Likelihood for  $\Sigma$ 

### On result basis:

•  $T_{1/2}$  measurements can't be compared directly, the conversion into  $m_{ee}$  is not trivial because of the uncertainties in NME



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On result basis:

- a pragmatic solution is to convert  $\ensuremath{\,T_{1/2}}\xspace \rightarrow m_{ee}$  range and compare the range



## 0v2β ranking experiments

## On result basis:

 neglect correlations (a by model comparison could be better) forget that sometimes a Nuclear Model works better for one isotope



# $\rm m_{ee}$ intervals from $\rm T_{_{1/2}}$



this is a personal selection of high sensitivity running or just closed expts.

Majorana mass [eV]

## On sensitivity basis (future expts):

- sensitivities or discovery potentials are defined by 3 parameters
  - Mass Time = Exposure
  - Energy Resolution (often extrapolated ..)
  - **Background** (guessed, desired ...)
- extrapolations, guess ... can be strongly biased !
- robustness of hypotheses on the achievement of energy resolution and background is hard to judge

 $\rightarrow$  experiments with a long history of precursors, with clear strategies for background suppression are more reliable (but this is not a number !)

## $0\nu 2\beta$ sensitivity

 $S_{0v}$ 

Half-life corresponding to the minimum detectable number of events over background at a given confidence level:

M × T = exposure
△ E = energy resolution (ROI)
B = background rate / keV

$$S_{0\nu\beta\beta} \propto \sqrt{\frac{M \times T}{B \times \Delta E}}$$

"zero bkg"  $\rightarrow M \times T \times B \times \Delta E < 1$  $S_{0 \nu \beta \beta} \propto M \times T$ 



Neutrino 2020- Agostini, Benato, Dewiler, Menendez, Vissani



## selected experiments (Lapologize for those not included)

running or with a plan for near/far the future
competitive

## Experiments

Collaboration	Isotope	Technique	mass (0vββ isotope)	Status
CANDLES-III	<sup>48</sup> Ca	305 kg CaF2 crystals in liquid scintillator	0.3 kg	Operating
CANDLES-IV	<sup>48</sup> Ca	CaF <sub>2</sub> scintillating bolometers	TBD	R&D
GERDA	<sup>76</sup> Ge	Point contact Ge in active LAr	44 kg	Complete
MAJORANA DEMONSTRATOR	<sup>76</sup> Ge	Point contact Ge in Lead	30 kg	Operating
LEGEND 200	<sup>76</sup> Ge	Point contact Ge in active LAr	200 kg	Construction
LEGEND 1000	<sup>76</sup> Ge	Point contact Ge in active LAr	1 tonne	R&D
SuperNEMO Demonstrator	<sup>82</sup> Se	Foils with tracking	7 kg	Construction
SELENA	<sup>82</sup> Se	Se CCDs	<1 kg	R&D
NvDEx	<sup>82</sup> Se	SeF <sub>6</sub> high pressure gas TPC	50 kg	R&D
ZICOS	<sup>96</sup> Zr	10% natZr in liquid scintillator	45 kg	R&D
AMoRE-I	<sup>100</sup> Mo	<sup>40</sup> CaMoO <sub>4</sub> scintillating bolometers	6 kg	Construction
AMoRE-II	<sup>100</sup> Mo	Li <sub>2</sub> MoO <sub>4</sub> scintillating bolometers	100 kg	Construction
CUPID	$^{100}$ Mo	Li <sub>2</sub> MoO <sub>4</sub> scintillating bolometers	250 kg	R&D
COBRA	<sup>116</sup> Cd/130Te	CdZnTe detectors	10 kg	Operating
CUORE	<sup>130</sup> Te	TeO <sub>2</sub> Bolometer	206 kg	Operating
SNO+	<sup>130</sup> Te	0.5% natTe in liquid scintillator	1300 kg	Construction
SNO+ Phase II	<sup>130</sup> Te	2.5% natTe in liquid scintillator	8 tonnes	R&D
Theia-Te	<sup>130</sup> Te	5% natTe in liquid scintillator	31 tonnes	R&D
KamLAND-Zen 400	<sup>136</sup> Xe	2.7% in liquid scintillator	370 kg	Complete
KamLAND-Zen 800	<sup>136</sup> Xe	2.7% in liquid scintillator	750 kg	Operating
KamLAND2-Zen	<sup>136</sup> Xe	2.7% in liquid scintillator	~tonne	R&D
EXO-200	<sup>136</sup> Xe	Xe liquid TPC	160 kg	Complete
nEXO	<sup>136</sup> Xe	Xe liquid TPC	5 tonnes	R&D
NEXT-WHITE	<sup>136</sup> Xe	High pressure GXe TPC	~5 kg	Operating
NEXT-100	<sup>136</sup> Xe	High pressure GXe TPC	100 kg	Construction
PandaX	<sup>136</sup> Xe	High pressure GXe TPC	~tonne	R&D
AXEL	<sup>136</sup> Xe	High pressure GXe TPC	~tonne	R&D
DARWIN	<sup>136</sup> Xe	natXe liquid TPC	3.5 tonnes	R&D
LZ	<sup>136</sup> Xe	natXe liquid TPC		R&D
Theia-Xe	<sup>136</sup> Xe	3% in liquid scintillator	50 tonnes	R&D
R&D	Const	truction Operating	Complete	

Updated from J. Wilkerson

#### Neutrino 2020- Dewiler

## GERDA

- <sup>76</sup>Ge  $Q_{\beta\beta}$  = 2.0 MeV
- enrichement ~ 90%
- 37 enrichGe HPGe in LAr





- resolution FWHM ~ 3.6 keV at  $Q_{BB}$
- **PSA** used to separate SingleSite from MultiSite
- status: CLOSED on Nov 2019
  - exposure I+II: 127.2 kg·yr
- background: ~ 5.2 10<sup>-4</sup> counts/(keV kg y)
  - T<sub>1/2</sub><sup>0νββ</sup> > 1.8 x 10<sup>26</sup> yr (90% C.L.)
  - **〈 m<sub>ββ</sub> 〉 <** (79 180) meV [NMEs]



#### Risultati di fase II PHYSICAL REVIEW LETTERS 125, 252502 (2020)



## **LEGEND-200**





Upgraded GERDA cryostat: new lock, cabling detector suspension



coupled fibers



New String layout, front end electronics

- **start** in 2022 (delay due to COVID emergency)
- 200 kg in upgraded GERDA infrastructure
  - → new dets = inverted coax  $\rightarrow$  large mass
- background
  - → 1-2 10<sup>-4</sup> counts/(keV kg y) (5 times improvement)
  - dominated by <sup>42</sup>K  $\beta$ -decays &  $\alpha$ -surface emitters

T<sub>1/2</sub><sup>0νββ</sup> ~ 10<sup>27</sup> y or m<sub>ee</sub> 35-75 meV



## LEGEND-1000



- @ where ? LNGS or SNOLAB
- **detectors =** inverted coaxial
- **mass =** 1000 kg
- **bkg** = 0.1 cts/(FWHM ton y)  $\rightarrow$  30 times better than GERDA-II



## CUORE



- <sup>130</sup>Te  $Q_{\beta\beta}$  = 2.5 MeV
- enrichment NO i.a.~ 30%
- 988 TeO, bolometers @ 10 mK
- FWHM ~ 7 keV at  $Q_{\beta\beta}$
- status running (> 1344 kg yr of raw exposure acquire









- exposure: 1038 kg yr
- background:
  - → 1.5 10<sup>-2</sup> counts/(keV kg y)
  - $\textbf{\textbf{+}}$  dominated by  $\alpha$  particles
- T<sub>1/2</sub><sup>0νββ</sup> > 2.2 x 10<sup>25</sup> yr (90% C.L.)
- **〈 m<sub>ββ</sub> 〉 <** (90 305) meV



## CUPID



- @ where ? LNGS in CUORE infrastructure
- <sup>100</sup>Mo Q<sub>bb</sub> = 3.0 MeV
- enrichment ~ 90%
- detectors ~1500 Li<sub>2</sub>MoO<sub>4</sub> scintillating bolometers





- mass = 250 kg of <sup>100</sup>Mo
- **bkg** = 10<sup>-4</sup> ckky
  - → CUORE without  $\alpha$  induced bkg
  - → dominated by  $2\nu\beta\beta$  pile-up

 $T_{1/2}^{0\nu\beta\beta} > 10^{27}$ 

• scintillating <sup>100</sup>Mo crystals deplCa<sup>100</sup>MoO<sub>4</sub> or Li<sub>2</sub><sup>100</sup>MoO<sub>4</sub> Photon channel MMC Gold wire Ge wafer CMO Crystal MMC Gold film Gold wire Phonon channel



- MMC sensors fast !! techn. challenging !!
- underground lab under excavation

	AMoRE-Pilot	AMoRE-I	AMoRE-II
Mass [kg]	1.9	~6.1	~200
Channels	12	36	~1000
BKG goal [ckky]	0.01	0.001	0.0001
Sensitivity [year]	~10 <sup>24</sup>	~10 <sup>25</sup>	~5×10 <sup>26</sup>
Sensitivity [meV]	380 to 640	120 to 200	17 to 29
Location	Y2L	Y2L	Yemilab
schedule	2017 to 2018	2019~	2021~



#### PHYSICAL REVIEW LETTERS 123, 161802 (2019)

#### **EXO-200** enrichment 200 kg LXe

•  ${}^{136}$ Xe  $Q_{\beta\beta}$  = 2.54 MeV

energy measured using two signals: ionization signal drifted to crossed wire planes scintillation (175nm) collected by APD

topology used to separate Signal/Background

separation of SingleSite/MultiSite events





- FWHM ~ 65 keV at  $Q_{RR}$
- status closed
- exposure: 234.1 kg·yr
- $T_{1/2}^{0\nu\beta\beta} > 3.5 \times 10^{25} \text{ yr} (90\% \text{ C.L.})$ •
- **〈 m<sub>βB</sub> 〉 < (78 239) me**∨ •



- 5000 kg LXe
- $T_{1/2}^{0\nu\beta\beta} > 9 \ 10^{27}$









- **gas TPC** with 2 dedicated readout planes
- event topology
- FWHM ~ 25 keV



## KamLAND-Zen

## LS with <sup>136</sup>Xe balloon



## KamLAND-Zen



## KamLAND-Zen 400

- closed
- limit 1.1 10<sup>26</sup> y



# KamLAND-Zen 800running since Jan 2019

• Very preliminary results with 132.7 days data  $T_{1/2}^{0\nu} > 4 \times 10^{25}$  year (90% C.L.) Sensitivity : 8 × 10<sup>25</sup> year

• target 5 yr sensitivity 5  $1\Omega^{26}$ 

## SNO+



- 780 tonne scintillator (47% filled, COVID stopped)
- loaded with Te(0.5%)-butanediol
   1220 kg 130Te
  - → 1330 kg <sup>130</sup>Te
  - → <sup>130</sup>Te  $Q_{\beta\beta}$  = 2.5 MeV
- predicted 460 p.e./MeV  $\rightarrow$  3%  $\sigma/E$  FWHM=180 keV
- challenge: stability (for 35 I > 26 months)







# we will likely have these experiments approaching their design sensitivity



SNO+ and AMORE are still in a very preliminary stage, schedule and sensitivity may be subject to delay ...



<sup>76</sup>Ge LEGEND-1000 10<sup>28</sup> yr

future leaders

~ 2030

<sup>100</sup> Mo AMORE II CUPID	5 10 <sup>26</sup> yr 10 <sup>27</sup> yr
<sup>130</sup> Te SNO+II	10 <sup>27</sup> yr
<sup>136</sup> Xe nEXO	10 <sup>28</sup> yr



# Thanks for your attention !

## Majorana D

Latest Release: First unblinding of data 16 kg-yr exposure [PRC 100 025501 (2019)]

Median  $T_{1/2}$  Sensitivity: 4.8 × 10<sup>25</sup> yr

Full Exposure Limit:  $T_{1/2} > 2.7 \times 10^{25}$  yr (90% CL)

Background Index at 2039 keV in lowest background config: 11.9 ± 2.0 cts/(FWHM t yr)



P-type Point-Contact (PPC)

## $0\nu2\beta$ candidates

- in most cases driven by **detector characteristic** 
  - → <sup>76</sup>Ge with Ge diodes
  - → <sup>136</sup>Xe with TPCs
  - > bolometers and scintillators have multiple choices
- isotopic abundance as high as possible

   → (not only) money issue

- Q-value as high as possible
  - → phase space
  - → background







the race to the exclusion of IO region started

- increase detector size ( $M_{\beta\beta}$  i.e. the source size)
- reduce background with active rejection techniques



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## **Discovery Sensitivity Comparison**

