FUTURE COLLIDERS FOR HEP

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CEPC Workshop October 24, 2022

1







and

DPF Community Planning Exercise

The Particle Physics Community Planning Exercise (a.k.a. "Snowmass") is organized by the Division of Particles and Fields (DPF) of the American Physical Society. Snowmass is a scientific study. It provides an opportunity for the entire particle physics community to come together to identify and document a scientific vision for the future of particle physics in the U.S. and its international partners. Snowmass will define the most important questions for the field of particle physics and identify promising opportunities to address them.

Snowmass frontiers:

Energy Frontier Neutrino Physics Frontier Rare Processes and Precision Cosmic Frontier Theory Frontier Accelerator Frontier Instrumentation Frontier **Computational Frontier Underground Facilities Community Engagement Snowmass Liaisons Snowmass Early Career**

https://snowmass21.org

With this year-long study, the Snowmass output will provide inputs for the prioritization of the research directions of the field in the decade to come: the "P5" process

(Particle Physics Project Prioritization Panel).

The P5 chair: Prof. Hitoshi Murayama





The Beginning of a new era

The underlying landscape is becoming visible The peak is fundamentally connected to the unseen mountain range The failure of minimal extensions to the Standard Model implies a much richer sector to explore.

With all the new experiments ready to push deeper into every frontier, we all feel the excitement of being poised for an explosion of new data and answers to the questions posed a decade ago.

ZAMAA

BSM as the guide for the future

We are not shooting in the dark.

Let's move away from "The SM is incomplete" and "we need to search everywhere" Instead, we might say

We have made the following discoveries, which will enlarge our model of the Universe once we figure out how it fits into what we already know.

- (1) Dark matter exists and its gravitational interactions are being mapped by a host of cosmic probes
- (2) Neutrinos interact with both gravity and the weak force, and we are measuring their interactions and flavor properties and narrowing down their masses.
- (3) Our Universe has an anti-matter deficit, is accelerating, and is filled with ancient thermal radiation, and we have new tools to measure matter, energy and space-time at all epochs

We have seen many examples in this summer study of how these are all connected. For example: Exploring the Dark Sector has the potential to address the hierarchy problem, neutrino masses, baryon asymmetry, and even strong CP. Signatures will show up in all the frontiers. Cosmic probes and GW are an orthogonal approach to some of the same questions tying the accelerating Universe and large scale structure to fundamental questions in particle physics

This type of message, coupled with specific *benchmark paths* is most compelling (a la T.Slatyer's talk)

July 26, 2022

Highlights and Messages from the Snowmass Summer Study. Prisca Cushman

LHC will continue at the energy frontier



- Already ~ $O(140 \text{ fb}^{-1})$ @ ATLAS/CMS
- Run 3 up-coming: 300 fb⁻¹
- HL-LHC: 3000 fb⁻¹
 → lead the energy/precision frontier!

Further searches at the LHC will be limited by

- Backgrounds
- Systematics
- New physics threshold

Snowmass'21 Accelerator Frontier Conveners



Steve Gourlay (LBNL)



Tor Raubenheimer (SLAC)



Vladimir Shiltsev (FNAL)

| Topical Group | | Topical Group co-Conveners | | | | |
|---------------|--|----------------------------|--------------------|--|----------------------|--|
| AF01 | Beam Phys & Accel. Education | Z. Huang (Stanford) | M. Bai (SLAC) | S. Lund (MSU) | | |
| AF02 | Accelerators for Neutrinos | J. Galambos (ORNL) | B. Zwaska (FNAL) | G. Arduini (CERN) | | |
| AF03 | Accelerators for EW/Higgs | F. Zimmermann (CERN) | Q. Qin (ESRF) | G.Hoffstaetter (Cornell) A.Faus-Golfe (IN2P3) | | |
| AF04 | Multi-TeV Colliders | M. Palmer (BNL) | A. Valishev (FNAL) | N. Pastrone (INFN) J.Tang (IHEP) | | |
| AF05 | Accelerators for PBC and Rare Processes | E. Prebys (UC Davis) | M. Lamont (CERN) | Richard Milner (MIT) | | |
| AF06 | Advanced Accelerator Concepts | C. Geddes (LBNL) | M. Hogan (SLAC) | P. Musumeci (UCLA) | R. Assmann (DESY) | |
| AF07 | Accelerator Technology R&D | | | | | |
| | Sub-Group RF | E. Nanni (LBNL) | H.Weise (DESY) | S. Belomestnykh (FNAL) | | |
| | Sub-Group Magnets | G. Sabbi (LBNL) | S. Zlobin (FNAL) | S. Izquierdo Bermudez (CERN) | | |
| | Sub-Group Targets/Sources | C. Barbier (ORNL) | Y. Sun (ANL) | Frederique Pellemoine (FNAL) | | |

Implementation Task Force

https://arxiv.org/abs/2208.06030

- The Accelerator Frontier Implementation
 Task Force (ITF) is charged with developing metrics and processes to facilitate a comparison between collider projects:
 - Higgs/EW factories (focus of slides below)
 - Lepton colliders with 3 TeV cme
 - Lepton and hh colliders 10+ TeV cme
 - FNAL site-fillers and eh colliders
- ITF addressed (four subgroups):
 - Physics reach (impact), beam parameters
 - Size, complexity, power, environment
 - Technical risk, readiness, and R&D required
 - Cost and schedule







Thomas Roser (BNL, Chair)

Philippe Lebrun (CERN)









Tor Raubenheimer (SLAC)

Katsunobu Oide (KEK)

Jim Strait (FNAL)







Vladimir Shiltsev (FNAL)

Reinhard Brinkmann (DESY)

John Seeman (SLAC)



FUTURE COLLIDERS UNDER DISCUSSIONS*

Snowmass 2021 Energy Frontier Collider Study Scenarios

| Ī | Collider | Type | \sqrt{s} | P [%] | Lint |
|---|---------------------------|------------------------|--------------------|------------------------------------|--|
| | | | | e^-/e^+ | ab^{-1} |
| | HL-LHC | $\mathbf{p}\mathbf{p}$ | 14 TeV | | 6 |
| | | | 250 C.M | $\pm 80 / \pm 20$ | 9 |
| q | ILC | ee | 250 GeV | $\pm 80/\pm 30$ $\pm 80/\pm 20$ | $\begin{pmatrix} 2 \\ 0 \end{pmatrix}$ |
| | | | 500 GeV | $\pm 80/\pm 30$ | 0.2 |
| | | | 1 TeV | $\pm 80/\pm 30$ | 4 |
| | | | 1 Iev | $\pm 60/ \pm 20$ | • |
| ľ | CLIC | ee | 380 GeV | $\pm 80/0$ | 1 |
| | | | 1.5 TeV | $\pm 80/0$ | 2.5 |
| | | | $3.0 \mathrm{TeV}$ | $\pm 80/0$ | 5 |
| | CEDC | 00 | M | | 16 |
| | CEFC | ee | M_Z | | 10 |
| | | | 2MW | | 2.0 |
| | | | 240 Gev | | 0.0 |
| ľ | FCC-ee | ee | M_Z | | 150 |
| | | | $2M_W$ | | 10 |
| | | | 240 GeV | | 5 |
| ļ | | | $2 M_{top}$ | - / - | 1.5 |
| Ì | FCC-hh | pp | 100 TeV | | 30 |
| | | | | | |
| | LHeC | ep | 1.3 TeV | | 1 |
| | FCC-eh | ep | 3.5 TeV | | 2 |
| | | | | | |
| | muon-collider (higgs) | $\mu\mu$ | 125 GeV | | 0.02 |
| | | | | | |
| (| High energy muon-collider | $\mu\mu$ | 3 TeV | | 1 |
| | | | 10 TeV | | 10 |
| | | | 14 TeV | | 20 |
| | | | 30 TeV | | 90 |

* Snowmass Energy Frontier: https://snowmass21.org

Proposals – Higgs/EW Physics

Higgs factory concepts (10)

| Name | CM energy range |
|---|--|
| FCC-ee | $e+e-, \sqrt{s} = 0.09 - 0.37 \text{ TeV}$ |
| CEPC | e+e-, $\sqrt{s} = 0.09 - 0.37$ TeV |
| ILC (Higgs factory) | e+e-, $\sqrt{s} = 0.09 - 1$ TeV |
| CLIC (Higgs factory) | e+e-, √s = 0.09 – 1 TeV |
| CCC (Cool Copper Collider) | e+e-, $\sqrt{s} = 0.25 - 0.55$ TeV |
| CERC (Circular ERL collider) | e+e-, $\sqrt{s} = 0.09 - 0.60$ TeV |
| ReLiC (Recycling Linear Collider) | e+e-, $\sqrt{s} = 0.25 - 1$ TeV |
| ERLC (ERL Linear Collider) | e+e-, $\sqrt{s} = 0.25 - 0.50$ TeV |
| XCC (FEL-based $\gamma\gamma$ collider) | ee ($\gamma\gamma$), $\sqrt{s} = 0.125 - 0.14$ TeV |
| MC (Higgs factory) | μ + μ -, \sqrt{s} = 0.13 TeV |





9

ILC (International Linear Collider) as a Higgs Factory & beyond Under serious consideration in Japan https://arxiv.org/abs/1901.09829



Ecm = 250 GeV / 2 ab⁻¹/yr: a Higgs factory = 500 GeV / 4 ab⁻¹/yr: a top-quark factory = 1000 GeV / 8 ab⁻¹/yr: new particle threshold

FCC (future circular collider): CERN



https://arxiv.org/abs/1607.01831, https://arxiv.org/abs/1606.00947; Arkani-Hamed, TH, Mangano, LT Wang, Phys. Rept. 1511.06495.

CEPC (Circular e⁻e⁺) / SppC (Super pp), China



Similar physics goals to FCC-ee, FCC-hh ! https://arxiv.org/abs/1811.10545

250 GeV cme Fermilab Site-Fillers



16-km collider e+e- ring

https://arxiv.org/abs/2203.08088

cool- or *SC-RF* e+e- linear colliders 7-km for 250 GeV, 12-km 0.5+ TeV



Implementation Task Force on Higgs Factories

Table I - ITF Report – T.Roser, et al, arXiv:2208.06030

| | | CME (TeV) | Lumi per IP@ Higgs (10^34) | Years, pre- project R&D | Years to 1 st Physics | Cost Range (2021 B\$) | Electric Power (MW) |
|-----------------------|---------------|--------------|----------------------------------|-------------------------------|--|-----------------------------|---------------------------|
| Circular <i>e+e</i> - | FCCee (4 IPs) | 0.24 | 7.7 | 0-2 | 13-18 | 12-18 | 290 |
| | CEPC (2 IPs) | 0.24 | 8.3 | 0-2 | 13-18 | 12-18 | 340 |
| | FermiHF | 0.24 | 1.2 | 3-5 | 13-18 | 7-12 | ~200 |
| Linear <i>e+e</i> - | ILC | 0.25 | 2.7 | 0-2 | <12 | 7-12 | 110 |
| | CLIC | 0.38 | 2.3 | 0-2 | 13-18 | 7-12 | 150 |
| | C^3 | 0.25 | 1.3 | 3-5 | 13-18 | 7-12 | 150 |
| | HELEN | 0.25 | 1.4 | 5-10 | 13-18 | 7-12 | ~110 |
| ERL-based | CERC | 0.24 | 78 | 5-10 | 19-24 | 12-30 | 90 |
| | ReLiC (2 IPs) | 0.24 | 165 | 5-10 | >25 | 7-18 | 315 |
| | ERLC | 0.24 | 90 | 5-10 | >25 | 12-18 | 250 |
| s-chan | ΧϹϹ-γγ | 0.125 | 0.1 | 5-10 | 19-24 | 4-7 | 90 |
| | μμ-Higgs | 0.13 | 0.01 | >10 | 19-24 | 4-7 | 200 |

ITF's Look Beyond Higgs Factories

| | CME (TeV) | Lumi per IP (10^34) | Years, pre- project R&D | Years to 1 st Physics | Cost Range (2021 B\$) | Electric Power (MW) |
|-------------------|--------------|---------------------------|-------------------------------|--|-----------------------------|---------------------------|
| FCCee-0.24 | 0.24 | 8.5 | 0-2 | 13-18 | 12-18 | 290 |
| ILC-0.25 | 0.25 | 2.7 | 0-2 | <12 | 7-12 | 140 |
| CLIC-0.38 | 0.38 | 2.3 | 0-2 | 13-18 | 7-12 | 110 |
| HELEN-0.25 | 0.25 | 1.4 | 5-10 | 13-18 | 7-12 | 110 |
| CCC-0.25 | 0.25 | 1.3 | 3-5 | 13-18 | 7-12 | 150 |
| E CERC(ERL) | 0.24 | 78 | 5-10 | 19-24 | 12-30 | 90 |
| CLIC-3 | 3 | 5.9 | 3-5 | 19-24 | 18-30 | ~550 |
| ILC-3 | 3 | 6.1 | 5-10 | 19-24 | 18-30 | ~400 |
| MC-3 | 3 | 2.3 | >10 | 19-24 | 7-12 | ~230 |
| MC-10-IMCC | 10-14 | 20 | >10 | >25 | 12-18 | O(300) |
| FCChh-100 | 100 | 30 | >10 | >25 | 30-50 | ~560 |
| Collider-in-Sea | 500 | 50 | >10 | >25 | >80 | »1000 |

Why muon colliders?

- (Snowmass Energy Frontier) HEP aspires 10+ TeV cme/parton
- Muon Collider is a viable option for the HEP future:
 - Combines discovery reach and precision physics
 - x7 energy reach vs pp eg 14 TeV μμ = 100 TeV pp
 - μ 's do not radiate when bent \rightarrow acceleration in rings:
 - Smaller(est) footprint 10-15 km vs 50-100 km
 - (Best) power efficiency Lumi/Power grows with energy
 - Low(est) cost due to compactness and power efficiency
- (ITF) 3-10 TeV Muon Collider can be designed in ~10-15 yrs and built in 20-25 yrs from now:
 - Past studies in the US and UK (+now in CERN) big advance
 - No insurmountable obstacles identified
 - But challenging technologies and design require R&D

International Muon Collider Collaboration



https://muoncollider.web.cern.ch

Fermilab on site:



Daniel Schulte; Mark Palmer; Katsuya Yonehara talk, March 2022

Muon collider on FNAL site: It may reach ~ m_h – 10 TeV



(Curtesy of Pushpa Bhat) Snowmass Day: https://indico.fnal.gov/event/50538

Accelerator Frontier Recommendation (to P5)

On Colliders: We need an integrated future collider R&D program to engage in the design and to coordinate the development of the next generation collider projects:

- to address in an integrated fashion the technical challenges of promising future collider concepts, that are not covered by the existing *General Accelerator R&D* (GARD) program.
- to enable synergistic U.S. engagement in ongoing global efforts (e.g., FCC, ILC, IMCC)
- to develop collider concepts and proposals for options feasible to be hosted in the U.S. (e.g., CCC, HELEN, Muon Collider, etc)

PHYSICS POTENTIAL @ FUTURE COLLIDERS Precision Higgs Physics: $(v/\Lambda)^2 < 6\%$

ILC: $E_{cm} = 250 (500) \text{ GeV}, 250 (500) \text{ fb}^{-1}$

- Model-independent measurement: $\Gamma_{\rm H} \sim 6\%$, $\Delta m_{\rm H} \sim 30$ MeV, $\Delta k_{{\rm W},Z} < 1\%$ (HL-LHC: assume SM, $\Gamma_{\rm H} \sim 5-8\%$, $\Delta m_{\rm H} \sim 50$ MeV)
- Higgs Factory: 10⁶ Higgs: $\Gamma_{\rm H} \sim 1\%$, $\Delta m_{\rm H} \sim 5$ MeV.
- FCC-hh / SPPC: $\Delta k_{HHH} \sim 5\%$ LARE W phase transition!
- 14 TeV muon collider: $\Delta k_{\rm HHH} \sim 3\%$, $\Delta k_{\rm W,Z} < 0.5\%$, $Y_{\mu} \sim 1\%$

ILC: arXiv:1710.07621; TLEP Report: 1308.6176; FCC: Arkani-Hamed, TH, Mangano, LT Wang, 1511.06495; muC: TH, D. Liu, I. Low, X. Wang, arXiv:2008.12204.

Pushing the "Naturalness" limit

contours of the two different search strategies.



han the stop. Top quark partners searches: The Higgs mass fine-tune: $\delta m_H/m_H \sim 1\% (1 \text{ TeV}/\Lambda)^2$ Thus, $m_{stop} > 8 \text{ TeV} \rightarrow 10^{-4}$ fine-tune! then, scalar TOPPARTNER FCC: Arkani-Hamed, TH, Mangano, LT Wang, 1511.06495; muC: The Muon Smasher's Guide, https://arxiv.org/abs/2103.14043

DM Searches

WIMP DM: mass bounded by the thermal relic

 $M_{\rm DM} < 1.8 \,\,{\rm TeV}\left(\frac{g_{\rm eff}^2}{0.3}\right)$

24

Muon Collider 2σ Reach ($\sqrt{s} = 3, 6, 10, 10, 30, 100$ TeV)

New Particle Searches

Electroweak Resonances: Z',W'

Colored Resonances:

Several Useful References:

- 1. Overview of accelerators V.Shiltsev, Physics Today 73, 4, 32 (2020).
- 2. RMP colliders V.Shiltsev, F.Zimmermann, <u>Rev.Mod.Phys. 93, 015006 (2021)</u>.
- 3. Ultimate limits of colliders V.Shiltsev, Proc. IPAC'21, WEPAB017 (2021).
- 4. Snowmass Accelerator Frontier report <u>arxiv:2209.14136</u>
- 5. ITF Report T.Roser, V.Shiltsev, et al, arXiv:2208.06030
- 6. αβγ cost model V.Shiltsev, <u>JINST 9 T07002 (2014).</u>
- 7. Crystal collider V.Shiltsev, Physics Uspekhi, 55 (10), 1033 (2012).
- 8. CPT-theorem V.Shiltsev, Mod. Phys. Lett. A, 26, 11, 761 (2011)