

# Yukawa Unification and Sparticle Spectroscopy at the LHC/Tevatron

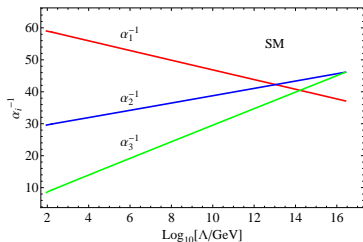
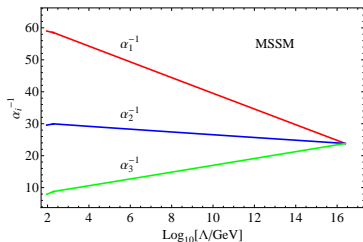
Qaisar Shafi

Bartol Research Institute  
Department Physics and Astronomy  
University of Delaware, USA

in collaboration with Ilia Gogoladze, Rizwan Khalid, Shabbar Raza,  
Adeel Ajaib, Tong Li and Kai Wang.

## Low Scale ( $\sim$ TeV) Supersymmetry (SUSY):

- Arguably the most compelling extension of the Standard Model;
- Resolves the gauge hierarchy problem;
- Provides cold dark matter candidate (LSP);
- Implements radiative electroweak symmetry breaking;
- Predicts new particles accessible at the LHC, and thereby enables unification of the SM gauge couplings;



- **Supersymmetric SO(10):**

- Fermion families reside in  $16_j (j=1,2,3)$  predicts 'right handed' neutrino  $\Rightarrow$  non-zero neutrino masses;

(Cf: SU(5) with families in  $\overline{10}_i + \overline{5}_i$ )

- Yukawa couplings provide masses to SM fermions. They include

$$16_i 16_j 10, 16_i 16_j 126, \text{ etc.}$$

- $16_3 16_3 10$  yields  $t - b - \tau$  unification

$$Y_t = Y_b = Y_\tau = Y_\nu$$

$\rightarrow$  In the old days it was used to predict the top quark mass!<sup>1</sup>

---

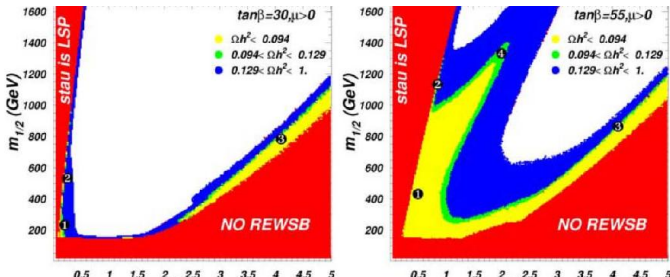
<sup>1</sup>B. Ananthanarayan, George Lazarides, Q. Shafi (1991);

- Nowadays, one employs  $t - b - \tau$  unification to make predictions, such as sparticle masses, which can be tested at the LHC/Tevatron (Baer et al.);
- $t - b - \tau$  unification can also be realized in  $SU(4)_c \times SU(2)_L \times SU(2)_R$ , a maximal subgroup of  $SO(10)$ ;

## CMSSM (mSUGRA):

- Unbroken  $Z_2$  matter parity  $\Rightarrow$  stable LSP, typically neutralino;
- Universal soft susy breaking parameters

$$m_0, m_{1/2}, A_0, \tan\beta, \text{sign}(\mu)$$



Alexander Belyaev, *Pramana* 72:143-160,2009.

## Supersymmetric SO(10)(Baer et al.)<sup>1</sup>

- $m_{16}, m_{10}, M_D, M_{1/2}, A_0, \tan \beta, \text{sign}(\mu)$
- $m_{16} \equiv$  Universal soft SUSY breaking sfermion mass
- $m_{10} \equiv$  Universal soft SUSY breaking MSSM Higgs mass
- $M_D \equiv$  The Higgs mass splitting  $M_{H_{u,d}}^2 = m_{10}^2 \mp 2M_D^2$
- $m_{1/2} \equiv$  Universal SSB gaugino mass
- $A_0 \equiv$  Universal SSB trilinear interaction
- $\tan \beta = \frac{v_u}{v_d}$
- $\mu \equiv$  SUSY bilinear Higgs parameter

<sup>1</sup>H. Baer, S. Kraml, S. Sekmen and H. Summy, JHEP 0803, 056 (2008)

- Random scans were performed over the parameter space

$m_{16}$ :	$0 \rightarrow 20$ TeV	(1 – 20 TeV),
$m_{10}/m_{16}$ :	$0 \rightarrow 1.5$	(0.8 – 1.4),
$m_{1/2}$ :	$0 \rightarrow 5$ TeV	(0 – 1 TeV),
$A_0/m_{16}$	$-3 \rightarrow 3$	(-2.5 – 1.9),
$M_D/m_{16}$ :	$0 \rightarrow 0.8$	(0.25 – 0.8),
$\tan \beta$ :	$40 \rightarrow 60$	(46 – 53).

- Quantify Yukawa unification by

$$R = \frac{\max(y_t, y_b, y_\tau)}{\min(y_t, y_b, y_\tau)}$$

## Constraints

$$m_{\tilde{\chi}_1^\pm} \text{ (chargino mass)} \geq 103.5 \text{ GeV},$$

$$m_h \text{ (lightest Higgs mass)} \geq 114.4 \text{ GeV},$$

$$m_{\tilde{\tau}} \text{ (stau mass)} \geq 86 \text{ GeV},$$

$$m_{\tilde{g}} \text{ (gluino mass)} \geq 220 \text{ GeV},$$

$$BR(B_s \rightarrow \mu^+ \mu^-) < 5.8 \times 10^{-8},$$

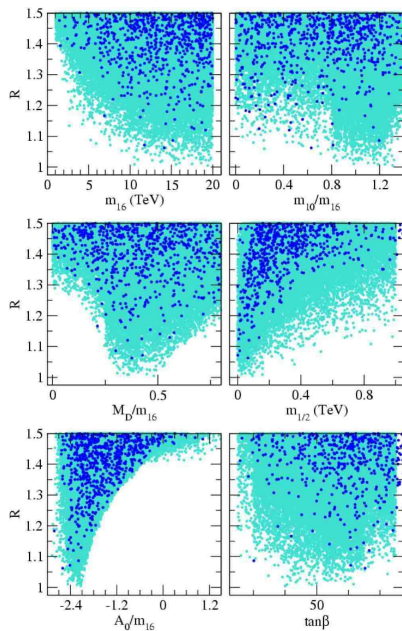
$$0.53 < \frac{BR(B_u \rightarrow \tau \nu_\tau)_{MSSM}}{BR(B_u \rightarrow \tau \nu_\tau)_{SM}} < 2.03 \text{ (} 2\sigma \text{)},$$

$$2.85 \times 10^{-4} \leq BR(b \rightarrow s \gamma) \leq 4.24 \times 10^{-4} \text{ (} 2\sigma \text{)},$$

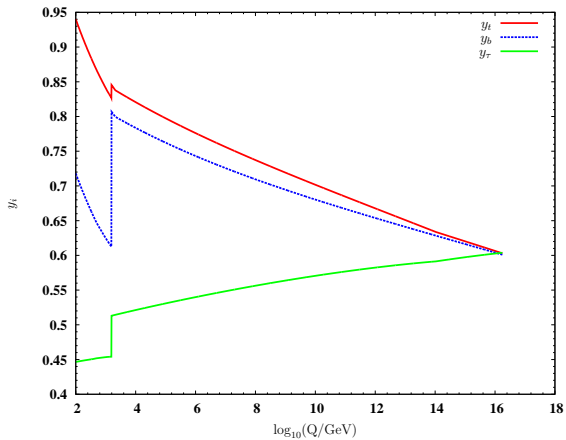
$$\Omega_{\text{CDM}} h^2 = 0.111_{-0.037}^{+0.028} \text{ (} 5\sigma \text{)},$$

$$3.4 \times 10^{-10} \leq \Delta\alpha_\mu \leq 55.6 \times 10^{-10} \text{ (} 3\sigma \text{)}.$$



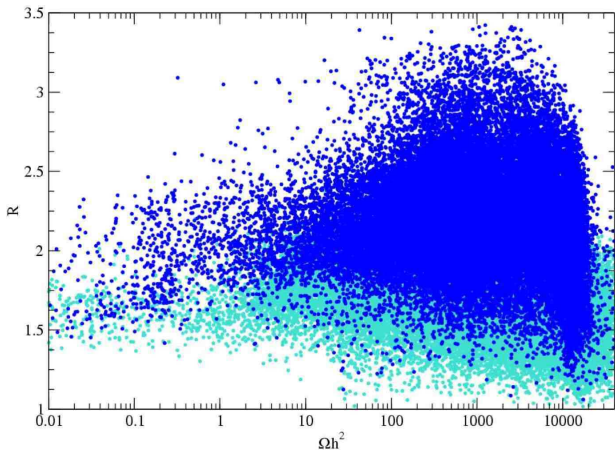


- SUSY and  $t - b - \tau$  Yukawa coupling unification



Radiative contributions to the bottom quark mass from the gluino and chargino loop

$$\frac{\delta m_b}{m_b} \approx \frac{g_3^2}{12\pi^2} \frac{\mu m_{\tilde{g}} \tan \beta}{m_{\tilde{b}}^2} - \frac{y_t^2}{32\pi^2} \frac{\mu A_t \tan \beta}{m_{\tilde{t}}^2} + \dots$$



H. Baer, S. Kraml, S. Sekmen and H. Summy, JHEP 0803, 056 (2008)

parameter	Pt. A	Pt. D
$m_{16}$	9202.9	2976.5
$m_{1/2}$	62.5	107.0
$A_0$	-19964.5	-6060.3
$m_{10}$	10966.1	3787.9
$\tan \beta$	49.1	49.05
$M_D$	3504.4	1020.8
$f_t$	0.51	0.48
$f_b$	0.51	0.47
$f_\tau$	0.52	0.52
$\mu$	4179.8	331.0
$m_{\tilde{g}}$	395.6	387.7
$m_{\tilde{u}_L}$	9185.4	2970.8
$m_{\tilde{t}_1}$	2315.1	434.5
$m_{\tilde{b}_1}$	2723.1	849.3
$m_{\tilde{e}_L}$	9131.9	2955.8
$m_{\tilde{\chi}_1^\pm}$	128.8	105.7
$m_{\tilde{\chi}_2}$	128.6	105.1
$m_{\tilde{\chi}_1}$	55.6	52.6
$m_A$	3273.6	776.8
$m_h$	125.4	111.1
$\sigma$ [fb]	75579.1	89666.1
% ( $\tilde{g}\tilde{g}$ )	86.8	80.5
% ( $\tilde{\chi}_1^\pm\tilde{\chi}_2^\pm$ )	8.8	12.8
% ( $\tilde{t}_1\tilde{t}_1$ )	0	1.1

- Lightest colored sparticle is gluino; But  $\Omega h^2 \gg 1$  !!
- DM: Axions, Axinos.

## Yukawa Unification and Neutralino DM in $SU(4)_c \times SU(2)_L \times SU(2)_R$ (4-2-2)

I.G. R. Khalid and Q. Shafi, Phys. Rev. D 79, 115004 (2009) .

- SM fermions:  $\psi_i = (\mathbf{4}, \mathbf{2}, \mathbf{1})$  and  $\psi_i^c = (\bar{\mathbf{4}}, \mathbf{1}, \mathbf{2})$
- MSSM Higgs:  $\mathbf{H} = (\mathbf{1}, \mathbf{2}, \mathbf{2})$

- Third family Yukawa coupling  $\psi \psi^c \mathbf{H}$  yields

$$Y_t = Y_b = Y_\tau = Y_\nu$$

- Asymptotic relation between the three MSSM gaugino masses

$$M_1 = \frac{3}{5} M_2 + \frac{2}{5} M_3$$

- One additional parameter compared to the SO(10) model  
(from gaugino non-universality)

We performed random scans for the following parameter range

$$\begin{aligned}0 &\leq m_{16} \leq 20 \text{ TeV}, \\0 &\leq M_2 \leq 1 \text{ TeV}, \\0 &\leq M_3 \leq 1 \text{ TeV}, \\-3 &\leq A_0/m_{16} \leq 0, \\0 &\leq M_D/m_{16} \leq 0.95, \\0 &\leq m_{10}/m_{16} \leq 1.5, \\40 &\leq \tan \beta \leq 58, \\\mu &> 0, \quad m_t = 172.6 \text{ GeV}.\end{aligned}$$

## Constraints

$$m_{\tilde{\chi}_1^\pm} \text{ (chargino mass)} \geq 103.5 \text{ GeV},$$

$$m_h \text{ (lightest Higgs mass)} \geq 114.4 \text{ GeV},$$

$$m_{\tilde{\tau}} \text{ (stau mass)} \geq 86 \text{ GeV},$$

$$m_{\tilde{g}} \text{ (gluino mass)} \geq 220 \text{ GeV},$$

$$BR(B_s \rightarrow \mu^+ \mu^-) < 5.8 \times 10^{-8},$$

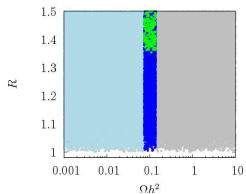
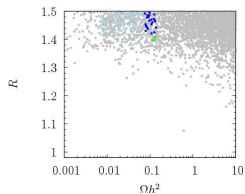
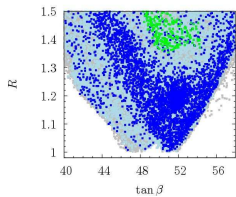
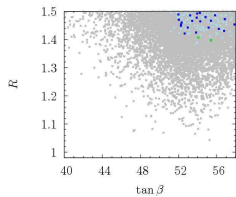
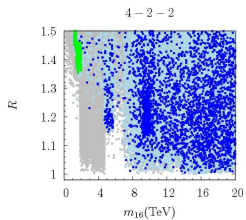
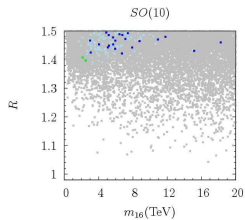
$$0.53 < \frac{BR(B_u \rightarrow \tau \nu_\tau)_{MSSM}}{BR(B_u \rightarrow \tau \nu_\tau)_{SM}} < 2.03 \text{ (} 2\sigma \text{)},$$

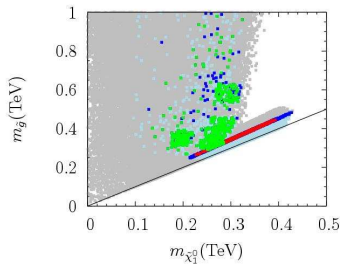
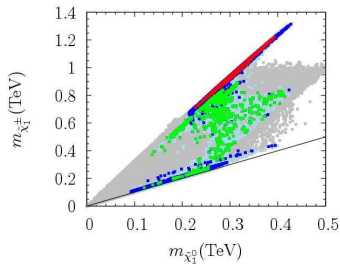
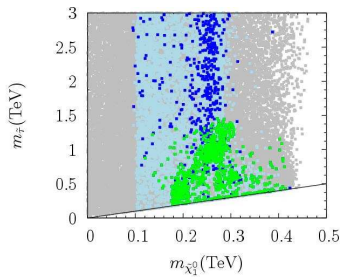
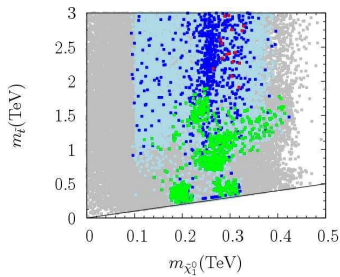
$$2.85 \times 10^{-4} \leq BR(b \rightarrow s \gamma) \leq 4.24 \times 10^{-4} \text{ (} 2\sigma \text{)},$$

$$\Omega_{\text{CDM}} h^2 = 0.111_{-0.037}^{+0.028} \text{ (} 5\sigma \text{)},$$

$$3.4 \times 10^{-10} \leq \Delta\alpha_\mu \leq 55.6 \times 10^{-10} \text{ (} 3\sigma \text{)}.$$







Points in green satisfy all experimental constraints. Red points represent  $R \leq 1.1$ , but do not satisfy g-2.

	Point 1	Point 2	Point 3
$m_{16}$	14110	8429	13124
$M_2$	832.03	1020.2	689.4
$M_3$	0.7945	60.542	9.6261
$\tan \beta$	50.82	46.41	51.17
$M_D/m_{16}$	0.4543	0.5595	0.3323
$m_{10}/m_{16}$	0.7741	1.1584	1.3048
$A_0/m_{16}$	-2.4487	-2.1527	-1.8226
$m_h$	123	126	127
$m_H$	7569	2163	9882
$m_A$	7520	2150	9818
$m_{H^\pm}$	7571	2175	9883
$m_{\tilde{\chi}_{1,2}^\pm}$	<b>887</b> ,13869	<b>975</b> ,4047	<b>712</b> ,3750
$m_{\tilde{\chi}_{1,2}^0}$	<b>283</b> , 885	<b>319</b> ,974	<b>228</b> ,712
$m_{\tilde{\chi}_{3,4}^0}$	13879,13879	4049,4049	3784,3785
$m_{\tilde{g}}$	<b>325</b>	<b>365</b>	<b>265</b>
$m_{\tilde{u}_{L,R}}$	14126,13916	8435,8361	13140,12841
$m_{\tilde{t}_{1,2}}$	5337,5726	<b>1911</b> ,2640	4931,5310
$m_{\tilde{d}_{L,R}}$	14126,14203	8435,8455	13141,13249
$m_{\tilde{b}_{1,2}}$	5237,5653	2521,2767	4115,5146
$m_{\tilde{\nu}_1}$	13988	8409	12926
$m_{\tilde{\nu}_3}$	10598	6577	9535
$m_{\tilde{e}_{L,R}}$	13988,14376	8408,8514	12926,13500
$m_{\tilde{\tau}_{1,2}}$	6412,10581	4270,6573	5580,9559
$\mu$	14100	4110	3840
$\Omega_{LSP} h^2$	0.095	0.112	0.116
$R$	<b>1.00</b>	1.07	1.09

## Yukawa unification with negative $\mu$ term

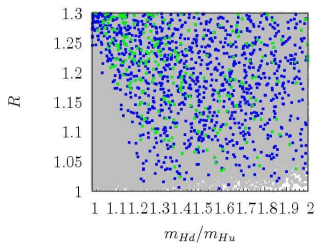
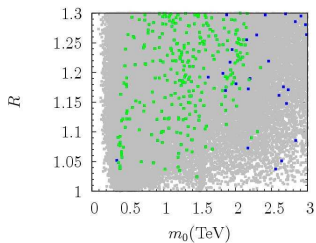
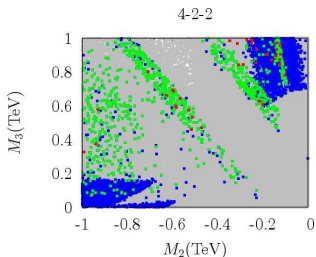
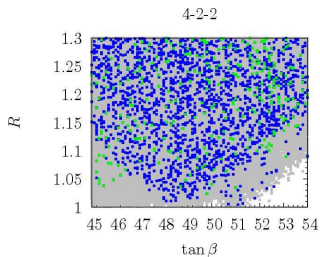
- Yukawa unification prefers  $\mu < 0$
- Dominant contributions to the bottom quark mass from the gluino and chargino loop

$$\delta m_b \approx \frac{g_3^2}{12\pi^2} \frac{\mu m_{\tilde{g}} \tan \beta}{m_b^2} - \frac{y_t^2}{32\pi^2} \frac{\mu A_t \tan \beta}{m_{\tilde{t}}^2} + \dots$$

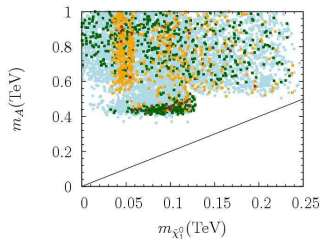
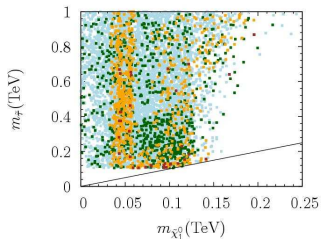
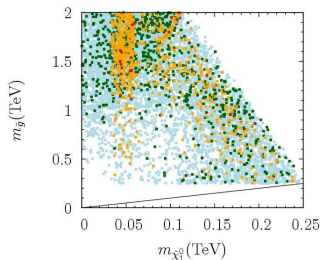
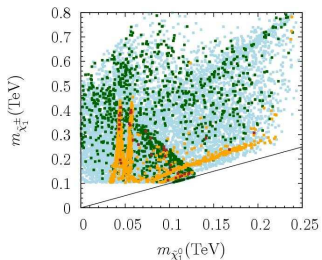
- Dominant contribution to the muon anomalous magnetic moment for large  $\tan \beta$  case is  $\Delta a_\mu^{SUSY} \propto \mu M_2 \tan \beta / \tilde{m}^4$
- In 4-2-2 model with left-right symmetry,  $M_2$  and  $M_3$  are free parameters
- We can have  $\mu < 0$ ,  $M_2 < 0$  and  $M_3 > 0$

We performed random scans for the following parameter range

$$\begin{aligned} 0 &\leq m_0, M_{H_u}, M_{H_d} \leq 20 \text{ TeV}, \\ 1 \text{ TeV} &\leq M_2 \leq 1 \text{ TeV}, \\ 0 &\leq M_3 \leq 1 \text{ TeV}, \\ -3 &\leq A_0/m_{16} \leq 3, \\ 45 &\leq \tan \beta \leq 55, \\ \mu &> 0, \quad \mu < 0, \quad m_t = 172.6 \text{ GeV}. \end{aligned}$$

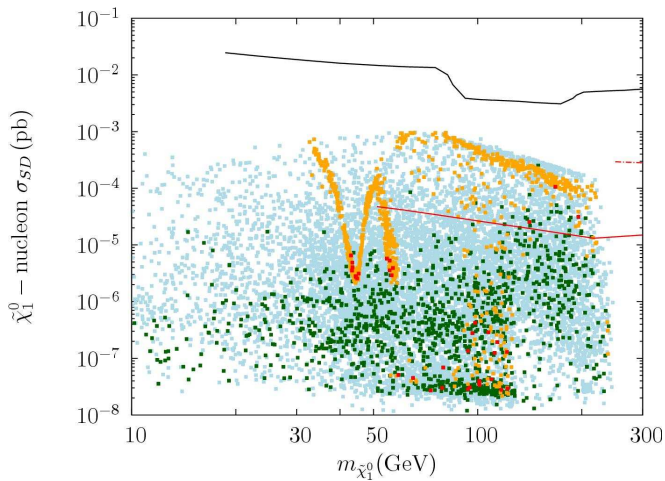


Green points satisfy all constraints. Points in red represent  $R \leq 1.1$



Brown points satisfy all constraints and  $R \leq 1.1$

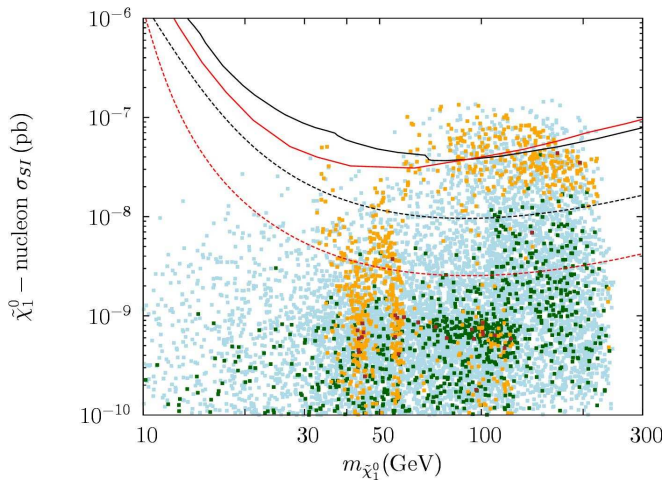
## Dark matter indirect detection



Brown points satisfy all constraints and  $R \leq 1.1$



## Dark matter direct detection



Brown points satisfy all constraints and  $R \leq 1.1$

	Point 1	Point 2	Point 3	Point 4	Point 5
$m_0$	1027	1800	1210	980	1720
$M_1$	-665	-81	-414	-126	-538
$M_2$	-1475	-543	-940	-517	-943
$M_3$	550	611	374	460	70
$\tan \beta$	49.1	52.8	50.6	47.0	47.6
$A_0/m_0$	0.26	1.06	-1.15	-1.08	-1.25
$m_{Hu}$	743	1919	1231	1090	295
$m_{Hd}$	1505	2395	1745	1869	1729
$m_h$	114	115	114	115	115
$m_H$	847	573	781	1100	1006
$m_A$	841	569	776	1090	1000
$m_{H^\pm}$	852	581	787	1100	1010
$m_{\tilde{\chi}_{1,2}^0}$	280,341	43,352	168,242	56,337	233,782
$m_{\tilde{\chi}_{3,4}^0}$	352,1236	380,513	246,795	371,476	1210,1216
$m_{\tilde{\chi}_{1,2}^\pm}$	342,1225	355,509	239,786	338,475	782,1217
$m_{\tilde{g}}$	1321	1470	955	1110	270
$m_{\tilde{u}_{L,R}}$	1771,1489	2170,2130	1550,1410	1400,1320	1818,1697
$m_{\tilde{t}_{1,2}}$	1053,1410	1400,1440	822,1040	826,965	1070,1248
$m_{\tilde{d}_{L,R}}$	1773,1512	2180,2160	1550,1440	1400,1370	1820,1730
$m_{\tilde{b}_{1,2}}$	954,1399	1350,1430	774,1020	724,906	992,1245
$m_{\tilde{\nu}_1}$	1391	1810	1340	1000	1807
$m_{\tilde{\nu}_3}$	1211	1420	1100	759	1550
$m_{\tilde{e}_{L,R}}$	1393,1096	1820,1820	1340,1250	1010,1040	1809,1763
$m_{\tilde{\tau}_{1,2}}$	500,1212	885,1420	641,1110	462,765	1170,1554
$\sigma_{SI}(\text{pb})$	$4.02 \times 10^{-8}$	$4.1 \times 10^{-9}$	$4.1 \times 10^{-8}$	$9.5 \times 10^{-10}$	$1.1 \times 10^{-10}$
$\sigma_{SD}(\text{pb})$	$8.4 \times 10^{-5}$	$7.5 \times 10^{-6}$	$1.7 \times 10^{-4}$	$8.2 \times 10^{-6}$	$2.9 \times 10^{-8}$
$\Omega_{CDM} h^2$	0.08	0.11	0.09	0.08	0.11
$R$	1.01	1.11	1.09	1.07	1.08
$g_3/g_1(M_{GUT})$	0.98	0.98	0.99	0.98	1.00

## Yukawa Unification & NLSP gluino search at Hadron Colliders

- Yukawa unification predicts light gluino, heavy scalars and is compatible with gluino-bino coannihilation with gluino as NLSP.

- Conventional gluino searches with small SM background

$$\tilde{g}\tilde{g} \rightarrow jets + \tilde{\chi}_1^\pm \tilde{\chi}_1^\pm \rightarrow jets + l^\pm l^\pm + \cancel{E}_T.$$

- For NLSP gluino these channels are absent and we consider the parameter space region with dominant contributions from gluino three body decay  $b\bar{b}\tilde{\chi}_1^0$

$$pp, p\bar{p} \rightarrow \tilde{g}\tilde{g} \rightarrow b\bar{b}b\bar{b} + \cancel{E}_T.$$

- Potential SM backgrounds considered

$$b\bar{b}b\bar{b}, b\bar{b}b\bar{b}Z \rightarrow b\bar{b}b\bar{b}\nu\bar{\nu}, jjb\bar{b}Z \rightarrow jjb\bar{b}\nu\bar{\nu}$$

- We choose two benchmark points from previously described 4-2-2 models

	$M_{\tilde{g}}$ (GeV)	$M_{\tilde{\chi}_1^0}$ (GeV)	$M_{\tilde{b}_1}$ (GeV)	$\text{Br}(\tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0)$
Model A ( $\mu > 0$ )	329	284	5294	76.3%
Model B ( $\mu < 0$ )	261	207	950	50.8%

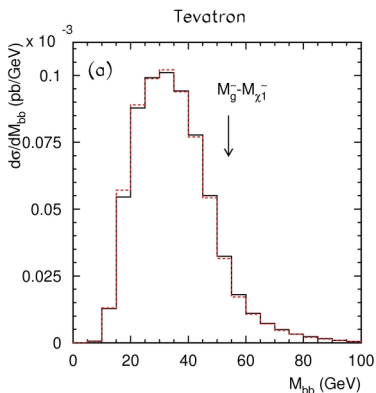
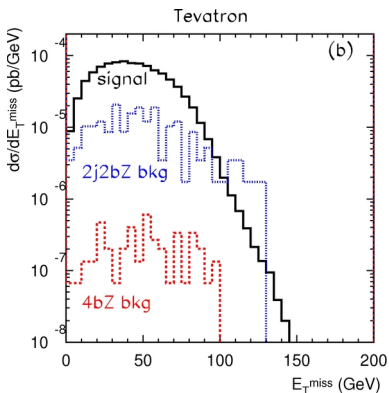
- For Tevatron, we employ the following event selection cuts

$$p_T^j > 15 \text{ GeV}, |\eta_j| < 1.0, \Delta R_{jj} > 0.4$$

and b tagging efficiency 50% and  $\cancel{E}_T > 30 \text{ GeV}$  cut for  $\tilde{\chi}_1^0$ .

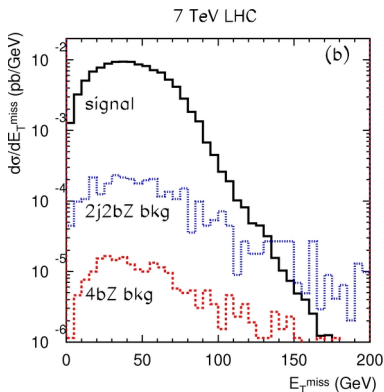
- The production cross section for the two points for Tevatron

$\sigma(\text{fb})@ \text{Tevatron}$	Model A	Model B	$b\bar{b}b\bar{b}$	$b\bar{b}b\bar{b}Z$	$jjb\bar{b}Z$	$S/\sqrt{B}$
basic cuts and 3b tagging	2.3	4.8	$2.7 \times 10^3$	0.02	1	
$\cancel{E}_T > 30 \text{ GeV}$	1.4	3.3	—	0.019	0.95	4.5(A)/11(B)



- The production cross section for the two points for LHC

$\sigma(\text{fb})@ 7 \text{ TeV LHC}$	Model A	Model B		$b\bar{b}b\bar{b}$	$b\bar{b}b\bar{b}Z$	$j\bar{j}b\bar{b}Z$
basic cuts and 3b tagging	286	541		$314 \times 10^3$	1.1	15
$\cancel{E}_T > 40 \text{ GeV}$	117	280		—	0.8	12



## Summary

- In supersymmetric and L-R symmetric  $SU(4)_c \times SU(2)_L \times SU(2)_R$  model with gravity mediated supersymmetry breaking,  $t - b - \tau$  Yukawa coupling unification is consistent with neutralino dark matter abundance and with all constraints from collider experiments (except  $(g - 2)_\mu$ ) for  $\mu > 0$ . For  $\mu < 0$  we can have Yukawa unification satisfying all current constraints.
- The model for  $\mu > 0$  predicts a very characteristic sparticle spectrum: very heavy sfermions ( $> 5 \text{ TeV}$ ) but relatively light gluinos ( $\gtrsim 300 \text{ GeV}$ ).
- For  $\mu < 0$ , Yukawa unification can be achieved with relatively light sparticle spectrum  $O(600) \text{ GeV}$ . NLSP gluino can be tested at LHC/Tevatron.

- NLSP gluino search at Tevatron and 7 TeV LHC through multi-b jets  $\tilde{g}\tilde{g} \rightarrow b\bar{b}b\bar{b}\tilde{\chi}_1^0\tilde{\chi}_1^0$ .
- With  $10 \text{ fb}^{-1}$  luminosity one can reach  $5\sigma$  at Tevatron after selection cuts. At 7 TeV LHC the signal is at least **one order of magnitude larger** than leading backgrounds.