量子计算与高能核物理交叉前沿讲习班@华南师范大学,广州 2022年11月13日至28日

Quantum Sensing and
searches for new particles— Sapphire: Spin Amplifier for Particle



<u>PHysIcs RE</u>searches

Xinhua Peng Nov. 22, 2022

Univ. of Sci. & Tech of China (USTC)

CAS Key Laboratory of Microscale Magnetic Resonance

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Nobel Prize in Physics 2022



"for experiments with entangled photons, establishing the violation of Bell inequalities and pioneering quantum information science"



Alain Aspect, John Clauser and Anton Zeilinger have each conducted groundbreaking experiments using entangled quantum states, where two particles behave like a single unit even when they are separated. Their results have cleared the way for new technology based upon quantum information.

The second quantum revolution



Break the limits of classical technologies

Quantum precision measurement



Dark energy



Monopoles



Dark matter



Magnetic phenomena and their strength

Sensitivity(T/Hz^{1/2})



Our research



Spin-based quantum precision measurement

Principle : The nucleus or electron spins placed in an external magnetic field, can absorb and release the electromagnetic radiation of the corresponding frequency, then the magnetic resonance phenomenon occurs.



Lamor precession

$$\omega = \gamma B$$

Outline

- Dark Matter and Axion: What is it and Why now?
- Quantum sensors: How to see it?
 - Noble-gas spin systems: intrinsic systems vs. Floquet systems
 - Noble-gas spin amplification effect
 - Noble-gas spin Masing effect
- Dark matter searches with spin-based-amplifier magnetometers
 - Axion dark matter
 - Exotic interactions mediated by new particles (the fifth force)

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Dark matter

Special Section

What is the nature of dark matter?



→ What kind of microscopic
particles are made of?
→ How do particles interact?

WHAT DON'T WE KNOW?

What Is the Universe Made Of

very once in a while, cosmologists are dragged, kicking and screaming, into a universe much more unsettling than they had any reason to expect. In the 1500s and 1600s, Copernicus, Kepler, and Newton showed that Earth is just one of many planets orbiting one of many stars, destroying the comfortable Medieval notion of a closed and tiny cosmos. In the 1920s, Edwin Hubble showed that our universe is constantly expanding and evolving, a finding that eventually shattered the idea that the universe is unchanging and eternal. And in the past few decades, cosmologists have discovered that the ordinary matter that makes up stars and galaxies and people is less than 5% of everything there is. Grappling with this new understanding of <text>

Axions and axion-like particles



Steven Weinberg

Volume 40, Number 4

PHYSICAL REVIEW LETTERS

23 JANUARY 1978

1 JULY 1984

A New Light Boson?

Steven Weinberg

Lyman Laboratory of Physics, Harvard University, Cambridge, Massachusetts 02138 (Received 6 December 1977)

It is pointed out that a global U(1) symmetry, that has been introduced in order to preserve the parity and time-reversal invariance of strong interactions despite the effects of instantons, would lead to a neutral pseudoscalar boson, the "axion," with mass roughly of order 100 keV to 1 MeV. Experimental implications are discussed.





Frank Wilczek

PHYSICAL REVIEW D VOLUME 30, NUMBER 1

New macroscopic forces?

J. E. Moody^{*} and Frank Wilczek Institute for Theoretical Physics, University of California, Santa Barbara, California 93106 (Received 17 January 1984)

The forces mediated by spin-0 bosons are described, along with the existing experimental limits. The mass and couplings of the invisible axion are derived, followed by suggestions for experiments to detect axions via the macroscopic forces they mediate. In particular, novel tests of the *T*-violating axion monopole-dipole forces are proposed.



How to see it?



Quantum sensors: sensitive low-energy tools



Dark matter is very weakly coupled to atomic molecular systems, causing small energy shifts

Science 357, 990 (2017)

Rev. Mod. Phys. 90,025008 (2018)

- Sensitive : quantum-noise limits
- **Tabletop** : small size and cheap
- Networks : low false-alarm rate Beyond the astrophysical limits !

Quantum sensors



Magnetic field: Magnetometers (or spin sensor)



Time standard: Optical clocks

Gravity: Interferometry

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How to see it?

The nonrelativistic Hamiltonian



Quantum spin metrology

$$\omega = \gamma B$$

Alkali-metal atoms K, Rb, Cs
Hyperpolarized noble gas ³He, ¹²⁹Xe

Precision limit:
$$\delta \omega = \frac{1}{P\gamma \sqrt{NT_2 t}}$$

Why we choose alkali metal atoms/noble gas?

- Alkali metal atoms: Well-developed pump and probe techniques
- Noble gas: Long coherence time (>10 s)
- Spin exchange interactions via collisions (Allow information transfer between them)

Coherence time (T_2^*)

³He: I=1/2, T2 >1 h, abundance: 0.01% ¹²⁹Xe: I=1/2, T2~20 s, abundance: 26.4%

²¹Ne: I=3/2, abundance: 21%

 83 Kr: I=9/2, abundance: 83%

¹³¹Xe: I=3/2, T2~5 s, abundance: 21.2%

Atomic magnetometer



Spin dynamics



 $\beta M_0^e \mathbf{P}^e \ll \beta M_0^n \mathbf{P}^n$

Coupled Bloch equations

 \mathbf{n}

$$\frac{\partial \mathbf{P}^{e}}{\partial t} = \frac{\gamma_{e}}{Q} (B_{z}^{0} \hat{z} + \mathbf{B}_{a} + \beta M_{0}^{n} \mathbf{P}^{n}) \times \mathbf{P}^{e} + \frac{P_{0}^{e} \hat{z} - \mathbf{P}^{e}}{T_{e}Q},$$
$$\frac{\partial \mathbf{P}^{n}}{\partial t} = \gamma_{n} (B_{z}^{0} \hat{z} + \mathbf{B}_{a}) \times \mathbf{P}^{n} + \frac{P_{0}^{n} \hat{z} - \mathbf{P}^{n}}{\{T_{2n}, T_{2n}, T_{1n}\}}.$$

Polarized noble-gas nuclear spins generate an effective field experienced by alkali-metal atoms due to the Fermi-contact interactions

 \cap

Fermi-contact factor
$$\kappa_0 \sim 540 \,(\text{Rb+Xe})$$
 $\beta = 8\pi\kappa_0/3$

Spin dynamics



Amplification-assisted magnetometer



Jiang et al., Nature Physics (2021). DOI: 10.1038/s41567-021-01392-z

Spin amplifier: Amplification of magnetic field



→ <u>Continous measurement</u>

Jiang et al., Nature Physics 17, pages 1402–1407 (2021)

Ultrasensitive magnetic field sensing



Jiang et al., Nature Physics 17, pages 1402–1407 (2021)

Periodically driven (Floquet) spin systems



Floquet spin amplification





Amplification factor:

$$\eta_{k,0}(u) = \frac{4\pi}{3} \kappa_0 M^n P_0^n \gamma_n T_{2n} J_k^2(u)$$

10 times improvement in the detection range, fT/Hz^{1/2} detect sensitivity

Jiang et al., PRL 128, 233201 (2022) Editors' suggestion

飞特斯拉 (fT) 级别的自旋传感装置



Noble-gas masing effect

Spin amplification combines with a "cavity-like" feedback



Spin amplification combines with a feedback

Nobel Prize: Maser and atomic clock





C. Townes

G.Basov M. Prokhorov



N. F. Ramsey



Optical frequency (Optical cavity)





Microwave frequency (Cavity)



Radio frequency feedback Coils)

Periodically driven (Floquet) spin maser



Jiang et al. Floquet maser, Science Advances 7, eabe0719 (2021)

Periodically driven (Floquet) spin maser



ScienceAdvances



Floquet maser

D Min Jiang^{1,2,3}, Haowen Su^{1,2,3}, Ze Wu^{1,2,3}, Kinhua Peng^{1,2,3,*} and D Dmitry Budker^{4,5,6}

Floquet-maser-based sensing

1-1000mHz: axion-mass 10⁻¹⁷eV-10⁻¹⁴eV



Jiang et al. Science Advances 7, eabe0719 (2021)

ScienceAdvances

Floquet maser

D Min Jiang^{1,2,3}, D Haowen Su^{1,2,3}, D Ze Wu^{1,2,3}, D Xinhua Peng^{1,2,3,*} and D Dmitry Budker^{4,5,6}

Science A masing ladder

A maser that amplifies emission of periodically modulated quantum states has uses in metrology

PHYS ORG

Extending maser techniques to Floquet systems

techniques to Floquet systems. In their paper published in the journal *Science Advances*, the group describes their approach to creating a new type of maser by amplifying radio frequencies in Floquet systems. Ren-Bao Liu, with the Chinese

physicsworld :: ESEARCH UPDATE

New Floquet maser is very good at detecting low frequency magnetic fields

«Science» Perspectives reports:

"... demonstrate a new type of maser... Conceivable applications of this work include precision clocks and detection of ultralight dark matter particles such as axions"

Spin amplification

SCIENCE CHINA Information Sciences

Review of noble-gas spin amplification via the spin-exchange collisions

Haowen Su^{1,2}, Min Jiang^{1,2} and Xinhua Peng^{1,2}



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Spin Amplifier for Particle PHysIcs Research



Sapphire project ("蓝宝石"计划)

Spin Amplifier for Particle PHysIcs Research



直接探测: 暗物质候选粒子 轴子、暗光子 "蓝宝石"计划: 自旋放大、fT灵敏度、 阵列探测、桌面式 **间接探测**: 新奇相互作用 新粒子作为传播子

Direct detection for ALP

Goal: to detect an oscillating nuclear-spin-dependent energy shift caused by the weak coupling between Spin and Axion

"pseudo-magnetic" field: $ec{B}_{
m ALP} \propto g_{
m aNN} \cos(m_{
m ALP} t) ec{v}$

Resonance frequency probes the mass m_{ALP} Amplitude B_{AIP} probes the coupling value of g_{ANN}



Beyond the astrophysical limits



Search for axion-like dark matter with spin-based amplifiers

Min Jiang^{1,2,3,7}, Haowen Su^{1,2,3,7}, Antoine Garcon^{4,5}, Xinhua Peng^{1,2,3} and Dmitry Budker^{4,5,6}



New constraints on axion-neutron coupling

Jiang et al., Nature Physics 17, pages 1402–1407 (2021)

Beyond the astrophysical limits



At least five orders of magnitude improvement, beyond the astrophysical limits

Jiang et al., Nature Physics 17, pages 1402–1407 (2021)

Exotic interactions beyond the standard model

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B. A. Dobrescu



I. Mocioiu



Extend axion to new mediator bosons and lead to 16 interactions

$$\begin{split} \mathcal{V}_{1} &= \frac{1}{r} y(r) \ , \\ \mathcal{V}_{2} &= \frac{1}{r} \ \vec{\sigma} \cdot \vec{\sigma}' y(r) \ , \\ \mathcal{V}_{3} &= \frac{1}{m^{2} r^{3}} \left[\vec{\sigma} \cdot \vec{\sigma}' \left(1 - r \frac{d}{dr} \right) - 3 \left(\vec{\sigma} \cdot \hat{\vec{r}} \right) \left(\vec{\sigma}' \cdot \hat{\vec{r}} \right) \left(1 - r \frac{d}{dr} + \frac{1}{3} r^{2} \frac{d^{2}}{dr^{2}} \right) \right] y(r) \\ \mathcal{V}_{4,5} &= -\frac{1}{2m r^{2}} \left(\vec{\sigma} \pm \vec{\sigma}' \right) \cdot \left(\vec{v} \times \hat{\vec{r}} \right) \left(1 - r \frac{d}{dr} \right) y(r) \ , \\ \mathcal{V}_{6,7} &= -\frac{1}{2m r^{2}} \left[\left(\vec{\sigma} \cdot \vec{v} \right) \left(\vec{\sigma}' \cdot \hat{\vec{r}} \right) \pm \left(\vec{\sigma} \cdot \hat{\vec{r}} \right) \left(\vec{\sigma}' \cdot \vec{v} \right) \right] \left(1 - r \frac{d}{dr} \right) y(r) \ , \\ \mathcal{V}_{8} &= \frac{1}{r} \left(\vec{\sigma} \cdot \vec{v} \right) \left(\vec{\sigma}' \cdot \vec{v} \right) y(r) \ , \end{split}$$

$$\begin{split} \mathcal{V}_{9,10} &= -\frac{1}{2m r^2} \left(\vec{\sigma} \pm \vec{\sigma}' \right) \cdot \hat{\vec{r}} \left(1 - r \frac{d}{dr} \right) y(r) \ , \\ \mathcal{V}_{11} &= -\frac{1}{m r^2} \left(\vec{\sigma} \times \vec{\sigma}' \right) \cdot \hat{\vec{r}} \left(1 - r \frac{d}{dr} \right) y(r) \ , \\ \mathcal{V}_{12,13} &= \frac{1}{2r} \left(\vec{\sigma} \pm \vec{\sigma}' \right) \cdot \vec{v} y(r) \ , \\ \mathcal{V}_{14} &= \frac{1}{r} \left(\vec{\sigma} \times \vec{\sigma}' \right) \cdot \vec{v} y(r) \ , \\ \mathcal{V}_{15} &= -\frac{3}{2m^2 r^3} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{\vec{r}} \right) \right] \left(\vec{\sigma}' \cdot \hat{\vec{r}} \right) + \left(\vec{\sigma} \cdot \hat{\vec{r}} \right) \left[\vec{\sigma}' \cdot \left(\vec{v} \times \hat{\vec{r}} \right) \right] \right\} \\ & \times \left(1 - r \frac{d}{dr} + \frac{1}{3} r^2 \frac{d^2}{dr^2} \right) y(r) \ , \\ \mathcal{V}_{16} &= -\frac{1}{2m r^2} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{\vec{r}} \right) \right] \left(\vec{\sigma}' \cdot \vec{v} \right) + \left(\vec{\sigma} \cdot \vec{v} \right) \left[\vec{\sigma}' \cdot \left(\vec{v} \times \hat{\vec{r}} \right) \right] \right\} \left(1 - r \frac{d}{dr} \right) y(r) \end{split}$$

B. Dobrescu, I. Mocioiu, JEHP 11, 005 (2006)

How to see it ?



Searched exotic interactions in our study

$$\begin{split} \mathcal{V}_{1} &= \frac{1}{r} y(r) , \\ \mathcal{V}_{2} &= \frac{1}{r} \vec{\sigma} \cdot \vec{\sigma}' y(r) , \\ \mathcal{V}_{3} &= \frac{1}{r} \vec{\sigma} \cdot \vec{\sigma}' (1 - r\frac{d}{dr}) - 3 \left(\vec{\sigma} \cdot \hat{r} \right) \left(\vec{\sigma}' \cdot \hat{r} \right) \left(1 - r\frac{d}{dr} + \frac{1}{3}r^{2}\frac{d^{2}}{dr^{2}} \right) \right] y(r) \\ \mathcal{V}_{3} &= \frac{1}{m^{2}r^{3}} \left[\vec{\sigma} \cdot \vec{\sigma}' \left(1 - r\frac{d}{dr} \right) - 3 \left(\vec{\sigma} \cdot \hat{r} \right) \left(\vec{\sigma}' \cdot \hat{r} \right) \left(1 - r\frac{d}{dr} + \frac{1}{3}r^{2}\frac{d^{2}}{dr^{2}} \right) \right] y(r) \\ \mathcal{V}_{4,5} &= -\frac{1}{2mr^{2}} \left(\vec{\sigma} \pm \vec{\sigma}' \right) \cdot \left(\vec{v} \times \hat{r} \right) \left(1 - r\frac{d}{dr} \right) y(r) , \\ \mathcal{V}_{4,5} &= -\frac{1}{2mr^{2}} \left[\left(\vec{\sigma} \cdot \vec{v} \right) \left(\vec{\sigma}' \cdot \hat{r} \right) \pm \left(\vec{\sigma} \cdot \hat{r} \right) \left(\vec{\sigma}' \cdot \vec{v} \right) \right] \left(1 - r\frac{d}{dr} \right) y(r) , \\ \mathcal{V}_{6,7} &= -\frac{1}{2mr^{2}} \left[\left(\vec{\sigma} \cdot \vec{v} \right) \left(\vec{\sigma}' \cdot \hat{r} \right) \pm \left(\vec{\sigma} \cdot \hat{r} \right) \left(\vec{\sigma}' \cdot \vec{v} \right) \right] \left(1 - r\frac{d}{dr} \right) y(r) , \\ \mathcal{V}_{8} &= \frac{1}{r} \left(\vec{\sigma} \cdot \vec{v} \right) \left(\vec{\sigma}' \cdot \vec{v} \right) y(r) , \\ \mathcal{V}_{8} &= \frac{1}{r} \left(\vec{\sigma} \cdot \vec{v} \right) \left(\vec{\sigma}' \cdot \vec{v} \right) y(r) , \\ \mathcal{V}_{8} &= \frac{1}{r} \left(\vec{\sigma} \cdot \vec{v} \right) \left(\vec{\sigma}' \cdot \vec{v} \right) y(r) , \\ \mathcal{V}_{16} &= -\frac{1}{2mr^{2}} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{r} \right) \right] \left(\vec{\sigma}' \cdot \vec{v} \right) \left[\vec{\sigma}' \cdot \left(\vec{v} \times \hat{r} \right) \right] \right\} \left(1 - r\frac{d}{dr} \right) y(r) \\ \mathcal{V}_{16} &= -\frac{1}{2mr^{2}} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{r} \right) \right] \left(\vec{\sigma}' \cdot \vec{v} \right) \left[\vec{\sigma}' \cdot \left(\vec{v} \times \hat{r} \right) \right] \right\} \left(1 - r\frac{d}{dr} \right) y(r) \\ \mathcal{V}_{16} &= -\frac{1}{2mr^{2}} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{r} \right) \right] \left(\vec{\sigma}' \cdot \vec{v} \right) \left[\vec{\sigma}' \cdot \left(\vec{v} \times \hat{r} \right) \right] \right\} \left(1 - r\frac{d}{dr} \right) y(r) \\ \mathcal{V}_{16} &= -\frac{1}{2mr^{2}} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{r} \right) \right] \left(\vec{\sigma}' \cdot \vec{v} \right] \right\} \left(1 - r\frac{d}{dr} \right) y(r) \\ \mathcal{V}_{16} &= -\frac{1}{2mr^{2}} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{r} \right) \right] \left(\vec{\sigma}' \cdot \vec{v} \right) \left[\vec{\sigma}' \cdot \left(\vec{v} \times \hat{r} \right) \right] \right\} \left(1 - r\frac{d}{dr} \right) y(r) \\ \mathcal{V}_{16} &= -\frac{1}{2mr^{2}} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{r} \right) \right] \left(\vec{\sigma}' \cdot \vec{v} \right) \right\} \left(1 - r\frac{d}{dr} \right) y(r) \\ \mathcal{V}_{16} &= -\frac{1}{2mr^{2}} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{r} \right] \right] \left(\vec{\sigma} \cdot \vec{v} \right) \right\} \right\} \left(1 - r\frac{d}{dr} \right) y(r) \\ \mathcal{V}_{16} &= -\frac{1}{2mr^{2}} \left\{ \left[\vec{\sigma} \cdot \left(\vec{v} \times \hat{r} \right] \right\} \left(\vec{\sigma} \cdot \vec{v} \right) \right\} \left(\vec{\sigma} \cdot \vec{v} \right) \right\}$$

V3 exp.	V11 exp.	V4,5 and V12,13 exp.			
polarized electron/proton (rubidium vapor)	polarized electron/protor (rubidium vapor)	n Unpolarized nucleon (BGO crystal)			
force mediator: axion	force mediator: Z' bosons				

Various spin sources

I: electron/proton (Rb) vapor high polarization Short force range (small size)



Electron/proton-neutron coupling

II: BGO crystal high nucleon density Non-magnetic effect



Nucleon-neutron coupling

Develop spin sources for different exotic interaction searches

Resonant detection with spin amplifiers



Searches for an axion-mediated interaction

Axion-mediated dipole-dipole interaction

$$V_{pp} = -\frac{g_p^1 g_p^2}{4} \left[(\hat{\sigma}_1 \cdot \hat{\sigma}_2) \left(\frac{m_a}{r^2} + \frac{1}{r^3} \right) - (\hat{\sigma}_1 \cdot \hat{r}) (\hat{\sigma}_2 \cdot \hat{r}) \left(\frac{m_a^2}{r} + \frac{3m_a}{r^2} + \frac{3}{r^3} \right) \right] \frac{e^{-m_a r}}{4\pi m_1 m_2}$$



Spurious dipole field: Search remains challenging in the axion window

How to shield the spurious ordinary field?

High-permeability materials: μ -metal



How to shield the spurious ordinary field?



Spurious dipole field is suppressed at least $\sim 10^4$

Search for an axion-mediated interaction



Search for an axion-mediated interaction

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Limits on Axions and Axionlike Particles within the Axion Window Using a Spin-Based Amplifier

Yuanhong Wang, Haowen Su, Min Jiang, Ying Huang, Yushu Qin, Chang Guo, Zehao Wang, Dongdong Hu, Wei Ji, Pavel Fadeev, Xinhua Peng, and Dmitry Budker Phys. Rev. Lett. **129**, 051801 – Published 25 July 2022



The most stringent constraints on $g_p^e g_p^n$ within the axion window

Search for exotic parity-violation interactions



Parity-odd spin-spin interaction mediated by Z' boson

Search for exotic parity-violation interactions



Experimental setup

Constraints on electron-neutron coupling



A five-order-of-magnitude improvement

Constraints on proton-neutron coupling



p – valence protons within the Rb nuclei in the source n – valence neutrons within the Xe nuclei in the sensor

Search for exotic spin-dependent interactions



H. Su,Y. Wang, M. Jiang[†], X. Peng[†] et al. Science Advances 7, eabi9535 (2021).

Search setup

Spin-mass coupling

Rotator



H. Su*, Y. Wang*, M. Jiang#, X. Peng# et al. Science Advances 7, eabi9535 (2021).

Constraints on spin-dependent interactions



At least 2 orders of magnitude improvement on constraints on Z' boson

H. Su,Y. Wang, M. Jiang[†], X. Peng[†] et al. Science Advances 7, eabi9535 (2021).



① DECEMBER 6, 2021

Ultra-high precision search for exotic interactions

by Liu Jia, Chinese Academy of Sciences



In a study published in *Science Advances*, the research team led by Prof. Peng Xinhua from University of Science and Technology of China of the Chinese Academy of Sciences, collaborating with Prof. Dmitry Budker from Helmholtz Institution, realized ultra-high precision search of exotic spin- and velocitydependent interactions beyond the standard model, and <u>amplified the magnetic field</u> signal of exotic interactions at least two technique to the investigation of exotic velocity-'eloped quantum spin-based amplifier.

SAPPHIRE projected sensitivity

³He-K spin amplifier



Solid-state spin source



4 orders of magnitude improvement

 10^{14} cm^{-3} 10^{18} cm^{-3}

<u>8 orders of magnitude improvements are ongoing!</u>

正在开展: aT级别弱磁测量新方法和技术

国际公开报道最高指标 ● <t

基于原子自旋新效应:
 自旋量子放大
 自旋协同效应
 多体量子相变等
 发展超低噪声器件及噪声抑制技术:
 高性能原子气室设计和加工
 (高气压、长寿命、低漏率)
 超导屏蔽技术、超稳激光等

正在开展: "蓝宝石" 计划及暗物质测量网络

✓ 研究趋势:从单个传感器到阵列式;从单一类型到多种类型
 提高灵敏度 ∝ 1/√N
 显著降低报警率∝ p^N
 结构信息



2022年

建成5台探测器阵列(合肥,杭州,苏州,哈尔滨) 已完成2个月数据采集和分析,即将公布结果!

2021年

2019-2020年

基金委原创探索项目(2021-2024)

2022-

Summary



Search for axions, axion-like particles, dark photon Develop sensitive spin amplifier for particle physics (SAPPHIRE)

Exotic fifth force between SM particles

Thanks to

Prof. Jiangfeng Du Prof. Zhengguo Zhao

Collaborators:

Prof. Dmitry Budker Prof. Jing Shu Dr. Yifan Chen Dr. Yue Zhao Dr. Xiao Xue



Dmitry Budker

NMR group (especially lowfield NMR group):

Dr. Min Jiang Haowen Su Yuanhong Wang Qing Li Shiming Song Ze Wu Minxiang Xu Yushu Qing



Min Jiang

Funding: USTC, CAS, NNSFC.....









Thank you for your attention!