# Spin transport (and hydrodynamics) with nonlocal collisions 

David Wagner<br>in collaboration with

Nora Weickgenannt, Enrico Speranza, and Dirk Rischke
based mainly on

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NW, ES, X.-L. Sheng, Q. Wang, DHR, Phys.Rev.D 104 (2021) 1, }01602
    NW, DW, ES, DHR, Phys.Rev.D 106 (2022) 9, 096014
        DW, NW, ES, Phys.Rev.Res. }5\mathrm{ (2023) 1,013187
        DW, NW, DHR, Phys.Rev.D 106 (2022) 11, 116021
        DW, NW, ES, 2306.05936 (2023)
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Chirality, Vorticity \& Magnetic Field in HIC | 17.07.2023

## Global $\Lambda$-polarization

- Global polarization: polarization of $\Lambda$-hyperons along angular-momentum direction

L. Adamczyk et al. (STAR), Nature 548 (2017) 62


## Global $\Lambda$-polarization

- Global polarization: polarization of $\Lambda$-hyperons along angular-momentum direction
- Can be well explained by considering local equilibrium on freeze-out hypersurface $S_{\varpi}^{\mu}=-\epsilon^{\mu \nu \alpha \beta} k_{\nu} \frac{\int \mathrm{d} \Sigma_{\lambda} k^{\lambda} f_{0}\left(1-f_{0}\right) \varpi_{\alpha \beta}}{8 m \int \mathrm{~d} \Sigma_{\lambda} k^{\lambda} f_{0}}$

L. Adamczyk et al. (STAR), Nature 548 (2017) 62

$$
\varpi_{\mu \nu}:=-\frac{1}{2}\left(\partial_{\mu} \beta_{\nu}-\partial_{\nu} \beta_{\mu}\right), \beta^{\mu}:=u^{\mu} / T, f_{0}=\left[\exp \left(u^{\mu} k_{\mu} / T\right)+1\right]^{-1}
$$

## Local $\Lambda$-polarization

- Local polarization:

Angle-dependent polarization of $\Lambda$-hyperons along beam-direction

F. Becattini, M. Buzzegoli, G. Inghirami, I. Karpenko, A.

Palermo, PRL 127 (2021) 272302
B. Fu, S. Y. F. Liu, L. Pang, H. Song, Y. Yin, PRL 127
(2021) 142301

## Local $\Lambda$-polarization

- Local polarization:

Angle-dependent polarization of $\Lambda$-hyperons along beam-direction

■ Could only be explained recently by incorporating shear effects (neglecting temperature gradients) $S_{\xi}^{\mu}=-\epsilon^{\mu \nu \alpha \beta} k_{\nu} \frac{\int \mathrm{d} \Sigma_{\lambda} k^{\lambda} f_{0}\left(1-f_{0}\right) \hat{t}_{\alpha} \frac{k^{\gamma}}{k^{0}} \Xi_{\gamma \beta}}{4 m T \int \mathrm{~d} \Sigma_{\lambda} k^{\lambda} f_{0}}$

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\omega_{\mu \nu}:=\frac{1}{2}\left(\partial_{\mu} u_{\nu}-\partial_{\nu} u_{\mu}\right), \Xi_{\mu \nu}:=\frac{1}{2}\left(\partial_{\mu} u_{\nu}+\partial_{\nu} u_{\mu}\right), \Delta^{\mu \nu}:=g^{\mu \nu}-u^{\mu} u^{\nu}
$$

## Alignment of $\phi$-mesons

- Spin-1 particles feature tensor polarization ( $\hat{=}$ alignment)


STAR collaboration, arXiv:2204.02302 (2022)

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- Spin-1 particles feature tensor polarization ( $\hat{=}$ alignment)
- Larger than expected
- Some theoretical developments, but no definitive answer yet
X.-L. Xia, H. Li, X-G. Huang, H.-Z. Huang,

PLB 817 (2021) 136325
X.-L. Sheng, L. Oliva, Z.-T. Liang, Q. Wang,
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- Can spin- 1 hydrodynamics help explain this?


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## Spin Hydrodynamics: Basics

- Hydrodynamics is based on conservation laws
- Consider a system of uncharged fields
$\rightarrow$ Should conserve energy-momentum and total angular momentum


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## Conservation laws

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\partial_{\mu} T^{\mu \nu} & =0  \tag{1a}\\
\partial_{\lambda} J^{\lambda \mu \nu}=: \partial_{\lambda} S^{\lambda \mu \nu}+T^{[\mu \nu]} & =0 \tag{1b}
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A^{[\mu} B^{\nu]}:=A^{\mu} B^{\nu}-A^{\nu} B^{\mu}
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- 10 equations for $16+24$ quantities
- Additional information about dissipative quantities has to be provided $\rightarrow$ Use kinetic theory with spin as effective microscopic model
- Rest of the presentation:
- Construct such a kinetic theory
- Perform hydrodynamic limit
- Obtain expressions for observables

$$
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## How to: Quantum kinetic theory

- Spin is a quantum property
$\rightarrow$ Start from quantum field theory
$\rightarrow$ Use Wigner-function formalism


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## Wigner function (Spin 1)

$$
W^{\mu \nu}(x, k):=-\frac{2}{(2 \pi \hbar)^{4} \hbar} \int \mathrm{~d}^{4} v e^{-i k \cdot y / \hbar}\left\langle: V^{\dagger \mu}(x+y / 2) V^{\nu}(x-y / 2):\right\rangle
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- Determines a quantum phase-space distribution function
- Equations of motion follow from field equations
- Determined by Lagrangian $\mathcal{L}_{0}+\mathcal{L}_{\text {int }}$


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- Determines a quantum phase-space distribution function
- Equations of motion follow from field equations
- Determined by Lagrangian $\mathcal{L}_{0}+\mathcal{L}_{\text {int }}$
- Independent components: scalar $f_{K}$, axial vector $G^{\mu}$ and traceless symmetric tensor $F_{K}^{\mu \nu}$

$$
\begin{aligned}
f_{K} & :=(1 / 3) K_{\mu \nu} W^{\mu \nu}, G^{\mu}:=-(i / 2 m) \epsilon^{\mu \nu \alpha \beta} k_{\nu} W_{\alpha \beta}, F_{K}^{\mu \nu}:=K_{\alpha \beta}^{\mu \nu} W^{\alpha \beta} \\
K^{\mu \nu} & :=g^{\mu \nu}-k^{\mu} k^{\nu} / m^{2}, K_{\alpha \beta}^{\mu \nu}:=\left(K_{\alpha}^{\mu} K_{\beta}^{\nu}+K_{\beta}^{\mu} K_{\alpha}^{\nu}\right) / 2-1 / 3 K^{\mu \nu} K_{\alpha \beta}
\end{aligned}
$$

## Extending phase space

## Boltzmann equations

- Not one, but nine equations in $(x, k)$-phase space

$$
k \cdot \partial f_{K}(x, k)=\mathcal{C}_{K}, \quad k \cdot \partial G^{\mu}(x, k)=\mathcal{C}_{G}^{\mu}, \quad k \cdot \partial F_{K}^{\mu \nu}(x, k)=\mathcal{C}_{K}^{\mu \nu}
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- Way to compactify this: Enlarge phase space from $(x, k)$ to $(x, k, \mathfrak{s})$
- Measure $\mathrm{d} S:=\frac{3 m}{2 \sigma \pi} \mathrm{~d}^{4} \mathfrak{s} \delta\left[\mathfrak{s}^{2}+\sigma^{2}\right] \delta(k \cdot \mathfrak{s})$


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## Boltzmann equation in extended phase space

$$
\begin{equation*}
\mathfrak{f}(x, k, \mathfrak{s}):=f_{K}-\mathfrak{s}_{\mu} G^{\mu}+\frac{5}{4} \mathfrak{s}_{\mu} \mathfrak{s}_{\nu} F_{K}^{\mu \nu} \tag{2}
\end{equation*}
$$

- Only on-shell parts $\mathfrak{f}(x, k, \mathfrak{s})=\delta\left(k^{2}-m^{2}\right) f(x, k, \mathfrak{s})$ contribute

$$
\begin{equation*}
k \cdot \partial f(x, k, \mathfrak{s})=\mathfrak{C}[f] \tag{3}
\end{equation*}
$$

$$
\mathfrak{C}:=\mathcal{C}_{K}-\mathfrak{s}_{\mu} \mathcal{C}_{G}^{\mu}+\frac{5}{4} \mathfrak{s}_{\mu} \mathfrak{s}_{\nu} \mathcal{C}_{K}^{\mu \nu}
$$

## Collision term

## Collision kernel

$$
\begin{align*}
\mathfrak{C}[f]= & \frac{1}{2} \int \mathrm{~d} \Gamma_{1} \mathrm{~d} \Gamma_{2} \mathrm{~d} \Gamma^{\prime} \mathrm{d} \bar{S}(k) \delta^{(4)}\left(k_{1}+k_{2}-k-k^{\prime}\right) \mathcal{W} \\
& \times\left[f\left(x+\Delta_{1}-\Delta, k_{1}, \mathfrak{s}_{1}\right) f\left(x+\Delta_{2}-\Delta, k_{2}, \mathfrak{F}_{2}\right)\right. \\
& \left.-f(x, k, \overline{\mathfrak{s}}) f\left(x+\Delta^{\prime}-\Delta, k^{\prime}, \mathfrak{s}^{\prime}\right)\right] \tag{4}
\end{align*}
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$$
\mathrm{d} \Gamma:=2 \mathrm{~d}^{4} k \delta\left(k^{2}-m^{2}\right) \mathrm{d} S(k)
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- Contributions inside the collision term have gradient corrections

$$
\begin{equation*}
f(x, k, \mathfrak{s})+\Delta^{\mu} \partial_{\mu} f(x, k, \mathfrak{s}) \approx f(x+\Delta, k, \mathfrak{s}) \tag{5}
\end{equation*}
$$

- A (momentum- and spin-dependent) spacetime shift $\Delta^{\mu}$ enters $\rightarrow$ Particles do not scatter at the same spacetime point!

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- A (momentum- and spin-dependent) spacetime shift $\Delta^{\mu}$ enters $\rightarrow$ Particles do not scatter at the same spacetime point!
- This enables a conversion of orbital and spin angular momenta

$$
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$$

## Nonlocal collisions

## Spacetime shifts

$$
\begin{align*}
\Delta^{\mu}:= & \frac{1}{3} \frac{1}{\mathcal{W}} \frac{(2 \pi \hbar)^{3}}{32} \frac{i \hbar}{m^{2}} M^{\gamma_{1} \gamma_{2} \delta_{1} \delta_{2}} M^{\zeta_{1} \zeta_{2} \eta_{1} \eta_{2}} h_{1, \gamma_{1} \eta_{1}} h_{2, \gamma_{2} \eta_{2}} h_{\zeta_{2} \delta_{2}}^{\prime} \\
& \times\left(H^{\mu}{ }_{\delta_{1}} k_{\zeta_{1}}-k_{\delta_{1}} H_{\zeta_{1}}{ }^{\mu}\right) \tag{6}
\end{align*}
$$

- Depend on the transfer-matrix elements

$$
\begin{equation*}
\left\langle 11^{\prime}\right| \widehat{t}\left|22^{\prime}\right\rangle=\epsilon_{1, \alpha}^{*} \epsilon_{1^{\prime}, \beta}^{*} \epsilon_{2, \gamma} \epsilon_{2^{\prime}, \delta} M^{\alpha \beta \gamma \delta} \tag{7}
\end{equation*}
$$

- Manifestly covariant
$\rightarrow$ no "no-jump" frame

$$
\begin{aligned}
h^{\mu \nu} & :=\frac{1}{3} K^{\mu \nu}+\frac{i}{2 m} \epsilon^{\mu \nu \alpha \beta} k_{\alpha} \mathfrak{s}_{\beta}+K_{\alpha \beta}^{\mu \nu} \mathfrak{s}^{\alpha} \mathfrak{s}^{\beta}, \\
H^{\mu \nu} & :=\frac{1}{3} K^{\mu \nu}+\frac{i}{2 m} \epsilon^{\mu \nu \alpha \beta} k_{\alpha} \mathfrak{s}_{\beta}+\frac{5}{8} K_{\alpha \beta^{\prime} \mathfrak{s}^{\alpha} \mathfrak{s}^{\beta}}
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## Local-equilibrium distribution function

$$
\begin{equation*}
f_{\mathrm{eq}}(x, k, \mathfrak{s})=\exp \left(-\beta_{0} E_{\mathbf{k}}+\frac{\hbar}{2} \Omega_{\mu \nu} \Sigma_{\mathfrak{s}}^{\mu \nu}\right) \tag{8}
\end{equation*}
$$

$$
\Sigma_{\mathfrak{s}}^{\mu \nu}:=-\frac{1}{m} \epsilon^{\mu \nu \alpha \beta} k_{\alpha} \mathfrak{s} \beta_{\beta}, E_{\mathbf{k}}:=k \cdot u
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- Necessary conditions on Lagrange multipliers $\beta_{0} u^{\mu}, \Omega^{\mu \nu}$ for a vanishing collision term: $\partial^{(\mu}\left(\beta_{0} u^{\nu)}\right)=0, \Omega^{\mu \nu}=-\frac{1}{2} \partial^{[\mu}\left(\beta_{0} u^{\nu]}\right)$
- Same conditions as for global equilibrium, where $k \cdot \partial f_{\text {eq }}=0$

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\sum_{\mathfrak{s}}^{\mu \nu}:=-\frac{1}{m} \epsilon^{\mu \nu \alpha \beta} k_{\alpha} \mathfrak{s}_{\beta}, E_{\mathbf{k}}:=k \cdot u
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- Same conditions as for global equilibrium, where $k \cdot \partial f_{\text {eq }}=0$
- However, we can relax these constraints if we only demand that the local part of the collision term vanishes!

$$
\sum_{\mathfrak{s}}^{\mu \nu}:=-\frac{1}{m} \epsilon^{\mu \nu \alpha \beta} k_{\alpha} \mathfrak{s}_{\beta}, E_{\mathbf{k}}:=k \cdot u
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## Moment expansion

- Split distribution function $f=f_{\text {eq }}+\delta f$


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## Irreducible moments

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## Irreducible moments

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\begin{equation*}
\rho_{r}^{\mu_{1} \cdots \mu_{\ell}}(x):=\int \mathrm{d} \Gamma E_{\mathbf{k}}^{r} k^{\left\langle\mu_{1}\right.} \cdots k^{\left.\mu_{\ell}\right\rangle} \delta f(x, k, \mathfrak{s}) \tag{9a}
\end{equation*}
$$

$k^{\left\langle\mu_{1}\right.} \cdots k^{\left.\mu_{\ell}\right\rangle}:=\Delta_{\nu_{1} \cdots \nu_{\ell}}^{\mu_{1} \cdots \mu_{\ell}} k^{\nu_{1}} \cdots k^{\nu_{\ell}}$

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\tau_{r}^{\mu, \mu_{1} \cdots \mu_{\ell}}(x) & :=\int \mathrm{d} \Gamma \mathfrak{s}^{\mu} E_{\mathbf{k}}^{r} k^{\left\langle\mu_{1}\right.} \cdots k^{\left.\mu_{\ell}\right\rangle} \delta f(x, k, \mathfrak{s}) \tag{9b}
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\psi_{r}^{\mu \nu, \mu_{1} \cdots \mu_{\ell}}(x) & :=\int \mathrm{d} \Gamma K_{\alpha \beta^{s}}^{\mu \nu} \mathfrak{s}^{\alpha} \mathfrak{s}^{\beta} E_{\mathrm{k}}^{r} k^{\left\langle\mu_{1}\right.} \cdots k^{\left.\mu_{\ell}\right\rangle} \delta f(x, k, \mathfrak{s}) \tag{9c}
\end{align*}
$$

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\end{align*}
$$

- Equations of motion can be derived from Boltzmann equation
- Knowing the evolution of all moments is equivalent to solving the Boltzmann equation

$$
k^{\left\langle\mu_{1}\right.} \cdots k^{\left.\mu_{\ell}\right\rangle}:=\Delta_{\nu_{1} \cdots \nu_{\ell}}^{\mu_{1} \cdots \mu_{\ell}} k^{\nu_{1}} \cdots k^{\nu_{\ell}}
$$

## Polarization observables in kinetic theory

## Vector Polarization (Pauli-Lubanski Pseudovector)

$$
\begin{equation*}
S^{\mu}(k):=\operatorname{Tr}\left[\hat{S}^{\mu} \hat{\rho}(k)\right]=\frac{1}{N(k)} \int \mathrm{d} \Sigma_{\lambda} k^{\lambda} \int \mathrm{d} S(k) \mathfrak{s}^{\mu} f(x, k, \mathfrak{s}) \tag{10}
\end{equation*}
$$

$$
N(k):=\int \mathrm{d} \Sigma_{\gamma} k^{\gamma} \int \mathrm{d} S(k) f(x, k, \mathfrak{s}), \quad \hat{S}^{\mu}:=-(1 / 2 m) \epsilon^{\mu \nu \alpha \beta} \hat{J}_{\nu \alpha} \hat{P}_{\beta}
$$

## Polarization observables in kinetic theory

## Vector Polarization (Pauli-Lubanski Pseudovector)

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$$

## Tensor Polarization

$$
\begin{equation*}
\rho_{00}(k)=\frac{1}{3}-\sqrt{\frac{2}{3}} \epsilon_{\mu}^{(0)}(k) \epsilon_{\nu}^{(0)}(k) \Theta^{\mu \nu}(k) \tag{11a}
\end{equation*}
$$

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\Theta^{\mu \nu}(k) & \left.:=\frac{1}{2} \sqrt{\frac{3}{2}} \operatorname{Tr}\left[\left(\hat{S}^{(\mu} \hat{S}^{\nu}\right)+\frac{4}{3} K^{\mu \nu}\right) \hat{\rho}(k)\right] \\
& =\frac{1}{2} \sqrt{\frac{3}{2}} \frac{1}{N(k)} \int \mathrm{d} \Sigma_{\lambda} k^{\lambda} \int \mathrm{d} S(k) K_{\alpha \beta^{\prime}}^{\mu \nu} \mathfrak{s}^{\alpha} \beta \tag{11b}
\end{array}\right)(x, k, \mathfrak{s})(t)
$$

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## Needed moments

$$
\begin{gather*}
\Pi:=-\frac{m^{2}}{3} \rho_{0}, \quad \pi^{\mu \nu}:=\rho_{0}^{\mu \nu} \quad\left(T^{\mu \nu}\right)  \tag{12a}\\
\mathfrak{p}^{\mu}:=\tau_{0}^{\langle\mu\rangle}, \quad z^{\mu \nu}:=\tau_{1}^{(\langle\mu\rangle,\langle\nu\rangle)}, \quad \mathfrak{q}^{\lambda \mu \nu}:=\tau_{0}^{\langle\lambda\rangle, \mu \nu} \quad\left(J^{\lambda \mu \nu}\right)  \tag{12b}\\
\psi_{1}^{\mu \nu}, \quad \psi_{0}^{\mu \nu, \lambda}\left(\Theta^{\mu \nu}\right) \tag{12c}
\end{gather*}
$$

## Results I: Dissipative Spin Hydro

## Dissipative Hydro: Evolution equations

$$
\begin{align*}
\tau_{\Pi} \dot{\Pi}+\Pi & =-\zeta \theta+\text { h.o.t. }  \tag{13a}\\
\tau_{\pi} \dot{\pi}^{\langle\mu \nu\rangle}+\pi^{\mu \nu} & =2 \eta \sigma^{\mu \nu}+\text { h.o.t. }  \tag{13b}\\
\tau_{\mathfrak{p}} \dot{\mathfrak{p}}^{\langle\mu\rangle}+\mathfrak{p}^{\langle\mu\rangle} & =\mathfrak{e}^{(0)}\left(\tilde{\Omega}^{\mu \nu}-\tilde{\varpi}^{\mu \nu}\right) u_{\nu}+\text { h.o.t. }  \tag{13c}\\
\tau_{\mathfrak{z}} \dot{\mathfrak{z}}^{\langle\mu\rangle\langle\nu\rangle}+\mathfrak{z}^{\langle\mu\rangle\langle\nu\rangle} & =\text { h.o.t. }  \tag{13d}\\
\tau_{\mathfrak{q}} \dot{\mathfrak{q}}^{(\lambda\rangle\langle\mu \nu\rangle}+\mathfrak{q}^{\langle\lambda\rangle\langle\mu \nu\rangle} & =\mathfrak{d}^{(2)} \beta_{0} \sigma_{\alpha}{ }^{\langle\mu} \epsilon^{\nu\rangle \lambda \alpha \beta} u_{\beta}+\text { h.o.t. }  \tag{13e}\\
\tau_{\psi_{1}} \dot{\psi}_{1}^{\langle\mu \nu\rangle}+\psi_{1}^{\langle\mu \nu\rangle} & =\xi \beta_{0} \pi^{\mu \nu}+\text { h.o.t. }  \tag{13f}\\
\tau_{\psi_{0}} \dot{\psi}_{0}^{\langle\mu \nu\rangle, \lambda}+\psi_{0}^{\langle\mu \nu\rangle, \lambda} & =\text { h.o.t. } \tag{13g}
\end{align*}
$$

$$
\varpi^{\mu \nu}:=-\frac{1}{2} \partial^{[\mu}\left(\beta_{0} u^{\nu]}\right), \tilde{A}^{\mu \nu}:=\epsilon^{\mu \nu \alpha \beta} A_{\alpha \beta}
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- Evaluate polarization and alignment in the Navier-Stokes limit

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$$

## Results II: Alignment

- Moments of spin-rank 2:

$$
\begin{equation*}
\psi_{1}^{\langle\mu \nu\rangle} \simeq \xi \beta_{0} \pi^{\mu \nu}, \quad \psi_{0}^{\langle\mu \nu\rangle, \lambda} \simeq 0 \tag{14}
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$$



## Results II: Alignment

## Alignment: Explicit expression

$$
\begin{align*}
\rho_{00}(k)= & \frac{1}{3} \\
- & \frac{4}{15}\left[\int \mathrm{~d} \Sigma_{\lambda} k^{\lambda} f_{0 \mathbf{k}}\left(1-3 \mathcal{H}_{\mathbf{k} 0}^{(0,0)} \Pi / m^{2}+\mathcal{H}_{\mathbf{k} 0}^{(0,2)} \pi^{\mu \nu} k_{\mu} k_{\nu}\right)\right]^{-1} \\
& \times \int \mathrm{d} \Sigma_{\lambda} k^{\lambda} \mathcal{H}_{\mathbf{k} 1}^{(2,0)} \xi \beta_{0} f_{0 \mathbf{k}} \epsilon_{\mu}^{(0)} \epsilon_{\nu}^{(0)} K_{\alpha \beta}^{\mu \nu} \Xi_{\gamma \delta}^{\alpha \beta} \pi^{\gamma \delta} \tag{15}
\end{align*}
$$

$$
\begin{aligned}
& f_{0 \mathbf{k}}:=\exp \left(-\beta_{0} E_{\mathbf{k}}\right) \\
& \Xi_{\alpha \beta}^{\mu \nu}:=\frac{1}{2} \Xi_{\alpha}^{(\mu} \Xi_{\beta}^{\nu)}-\frac{1}{\Xi^{2}} \Xi^{\mu \gamma} \Xi_{\gamma}^{\nu} \Xi_{\alpha \delta} \Xi_{\beta}^{\delta} \\
& \Xi^{\mu \nu}:=\Delta^{\mu \nu}+k^{\langle\mu\rangle} k^{\langle\nu\rangle} / E_{\mathbf{k}}^{2}
\end{aligned}
$$

- Polarization is determined by the Pauli-Lubanski (pseudo)vector


## Results III: Polarization

- Polarization is determined by the Pauli-Lubanski (pseudo)vector


## Pauli-Lubanski pseudovector (spin 1/2)

$$
\begin{aligned}
S^{\mu}(k)= & \frac{1}{2 \mathcal{N}} \int \mathrm{~d} \Sigma_{\lambda} k^{\lambda} \mathrm{d} S(k) \mathfrak{s}^{\mu} f(x, k, \mathfrak{s}) \\
\simeq & \int \mathrm{d} \Sigma_{\lambda} k^{\lambda} \frac{f_{0}}{2 \mathcal{N}}\left\{-\frac{\hbar}{2 m} \tilde{\Omega}^{\mu \nu} k_{\nu}+\left(\delta_{\nu}^{\mu}-\frac{u^{\mu} k_{\langle\nu\rangle}}{E_{\mathbf{k}}}\right)\right. \\
& \left.\times\left[\mathfrak{e} \chi_{\mathfrak{p}}\left(\tilde{\Omega}^{\nu \rho}-\tilde{\varpi}^{\nu \rho}\right) u_{\rho}-\chi_{\mathfrak{q}} \mathfrak{d} \beta_{0} \sigma_{\rho}{ }^{\langle\alpha} \epsilon^{\beta\rangle \nu \sigma \rho} u_{\sigma} k_{\langle\alpha} k_{\beta\rangle}\right]\right\}(16 \mathrm{~b})
\end{aligned}
$$

$$
\mathcal{N}:=\int \mathrm{d} \Sigma_{\lambda} k^{\lambda} \mathrm{d} S(k) f(x, k, \mathfrak{s})
$$

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& \left.\times\left[\mathfrak{e} \chi_{\mathfrak{p}}\left(\tilde{\Omega}^{\nu \rho}-\tilde{\varpi}^{\nu \rho}\right) u_{\rho}-\chi_{\mathfrak{q}} \mathfrak{d} \beta_{0} \sigma_{\rho}^{\langle\alpha} \epsilon^{\beta\rangle \nu \sigma \rho} u_{\sigma} k_{\langle\alpha} k_{\beta\rangle}\right]\right\}(16 \mathrm{~b})
\end{aligned}
$$

- Contains novel contributions from fluid shear

■ Only sourced by nonlocal collisions

$$
\mathcal{N}:=\int \mathrm{d} \Sigma_{\lambda} k^{\lambda} \mathrm{d} S(k) f(x, k, \mathfrak{s})
$$

## Summary

- Developed quantum kinetic theory and dissipative spin hydrodynamics
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- Kinetic formulation can be derived rigorously from QFT via the Wigner-function formalism
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- Employed method of moments to extract hydrodynamic limit
$\rightarrow$ Introduce multiple sets of moments dependent on spin
$\rightarrow$ Follow standard procedure to obtain equations of motion
$\rightarrow$ Truncate such that the evolution of $S^{\lambda \mu \nu}$ can be described
- Developed quantum kinetic theory and dissipative spin hydrodynamics
- Kinetic formulation can be derived rigorously from QFT via the Wigner-function formalism
- Quantum effects result in nonlocal collisions
- Employed method of moments to extract hydrodynamic limit $\rightarrow$ Introduce multiple sets of moments dependent on spin $\rightarrow$ Follow standard procedure to obtain equations of motion $\rightarrow$ Truncate such that the evolution of $S^{\lambda \mu \nu}$ can be described
- Connected polarization and alignment to fluid quantities in the Navier-Stokes limit


## Future perspectives

- Evaluate expressions for polarization and alignment with hydrodynamic simulations
- Implement full spin hydrodynamics numerically
- Include electric and magnetic fields


## Appendix

## Relevant time scales: An estimation



- Simplest interaction: constant cross section
- Spin-related relaxation times shorter than standard dissipative time scales, but not much


## Moment equations: Spin-rank 0

- Moments follow relaxation-type equations


## Moment equation for $\ell=0$

$$
\begin{align*}
\dot{\rho}_{r}-\mathfrak{C}_{r-1}= & {\left[(1-r) I_{r 1}-I_{r 0}\right] \theta-I_{r 0} \dot{\alpha}_{0}+I_{r+1,0} \dot{0}_{0} } \\
& +(r-1) \rho_{r-2}^{\mu \nu} \sigma_{\mu \nu}+r \rho_{r-1}^{\mu} \dot{u}_{\mu}-\nabla_{\mu} \rho_{r-1}^{\mu} \\
& -\frac{1}{3}\left[(r+2) \rho_{r}-(r-1) m^{2} \rho_{r-2}\right] \theta \tag{17}
\end{align*}
$$

- Depend both on equilibrium and dissipative quantities
- Not a closed system
- Blue terms will become Navier-Stokes values

$$
\begin{aligned}
& \dot{A}:=u \cdot \partial A, \nabla^{\mu}:=\Delta^{\mu \nu} \partial_{\nu} \\
& \theta:=\nabla \cdot u, \sigma^{\mu \nu}:=\nabla^{\langle\mu} u^{\nu\rangle}, E_{k}:=k \cdot u \\
& I_{n q}:=[(2 q+1)!!]^{-1} \int \mathrm{~d} \Gamma E_{k}^{n-2 q}\left(-k^{\langle\alpha\rangle} k_{\alpha}\right)^{q} \\
& \quad \text { David Wagner } \\
& \text { Nonlocal spin transport }
\end{aligned}
$$

## Moment equations: Spin-rank 1

- Same procedure as for the moments of spin-rank 0


## Moment equation for $\ell=0$

$$
\begin{align*}
\dot{\tau}_{r}^{\langle\mu\rangle}-\mathfrak{C}_{r-1}^{\langle\mu\rangle}= & \frac{\hbar}{2 m}\left\{\left[I_{r+1,0}+r I_{r+1,1}\right] \theta+I_{r+1,0} \dot{\alpha}_{0}-I_{r+2,0} \dot{\beta}_{0}\right\} \omega_{0}^{\mu} \\
& -\frac{\hbar}{4 m} I_{r+1,1} \Delta_{\lambda}^{\mu} \nabla_{\nu} \tilde{\Omega}^{\lambda \nu}-\frac{\hbar}{4 m} I_{r+1,0} \epsilon^{\mu \nu \alpha \beta} u_{\nu} \dot{\Omega}_{\alpha \beta} \\
& -\frac{\hbar}{4 m} \tilde{\Omega}^{\langle\mu\rangle \nu}\left[I_{r+1,1} \nabla_{\nu} \alpha_{0}-I_{r+2,1}\left(\nabla_{\nu} \beta_{0}+\beta_{0} \dot{u}_{\nu}\right)\right] \\
& +r \dot{u}_{\nu} \tau_{r-1}^{\langle\mu\rangle, \nu}+(r-1) \sigma_{\alpha \beta} \tau_{r-2}^{\langle\mu\rangle, \alpha \beta}-\Delta_{\lambda}^{\mu} \nabla_{\nu} \tau_{r-1}^{\lambda, \nu} \\
& -\frac{1}{3}\left[(r+2) \tau_{r}^{\langle\mu\rangle}-(r-1) m^{2} \tau_{r-2}^{\langle\mu\rangle}\right] \theta \tag{18}
\end{align*}
$$

- Determine the (vector) polarization of particles

$$
\tilde{\Omega}^{\mu \nu}:=\epsilon^{\mu \nu \alpha \beta} \Omega_{\alpha \beta}, \Omega^{\mu \nu}=u^{[\mu} \kappa_{0}^{\nu]}+\epsilon^{\mu \nu \alpha \beta} u_{\alpha} \omega_{0, \beta}
$$

## Moment equations: Spin-rank 2

Moment equation for $\ell=0$

$$
\begin{align*}
\dot{\psi}_{r}^{\langle\mu \nu\rangle}-\mathfrak{C}_{r-1}^{\langle\mu \nu\rangle}= & -\frac{\theta}{3}\left[(r+2) \psi_{r}^{\langle\mu \nu\rangle}-(r-1) m^{2} \psi_{r-2}^{\langle\mu \nu\rangle}\right]+r \psi_{r-1}^{\langle\mu \nu\rangle, \alpha} \dot{u}_{\alpha} \\
& -\Delta_{\alpha \beta}^{\mu \nu} \nabla_{\gamma} \psi_{r-1}^{\alpha \beta, \gamma}+(r-1) \psi_{r-2}^{\langle\mu \nu\rangle, \alpha \beta} \sigma_{\alpha \beta} \tag{19}
\end{align*}
$$

- No dependence on equilibrium quantities appears because moments of spin-rank 2 do not appear in any conserved current
- Nonetheless, they determine the tensor polarization of spin-1 particles

$$
\Delta_{\alpha \beta}^{\mu \nu}:=\left(\Delta_{\alpha}^{(\mu} \Delta_{\beta}^{\nu)}\right) / 2-(1 / 3) \Delta^{\mu \nu} \Delta_{\alpha \beta}
$$

## Truncation

- Spin-1: Which moments are contained in the total tensor polarization?


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## Total tensor polarization

$$
\begin{equation*}
\bar{\Theta}^{\mu \nu}:=\int \mathrm{d} K N(k) \Theta^{\mu \nu}(k)=\frac{1}{2} \sqrt{\frac{3}{2}} \int \mathrm{~d} \Sigma_{\lambda}\left(u^{\lambda} \psi_{1}^{\mu \nu}+\psi_{0}^{\mu \nu, \lambda}\right) \tag{20}
\end{equation*}
$$

[^0]
## Truncation

- Spin-1: Which moments are contained in the total tensor polarization?


## Total tensor polarization

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\begin{equation*}
\bar{\Theta}^{\mu \nu}:=\int \mathrm{d} K N(k) \Theta^{\mu \nu}(k)=\frac{1}{2} \sqrt{\frac{3}{2}} \int \mathrm{~d} \Sigma_{\lambda}\left(u^{\lambda} \psi_{1}^{\mu \nu}+\psi_{0}^{\mu \nu, \lambda}\right) \tag{20}
\end{equation*}
$$

- Lowest-order approximation: Keep only these moments in the employed basis, i.e.,

$$
\begin{equation*}
\delta f(x, k) \hat{=} \delta f\left(\Pi, \pi^{\mu \nu}, \mathfrak{p}^{\mu}, \mathfrak{z}^{\mu \nu}, \mathfrak{q}^{\lambda \mu \nu}, \psi_{1}^{\mu \nu}, \psi_{0}^{\mu \nu, \lambda}, k\right) \tag{21}
\end{equation*}
$$

$$
\mathrm{d} K:=\mathrm{d}^{3} \mathbf{k} /\left[(2 \pi \hbar)^{3} 2 k^{0}\right]
$$


[^0]:    $\mathrm{d} K:=\mathrm{d}^{3} \mathbf{k} /\left[(2 \pi \hbar)^{3} 2 k^{0}\right]$

