

Constraining nonstandard neutrino interactions at electron colliders

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based on work in

Phys. Rev. D 104, 035043 (2021)

[arXiv:2105.11215]



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Portal to New Physics



Higgs Portal

$$H^\dagger H \phi^\dagger \phi$$

Neutrino Portal

$$H L N_R$$

Vector Portal

$$F_{\mu\nu} X^{\mu\nu}$$

Only three *renormalizable* portals in the Standard Model

Nonstandard Interactions

$$\mathcal{L}^{eff} = \mathcal{L}_{SM} + \frac{c^{d=5}}{\Lambda} \mathcal{O}^{d=5} + \frac{c^{d=6}}{\Lambda^2} \mathcal{O}^{d=6} + \dots$$

Neutrino mass

$$\frac{c^{d=5}}{\Lambda} (\bar{L}_L^c \tilde{\phi}^*) (\tilde{\phi}^\dagger L_L)$$

[P. Minkowski (1977), T. Yanagida (1979), Weinberg (1979)
S.L. Glashow (1980)]



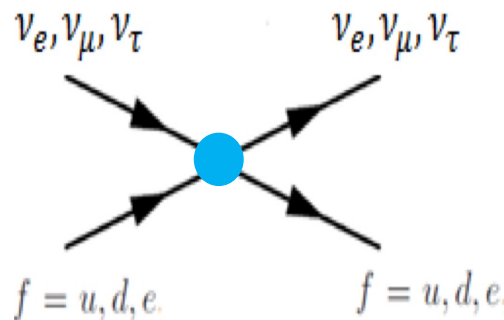
Seesaw mechanism

Nonstandard Interactions (NSI)

$$\mathcal{L}_{\text{NC-NSI}} = -2\sqrt{2}G_F \epsilon_{\alpha\beta}^{fC} [\bar{\nu}_\alpha \gamma^\rho P_L \nu_\beta] [\bar{f} \gamma_\rho P_C f] ,$$

$$\mathcal{L}_{\text{CC-NSI}} = -2\sqrt{2}G_F \epsilon_{\alpha\beta}^{ff'C} [\bar{\nu}_\beta \gamma^\rho P_L \ell_\alpha] [\bar{f}' \gamma_\rho P_C f] ,$$

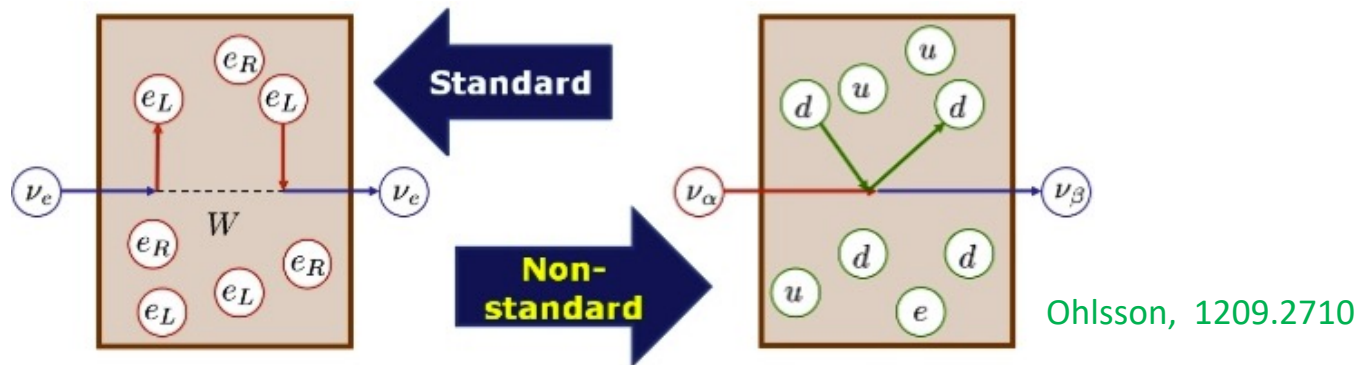
[Wolfenstein, Phys. Rev., D17, 2369 (1978)]



Matter Effects

$$\mathcal{L}_{\text{NSI}} = -2\sqrt{2}G_F \sum_{\alpha\beta fC} \epsilon_{\alpha\beta}^{fC} [\bar{\nu}_{\alpha L} \gamma^\rho \nu_{\beta L}] [\bar{f} \gamma_\rho P_C f] \quad \begin{array}{l} C = L, R, \\ f = u, d, e \end{array}$$

Wolfenstein, Phys. Rev. D 17, 2369 (1978)



Modification of matter potential

$$i \frac{d}{dt} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \frac{1}{2E} \left[U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} U^\dagger + A \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{pmatrix} \right] \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$$

SM

Effective coefficient

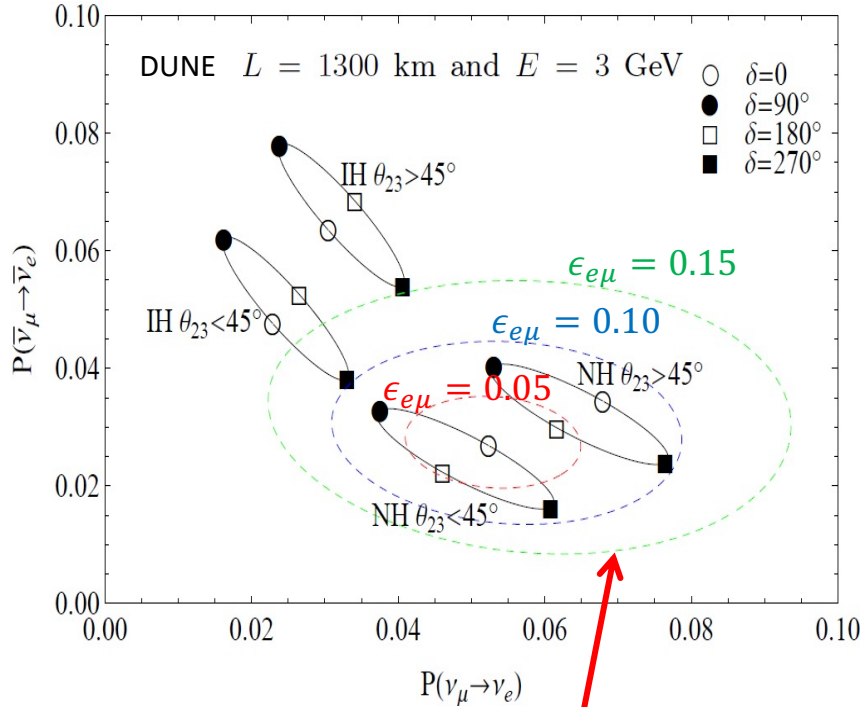
$$\epsilon_{\alpha\beta} \equiv \sum_{f,C} \epsilon_{\alpha\beta}^{fC} \frac{N_f}{N_e}$$

$$A \equiv 2\sqrt{2}G_F N_e E$$

On earth $N_u = N_d = 3N_e$

Oscillation experiments w/ NSI

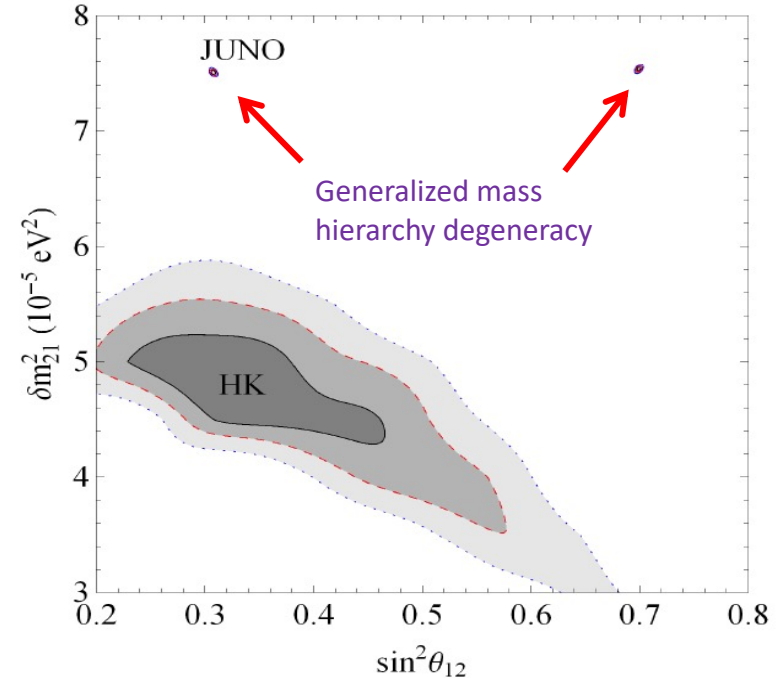
JL, Marfatia, Whisnant 1612.00927



Continuous four-fold degeneracy

The determination of CP violation, θ_{23} octant, and the mass hierarchy at DUNE is strongly affected by NSI.

JL, Marfatia, Whisnant 1704.04711



Bakhti, Farzan 1403.0744

Coloma, Schwetz 1604.05772

$$\delta m_{31}^2 \rightarrow -\delta m_{32}^2, \quad \theta_{12} \rightarrow 90^\circ - \theta_{12}, \quad \delta \rightarrow 180^\circ - \delta,$$

$$\epsilon_{ee} \rightarrow -\epsilon_{ee} - 2, \quad \epsilon_{\alpha\beta} e^{i\phi_{\alpha\beta}} \rightarrow -\epsilon_{\alpha\beta} e^{-i\phi_{\alpha\beta}} \quad (\alpha\beta \neq ee),$$

$$H \rightarrow -H^*$$

Oscillation probabilities are unchanged

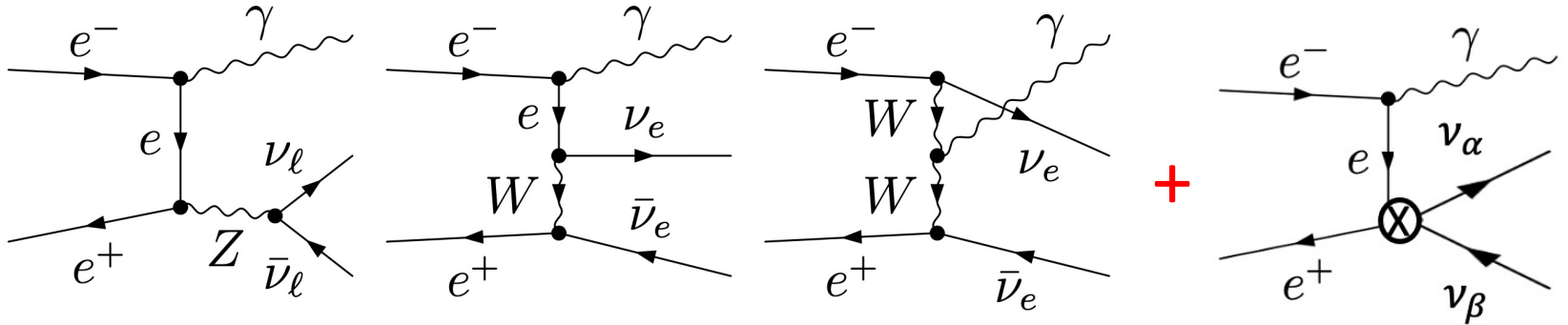
Global constraints on NSI with u,d

	Total Rate	Data Release t+E	Our Fit t+E Chicago	Our Fit t+E Duke
ε_{ee}^u	$[-0.012, +0.621]$	$[+0.043, +0.384]$	$[-0.032, +0.533]$	$[-0.004, +0.496]$
$\varepsilon_{\mu\mu}^u$	$[-0.115, +0.405]$	$[-0.050, +0.062]$	$[-0.094, +0.071] \oplus [+0.302, +0.429]$	$[-0.045, +0.108] \oplus [+0.290, +0.399]$
$\varepsilon_{\tau\tau}^u$	$[-0.116, +0.406]$	$[-0.050, +0.065]$	$[-0.095, +0.125] \oplus [+0.302, +0.428]$	$[-0.045, +0.141] \oplus [+0.290, +0.399]$
$\varepsilon_{e\mu}^u$	$[-0.059, +0.033]$	$[-0.055, +0.027]$	$[-0.060, +0.036]$	$[-0.060, +0.034]$
$\varepsilon_{e\tau}^u$	$[-0.250, +0.110]$	$[-0.141, +0.090]$	$[-0.243, +0.118]$	$[-0.222, +0.113]$
$\varepsilon_{\mu\tau}^u$	$[-0.012, +0.008]$	$[-0.006, +0.006]$	$[-0.013, +0.009]$	$[-0.012, +0.009]$
ε_{ee}^d	$[-0.015, +0.566]$	$[+0.036, +0.354]$	$[-0.030, +0.468]$	$[-0.006, +0.434]$
$\varepsilon_{\mu\mu}^d$	$[-0.104, +0.363]$	$[-0.046, +0.057]$	$[-0.083, +0.077] \oplus [+0.278, +0.384]$	$[-0.037, +0.099] \oplus [+0.267, +0.356]$
$\varepsilon_{\tau\tau}^d$	$[-0.104, +0.363]$	$[-0.046, +0.059]$	$[-0.083, +0.083] \oplus [+0.279, +0.383]$	$[-0.038, +0.104] \oplus [+0.268, +0.354]$
$\varepsilon_{e\mu}^d$	$[-0.058, +0.032]$	$[-0.052, +0.024]$	$[-0.059, +0.034]$	$[-0.058, +0.034]$
$\varepsilon_{e\tau}^d$	$[-0.198, +0.103]$	$[-0.106, +0.082]$	$[-0.196, +0.107]$	$[-0.181, +0.101]$
$\varepsilon_{\mu\tau}^d$	$[-0.008, +0.008]$	$[-0.005, +0.005]$	$[-0.008, +0.008]$	$[-0.007, +0.008]$
ε_{ee}^p	$[-0.035, +2.056]$	$[+0.142, +1.239]$	$[-0.095, +1.812]$	$[-0.024, +1.723]$
$\varepsilon_{\mu\mu}^p$	$[-0.379, +1.402]$	$[-0.166, +0.204]$	$[-0.312, +0.138] \oplus [+1.036, +1.456]$	$[-0.166, +0.337] \oplus [+0.952, +1.374]$
$\varepsilon_{\tau\tau}^p$	$[-0.379, +1.409]$	$[-0.168, +0.257]$	$[-0.313, +0.478] \oplus [+1.038, +1.453]$	$[-0.167, +0.582] \oplus [+0.950, +1.382]$
$\varepsilon_{e\mu}^p$	$[-0.179, +0.112]$	$[-0.174, +0.086]$	$[-0.179, +0.120]$	$[-0.187, +0.131]$
$\varepsilon_{e\tau}^p$	$[-0.877, +0.340]$	$[-0.503, +0.295]$	$[-0.841, +0.355]$	$[-0.817, +0.386]$
$\varepsilon_{\mu\tau}^p$	$[-0.041, +0.025]$	$[-0.020, +0.019]$	$[-0.044, +0.026]$	$[-0.048, +0.030]$

Table 2. 2σ allowed ranges for the NSI couplings $\varepsilon_{\alpha\beta}^u$, $\varepsilon_{\alpha\beta}^d$ and $\varepsilon_{\alpha\beta}^p$ as obtained from the global analysis of oscillation plus COHERENT data. See text for details.

[Coloma et.al., 1911.09109](#)

NSI@electron collider



$$e^+e^- \rightarrow \nu\bar{\nu}\gamma \quad \frac{d^2\sigma}{dx_\gamma dz_\gamma} = H(x_\gamma, z_\gamma; s) \sigma_0(s_\gamma)$$

Nicrosini, Trentadue, PLB 231, 487 (1989)

Altarelli-Parisi radiator function $H(x_\gamma, z_\gamma; s) = \frac{\alpha}{\pi} \frac{1}{x_\gamma} \left[\frac{1 + (1 - x_\gamma)^2}{1 - z_\gamma^2} - \frac{x_\gamma^2}{2} \right]$

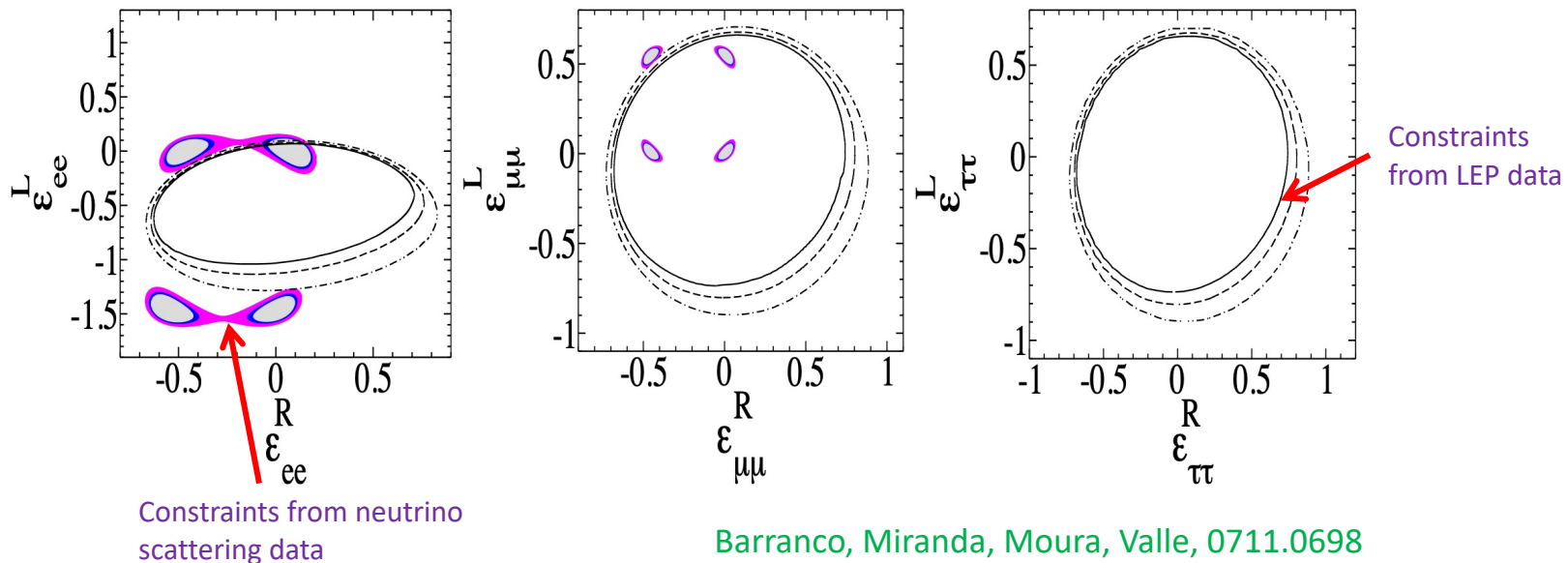
$$e^+e^- \rightarrow \nu\bar{\nu} \quad \sigma_0^{\text{SM}}(s) = \frac{N_\nu G_F^2}{6\pi} M_Z^4 (g_L^2 + g_R^2) \frac{s}{[(s - M_Z^2)^2 + (M_Z \Gamma_Z)^2]} \\ + \frac{G_F^2}{\pi} M_W^2 \left\{ \frac{s + 2M_W^2}{2s} - \frac{M_W^2}{s} \left(\frac{s + M_W^2}{s} \right) \log \left(\frac{s + M_W^2}{M_W^2} \right) \right. \\ \left. - g_L \frac{M_Z^2 (s - M_Z^2)}{(s - M_Z^2)^2 + (M_Z \Gamma_Z)^2} \times \left[\frac{(s + M_W^2)^2}{s^2} \log \left(\frac{s + M_W^2}{M_W^2} \right) - \frac{M_W^2}{s} - \frac{3}{2} \right] \right\}$$

Electron NSI

Berezhniana , Rossi, Phys.Lett.B 535 (2002)

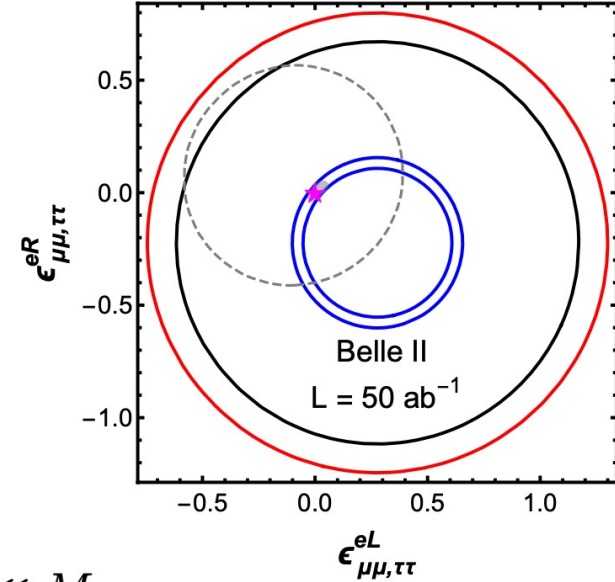
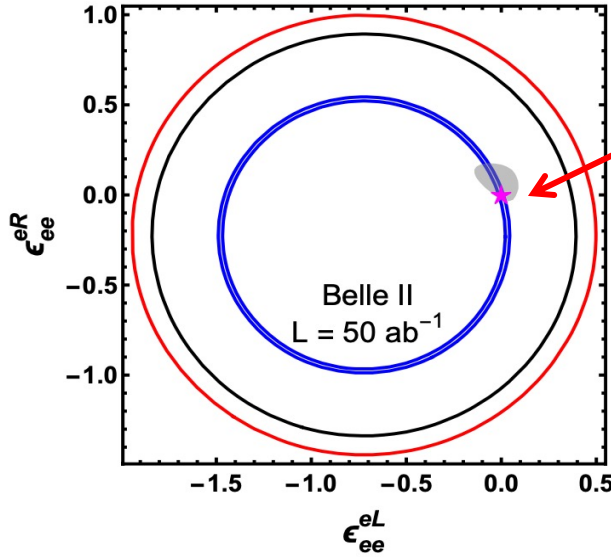
$$\sigma_0^{\text{NSI}}(s) = \sum_{\alpha,\beta=e,\mu,\tau} \frac{G_F^2}{6\pi} s \left[((\epsilon_{\alpha\beta}^{eL})^2 + (\epsilon_{\alpha\beta}^{eR})^2) - 2 (g_L \epsilon_{\alpha\beta}^{eL} + g_R \epsilon_{\alpha\beta}^{eR}) \frac{M_Z^2 (s - M_Z^2)}{(s - M_Z^2)^2 + (M_Z \Gamma_Z)^2} \right] \\ + \frac{G_F^2}{\pi} \epsilon_{ee}^{eL} M_W^2 \left[\frac{(s + M_W^2)^2}{s^2} \log \left(\frac{s + M_W^2}{M_W^2} \right) - \frac{M_W^2}{s} - \frac{3}{2} \right].$$

12 independent NSI parameters



Barranco, Miranda, Moura, Valle, 0711.0698

NSI@ Belle II



$$e^+e^- \rightarrow \nu\bar{\nu}\gamma \quad \text{with } \sqrt{s} \ll M_Z$$

$$\frac{d\sigma^{\text{SM}}}{dx_\gamma dz_\gamma} = H(x_\gamma, z_\gamma, s) \frac{G_F^2 s_\gamma}{2\pi} C^{\text{SM}}$$

$$C^{\text{SM}} \equiv g_L^2 + g_R^2 + \frac{2}{3}g_L + \frac{1}{3}$$

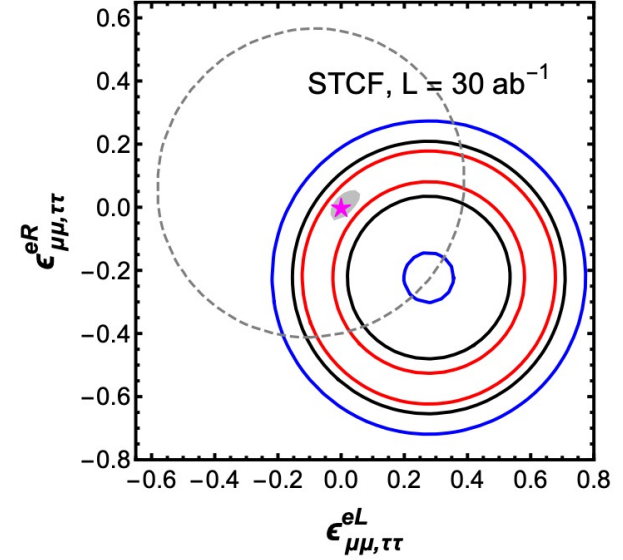
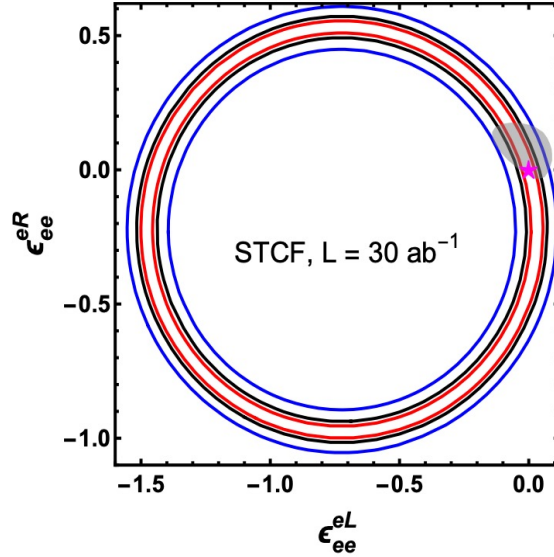
$$\frac{d\sigma^{\text{NSI}}}{dx_\gamma dz_\gamma} = H(x_\gamma, z_\gamma, s) \frac{G_F^2 s_\gamma}{2\pi} C^{\text{NSI}}$$

$$C^{\text{NSI}} \equiv \frac{1}{3} \sum_{\alpha,\beta=e,\mu,\tau} \left[(\epsilon_{\alpha\beta}^{eL})^2 + (\epsilon_{\alpha\beta}^{eR})^2 + 2(g_L \epsilon_{\alpha\beta}^{eL} + g_R \epsilon_{\alpha\beta}^{eR}) \right] + \frac{2}{3} \epsilon_{ee}^{eL}$$

$$\sigma = \sigma^{\text{SM}} + \sigma^{\text{NSI}}$$

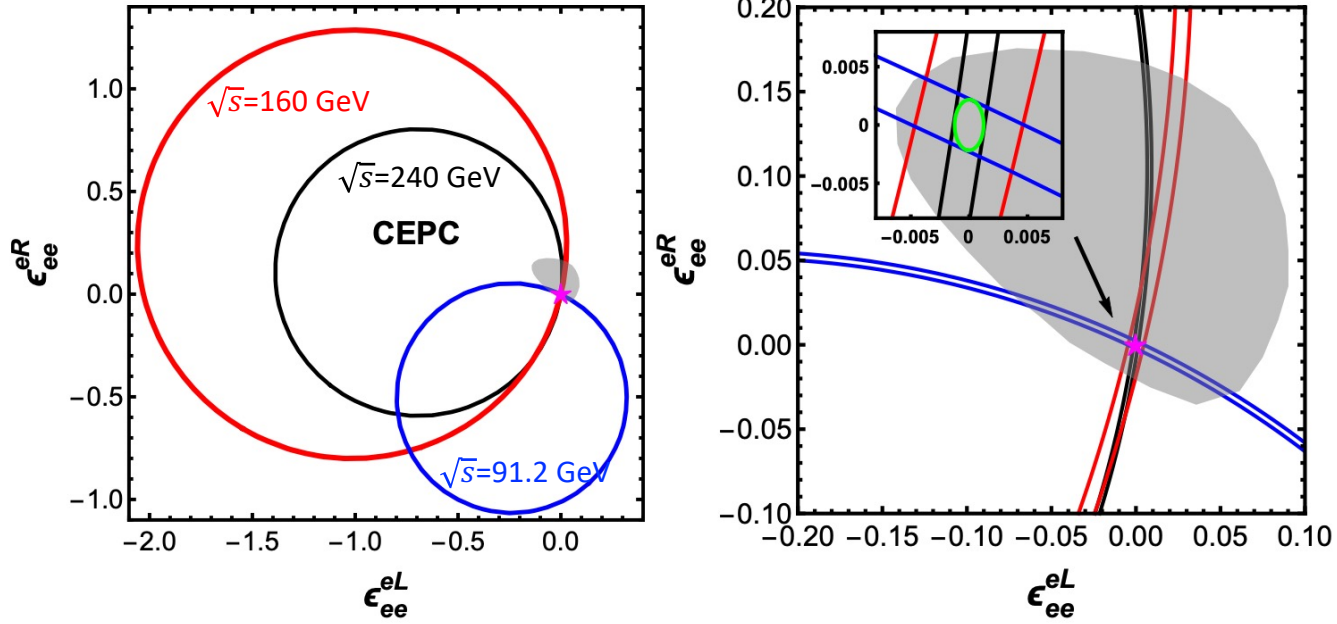
Iso- σ contours form concentric circles with center at $(\epsilon_{\alpha\alpha}^{eL}, \epsilon_{\alpha\alpha}^{eR}) = (-g_L - \delta_{\alpha e}, -g_R)$

NSI@STCF



	STCF-2 $L = 30 \text{ ab}^{-1}$	STCF-4 $L = 30 \text{ ab}^{-1}$	STCF-7 $L = 30 \text{ ab}^{-1}$	Belle II $L = 50 \text{ ab}^{-1}$	Previous Limit 90% Allowed [45]
ϵ_{ee}^{eL}	$[-0.067, 0.061]$	$[-0.033, 0.031]$	$[-0.018, 0.018]$	$[-0.0091, 0.0089]$	$[-0.03, 0.08]$
ϵ_{ee}^{eR}	$[-0.60, 0.15]$	$[-0.163, 0.087]$	$[-0.070, 0.053]$	$[-0.031, 0.028]$	$[0.004, 0.15]$
$\epsilon_{\mu\mu/\tau\tau}^{eL}$	$[-0.13, 0.69]$	$[-0.073, 0.101]$	$[-0.044, 0.052]$	$[-0.023, 0.025]$	$[-0.03, 0.03]/[-0.5, 0.2]$
$\epsilon_{\mu\mu/\tau\tau}^{eR}$	$[-0.60, 0.15]$	$[-0.163, 0.087]$	$[-0.070, 0.053]$	$[-0.031, 0.028]$	$[-0.03, 0.03]/[-0.3, 0.4]$

NSI@CEPC



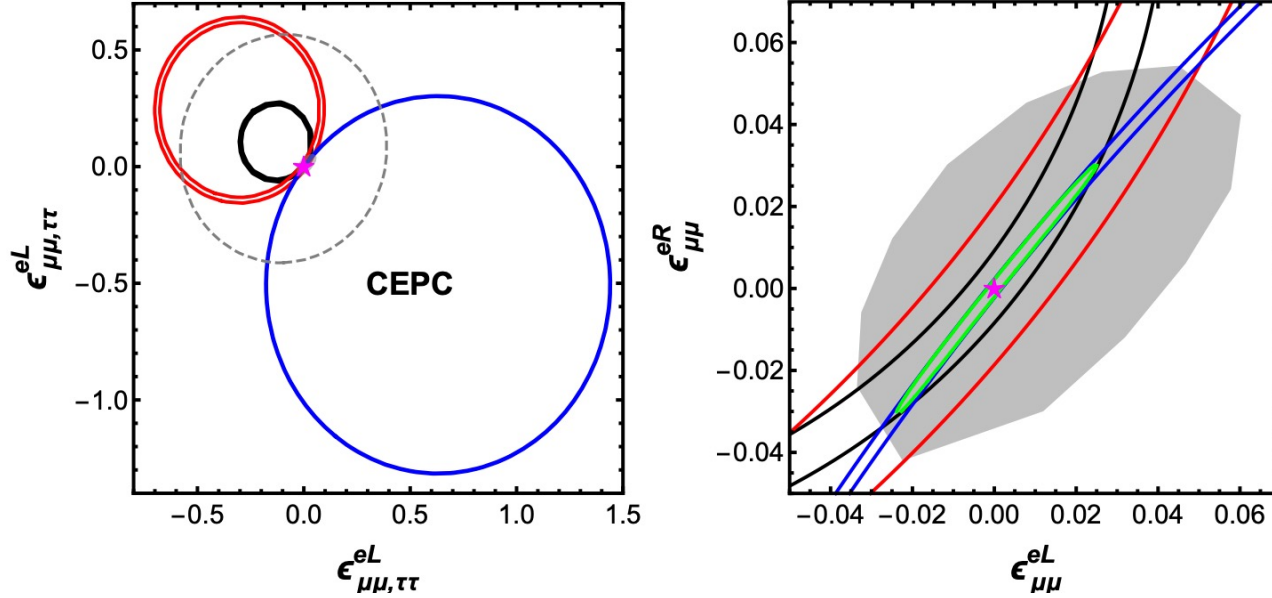
$$\sqrt{s} \geq M_Z \quad \sigma^{\text{NSI}}(\epsilon_{ee}^{eL}, \epsilon_{ee}^{eR}) = I_1 \left((\epsilon_{ee}^{eL})^2 + (\epsilon_{ee}^{eR})^2 \right) + (I_2 + I_3) \epsilon_{ee}^{eL} + I_2 \frac{g_R}{g_L} \epsilon_{ee}^{eR}$$

$$I_i \equiv \int dx \int dz_\gamma H(x_\gamma, z_\gamma, s) [\sigma_0^{\text{NSI}}(s_\gamma)]_i \quad \text{is a function of } \sqrt{s}$$

Allowed regions lie between two concentric circles with the center at

$$(\epsilon_{ee}^{eL}, \epsilon_{ee}^{eR}) = \left(-\frac{I_2 + I_3}{2I_1}, -\frac{I_2 g_R}{2I_1 g_L} \right)$$

NSI@CEPC



	CEPC-91.2 $L = 16 \text{ ab}^{-1}$	CEPC-160 $L = 2.6 \text{ ab}^{-1}$	CEPC-240 $L = 5.6 \text{ ab}^{-1}$	CEPC-combined $L = 24.2 \text{ ab}^{-1}$	Previous Limit 90% Allowed [45]
ϵ_{ee}^{eL}	[-0.0037,0.0037]	[-0.0036,0.0035]	[-0.0010,0.0010]	[-0.00095,0.00095]	[-0.03,0.08]
ϵ_{ee}^{eR}	[-0.0017,0.0017]	[-0.014,0.015]	[-0.0065,0.0070]	[-0.0017,0.0017]	[0.004,0.15]
$\epsilon_{\mu\mu/\tau\tau}^{eL}$	[-0.0014,0.0014]	[-0.012,0.012]	[-0.0055,0.0053]	[-0.0013,0.0013]	[-0.03,0.03]/[-0.5,0.3]
$\epsilon_{\mu\mu/\tau\tau}^{eR}$	[-0.0017,0.0017]	[-0.014,0.015]	[-0.0065,0.0070]	[-0.0017,0.0017]	[-0.03,0.03]/[-0.3,0.4]

Summary

- Neutrino oscillation experiments are strongly affected by NSI, it is natural to seek complementary constraints on NSI from other type experiments.
- Both Belle II and STCF can provide competitive and complementary bounds on $\varepsilon_{ee}^{eL,R}$ as compared to current global analysis, and strong improvements in the constraints on $\varepsilon_{\tau\tau}^{eL,R}$.
- CEPC running with three CM energies $\sqrt{s} = 240, 160, 91.2$ GeV will break the degeneracy between the left- and right-handed NSI, and can reach a sub-percent level sensitivity on electron NSI.

Thank you!

Backup slides

Coordinates of the circle centers

