Top quark properties measurements at CEPC

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Motivation

- Top properties are fundamental in the Standard Model, also can be stringent check of internal consistency of SM.
- Top properties are measured using top reconstruction at hadron colliders. But it is difficult to further improve the precision given dominant systematic uncertainties at hadron colliders.
- e⁺e⁻ colliders can provide not only the top reconstruction method but also the ttbar threshold scan.

Motivation

- ttbar threshold scan is made against \sqrt{s} and cross section, which is direct observable.
- It brings measurements of such parameters:
 - Top mass
 - Top width
 - Top Yukawa coupling
 - $\alpha_{\rm s}$ (strong coupling)



Our setup

- Use the package "QQbar_threshold" to calculate cross section near threshold in ee-colliders at N3LO in resummed non-relativistic perturbation theory
 - Coulomb interactions between the quark and the antiquark leading to a strong enhancement of the cross section is included
 - Initial state radiation (ISR) effects are also included in the package
- We incorporate Luminosity Spectrum(LS) by a simple Gaussian function with CEPC LS (~0.5GeV, provided by Yiwei Wang) as the energy resolution at the moment.



Fig. 4 Top pair production cross section from theory calculations, with the luminosity spectrum (LS) of CLIC at 350 GeV and ISR as well as for all effects combined

Method: \sqrt{s} scan

- Since we are interested in the precise measurement of top mass/width by using threshold scan, we can try to fit the calculated models to experiment data.
- We can construct our likelihood function with 1 energy point in the following way:

$$L = P(D|E(\sigma(m_{top}, \Gamma_{top}, \alpha_{s'}, \sqrt{s})), \mathcal{L}, \theta)$$

- Since we do not have real experiment data, we use <u>QQbar_threshold</u> to generate pseudo data instead.
 - In this set of pseudo data, top mass is set to be 171.5 GeV, and top width is set to be 1.33 GeV.

Method: \sqrt{s} scan

 For different top mass/width, we select multiple center mass energy points. To combine the statistical power of all scan points, we can multiply 1-point likelihood functions together:

$$L = \prod_{i} P(D_i | E_i(\sigma(m_{top'} \Gamma_{top'} \alpha_{s'} \sqrt{s})), \mathcal{L}_i, \theta)$$

• i corresponds to the i-th scan point



Reminder: last status at Yangzhou Workshop

- We use these setup:
 - The acceptance and selection efficiency are assumed to be 100%.
 - Background events are not considered.
 - ISR is considered, but LS is not included.
 - Luminosity per scan point is assumed to range from 25fb⁻¹ to 100fb⁻¹.
 - Systematic uncertainties are not considered.
 - Top mass is set to be 171.5GeV. Top width is set to be 1.33GeV.
- We use these 3 following schemes:

4 points scheme, 6 points scheme, and 8 points scheme.

$6\sqrt{s}$ scheme={341,342,342.5,343,343.5,344.5}



25fb⁻¹ sigma: -0.004125 +0.004125

100fb⁻¹ sigma: -0.002189 +0.002189

\sqrt{s} NLL scan

- We pick the 6- \sqrt{s} scheme for its better performance.
- And we tested more luminosity assumptions.
- The curve is consistent with our expectation.



 $4\sqrt{s}$ scheme={341.5,342.5,343,344.5}

Summary of Previous Results

scheme	4 points	6 points	8 points
σ(m _t)/MeV 100fb ⁻¹	2.9	2.2	2.2
σ(m _t)/MeV 25fb ⁻¹	5.1	4.1	4.1

- Build up the machinery of this \sqrt{s} scan to estimate measurement uncertainties.
- Test with a few set of parameters and schemes.
- The way of selecting points is crucial if we want lower error.



Fisher information



Setup

- ISR is considered, but LS is not considered. Systematic uncertainties are not taken into account.
- Acceptance and selection efficiency are added.
- Background events are included.
- We consider these 2 channels: semi-leptonic and fully-hadronic.
 - The selection efficiency and background events obtained from Eur. Phys. J. C (2013) 73:2530

Luminosity and scan \sqrt{s} range

- In reality, the total operation time is limited, so the total luminosity is limited.
- We study the scanning schemes with total lumi fixed.
- Total luminosity will be 100 fb^{-1} . The luminosity of each point is the same. \sqrt{s} scan ranges from 340GeV to 345GeV.
 - Drop less sensitive points step by step from 8 points to 1 point.



$8\sqrt{s}$ scheme ={340,341,342,342.5,343,343.5,344.5,345}



12.5fb⁻¹ per point $\sigma(m_t)$: -0.0140273 0.0139727

$6\sqrt{s}$ scheme={341,342,342.5,343,343.5,344.5}

We dropped 340 and 345.



 $4\sqrt{s}$ scheme={342,342.5,343,343.5}

We dropped 341 and 344.5.



$1\sqrt{s}$ scheme={343}



Results

scheme	8 points	6 points	4 points	1 point
σ (m _t) /MeV	13.97	10.43	12.06	8.40

- For 171.5GeV top mass, 343 GeV center mass energy is the best point, given the total luminosity 100 fb⁻¹.
- Top mass is known as 171.5GeV, so we can get the best point through its known fisher information. But for unknown top mass, we need to first locate a proper range.



$1-\sqrt{s}$ scheme gives the best result

- We use 343GeV as the 1 point scheme.
- We used equal luminosity per point.
- So we have 2 questions:
 - How about unequal luminosity per point?
 - Is 343GeV really the best point?

343GeV

 $\sqrt{s} = \{342, 342.5, 343, 343.5\}$

Exhaustion on 4 sqrts scheme

- Total lumi = $100 fb^{-1}$
- $\sqrt{s} = \{342, 342.5, 343, 343.5\}$
- Calculate the top mass errors of all schemes with possible lumi fractions while keeping the total lumi fixed.
- 286 lumi combinations in total
 - List from low error to high error
 - Top 30 are listed
- Conclusion: 343GeV is the best point.

v3 ={0	+2, 0+2.0, 0+0, 0+0.0}	
lumi ratio= ·	$\{0, 0, 100, 0\},\$	err= 0.00840344
lumi ratio= ·	{0, 10, 90, 0},	err= 0.00852734
lumi ratio= ·	{0, 10, 80, 10},	err= 0.0085625
lumi ratio= ·	{0, 20, 80, 0},	err= 0.0085625
lumi ratio= ·	{10, 0, 90, 0},	err= 0.0085625
lumi ratio= ·	{0, 0, 90, 10},	err= 0.00861523
lumi ratio= ·	{10, 0, 80, 10},	err= 0.00889108
lumi ratio= ·	{0, 40, 60, 0},	err= 0.00889648
lumi ratio= ·	{0, 0, 80, 20},	err= 0.0089375
lumi ratio= ·	{0, 20, 70, 10},	err= 0.0089375
lumi ratio= ·	{0, 30, 70, 0},	err= 0.0089375
lumi ratio= ·	{10, 10, 80, 0},	err= 0.0089375
lumi ratio= ·	{10, 20, 70, 0},	err= 0.0089375
lumi ratio= ·	{20, 0, 80, 0},	err= 0.00896973
lumi ratio= ·	{0, 10, 70, 20},	err= 0.00897742
lumi ratio= ·	{0, 30, 60, 10},	err= 0.00900123
lumi ratio=	{10, 10, 70, 10},	err= 0.00902295
lumi ratio= ·	{0, 10, 60, 30},	err= 0.0090625
lumi ratio=	{0, 40, 50, 10},	err= 0.0090625
lumi ratio= ·	{0, 50, 50, 0},	err= 0.0090625
lumi ratio=	{0, 60, 40, 0},	err= 0.0090625
lumi ratio=	{10, 20, 60, 10},	err= 0.0090625
lumi ratio= ·	{10, 40, 50, 0},	err= 0.0090625
lumi ratio= ·	{20, 0, 70, 10},	err= 0.0090625
lumi ratio= ·	{0, 0, 70, 30},	err= 0.00906836
lumi ratio= ·	{10, 0, 70, 20},	err= 0.0090918
lumi ratio= ·	{0, 20, 60, 20},	err= 0.00910352
lumi ratio= ·	{10, 30, 60, 0},	err= 0.00910937
lumi ratio=	$\{20, 10, 70, 0\},\$	err= 0.00911993
lumi ratio=	$\{0, 30, 50, 20\},$	err= 0.00925
lumi ratio= ·	$\{0, 40, 40, 20\},$	err= 0.00939502

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Proposal to Find the Best Point

- Running at a low luminosity (1fb⁻¹) to find out if 343GeV is the best point.
- The discriminant value is much smaller than the one used for deriving $\sigma(m_t)$

sqrts = {340, 341,	342, 342.5,	343, 343.5	, 344.5, 345}
lum= 1, discrimina	ant value = 10	e-4	
lumi ratio= {0, 0,	0, 0, 1, 0,	0,0}, er	r= 0.0010625
lumi ratio= {0, 0,	0, 1, 0, 0,	0,0}, er	r= 0.00138158
lumi ratio= {0, 0,	0, 0, 0, 1,	0,0}, er	r= 0.00159912
lumi ratio= {0, 0,	1, 0, 0, 0,	0,0}, er	r= 0.00190234
lumi ratio= {0, 1,	0, 0, 0, 0, 0,	0,0}, er	r= 0.0039375
lumi ratio= {0, 0;	0, 0, 0, 0, 0,	1, 0}, er	r= 0.00447998
lumi ratio= {0, 0,	0, 0, 0, 0, 0,	0, 1}, er	r= 0.00678113
lumi ratio= {1, 0,	0, 0, 0, 0, 0 <u>,</u>	0,0}, er	r= 0.0069375

Adding Luminosity Spectrum

- Total luminosity will be 100fb⁻¹.
 - We would like to compare our results with CLIC, so we are trying to keep these parameters close to CLIC's.
- The luminosity spectrum of CEPC



Provided by Yiwei Wang

Eur. Phys. J. C (2013) 73:2530

Setup

- ISR and LS are considered. Systematic uncertainties are not taken into account.
- Top mass is set to be 171.5GeV. Top width is set to be 1.33GeV.
- Acceptance and selection efficiency are added.
- Background events are included.
- We consider these 2 channels: semi-leptonic and fully-hadronic.
 - The selection efficiency and background events obtained from Eur. Phys. J. C (2013) 73:2530

$8\sqrt{s}$ scheme ={340,341,342,342.5,343,343.5,344.5,345}



$6\sqrt{s}$ scheme={341,342,342.5,343,343.5,344.5}

We dropped 340 and 345.



$4\sqrt{s}$ scheme={342,342.5,343,343.5}

We dropped 341 and 344.5.



$1\sqrt{s}$ scheme={343}



 $100 \text{fb}^{-1} \text{ per point } \sigma(m_t) : -0.0129375 \ 0.0129375$

Results

scheme	8 points	6 points	4 points	1 point
σ (m _t) /MeV	20.06	17.56	14.93	12.93

- How about unequal lumi per point?
- What is the best point when considering LS added?

L(342.75GeV)=40%	lumi ratio= {0, 40, 60, 0},	err= 0.0125625
L(343GeV)=60%	$lumi ratio = \{0, 50, 50, 0\},$	err= 0.0125625
	lumi ratio= {0, 60, 40, 0},	err= 0.0125625
Exhaustion on 1 carts schome	$1 \text{ umi ratio} = \{0, 70, 50, 0\},$	arr = 0.0125025
LXHAUSUUH UH 4 SYILS SCHEITIE	$1 \text{ mir} = \{10, 60, 20, 0\},\$	err= 0.0125625
	$lumi ratio = \{10, 70, 20, 0\},$	err= 0.0125625
	lumi ratio= {10, 80, 10, 0},	err= 0.0125625
\mathbf{T} , \mathbf{I} , \mathbf{I} , \mathbf{I}	lumi ratio= {10, 90, 0, 0},	err= 0.0125625
• $1 \text{ otal lum} = 100$	lumi ratio= {20, 80, 0, 0},	err= 0.0125625
542.75Ge <u>v</u>	lumi ratio= {0, 100, 0, 0},	err= 0.012605
• $\sqrt{s} = \{342.5, 342.75, 343, 343, 5\}$	lumi ratio= {0, 90, 10, 0},	err= 0.0126211
	lumi ratio= {0, 20, 80, 0},	err= 0.01275
 Calculate the top mass error of all 	lumi ratio= {0, 30, 70, 0},	err= 0.01275
	$lumi ratio = \{10, 50, 40, 0\},$	err= 0.012/5
possible lumi traction	$1 \text{ min} [73110 = \{20, 70, 10, 0\}, 10 \text{ min} [73110 = \{20, 70, 10, 0\}]$	err = 0.01275
	$1 \text{ mir ratio} = \{20, 0, 30, 40, 0\}$	err= 0.0128777
 286 Iumi combinations in total 	$lumi ratio = \{40, 30, 30, 0\},$	err= 0.0128906
· List from low arror to bigh arror	lumi ratio= {30, 20, 50, 0},	err= 0.0128965
 List from low error to high error 	lumi ratio= {40, 40, 20, 0},	err= 0.0128965
 Top 30 are listed 	lumi ratio= {20, 20, 60, 0},	err= 0.0129067
	lumi ratio= {40, 60, 0, 0},	err= 0.0129067
 Conclusion: with cenc LS the best 	lumi ratio= {20, 10, 70, 0},	err= 0.0129141
	lumi ratio= {10, 0, 90, 0},	err= 0.0129181
point shift.	$lumi ratio = \{40, 50, 10, 0\},$	err= 0.0129199
•	$1 \text{ umi ratio} = \{30, 40, 10, 0\},$	err = 0.0129290
3/3GeV	$1 \text{ mit} ratio = \{0, 0, 100, 0\},$	err= 0.0129331
11/8/2021	$lumi ratio = \{0, 10, 90, 0\},$	err= 0.0129375
11/0/2021	lumi ratio= {0, 90, 0, 10},	err= 0.0129375

Proposal to find the Best Point

- Running at a low luminosity (1fb⁻¹)
- The discriminant value is $^{342.5Ge}_{342.75G}$ much smaller than the 343GeV one used for deriving $\sigma(m_t)$

	sqrts	s = {342	2, 34	12.5	5, 3	342.	.75,	, 34	43,	343.25	, 34 3	3.5,	344,	0}
	lum=	1, disc	rimi	inar	۱t ۱	/alı	ie =	= 10	<u>9-4</u>					
eV—	lumi	ratio=	{0,	1,	0,	0,	0,	0,	0,	0},	err=	0.00	91937	5
eV—	lumi	ratio=	{0,	0,	1,	0,	0,	0,	0,	0},	err=	0.00	91937	5
<i>i</i> —	lumi	ratio=	{0,	0,	0,	1,	0,	0,	0,	0},	err=	0.00	91937	5
V	lumi	ratio=	{0,	0,	0,	0,	1,	0,	0,	0},	err=	0.00	020258	88
	lumi	ratio=	{1,	0,	0,	0,	0,	0,	0,	0},	err=	0.00	923906	62
	lumi	ratio=	{0,	0,	0,	0,	0,	1,	0,	0},	err=	0.00	92562	5
	lumi	ratio=	{0,	0,	0,	0,	0,	0,	1,	0},	err=	0.00	984692	24
	lumi	ratio=	{0,	0,	0,	0,	0,	0,	0,	1},	err=	0.00	084692	24

• Cannot distinguish 342.5, 342.75 and 343

Proposal to find the Best Point

- Running at 10fb⁻¹
- The discriminant value is much smaller than the one used for deriving $\sigma(m_t)$

<pre>342.75GeV 343GeV 343GeV 343GeV 342.5GeV 343GeV 342.5GeV 342.5</pre>	U,
342.75GeV lumi ratio= {0, 0, 10, 0, 0, 0, 0, 0}, err= 0.0055625 343GeV lumi ratio= {0, 0, 0, 10, 0, 0, 0, 0}, err= 0.0055625 lumi ratio= {0, 10, 0, 0, 0, 0, 0, 0}, err= 0.0059375 lumi ratio= {0, 0, 0, 0, 0, 10, 0, 0, 0}, err= 0.0064140	
343GeV 342.5GeV lumi ratio= {0, 0, 0, 10, 0, 0, 0, 0}, err= 0.0055625 lumi ratio= {0, 10, 0, 0, 0, 0, 0, 0}, err= 0.0059375 lumi ratio= {0, 0, 0, 0, 0, 10, 0, 0, 0}, err= 0.0064140	5
342.5GeV → lumi ratio= {0, 10, 0, 0, 0, 0, 0, 0}, err= 0.0059375 lumi ratio= {0, 0, 0, 0, 0, 10, 0, 0, 0}, err= 0.0064140	5
lumi ratio= {0, 0, 0, 0, 10, 0, 0}, err= 0.0064140	5
)6
lumi ratio= {10, 0, 0, 0, 0, 0, 0, 0}, err= 0.0075683	36
lumi ratio= {0, 0, 0, 0, 0, 10, 0, 0}, err= 0.0080390)6
lumi ratio= {0, 0, 0, 0, 0, 0, 10, 0}, err= 0.02775	
lumi ratio= {0, 0, 0, 0, 0, 0, 0, 10}, err= 0.02775	

 cannot distinguish 342.75 and 343

Compare with CLIC and FCC-ee

scheme	8 points	6 points	4 points	1 point
$\sigma(m_t)/MeV$	20.06	17.56	14.93	12.93

Comparable with FCC-ee under similar conditions (lumi differ by a factor of 2)

- The estimation of FCC-ee:
 - ~17 MeV for top mass (stat. uncert.)
 - ~45 MeV for top width (stat. uncert.)
 - with 25fb⁻¹ taken at each of the 8 centre-of-mass energy points N3LO cross-section calculation brings 40 MeV uncertainty additionally

2d fit results of CLIC Eur. Phys. J. C (2013) 73:2530

Table 4 Summary of the 2D simultaneous top mass and α_s determination with a threshold scan at ILC for 10 points with a total integrated luminosity of 100 fb⁻¹. Event selection and background rejection from CLIC_ILD is used

1S top mass and α_s combined 2D fit

m_t stat. error	27 MeV
m_t theory syst. (1 %/3 %)	5 MeV/9 MeV
α_s stat. error	0.0008
α_s theory syst. (1 %/3 %)	0.0007/0.0022

Summary

- The uncertainty of this method is tightly related to our points selection scheme.
- 1-point scheme has the best performance, if we have already found the best point.
- We proposed to scan the point with low luminosity to identify the best point.
 - the method should be further investigated, considering the effects of systematics, etc.
- We can scan some points in non-sensitive area (e.g. 320GeV) to do background study.

Next

- We should find the lowest luminosity where we can distinguish similar \sqrt{s} .
- LS can cause the shift of the best point. We should find a proper way to deal with it.
- Width study is working in progress.
 - The best points for width and for mass are different.
- 2D-scan should be applied, considering the influence of both mass and width.

Back-up

How do we get the Fisher information

- We assume that the number of events obeys $\sim N(mu,sigma)$.
- We believe it is an extreme situation of Poisson distribution, so here mu = sigma.
- Then we construct its likelihood function, but only pick up 1 point in sample space.

$$egin{aligned} l(x| heta) &= \log f(x| heta) \ l'(x| heta) &= rac{\partial}{\partial heta} \log f(x| heta) = rac{f'(x| heta)}{f(x| heta)} \ I(heta) &= E[l'(X| heta)^2] = \int [l'(X| heta)^2] f(x| heta) dx \end{aligned}$$

Acceptance and selection efficiency for signal

- The number read from CLIC Eur. Phys. J. C (2013) 73:2530
- semi-leptonic :
 - Data: 8296, Bkg: 643, extracted signal: 7653, acceptance*selection efficiency = 48.13%, Branch ratio=30%
- Full-hadronic
 - Data: 11396, Bkg: 1393, extracted signal: 10003, acceptance*selection efficiency = 41.0%, Branch ratio=46%
- These parameters are under 500 GeV situation. At the moment we assume that acceptance and selection efficiency will not change under 352 GeV situation.
- The signal yields of our pseudo data: at 343GeV, 100 $\rm fb^{-1}$
 - semi leptonic 4009.14
 - fully hadronic 5236.67

Background events

- Background events are directly scaled from 500GeV to 352GeV, according to their cross section estimated by CLIC paper.
 - For CLIC's 500GeV situation, the luminosity is 100 fb^{-1}
 - Because there is no information about background yields under 352GeV in the paper of CLIC. Table 1 Signal and considered physics background processes, with their approximate cross section calculated for CLIC at 500 GeV and at

Background

- Result:
 - semi leptonic bkg event number:2380
 - fully hadronic bkg event number:5156

352 GeV Type Final σ σ 500 GeV 352 GeV state Signal ($m_{top} = 174 \text{ GeV}$) tī 530 fb 450 fb Background WW 7.1 pb 11.5 pb Background 410 fb 865 fb ZZ Background 25.2 pb

 $q\bar{q}$

WWZ

2.6 pb

40 fb

10 fb