

# Light-ray OPE in QCD

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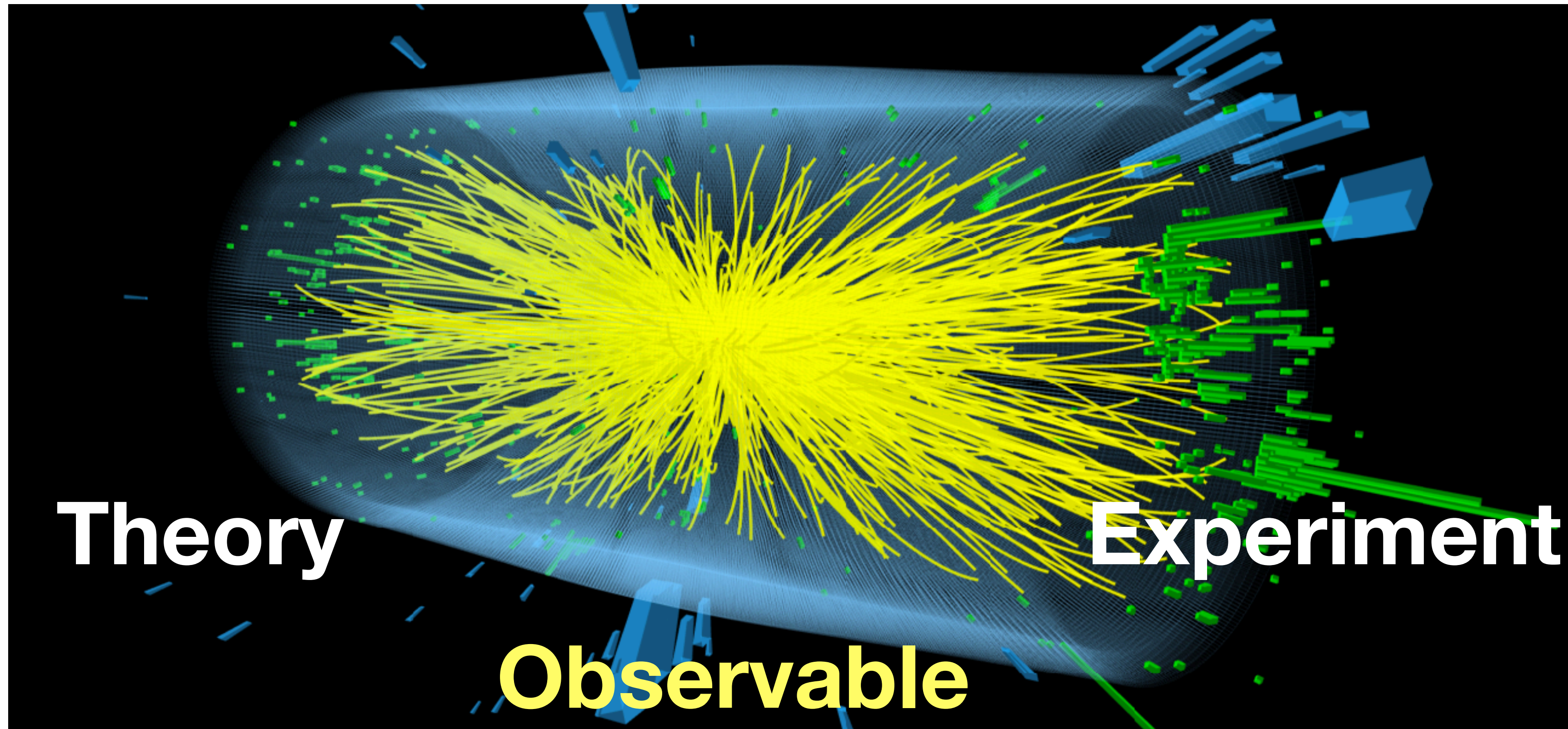
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2021-10-14

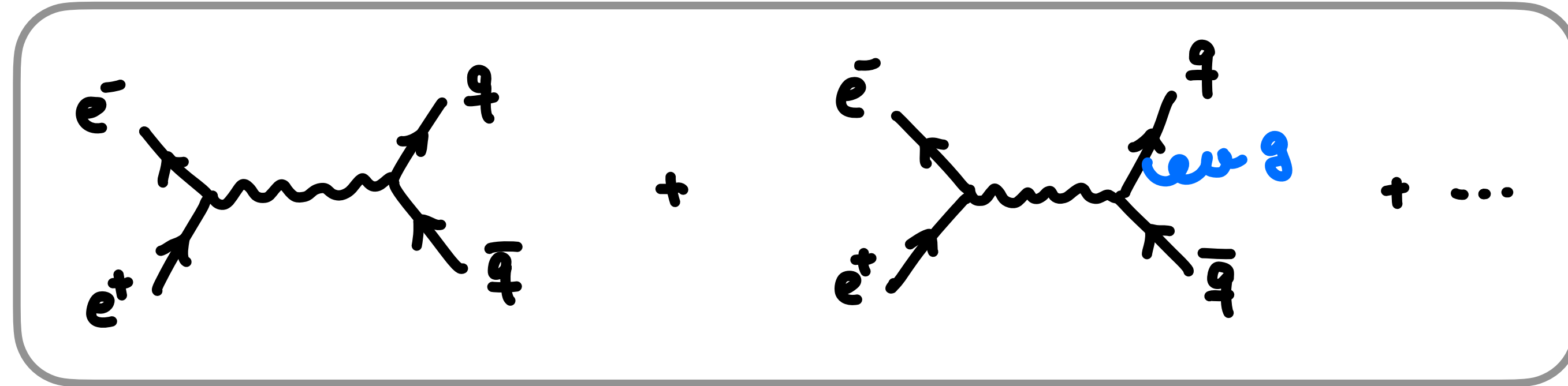
# Collider Physics



- Experimentally viable
- Easy to calculate
- Have clean theoretical understanding

# Event Shapes

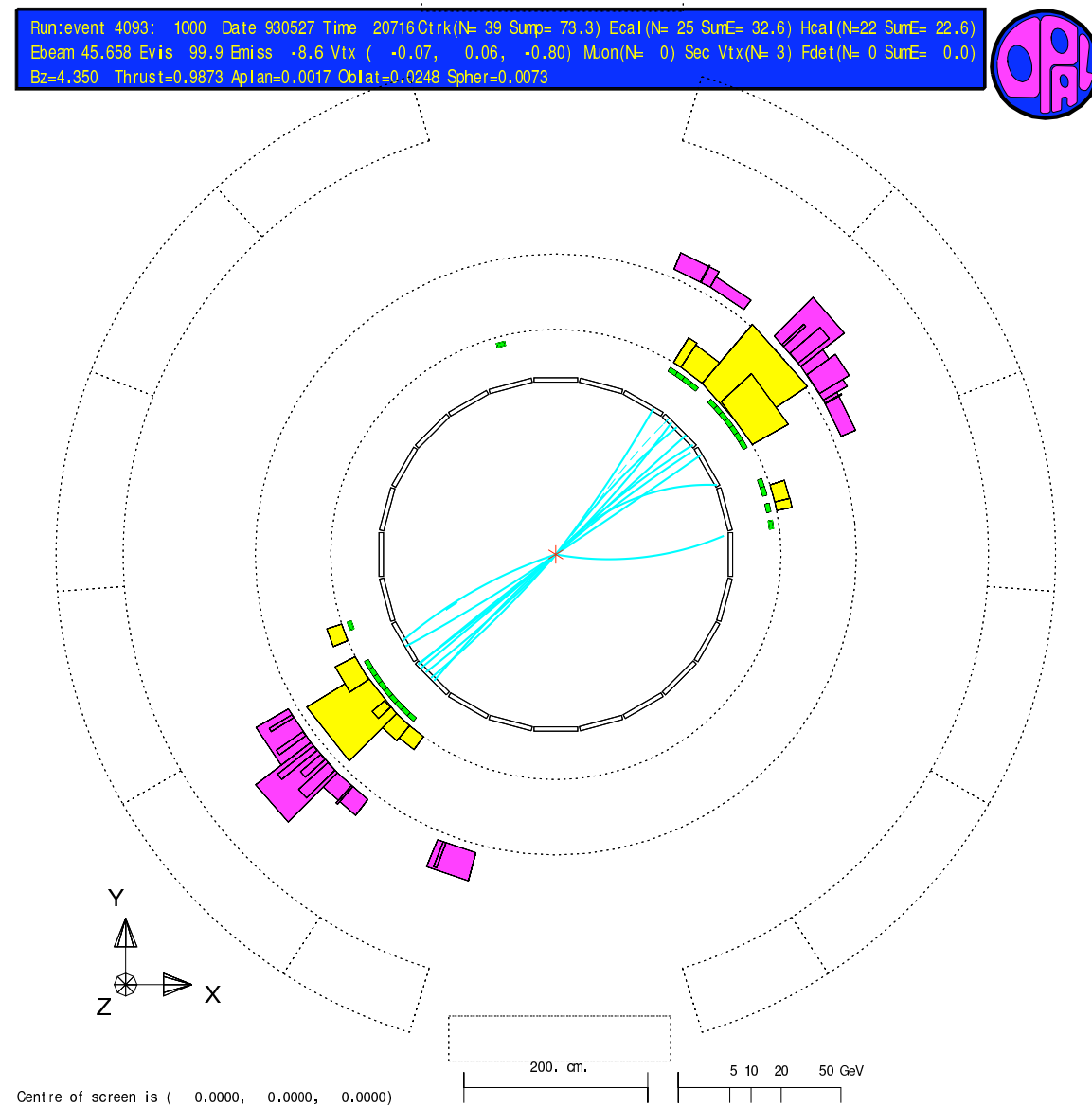
$e^+e^-$  annihilation



pencil-like distribution

vs

(more) spherical distribution



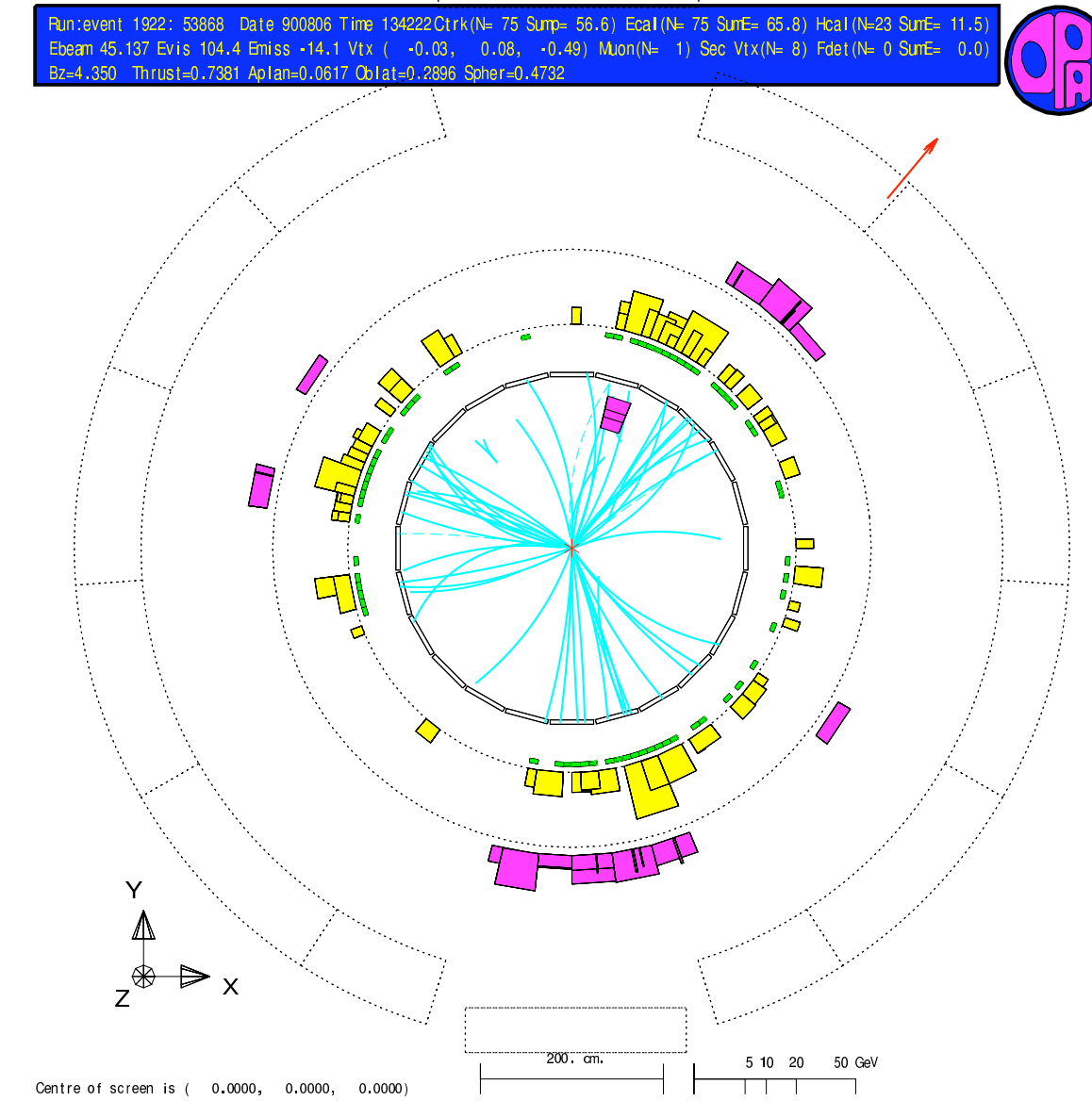
**Example: Thrust**

[Farhi, 1977]

$$T = \max_{\vec{n}} \frac{\sum_i |\vec{n} \cdot \vec{p}_i|}{Q}$$

pencil-like  $T \sim 1$

spherical  $T \sim 1/2$



# Energy-energy correlator

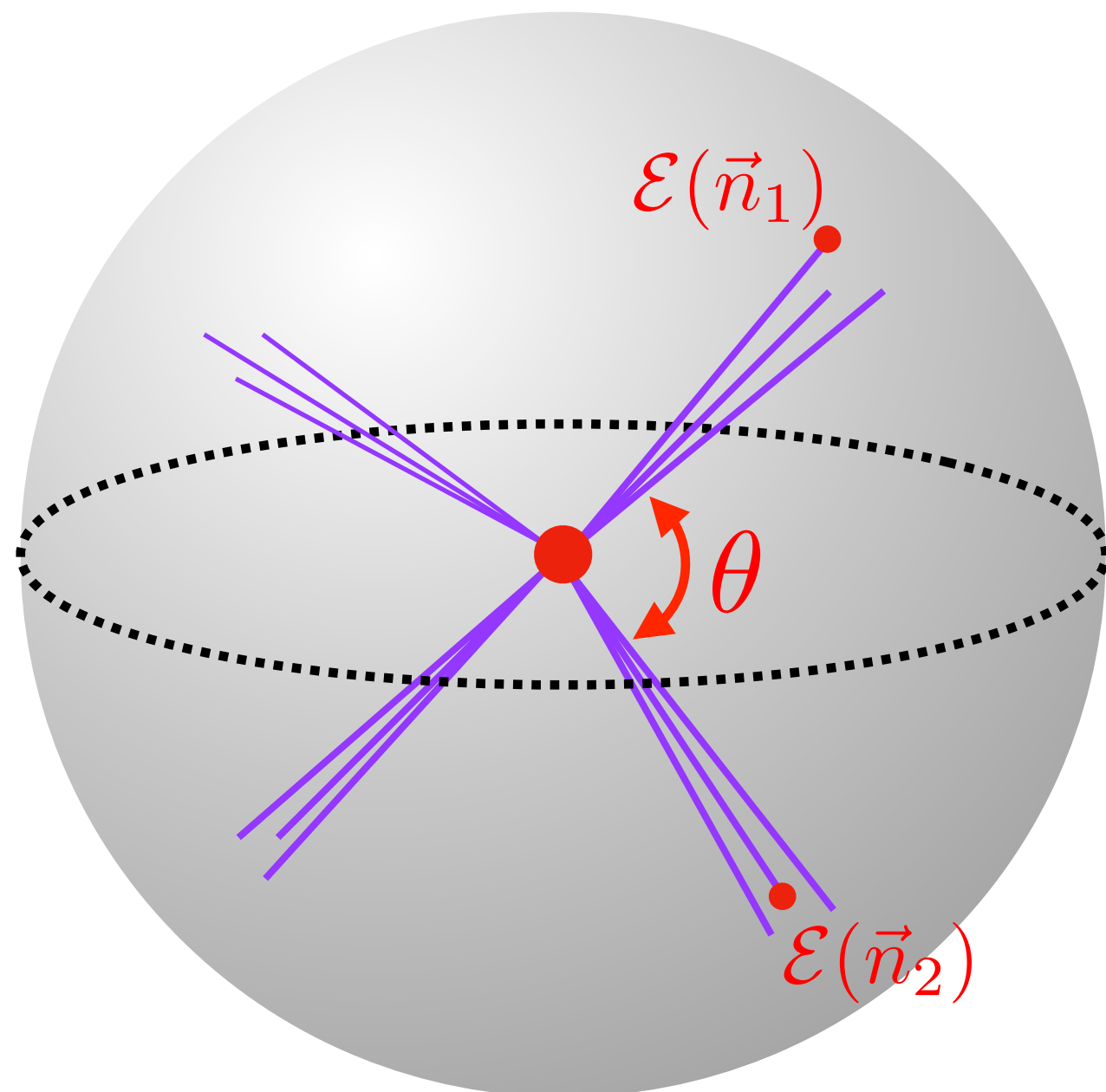
[Basham, Brown, Ellis and Love, 1978]

introduced energy-energy correlation

$$\frac{d\Sigma}{dz} = \sum_{i,j} \int d\sigma \frac{E_i E_j}{Q^2} \delta \left( z - \frac{1 - \cos \theta_{ij}}{2} \right)$$

which characterizes the correlation of two **energy detectors** at spatial infinity (**celestial sphere**).

**Energy Correlation  
on the celestial sphere**



**Familiar concept in statistical mechanics!**

## Probability Distribution

differential cross section

$d\sigma$

Boltzmann factor

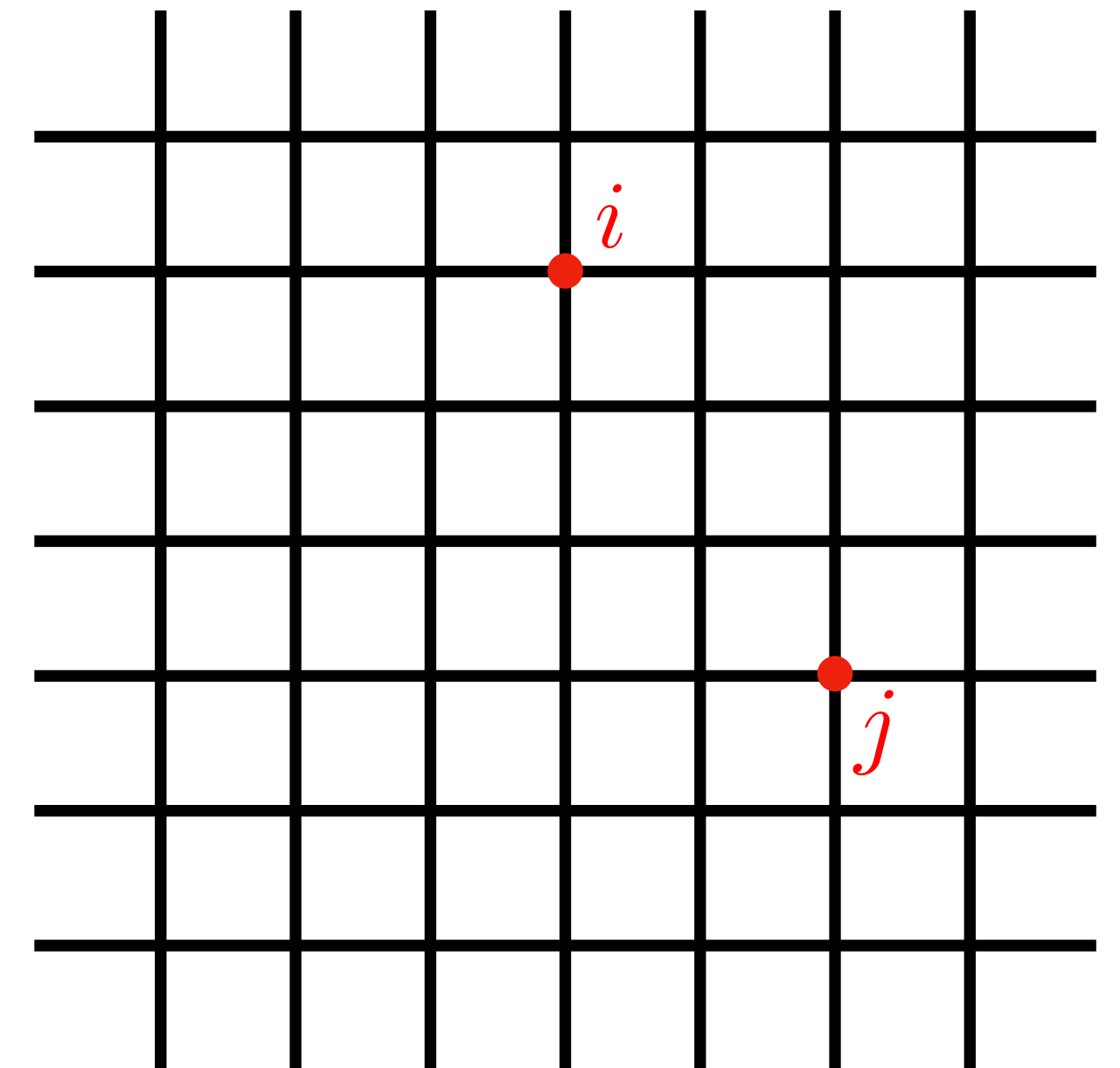
$e^{-\beta H}$

## Weighting Factor

eigenvalues of energy

eigenvalues of spin

**Spin Correlation  
on the plane (2D Ising)**

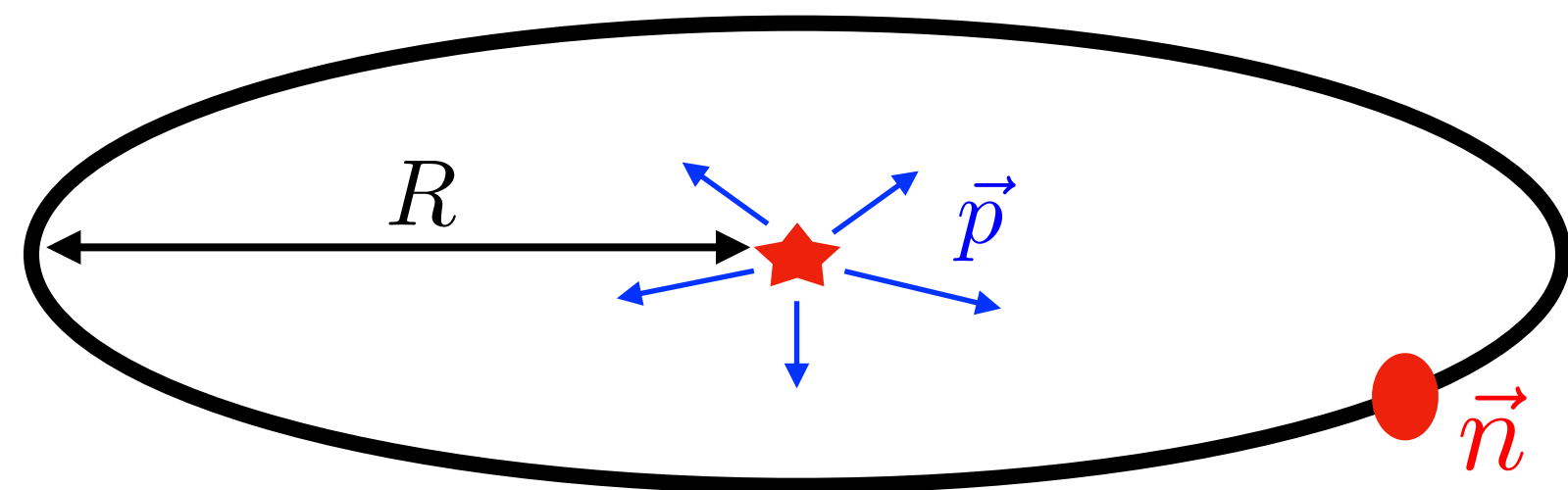


# Energy Flow Operator $\mathcal{E}(\vec{n})$

“Perturbative” vs “Non-perturbative”

A calorimeter only detects **particles** flowing along direction  $\vec{n}$ , and weight with its energy  $E$ , e.g.

$$\int \frac{d^3 p}{(2\pi)^3 2E_{\vec{p}}} E_{\vec{p}} a_{\vec{p}}^\dagger a_{\vec{p}} \delta^{(2)}(\vec{n} - \hat{p})$$



The radiation power passing the detector (located at  $R\vec{n}$ ) at time  $t$  is

$$n^i T_i^0(t, R\vec{n}) R^2 d\Omega$$

[Korchemsky, Sterman, 1999;  
Hofman, Maldacena, 2008;  
Bauer, Fleming, Lee, Sterman, 2008; ...]

- Integrate  $t$  to get the total received energy
- Detector is effectively located at infinity

$$\mathcal{E}(\vec{n}) = \lim_{R \rightarrow \infty} R^2 \int_0^\infty dt n_i T^{0i}(t, R\vec{n})$$

non-perturbative definition

# Energy Flow Operator

- For free theory, we can use mode expansion [see, e.g. Bauer, Fleming, Lee and Sterman, 2008]

$$\mathcal{E}(\vec{n}) = \lim_{r \rightarrow \infty} r^2 \int_0^\infty dt \vec{n}_i T^{0i}(t, r\vec{n}) \xrightarrow[T_{0i} = \partial_t \phi \partial_i \phi]{\text{Free scalar}} \int \frac{d^3 p}{(2\pi)^3 2E_{\vec{p}}} E_{\vec{p}} a_{\vec{p}}^\dagger a_{\vec{p}} \delta^{(2)}(\vec{n} - \hat{p})$$

so they are equivalent when acting on asymptotic Fock state.

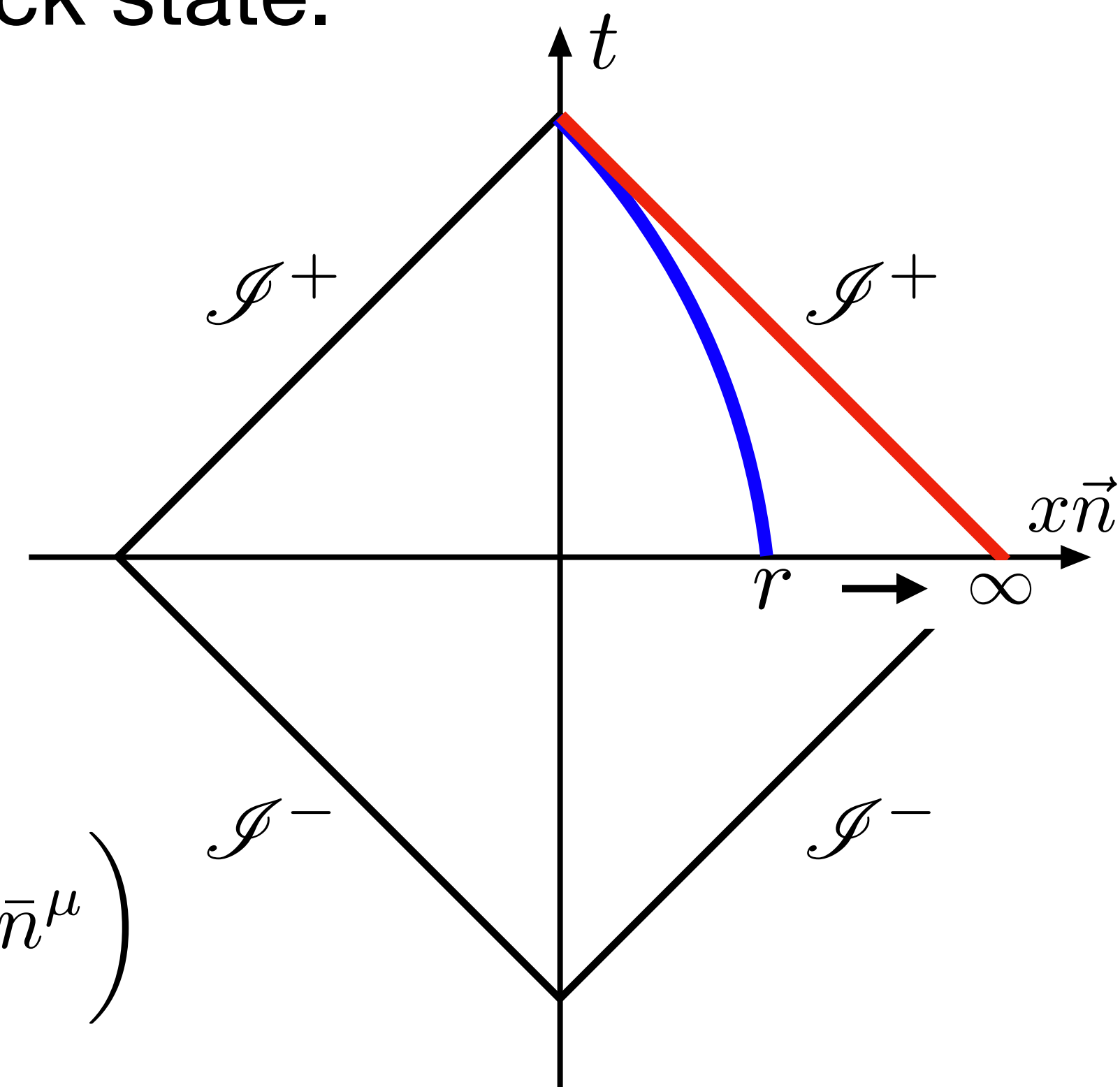
- The energy flow operator is a non-local operator defined on a **light-ray** located at **future null infinity**

$$\mathcal{E}(\vec{n}) = \lim_{r \rightarrow \infty} r^2 \int_0^\infty dt \vec{n}_i T^{0i}(t, r\vec{n})$$

- Equivalent form in lightcone coordinate

$$\mathcal{E}(\vec{n}) = \frac{1}{4} \lim_{x^+ \rightarrow \infty} \left( \frac{x^+}{2} \right)^2 \int_{-\infty}^\infty dx^- \bar{n}_{\mu_1} \bar{n}_{\mu_2} T^{\mu_1 \mu_2} \left( \frac{x^+}{2} n^\mu + \frac{x^-}{2} \bar{n}^\mu \right)$$

- This is an example of light-ray operators.



# Energy Correlators

Energy correlators are correlation function of multiple **energy flow operators** inside some **non-vacuum** states

$$\langle \Psi' | \mathcal{E}(\vec{n}_1) \mathcal{E}(\vec{n}_2) \dots | \Psi \rangle$$

Looks like a correlation function in a fictitious 2D field theory on  $S^2$

e.g. created by local operators

[momentum space]

$$\langle O'(-q) | \mathcal{E}(\vec{n}_1) \mathcal{E}(\vec{n}_2) \dots | O(q) \rangle$$

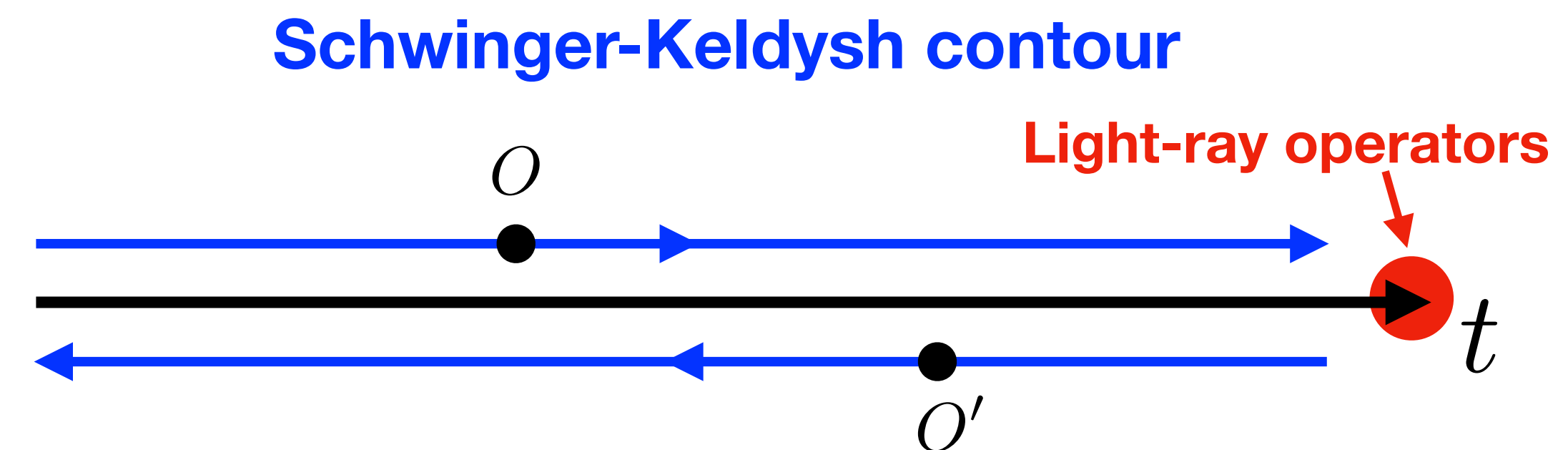
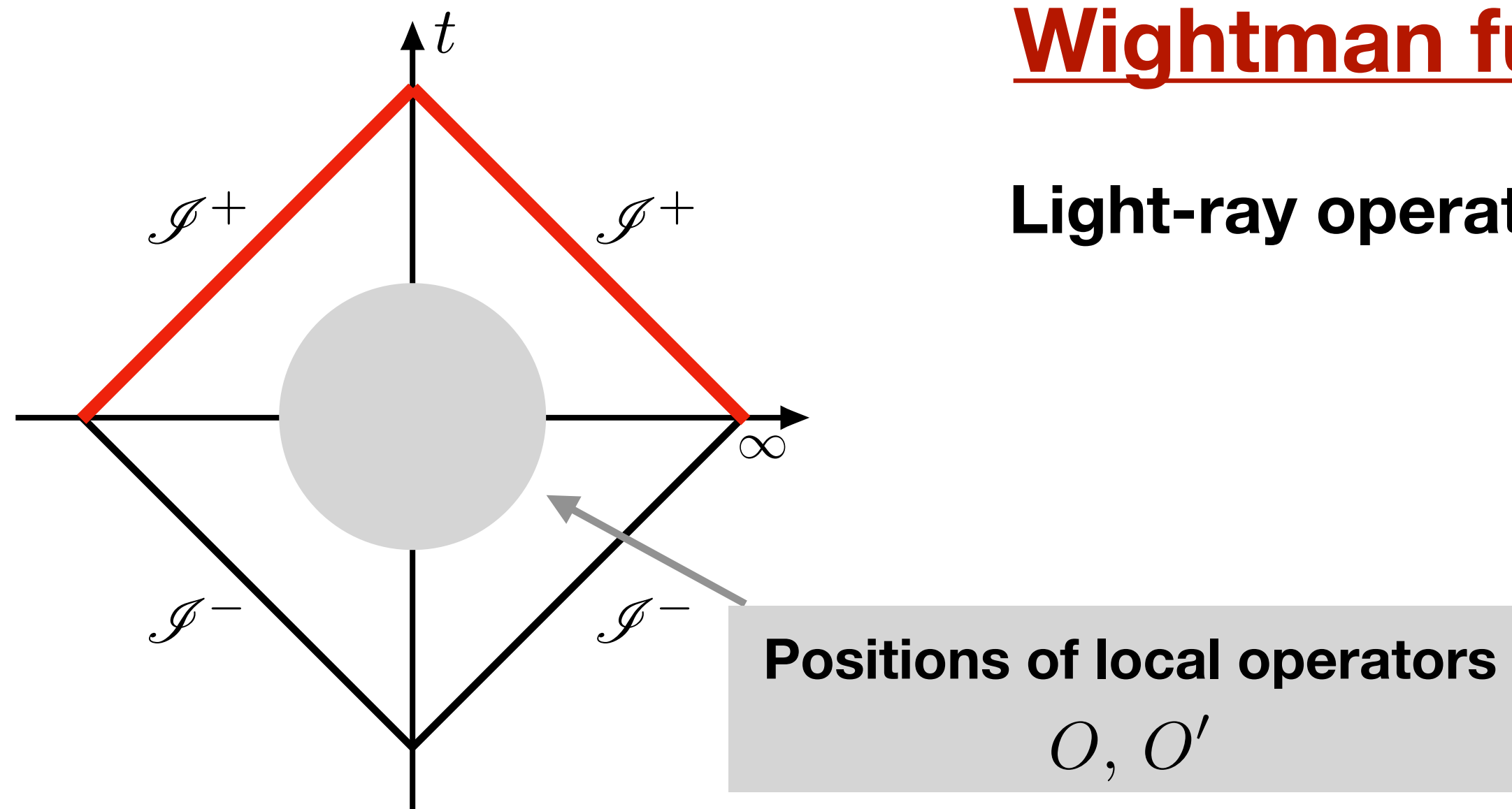
Source with total momentum  $q = (Q, 0, 0, 0)$

[coordinate space]

$$\langle \Omega | O'(x) \mathcal{E}(\vec{n}_1) \mathcal{E}(\vec{n}_2) \dots O(0) | \Omega \rangle$$

**Wightman function, not time-ordered function**

Light-ray operators lie in the future of all local sources



# Light-ray Operators

light transform of local operators

Energy flow operator

$$\mathcal{E}(\vec{n}) = \lim_{r \rightarrow \infty} r^2 \int_0^\infty dt \vec{n}_i T^{0i}(t, r\vec{n})$$

Energy radiation fall with inverse square law,  $r^2$  compensates this effect to be non-vanishing.

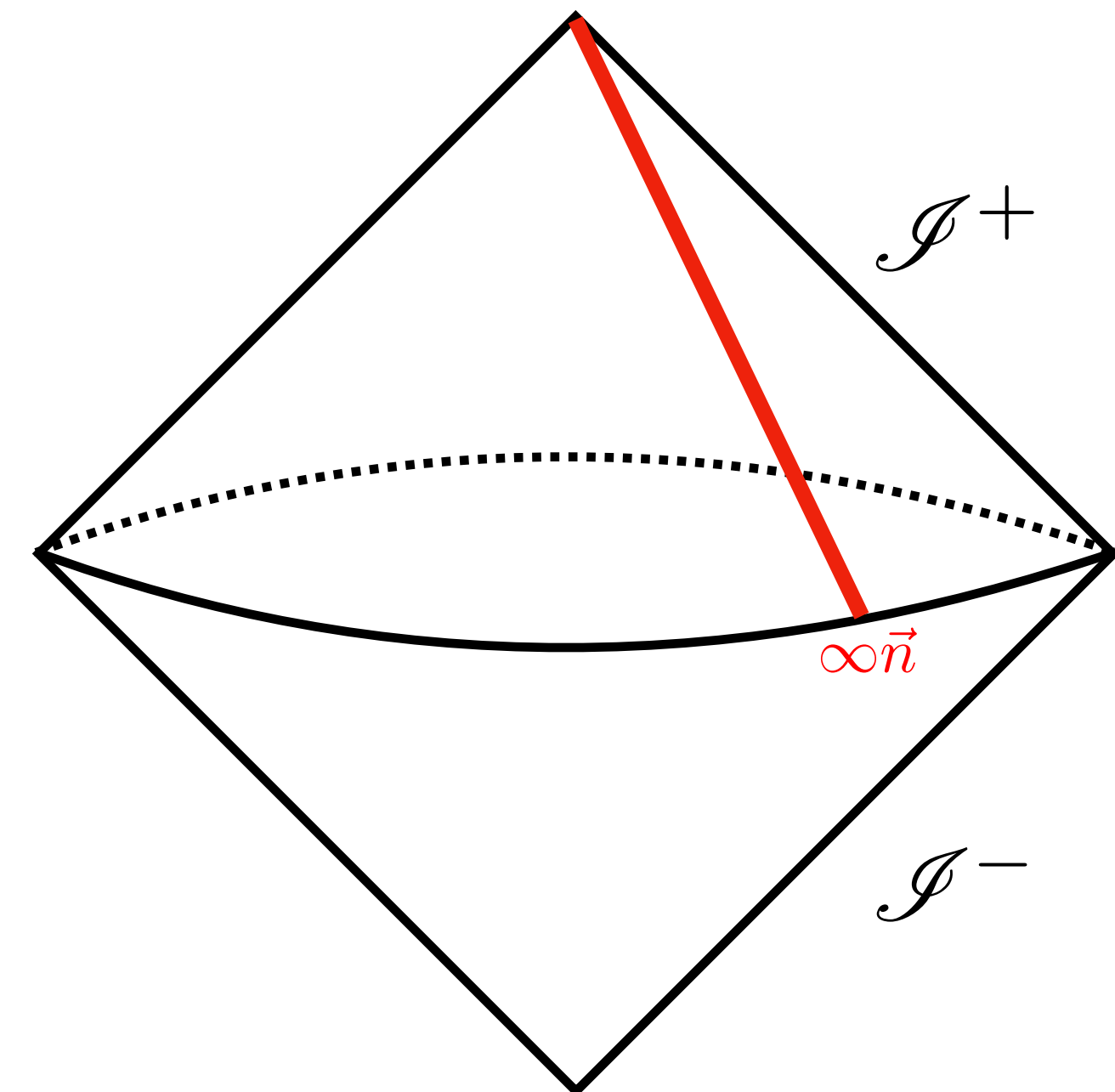
More general light-ray operators

$$\mathbb{O}(\vec{n}) = \lim_{r \rightarrow \infty} r^{\Delta - J} \int_0^\infty dt O^{\mu_1 \dots \mu_J}(t, r\vec{n}) \bar{n}_{\mu_1} \dots \bar{n}_{\mu_J}$$

Lesson from CFT :  $O(x, z) = O^{\mu_1 \mu_2 \dots \mu_n}(x) z_{\mu_1} z_{\mu_2} \dots z_{\mu_n}$

$$\langle O(0, z_1) O(x, z_2) \rangle = \frac{(z_1 \cdot z_2 - 2 \frac{z_1 \cdot x z_2 \cdot x}{x^2})^J}{(x^2)^\Delta}$$

$\xrightarrow{(x^+ \rightarrow +\infty)}$   
 send  $x$  to null infinity  $\sim (x^+)^{J-\Delta}$





# Light-ray Operators

light transform of local operators

Energy flow operator

$$\mathcal{E}(\vec{n}) = \lim_{r \rightarrow \infty} r^2 \int_0^\infty dt \vec{n}_i T^{0i}(t, r\vec{n})$$

Energy radiation fall with inverse square law,  $r^2$  compensates this effect to be non-vanishing.

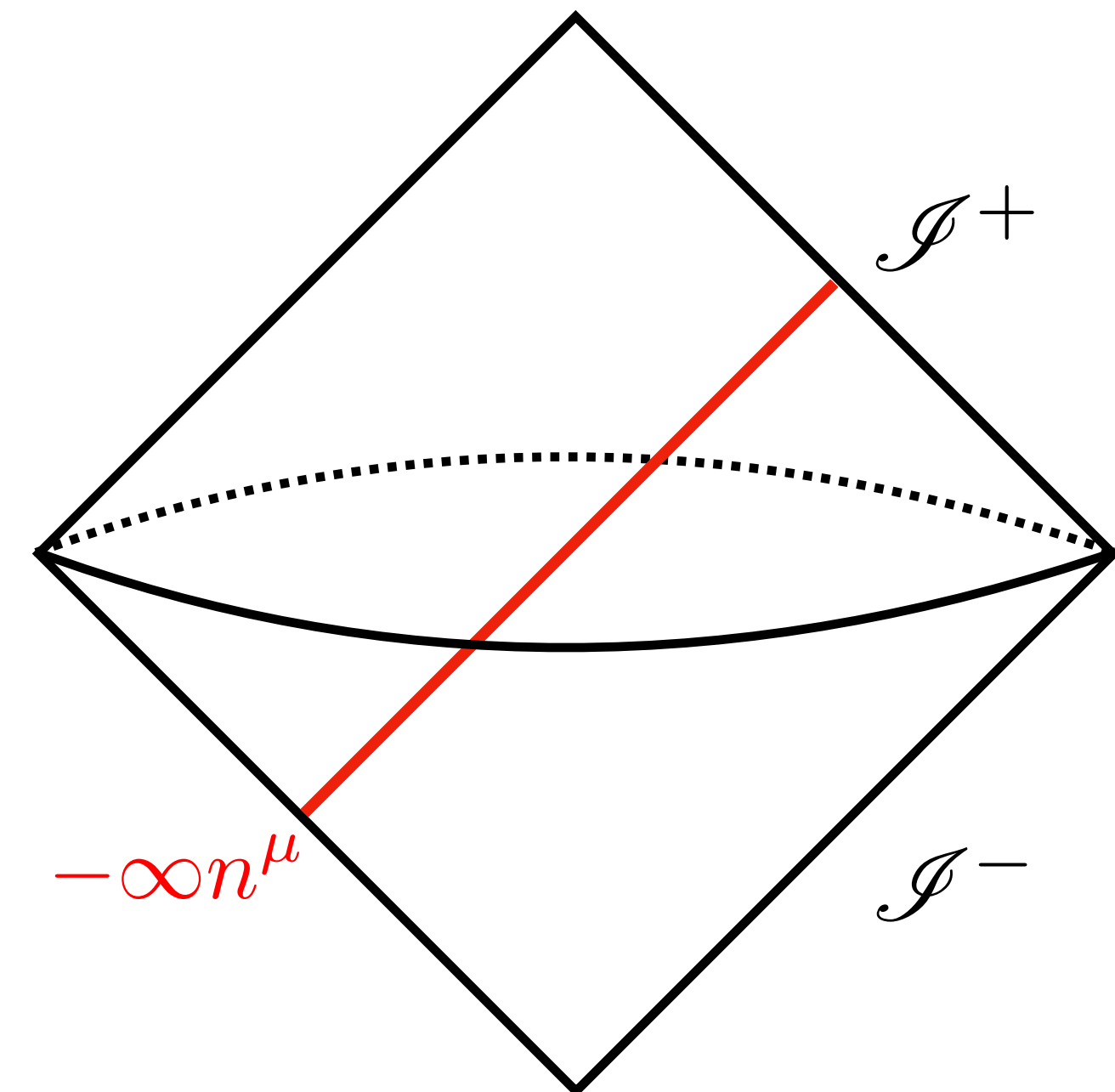
More general light-ray operators

$$\mathbb{O}(\vec{n}) = \lim_{r \rightarrow \infty} r^{\text{twist} - J} \int_0^\infty dt O^{\mu_1 \dots \mu_J}(t, r\vec{n}) \bar{n}_{\mu_1} \dots \bar{n}_{\mu_J}$$

In other contexts of physics, light-ray operators are not necessarily at null infinity—they can live on any light-ray.

$$\mathbf{L}[\mathcal{O}](\mathbf{x}, \mathbf{n}) = \int_{-\infty}^{\infty} d\alpha (-\alpha)^{-\Delta - J} \mathcal{O}\left(x - \frac{n}{\alpha}, n\right)$$

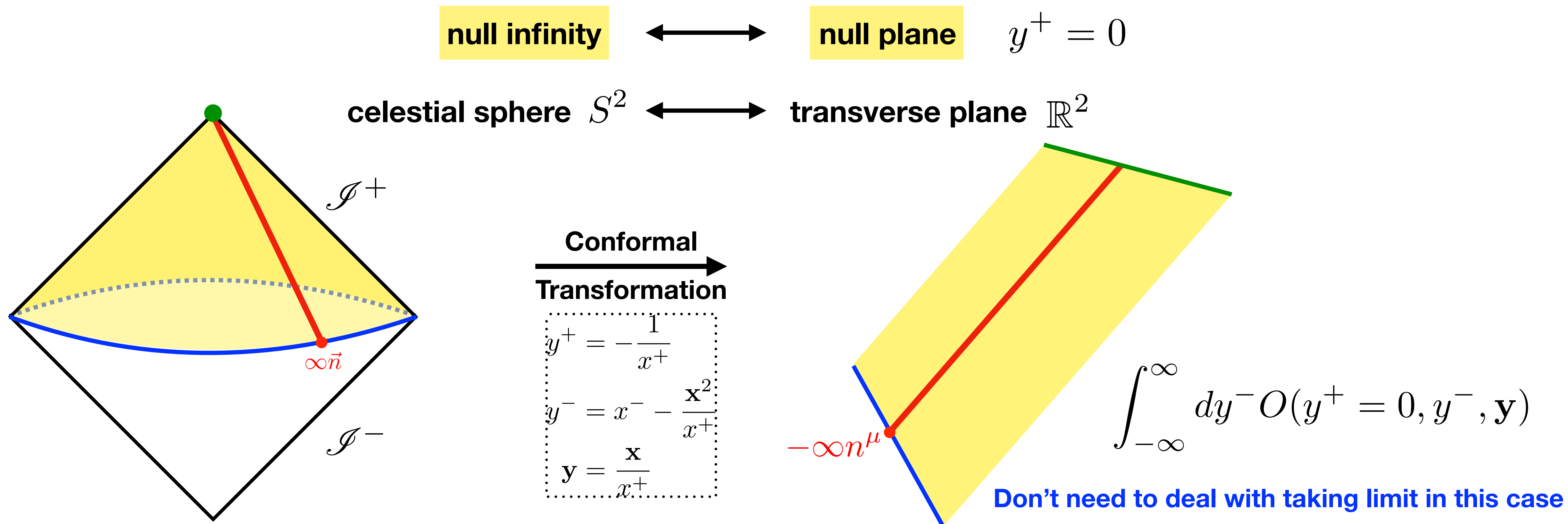
[Kravchuk, Simmons-Duffin, 2018]



# Light-ray Operators

In CFT, different configurations are related by conformal transformation.

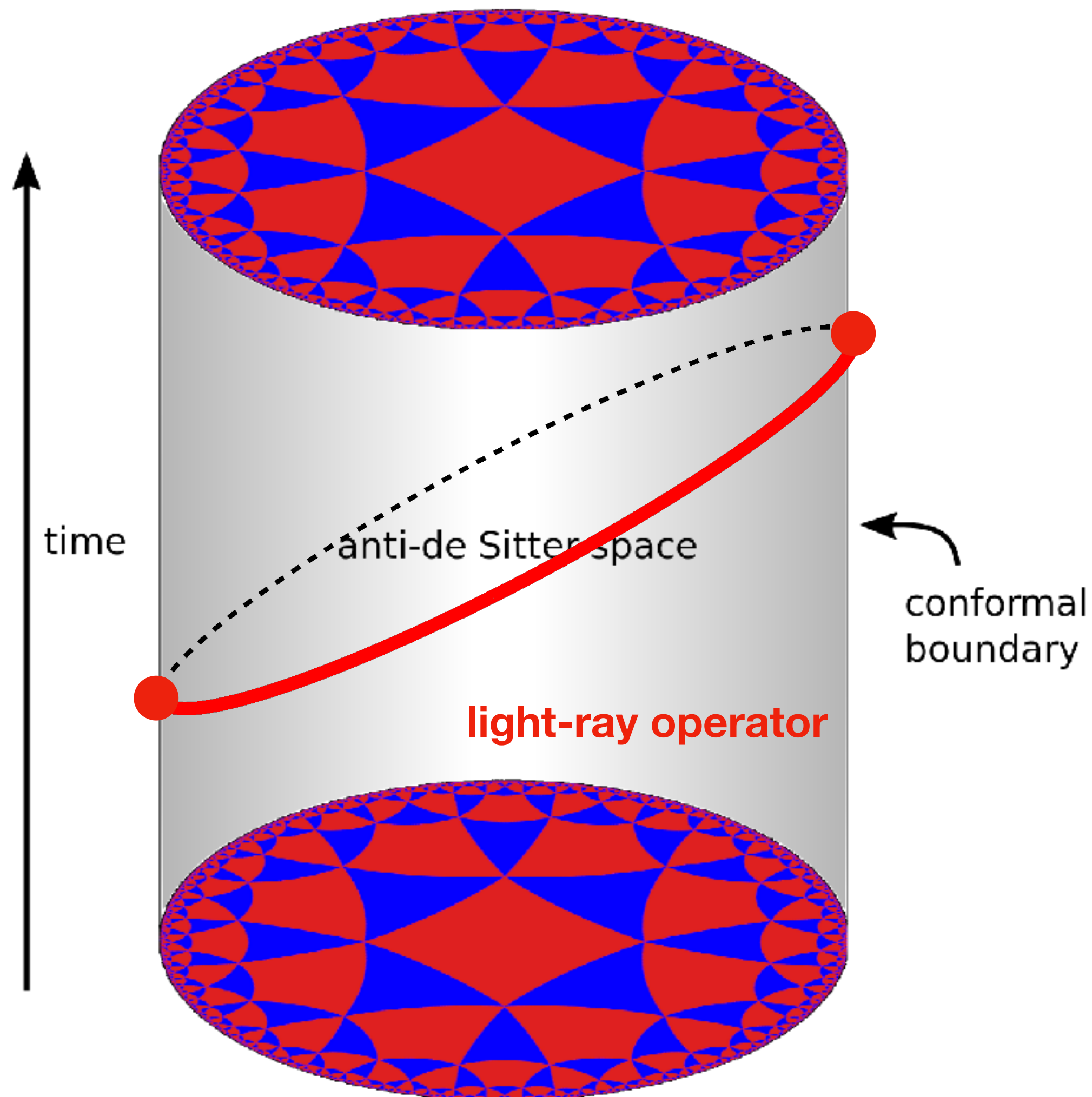
[Hofman and Maldecena, 2008; Kravchuk, Simmons-Duffin, 2018]



In embedding space formalism, they correspond to different gauge fixing.

[see Kologlu, Kravchuk, Simmons-Duffin and Zhiboedov, 2019]

# AdS/CFT Correspondence



CFT

AdS

strong coupling

weak coupling

stress tensor insertion

metric perturbation

$$T^{\mu\nu}$$

$$h_{\mu\nu}$$

energy flow operator  $\mathcal{E}$

gravitational shockwave

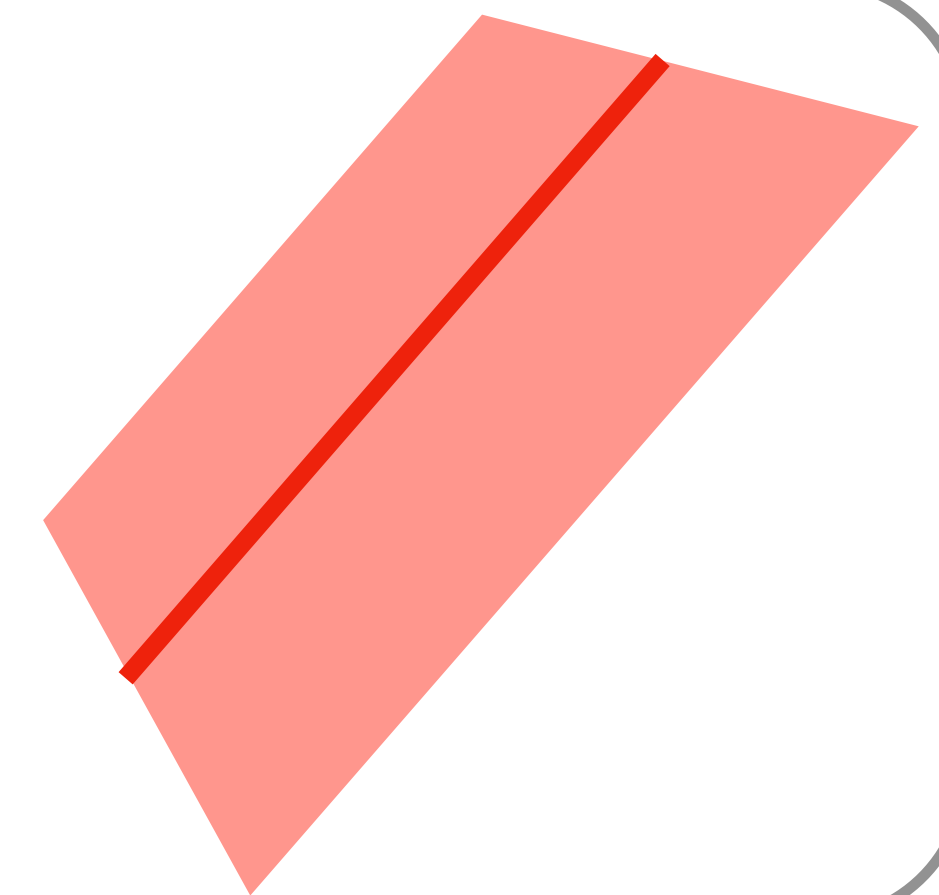
energy correlators

propagation through shockwaves

shockwave (flat space version)

$$h_{--} \propto \delta(y^-) \frac{1}{|\vec{y}_\perp|^{d-4}}$$

localized on the null plane



In strong coupling limit, energy correlations are uniform distributions. [Hofman, Maldacena, 2008]

# Spacetime Symmetry

**Little group** that fixes a **light-ray** in **future null infinity** consists of

translations, collinear boost, transverse rotations, dilatation

Poincare group part

- Dimension =  $J - 1$

$$\mathbb{O}(\vec{n}) = \lim_{r \rightarrow \infty} \boxed{r^{\Delta - J}} \boxed{\int_0^\infty dt} \boxed{O^{\mu_1 \dots \mu_J}(t, r\vec{n}) \bar{n}_{\mu_1} \dots \bar{n}_{\mu_J}}$$

$-(\Delta - J)$ 
 $-1$ 
 $+\Delta$

- Collinear Spin =  $1 - \Delta$

**Boost quantum number**

$$\lim_{\bar{n} \cdot x \rightarrow \infty} (\bar{n} \cdot x)^{\Delta - J} \int_{-\infty}^{\infty} d(n \cdot x) O^{\mu_1 \dots \mu_J}(x) \bar{n}_{\mu_1} \dots \bar{n}_{\mu_J}$$

Boost along  $\vec{n}$

$$n^\mu \rightarrow \lambda n^\mu, \quad \bar{n}^\mu \rightarrow \lambda^{-1} \bar{n}^\mu$$

$$\mathbb{O}(\vec{n}) \rightarrow \lambda^{1-\Delta} \mathbb{O}(\vec{n}) \xrightarrow[\vec{n} \rightarrow n^\mu]{\text{promote}}$$

$$\boxed{\mathbb{O}(\lambda n^\mu) = \lambda^{1-\Delta} \mathbb{O}(n^\mu)}$$

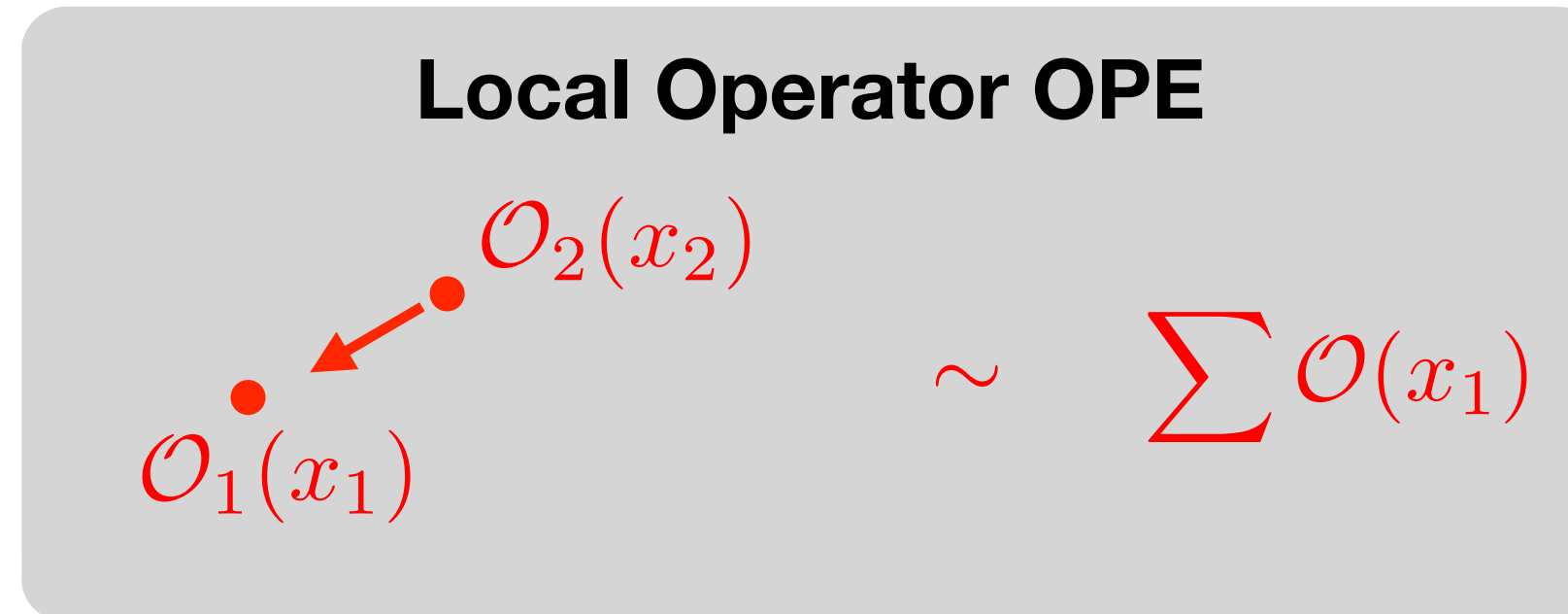
**homogeneous**

- Transverse Spin = transverse spin of the local operator
- Momentum = 0 (invariant under translations)

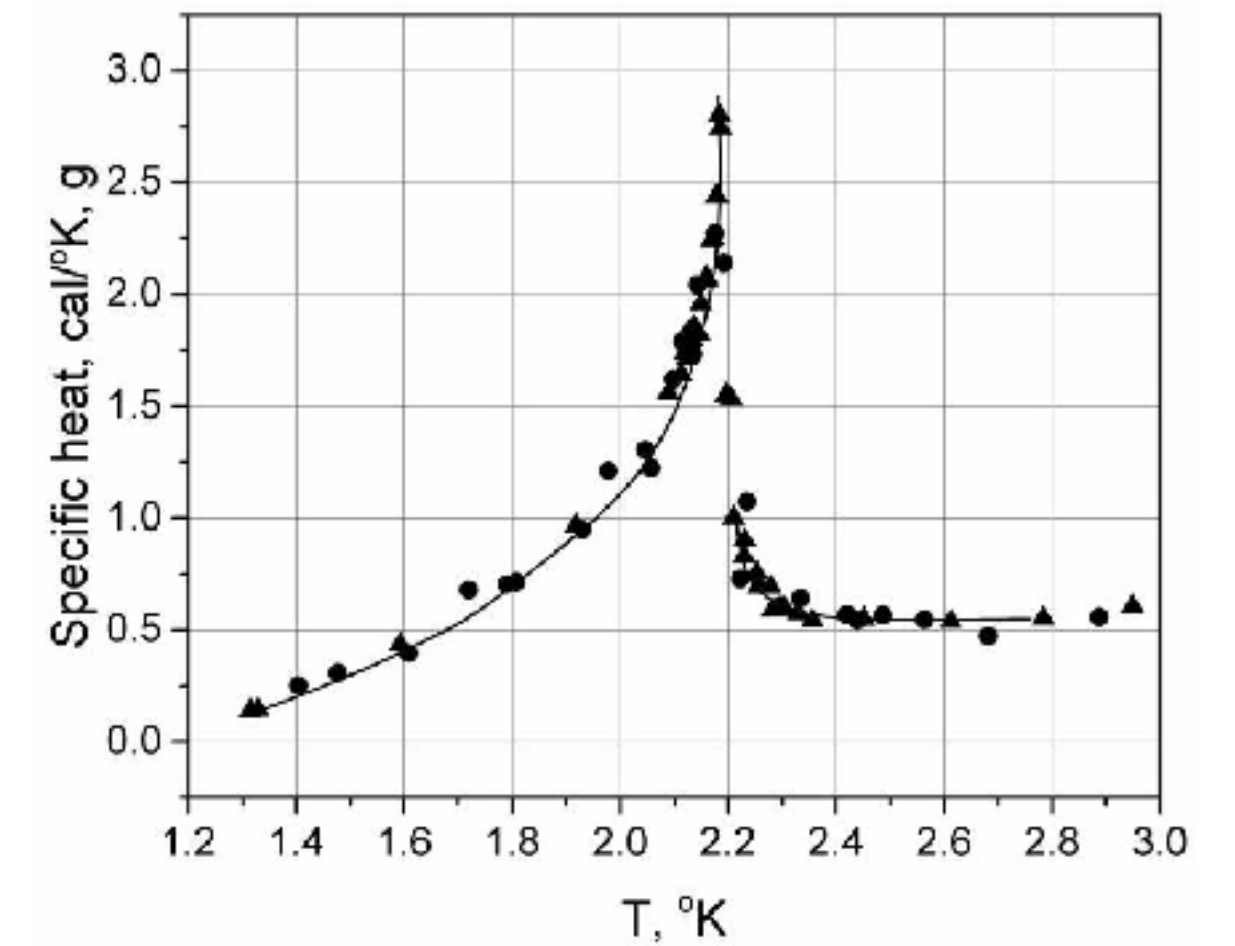
# Light-ray OPE



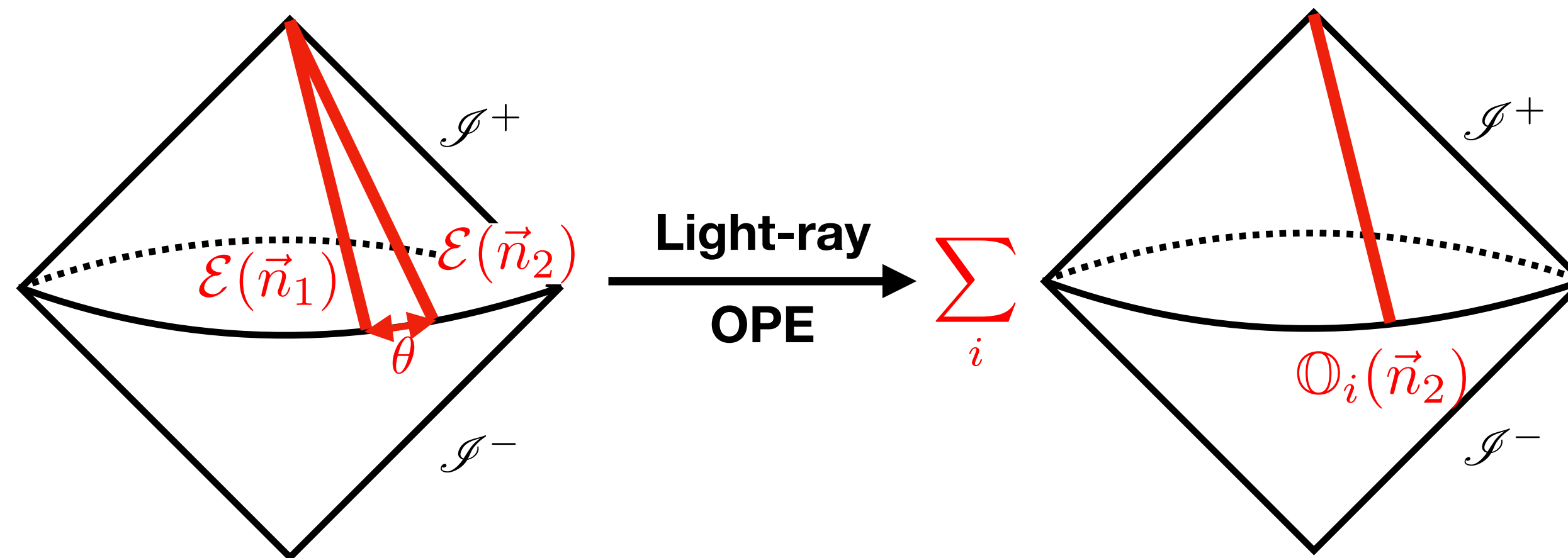
Short distance **scaling behavior** is determined by **local Operator Product Expansion (OPE)**.



**liquid helium critical behavior**

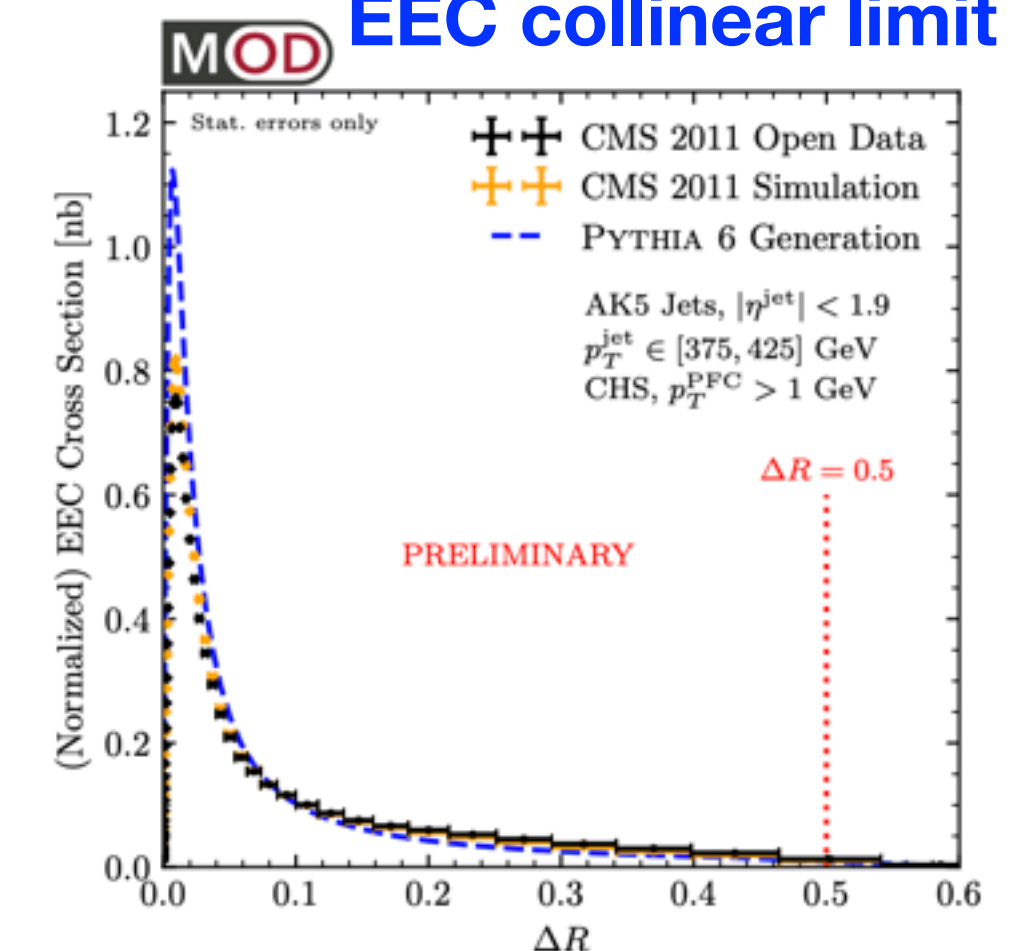


**Small angle behavior** is controlled by the **OPE** of these **light-ray operators**.



“local OPE on the celestial sphere”

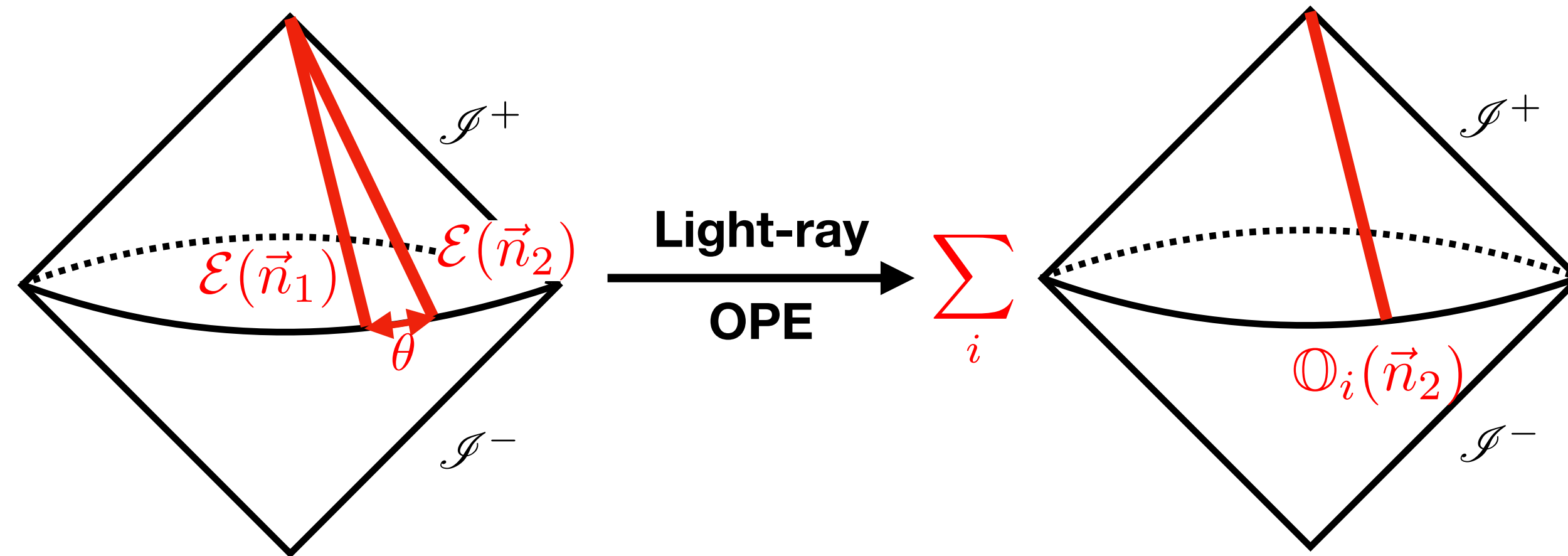
**EEC collinear limit**



[Komiske, Mout, Thaler, Zhu, in preparation]

# Light-ray OPE

## Symmetry and Power Counting



Expansion parameter  $\theta$  is dimensionless, the dimensions simply add up.  
(of light-ray operators)

Recall that the dimension of a light-ray operator is  $J - 1$   $J$  is the collinear spin of the local operator

$$\mathbb{O}_1 \mathbb{O}_2 \supset \mathbb{O} \quad (J_1 - 1) + (J_2 - 1) = J - 1 \quad \text{(Exact in CFT)}$$

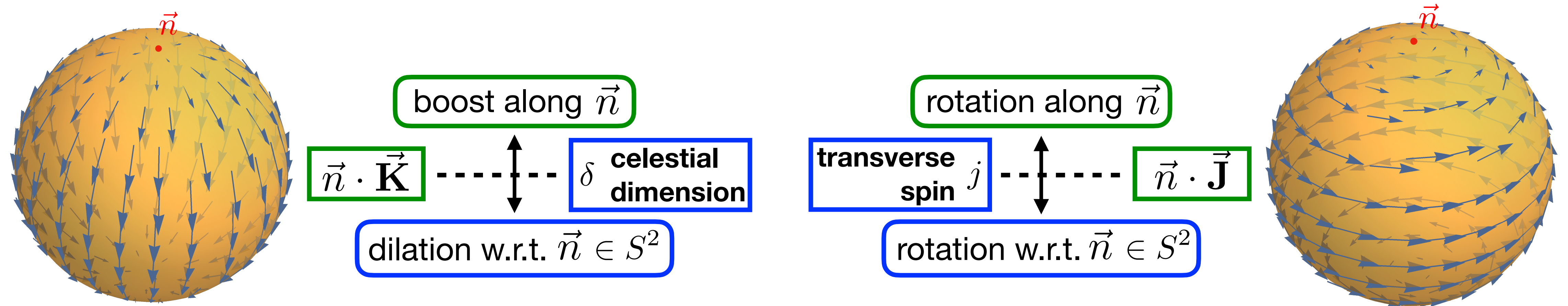
Example: in  $\mathcal{E}\mathcal{E}$  OPE,  $J_1 = J_2 = 1 \Rightarrow J = 3$

different from local OPE:  $J = 3$  operator is absent in the  $TT$  OPE

# Light-ray OPE

## Symmetry and Power Counting

**Lorentz group** is equivalent to the **conformal group** on the **celestial sphere**.



Angle  $\theta$  plays the role of length on the celestial sphere

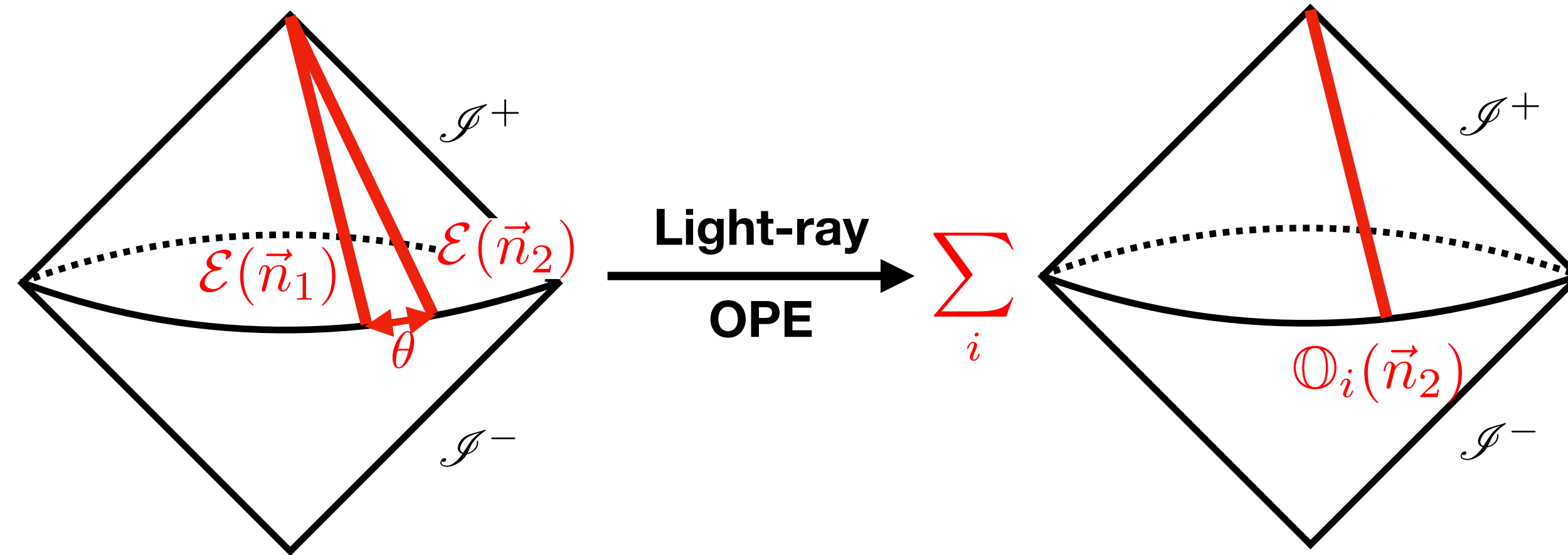
→ power counting on  $\theta$  is related to **celestial dimension (boost quantum number)**

**twist  $\tau_i \equiv \dim \Delta_i - \text{spin } (J_i = 3)$**

**Light-ray OPE**  $\mathcal{E}(\vec{n}_1) \mathcal{E}(\vec{n}_2) \sim \sum_i c_i \theta^{\tau_i - 4} \mathbb{O}_i(\vec{n}_2)$  **[Hofman, Maldacena, 2008]**

collinear spin  $(1 - 4) + (1 - 4) = (\tau_i - 4) + (1 - \Delta_i)$

# Light-ray OPE



**Light-ray OPE**  $\mathcal{E}(\vec{n}_1) \mathcal{E}(\vec{n}_2) \sim \sum_i c_i \theta^{\tau_i - 4} \mathbb{O}_i(\vec{n}_2)$

Small angle scaling is dominated by the leading twist operators.

Light-ray OPE in CFT is rigorous and convergent. [Kologlu, Kravchuk, Simmons-Duffin, Zhiboedov, 2019]  
[Chang, Kologlu, Kravchuk, Simmons-Duffin, Zhiboedov, 2020]

In QCD, things are less understood, but the leading power contribution is. [HC, Mault, Zhu, 2020]





Polchinski: There is a lot of QCD data, can you see this (scaling behavior) there?

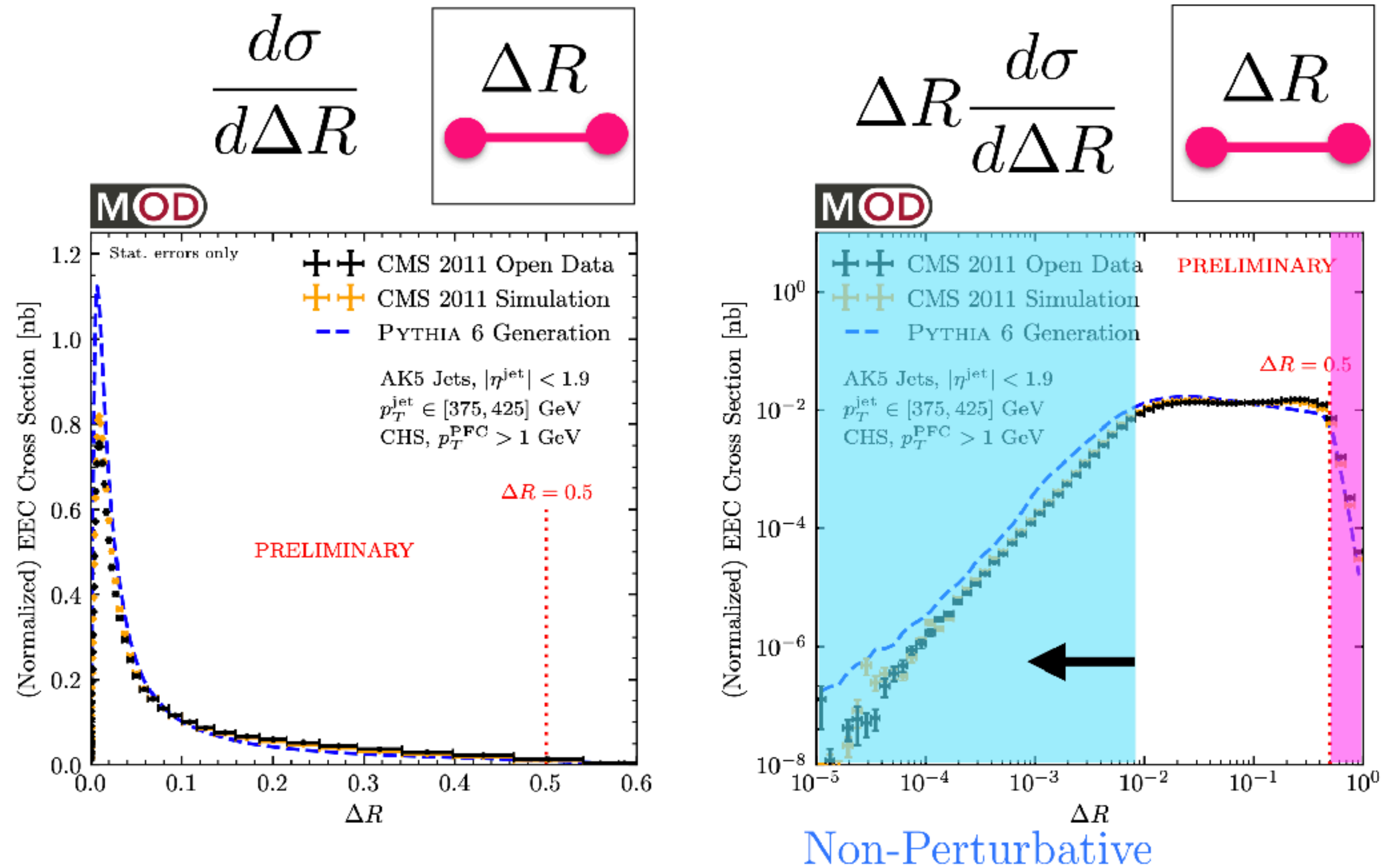
Maldacena: People do not do this. I haven't figured out why they don't. I think they just haven't thought about this. I was talking to people who did this calculation of two-point function at LEP, computing  $\alpha_s$  and so on, and they focused mostly on the large angles. But they didn't study the small angles. And I asked him whether they had a good reason for not studying the small angles and they said well we didn't know the resummation formula, didn't study it.

# EEC with CMS Open Data

[Komiske, Mout, Thaler, Zhu, in preparation]

Packaged in “MIT Open Data”, provided by Jesse Thaler and Patrick Komiske

Nice scaling behavior in perturbative regime



**Application in QCD**

# Leading Twist Operators in QCD

## Local Operators

transverse spin-0

$$\mathcal{O}_q^{[J]} = \frac{1}{2^J} \bar{\psi} \gamma^+ (iD^+)^{J-1} \psi$$

$$\mathcal{O}_g^{[J]} = -\frac{1}{2^J} F_a^{\mu+} (iD^+)^{J-2} F_a^{\mu+}$$

transverse spin-2

helicity  $\pm$

$$\mathcal{O}_{\tilde{g},\lambda}^{[J]} = -\frac{1}{2^J} F_a^{\mu+} (iD^+)^{J-2} F_a^{\nu+} \epsilon_{\lambda,\mu} \epsilon_{\lambda,\nu}$$

$$\lim_{r \rightarrow \infty} r^2 \int_0^\infty dt$$

$$\vec{\mathbb{O}}^{[J]}(\vec{n}) =$$

$$\mathbb{O}_q^{[J]}(\vec{n})$$

$$\mathbb{O}_g^{[J]}(\vec{n})$$

$$\mathbb{O}_{\tilde{g},+}^{[J]}(\vec{n})$$

$$\mathbb{O}_{\tilde{g},-}^{[J]}(\vec{n})$$

unpolarized

polarized

## Mode Expansion

$$\mathbb{O}_q^{[J]}(\vec{n}) = \sum_s \int \frac{d^3 p}{(2\pi)^3 2E} \delta^{(2)}(\hat{p} - \vec{n}) E^{J-1} (b_{\vec{p},s}^\dagger b_{\vec{p},s} + d_{\vec{p},s}^\dagger d_{\vec{p},s})$$

$$\mathbb{O}_g^{[J]}(\vec{n}) = \sum_{\lambda,c} \int \frac{d^3 p}{(2\pi)^3 2E} \delta^{(2)}(\hat{p} - \vec{n}) E^{J-1} a_{\vec{p},\lambda,c}^\dagger a_{\vec{p},\lambda,c}$$

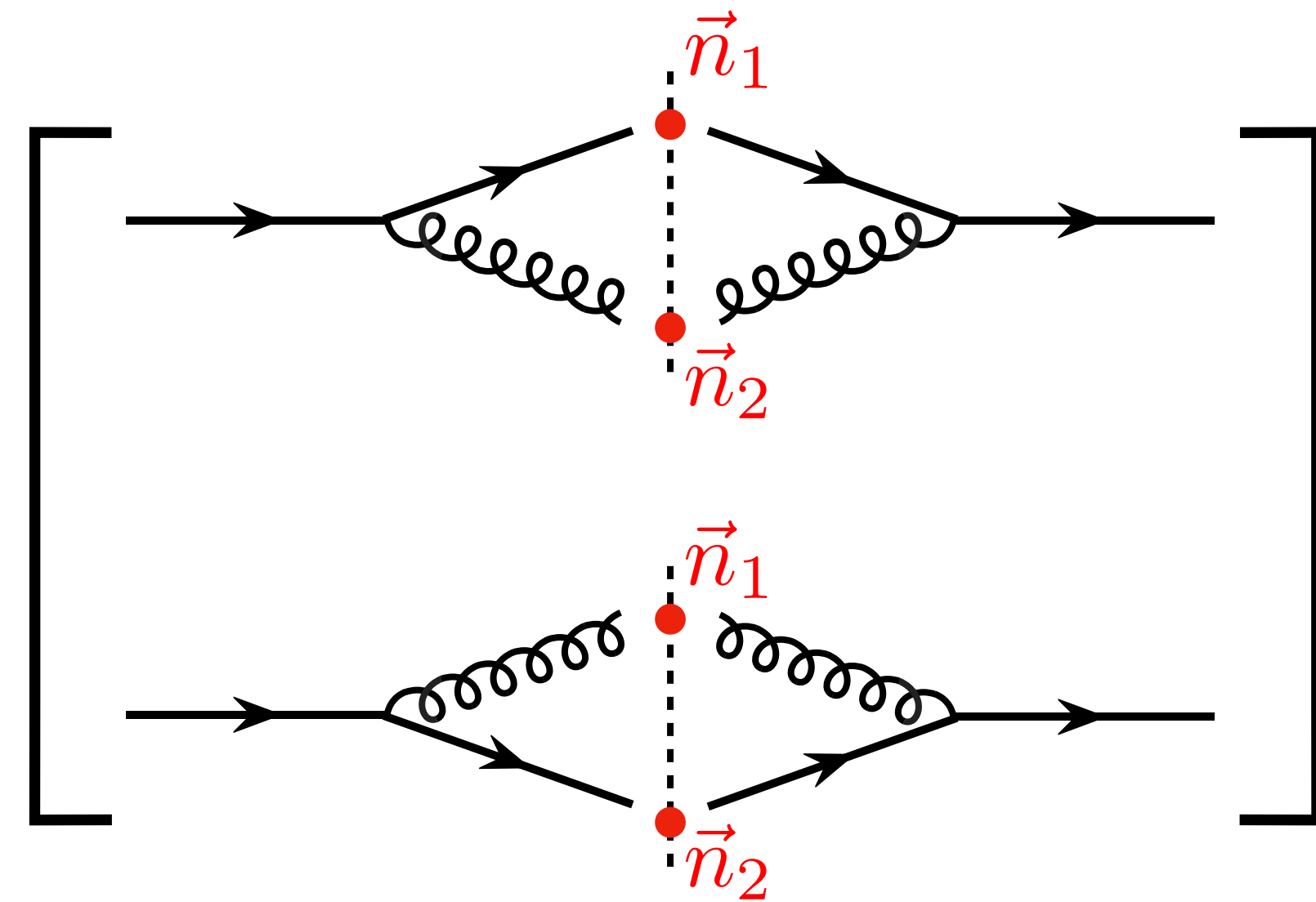
$$\mathbb{O}_{\tilde{g},\lambda}^{[J]}(\vec{n}) = - \sum_c \int \frac{d^3 p}{(2\pi)^3 2E} \delta^{(2)}(\hat{p} - \vec{n}) E^{J-1} a_{\vec{p},\lambda,c}^\dagger a_{\vec{p},-\lambda,c}$$

# $\mathcal{E}(\vec{n}_1)\mathcal{E}(\vec{n}_2)$ OPE

in quark state

$$\langle \Omega | \psi(x) \mathcal{E}(\vec{n}_1)\mathcal{E}(\vec{n}_2) \bar{\psi}(0) | \Omega \rangle$$

$$= \int \frac{E_1^2 dE_1}{(2\pi)^3 2E_1} \frac{E_2^2 dE_2}{(2\pi)^3 2E_2} E_1 E_2 e^{-i(p_1+p_2)\cdot x}$$



Wilson coefficient

$$\xrightarrow{\theta \rightarrow 0} -\frac{1}{2\pi} \frac{2}{\theta^2} \left[ (\gamma_{qq}(2) - \gamma_{qq}(3)) + (\gamma_{gq}(2) - \gamma_{gq}(3)) \right] \langle \Omega | \psi(x) \mathcal{O}_q^{[3]}(\vec{n}_2) \bar{\psi}(0) | \Omega \rangle$$

related to twist-2 Anom. Dim.

$$\int \frac{d^3 p}{(2\pi)^3 2E} \delta^{(2)}(\hat{p} - \vec{n}) \not{n} E^3 e^{-ip\cdot x}$$

Confirms the general analysis: (1) spin-3 operator, (2) correct scaling behavior

# $\mathcal{E}(\vec{n}_1)\mathcal{E}(\vec{n}_2)$ OPE

in gluon state

$$\langle \Omega | A_b^\nu(x) \mathcal{E}(\vec{n}_1)\mathcal{E}(\vec{n}_2) A_a^\mu(0) | \Omega \rangle$$

$$= \int \frac{E_1^2 dE_1}{(2\pi)^3 2E_1} \frac{E_2^2 dE_2}{(2\pi)^3 2E_2} E_1 E_2 e^{-i(p_1+p_2)\cdot x} \left[ \begin{array}{c} \text{Diagram 1: } \mu, a \text{ and } \nu, b \text{ gluons meeting at } \vec{n}_1, \vec{n}_2 \\ \text{Diagram 2: } \mu, a \text{ and } \nu, b \text{ gluons meeting at } \vec{n}_1, \vec{n}_2 \end{array} \right]$$

$$\xrightarrow{\theta \rightarrow 0} -\frac{1}{2\pi} \frac{2}{\theta^2} \left[ \textcircled{c_g} \langle \Omega | A_b^\nu(x) \mathbb{O}_g^{[3]} A_a^\mu(0) | \Omega \rangle + \textcircled{c_{\tilde{g}}} \left( e^{2i\phi} \langle \Omega | A_b^\nu(x) \mathbb{O}_{\tilde{g},-}^{[3]} A_a^\mu(0) | \Omega \rangle + e^{-2i\phi} \langle \Omega | A_b^\nu(x) \mathbb{O}_{\tilde{g},+}^{[3]} A_a^\mu(0) | \Omega \rangle \right) \right]$$

Wilson coefficients

unpolarized  $\textcircled{c_g} = (\gamma_{gg}(2) - \gamma_{gg}(3)) + 2n_f (\gamma_{qg}(2) - \gamma_{qg}(3))$

polarized  $\textcircled{c_{\tilde{g}}} = (\gamma_{g\tilde{g}}(2) - \gamma_{g\tilde{g}}(3)) + 2n_f (\gamma_{q\tilde{g}}(2) - \gamma_{q\tilde{g}}(3))$

# $\vec{\mathbb{O}}^{[J]}(\vec{n}_1)\mathcal{E}(\vec{n}_2)$ **OPE**

Introduce the Wilson coefficient matrix

$$\hat{C}_\phi(J) = \begin{pmatrix} \gamma_{qq}(J) & 2n_f\gamma_{qg}(J) & 2n_f\gamma_{q\tilde{g}}(J)e^{-2i\phi}/2 & 2n_f\gamma_{q\tilde{g}}(J)e^{2i\phi}/2 \\ \gamma_{gq}(J) & \gamma_{gg}(J) & \gamma_{g\tilde{g}}(J)e^{-2i\phi}/2 & \gamma_{g\tilde{g}}(J)e^{2i\phi}/2 \\ \gamma_{\tilde{g}q}(J)e^{2i\phi} & \gamma_{\tilde{g}g}(J)e^{2i\phi} & \gamma_{\tilde{g}\tilde{g}}(J) & \gamma_{\tilde{g}\tilde{g},\pm}(J)e^{4i\phi} \\ \gamma_{\tilde{g}q}(J)e^{-2i\phi} & \gamma_{\tilde{g}g}(J)e^{-2i\phi} & \gamma_{\tilde{g}\tilde{g},\pm}(J)e^{-4i\phi} & \gamma_{\tilde{g}\tilde{g}}(J) \end{pmatrix}$$

**OPE leading contribution**

$$\vec{\mathbb{O}}^{[J]}(\hat{n}_1)\mathcal{E}(\hat{n}_2) = -\frac{1}{2\pi}\frac{2}{\theta^2} \left[ \hat{C}_\phi(J) - \hat{C}_\phi(J+1) \right] \vec{\mathbb{O}}^{[J+1]}(\hat{n}_1) + \text{higher twist}$$

Energy flow operator corresponds to  $J = 2$ , and doesn't distinguish quark and gluon.

$$\mathcal{E} \sim \mathbb{O}_q^{[2]} + \mathbb{O}_g^{[2]}$$

$$\mathcal{E}\mathcal{E} \text{ OPE} \quad \mathcal{E}(\hat{n}_1)\mathcal{E}(\hat{n}_2) = -\frac{1}{2\pi}\frac{2}{\theta^2} \vec{\mathcal{J}} \left[ \hat{C}_\phi(2) - \hat{C}_\phi(3) \right] \vec{\mathbb{O}}^{[3]}(\hat{n}_1) + \text{higher twist}$$

$$\vec{\mathcal{J}} = (1, 1, 0, 0)$$

# 3-point Energy Correlator

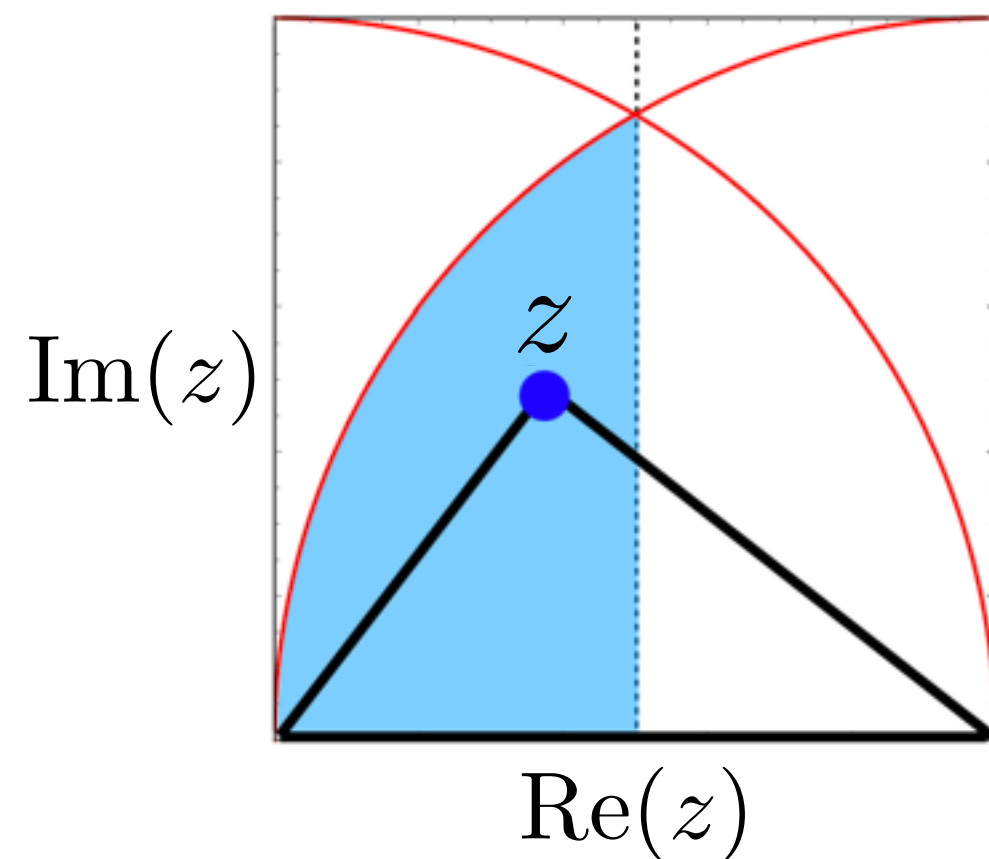
in the collinear limit

## Kinematics

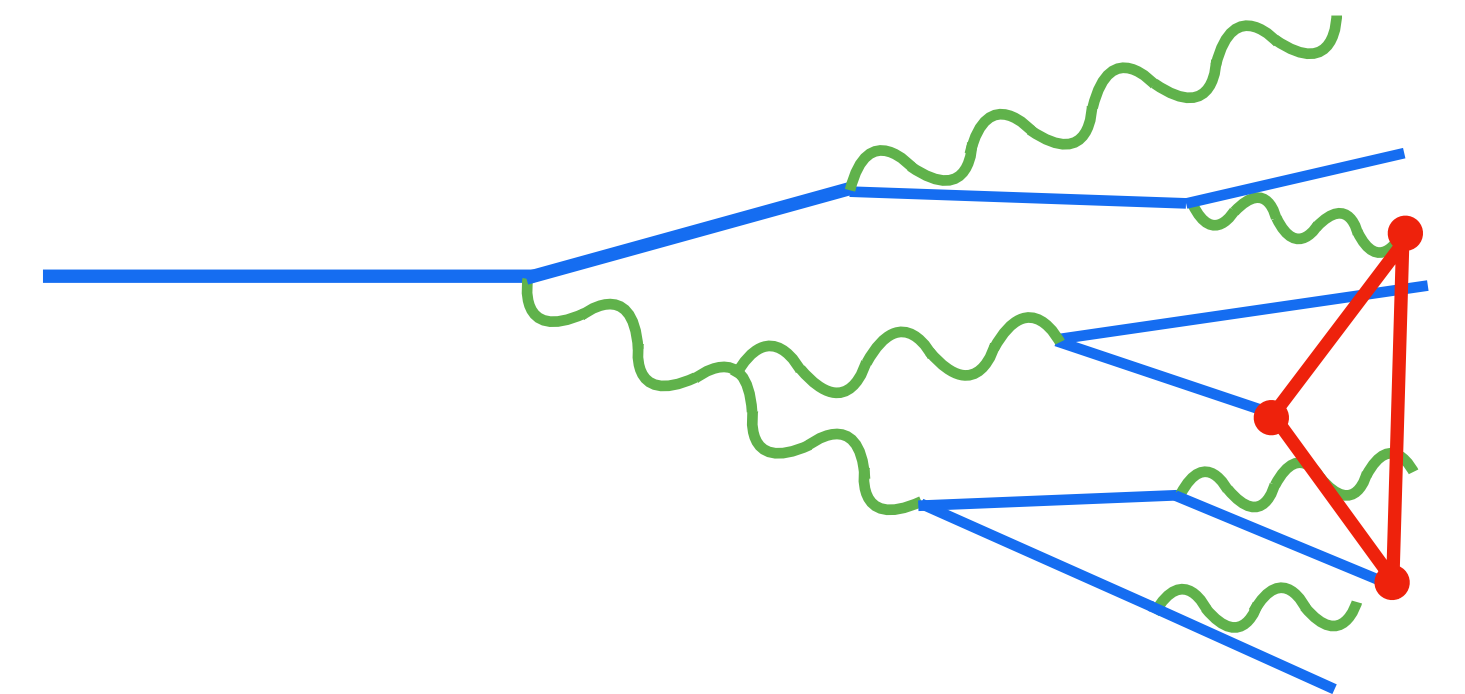
In the collinear limit, EEEC configuration can be approximated by a triangle.

[HC, Luo, Mout, Yang, Zhang, Zhu, 2019]

Moduli space of triangle shape



Parameterized in terms of  
 (1) the longest side<sup>2</sup>  $x_L$  [Size]  
 (2) a complex number  $z$  [Shape]



$$\frac{E_a E_b E_c}{(Q/2)^3} \delta\left(x_1 - \frac{s_{ab}}{4E_a E_b}\right) \delta\left(x_2 - \frac{s_{bc}}{4E_b E_c}\right) \delta\left(x_3 - \frac{s_{ac}}{4E_a E_c}\right)$$

**Definition**

$$\frac{1}{\sigma_{\text{tot}}} \frac{d^3 \Sigma_i}{dx_1 dx_2 dx_3} = \sum_{a, b, c} \int \boxed{d\Phi_c^{(3)}} \boxed{P_{abc}^{(i)}} \boxed{\mathcal{M}_{\text{EEEC}}}$$

collinear phase space      measurement

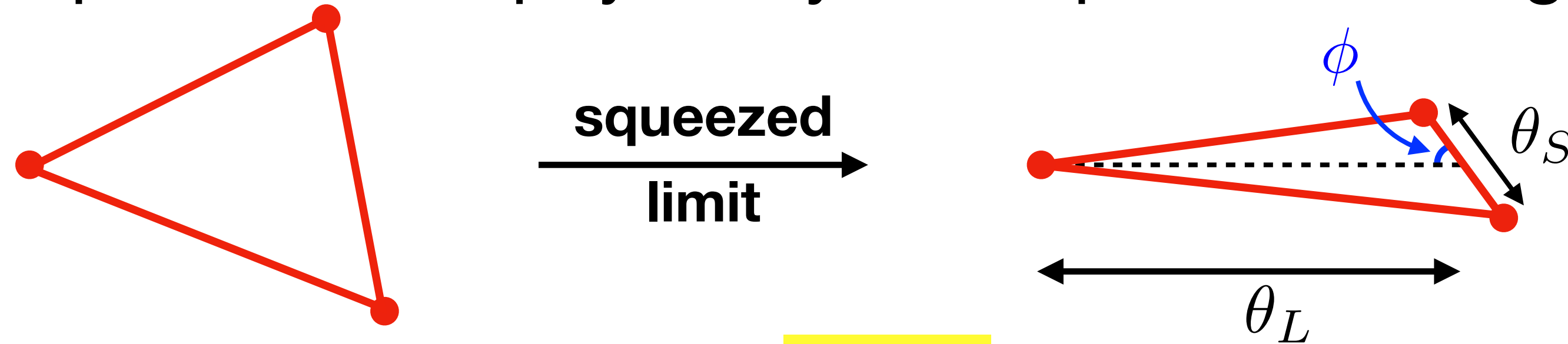
1 → 3 splitting probability

$x_1, x_2, x_3$  length<sup>2</sup> of each side ~ angle<sup>2</sup>



# Squeezed Limit

Squeezed limit physically corresponds to bringing two detectors very close.

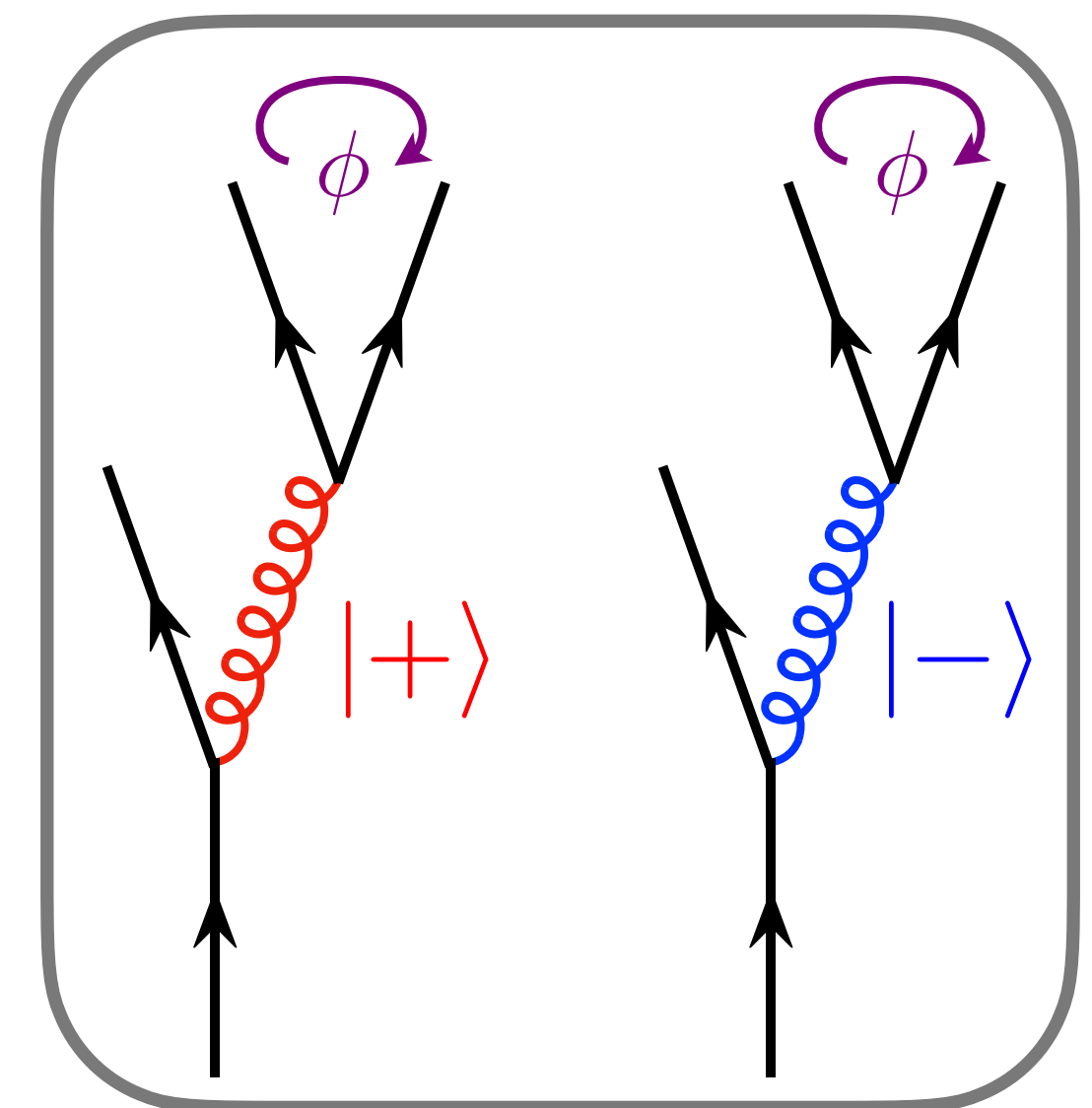


$$\frac{d^3 \Sigma_i}{d\theta_L^2 d\theta_S^2 d\phi} \simeq \frac{1}{\pi} \left(\frac{\alpha_s}{4\pi}\right)^2 \frac{\text{Sq}_i^{(0)}(\phi)}{\theta_L^2 \theta_S^2} + \dots$$

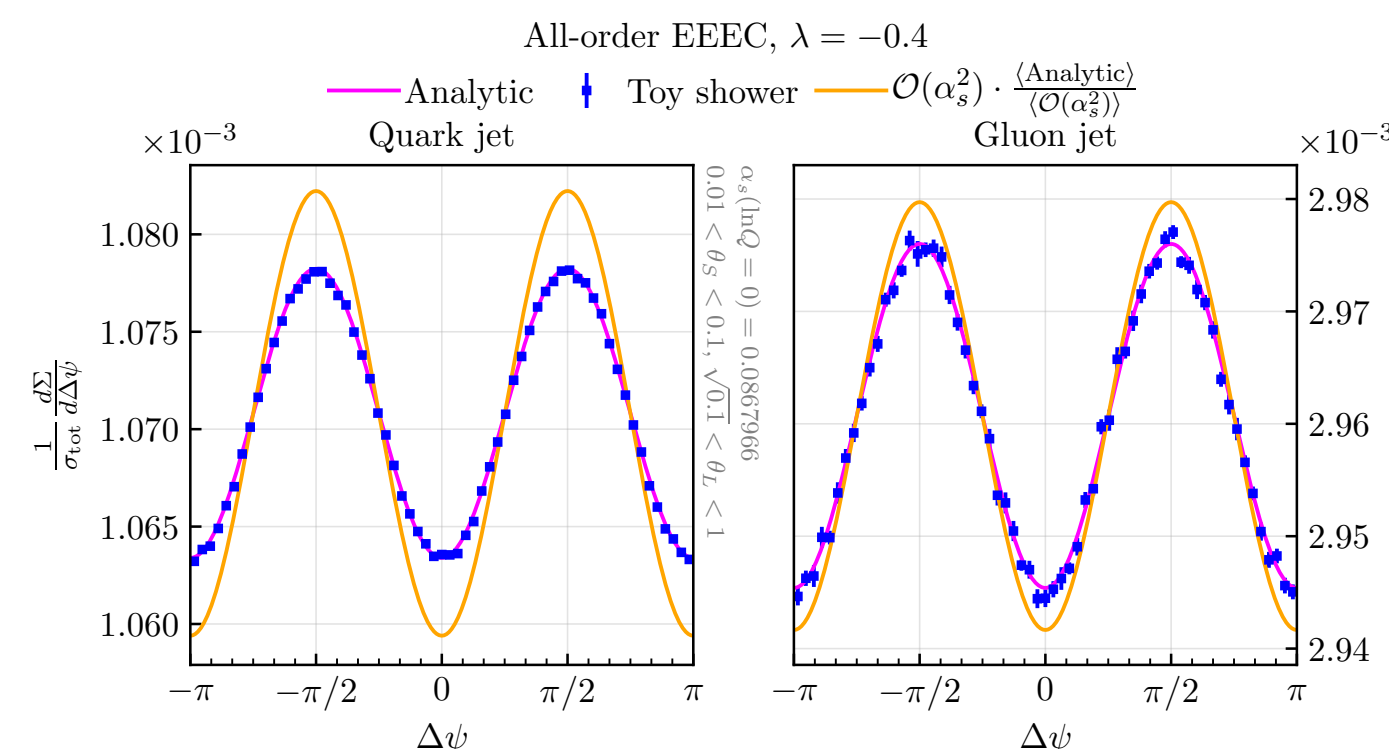
$$\text{Sq}_q^{(0)}(\phi) = C_F n_f T_F \left( \frac{39 - 20 \cos(2\phi)}{225} \right) + C_F C_A \left( \frac{273 + 10 \cos(2\phi)}{225} \right) + C_F^2 \frac{16}{5}$$

$$\text{Sq}_g^{(0)}(\phi) = C_A n_f T_F \left( \frac{126 - 20 \cos(2\phi)}{225} \right) + C_A^2 \left( \frac{882 + 10 \cos(2\phi)}{225} \right) + C_F n_f T_F \frac{3}{5}$$

Interference Effect



Squeezed limit encodes spin correlation information and the Leading Power resummation is done. **[HC, Mout, Zhu, 2020]**



When collinear spin correlation is included in the **PanScales** family of parton showers, our resummed result provides validation of shower results.

**[Karlberg, Salam, Scyboz, Verheyen, 2021]**

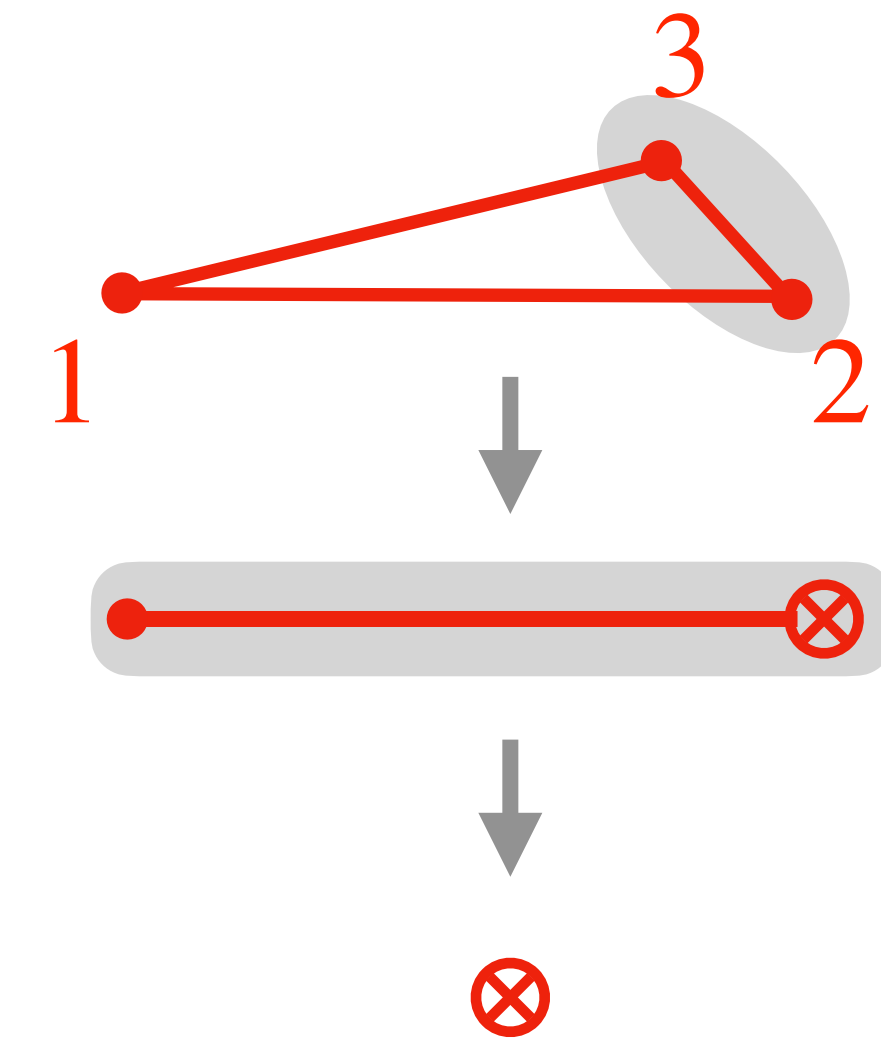
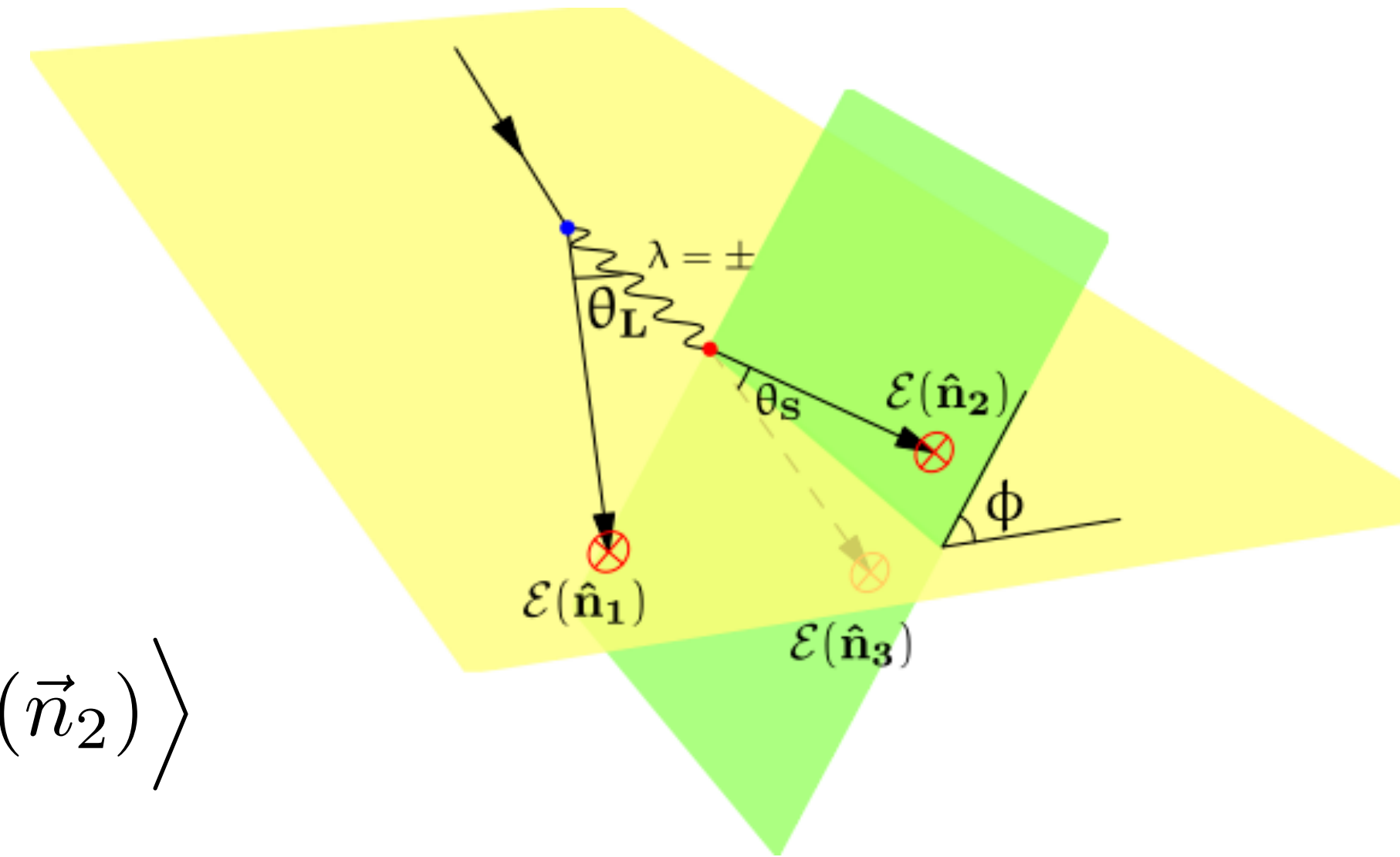
Contributing Operator

$$F_a^{\mu+} (iD^+) F_a^{\nu+} \epsilon_{\lambda,\mu} \epsilon_{\lambda,\nu}$$

**twist-2, transverse spin-2 gluonic operator**

# Squeezed Limit

from light-ray OPE



$$\langle \mathcal{E}(\vec{n}_1) \mathcal{E}(\vec{n}_2) \mathcal{E}(\vec{n}_3) \rangle$$

$$\longrightarrow -\frac{1}{2\pi} \frac{2}{\theta_S^2} \vec{\mathcal{J}} \left[ \hat{C}_{\phi_S}(2) - \hat{C}_{\phi_S}(3) \right] \langle \mathcal{E}(\vec{n}_1) \vec{\mathcal{O}}^{[3]}(\vec{n}_2) \rangle$$

$$\longrightarrow \frac{1}{(2\pi)^2} \frac{2}{\theta_S^2} \frac{2}{\theta_L^2} \vec{\mathcal{J}} \left[ \hat{C}_{\phi_S}(2) - \hat{C}_{\phi_S}(3) \right] \left[ \hat{C}_{\phi_L}(3) - \hat{C}_{\phi_L}(4) \right] \boxed{\langle \vec{\mathcal{O}}^{[4]}(\vec{n}_1) \rangle} \text{ Hard function}$$

This correctly reproduces the previous fixed order squeezed limit results.

**Hierarchy**  $\theta_L \gg \theta_S$

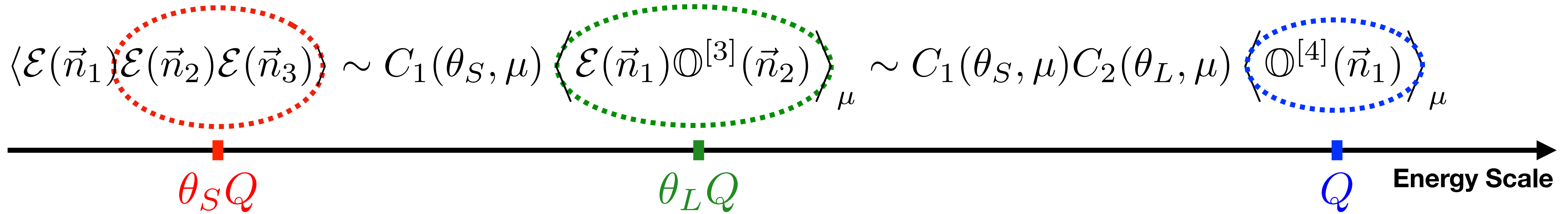
**fixed order result needs to be resummed !**

**unpolarized jet**  $\langle \mathcal{O}_{\tilde{g}, \pm}^{[J]} \rangle = 0$

**quark jet**  $\langle \vec{\mathcal{O}}^{[4]}(\vec{n}_1) \rangle = (1, 0, 0, 0)$

**gluon jet**  $\langle \vec{\mathcal{O}}^{[4]}(\vec{n}_1) \rangle = (0, 1, 0, 0)$

# RG evolution



transverse  
spin-0

$$\mathcal{O}_q^{[J]} = \frac{1}{2^J} \bar{\psi} \gamma^+ (iD^+)^{J-1} \psi$$

$$\mathcal{O}_g^{[J]} = -\frac{1}{2^J} F_a^{\mu+} (iD^+)^{J-2} F_a^{\mu+}$$

transverse  
spin-2

$$\mathcal{O}_{\tilde{g},ij}^{[J]} = -\frac{1}{2^J} F_a^{(i+} (iD^+)^{J-2} F_a^{j)+}$$

$$\left[ \mathcal{O}_{\tilde{g},\lambda}^{[J]} = -\frac{1}{2^J} F_a^{\mu+} (iD^+)^{J-2} F_a^{\nu+} \epsilon_{\lambda,\mu} \epsilon_{\lambda,\nu} \right]$$

CANNOT  
MIX

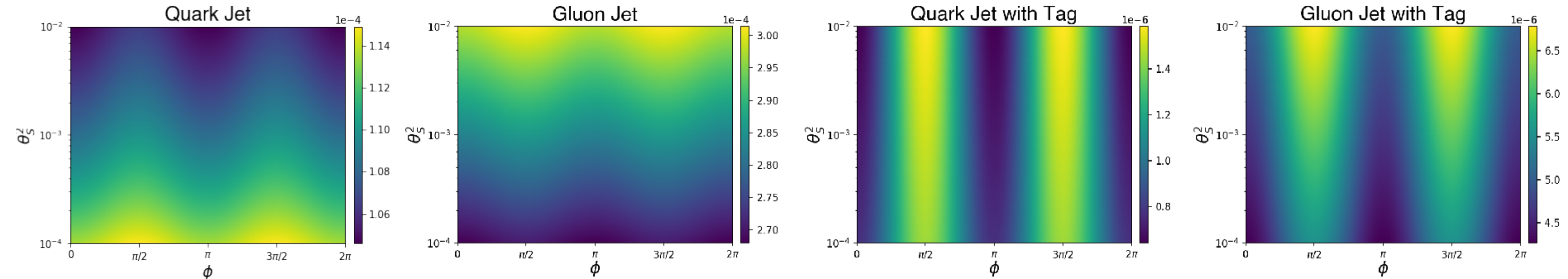
RG equation:

$$\frac{d}{d \ln \mu^2} \vec{\mathcal{O}}^{[J]} = -\hat{\gamma}(J) \cdot \vec{\mathcal{O}}^{[J]}$$

$$\hat{\gamma}(J) = \begin{pmatrix} \gamma_{qq}(J) & 2n_f \gamma_{qg}(J) & 0 \\ \gamma_{gq}(J) & \gamma_{gg}(J) & 0 \\ 0 & 0 & \gamma_{\tilde{g}\tilde{g}}(J) \mathbf{1} \end{pmatrix}$$

# LL Results for Squeezed Limit

$$\begin{aligned}
 & \mathcal{E}(\hat{n}_1)\mathcal{E}(\hat{n}_2)\mathcal{E}(\hat{n}_3) \\
 &= \frac{1}{(2\pi)^2} \frac{2}{\theta_S^2} \frac{2}{\theta_L^2} \vec{\mathcal{J}} \left[ \hat{C}_{\phi_S}(2) - \hat{C}_{\phi_S}(3) \right] \left[ \frac{\alpha_s(\theta_L Q)}{\alpha_s(\theta_S Q)} \right]^{\frac{\hat{\gamma}(3)}{\beta_0}} \left[ \hat{C}_{\phi_L}(3) - \hat{C}_{\phi_L}(4) \right] \left[ \frac{\alpha_s(Q)}{\alpha_s(\theta_L Q)} \right]^{\frac{\hat{\gamma}(4)}{\beta_0}} \vec{\mathcal{O}}^{[4]}(\hat{n}_1)
 \end{aligned}$$

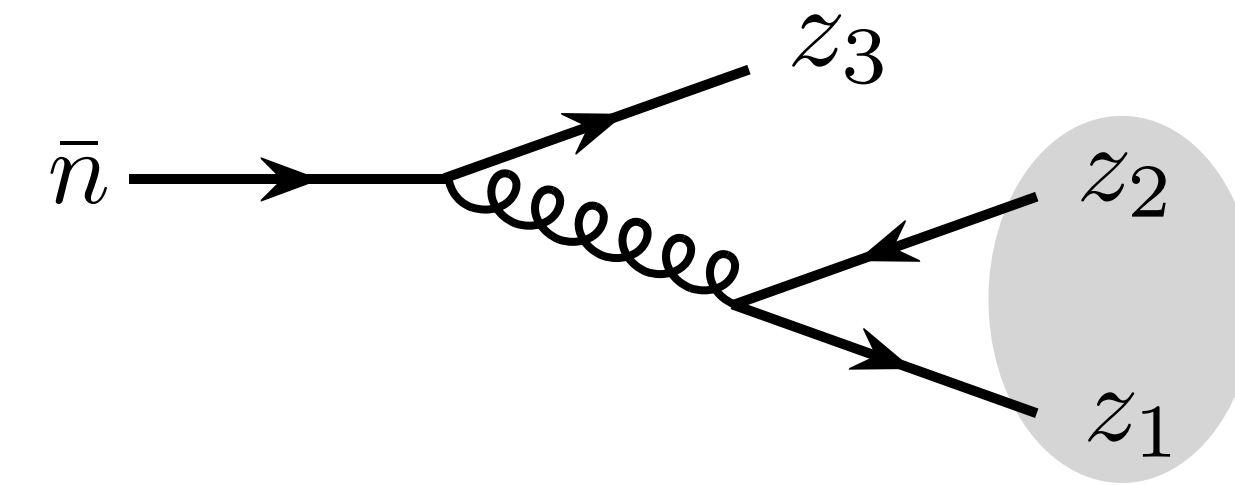


# More Structures...

higher power expansion

For simplicity, tagging final state quarks

Squeezed limit:  $z_1 \cdot z_2 \rightarrow 0$



Expanding the full result:

highest transverse spin series

[HC, Moutl, Sandor, Zhu, forthcoming]

$$g(u, v) \equiv g(z, \bar{z}) \propto -z^3 \bar{z} \cos 2\phi + \frac{39}{10} z^2 \bar{z}^2 - z \bar{z}^3 \quad \text{LP}$$

$$-z^4 \bar{z} \cos 3\phi + \frac{39}{20} z^3 \bar{z}^2 + \frac{39}{20} z^2 \bar{z}^3 - z \bar{z}^4 \quad \text{NLP}$$

$$-\frac{6}{7} z^5 \bar{z} \cos 4\phi + \frac{229}{140} z^4 \bar{z}^2 - \frac{211}{140} z^3 \bar{z}^3 + \frac{229}{140} z^2 \bar{z}^4 - \frac{6}{7} z \bar{z}^5 \quad \text{NNLP}$$

$$-\frac{5}{7} z^6 \bar{z} \cos 5\phi + \frac{207}{140} z^5 \bar{z}^2 - \frac{233}{140} z^4 \bar{z}^3 - \frac{233}{140} z^3 \bar{z}^4 + \frac{207}{140} z^2 \bar{z}^5 - \frac{5}{7} z \bar{z}^6 \quad \text{NNNLP}$$

$$-z^3 \bar{z} {}_2F_1(3, 2, 6, z)$$



Block Structure

Conformal Symmetry  
on the Celestial Sphere

...

...

...

# Summary

- Light-ray operators play an important role in collider physics.
- Light-ray OPE, organized as twist expansion, governs the small angle scaling behavior.
- As an application in QCD, light-ray OPE correctly predicts the perturbative 3-point energy correlator in the squeezed limit and nicely organizes the RG.

**Thanks!**