## Lessons from Isobar and Event Shape Selection Analysis for Search of CME with RHIC Beam Energy Scan Data

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## Chiral Magnetic Effect



- Chirality imbalance coupled with strong magnetic field induces a charge separation along the B field direction (violates Parity Symmetry dynamically in strong interaction!)
- Heavy-Ion Collision provides an opportunity to observe an intrinsic QCD toplogical phenomenon experimentally.


## Chiral Magnetic Effect



- To quantify the collective motions including the charge separation, we expand the particle azimuthal angle distribution as:

$$
\frac{d N_{ \pm}}{d \phi} \propto 1+2 v_{1} \cos \left(\phi-\Psi_{\mathrm{RP}}\right)+2 v_{2} \cos \left[2\left(\phi-\Psi_{\mathrm{RP}}\right)\right]+\cdots+2 a_{ \pm} \sin \left(\phi-\Psi_{\mathrm{RP}}\right)+\cdots,
$$

- Experiment observable:

$$
\Delta \gamma^{\mathrm{CME}}=\gamma^{\mathrm{OS}}-\gamma^{\mathrm{SS}}>0
$$



## Lessons from Isobar Collisions

- Isobar data did not observe the predefined CME signatures.
- Why? BKG difference: multiplicity mismatch. $\mathrm{v}_{2}{ }^{\sim} 2 \%$



Isobar collisions sensitive to potential difference in CME observable due to different Magnetic Field
Ru+Ru/Zr+Zr isobars are not exactly the same within sub-percent level to achieve the accuracy needed
These isobars are also small in Z, not favorable for CME searches
Isobar data $\rightarrow$ CME signal is probably small, not necessarily zero ! $\rightarrow$ how small?

## Event Shape Selection (Engineering)

## Large Colliding Nuclei $\mathrm{Pb}+\mathrm{Pb}$ at LHC and $\mathrm{Au}+\mathrm{Au}$ at RHIC




Unable to select a spherical event shape sample AND with finite Magnetic Field to search for CME
$q_{2}$ from an eta region different from particle of interest
So selected $\mathrm{q}_{2}$ not effective in selecting the shape of emission for particles of interest
If the selection is such that $\mathrm{v} 2 \rightarrow 0$ when the events corresponding to the most central collisions, then the number of spectator protons are minimum and not favorable for CME

## Event Shape Selection to control $\mathrm{v}_{2}$ ?

q2 or v2 has contributions:
participant shape distribution - likely long range and correlated over large eta gaps emission pattern fluctuations - short eta range, uncorrelated for different eta regions
$0 \longleftarrow\left\langle v_{2}\right\rangle$
CME Search: POI - spherical B field -- finite


## Event Shape Selection Variables

- Single $q^{2}$ and single v2 are constructed from final particles
- Pair $q^{2}$ and pair $\mathrm{v}_{2}$ are calculated based on pair momentum
- Pair momentum is obtained from adding momenta of two particles to mimic decay kinematics $\quad \sim$ more related to $\Delta \gamma$ background


Event shape variables
Elliptic flow variables

(b)

## Origins of BKG $\sim \mathrm{v}_{2}{ }^{\text {res }}$

- The BKG from resonance flowing decay in $\Delta \gamma$ is well-represented by product of


$$
\Delta \gamma\{B G\}=v_{2}^{r e s}\left\langle\cos \left(\phi_{\alpha}+\phi_{\beta}-2 \Psi_{R P}\right)\right\rangle
$$

- A toy model targeted $\rho \rightarrow \pi^{+} \pi^{-}$resonance decay confirms the above relation.
- Can we use $\mathrm{V}_{2}{ }^{\text {res }}$ directly to control the BG ?
- NO. Why? $\mathrm{v}_{2}{ }^{\text {res }}$ is modified by the CME existence.


## Optimal ESS Approach





- Resonance $\mathbf{v}_{2}{ }^{\text {res }}$ is significantly modified under the CME.
- The increase is proportional to $\mathbf{a}_{1}{ }^{2}$
- Single $\mathrm{v}_{2}$ and pair $\mathrm{v}_{2}$ are almost constant.


## Optimal ESS Approach



Event shape variables
(a)


Elliptic flow variables
(b)

- Unmixed recipes cause residual background near zero-flow region
- Mixed recipes have advantage that the $\mathrm{v}_{2}$ and binning $\mathrm{q}^{2}$ are less correlated.

。However, pair v2 contains true CME signal, which may lead to over subtraction.
Scenario (c) - pair $\mathrm{q}^{2}$, single $\mathrm{v}_{2}$ is the optimal solution.

## ESS for AVFD Events: n5/s = 0 (pure BKG)

- AVFD model confirms possible residual background in (a) and (b) unmixed recipes.
- Mixed combination can remove residual BKG and predict over-subtraction in scenario (d)
- ESS (c) using single v2 and binning by pair q2 can well reproduce BKG.



## ESS for AVFD: $\mathrm{n}_{5} / \mathrm{s}=0.1$ (moderate CME)

- With CME signal, residual background preserves in (a) and (b) unmixed recipes.
- AVFD confirms that ESS (c) suppress the residual BKG, and successfully match the true signal.
- Over-subtraction of BKG as predicted in (d) projection






## Why Over-subtraction in ESS using $\mathrm{v}_{\mathbf{2}}$ (resonance)

- Using $\mathrm{v}_{2}{ }^{\text {res }}$ will cause severe over-subtraction.
- Explains that ESS (d) pair v2 that contains possible CME signal also cause over-subtraction.




## Using separate region Qb - Not Effective

- Separate region $\mathrm{q}_{2}$ has weak correlation with the POl's $\mathrm{V}_{2}$
- This cause a gap in $\Delta \gamma-v_{2}$ plot that leads to less reliable results:
- Statistic errors are 3 times larger than ESS involving POI.
- Systematic uncertainties demonstrated by 2nd-order polynomial have large variation, even exceed statistic errors.





## CME changes the invariant mass distribution



There is no clear region of signal vs background in invariant mass distribution

## Application to STAR data



- Both ESS approaches can extrapolate $\Delta \gamma_{E S S}^{112}=\left(1-2 v_{2}\right) \cdot$ Intercept




## 27 GeV : EPD spectator plane




- Spectator plane is more correlated to the magnetic field direction.
- Finite $\Delta \gamma_{E S S}^{112}$ in mid central events. $\Delta \gamma_{E S S}^{132}$ consistent with zero for all centralities.
- The precision of STAR measurement after ESS is controlled to be $5.4 \%$ of ensemble average $\Delta \gamma^{112}$.


### 19.6 GeV : EPD spectator plane




- Spectator plane is more correlated to the magnetic field direction.
- Finite $\Delta \gamma_{E S S}^{112}$ in mid central events; $\Delta \gamma_{E S S}^{132}$ consistent with zero for all centralities.
- The precision of STAR measurement after ESS $i_{1}$ controlled to be $3.6 \%$ of ensembled average $\Delta \gamma^{112}$.


## Perspectives

- Resonance v2 turns out to be a CME sensitive observable
- We developed an optimized Event Shape Selection method -single v2 and pair $q^{2}$, that utilize pair particle information to further suppress residual BKG
- We demonstrate that event shape selection (ESS) approach substantially suppresses (over five-fold) $\nu_{2}$ related backgrounds, enhancing the CME search sensitivity considerably.
- Using 1st-order EPD spectator plane, we can achieve a 4-5\% precision in ESS measurement of $\Delta \gamma^{112}\{h h\}$ in RHIC' s BES-II data.



