

Aug, 10, 2022 @ CHEP

Gravitational waves from axion clumps and axion searches

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Based on : arXiv: [2003.10527](#), [2206.13543](#)

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Research Scientific Calenda

Search Challenges and Opportunities of High Frequency
Gravitational Wave Detection | (smr 3493)

Overview
Programme
Speakers
Practical info

Starts 14 Oct 2019
Ends 16 Oct 2019
Central European Time

ICTP
Kastler Lecture Hall (AGH)
Via Grignano, 9
I - 34151 Trieste (Italy)

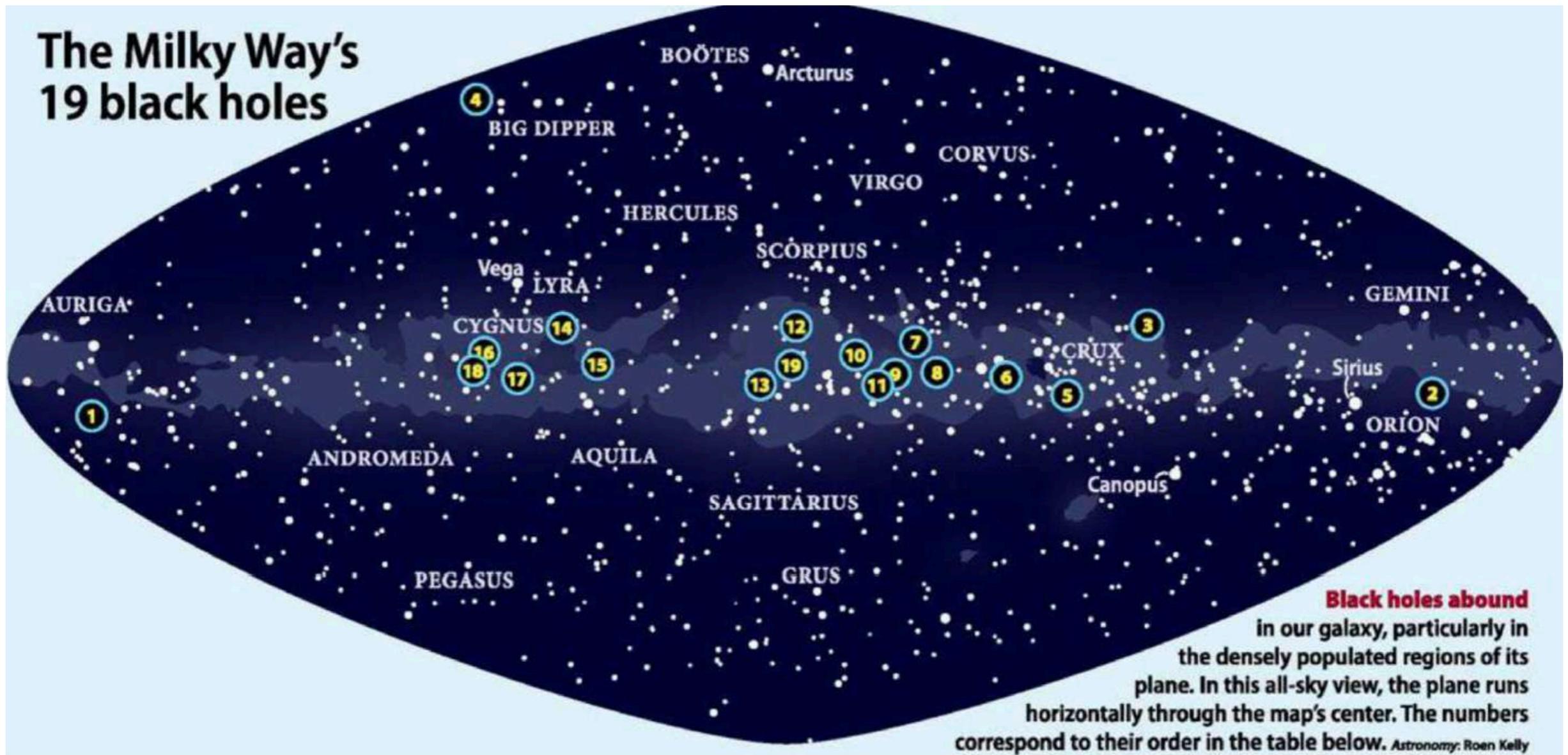


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Outline

- Axion resonance effects
- Induced Gravitational waves in difference frequencies
- Superradiance around the primordial blackhole.
- Axion detection (RE-LEAP)

The Milky Way's 19 black holes



Active Galactic Nuclei (AGN) and Black hole binary(BHB)

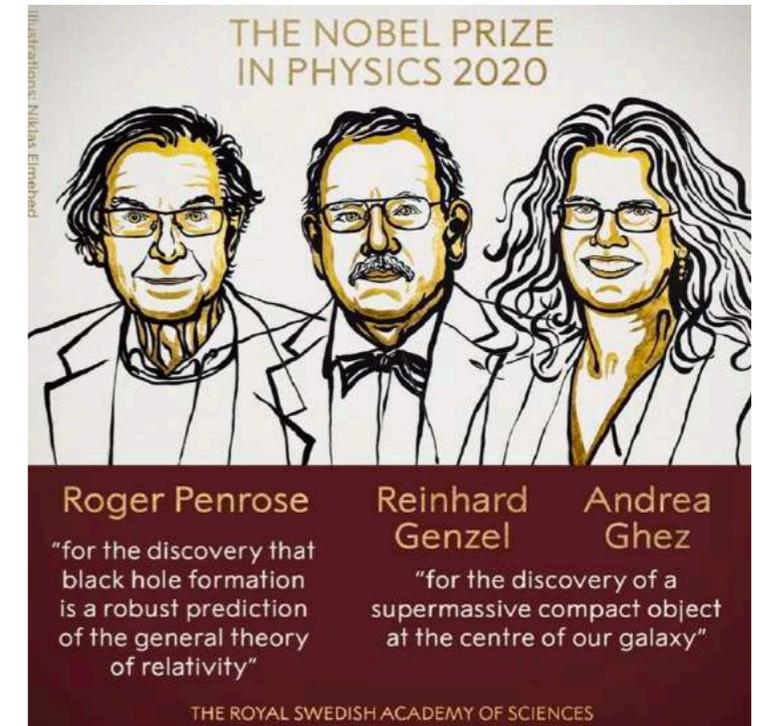
- Sgr A (Supermassive Compact Object) in the galaxy center, Nobel prize in physics(2020)
- Active Galactic Nuclei



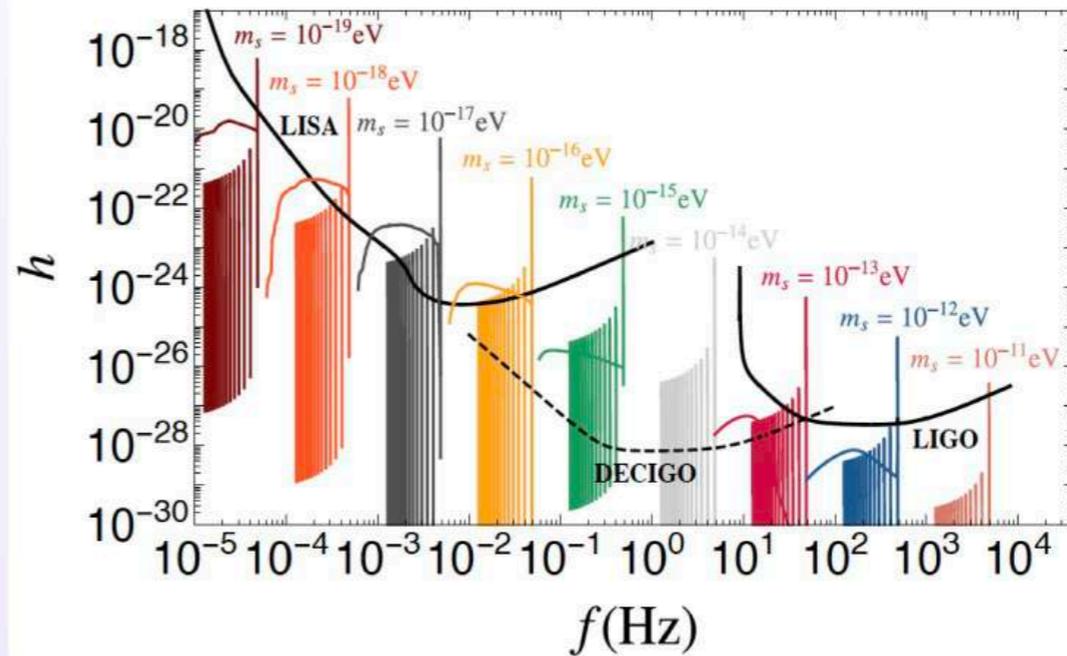
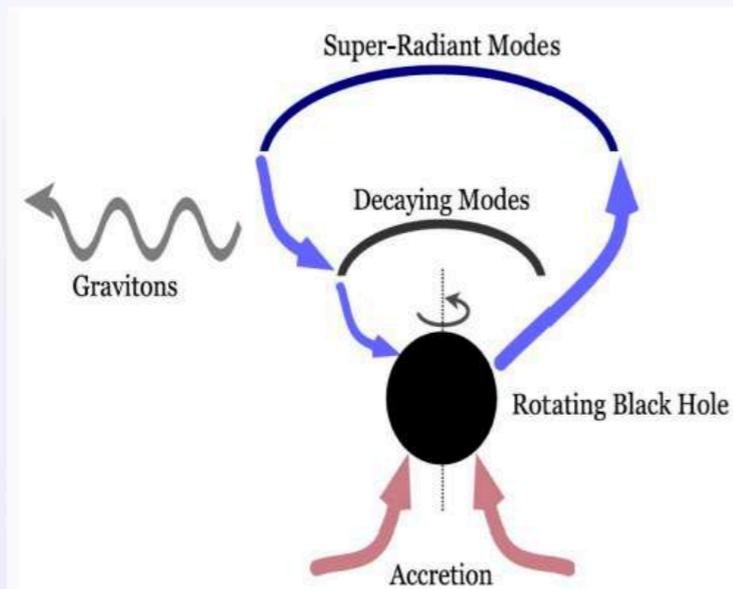
Credit: NASA/JPL-Caltech



Illustration of MAXI J1820+070. Credit: John Paice.



Black hole energy levels and gravitational waves



- Energy transition $\vartheta^+ \rightarrow \vartheta^- + h$, Strain $h \sim 10^{-19} - 10^{-27}$
- [cf. Brito-Cardoso-Pani, Superradiance 2020]

Comparing different channels:

- Summary of different channels with different Chern-Simons types of couplings:

$$\mathbf{a}_+ \rightarrow \mathbf{a}_- + \mathbf{g}$$

$$\mathbf{a} \rightarrow \mathbf{g}\mathbf{g}$$

$$\mathbf{a}\mathbf{a} \rightarrow \mathbf{g}$$

$$\mathbf{a} \rightarrow \gamma\gamma \text{ or } \mathbf{a}\mathbf{a} \rightarrow \gamma\gamma$$

- A conference white paper
- (Living Reviews in Relativity (2021) 1, 4 • e-Print: [2011.12414](https://arxiv.org/abs/2011.12414), Impact Factor: 40.43 (2020)):

Ultra-High-Frequency GWs: A Theory and Technology Roadmap

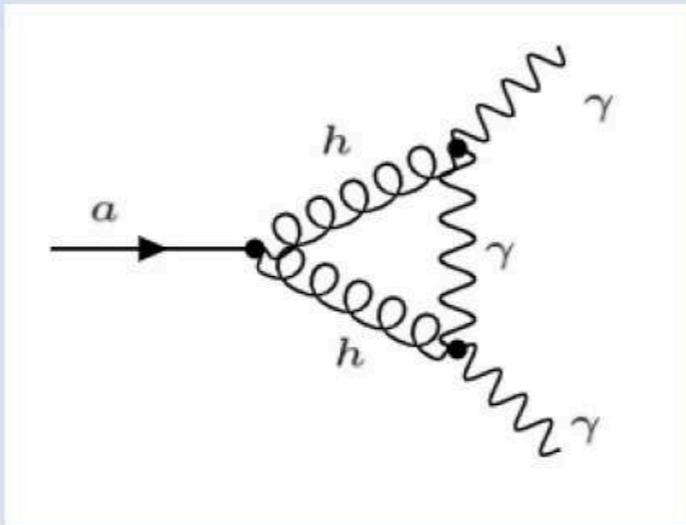
Oct 12 – 15, 2021
CERN
Europe/Zurich timezone

Challenges and Opportunities of Gravitational Wave Searches at MHz to GHz Frequencies

N. Aggarwal^{a,*}, O.D. Aguiar^b, A. Bauswein^c, G. Cella^d, S. Clesse^e, A.M. Cruise^f, V. Domcke^{g,h,i,*},
D.G. Figueroa^j, A. Geraci^k, M. Goryachev^l, H. Grote^m, M. Hindmarsh^{n,o}, F. Muia^{p,i,*}, N. Mukund^q,
D. Ottaway^{r,s}, M. Peloso^{t,u}, F. Quevedo^{p,*}, A. Ricciardone^{t,u}, J. Steinlechner^{v,w,x,*}, S. Steinlechner^{v,w,*},
S. Sun^{y,z}, M.E. Tobar^l, F. Torrenti^α, C. Unal^β, G. White^γ

Branch Ratio and Triangle Feynman Diagram

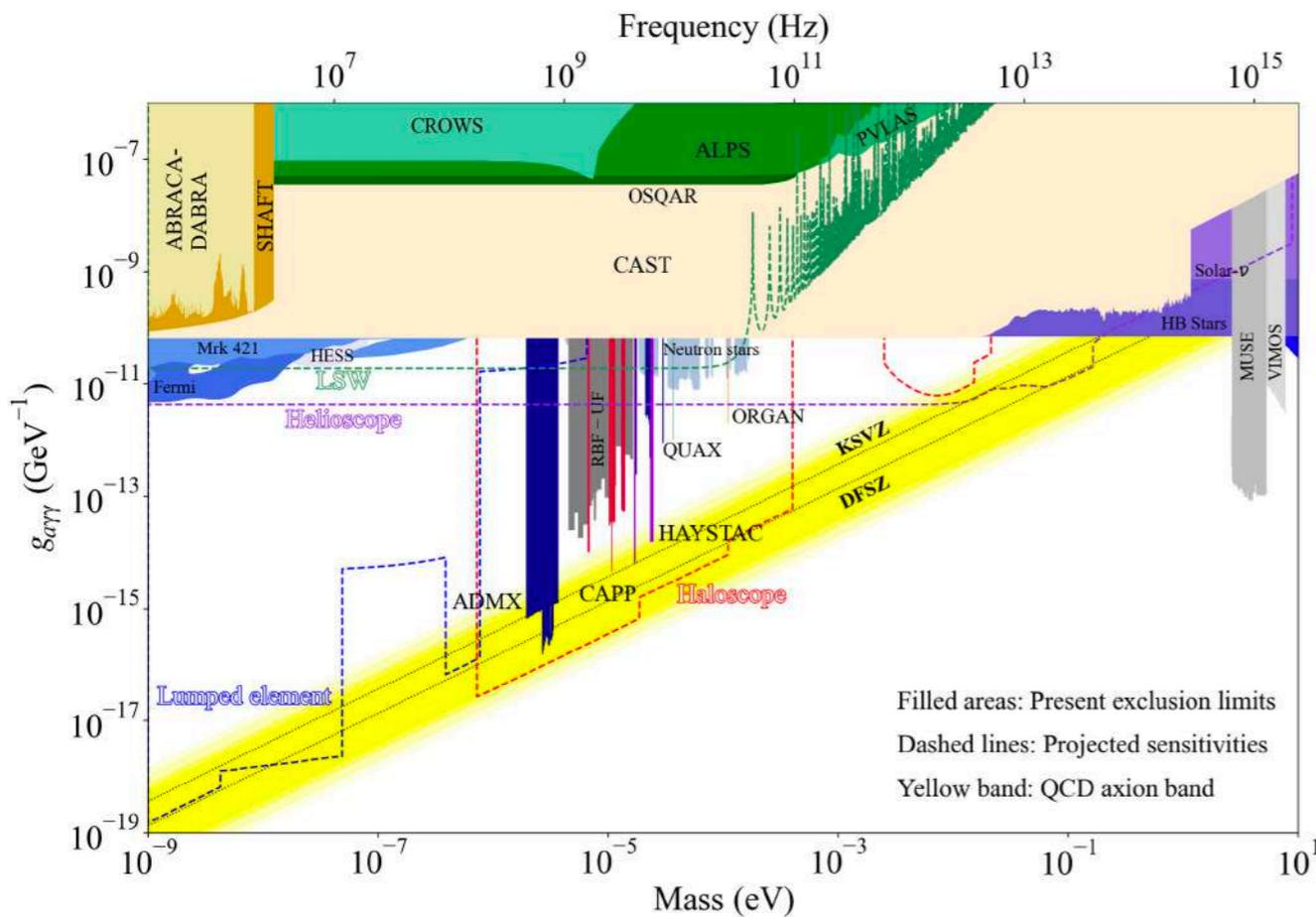
- The triangle Feynman diagram, where the axion-photon coupling is generated from CS gravity coupling.



- $\mathcal{L}_{\partial F \tilde{F}} = -\frac{\alpha_\gamma}{4} \partial F_{\mu\nu} \tilde{F}^{\mu\nu}, \quad \mathcal{L}_{\partial R \tilde{R}} = \frac{\alpha_g}{4} \partial R^\beta_{\alpha\gamma\delta} \tilde{R}^{\alpha\gamma\delta}_\beta.$
- This diagram is quadratically divergent, with 4 extra powers of momentum from 3 vertices in the loop and two powers of M_{pl} from $h_{\mu\nu} T^{\mu\nu}$ coupling
- It is evaluated as $\alpha_\gamma \sim \alpha_g (\Lambda/M_{pl})^4$, where Λ is the cut-off for Chern-Simons theory.

Resonant Electric Probe to Axionic Dark Matter 2206.13543

- Axion detection communities:
 - Cavities: ADMX [26], HAYSTAC [27], CAPP [28], QUAX- $\alpha\gamma$ [29], TASEH [30], DANCE [31], ORGAN [32], CAST-RADES [33],
 - non-cavity high-frequency designs: MADMAX [34], BREAD [35], etc
 - lower axion masses : ADMX-SLIC [37], DM-Radio [38], BASE [39], ABRA CADARA [40, 41]

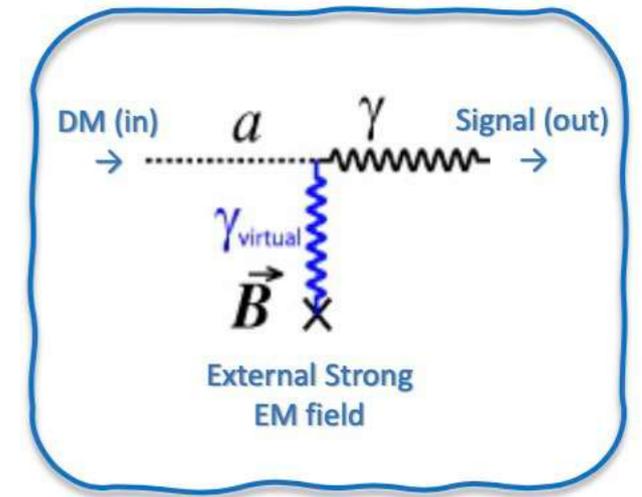


$$\mathcal{L}_{a\gamma\gamma} = -g_{a\gamma} a \vec{E} \cdot \vec{B},$$

Axion modified Electromagnetism

$$\vec{\nabla} \cdot \vec{E} = \rho_e - g_{a\gamma} \vec{B} \cdot \vec{\nabla} a,$$

$$\vec{\nabla} \times \vec{B} - \frac{\partial \vec{E}}{\partial t} = \vec{j}_e - g_{a\gamma} \vec{E} \times \vec{\nabla} a + g_{a\gamma} \vec{B} \frac{\partial a}{\partial t},$$



- Axion dark matter: $a(x, t) \approx a_0 \cos(m_a \vec{v} \cdot \vec{x} - \omega t + \phi_0)$,

$$\rho_a = -g_{a\gamma} B_0 (\vec{v} \cdot \hat{z}) \sqrt{2\rho_{\text{DM}}} \cos(\omega t),$$

$$\vec{j}_a = \hat{z} g_{a\gamma} B_0 \left(1 + \frac{1}{2} v^2\right) \sqrt{2\rho_{\text{DM}}} \cos(\omega t),$$

- Axion effective charge and current satisfies:

$$\partial_t \rho_a + \nabla \cdot \vec{j}_a = 0,$$

W/O cavity? – ‘aQED’ induction effects

➤ axion-modified Maxwell equations:

Effective charge: (suppressed as $v_a \ll 1$)
(j^0 of the locally conserved 4-current $\partial_\mu j_a^\mu = 0$)

$$\vec{\nabla} \cdot \vec{E} = \rho_e + g \vec{B} \cdot \nabla a$$

$$\vec{\nabla} \times \vec{B} - \frac{\partial \vec{E}}{\partial t} = g \vec{E} \times \vec{\nabla} a - g \vec{B} \frac{\partial a}{\partial t} + \vec{j}_e$$

$$\vec{\nabla} \cdot \vec{B} = 0$$

$$\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t},$$

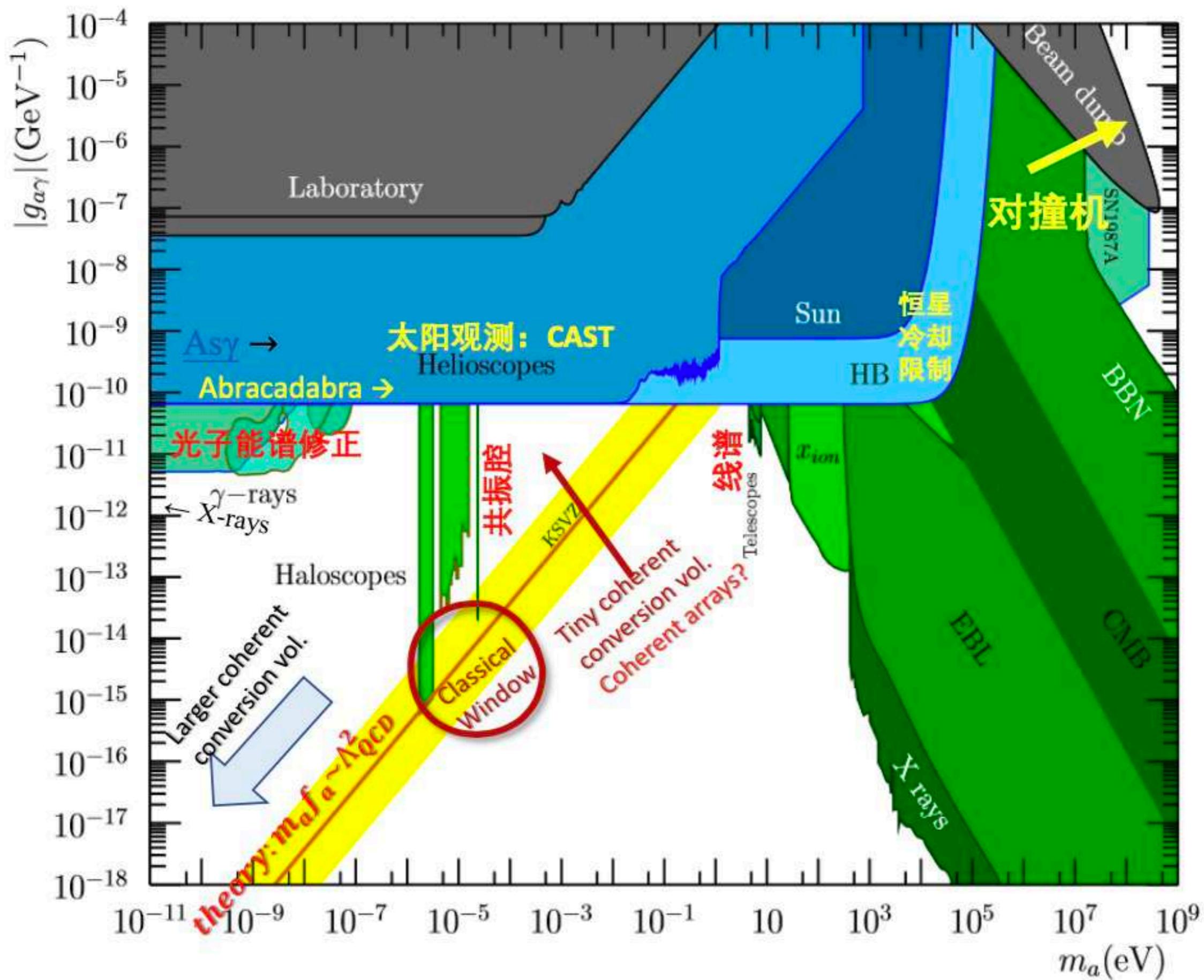
Axio-magnetic current:
LC, Abracadabra, ADMX-SLIC, DM-Radio, etc

Axio-electric current $\vec{j}_a = g \vec{E} \times \vec{\nabla} a$

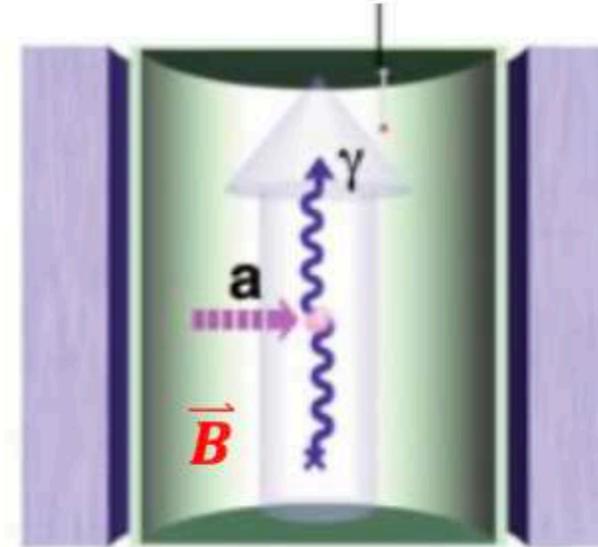
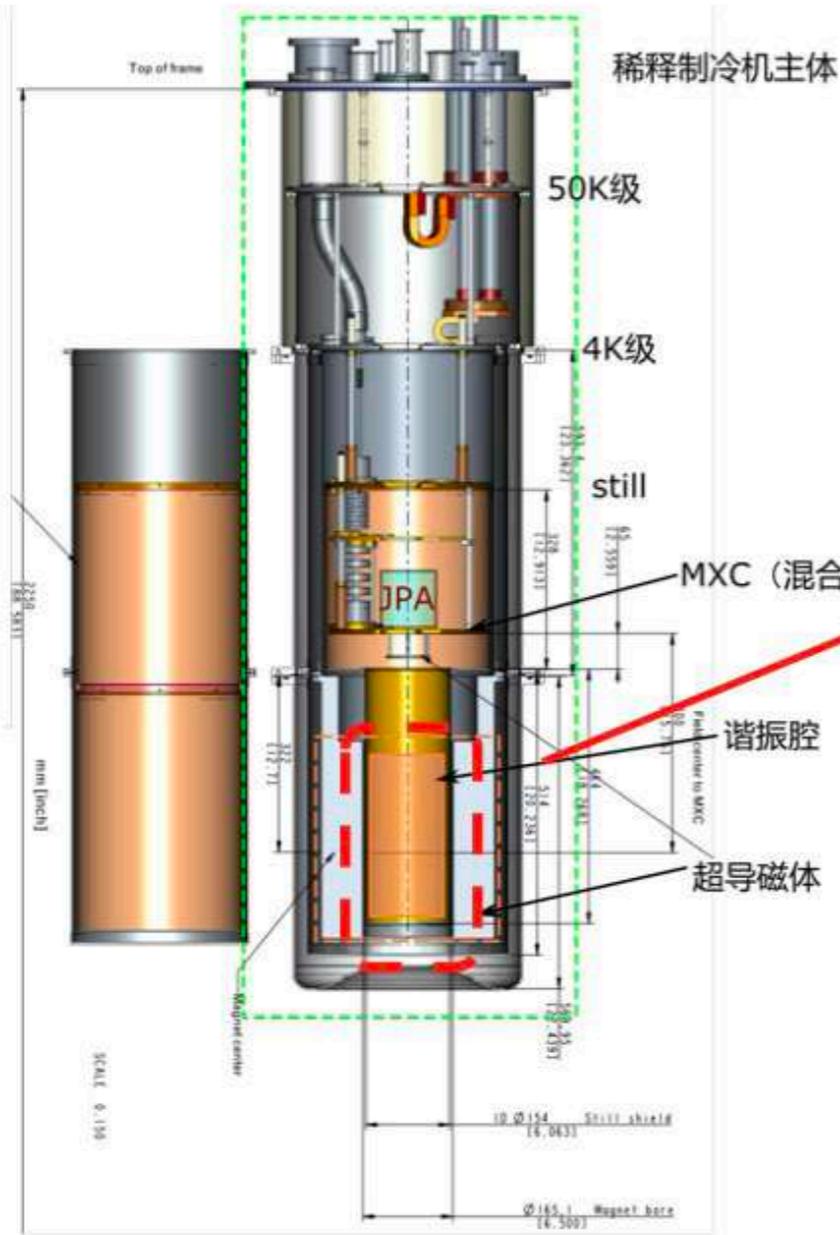
Axion's effective sources:

effective (moving) charge &
effective displacement currents

DM axion flow induces a magnetic signal
inside E field: see [2012.13946](#) (broad-band)
& [2204.14033](#) (narrow-band)



Cryogenic resonant EM cavity



Cavity frequency tuned to expected axion decay signal frequency

QCD axion DM: emergence of a microwave signal

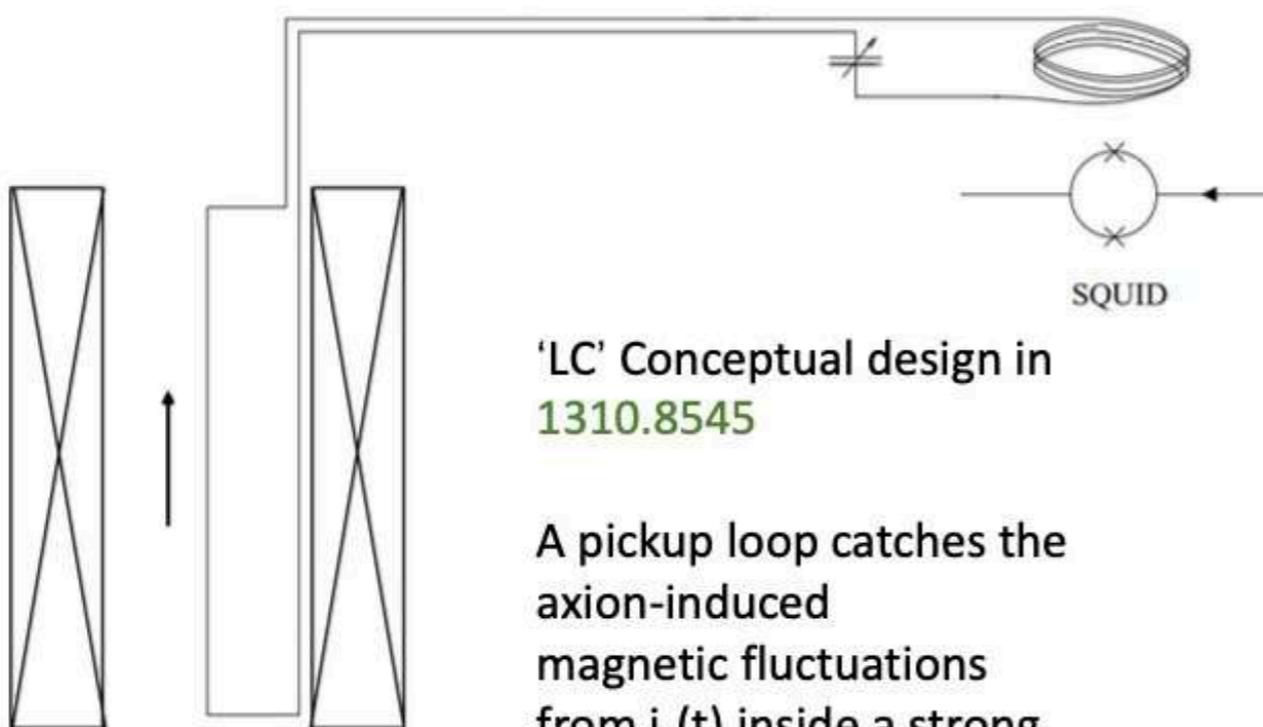
$$P_{\text{axion}} = 2.2 \cdot 10^{-23} \text{ W} \left(\frac{V}{136 \text{ L}} \right) \left(\frac{B}{7.6 \text{ T}} \right)^2 \left(\frac{C}{0.4} \right) \cdot \left(\frac{g_\gamma}{0.36} \right)^2 \left(\frac{\rho_a}{0.45 \text{ GeV cm}^{-3}} \right) \left(\frac{f}{740 \text{ MHz}} \right) \left(\frac{Q}{30000} \right)$$

➤ single photon level: $O(10)$ photons s^{-1}

Magnetic signal from B field

➤ 'LC'-type designs: ADMX-SLIC, Abracadabra, DM-Radio, etc.

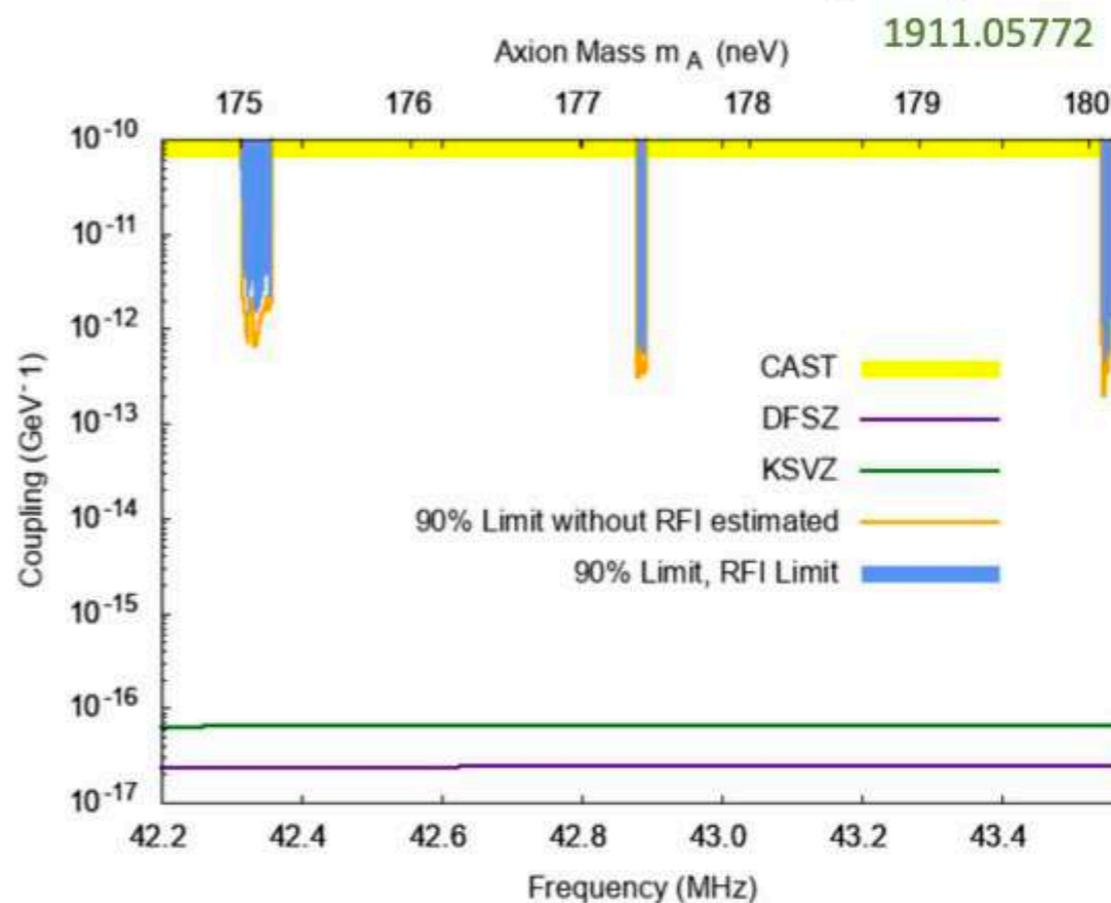
Enhanced by LC resonance
and measured with a
quantum magnetometer



'LC' Conceptual design in
1310.8545

A pickup loop catches the
axion-induced
magnetic fluctuations
from $j_a(t)$ inside a strong
lab B field

Realization in ADMX-SLIC (2019)



Long solenoid solutions

$$E_{a,r} = -\cos \omega t \hat{r} \cdot \begin{cases} \frac{1}{2} g_{a\gamma} B_0 v_z \sqrt{2\rho_{DM} r}, & r \leq R, \\ \frac{1}{2} g_{a\gamma} B_0 v_z \sqrt{2\rho_{DM}} \frac{R^2}{r}, & r \geq R, \end{cases}$$

$$E_{a,z} = \frac{\dot{j}_a}{\omega} e^{i\omega t} \cdot \begin{cases} \alpha(\omega) J_0(\omega r) - 1, & r \leq R, \\ \beta(\omega) H_0^+(\omega r), & r \geq R, \end{cases} \quad \begin{aligned} \alpha(\omega) &= \frac{i\pi\omega R}{2} H_1^+(\omega R), \\ \beta(\omega) &= \frac{i\pi\omega R}{2} J_1(\omega R), \end{aligned}$$

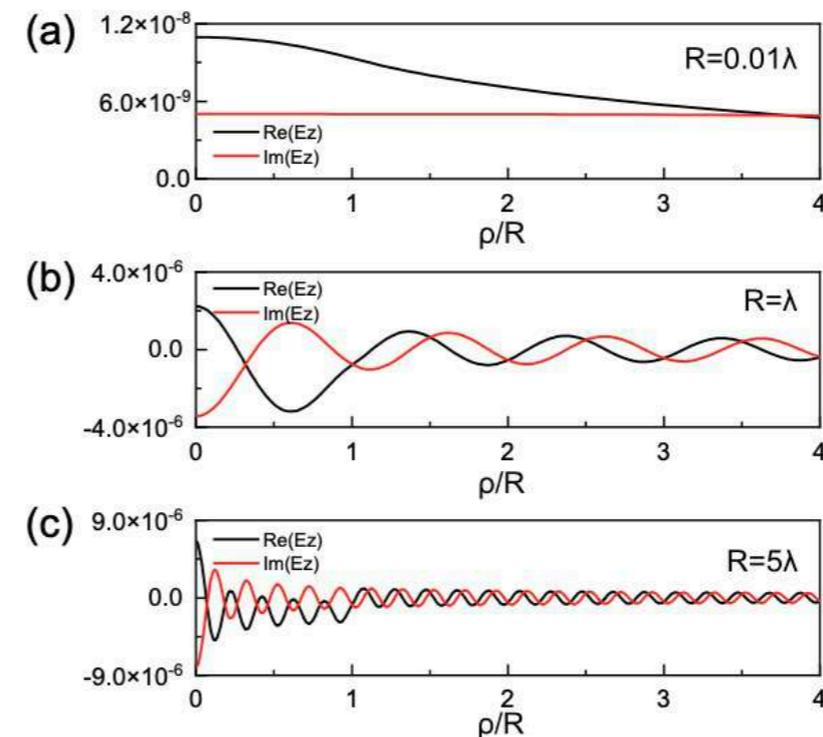
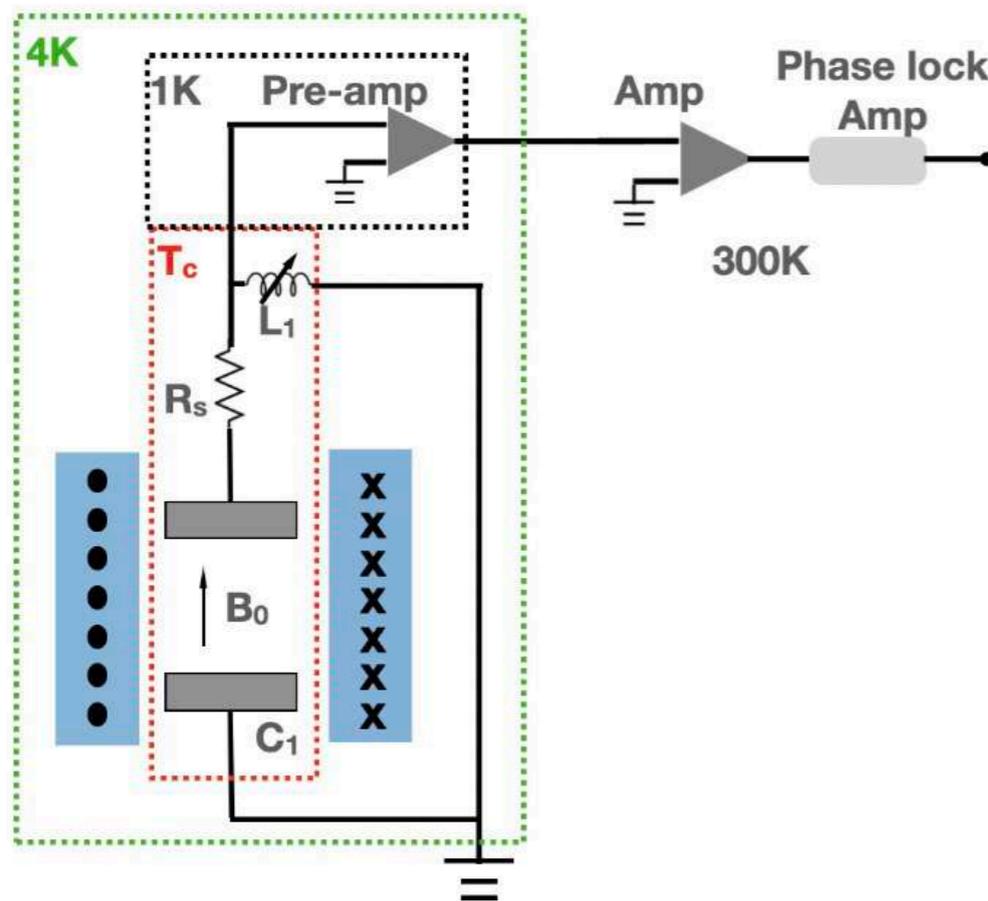
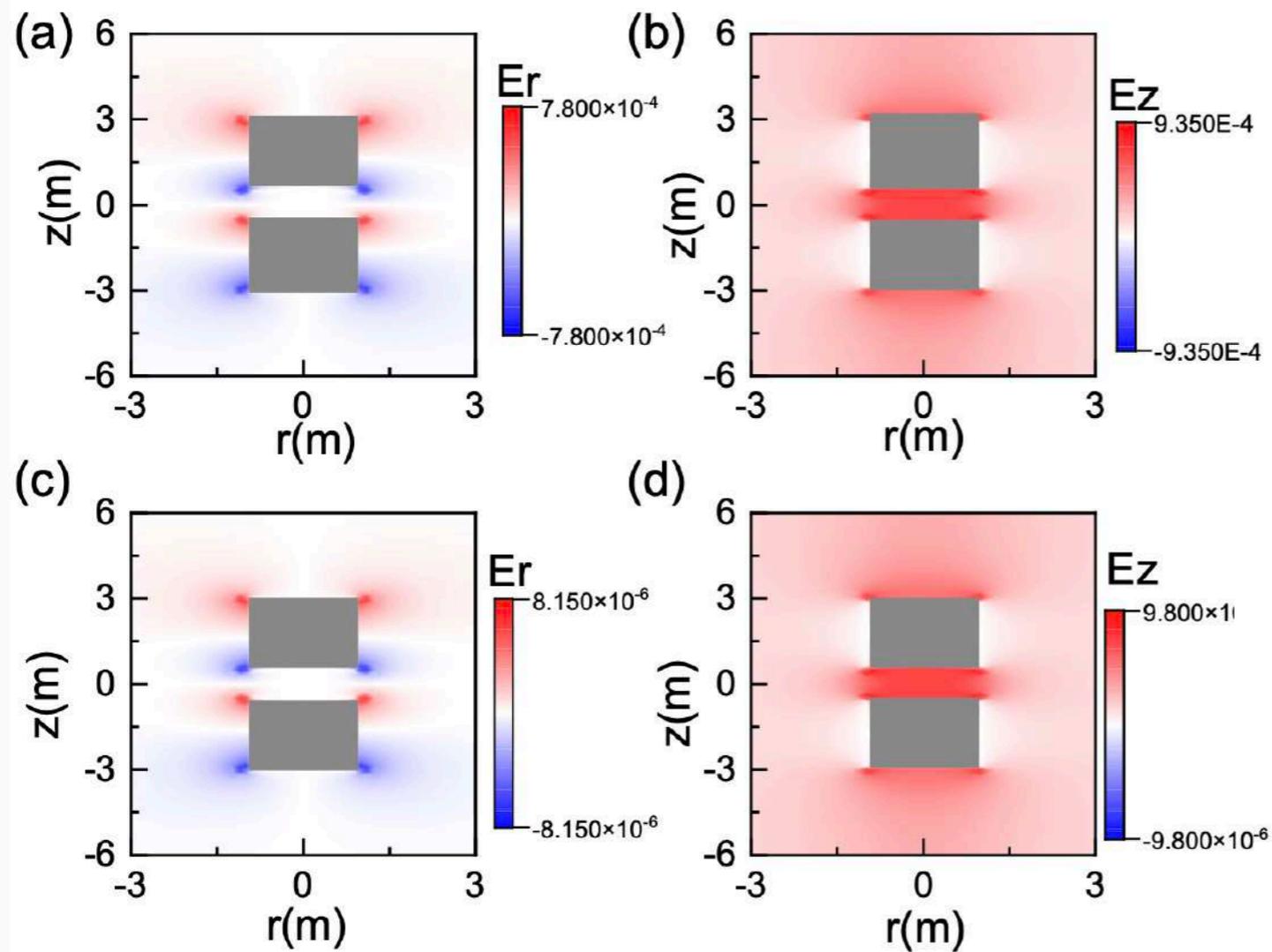


FIG. 4. Validation 2D plots by EM simulation to compare with analytic ansatz. E_z are in the unit of V/m

Numerical solutions:

- With COMSOL Multiphysics:



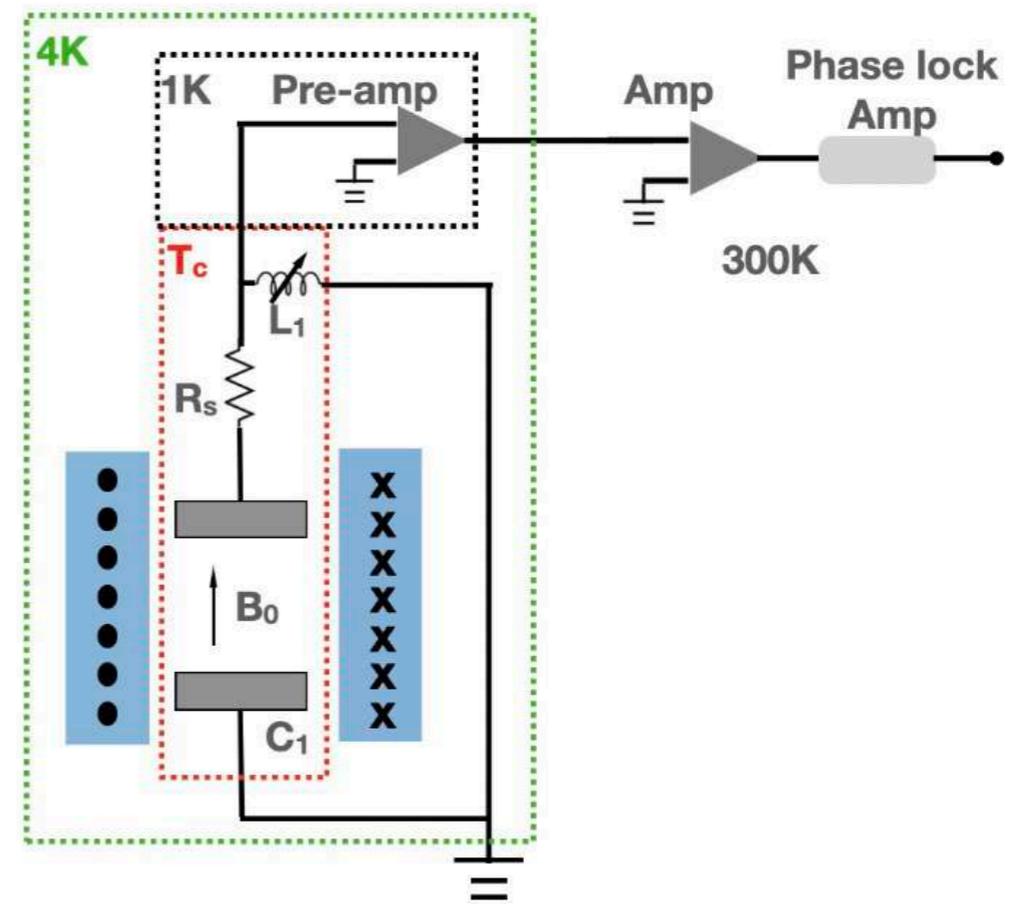
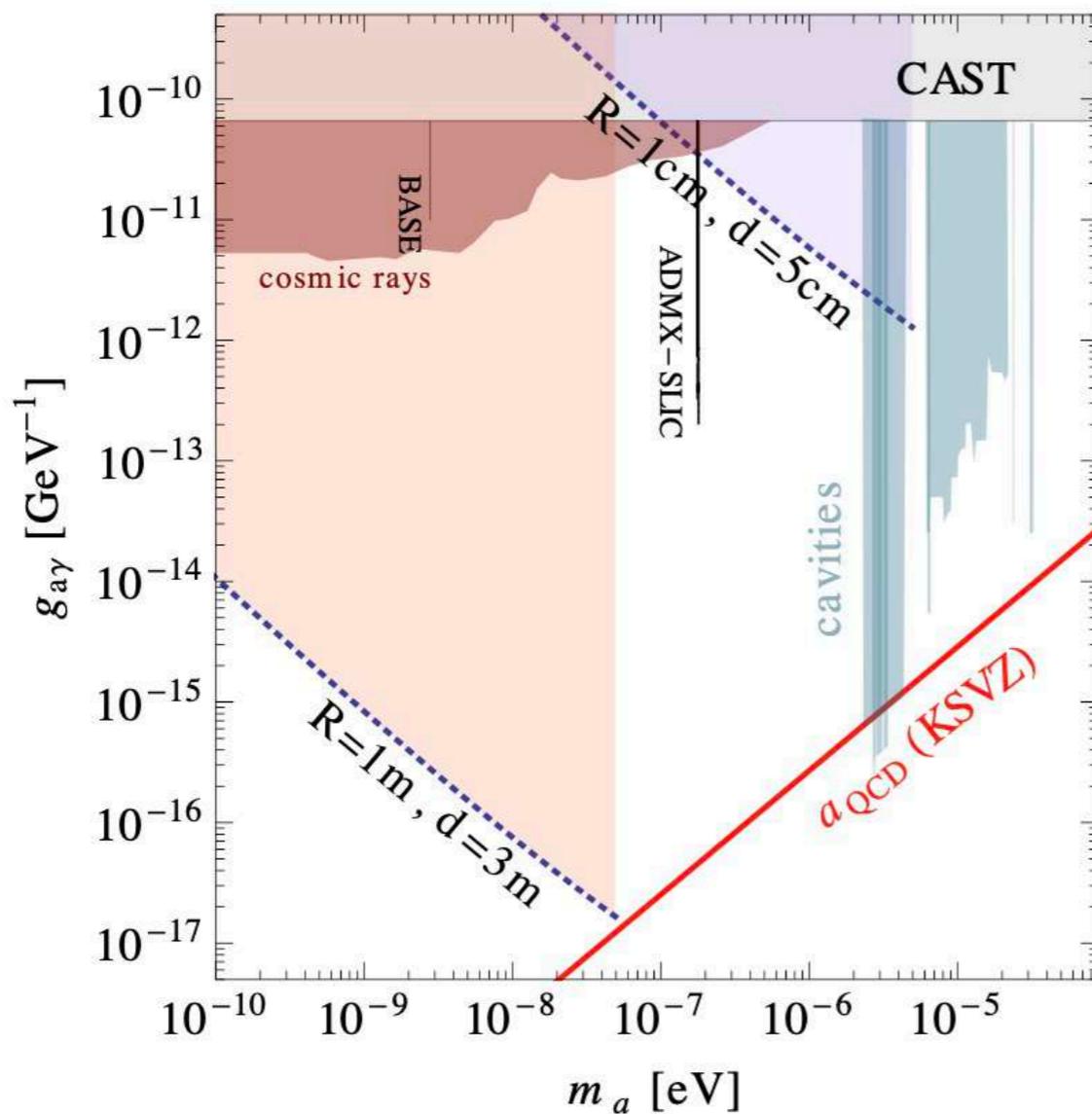
$$\eta(\omega) \equiv \frac{\int \vec{E}_a \cdot d\vec{A}}{\int g_{a\gamma} a \vec{B}_0 \cdot d\vec{A}}$$

ωR	$2R : d = 2 : 1$	$2R : d = 2 : 5$
1	1.81×10^{-2}	1.48×10^{-2}
0.1	1.87×10^{-4}	2.31×10^{-4}
0.01	5.04×10^{-6}	2.58×10^{-6}

Experimental setups: Resonant **E**lectric **A**xion **P**robe(RE-LEAP)

$$\text{SNR} = \frac{P_{\text{sig.}}}{k_B T_N} \sqrt{\frac{\Delta t}{\Delta f}}$$

$$= \frac{Q_c g_{a\gamma}^2 \eta^2 f_c^{-1} \rho_{\text{DM}} B_0^2 (\pi R^2 d)}{k_B T_N} \sqrt{Q_c \frac{2\pi}{m_a^3} \Delta t.}$$



Summary

- We have many ways to detect axions.

Direct (Lab) vs indirect (through gravitational waves!)

Thank you



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Challenges and Opportunities of High Frequency Gravitational Wave Detection

14 - 16 October 2019, Miramare - Trieste, Italy