### Fast & rigorous constraints on chiral threenucleon forces from few-body observables

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# The importance of the three-nucleon force in Chiral Effective Field Theory



- systems and symmetric nuclear matter

In EFTs three-body forces inevitably arise because degrees of freedom are integrated out

In  $\chi$ EFT without an explicit  $\Delta(1232)$  three-nucleon forces (3NFs) appear at  $O(Q^3)$  (NNLO)



van Kolck (1994); Epelbaum et al. (2002)

CE

Depends on two parameters,  $c_D$  and  $c_E$ , once  $\pi N$  low-energy constants (LECs) are fixed

This 3NF has small but important effects in light nuclei and helps drive saturation in heavier

<u>Goal here</u>: estimate  $c_D$  and  $c_E$  from few-nucleon data including  $\chi$ EFT truncation error

### Few-nucleon physics implementation

- No-Core Shell Model calculations of A=3 and A=4 bound-state observables
- Binding energy of <sup>3</sup>H, <sup>4</sup>He, Charge radius of <sup>4</sup>He, β-decay half-life of <sup>3</sup>H, aka "GT matrix element"
- Fully converged for A=4 with  $\hbar\omega = 36$  MeV, N<sub>max</sub>=18 due to use of relatively soft interaction
- New fit to np and pp scattering data for  $0 \le E_{lab} \le 290$  MeV. Truncation error included
- πN LECs fixed at central values of Roy-Steiner analysis of Siemens et al.

	LO	NLO	$\langle \mathrm{NNLO} \rangle_{\mathrm{ppd}}$	Experiment	Adopted uncertainty
$E(^{3}\mathrm{H})$ [MeV]	-5.65	-8.38	-8.52	-8.482 [40]	0.015
$E(^{4}\text{He})$ [MeV]	-24.08	-30.21	-28.19	-28.296 [41]	0.005
$r(^{4}\text{He})$ [fm]	1.27	1.33	1.45	1.4552(62) [42]	0.0062
$fT_{1/2}$ [s]			1127.3	1129.6(3.0) [43]	3.0



### 3N error model

$$y_{exp} = y_{th} + \delta y_{exp} + \delta y_{th}$$

$$y_{th} = y_{ref} \sum_{i=0}^{k} c_i(\{a_i\})Q^i \qquad Q = \frac{p_{typ}}{\Lambda_b} \qquad y_{ref} = y_{LO} \text{ here}$$

$$c_i\text{'s are Gaussian random variables with mean zero \Rightarrow \delta y_{th} = y_{ref} \bar{c} \frac{Q^{k+1}}{\sqrt{1-Q^2}}$$

Assume

- mean-square value of the higher-order coefficients
- $\bar{c}^2$  and Q are also constrained by information from the lower-order calculations

• Q is not obvious: we will actually make it a parameter and sample it. We will also sample  $\bar{c}^2$ , the

As a first go we will take the uncertainties in the different observables to be uncorrelated



 $\operatorname{pr}(c_D, c_E, \overline{c}^2, Q | D, I) \propto \exp\left(-\frac{1}{2}\mathbf{r}^T(\mathbf{\Sigma}_{\exp} + \mathbf{\Sigma}_{\exp})\right)$ 

$$\mathbf{E}_{\text{th}}^{-1}\mathbf{r} \exp\left(-\frac{c_D^2 + c_E^2}{2\bar{a}^2}\right) \operatorname{pr}(\bar{c}^2 | Q, \bar{a}, I) \operatorname{pr}(Q | c_D, c_E, I)$$
$$\mathbf{r} = \mathbf{y}_{\exp} - \mathbf{y}_{\text{th}}$$

 $\operatorname{pr}(c_D, c_E, \bar{c}^2, Q \mid D, I) \propto \exp\left(-\frac{1}{2}\mathbf{r}^T (\boldsymbol{\Sigma}_{\exp} + (\boldsymbol{\Sigma}_{th})^{-1}\mathbf{r})\right) \exp\left(-\frac{c_D^2 + c_E^2}{2\bar{a}^2}\right) \operatorname{pr}(\bar{c}^2 \mid Q, \bar{a}, I) \operatorname{pr}(Q \mid c_D, c_E, I)$ **Truncation errors**  $\mathbf{r} = \mathbf{y}_{exp} - \mathbf{y}_{th}$ 

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#### Truncation errors

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$$(\boldsymbol{\Sigma}_{\text{th,uncorr}})_{ij} = (\mathbf{y}_{\text{ref}})^2 \bar{c}^2 \delta_{ij} \sum_{n=k+1}^{\infty} Q^{2n}$$
; experiment

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Can include NN in "fit" by expanding meaning of  $\vec{a}$  to include NN parameters. Incorporate NN information by using posterior from that analysis as a prior on  $\vec{a}_{NN}$ , the NN piece of  $\vec{a}$ , here

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•  $pr(Q | \mathbf{a}, I)$  then also affected by that information. Starts as weakly informative Beta distribution.

# Emulation using Eigenvector Continuation

- We use Eigenvector Continuation to emulate few-nucleon observables
- Emulator is built in 11-dimensional parameter space:  $\overrightarrow{a}$ includes  $c_D$  and  $c_E$  and 9 NN parameters at NNLO: " $\vec{a}_{NN}$ "
- Solve  $H(\vec{a}) | \psi(\vec{a}) \rangle = E(\vec{a}) | \psi(\vec{a}) \rangle$  at N<sub>EC</sub> points in parameter space
- Project  $H(\vec{a})$  onto subspace spanned by these N<sub>EC</sub> wave functions; solve generalized eigenvalue problem in subspace
- Eigenvector at  $\overrightarrow{a}$  obtained as linear combination of N<sub>EC</sub> vectors in subspace. Denote coefficients of linear combination by  $\beta(\vec{a})$
- Observables at  $\vec{a}$  then reconstructed from  $\beta(\vec{a})$  and projection of observable to subspace

Frame et al., Phys. Rev. Lett. 121, 032501 (2018); König et al., Phys. Lett. B 810, 135814 (2020)







### Emulation results

- First EC emulation of transition matrix element
- Eigenvector continuation with N<sub>EC</sub>=50 training points is very accurate for all observables considered.



• We obtain the I3-d posterior  $pr(\vec{a}_{NN}, c_D, c_E, \bar{c}, Q | D, I)$ 

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- Not Gaussian! t-distribution with v=2.6 degrees of freedom















Truncation errors essential to get consistency

- Constraints other than GT matrix
   element essentially
   degenerate
- Posterior almost obtained with just fT<sub>1/2</sub> and E(<sup>4</sup>He)



# Results for Q, $\bar{c}$

- $pr(Q \mid \overrightarrow{a}, I)$  starts as a weakly informative Beta distribution
- Can use information on size of  $c_3$  (NNLO-NLO shift) to update Q posterior (and  $\bar{c}$  posterior too)
- But then we sample the likelihood (times the prior) allowing size of theory errors (i.e.  $\bar{c}$  and Q) to vary
- Q and  $\bar{c}$  then include information on how far NNLO result is from experiment
- $\bar{c}$  natural; Q = 0.33(6)





# Posterior predictive distribution

Chiral Effective Field Theory can describe all these data at NNLO [O(Q<sup>3</sup>)] once truncation errors are accounted for

These are also t-distributions



### Linear models with variance estimation: time for t

- This problem is linear in  $c_D$  and  $c_E$  in the region that matters for the final posterior
- So posterior for  $c_D$  and  $c_E$  would be Gaussian if we kept  $\overline{c}$  and Q fixed
- By sampling  $\bar{c}$  and Q we end up doing "variance estimation" in our statistical model
- If 7⁄

$$= \bar{c}^2 Q^{2(k+1)} \text{ we have (approximately):}$$
  

$$\operatorname{pr}(\vec{a} \mid D, \mathcal{V}) \propto \frac{1}{\sqrt{V}} \exp\left(-\frac{(y_{\exp} - y_{\operatorname{th}}(\vec{a}))^2}{2\mathcal{V}}\right) \qquad \operatorname{pr}(\mathcal{V}) \propto \frac{1}{\mathcal{V}^{n/2+1}} \exp\left(-\frac{ns^2}{2\mathcal{V}}\right)$$

• Since  $y_{th}(\overrightarrow{a})$  is linear in  $\overrightarrow{a}$  marginalizing over  $\mathcal{V}$  then yields a t-distribution



# Summary and Outlook

- Part of an ongoing effort to develop, apply, and evaluate Bayesian statistical methods for EFTs of nuclei See also: P. Maris et al., Phys. Rev. C (2021)
- Truncation errors are included in extraction of c<sub>D</sub> and c<sub>E</sub> from few-nucleon observables
- Parameters of statistical model of truncation errors estimated simultaneously:  $Q = 0.33(6), \bar{c} \in [0.87, 1.44]$
- The LECs cD and CE are strongly correlated. Joint pdf best represented by a multivariate t distribution
- For 3NF parameter estimation you should not only use observables that are related by universality Lupu, Barnea, Gazit, arXiv:1508.05654
- Impact of NN uncertainties in the posterior is small; that of πN uncertainties remains to be assessed
- Future work: comparing  $\chi$ EFT Hamiltonians with different regulators and with  $\Delta(1232)$  degrees of freedom; different assumptions for correlations of theory uncertainties
- Extending the results is straightforward via open-source Python package <u>fit3bf</u>

See also: K. Kravvaris et al., Phys. Rev. C (2020)

