

The 7th International Conference on Chirality, Vorticity and Magnetic Field in
Heavy Ion Collisions

Interplay between nuclear structure and the chiral magnetic effect in isobaric collisions



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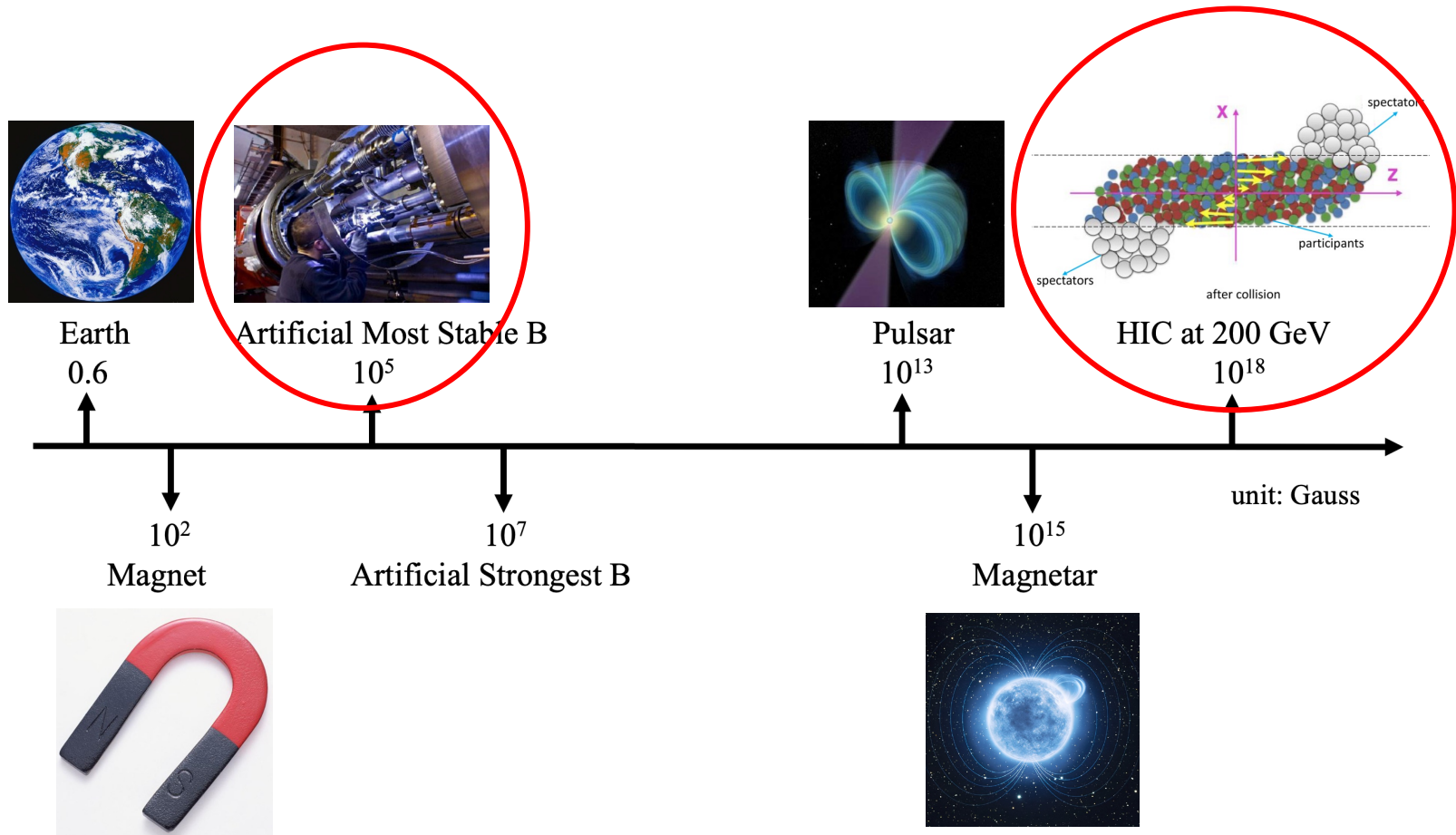
Jul¹⁵/₁₉ 2023 ◀◀◀◀

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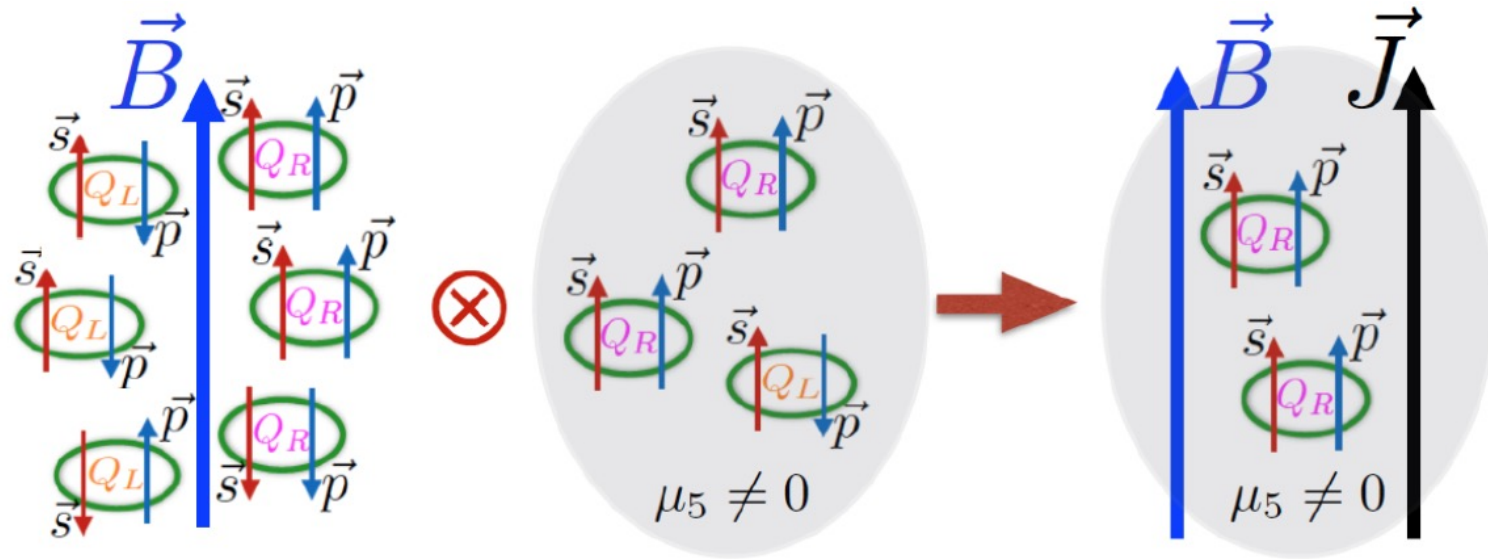
Outline

- 1) Motivation
- 2) Results and Discussions
 - ✓ N_{ch} & v_2
 - ✓ $\Delta\delta$ & $\Delta\gamma$
- 3) Summary

Strong Magnetic Field in HIC



Chiral Magnetic Effect (CME)



Prog. Part. Nucl. Phys. 88 (2016) 1-28.

$$\vec{J} = \sigma_5 \vec{B} = \left(\frac{(Qe)^2}{2\pi^2} \mu_5 \right) \vec{B}$$

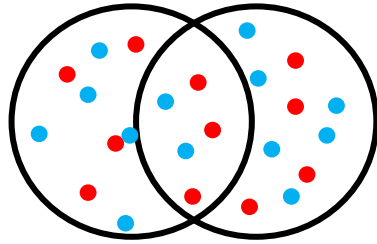
Observables: $\gamma \equiv \langle \cos(\phi_\alpha + \phi_\beta - 2\Psi_{\text{RP}}) \rangle$

$$= \langle \cos \Delta\phi_\alpha \cos \Delta\phi_\beta \rangle - \langle \sin \Delta\phi_\alpha \sin \Delta\phi_\beta \rangle \implies \Delta\gamma \equiv \gamma^{OS} - \gamma^{SS}$$

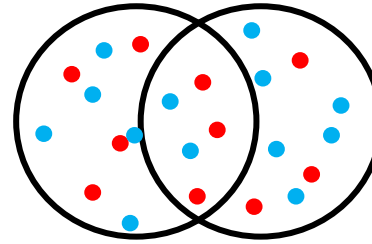
$$= (\langle v_{1,\alpha} v_{1,\beta} \rangle + B_{\text{in}}) - (\langle a_{1,\alpha} a_{1,\beta} \rangle + B_{\text{out}})$$

$$\delta \equiv \langle \cos(\phi_\alpha - \phi_\beta) \rangle \implies \Delta\delta = \delta^{OS} - \delta^{SS}$$

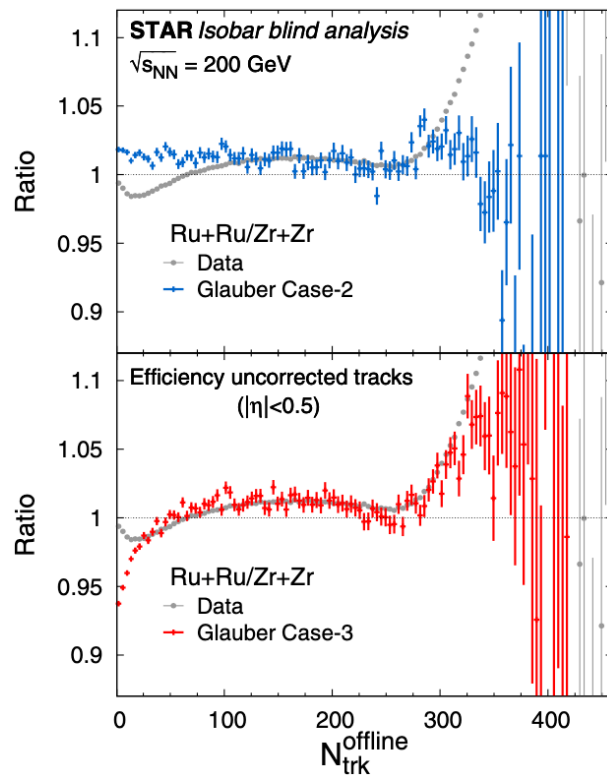
Isobaric Collisions



$^{96}_{44}\text{Ru} + ^{96}_{44}\text{Ru}$



$^{96}_{40}\text{Zr} + ^{96}_{40}\text{Zr}$



- Ratio $\neq 1 \implies$ Different background.
- Different initial nuclear structures, deformation or neutron skin?

Nuclear structure introduced in AMPT

A+B



HIJING

*energy in
excited strings and minijet partons*

nucleon
spectators

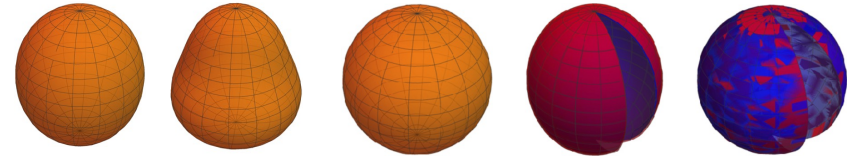
fragment into partons

ZPC (Zhang's Parton Cascade)

till parton freezeout

Quark Coalescence

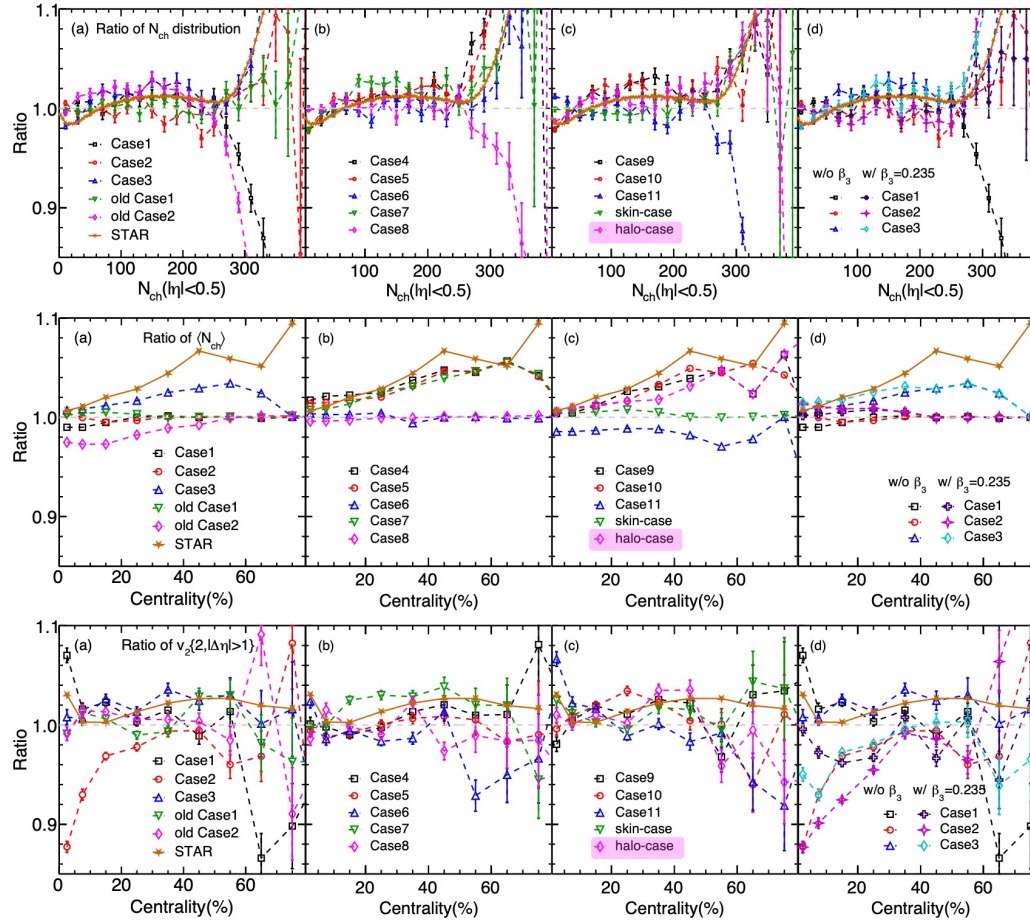
ART (A Relativistic Transport model for hadrons)



$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + \exp\left\{\frac{1}{a} (r - R[\theta, \phi])\right\}}$$

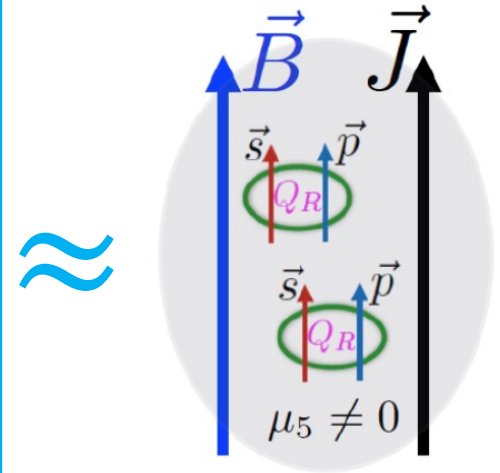
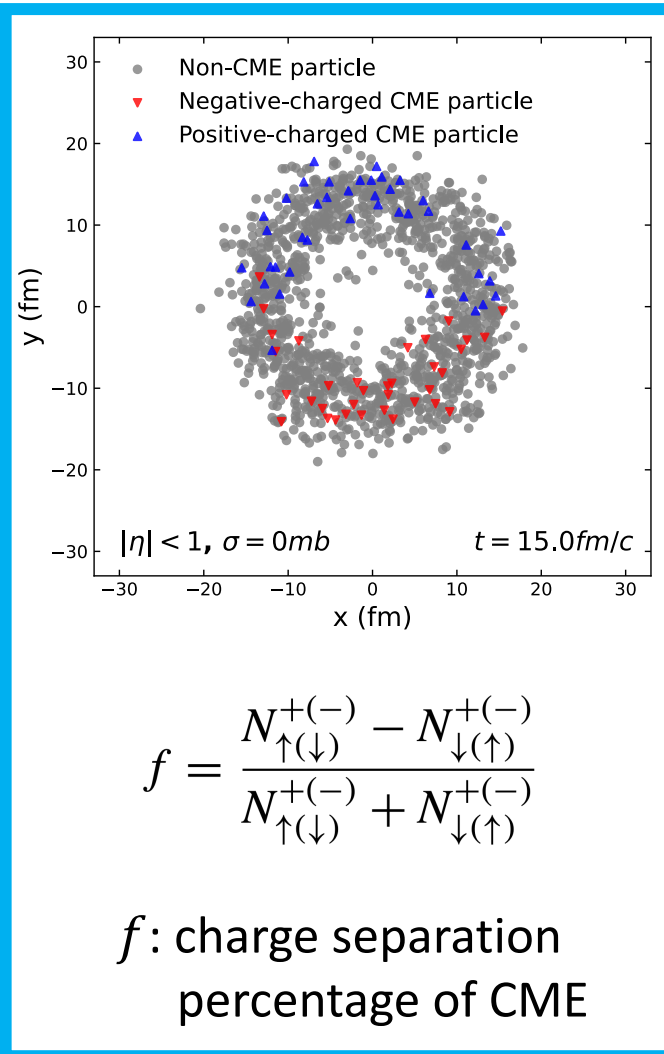
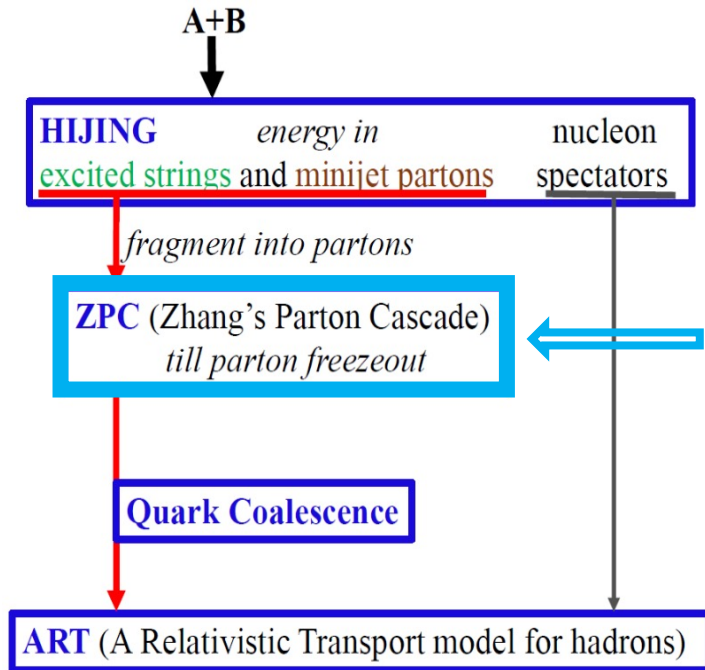
$$R[\theta, \phi] = R_0 \left(1 + \beta_2 [\cos\gamma Y_{2,0} + \sin\gamma Y_{2,2}] + \beta_3 \sum_{m=-3}^3 \alpha_{3,m} Y_{3,m} + \beta_4 \sum_{m=-4}^4 \alpha_{4,m} Y_{4,m} \right)$$

Results for 18 Cases VS STAR



	old Case 1	old Case 2	Case 1	Case 2	Case 3	Case 4
$\langle \chi^2 \rangle$	0.204	0.682	0.255	0.400	0.097	0.053
	Case 5	Case 6	Case 7	Case 8	Case 9	Case 10
$\langle \chi^2 \rangle$	0.049	0.224	0.051	0.227	0.057	0.048
	Case 11	skin-type	halo - type	Case 1 w/ β_3	Case 2 w/ β_3	Case 3 w/ β_3
$\langle \chi^2 \rangle$	0.430	0.177	0.047	0.247	0.506	0.166

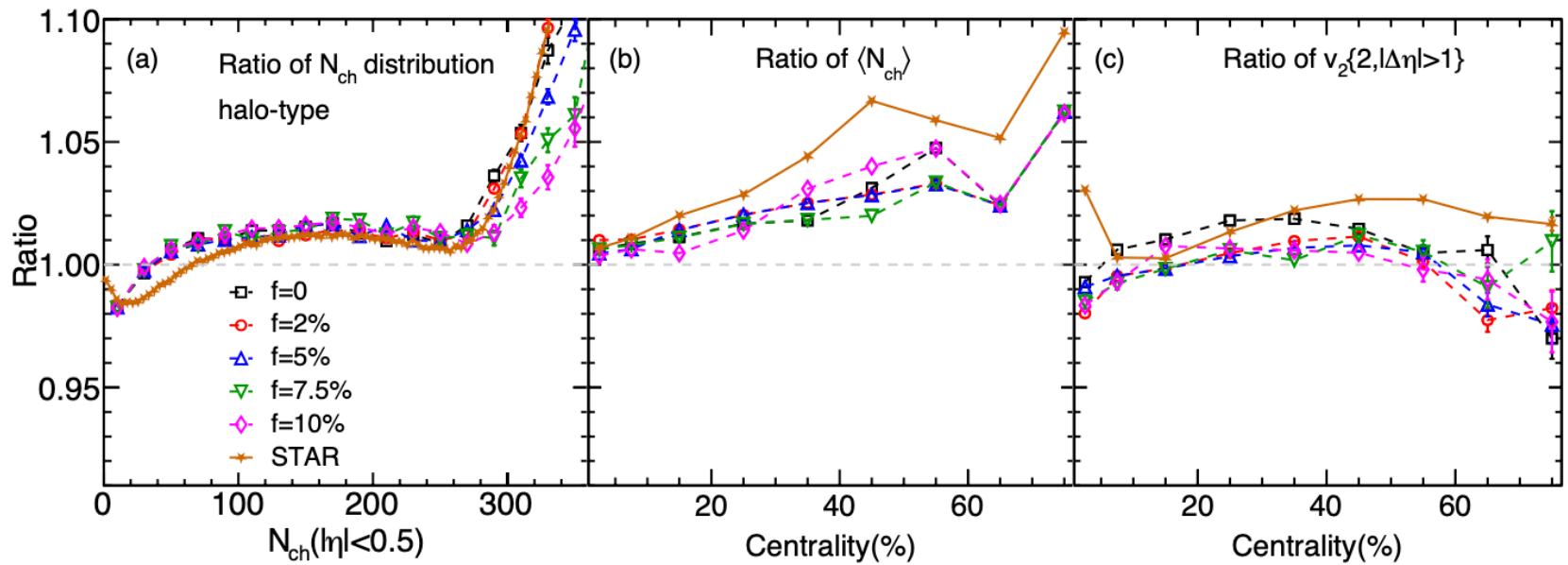
Charge Separation in AMPT



$$\vec{J} = \sigma_5 \vec{B} = \left(\frac{(Qe)^2}{2\pi^2} \mu_5 \right) \vec{B}$$

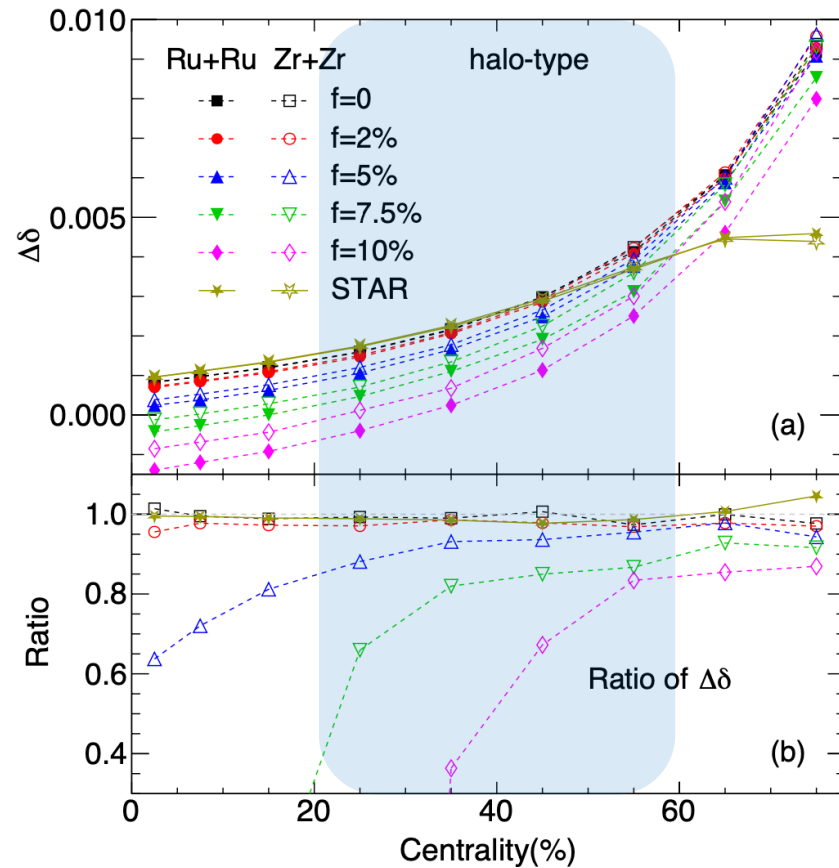
$f = 0, 2\%, 5\%, 7.5\%, 10\%$ in AMPT.

Halo-Type Results VS STAR



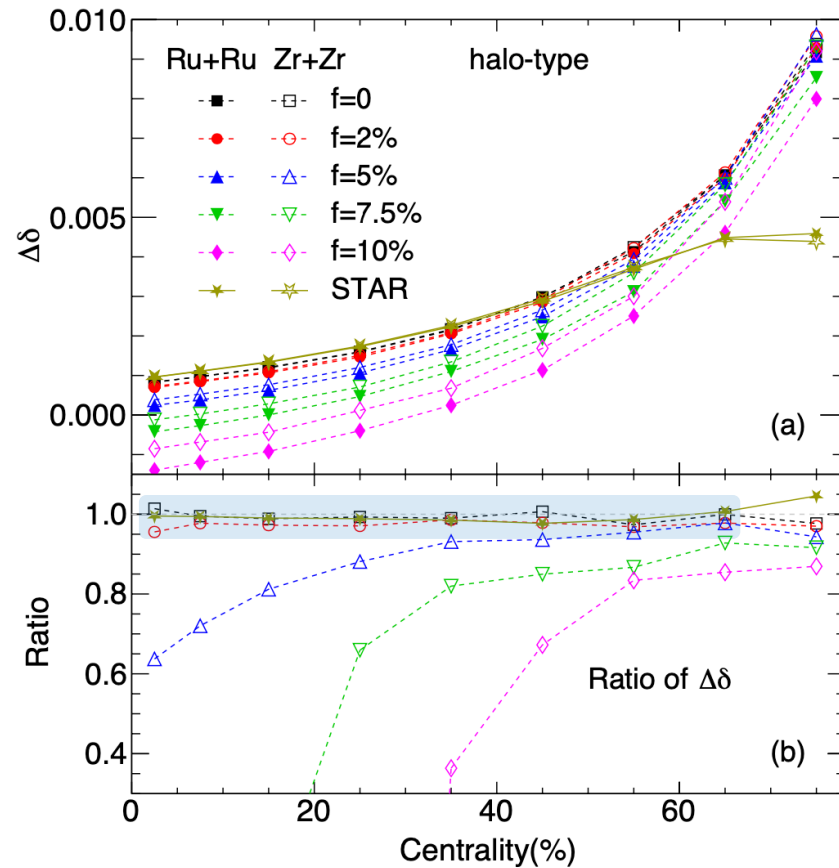
- At central and mid-central collisions, the N_{ch} & v_2 ratios are close to data.
- The ratios have no obvious dependences with respect to f .
- Overall, the results at small f are closer to data.

$\Delta\delta$ with Different f



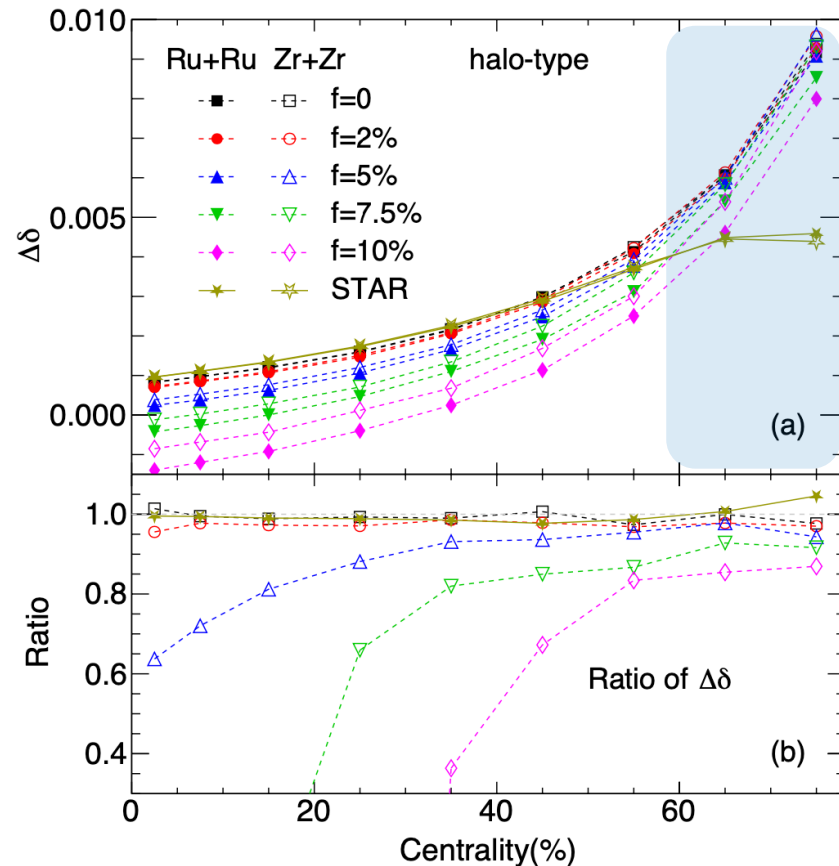
- $\delta \equiv \langle \cos(\phi_\alpha - \phi_\beta) \rangle$, $\Delta\delta = \delta^{OS} - \delta^{SS}$
- At mid-central collisions, $\Delta\delta \propto f$.

$\Delta\delta$ with Different f



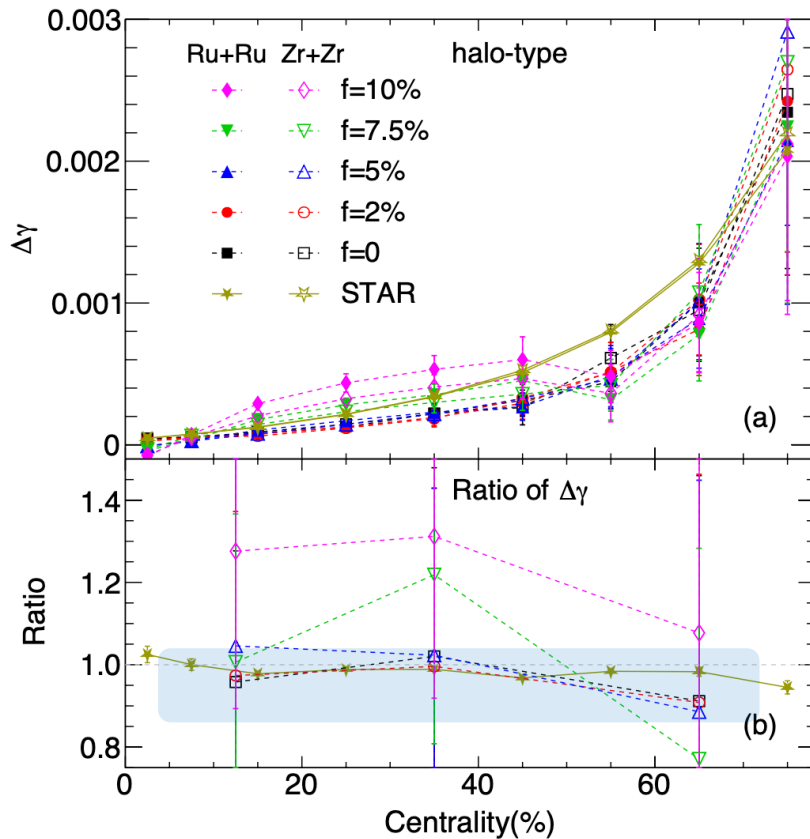
- $\delta \equiv \langle \cos(\phi_\alpha - \phi_\beta) \rangle$, $\Delta\delta = \delta^{OS} - \delta^{SS}$
- At mid-central collisions, $\Delta\delta \propto f$.
- At central & mid-central collisions, $\Delta\delta$ & its ratios are closer to data for $f < 5\%$.

$\Delta\delta$ with Different f



- $\delta \equiv \langle \cos(\phi_\alpha - \phi_\beta) \rangle$, $\Delta\delta = \delta^{OS} - \delta^{SS}$
- At mid-central collisions, $\Delta\delta \propto f$.
- At central & mid-central collisions, $\Delta\delta$ & its ratios are closer to data for $f < 5\%$.
- At peripheral collisions, model results are above the data.

$\Delta\gamma$ w/ Different f

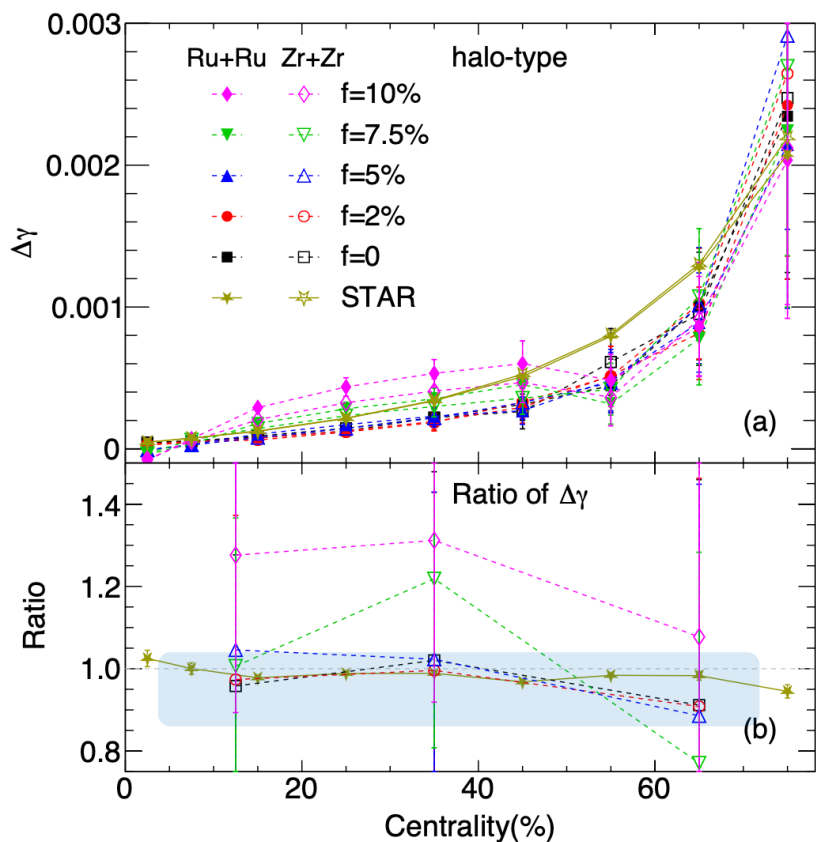


➤ $\gamma = \langle \cos(\phi_\alpha + \phi_\beta - 2\Psi_{\text{RP}}) \rangle,$

$\Delta\gamma = \gamma^{OS} - \gamma^{SS}$

➤ At central & mid-central collisions, $\Delta\gamma$ & its ratios are much closer to data for $f \leq 5\%$.

$\Delta\gamma$ w/ Different f



➤ $\gamma = \langle \cos(\phi_\alpha + \phi_\beta - 2\Psi_{\text{RP}}) \rangle,$

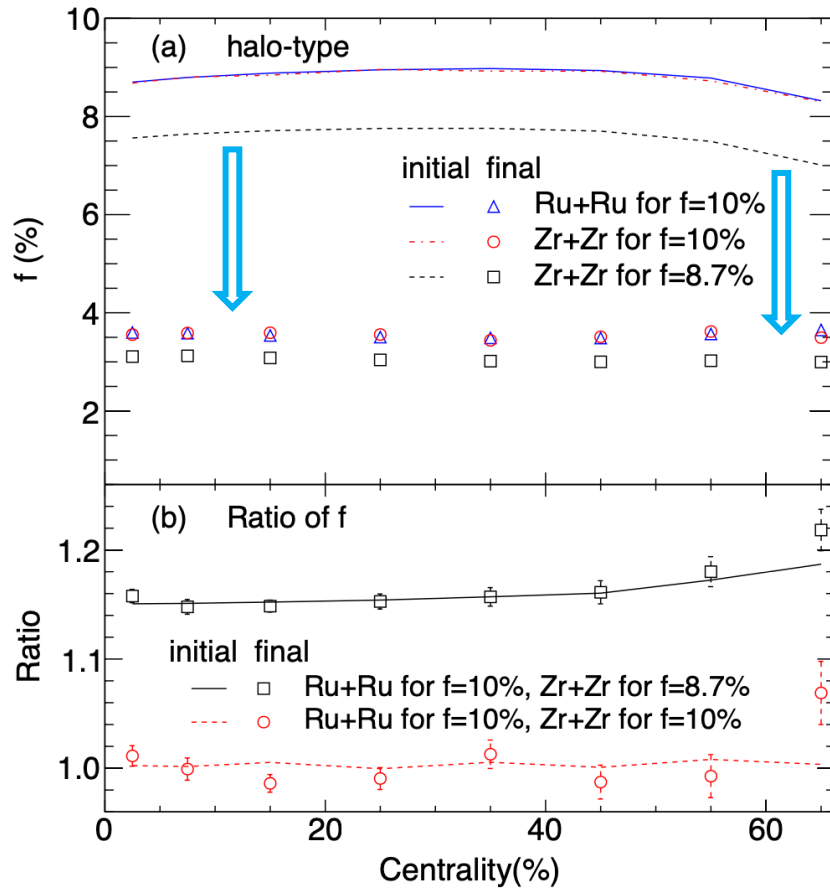
$\Delta\gamma = \gamma^{OS} - \gamma^{SS}$

➤ At central & mid-central collisions, $\Delta\gamma$ & its ratios are much closer to data for $f \leq 5\%$.

➤ The results at $f = 2\%$ & $f = 5\%$ are similar to $f = 0$.

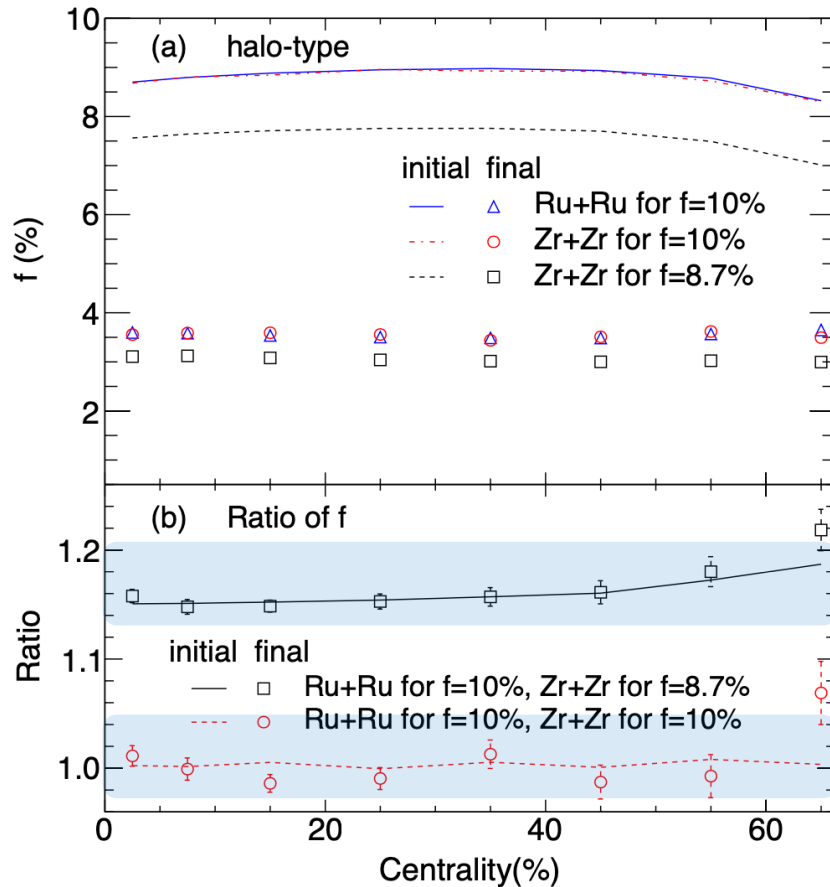
➤ In isobaric collisions, the CME signal is weak and need to more sensitive observables.

Final State Interaction



➤ Final state interactions weaken the CME signal.

Final State Interaction



- Final state interactions weaken the CME signal.
- The ratios at initial state are consistent with final state.
- The differences of CME can keep to the final state in isobaric collisions.

Summary

- The neutron-skin has big effect in isobaric collisions.
- $\Delta\delta$ & $\Delta\gamma$ results can be reproduced by AMPT w/o or w/ small CME strength.
- In isobaric collisions, initial CME signal is absent or small.
- Final state interactions significantly weaken the initial CME.
- More sensitive observables are required for searching the possible small CME signal in isobaric collisions.

Summary

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- $\Delta\delta$ & $\Delta\gamma$ results can be reproduced by AMPT w/o or w/ small CME strength.
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Thank you for your attention!