

Introduction of CAS & ITP

🎵 Chinese Academy of Sciences (CAS)

- ✧ Independent of Ministry of Education, but award degrees
- ✧ >100 institutes in China; ~50 in Beijing

🎵 Institute of Theoretical Physics (ITP)

- ✧ Smallest institute in CAS; founded in 1978
- ✧ ~35 permanent staffs; ~20 postdocs; ~100 students
- ✧ Atomic, nuclear, particle, cosmology, condensed matter, biophysics, statistics, quantum physics & quantum information

- ✧ Kavli Institute for Theoretical Physics China: running programs

🎵 Theoretical Nuclear Physics Group

- ✧ En-Guang Zhao, SGZ, 1 postdoc fellow, 4 students
- ✧ Heavy-ion reactions & synthesis mechanism of superheavy nuclei
- ✧ Structure of exotic nuclei & superheavy nuclei

Neutron halo in deformed nuclei

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International Symposium on Nuclear Physics in Asia*

Outline

🎵 Introduction

🎵 Relativistic Hartree (Bogoliubov) model for exotic nuclei

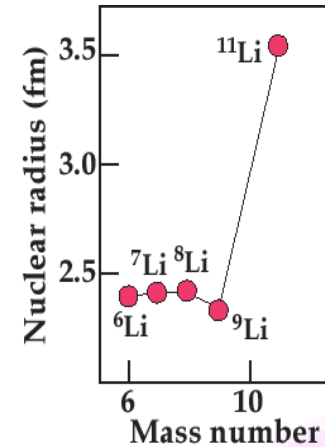
- ✧ RMF - a brief introduction
- ✧ Deformed relativistic Hartree-Bogoliubov model in a Woods-Saxon basis

🎵 Neutron halo in deformed nuclei: ^{44}Mg

- ✧ **Density distributions**; **single particle states** in canonical basis
- ✧ **Decoupling** between deformations of core & halo
- ✧ **Mechanism of the decoupling**

🎵 Summary & perspectives

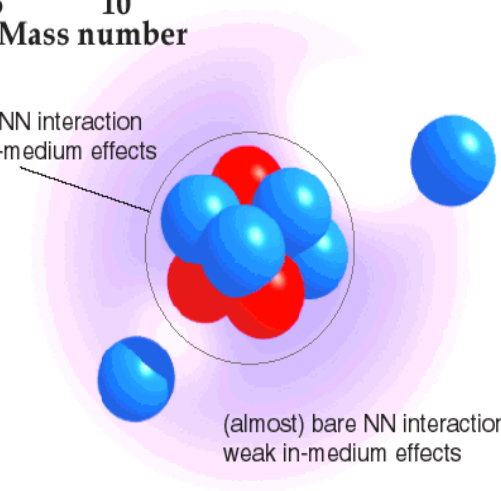
Halo in spherical nuclei



I. Tanihata et al.
Phys. Rev. Lett. 55, 2676 (1985)

Interaction cross section
measurements at Bevalac
(790 MeV/u)

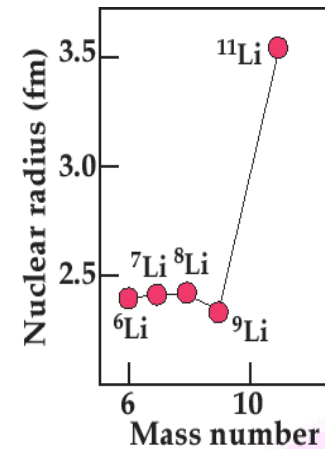
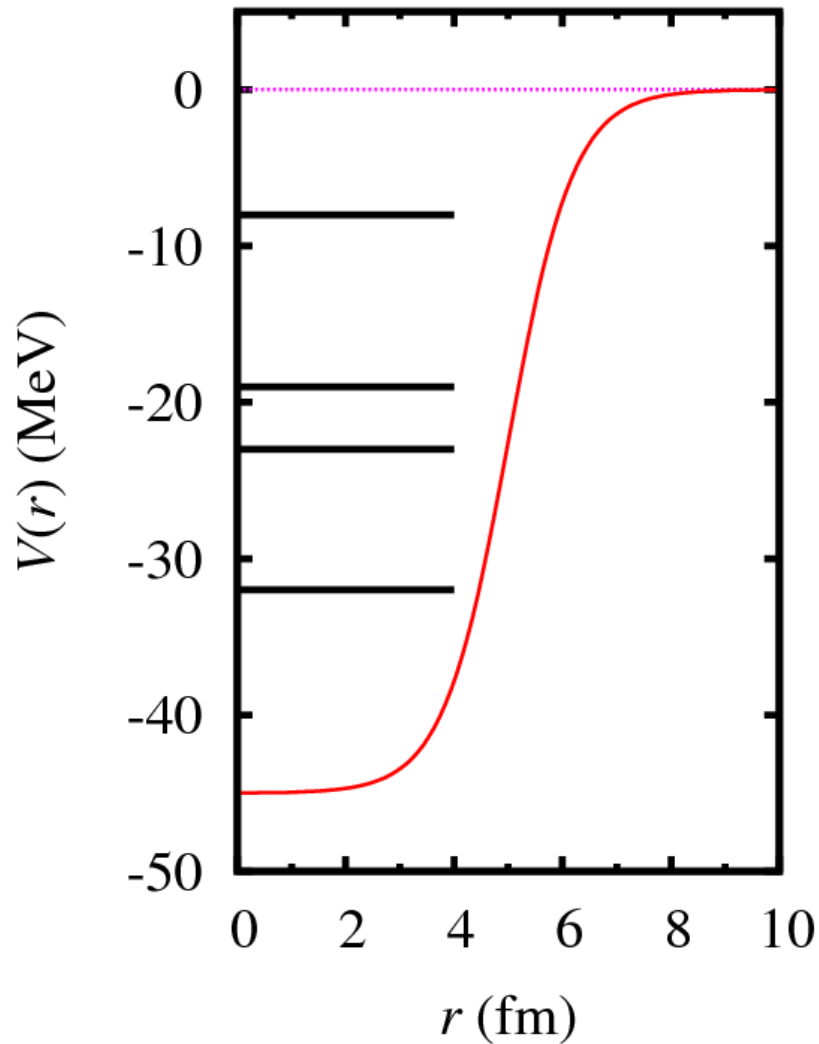
effective NN interaction
strong in-medium effects



(almost) bare NN interaction
weak in-medium effects

Halo in spherical nuclei

SPL in mean field model



I. Tanihata et al.
Phys. Rev. Lett. 55, 2676 (1985)

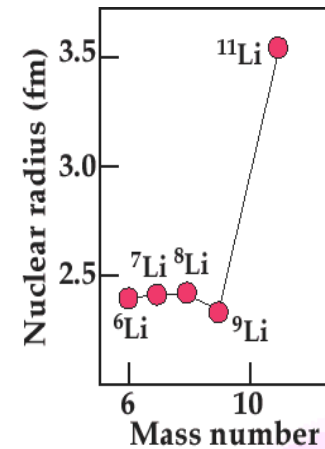
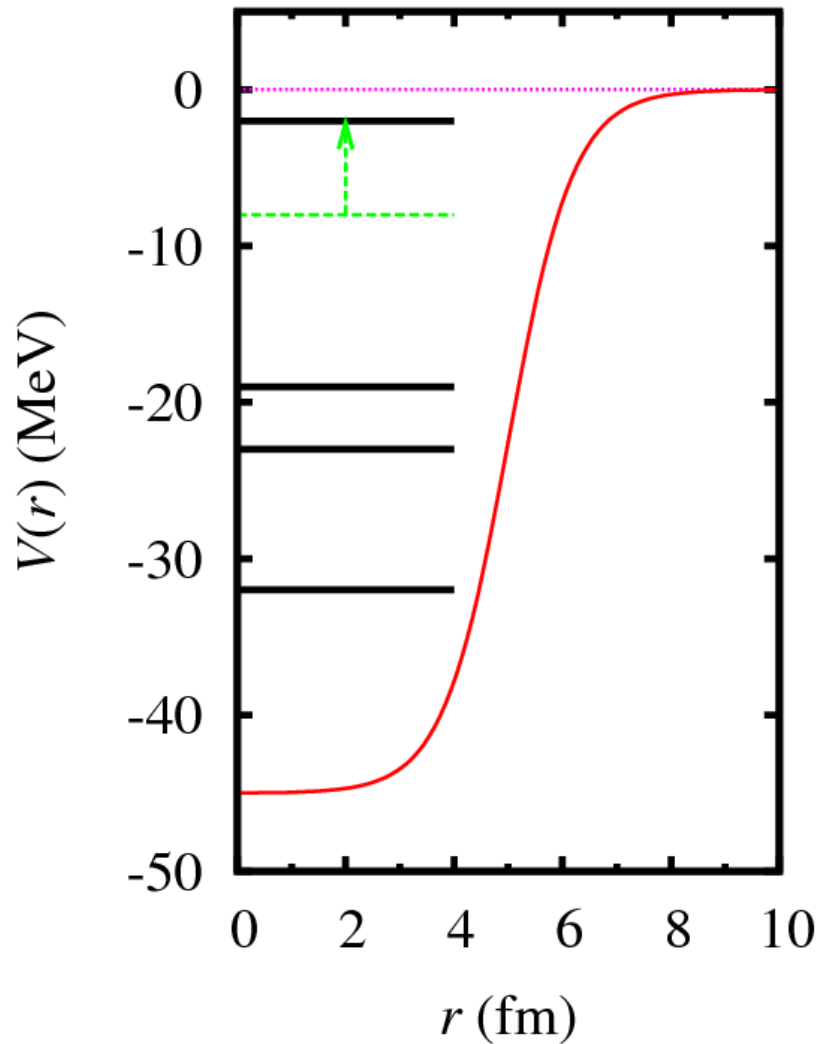
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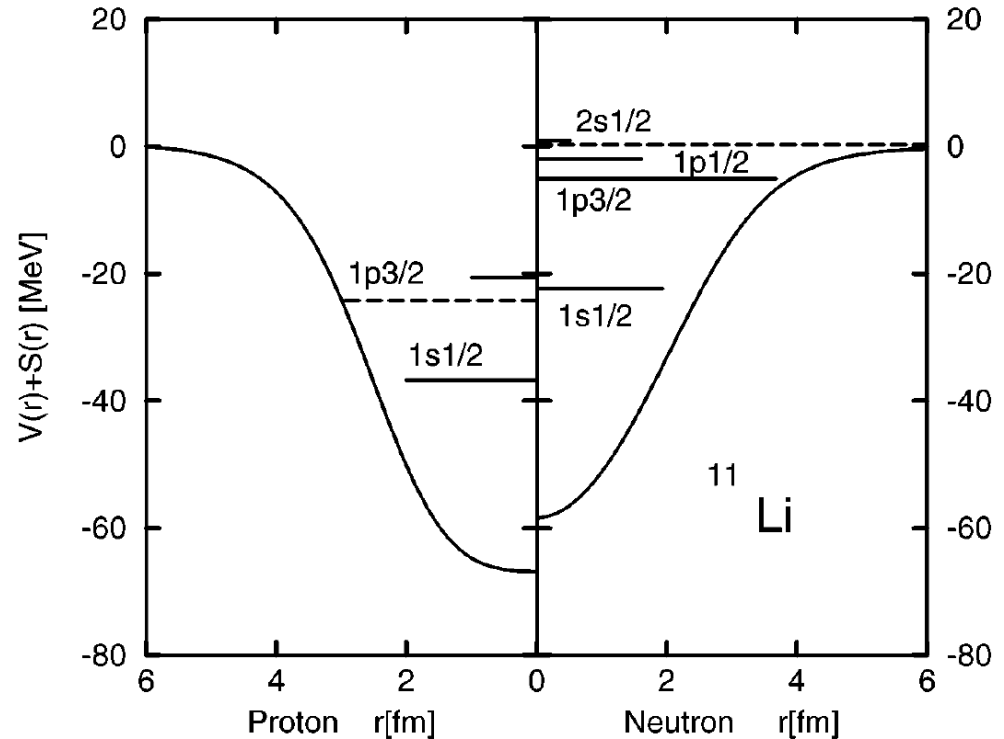
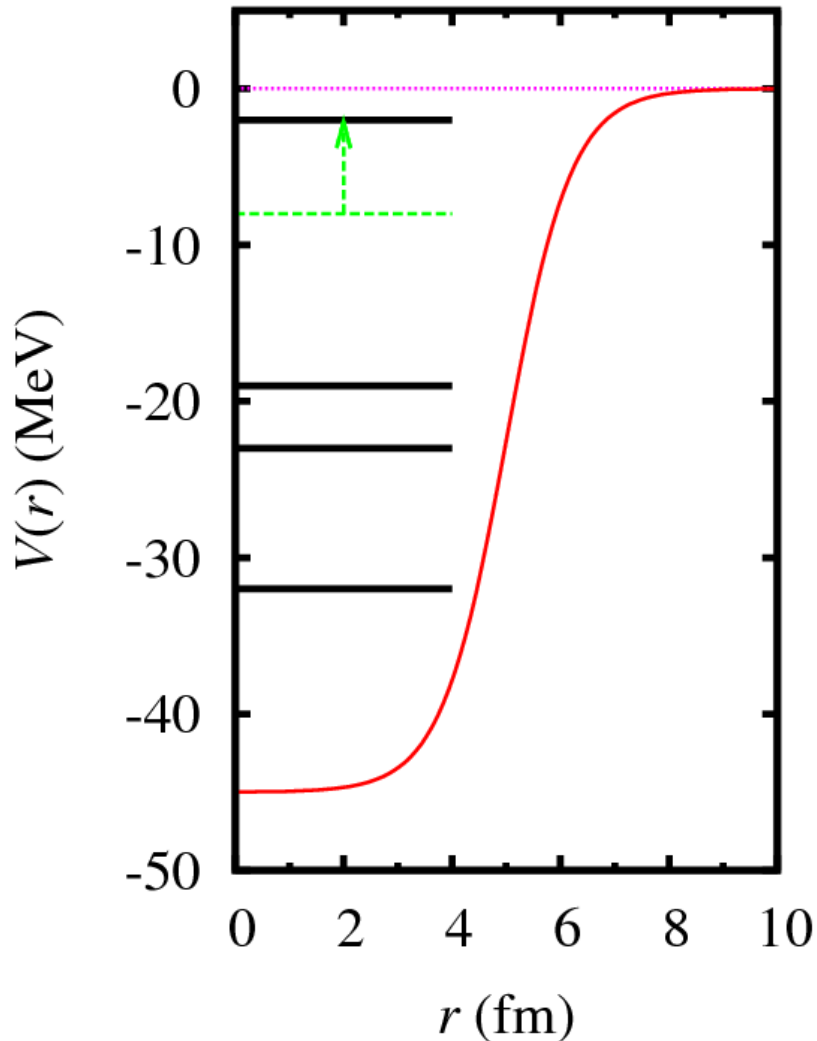
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SPL in mean field model



Canonical SPL in self-consistent r -space relativistic Hartree-Bogoliubov model

Meng & Ring, PRL77(1996)3963

Hal o i n deformed nuclei? deformed hal o?

Halo in deformed nuclei? deformed halo?

- ♪ **Deformed single particle model: square well, spin-orbit interaction neglected**

“in the limit of very small binding energy the valence particles in specific orbitals, characterized by a very small Λ , can give rise to the halo structure which is completely decoupled from the rest of the system”

Misu, Nazarewicz, Aberg, NPA614(97)44

- ♪ **Deformed Skyrme Hartree-Fock model: by blocking the first 1/2+ neutron level, a large deformation is found**

“Such a large deformation can result in a good agreement between the calculated and experimental density distributions for the nucleus”

Pei, Xu & Stevenson, NPA765(06)29

♪ ...

What we aim at

A self-consistent description of :

- ✓ Deformation
- ✓ Continuum contribution
- ✓ Large spatial distribution
- ✓ Interplay between them

Relativistic mean field model

$$\begin{aligned} L = & \bar{\psi}_i (i\gamma_\mu \partial^\mu - M) \psi_i + \frac{1}{2} \partial_\mu \sigma \partial^\mu \sigma - U(\sigma) - g_\sigma \bar{\psi}_i \sigma \psi_i \\ & - \frac{1}{4} \Omega_{\mu\nu} \Omega^{\mu\nu} + \frac{1}{2} m_\omega^2 \omega_\mu \omega^\mu - g_\omega \bar{\psi}_i \gamma_\mu \omega^\mu \psi_i \\ & - \frac{1}{4} \bar{R}_{\mu\nu} \bar{R}^{\mu\nu} + \frac{1}{2} m_\rho^2 \bar{\rho}_\mu \bar{\rho}^\mu - g_\rho \bar{\psi}_i \gamma_\mu \bar{\rho}^\mu \bar{\tau} \psi_i \\ & - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - e \bar{\psi}_i \frac{1-\tau_3}{2} \gamma_\mu A^\mu \psi_i \end{aligned}$$

References:

Serot & Walecka, Adv. Nucl. Phys. 16 (86) 1

Reinhard, Rep. Prog. Phys. 52 (89) 439

Ring, Prog. Part. Nucl. Phys. 37 (96) 193

Vretenar, Afanasjev, Lalazissis & Ring, Phys. Rep. 409 (05) 101

Meng, Toki, SGZ, Zhang, Long & Geng, Prog. Part. Nucl. Phys. 57 (06) 470

Deformed RHB in a Woods-Saxon basis

Kucharek&Ring, ZPA339(91)23

$$\sum_{\sigma p} \int d^3 r' \begin{pmatrix} h(\mathbf{r}\sigma p; \mathbf{r}'\sigma' p') - \lambda & \Delta(\mathbf{r}\sigma p; \mathbf{r}'\sigma' p') \\ -\Delta^*(\mathbf{r}\sigma p; \mathbf{r}'\sigma' p') & -h(\mathbf{r}\sigma p; \mathbf{r}'\sigma' p') + \lambda \end{pmatrix} \begin{pmatrix} U_E(\mathbf{r}'\sigma' p') \\ V_E(\mathbf{r}'\sigma' p') \end{pmatrix} = E \begin{pmatrix} U_E(\mathbf{r}\sigma p) \\ V_E(\mathbf{r}\sigma p) \end{pmatrix}$$

SGZ, Meng & Ring, PRC68(03)034323

Axially deformed nuclei

$$\beta_{km}^+ = \sum_{(i\kappa)} u_{k,(i\kappa)}^{(m)} a_{i\kappa m}^+ + v_{k,(i\tilde{\kappa})}^{(m)} \tilde{a}_{i\kappa m}$$

$$\begin{pmatrix} U_k^{(m)}(\mathbf{r}\sigma p) \\ V_k^{(m)}(\mathbf{r}\sigma p) \end{pmatrix} = \sum_{i\kappa} \begin{pmatrix} u_{k,(i\kappa)}^{(m)} \varphi_{i\kappa m}(\mathbf{r}\sigma p) \\ v_{k,(i\tilde{\kappa})}^{(m)} \tilde{\varphi}_{i\kappa m}(\mathbf{r}\sigma p) \end{pmatrix}$$

$$\varphi_{i\kappa m}(\mathbf{r}\sigma p) = \frac{1}{r} \begin{pmatrix} iG_{i\kappa}(r) Y_{\kappa m}(\Omega\sigma) \\ -F_{i\kappa}(r) Y_{\kappa m}(\Omega\sigma) \end{pmatrix}$$

$$\begin{pmatrix} \mathbf{A} & \mathbf{B} \\ \mathbf{C} & \mathbf{D} \end{pmatrix} \begin{pmatrix} \mathbf{U} \\ \mathbf{V} \end{pmatrix} = E \begin{pmatrix} \mathbf{U} \\ \mathbf{V} \end{pmatrix}$$

$$\mathbf{U} = \left(u_{k,(i\kappa)}^{(m)} \right)$$

$$\mathbf{V} = \left(v_{k,(i\tilde{\kappa})}^{(m)} \right)$$

Parameter set for ph & pp channels

NL3, $R_{\max} = 20$ fm, $\Delta r = 0.1$ fm

$$V^{\text{pair}} = \frac{1}{4} V_0 \delta(\mathbf{r}_1 - \mathbf{r}_2) \left(1 - \frac{\rho(\mathbf{r}_1)}{\rho_0}\right)^\gamma$$

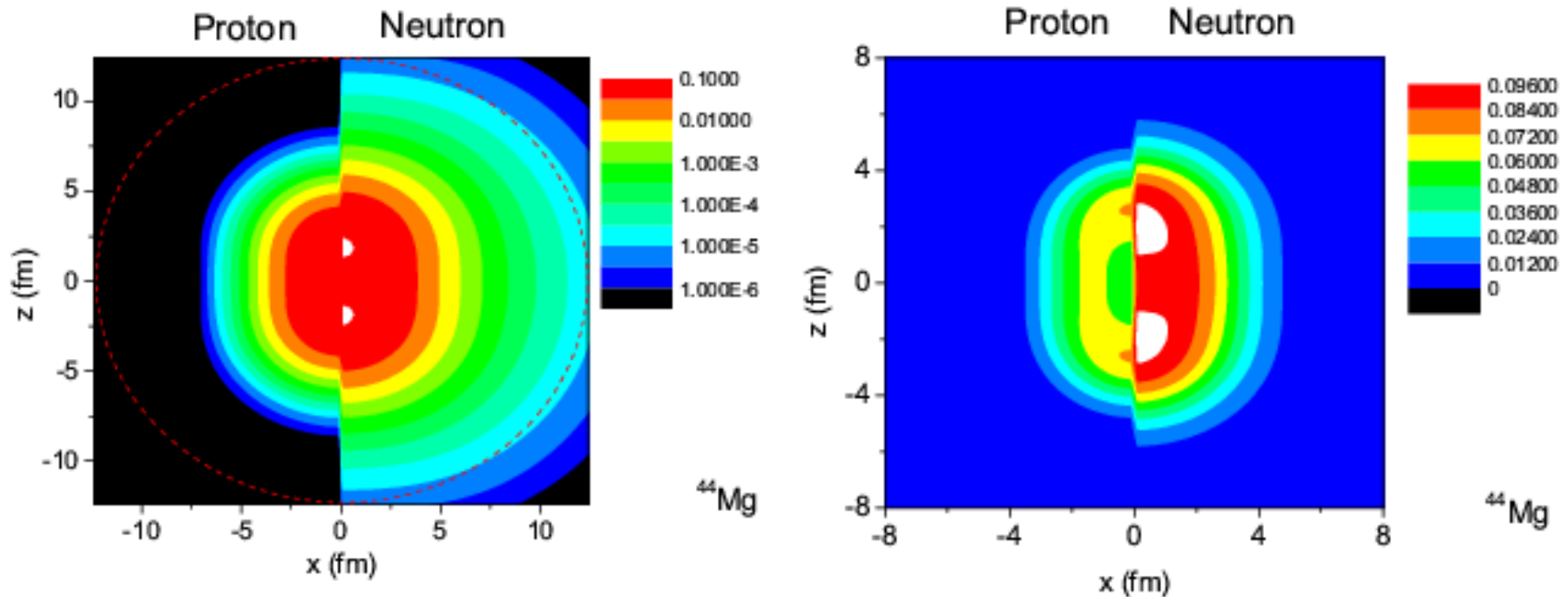
^{20}Mg : spherical from DRHBWS calculation

| Model | Pairing force | Parameters | $E_{\text{pair}}^{\text{p}}$ (MeV) |
|--------|------------------|---|------------------------------------|
| SRHBHO | Gogny | D1S | -9.2382 |
| RCHB | Surface δ | $V_0 = 374$ MeV fm ³ $\rho_0 = 0.152$ fm ³ | -9.2387 |
| | Sharp cutoff | $E_{\text{cut}}^{\text{q.p.}} = 60$ MeV | |
| DRHBWS | Surface δ | $V_0 = 380$ MeV fm ³ $\rho_0 = 0.152$ fm ³ | -9.2383 |
| | Smooth cutoff | $E_{\text{cut}}^{\text{q.p.}} = 60$ MeV $\Gamma = 5.65$ MeV | |

SGZ, Meng, Ring & Zhao, PRC82(10)011301R

^{44}Mg from DRHBWS

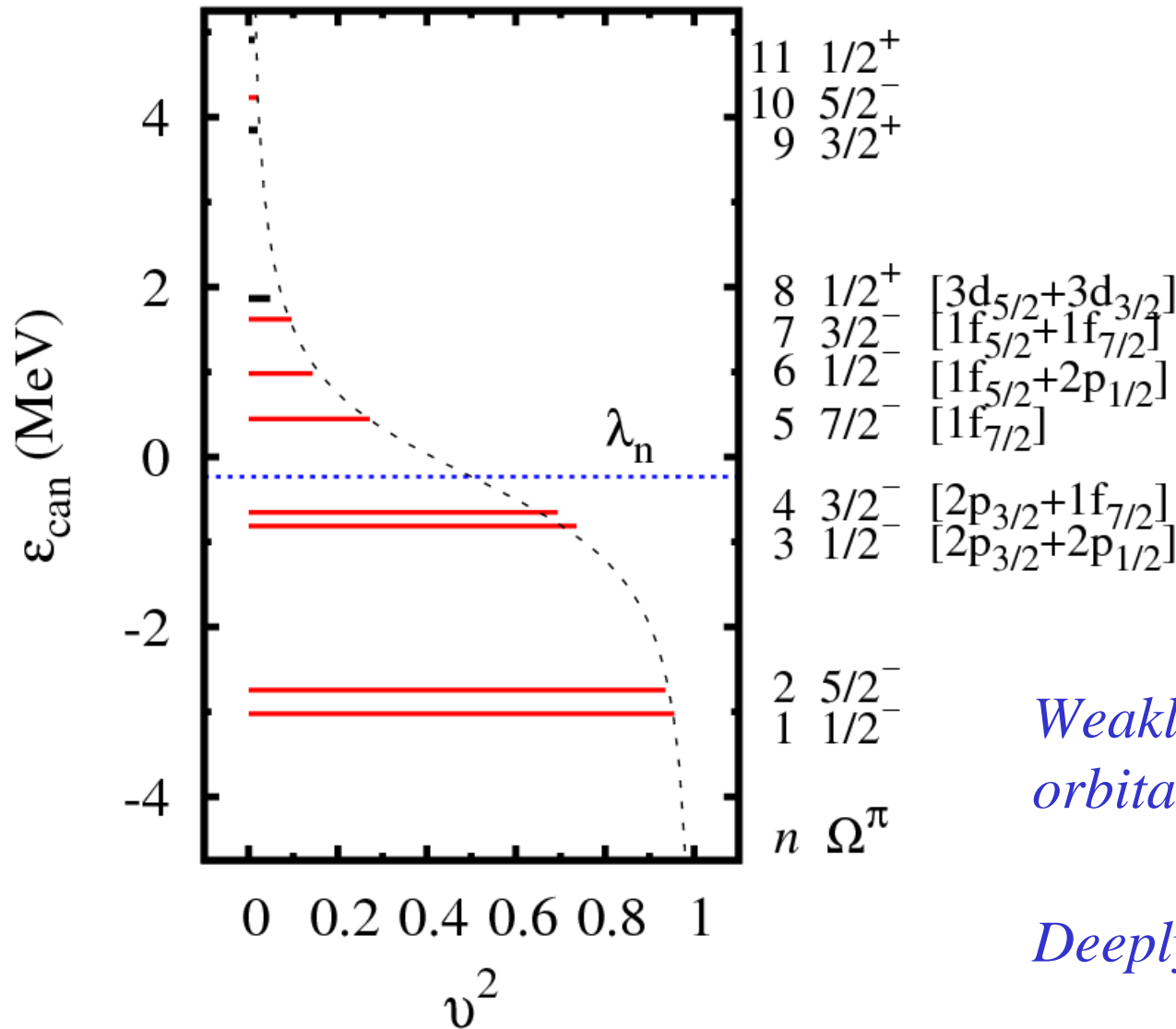
SGZ, Meng, Ring & Zhao, PRC82(10)011301R



♪ Prolate deformation

♪ Large spatial extension in neutron density distribution

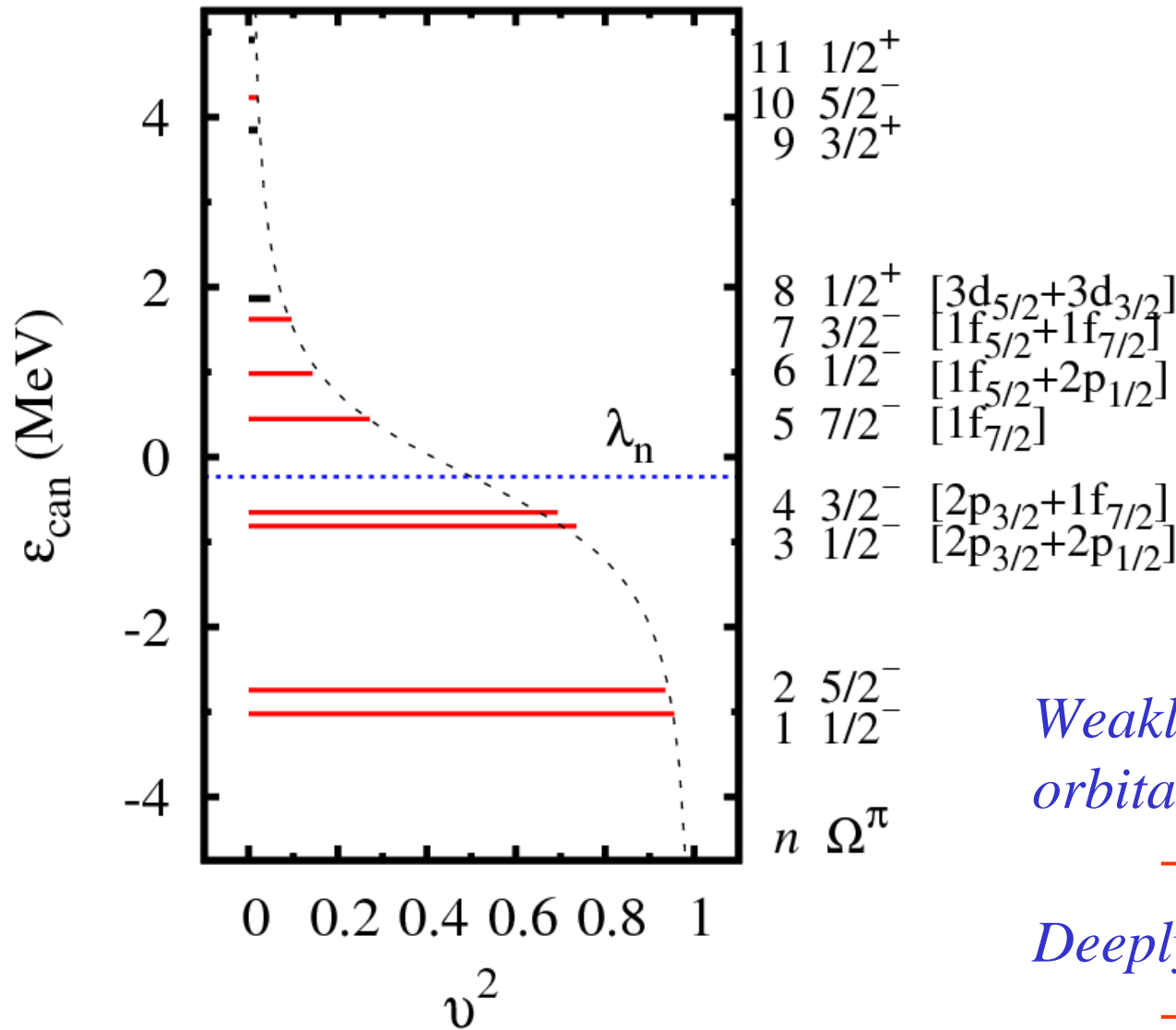
Single neutron states in canonical basis



Weakly bound & continuum orbitals

Deeply bound orbitals

Single neutron states in canonical basis



Weakly bound & continuum orbitals

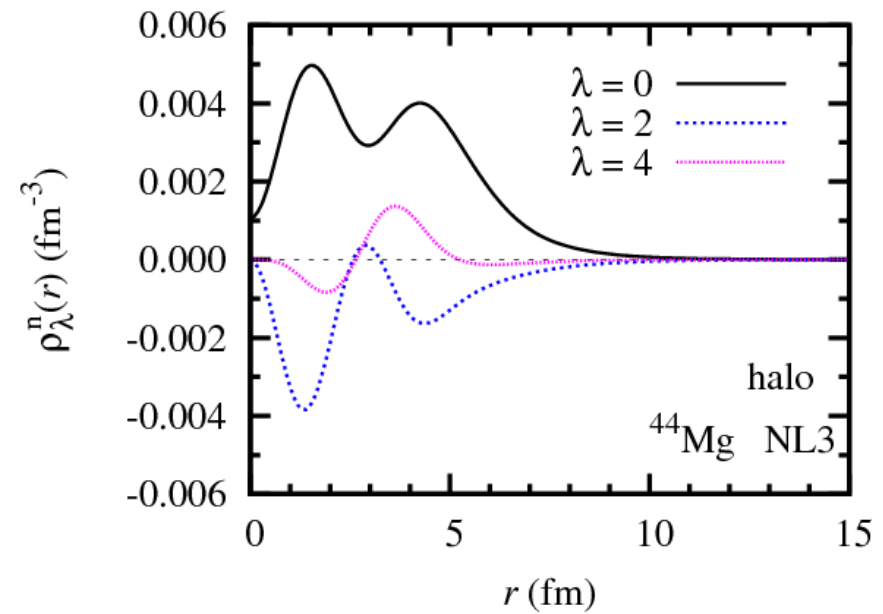
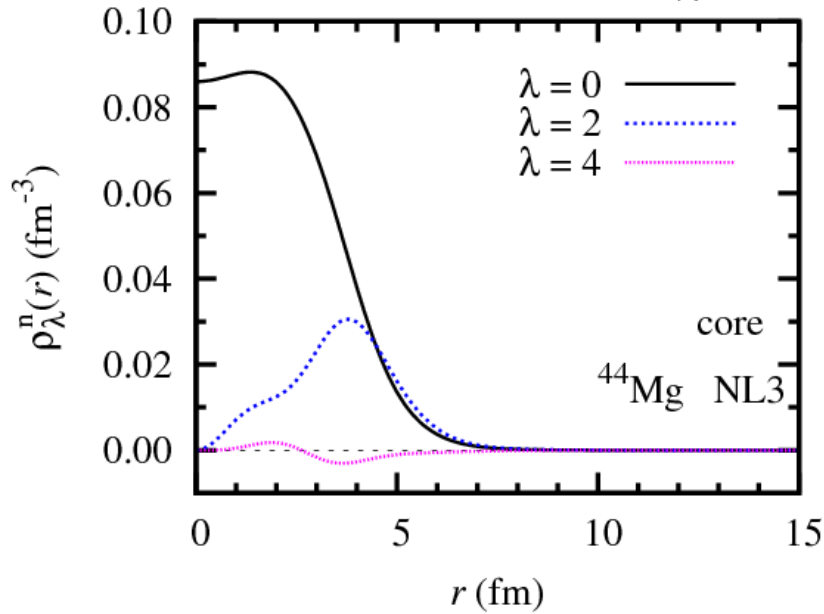
→ halo

Deeply bound orbitals

→ core

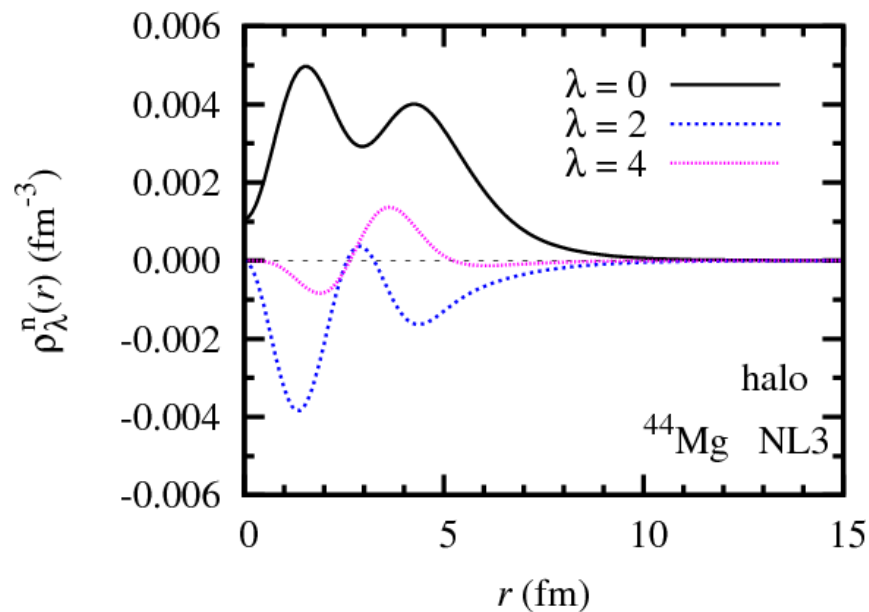
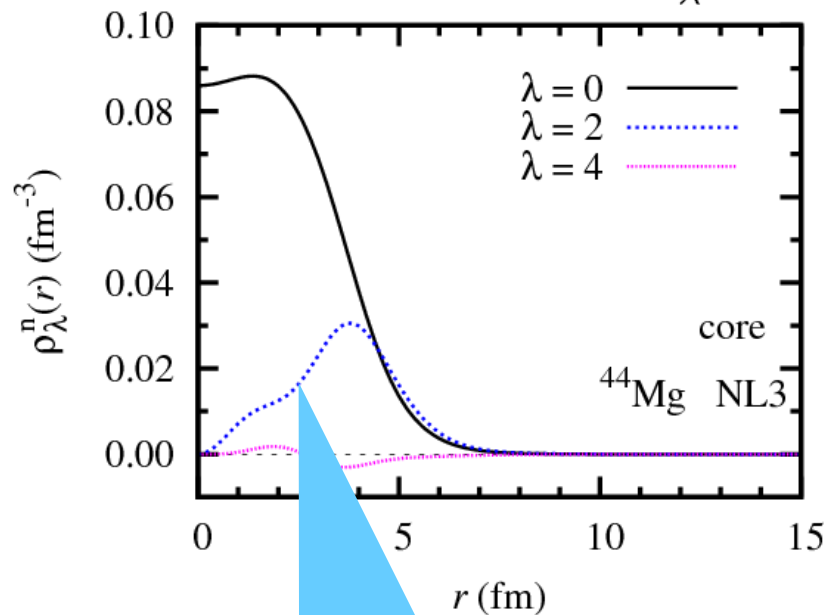
Density of core & halo: shape decoupling

$$\rho(\mathbf{r}) = \sum_{\lambda} \rho_{\lambda}(r) P_{\lambda}(\cos \theta), \quad \lambda = 0, 2, 4, \dots$$



Density of core & halo: shape decoupling

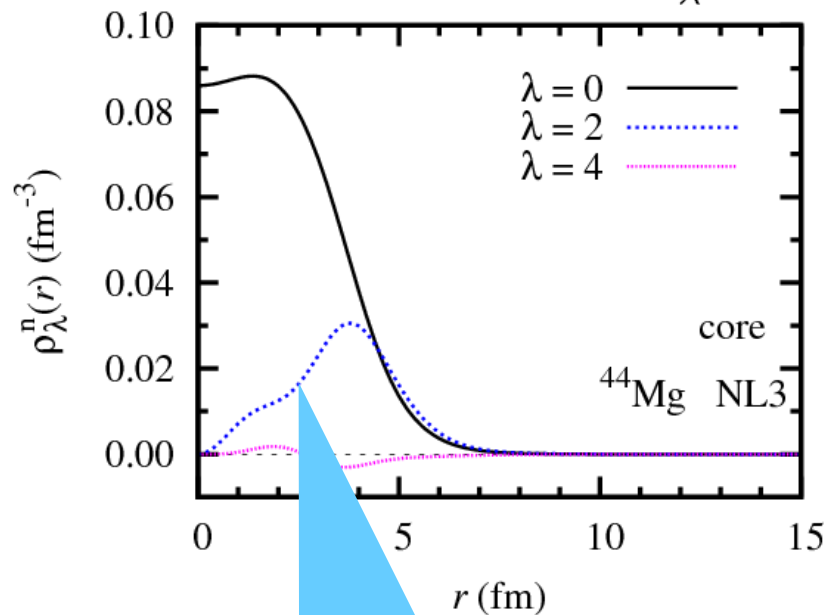
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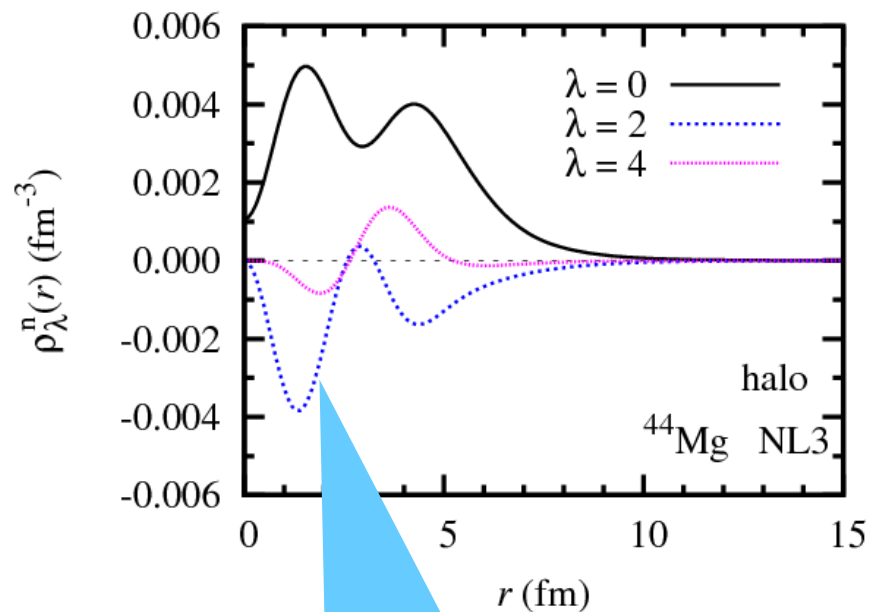
Core: prolate

Density of core & halo: shape decoupling

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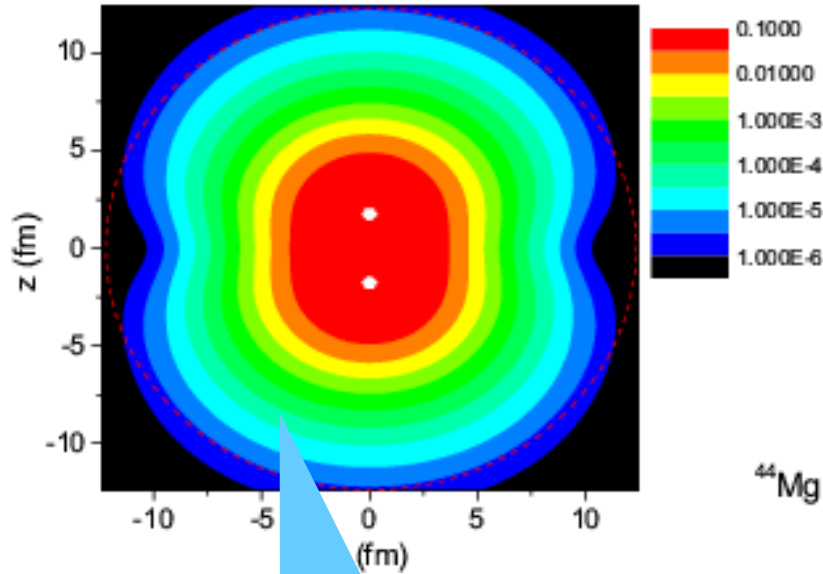


Core: prolate

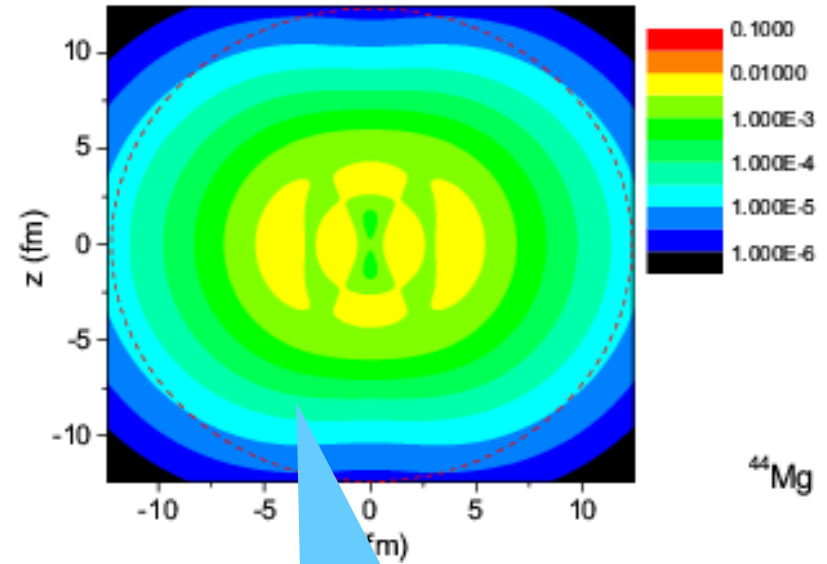


halo: oblate

Density of core & halo: shape decoupling



Core: prolate



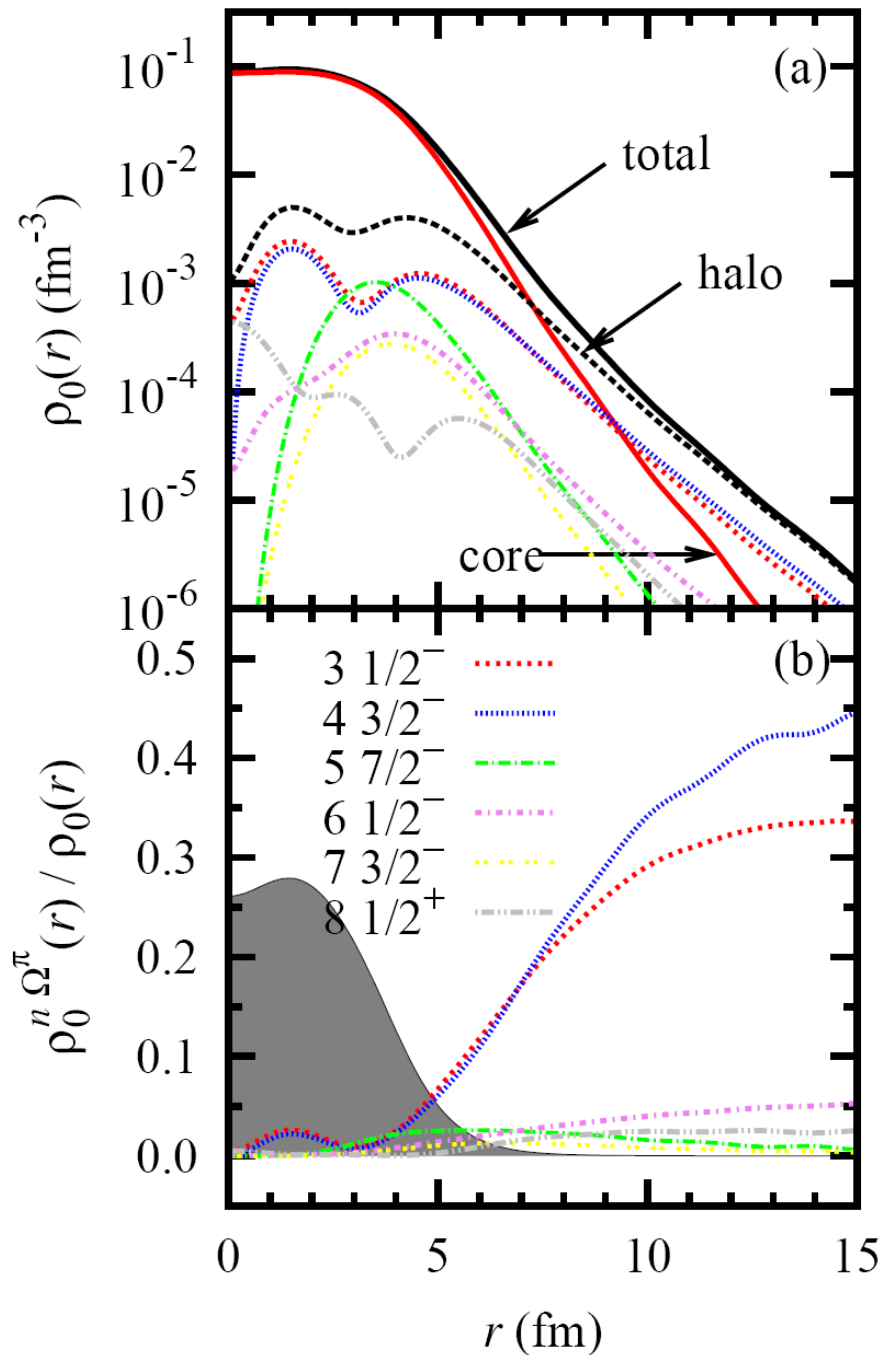
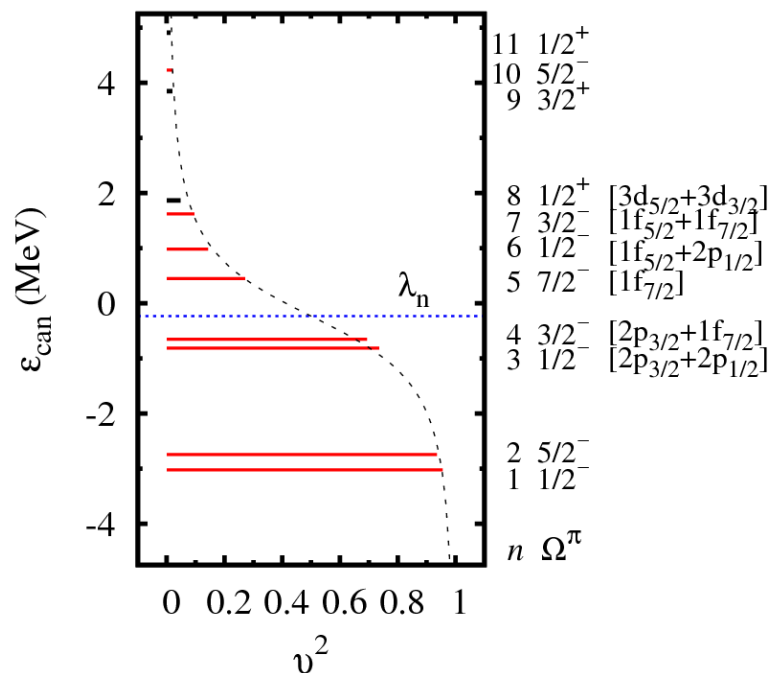
halo: oblate

Decomposition of neut. density distri.

♪ The 3rd & 4th states contribute to tail part of neutron density distribution

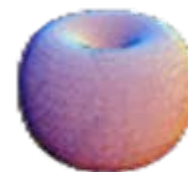
♪ Main component: $2p_{3/2}$

♪ $R_{\text{core}} = 3.72 \text{ fm}$, $R_{\text{halo}} = 5.86 \text{ fm}$



Shape of low- l single particle orbital

$$l = 1, \Lambda = \pm 1 \quad |Y_{1\pm 1}(\theta, \phi)|^2 \propto \sin^2(\theta)$$

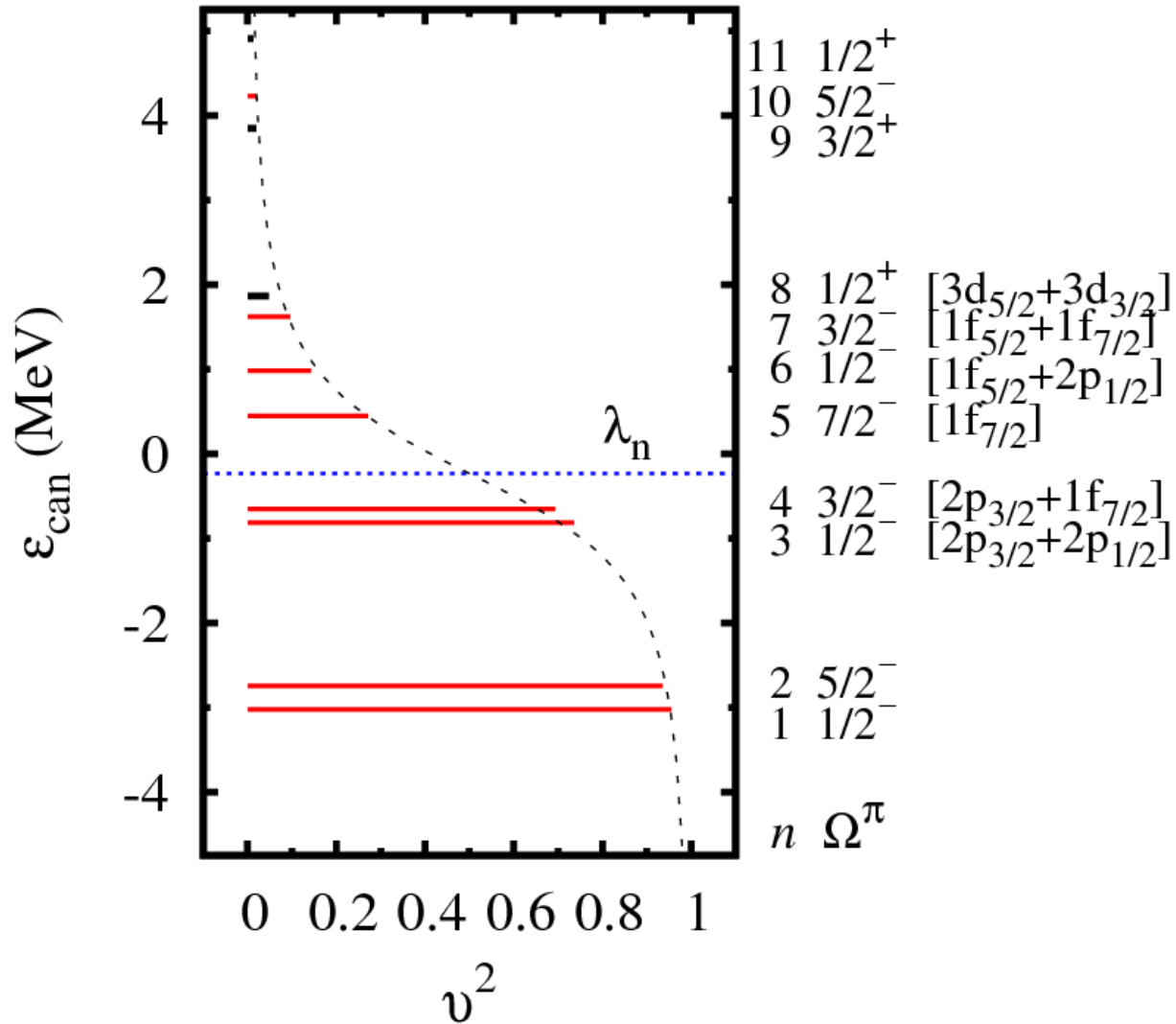


$$l = 1, \Lambda = 0 \quad |Y_{10}(\theta, \phi)|^2 \propto \cos^2(\theta)$$

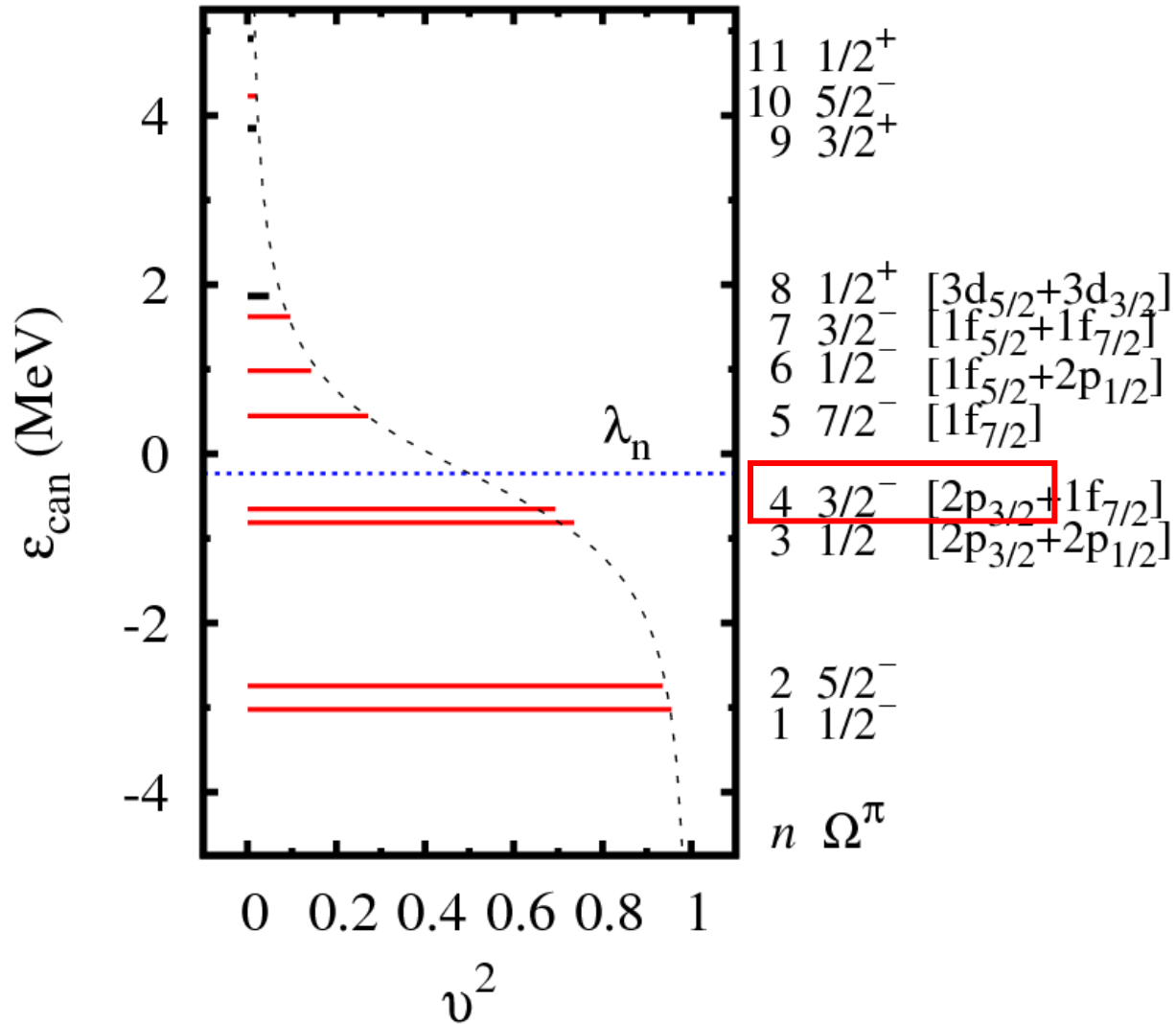


Misu, Nazarewicz, Aberg, NPA614(97)44

Mechanism of shape decoupling



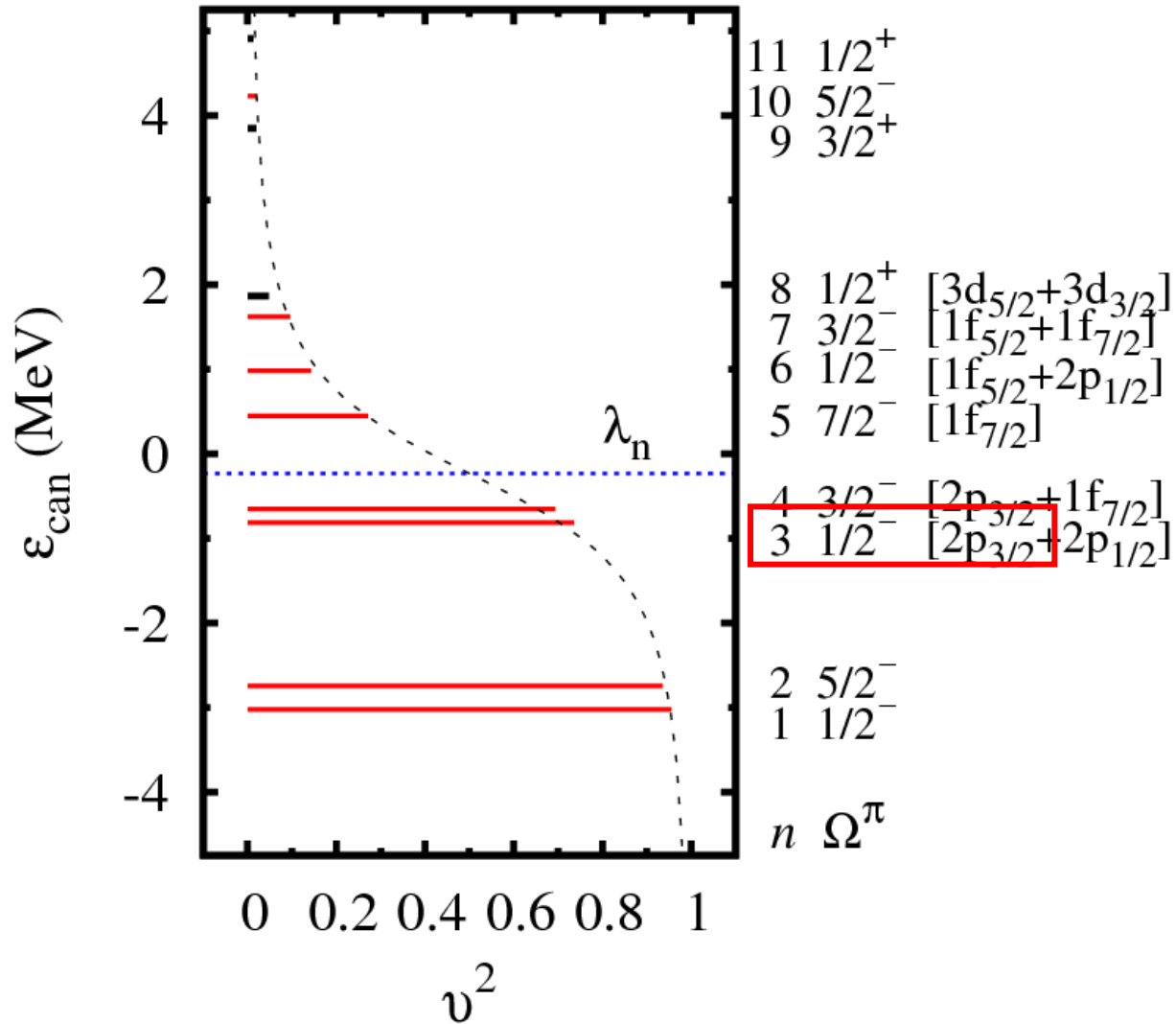
Mechanism of shape decoupling



$\Lambda = \pm 1$



Mechanism of shape decoupling



$\Lambda = \pm 1$
 $\Lambda = 0$

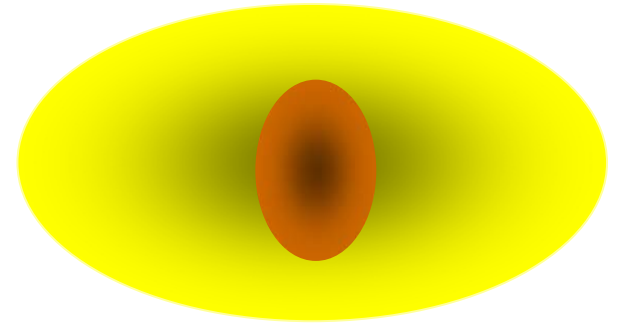


How to probe the decoupling?

♪ Larger cross section

♪ Narrower momentum distribution

✧ Double-hump ! ?



How to measure?

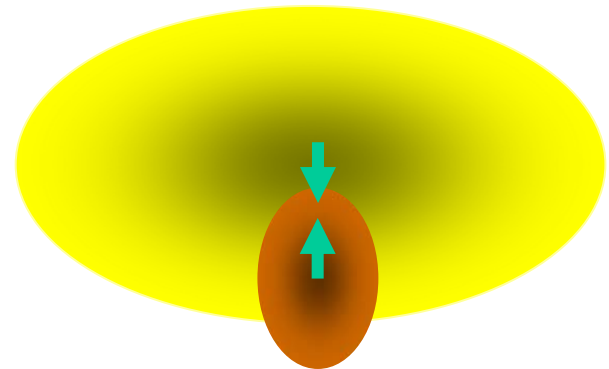
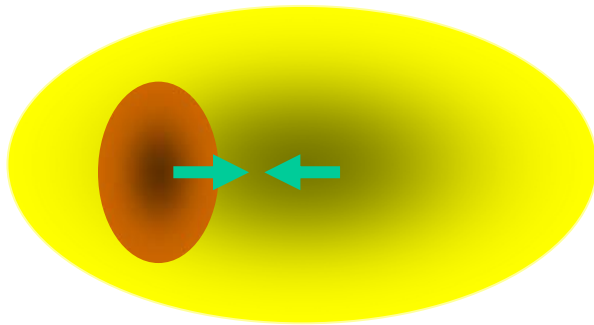
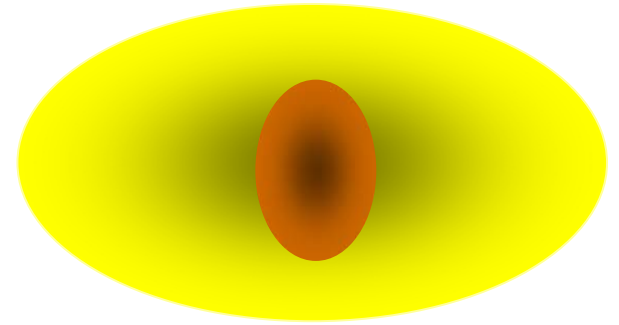
♪ Larger cross section

♪ Narrower momentum distribution

✧ Double-hump ! ?

♪ New dipole modes

♪ ...



Summary & perspectives

- ♪ Is there a halo in deformed nuclei or a deformed halo?
- ♪ **A deformed relativistic Hartree-Bogoliubov model in a Woods-Saxon basis** which describes self-consistently
 - ✧ Deformation; Large spatial distribution
 - ✧ Weakly bound & continuum
- ♪ **In deformed nuclei close to neutron drip line**
 - ✧ Halo may occur, depending on intrinsic properties of orbitals around the Fermi level
 - ✧ There might be **shape decoupling** between core and halo
- ♪ **How to observe?**
 - ✧ Momentum distribution
 - ✧ Soft dipole modes

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