Lessons from Isobar and Event Shape Selection Analysis for Search of CME with RHIC Beam Energy Scan Data Huan Zhong Huang (黄焕中)

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Chiral Magnetic Effect





- along the B field direction (violates Parity Symmetry dynamically in strong interaction!)
- phenomenon experimentally.



• Chirality imbalance coupled with strong magnetic field induces a charge separation

• Heavy-Ion Collision provides an opportunity to observe an intrinsic QCD toplogical

Chiral Magnetic Effect



- Experiment observable: $\Delta \gamma^{\rm CME}$

 $\gamma^{112} = \langle \cos(\phi_1 + \phi_2 - 2\psi_{\rm RP}) \rangle = \langle \cos(\phi_1 + \phi_2 - 2\psi_{\rm RP}) \rangle$ Non-CME Decay of flowing resona \mathcal{P}_2

Flowing resonance decay



D. E. Kharzeev, J. Liao, S. A. Voloshin, and G. Wang, Prog. Part. Nucl. Phys. 88, 1 (2016).

• To quantify the collective motions including the charge separation, we expand the particle azimuthal angle distribution as:

 $\frac{dN_{\pm}}{d\phi} \propto 1 + 2v_1 \cos(\phi - \Psi_{\rm RP}) + 2v_2 \cos[2(\phi - \Psi_{\rm RP})] + \dots + 2a_{\pm} \sin(\phi - \Psi_{\rm RP}) + \dots$

$$\Delta \gamma^{\text{CME}} = \gamma^{\text{OS}} - \gamma^{\text{SS}} > 0$$

$$(\alpha_{RP}) = \langle \cos(\phi_1 - \psi_{RP})\cos(\phi_2 - \psi_{RP}) \rangle - \langle \sin(\phi_1 - \psi_{RP})\sin(\phi_2 - \psi_{RP}) \rangle$$

$$(\alpha_{RP}) = \gamma^{\text{OS}} - \gamma^{\text{SS}} \propto \frac{v_2}{N}$$

$$\Delta \gamma^{\text{reso}} = \gamma^{\text{OS}} - \gamma^{\text{SS}} \propto \frac{v_2}{N}$$

$$\Delta \gamma^{112} = \Delta \gamma^{\text{CME}} + k \frac{v_2}{N} + \Delta \gamma^{\text{non-flow}}$$

$$(\alpha_{RP}) = \lambda^{\text{Measured}}$$

$$(\alpha_{RP}) = \lambda^{\text{SS}} =$$

 $\Delta \gamma^{\rm reso}$

(RP)

Lessons from Isobar Collisions

- Isobar data did not observe the predefined CME signatures.
- Why? BKG difference: multiplicity mismatch. $v_2 \sim 2\%$



- Isobar collisions sensitive to potential difference in CME observable due to different Magnetic Field Ru+Ru/Zr+Zr isobars are not exactly the same within sub-percent level to achieve the accuracy needed These isobars are also small in Z, not favorable for CME searches
 - Isobar data \rightarrow CME signal is probably small, not necessarily zero ! \rightarrow how small?

Phys. Rev. C 105, 014901

Event Shape Selection (Engineering) Phys. Lett. B 777, 151 (2018) Large Colliding Nuclei Pb+Pb at LHC and Au+Au at RHIC



q₂ from an eta region different from particle of interest

So selected q₂ not effective in selecting the shape of emission for particles of interest

If the selection is such that $v2 \rightarrow 0$ when the events corresponding to the most central collisions, then the number of spectator protons are minimum and not favorable for CME

γ_{ab} *dN/d η (opp-same) vs v_2

Unable to select a spherical event shape sample **AND** with finite **Magnetic Field** to search for CME

Event Shape Selection to control v₂?

q2 or v2 has contributions:





participant shape distribution – likely long range and correlated over large eta gaps emission pattern fluctuations – short eta range, uncorrelated for different eta regions $0 \leftarrow \langle v_2 \rangle$



Event Shape Selection Variables

- $^{\bullet}$ Single q^2 and single v_2 are constructed from final particles
- Pair q^2 and pair v_2 are calculated based on pair momentum





Which one of the above four ESS is the best?



Origins of BKG ~ V₂^{res}

• The BKG from resonance flowing decay in $\Delta \gamma$ is well-represented by product of



S. A. Voloshin, Phys. Rev. C 70, 057901 (2004)

$$\Delta\gamma\{BG\} = v_2^{res} \langle \cos(\phi_{\alpha} + \phi_{\beta} - 2\Psi_{RP}) \rangle$$

• A toy model targeted $\rho \rightarrow \pi^+ \pi^-$ resonance decay confirms the above relation.

• Can we use v_2^{res} directly to control the BG?

• NO. Why? v_2^{res} is modified by the CME existence.



Optimal ESS Approach



- Resonance v₂^{res} is significantly modified under the CME.
- The increase is proportional to a_1^2
- Single v₂ and pair v₂ are almost constant.







- Unmixed recipes cause residual background near zero-flow region 0
- Mixed recipes have advantage that the v_2 and binning q^2 are less correlated. 0

However, pair v2 contains true CME signal, which may lead to over subtraction.

Scenario (c) – pair q^2 , single v_2 is the optimal solution.

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ESS for AVFD Events: $n_5/s = 0$ (pure BKG)

- AVFD model confirms possible residual background in (a) and (b) unmixed recipes.
- Mixed combination can remove residual BKG and predict over-subtraction in scenario (d)
- using single v2 and • ESS **(C)** binning by pair q2 can well reproduce BKG.







ESS for AVFD: $n_5/s = 0.1$ (moderate CME)

- With CME signal, residual background preserves in (a) and (b) unmixed recipes.
- AVFD confirms that ESS (c) suppress the residual BKG, and successfully match the true signal.
- Over-subtraction of BKG as predicted in (d) projection





Why Over-subtraction in ESS using v₂(resonance)

- Using v_2^{res} will cause severe over-subtraction.
- Explains that ESS (d) pair v2 that contains possible CME signal also cause over-subtraction.



Using separate region Qb – Not Effective

- Separate region q₂ has weak correlation with the POI's v₂
- This cause a gap in $\Delta \gamma$ v₂ plot that leads to less reliable results:
 - Statistic errors are 3 times larger than ESS involving POI.
 - Systematic uncertainties demonstrated by 2nd-order polynomial have large variation, even exceed statistic errors.

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CME changes the invariant mass distribution

There is no clear region of signal vs background in invariant mass distribution

- Spectator plane is more correlated to the magnetic field direction.
- Finite $\Delta \gamma_{ESS}^{112}$ in mid central events. $\Delta \gamma_{ESS}^{132}$ consistent with zero for all centralities.
- The precision of STAR measurement after ESS is controlled to be 5.4% of ensemble average $\Delta \gamma^{112}$.

19.6 GeV : EPD spectator plane

- Spectator plane is more correlated to the magnetic field direction.
- Finite $\Delta \gamma_{ESS}^{112}$ in mid central events; $\Delta \gamma_{ESS}^{132}$ consistent with zero for all centralities.

• The precision of STAR measurement after ESS i_1 s controlled to be 3.6% of ensembled average $\Delta \gamma^{112}$.

Perspectives

- Resonance v_2 turns out to be a CME sensitive observable
- We developed an optimized Event Shape Selection method —single v_2 and pair q^2 , that utilize pair particle information to further suppress residual BKG
- of $\Delta \gamma^{112} \{hh\}$ in RHIC' s BES-II data.

• We demonstrate that event shape selection (ESS) approach substantially suppresses (over five-fold) v_2 related backgrounds, enhancing the CME search sensitivity considerably.

Using 1st-order EPD spectator plane, we can achieve a 4-5% precision in ESS measurement

$$\Delta \gamma^{112} \xrightarrow{\text{ESS}} \text{spectator plan}$$
$$= |\Delta \gamma^{\text{CME}}| + k \frac{\nu_2}{N} + \Delta \chi^{\text{non-flow}}$$

Measured Signal Backgrounds

Thank you!