

Sci-W ECAL Status for CEPC

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On behalf of CEPC Calorimeter working group



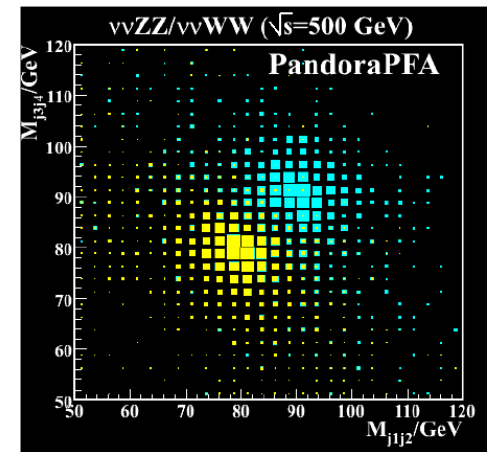
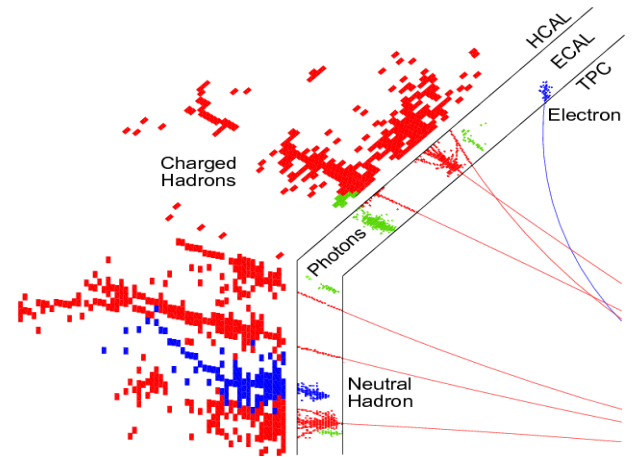
Outline

- Brief review of Sci-W ECAL of CEPC
- CEPC Sci-ECAL Status
 - LED test batch analysis
 - MIPs spectra batch analysis
 - DAC calibration batch analysis
 - Low gain and high ratio
 - Stability of Sci-ECAL
- Summary and outlook



PFA Calorimeter

- Good separation of particles
- Compact showers to minimize overlap
 - Small moliere radius
- Minimum amount dead material
 - inside the magnet coil
- Detailed information of showers
 - High granularity

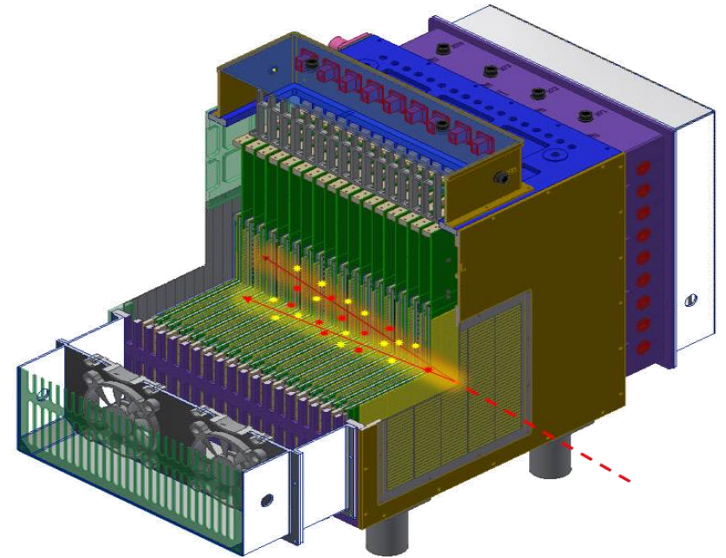


An Imaging calorimeter could meet the requirements



Sci-W PFA ECAL of CEPC

- A sampling calorimeter based on plastic scintillator and W absorber was built last year
 - 32 layers, ~ 22 r.l
 - Each layer has 210 channels, total number of readout is 6720
 - Layer area: $226\text{mm} \times 226\text{ mm}$
 - Cell size: $5\text{mm} \times 45\text{mm} \times 2\text{mm}$
 - power consumption: 4.8 w/layer



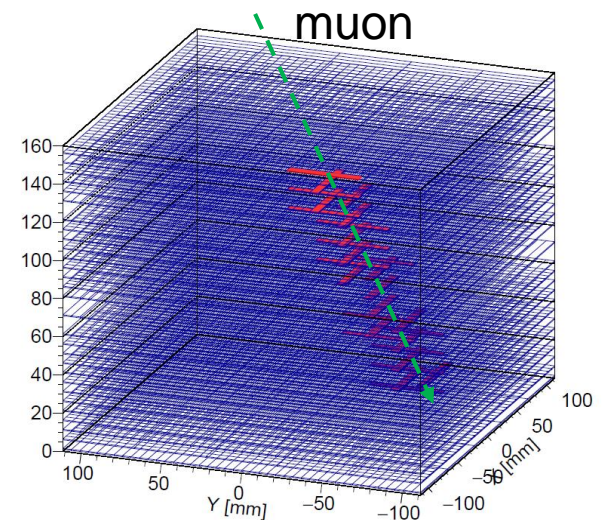
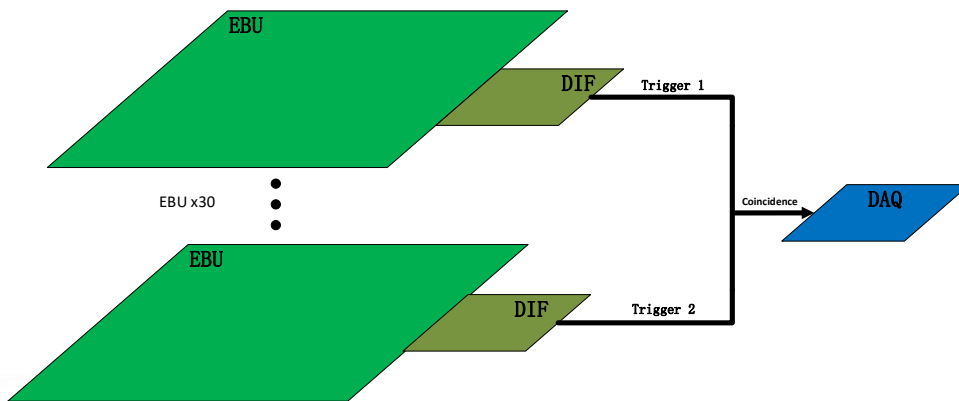
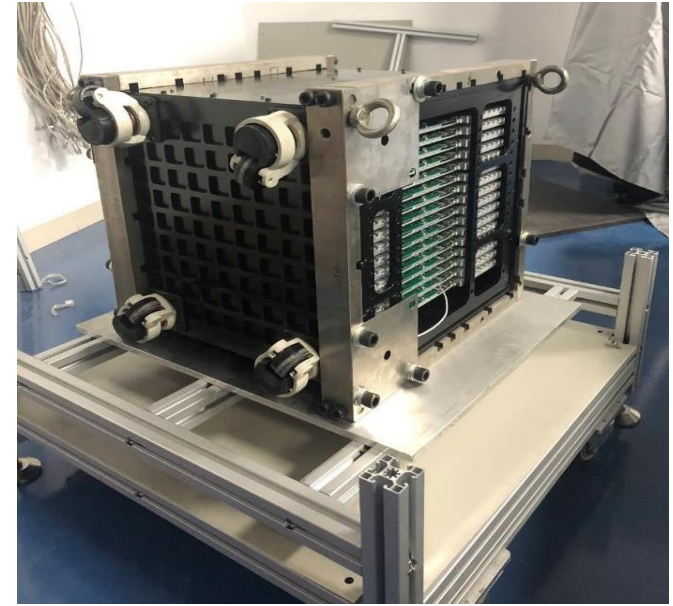
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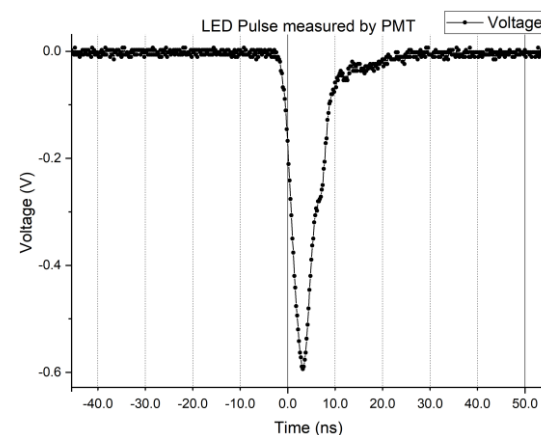
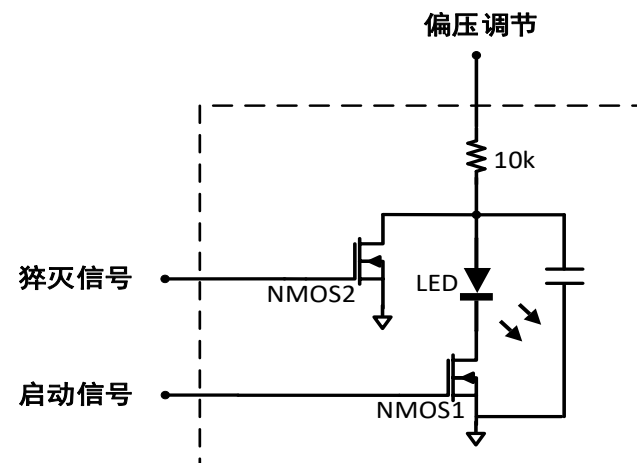
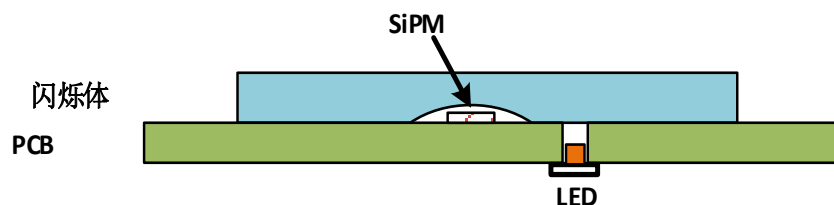
Cosmic Ray test

- Long term cosmic ray test: ~3 months
 - ScECAL has been rotated by 90 degree
 - Coincidence trigger of Layer1 & Layer29
 - Event rate : ~ 16 per minute
 - ~1.5 million cosmic ray events collected
- Purpose
 - Function verification (stability, temperature correction, etc)
 - EBU efficiency and Position resolution
 - Cell-to-cell MIP calibration



LED test

- The S.PE is a very important parameter to monitor the stability of the SiPM gain
- A circuit was designed to drive LED to calibrate S.PE of SiPM, also it could be used to test the ratio of low gain and high channels

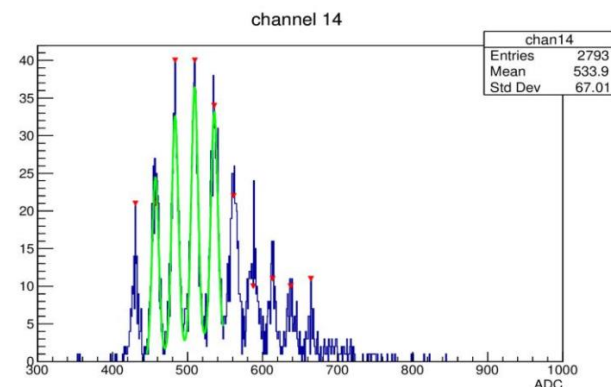
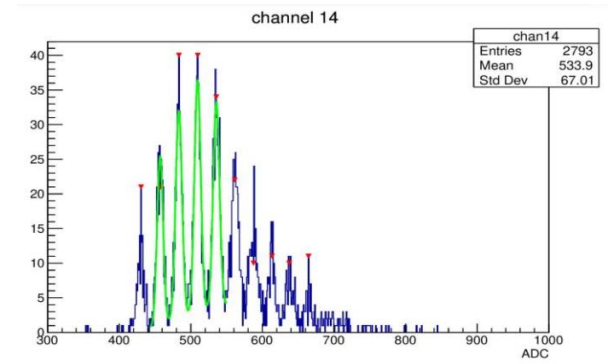
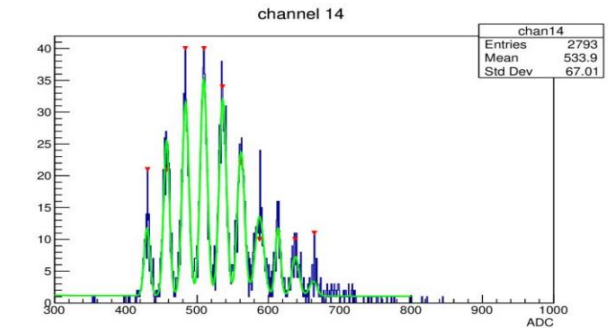


LED light spectrum

How to batch fit the LED calibration results

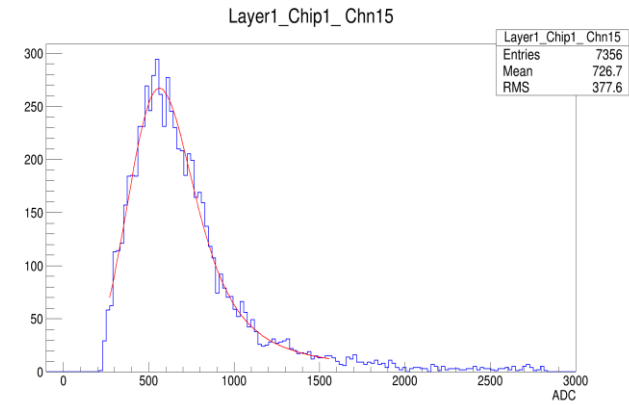
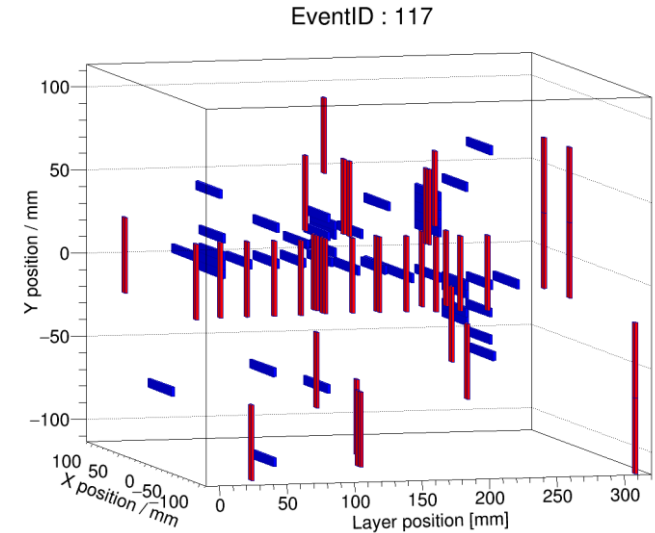
- Multi-gaussian fit
- Three different fitting methods
 - All the parameters are free
 - Considering the correlation of photoelectric peak position
 - It is approximately considered that the width of photoelectric peaks are consistent

From the total number of channels successfully fitted, we selected the 3rd method



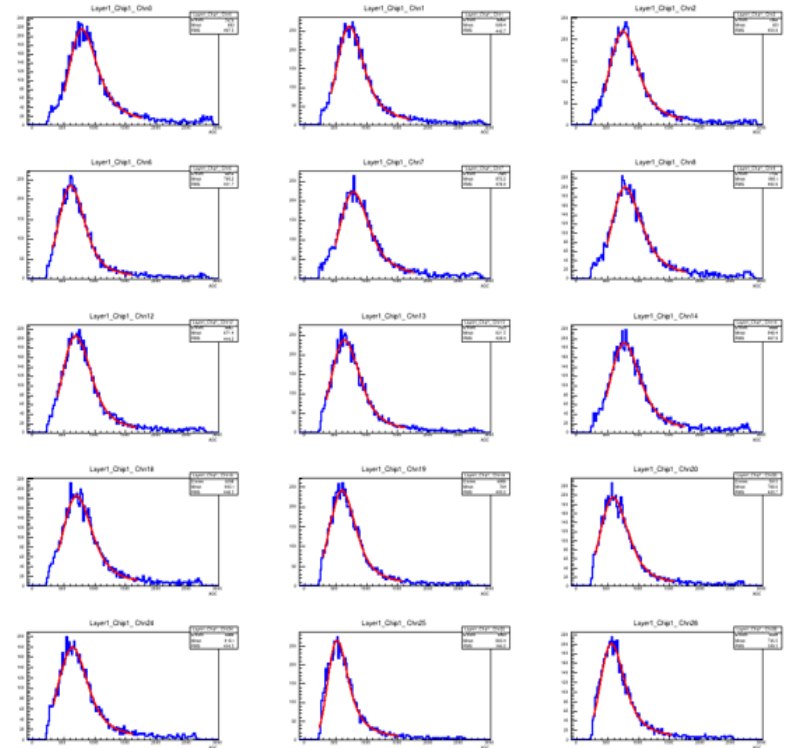
MIPs spectra

- In order to reconstruct the total energy deposition in calorimeter of incident event, we should know the deposition in each SD element
- MPV value of MIPs is the reference for SD energy reconstruction
- Landau convolution Gaussian function is used to fit
 - Landau describes the energy fluctuation
 - Gauss describes the fluorescence process, electronic gain and so on



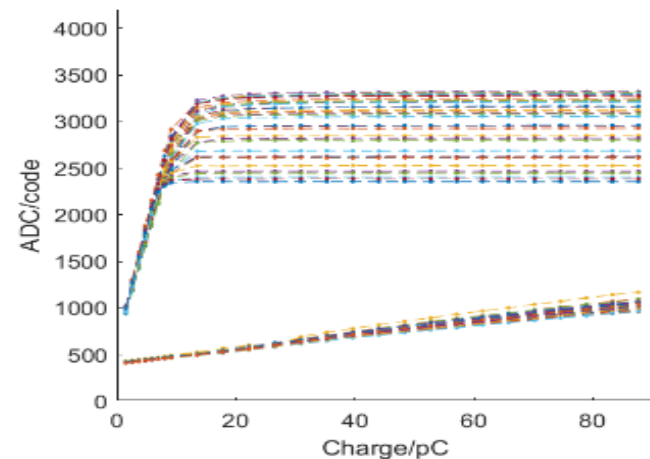
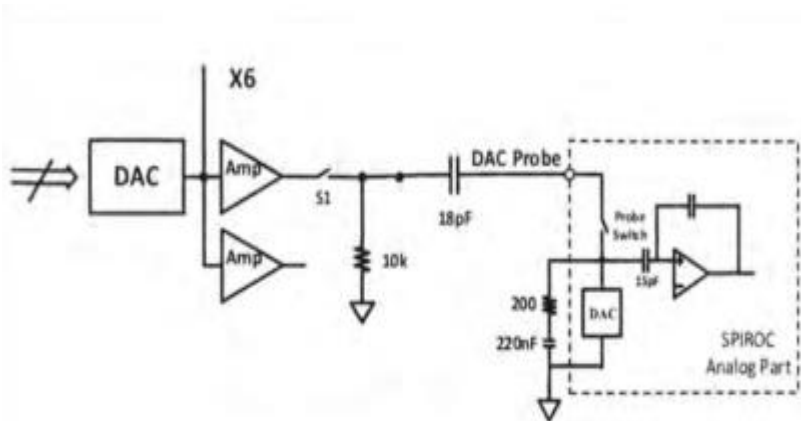
How to batch fit the MIPs spectra

- Because the signal amplitude of each channel is different, it introduces trouble for channel by channel fitting
- We put all channels of one layer into a hist for fitting, and let the parameters free.
- Then, the parameters obtained by fitting are sent to these channels for the second step of fitting



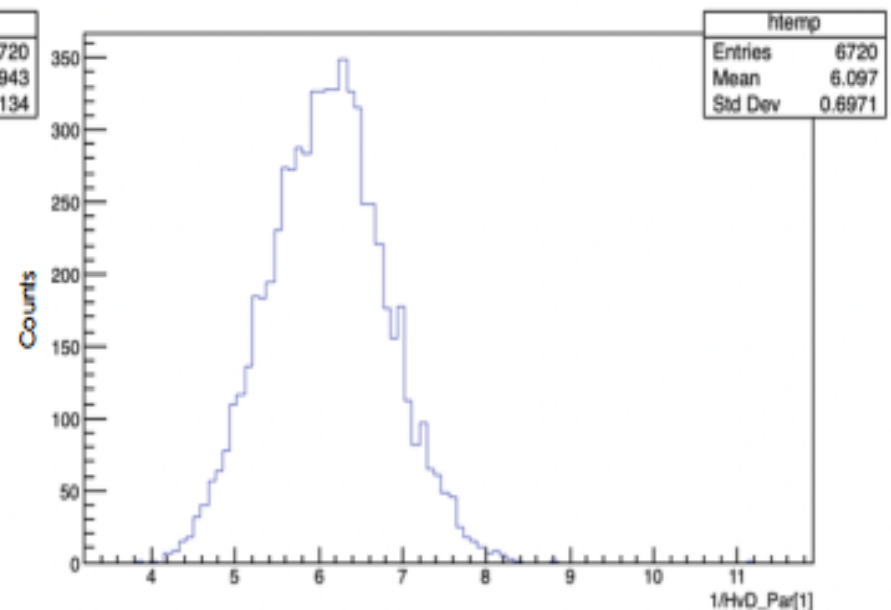
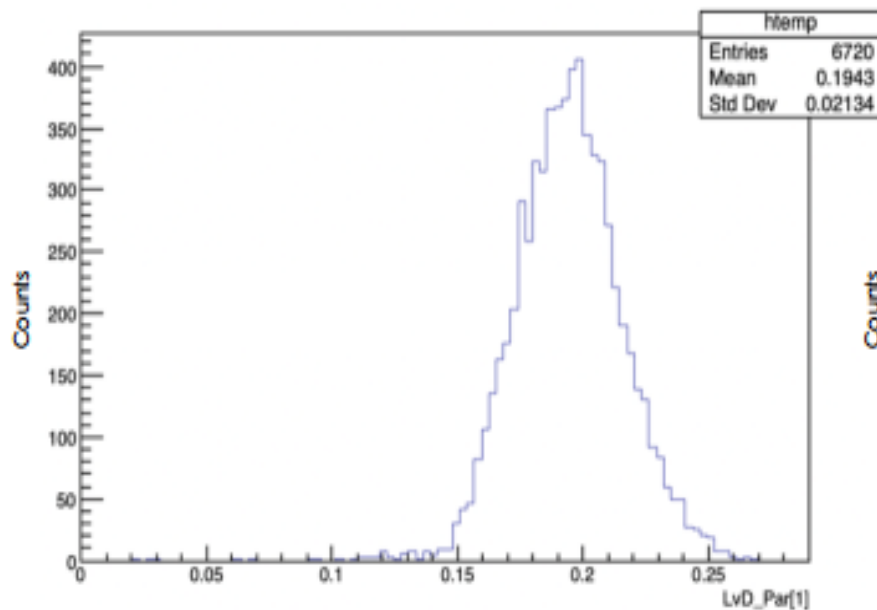
DAC Calibration

- Digital-to-Analog Conversion (0 – 5 V) driver was designed to test the linearity of readout channels
- The input signal waveform is similar to SiPM signal, $\tau \sim 3.5 \text{ ns}$
- One SP2E chip has one probe pin, can select to connect with 36 channels



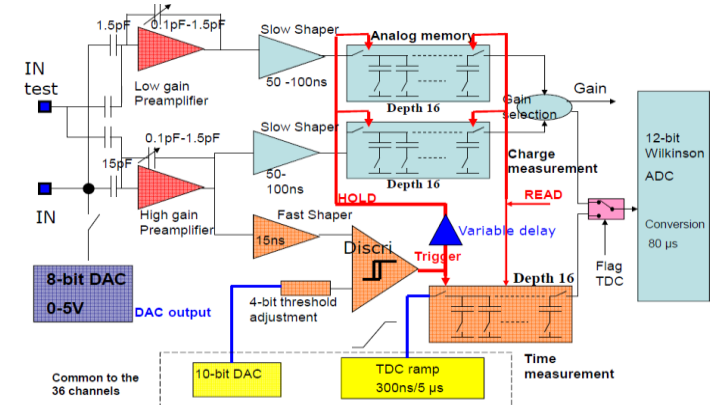
Linearity Calibration based-on DAC

- The DAC test results are fitted channel by channel
- The high gain channel factor is 0.19 ADC/pC, and the low gain channel is 6.1 ADC/pC

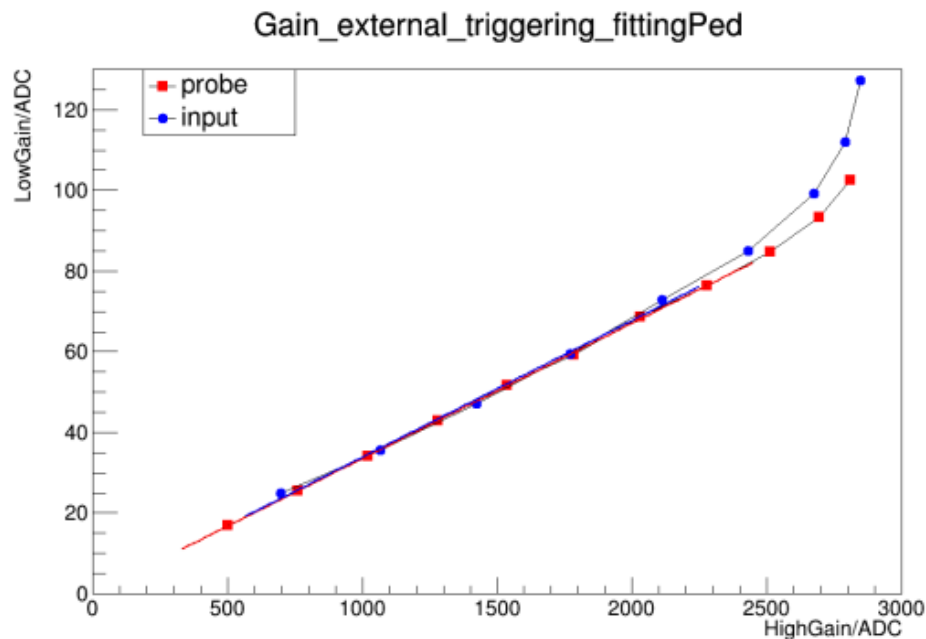


High gain and low gain ratio

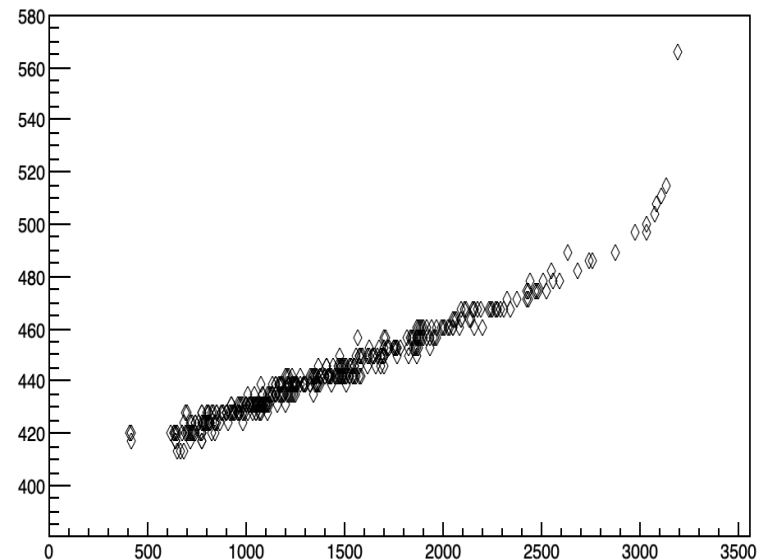
- SPIROC2E chip has two different gain channel
- Cosmic ray test can only obtain the relationship between energy and ADC in high gain channel, but there is no signal in low gain channel
- The coefficients of high and low gain are very important to realize the conversion of high and low gain channels
- Three different methods were used to calibrate the ratio
 - DAC test
 - LED test
 - Cosmic ray test



High gain and low gain ratio



The ratio of LED (blue) is 29.5,
and DAC (red) is 29.8, the two
test results are almost the same

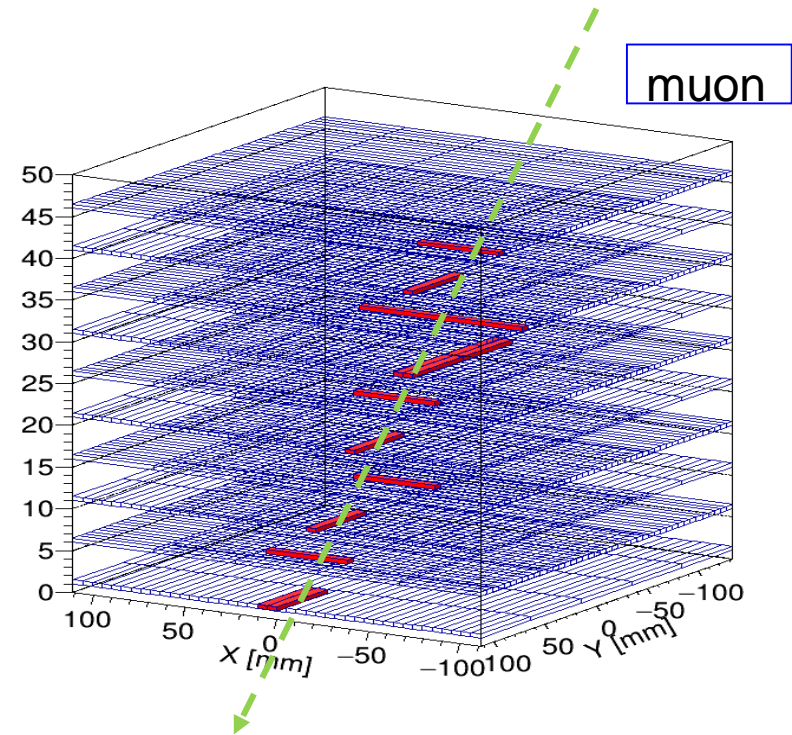


The cosmic ray test results
were analyzing now

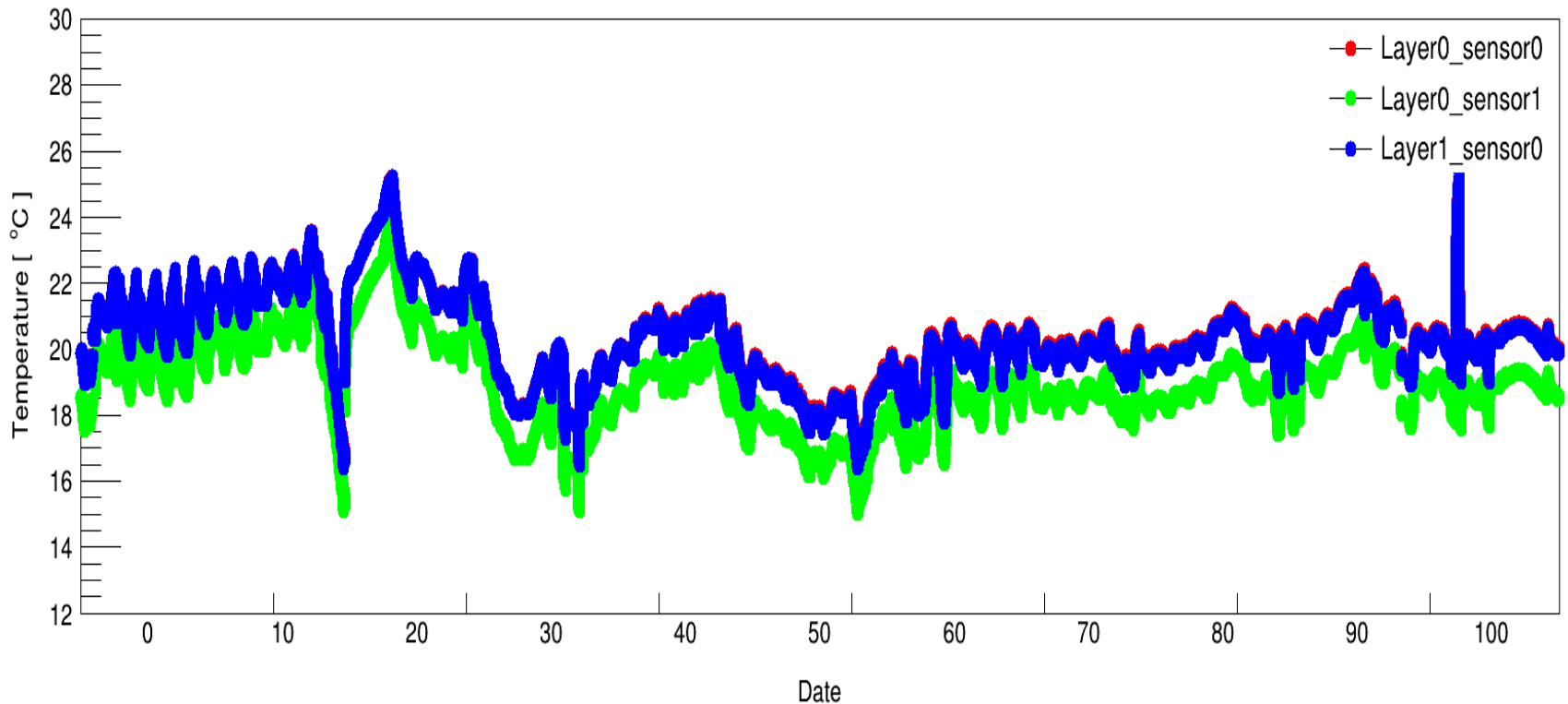


The stability of the cosmic test

- Stability is an important index of the detector
- We monitor some parameters in this test period, ~3 month
 - Temperature
 - Pedestal, MIPs spectra, detection efficiency..



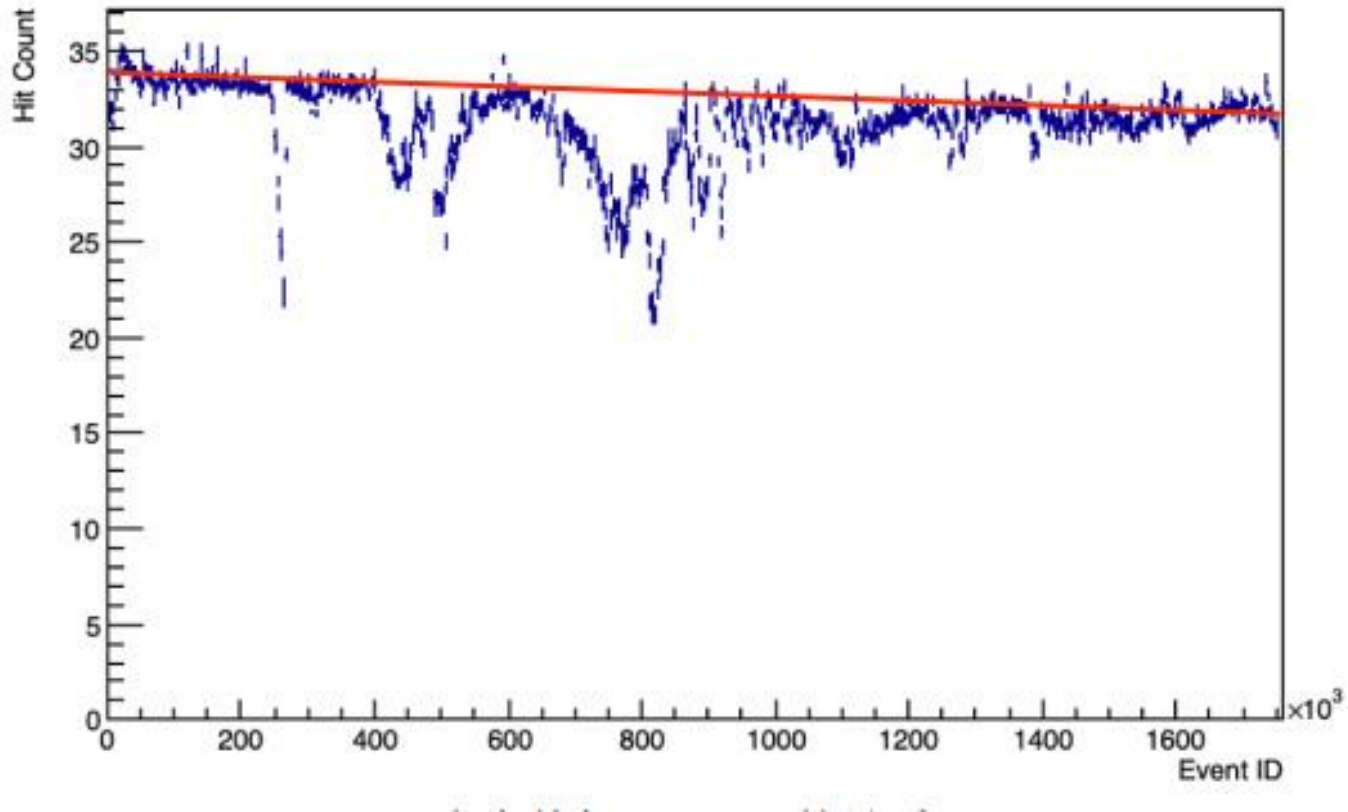
Temperature



- The temperature is between 14 and 26 degrees, with an average of 20 degrees
- At first the test room with relatively good temperature control conditions
- Most of the time, the temperature control condition of the room is not very good



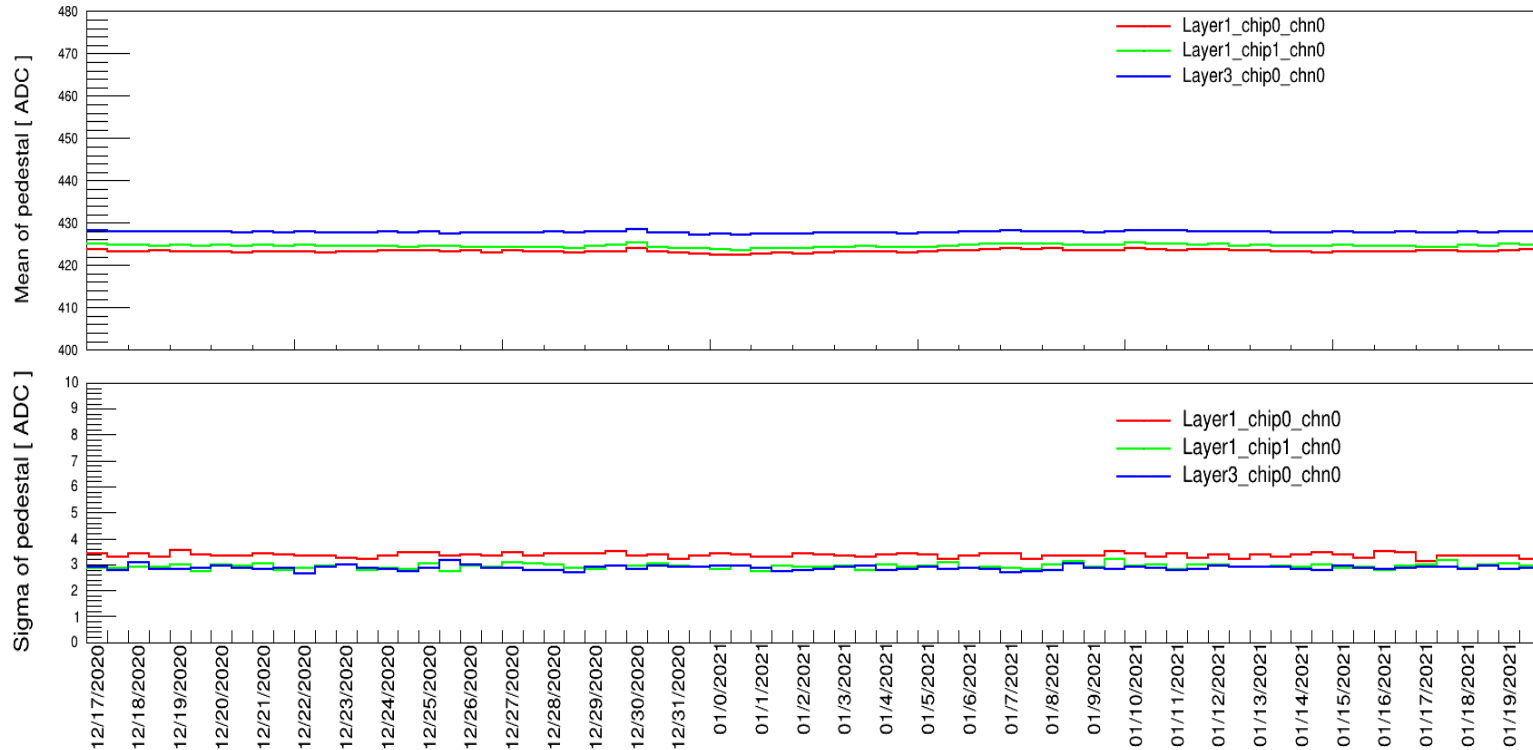
Hit Number stability



- The hit number is between 20 and 40 hits, with an average of 32
- There are two obvious decreases in the middle, which are related to the temperature



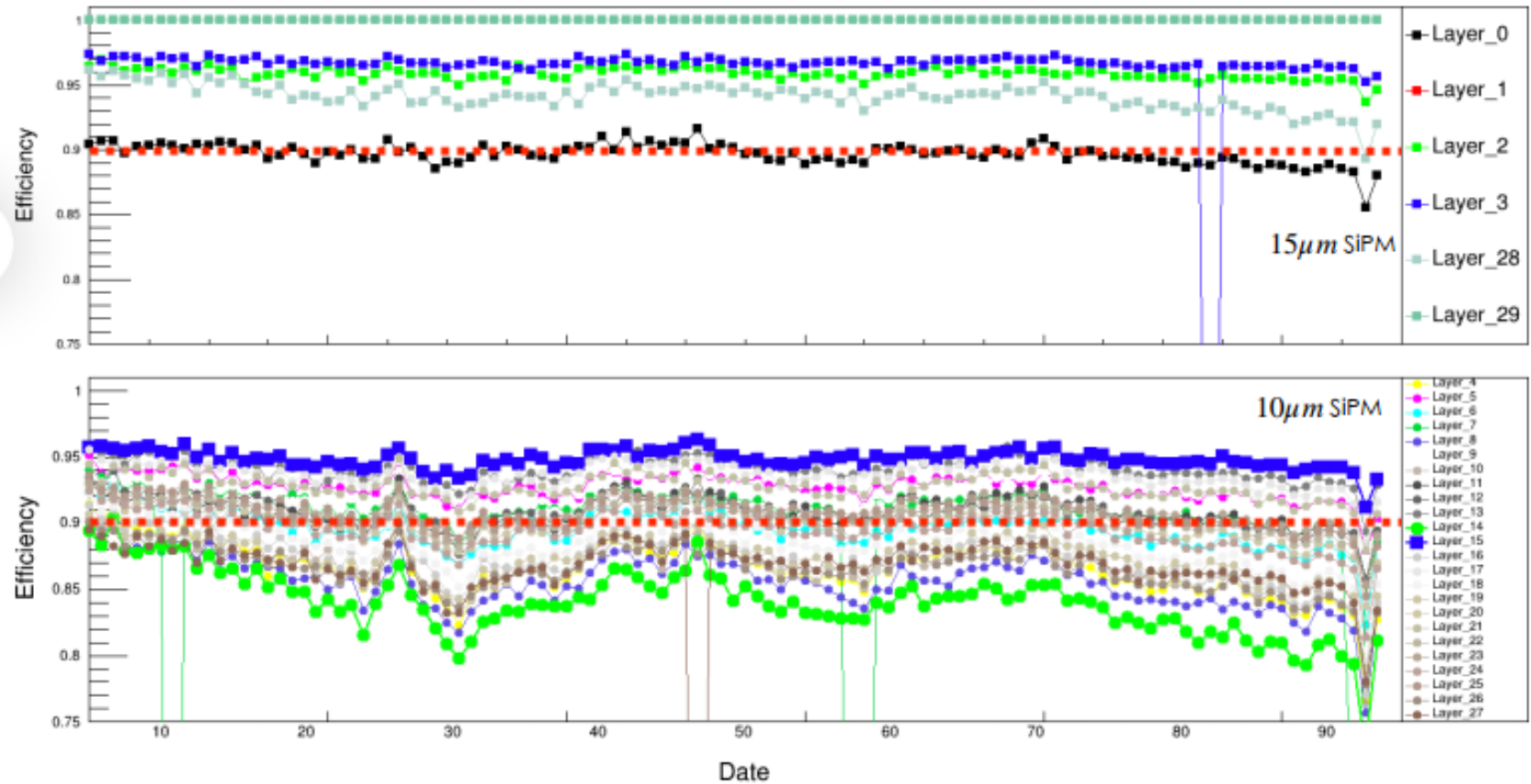
Pedestal stability



- The pedestal position and width are both very stable
- The noise (pedestal width) are about 4 ADC



Detection efficiency



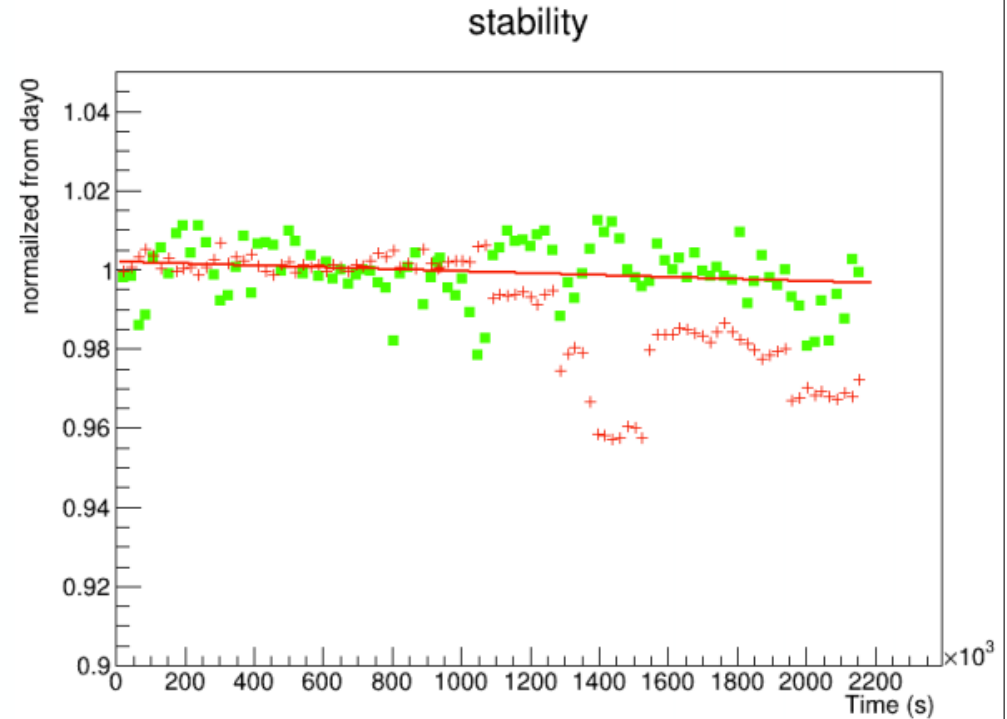
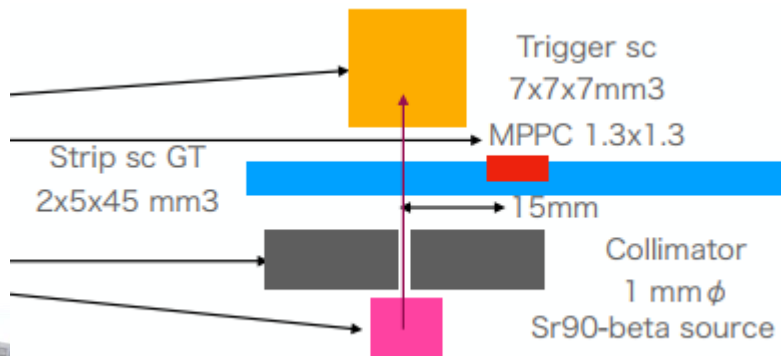
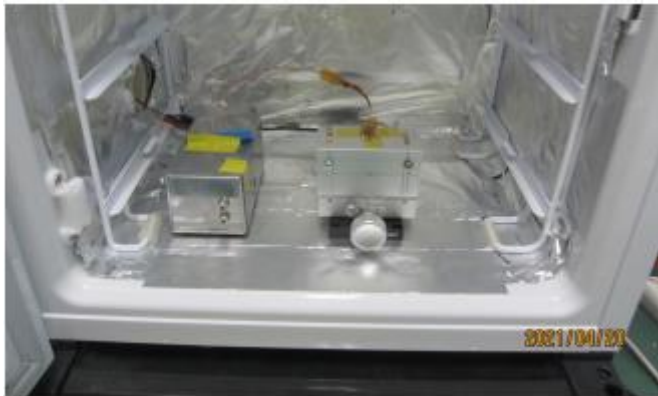
- The detection efficiency is generally stable
- the efficiency seems to decrease gradually for 10 um SiPM



Scintillator/SiPM stability

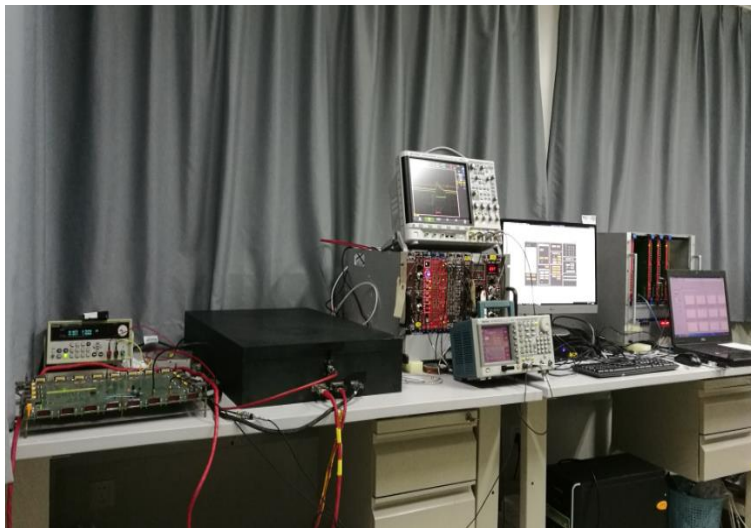
Shinshu group

A Sr-90 source was used to monitor the MIPs peak position stability

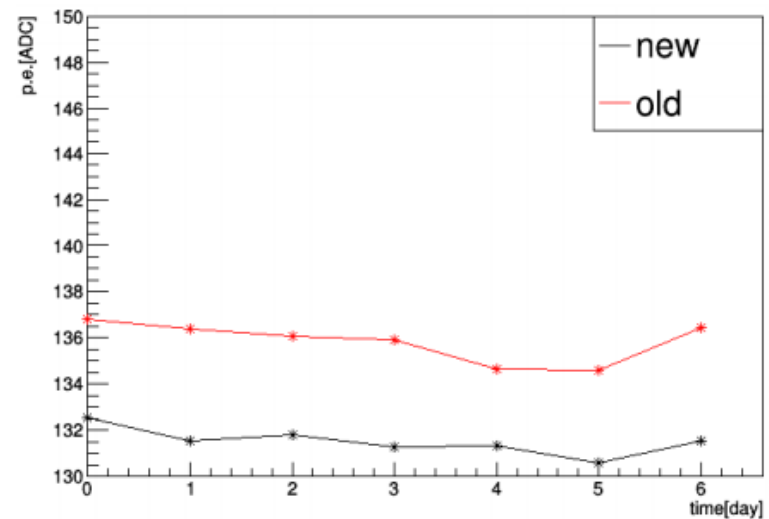
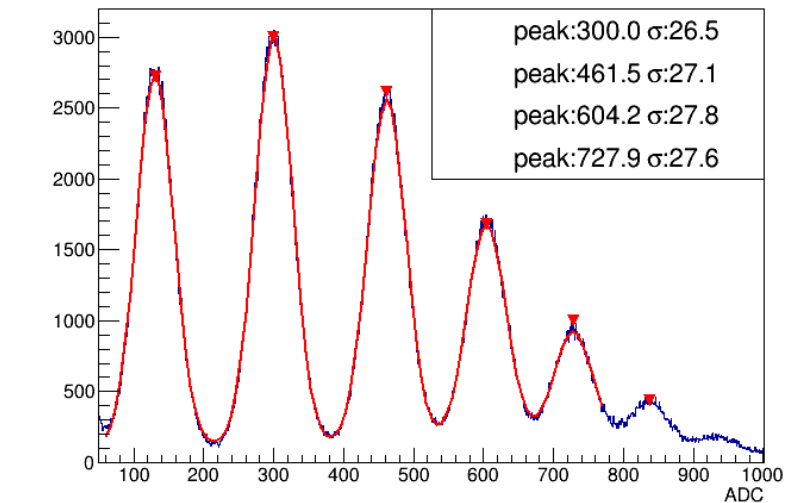


SiPM stability

USTC group



An LED test system was used to calibrate the SPE

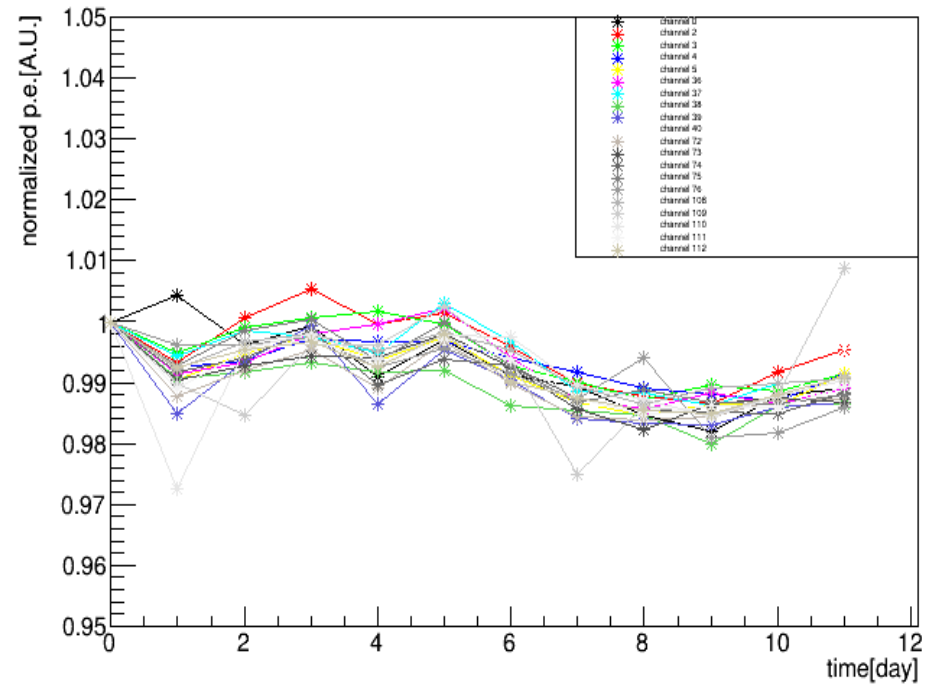


SiPM stability

USTC group

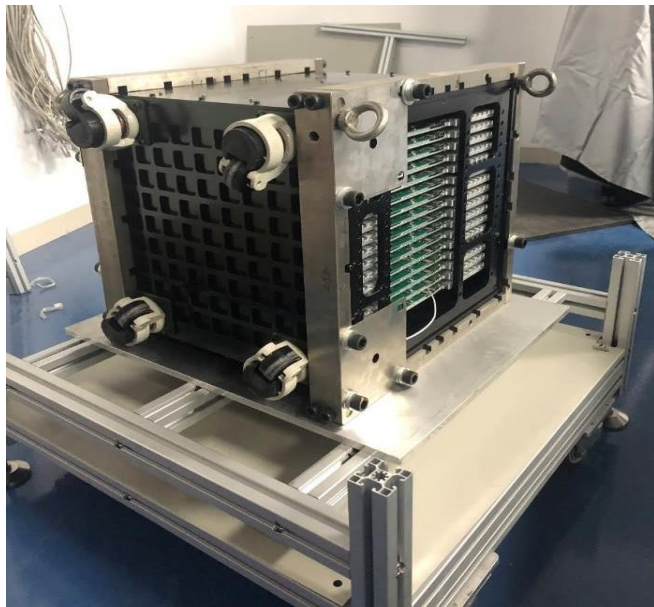


The AHCAL scintillator batch test system was also used to test the stability of SiPM

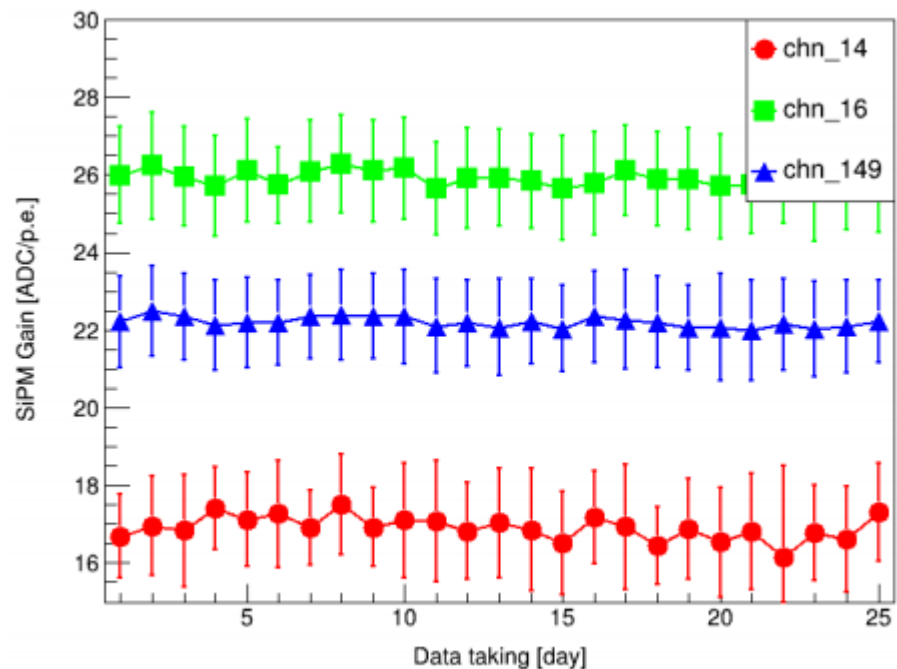


SiPM stability

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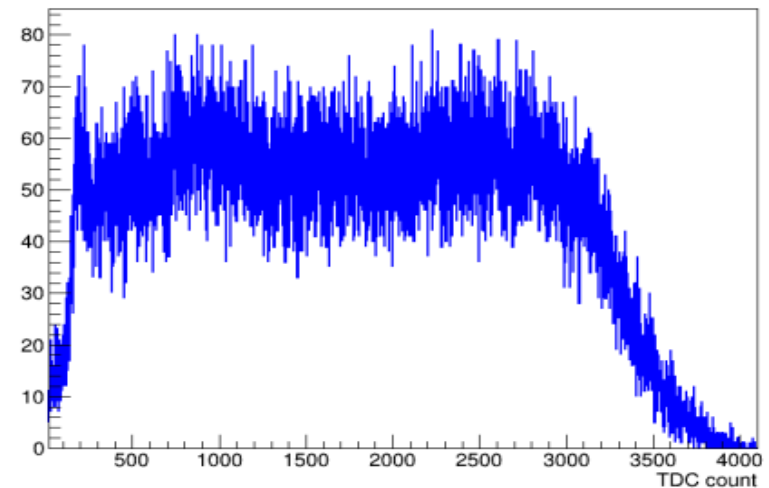
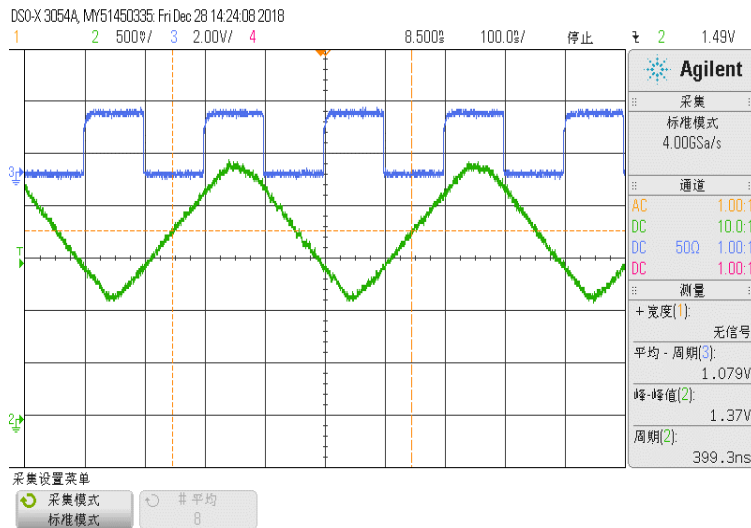


The Sci-ECAL was used LED driver to monitor the SiPM gain day by day



Time Mode Test

- ◆ SP2E chip has a function of time measurement
- ◆ In the previous test, we used high/low gain mode for test
- ◆ and now we use time measurement mode



Cosmic ray events fall randomly in the time measurement interval

SPIROC2E chip

Summary and outlook

- The cosmic ray test provides us with a lot of data
- Develop analysis methods to facilitate channel by channel analysis of calorimeter, and batch analysis of data
- Some important parameters are analyzed channel by channel
- Next step, we hope to reconstruct the beam test data of last year in IHEP according to these parameters
- And, continue to carry out time mode testing in the laboratory



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THANKS



backup



ECAL test trigger

