





- □ The official samples are finally ready!
- $\hfill\square$ So now we will be using MELA as observable.
- □ Show kinematic distribution at the SR; only for VBF related samples;
- \Box And some plots in the CR for $m_{4\ell}$ [130, 220] GeV.
- $\hfill\square$ Will be studying the fit at SR with 13 bins in MELA.
- $\hfill\square$ And the fit with SR+CR, see next slides.



 \Box 220 < $m_{4\ell}$ < 2000 GeV, $N_{
m jets}$ \geq 2, $\Delta\eta(jj)$ > 4.0

 \Box Divide MELA into 13 bins for each channel — namely 4 μ , 4e and 2 μ 2e.

Off-shell couplings measurements: 4ℓ channel Kinematic distribution on the SR



Off-shell couplings measurements: 4ℓ channel

5



Off-shell couplings measurements: 4 ℓ channel Kinematic distribution on the CR



6

 $\Box~$ 130 $< m_{4\ell} <$ 220 GeV, $\textit{N}_{\rm jets} \geq$ 2

- $\hfill\square$ This region will be divided into 8 bins and then added to the SR.
- $\hfill\square$ Currently, we use the MC number in the CR instead of the data.

Off-shell couplings measurements: 4ℓ channel

Likelihood scan vs $\mu_{
m off-shell}$



□ SR+CR: best fit $\mu_{\text{off-shell}} = 1.0^{+4.46}_{-1.1}$ at 95% CLs 8.42. □ SR: best fit $\mu_{\text{off-shell}} = 1.0^{+4.92}_{-1.2}$ at 95% CLs 10.45.

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Off-shell couplings measurements: 4ℓ channel Likelihood scan vs μ_{aazz}



8

□ SR+CR: best fit $\mu_{qqZZ} = 1.0^{+0.1}_{-1.0}$. □ SR: best fit $\mu_{qqZZ} = 1.0^{+0.33}_{-0.22}$.



Including the CR to the SR improve the fit by ~ 19%.
 No systematic is considered in the fit, only luminosity.
 The normalisation of the *qqZZ* background is set free.

To do . . .

□ Adding all the systematic and then try to check the Ranking, pull etc.

□ Then see the background normalisation effect.

Thank you!

