

Theory review on Charged Lepton Flavor Violation

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Introduction

Before the LHC started operating we all hoped for **great discoveries...**

A dense tropical rainforest with sunlight filtering through the canopy. The scene is filled with various types of green plants, including palm trees and broad-leafed species. The lighting is bright, creating a high-contrast scene with deep shadows and bright highlights where the sun hits the leaves.

Microscopic
black holes

Extra dimensions

Supersymmetry

Compositeness

LHC expectations

LHC results...

**125 GeV
palm tree**

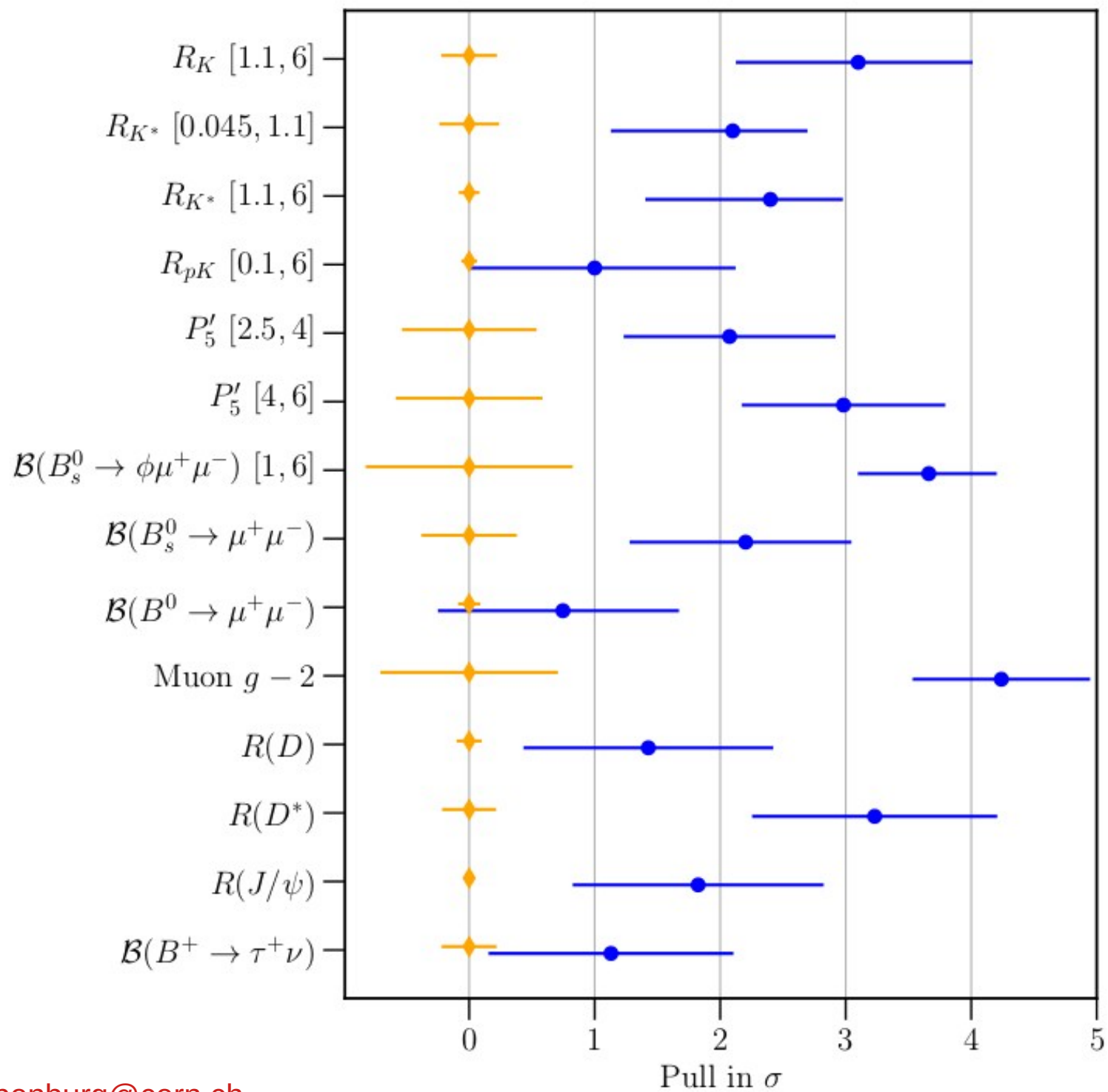


LHC results...

**125 GeV
palm tree**

**Flavor
anomalies**





Introduction

Do we have a good reason to go **Beyond the Standard Model?**

Introduction

Do we have a good reason to go **Beyond the Standard Model?**



Neutrinos!

The lepton sector is still
to be understood!

Neutrinos and the lepton sector

Dear radioactive Ladies and Gentlemen...

Physikalisches Institut
der Eidg. Technischen Hochschule
Zürich

Zürich, 4. Dez. 1930
Gloriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich halbvollst anzuhören bitte, Ihnen des näheren auseinandersetzen wird, bin ich angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie des kontinuierlichen beta-Spektrums auf einen verzweifelten Ausweg verfallen um den "Wechselsatz" (1) der Statistik und den Energiesatz zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale Teilchen, die ich Neutronen nennen will, in den Kernen existieren, welche den Spin $1/2$ haben und das Ausschliessungsprinzip befolgen und sich von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen dürfte von derselben Grössenordnung wie die Elektronenmasse sein und jedenfalls nicht grösser als $0,01$ Protonenmasse.- Das kontinuierliche beta-Spektrum wäre dann verständlich unter der Annahme, dass beim beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert wird, derart, dass die Summe der Energien von Neutron und Elektron konstant ist.



December 4th, 1930

Letter to his colleagues in Tübingen

1930

Pauli's neutrino hypothesis

Open questions

What is the origin of neutrinos masses?

Are they Dirac or Majorana?

What is the absolute scale of neutrino masses?

What is the mass ordering?

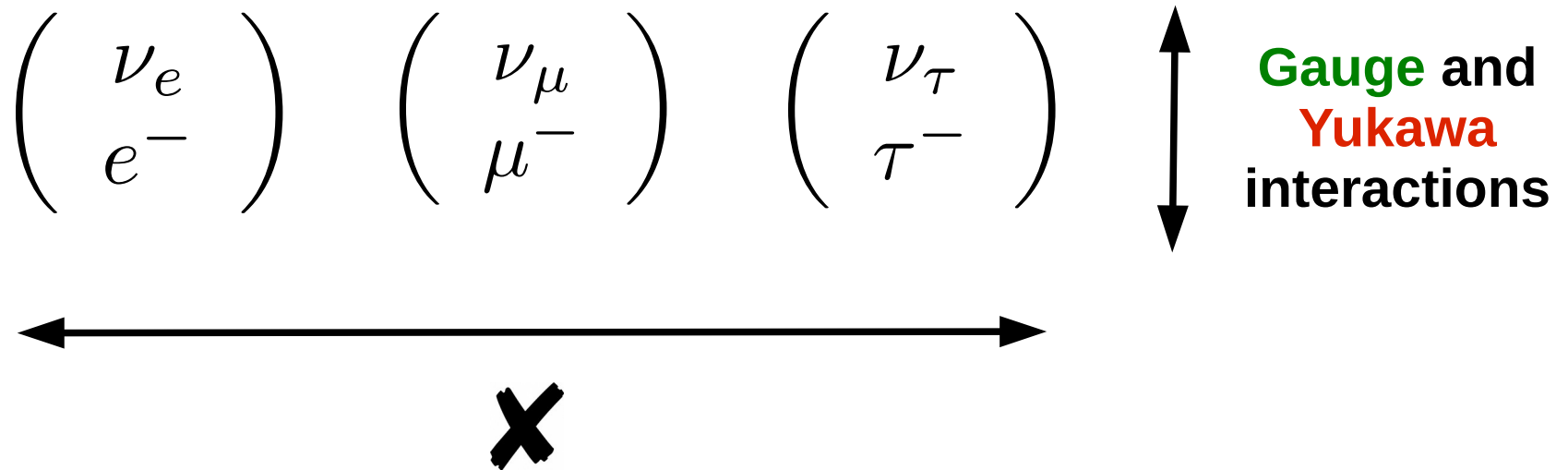
Are there more than three neutrinos? Maybe sterile?

Is there CP violation in the lepton sector?

Is lepton flavor universality violated?

Lepton flavor violation

In the **Standard Model**, three copies of the leptonic SU(2) doublet are introduced

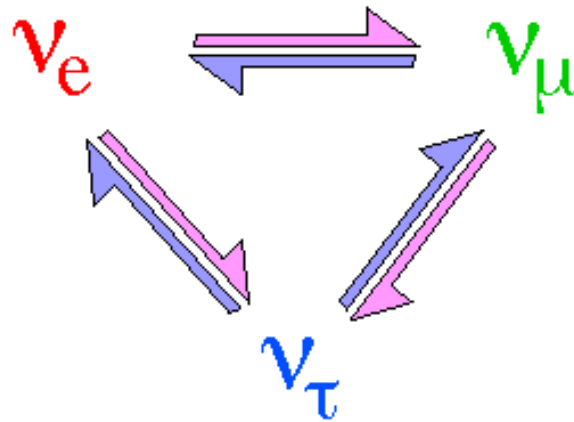
$$\begin{pmatrix} \nu_e \\ e^- \end{pmatrix} \quad \begin{pmatrix} \nu_\mu \\ \mu^- \end{pmatrix} \quad \begin{pmatrix} \nu_\tau \\ \tau^- \end{pmatrix} \quad \begin{array}{c} \updownarrow \\ \text{Gauge and} \\ \text{Yukawa} \\ \text{interactions} \end{array}$$


Is **lepton flavor** a conserved quantity?

Neutrino oscillations: LFV

We already know the answer: **NO**

Neutrino flavor oscillations: flavor violating process!



Then... is this process a sign of BSM physics?

Yes it is!



What about CLFV?

In conclusion, lepton flavor is **not** conserved: there is **lepton flavor violation (LFV)**

However... what about **charged lepton flavor violation (CLFV)**?

$$\mu^- \rightarrow e^- \gamma$$

$$h \rightarrow \mu^- \tau^+$$

$$\tau^- \rightarrow \mu^- \mu^+ \mu^-$$

$$\pi^0 \rightarrow e^- \mu^+$$

$$K_L^0 \rightarrow \pi^0 e^- \mu^+$$

...

Never observed...

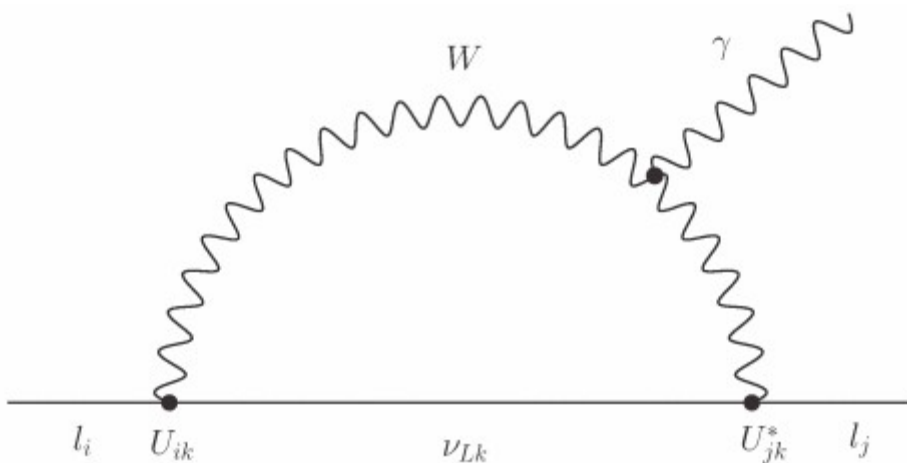
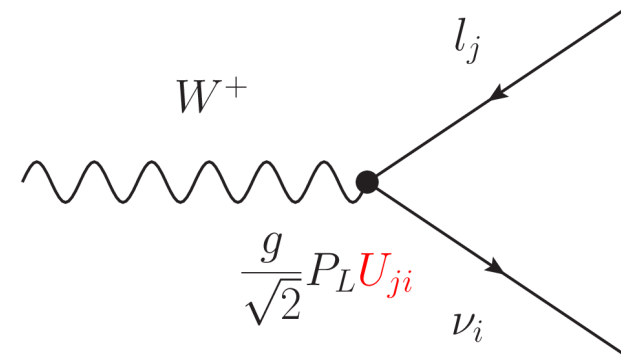
What about CLFV?

SM + Dirac neutrino masses

$$-\mathcal{L}_{cc} = \frac{g}{\sqrt{2}} U_{ji} \bar{l}_j \gamma^\mu P_L \nu_i W_\mu^- + \text{c.c.}$$

U : lepton mixing matrix

[analog of the CKM matrix in the lepton sector]



$$\text{Br}(\mu \rightarrow e \gamma) = \frac{3\alpha}{32\pi} \left| \sum_k U_{ek} U_{\mu k}^* \frac{m_{\nu k}^2}{m_W^2} \right|^2 \lesssim 10^{-54}$$

[Petcov, 1977]

Since neutrino masses are the **only source** of LFV, all cLFV amplitudes are strongly suppressed (in fact, **GIM** suppressed)

Why do we care about LFV?

The observation of **CLFV** would be a clear signal of (non-trivial) physics beyond the Standard Model

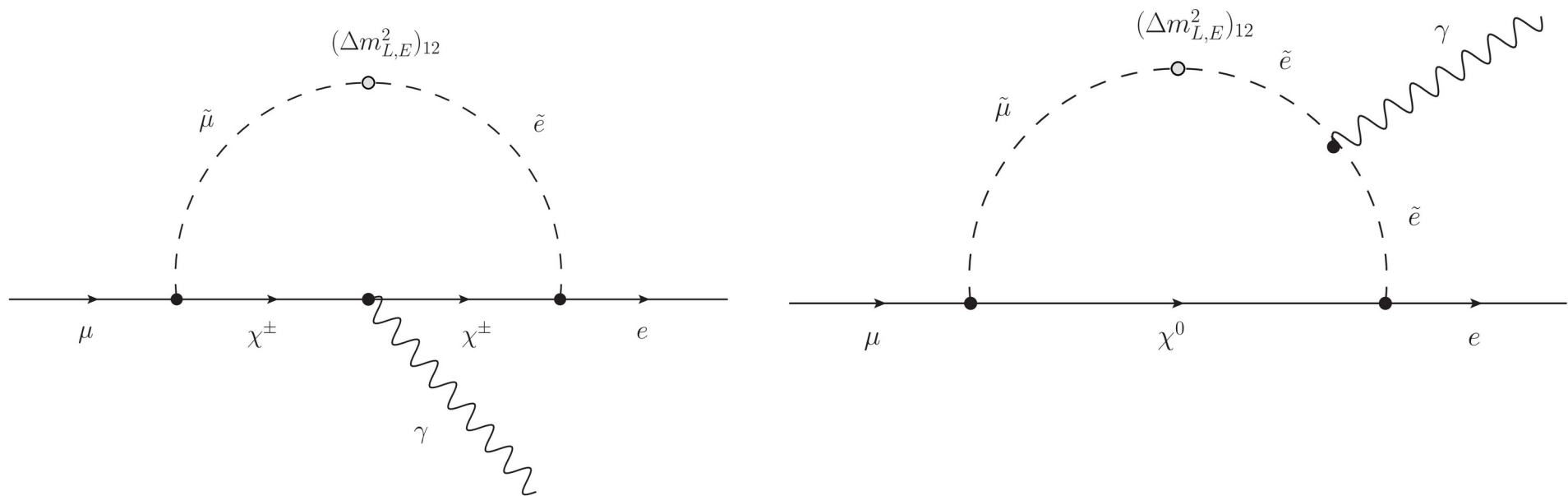
In fact, most **BSM models** predict **large CLFV rates**

We can probe **very high energy scales!**

$$\mathcal{O} = \frac{c_{e\mu}}{\Lambda^2} \bar{\mu} e \bar{e} e \quad \Rightarrow \quad \frac{\Lambda}{\sqrt{c_{e\mu}}} \gtrsim 100 \text{ TeV}$$

Why do we care about LFV?

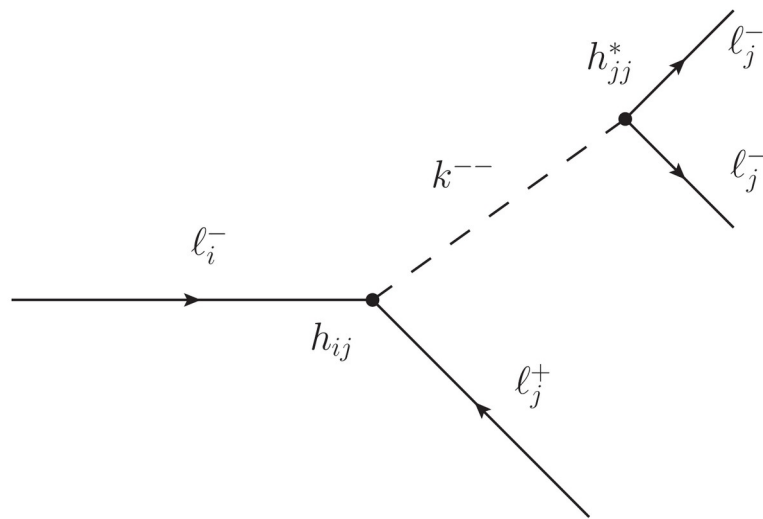
Example 1: Supersymmetric models



Sleptons: a whole new sector coupled to the **SM leptons**
Strong constraints on the **off-diagonal** soft terms

Why do we care about LFV?

Example 2: Babu-Zee model



$$\text{BR} \sim \left| \frac{h_{ij} h_{jj}^*}{m_k^2} \right|^2$$

Small **off-diagonal h couplings** and/or **heavy k's** are required

Experimental projects

LFV Process	Present Bound	Future Sensitivity
$\mu \rightarrow e\gamma$	4.2×10^{-13}	6×10^{-14} (MEG)
$\tau \rightarrow e\gamma$	3.3×10^{-8}	$\sim 10^{-8} - 10^{-9}$ (B factories)
$\tau \rightarrow \mu\gamma$	4.4×10^{-8}	$\sim 10^{-8} - 10^{-9}$ (B factories)
$\mu \rightarrow 3e$	1.0×10^{-12}	$\sim 10^{-16}$ (Mu3e)
$\tau \rightarrow 3e$	2.7×10^{-8}	$\sim 10^{-9} - 10^{-10}$ (B factories)
$\tau \rightarrow 3\mu$	2.1×10^{-8}	$\sim 10^{-9} - 10^{-10}$ (B factories)
$\mu^-, \text{Au} \rightarrow e^-, \text{Au}$	7.0×10^{-13}	—
$\mu^-, \text{SiC} \rightarrow e^-, \text{SiC}$	—	2×10^{-14} (DeeMe)
$\mu^-, \text{Al} \rightarrow e^-, \text{Al}$	—	$10^{-15} - 10^{-17}$ (COMET)
$\mu^-, \text{Ti} \rightarrow e^-, \text{Ti}$	4.3×10^{-12}	$10^{-17} - 10^{-18}$ (Mu2e)
		$\sim 10^{-18}$ (PRISM/PRIME)

Experimental projects

History of $\mu \rightarrow e\gamma$, $\mu N \rightarrow eN$, and $\mu \rightarrow 3e$

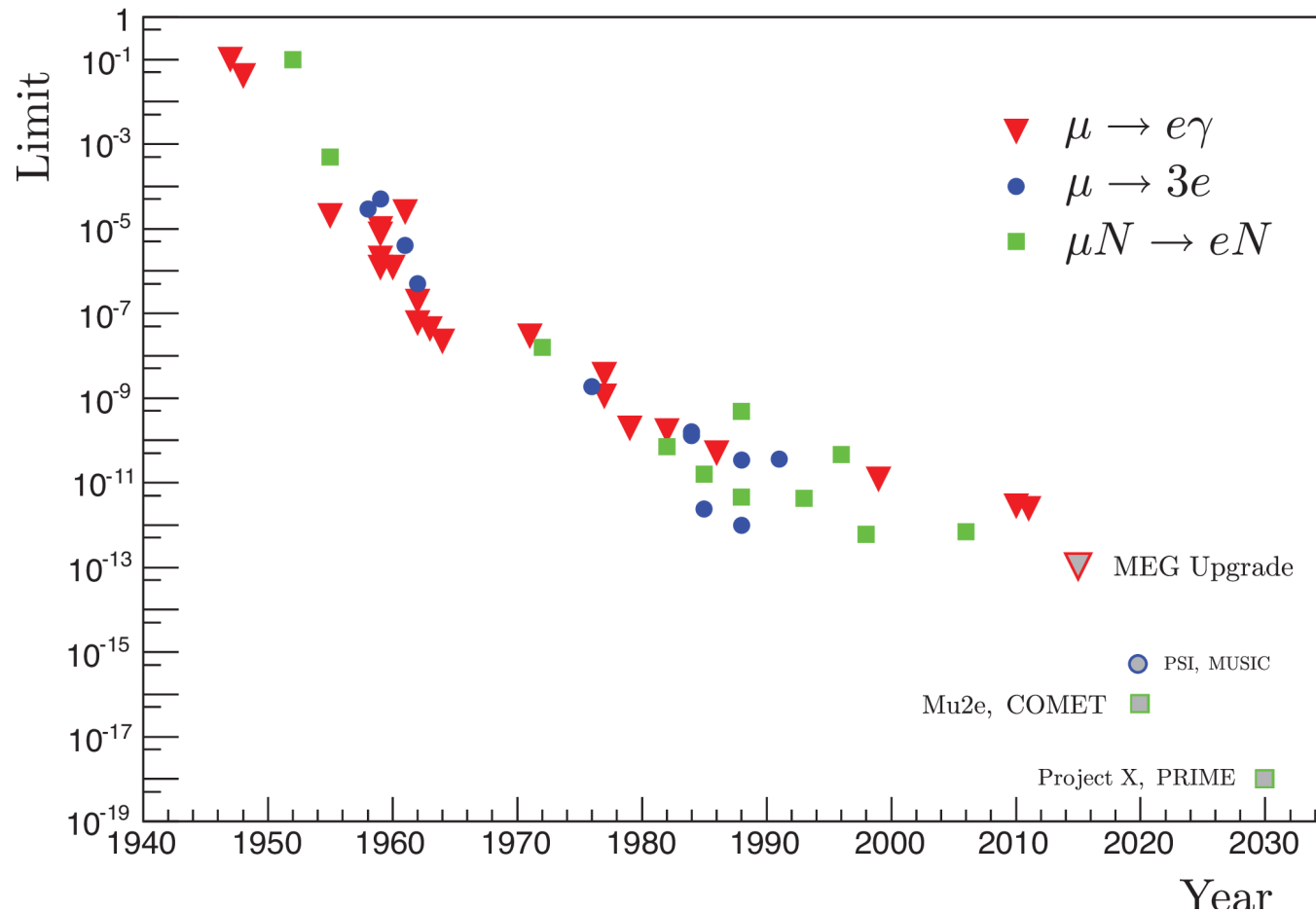
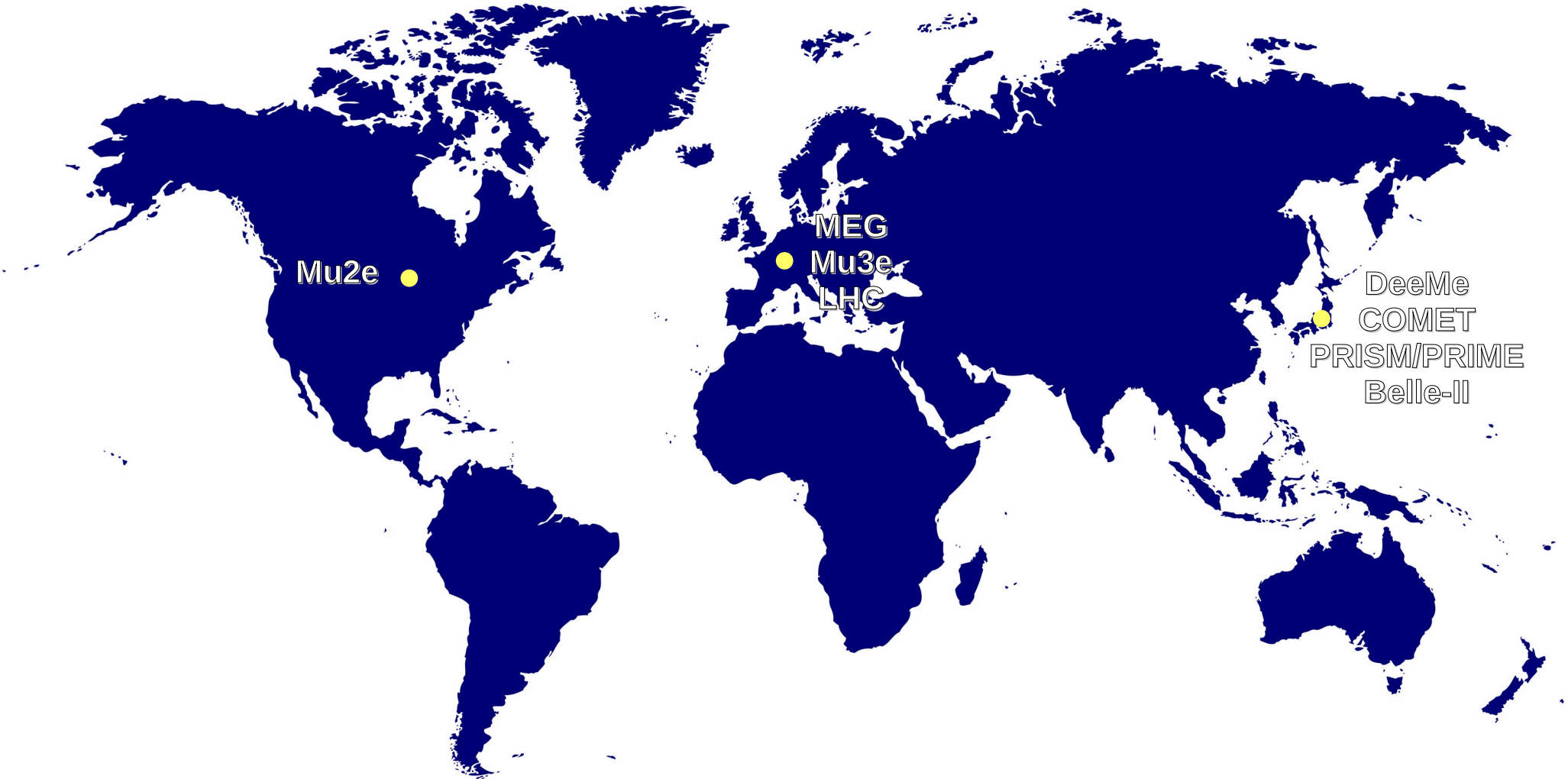


Figure taken from Bernstein & Cooper [arXiv:1307.5787]

Experimental projects



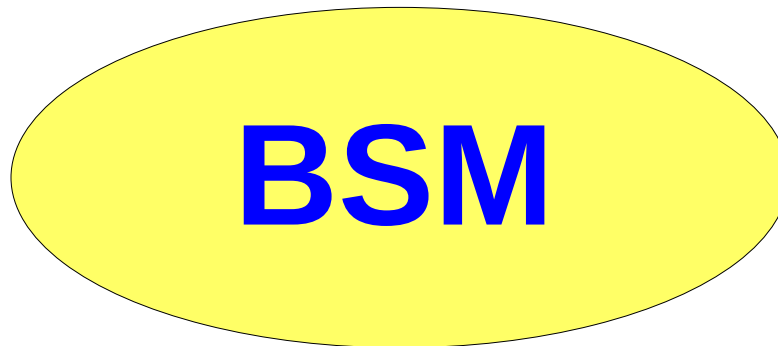
LFV : Where to look for?

$$l_i \rightarrow l_j \gamma$$

$$l_i \rightarrow 3 l_j$$

$$l_i \rightarrow l_j l_k l_k$$

$\mu - e$
conversion in nuclei



LFV at colliders

$$M \rightarrow l_i l_j$$

LFV : Where to look for?

Everywhere!

ratio	LHT	MSSM (dipole)	MSSM (Higgs)	SM4
$\frac{\text{Br}(\mu^- \rightarrow e^- e^+ e^-)}{\text{Br}(\mu \rightarrow e \gamma)}$	0.02... 1	$\sim 6 \cdot 10^{-3}$	$\sim 6 \cdot 10^{-3}$	0.06... 2.2
$\frac{\text{Br}(\tau^- \rightarrow e^- e^+ e^-)}{\text{Br}(\tau \rightarrow e \gamma)}$	0.04... 0.4	$\sim 1 \cdot 10^{-2}$	$\sim 1 \cdot 10^{-2}$	0.07... 2.2
$\frac{\text{Br}(\tau^- \rightarrow \mu^- \mu^+ \mu^-)}{\text{Br}(\tau \rightarrow \mu \gamma)}$	0.04... 0.4	$\sim 2 \cdot 10^{-3}$	0.06... 0.1	0.06... 2.2
$\frac{\text{Br}(\tau^- \rightarrow e^- \mu^+ \mu^-)}{\text{Br}(\tau \rightarrow e \gamma)}$	0.04... 0.3	$\sim 2 \cdot 10^{-3}$	0.02... 0.04	0.03... 1.3
$\frac{\text{Br}(\tau^- \rightarrow \mu^- e^+ e^-)}{\text{Br}(\tau \rightarrow \mu \gamma)}$	0.04... 0.3	$\sim 1 \cdot 10^{-2}$	$\sim 1 \cdot 10^{-2}$	0.04... 1.4
$\frac{\text{Br}(\tau^- \rightarrow e^- e^+ e^-)}{\text{Br}(\tau^- \rightarrow e^- \mu^+ \mu^-)}$	0.8... 2	~ 5	0.3... 0.5	1.5... 2.3
$\frac{\text{Br}(\tau^- \rightarrow \mu^- \mu^+ \mu^-)}{\text{Br}(\tau^- \rightarrow \mu^- e^+ e^-)}$	0.7... 1.6	~ 0.2	5... 10	1.4... 1.7
$\frac{\text{R}(\mu \text{Ti} \rightarrow e \text{Ti})}{\text{Br}(\mu \rightarrow e \gamma)}$	$10^{-3} \dots 10^2$	$\sim 5 \cdot 10^{-3}$	0.08... 0.15	$10^{-12} \dots 26$

Table taken from Buras et al [arXiv:1006.5356]

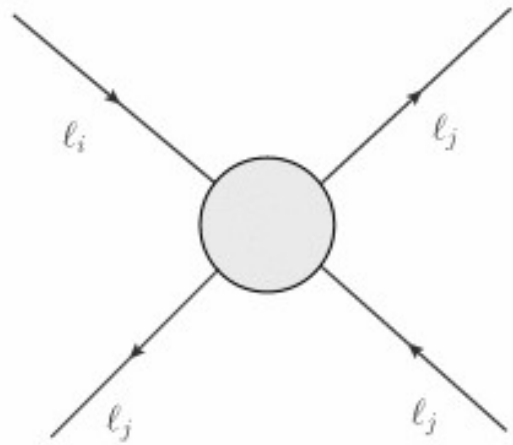
$$l_i \rightarrow 3 l_j \text{ VS } l_i \rightarrow l_j \gamma$$

What contribution dominates $l_i \rightarrow 3 l_j$?

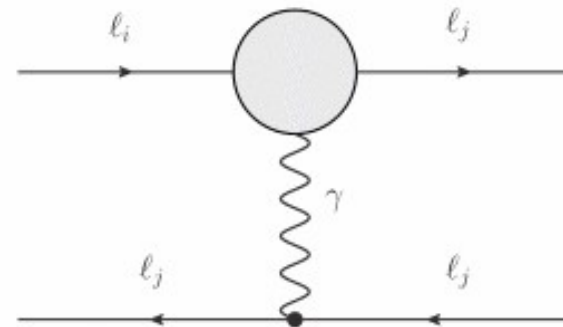
In many models of interest: **Photonic dipole contributions**

Most popular example: MSSM

[Hisano et al 1996; Arganda, Herrero 2006]



\simeq



Dipole dominance

$$\frac{BR(l_i \rightarrow 3 l_j)}{BR(l_i \rightarrow l_j \gamma)} = \frac{\alpha}{3\pi} \left(\log \frac{m_{l_i}^2}{m_{l_j}^2} - \frac{11}{4} \right) \Rightarrow BR(l_i \rightarrow l_j \gamma) \gg BR(l_i \rightarrow 3 l_j)$$

The LFV program

In order to unravel the **physics behind LFV** (and perhaps neutrino masses!) we must:

- **Search for LFV in as many observables as possible:** they might have information about different sectors of the theory
- **Study the relations among different observables** (ratios, correlations, hierarchies...)
- **Understand the origin of such relations:** what is the underlying physics?

Outline of the talk

- Introduction: Lepton Flavor Violation **FINISHED!**
- Selected topics
 - 3-loop neutrino mass models and CLFV
 - Ultralight scalars and CLFV
- Final remarks



Chuck Norris fact of the day

*Chuck Norris counted to
infinity. Twice.*



3-loop ν mass models and CLFV

**With Ricardo Cepedello, Martin Hirsch and
Paulina Rocha-Morán**

[JHEP 08 \(2020\) 067 \[arXiv:2005.00015\]](#)

3-loop neutrino mass models

3-loop Majorana neutrino mass models
can actually be very simple

Minimal models

KNT model: [Krauss, Nasri & Trodden, 2002](#)

AKS model: [Aoki, Kanemura & Seto, 2008](#)

Cocktail model: [Gustafsson, No & Rivera, 2012](#)

[Full classification in [Cepedello et al, 2018](#)]

3-loop neutrino mass models

3-loop Majorana neutrino mass models
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KNT model: Krauss, Nasri & Trodden, 2002

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Focus on **AKS**...

← ... but similar
(qualitative)
conclusions hold
for the other two

[Full classification in Cepedello et al, 2018]

The AKS model

[Aoki, Kanemura, Seto, 2008]

	gen	$SU(2)_L$	$U(1)_Y$	\mathbb{Z}_2
Φ	1	2	1/2	+
φ	1	1	0	-
S	1	1	1	-
N	3	1	0	-

← 2nd Higgs doublet

← real

Conserved \mathbb{Z}_2 parity

Dark Matter!

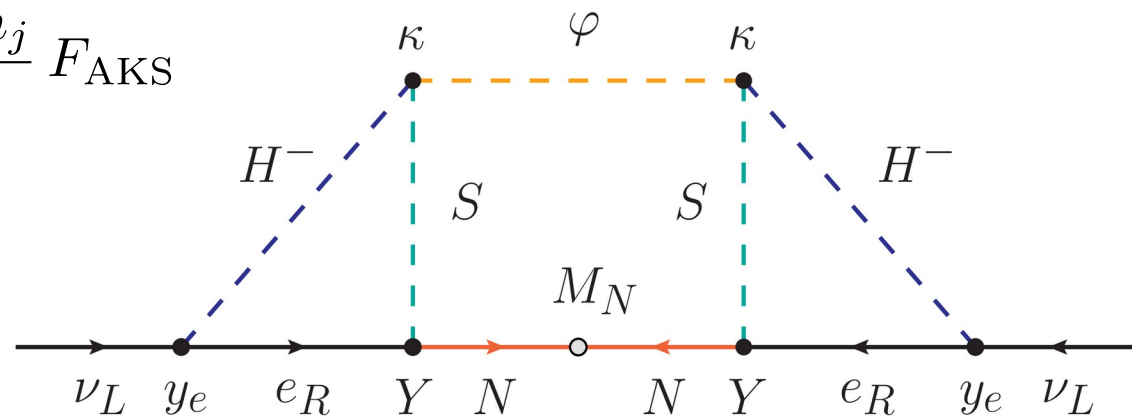
$$-\mathcal{L} \supset Y \overline{e_R^c} N S^* + \frac{1}{2} M_N \overline{N^c} N + \kappa \phi \Phi S^* \varphi + \text{h.c.}$$

$$(m_\nu)_{ij} = C_{\text{AKS}} \frac{\kappa^2}{(16\pi^2)^3} \frac{m_i Y_{i\alpha} Y_{j\beta} m_j}{(M_N)_{\alpha\beta}} F_{\text{AKS}}$$

Y : parametrized *à la* Casas-Ibarra

Master parametrization

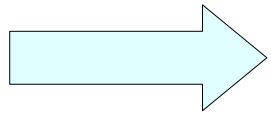
[Cordero-Carrión, Hirsch, AV, 2018, 2019]



(+ diagram with crossed S lines)

Neutrino mass in the AKS model

$$(m_\nu)_{ij} = C_{\text{AKS}} \frac{\kappa^2}{(16\pi^2)^3} \frac{m_i Y_{i\alpha} Y_{j\beta} m_j}{(M_N)_{\alpha\beta}} F_{\text{AKS}}$$

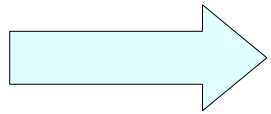


[Casas, Ibarra,
2001]

$$Y = \frac{i (16\pi^2)^{3/2}}{\kappa \tan \beta} \mathcal{R} \sqrt{M_N / F_{\text{AKS}}} \sqrt{\widehat{\mathcal{M}}_\nu} U^\dagger \widehat{\mathcal{M}}_e^{-1}$$

Neutrino mass in the AKS model

$$(m_\nu)_{ij} = C_{\text{AKS}} \frac{\kappa^2}{(16\pi^2)^3} \frac{m_i Y_{i\alpha} Y_{j\beta} m_j}{(M_N)_{\alpha\beta}} F_{\text{AKS}}$$



[Casas, Ibarra, 2001]

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Fit to oscillation data

Simple estimate



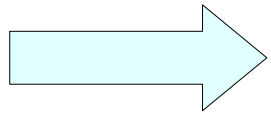
$$R = \mathbb{I}$$

$$m_\nu \sim 0.1 \text{ eV}$$

$$Y \sim \begin{matrix} & \begin{matrix} \uparrow & \uparrow & \uparrow \\ 1/m_e & 1/m_\mu & 1/m_\tau \end{matrix} \\ \begin{pmatrix} 100 & 1 & 0.1 \\ 100 & 1 & 0.1 \\ 100 & 1 & 0.1 \end{pmatrix} \end{matrix}$$

Neutrino mass in the AKS model

$$(m_\nu)_{ij} = C_{\text{AKS}} \frac{\kappa^2}{(16\pi^2)^3} \frac{m_i Y_{i\alpha} Y_{j\beta} m_j}{(M_N)_{\alpha\beta}} F_{\text{AKS}}$$



[Casas, Ibarra, 2001]

$$Y = \frac{i (16\pi^2)^{3/2}}{\kappa \tan \beta} \mathcal{R} \sqrt{M_N / F_{\text{AKS}}} \sqrt{\widehat{\mathcal{M}}_\nu} U^\dagger \widehat{\mathcal{M}}_e^{-1}$$

Fit to oscillation data

Simple estimate



$$R = \mathbb{I}$$

$$m_\nu \sim 0.1 \text{ eV}$$

$$Y \sim \begin{pmatrix} 100 & 1 & 0.1 \\ 100 & 1 & 0.1 \\ 100 & 1 & 0.1 \end{pmatrix}$$

\uparrow \uparrow \uparrow
 $1/m_e$ $1/m_\mu$ $1/m_\tau$

LFV constraints

Perturbativity



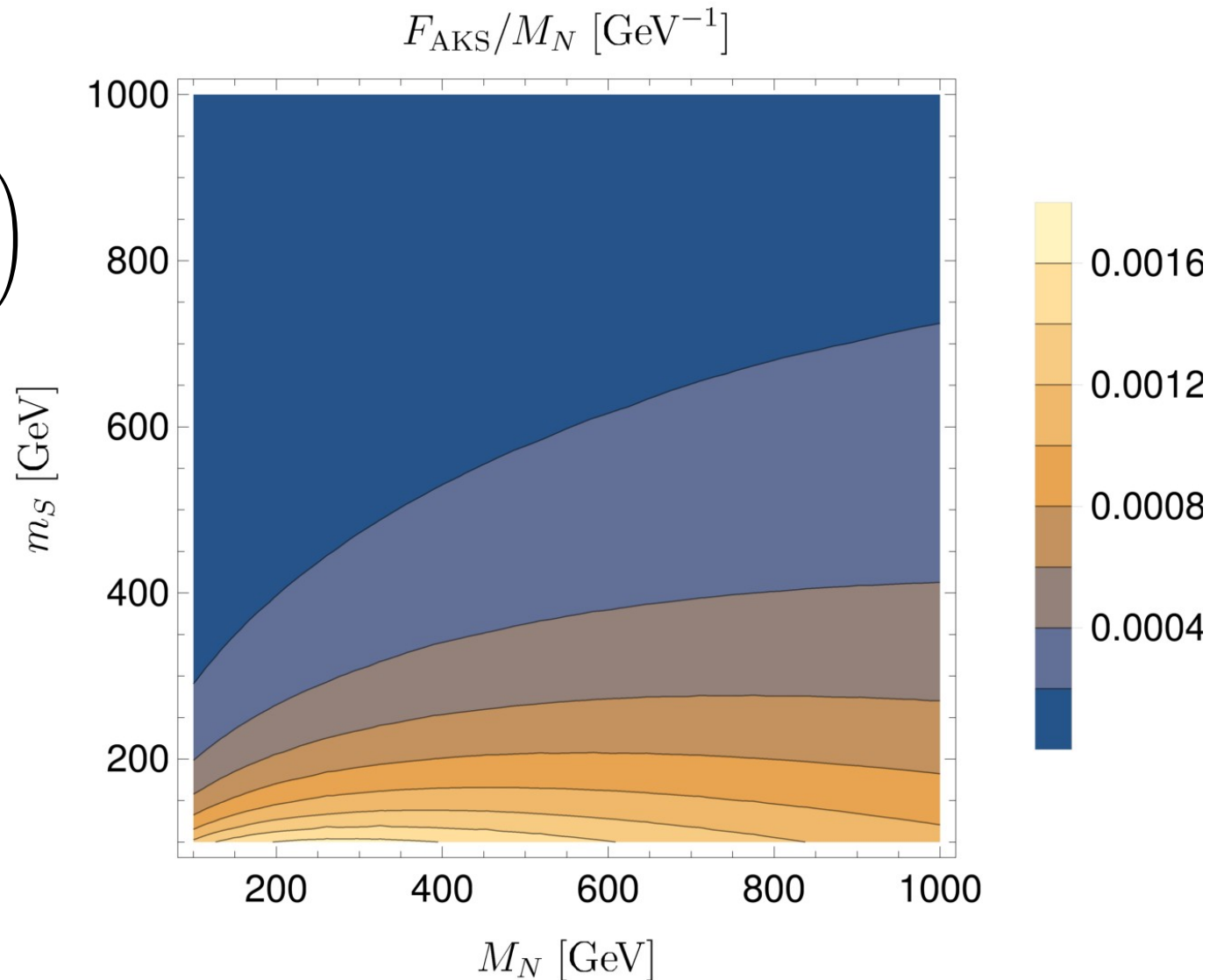
AKS loop function

Loop function

$$F_{\text{AKS}} \left(\frac{M_S^2}{M_N^2}, \frac{M_\varphi^2}{M_N^2}, \frac{M_{H^\pm}^2}{M_N^2} \right)$$

Common scalar mass

$$M_S = M_\varphi = M_{H^\pm} \equiv m_S$$



AKS loop function

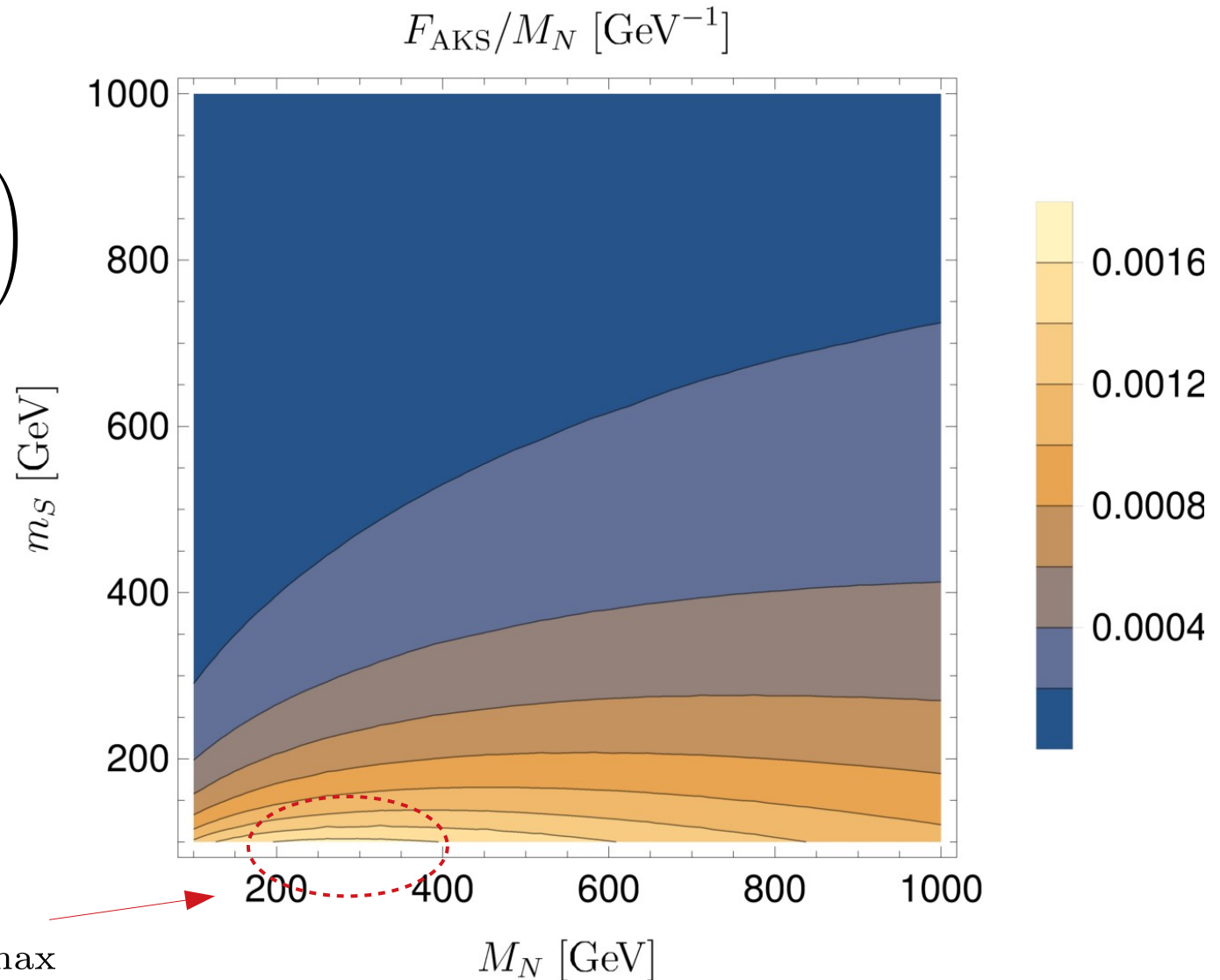
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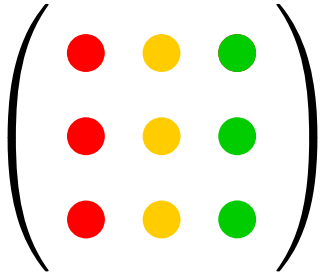
Common scalar mass

$$M_S = M_\varphi = M_{H^\pm} \equiv m_S$$

$(F_{\text{AKS}}/M_N)_{\text{max}}$



Fine-tuning Υ in the AKS model

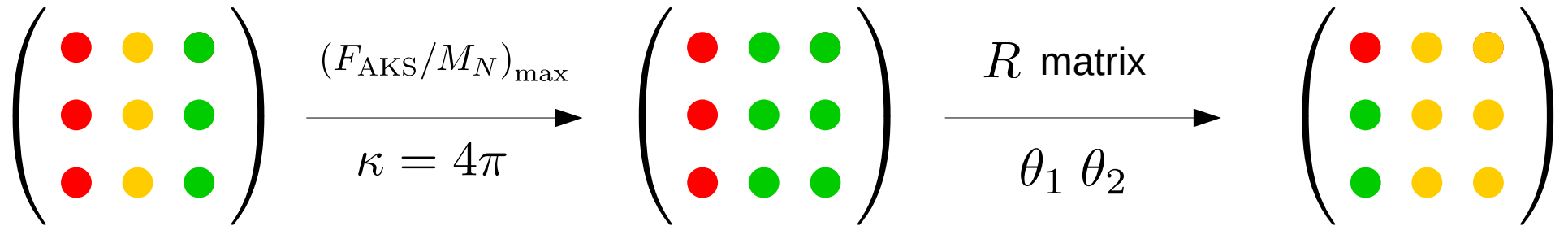


Fine-tuning Y in the AKS model

$$\begin{pmatrix} \bullet & \bullet & \bullet \\ \bullet & \bullet & \bullet \\ \bullet & \bullet & \bullet \end{pmatrix} \xrightarrow[\kappa = 4\pi]{(F_{\text{AKS}}/M_N)_{\text{max}}} \begin{pmatrix} \bullet & \bullet & \bullet \\ \bullet & \bullet & \bullet \\ \bullet & \bullet & \bullet \end{pmatrix}$$

The diagram illustrates a transformation of a 3x3 matrix. The left matrix has a red dot in the first column, a yellow dot in the second column, and a green dot in the third column for each row. The right matrix has a red dot in the first column and green dots in the second and third columns for each row. The transformation is labeled with $(F_{\text{AKS}}/M_N)_{\text{max}}$ above the arrow and $\kappa = 4\pi$ below it.

Fine-tuning Y in the AKS model



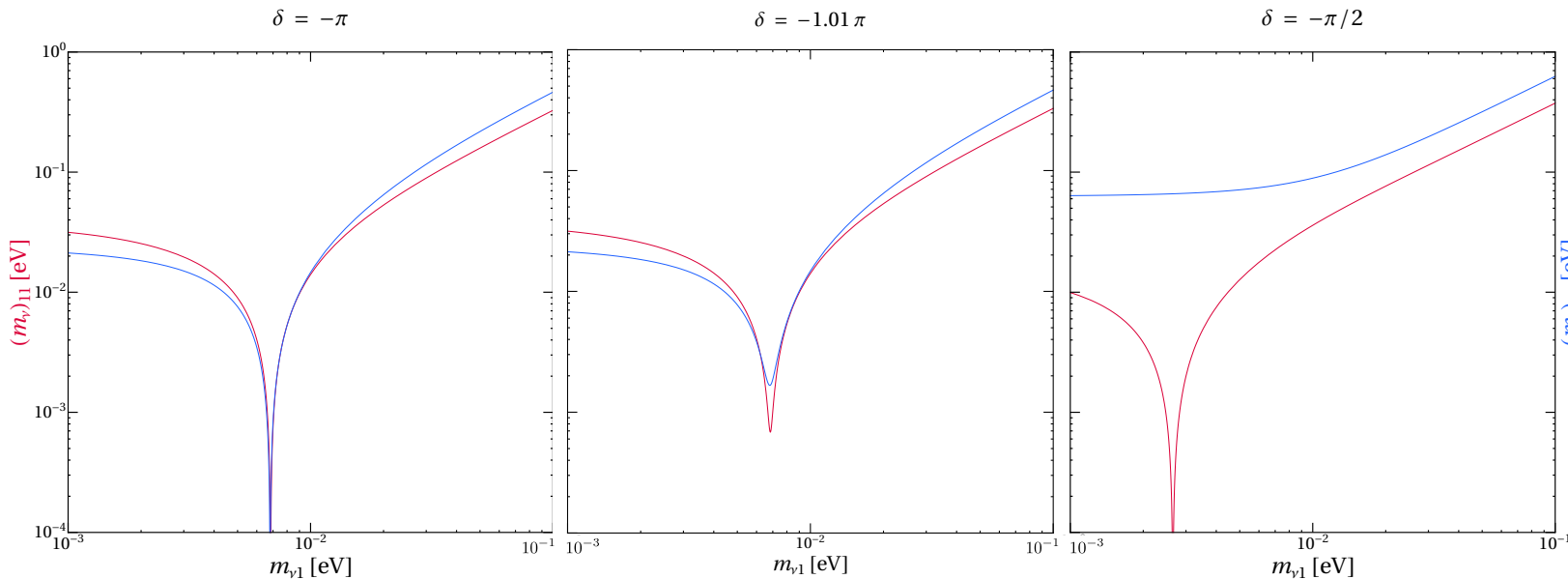
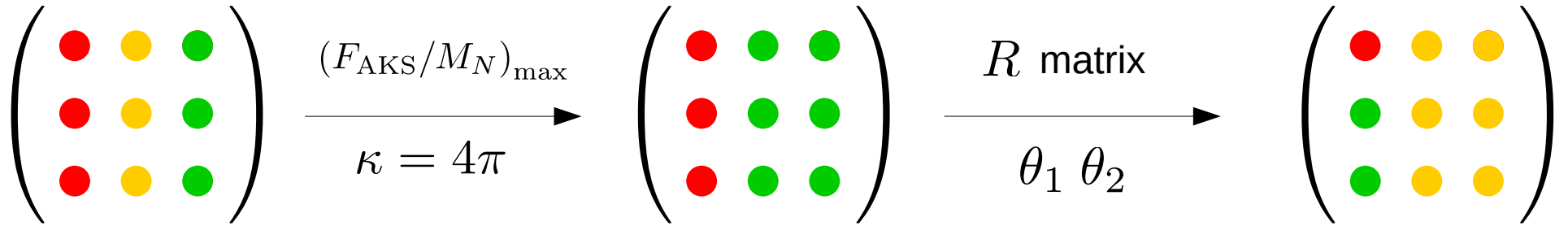
$$Y = \frac{i (16\pi^2)^{3/2}}{\kappa \tan \beta} \mathcal{R} \sqrt{M_N/F_{\text{AKS}}} \sqrt{\widehat{\mathcal{M}}_\nu} U^\dagger \widehat{\mathcal{M}}_e^{-1}$$

$$R = R(\theta_1, \theta_2, \theta_3)$$

θ_i : complex angles

$$\theta_1 \theta_2 \Rightarrow Y_{21,31} \rightarrow 0$$

Fine-tuning γ in the AKS model

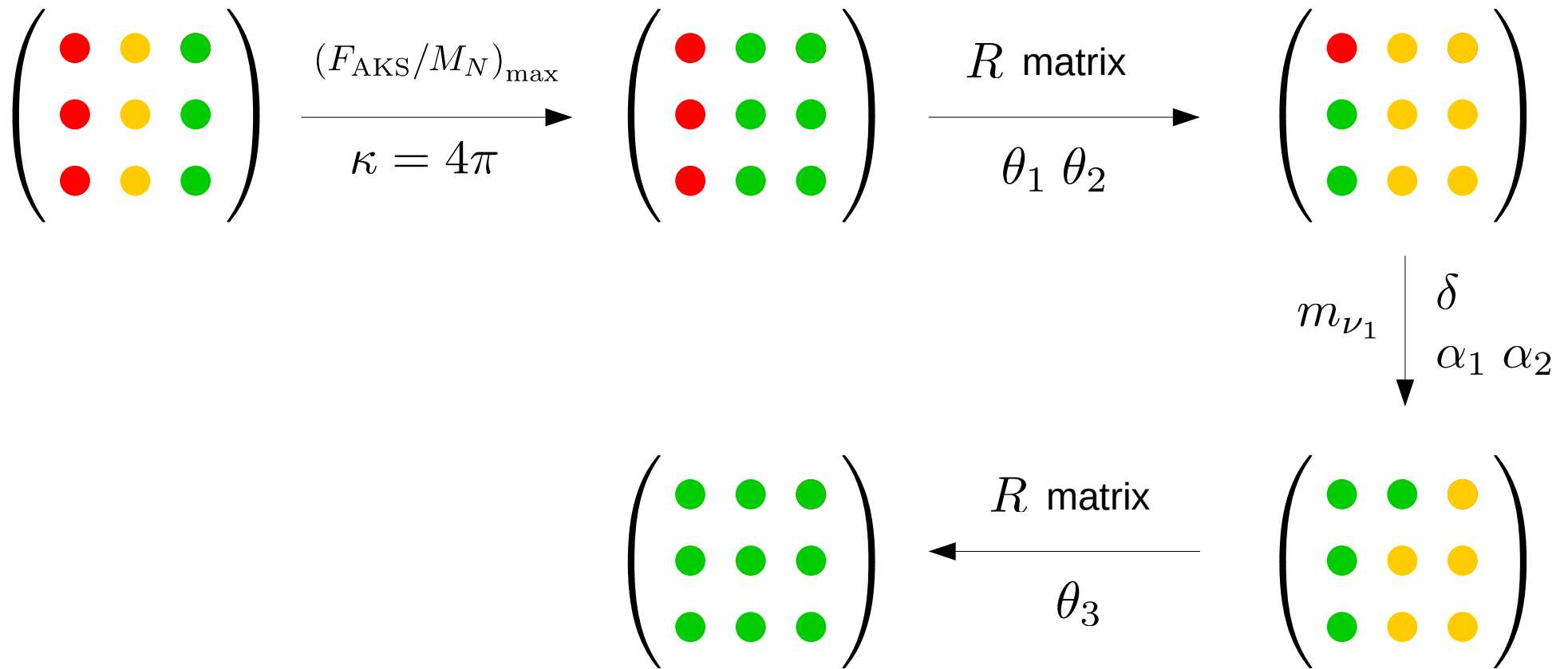


2σ tension with fits
 [De Salas et al, 2020]

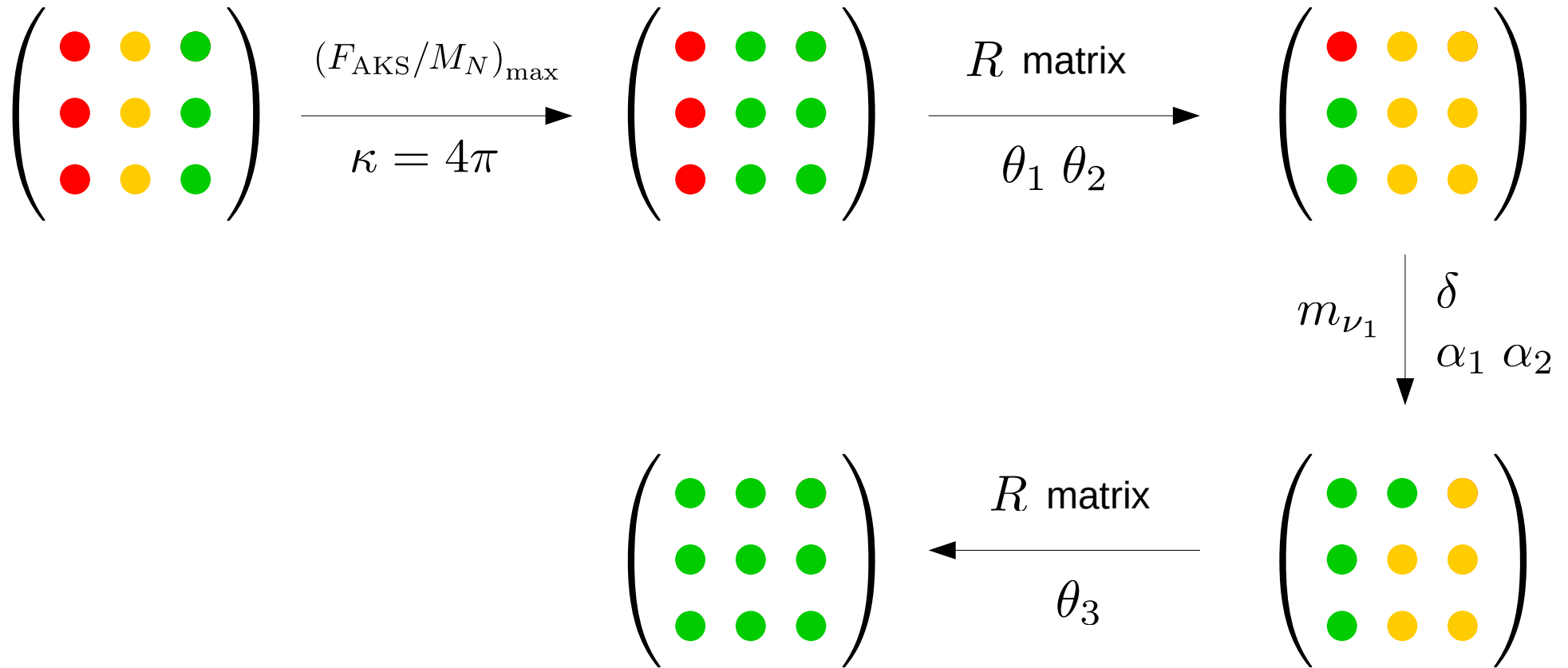
Fine-tuned

Partially working for $\delta = -\frac{\pi}{2}$

Fine-tuning γ in the AKS model



Fine-tuning Υ in the AKS model



Perturbativity + flavor = (fine-tuning)⁴

+
largish LFV
effects

Ultralight scalars and CLFV

With Pablo Escribano

JHEP 03 (2021) 240 [[arXiv:2008.01099](https://arxiv.org/abs/2008.01099)]

Ultralight scalars

ϕ : ultralight scalar that couples to charged leptons

$$m_\phi \ll m_e$$

($m_\phi = 0$ included)

Motivation: axion, axion-like particles, majoron, familon, Goldstone bosons...

Ultralight scalars

ϕ : ultralight scalar that couples to charged leptons

$$m_\phi \ll m_e$$

($m_\phi = 0$ included)

Effective Lagrangian

$$\mathcal{L} = \mathcal{L}_{\ell\ell\phi} + \mathcal{L}_{\ell\ell\gamma} + \mathcal{L}_{4\ell}$$

Ultralight scalars

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$$m_\phi \ll m_e \\ (m_\phi = 0 \text{ included})$$

Effective Lagrangian

$$\mathcal{L} = \mathcal{L}_{\ell\ell\phi} + \mathcal{L}_{\ell\ell\gamma} + \mathcal{L}_{4\ell}$$

$$\mathcal{L}_{\ell\ell\phi} = \phi \bar{\ell}_\beta \left(S_L^{\beta\alpha} P_L + S_R^{\beta\alpha} P_R \right) \ell_\alpha + \text{h.c.}$$

Ultralight scalars

ϕ : ultralight scalar that couples to charged leptons

$$m_\phi \ll m_e$$

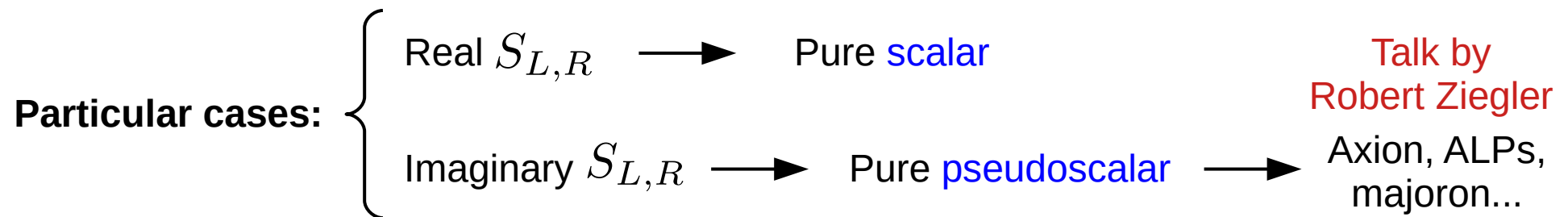
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Effective Lagrangian

$$\mathcal{L} = \mathcal{L}_{\ell\ell\phi} + \mathcal{L}_{\ell\ell\gamma} + \mathcal{L}_{4\ell}$$

$$\mathcal{L}_{\ell\ell\phi} = \phi \bar{\ell}_\beta \left(S_L^{\beta\alpha} P_L + S_R^{\beta\alpha} P_R \right) \ell_\alpha + \text{h.c.}$$

General dimensionless complex coefficients



Ultralight scalars

ϕ : ultralight scalar that couples to charged leptons

$$m_\phi \ll m_e \\ (m_\phi = 0 \text{ included})$$

Effective Lagrangian

$$\mathcal{L} = \mathcal{L}_{\ell\ell\phi} + \mathcal{L}_{\ell\ell\gamma} + \mathcal{L}_{4\ell}$$

$$\mathcal{L}_{\ell\ell\phi} = \phi \bar{\ell}_\beta \left(S_L^{\beta\alpha} P_L + S_R^{\beta\alpha} P_R \right) \ell_\alpha + \text{h.c.}$$

$$\mathcal{L}_{\ell\ell\gamma} = \frac{e m_\alpha}{2} \bar{\ell}_\beta \sigma^{\mu\nu} \left[(K_2^L)^{\beta\alpha} P_L + (K_2^R)^{\beta\alpha} P_R \right] \ell_\alpha F_{\mu\nu} + \text{h.c.}$$

$$\mathcal{L}_{4\ell} = \sum_{\substack{I=S,V,T \\ X,Y=L,R}} (A_{XY}^I)^{\beta\alpha\delta\gamma} \bar{\ell}_\beta \Gamma_I P_X \ell_\alpha \bar{\ell}_\delta \Gamma_I P_Y \ell_\gamma + \text{h.c.}$$

Leptonic observables

ϕ : ultralight scalar that couples to charged leptons

$$m_\phi \ll m_e$$

($m_\phi = 0$ included)

- Produced in the final state
- $l_\alpha \rightarrow l_\beta \phi$
 - $l_\alpha \rightarrow l_\beta \phi \gamma$
- Virtual particle
- $l_\alpha \rightarrow l_\beta \gamma$
 - $l_\alpha^- \rightarrow l_\beta^- l_\beta^- l_\beta^+$
 - $l_\alpha^- \rightarrow l_\beta^- l_\gamma^- l_\gamma^+$
 - $l_\alpha^- \rightarrow l_\beta^+ l_\gamma^- l_\gamma^+$
 - Lepton dipole moments (AMMs & EDMs)

[Escribano, AV, 2020]

Leptonic observables

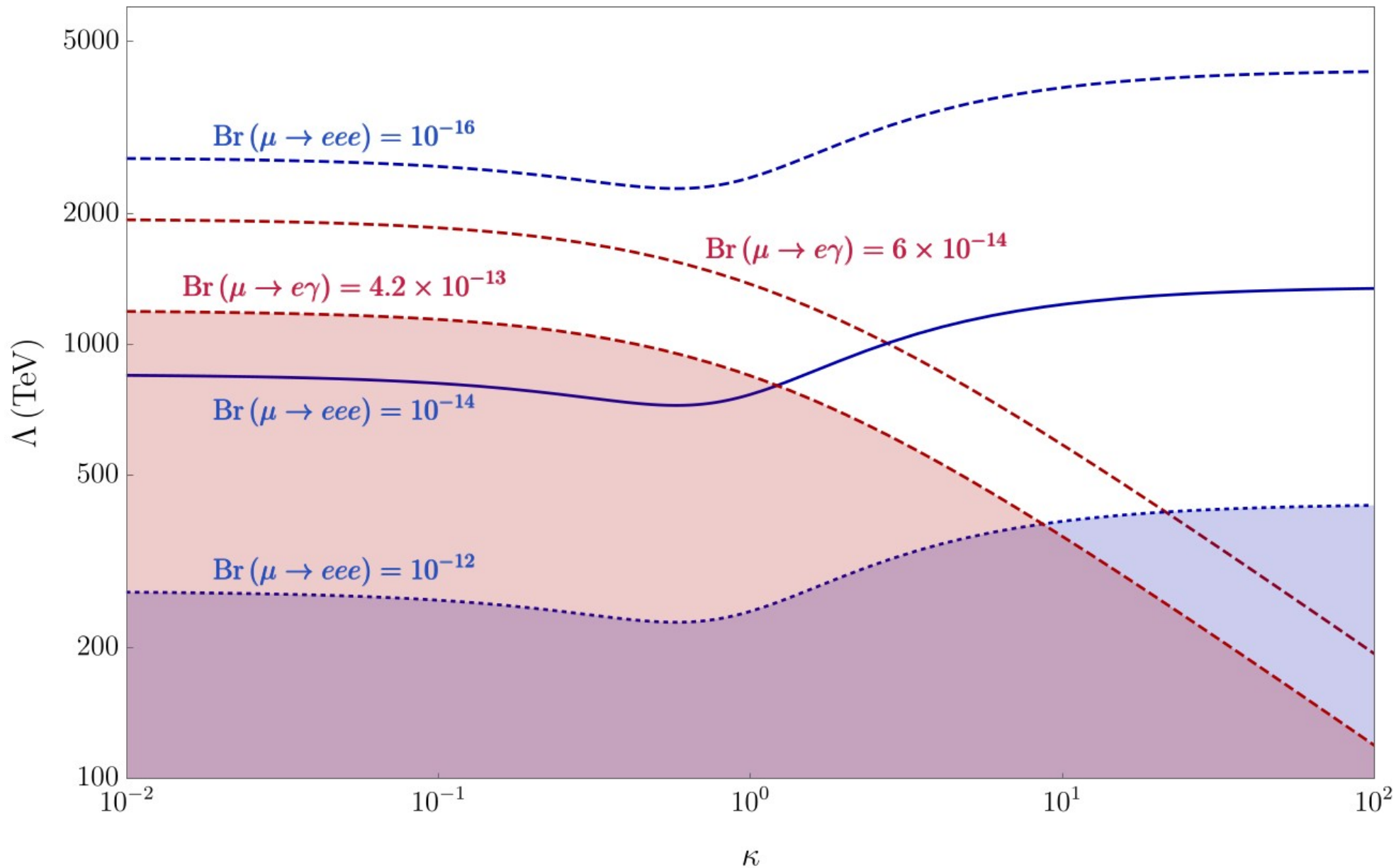
NEW!

$$\begin{aligned}
 & \Gamma_\phi \left(l_\alpha^- \rightarrow l_\beta^- l_\beta^- l_\beta^+ \right) = \\
 & \frac{m_\alpha}{512\pi^3} \left\{ \left(|S_L^{\beta\alpha}|^2 + |S_R^{\beta\alpha}|^2 \right) \left\{ |S^{\beta\beta}|^2 \left(4 \log \frac{m_\alpha}{m_\beta} - \frac{49}{6} \right) - \frac{2}{6} \left[(S^{\beta\beta*})^2 + (S^{\beta\beta})^2 \right] \right\} \right. \\
 & - \frac{m_\alpha^2}{6} \left\{ S_L^{\beta\alpha} S^{\beta\beta} A_{LL}^{S*} + 2S_L^{\beta\alpha} S^{\beta\beta*} A_{LR}^{S*} + 2S_R^{\beta\alpha} S^{\beta\beta} A_{RL}^{S*} + S_R^{\beta\alpha} S^{\beta\beta*} A_{RR}^{S*} \right. \\
 & - 12 \left(S_L^{\beta\alpha} S^{\beta\beta} A_{LL}^{T*} + S_R^{\beta\alpha} S^{\beta\beta*} A_{RR}^{T*} \right) - 4 \left(S_R^{\beta\alpha} S^{\beta\beta} A_{RL}^{V*} + S_L^{\beta\alpha} S^{\beta\beta*} A_{LR}^{V*} \right) \\
 & \left. \left. + 6e^2 \left[S_R^{\beta\alpha} S^{\beta\beta} (K_2^L)^{\beta\alpha*} + S_L^{\beta\alpha} S^{\beta\beta*} (K_2^R)^{\beta\alpha*} \right] + \right\} \right\}
 \end{aligned}$$

Infrared divergence



$$S^{\beta\beta} = S_L^{\beta\beta} + S_R^{\beta\beta*}$$



$$e (K_2^L)^{\beta\alpha} \equiv \frac{1}{(\kappa + 1) \Lambda^2} \quad S_L^{\beta\alpha} \equiv m_\alpha \frac{\kappa}{(\kappa + 1) \Lambda}$$

Final remarks

Final remarks

LFV is going to live a **golden age**

Many LFV observables. **Correlations** are not only possible, but in fact expected!

We must be **ready**: understand the LFV anatomy, patterns, correlations, hierarchies...

Thank you!

谢谢！



Backup slides

The KNT model

[Krauss, Nasri, Trodden, 2002]

	gen	$SU(2)_L$	$U(1)_Y$	\mathbb{Z}_2
h	1	1	1	+
S	1	1	1	-
N	3	1	0	-

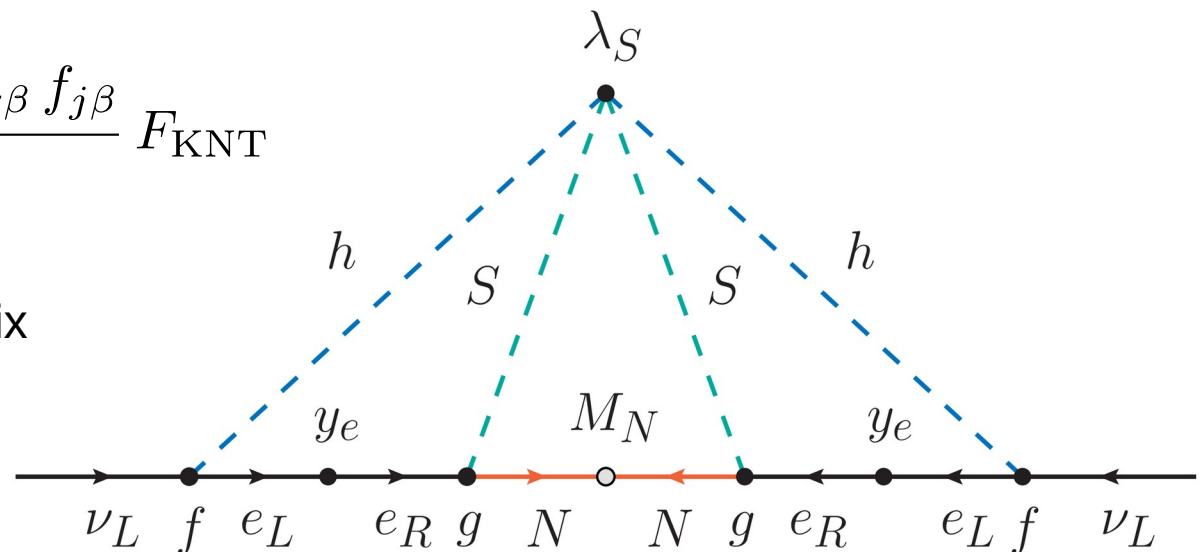
← Singly
charged
scalars

Conserved \mathbb{Z}_2 parity
Dark Matter!

$$-\mathcal{L} \supset f \bar{\ell}^c \ell h + g \bar{N}^c e_R S + \frac{1}{2} M_N \bar{N}^c N + \lambda_S (h S^*)^2 + \text{h.c.}$$

$$(m_\nu)_{ij} = \frac{\lambda_S}{(16\pi^2)^3} \frac{f_{i\alpha} m_\alpha g_\alpha^* g_\beta^* m_\beta f_{j\beta}}{(M_N)_{\alpha\beta}} F_{\text{KNT}}$$

f : antisymmetric Yukawa matrix



The Cocktail model

[Gustafsson, No, Rivera, 2012]

	gen	$SU(2)_L$	$U(1)_Y$	\mathbb{Z}_2
S	1	1	1	-
ρ	1	1	2	+
η	1	2	1/2	-

Conserved \mathbb{Z}_2 parity

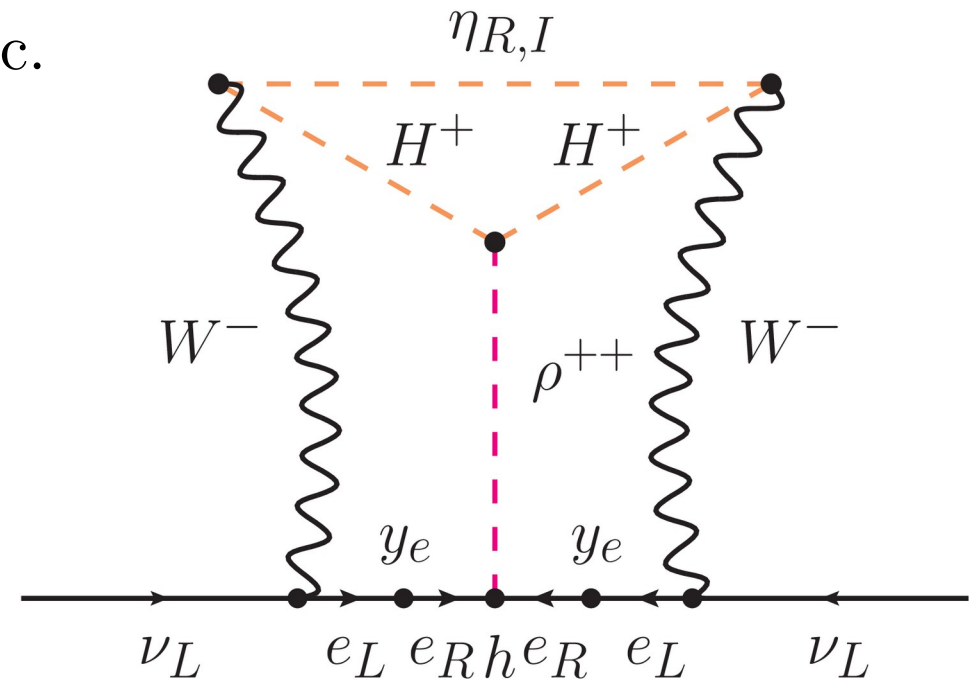
Dark Matter!

← Inert doublet

$$-\mathcal{L} \supset h \overline{e_R^c} e_R \rho + \frac{1}{2} \lambda_5 (\phi \eta^*)^2 + \text{h.c.}$$

$$(m_\nu)_{ij} = \frac{\lambda_5}{(16\pi^2)^3} \frac{m_i h_{ij} m_j}{v} F_{\text{Cocktail}}$$

h : symmetric Yukawa matrix
(type-II seesaw-like)



FlavorKit

[Porod, Staub, AV, 2014]

A computer tool that provides automatized analytical and numerical computation of flavor observables. It is based on **SARAH**, **SPheno** and **FeynArts/FormCalc**.

Lepton flavor	Quark flavor
$l_\alpha \rightarrow l_\beta \gamma$	$B_{s,d}^0 \rightarrow l^+ l^-$
$l_\alpha \rightarrow 3 l_\beta$	$\bar{B} \rightarrow X_s \gamma$
$\mu - e$ conversion in nuclei	$\bar{B} \rightarrow X_s l^+ l^-$
$\tau \rightarrow P l$	$\bar{B} \rightarrow X_{d,s} \nu \bar{\nu}$
$h \rightarrow l_\alpha l_\beta$	$B \rightarrow K l^+ l^-$
$Z \rightarrow l_\alpha l_\beta$	$K \rightarrow \pi \nu \bar{\nu}$
	$\Delta M_{B_{s,d}}$
	ΔM_K and ε_K
	$P \rightarrow l \nu$

Not limited to a single model: use it for the **model of your choice**

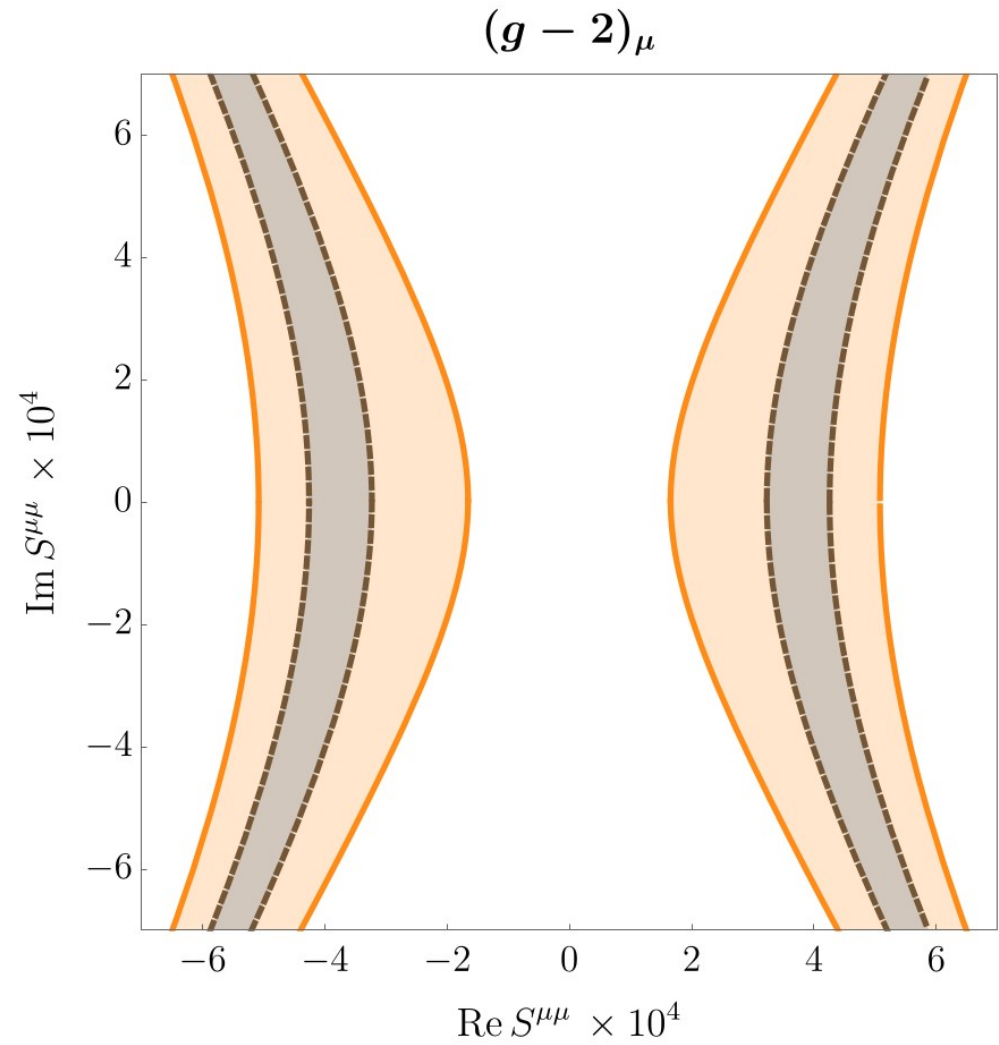
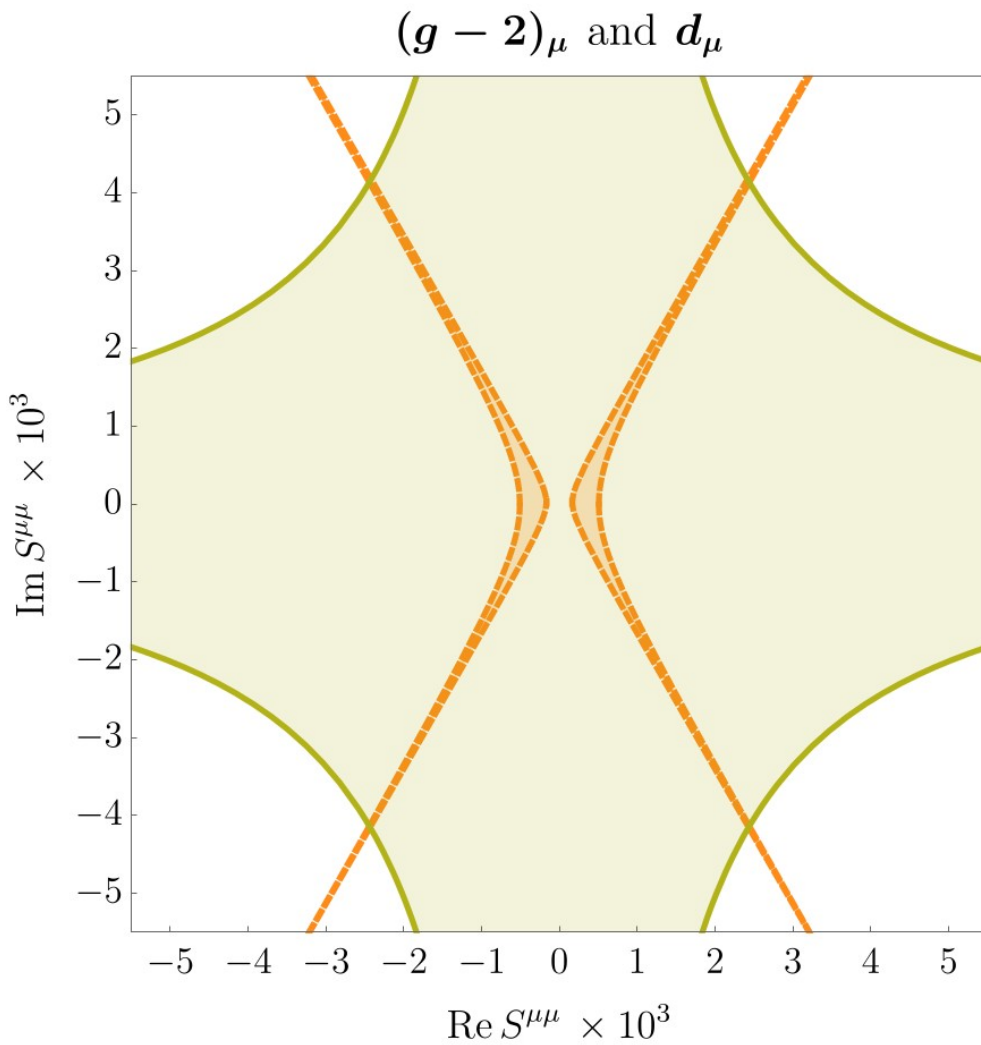
Easily **extendable**

Many observables ready to be computed in your favourite model!

Manual: [arXiv:1405.1434](https://arxiv.org/abs/1405.1434)

Website: <http://sarah.hepforge.org/FlavorKit.html>

Muon g-2



Flavor diagonal scenario

$$\Delta a_\alpha = \frac{1}{16\pi^2} \left[3 (\text{Re } S^{\alpha\alpha})^2 - (\text{Im } S^{\alpha\alpha})^2 \right]$$

[Escribano, AV, 2020]