

Quantum kinetic theory and its applications to chiral transports and spin polarizations

浦 实
中国科学技术大学

第十届华大QCD讲习班 Online, 2022.10.31–11.04, 华中师范大学, 武汉

Recent review:

- Y. Hidaka, SP, Q.Wang, D.L. Yang, *Foundations and Applications of Quantum Kinetic Theory*, Progress in Particle and Nuclear Physics, 127, 103989 (2022).

Outline

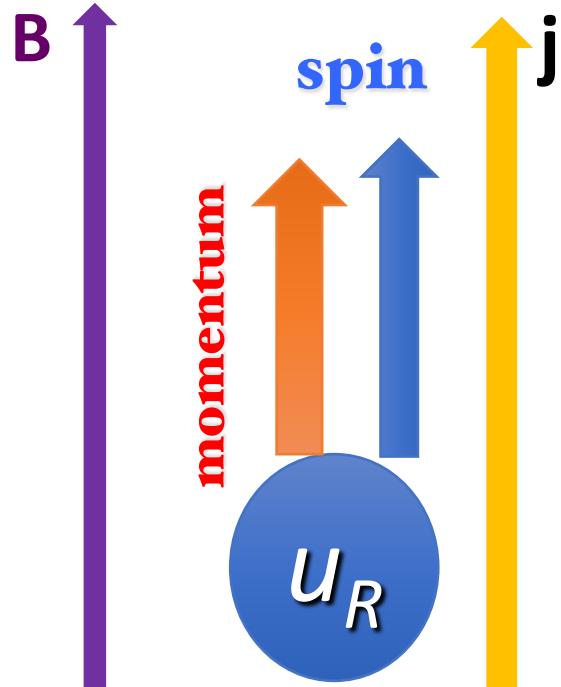
- Part 1:
Chiral magnetic effect, Berry phase and kinetic theory
- Part 2:
Wigner functions and the master equations
- Part 3:
Quantum kinetic theory in massless limit and collisions
- Part 4:
Applications to heavy ion physics

Part 2

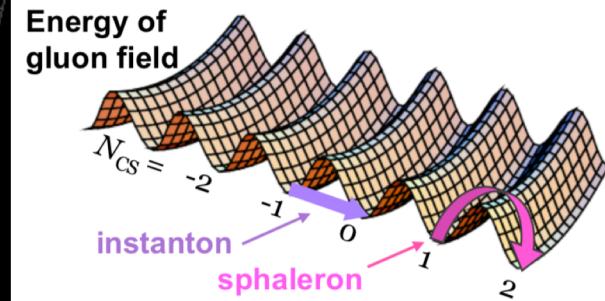
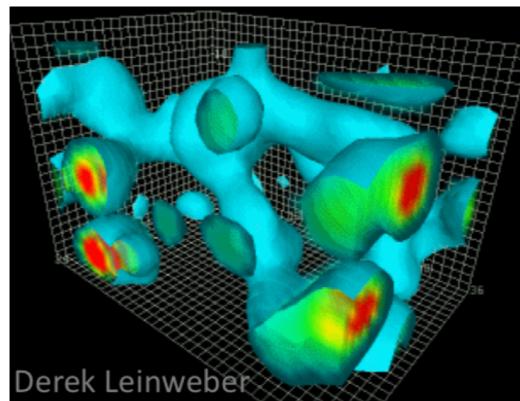
Wigner functions and the master equations

1. **Definition of gauge invariant covariant Wigner function**
 - (1a) Gauge invariant covariant Wigner function
 - (1b) Vector, chiral currents
 - (1c) The choice of gauge link
2. **Master equation for covariant Wigner function**
 - (2a) Closed-Time-Path formalism
 - (2b) From Dirac equation to master equations for chiral fermions
 - (2c) Master equations in general case (massive case)
3. **Equal-time formulism for Wigner function**

Chiral Magnetic Effect



- Magnetic fields
- Nonzero axial chemical potential
- Number of **Left** handed fermions \neq Number of **Right** handed fermions



- Charge current: charge separation

$$\mathbf{j} = \frac{e^2}{2\pi^2} \mu_5 \mathbf{B},$$

Kharzeev, Fukushima, Warrigna, (08,09), etc. ...

Chiral vortical effect Vs CME

$$\begin{aligned} j^\mu &= \xi_B B^\mu + \xi \omega^\mu, \\ j_5^\mu &= \xi_{5B} B^\mu + \xi_5 \omega^\mu, \end{aligned}$$

Son, Surowka, PRL 2009;...

$$\begin{aligned} \xi &= \frac{1}{\pi^2} \mu \mu_5, \quad \xi_B = \frac{e}{2\pi^2} \mu_5, \\ \xi_5 &= \boxed{\frac{T^2}{6}} + \frac{1}{2\pi^2} (\mu^2 + \mu_5^2), \quad \xi_{B5} = \frac{e}{2\pi^2} \mu. \end{aligned}$$

QFT: Vilenkin, PRD 22 (1980) 3080

- If we replace the Lorentz force by the Coriolis force,
Stephanov, Yin, PRL 2012

$$\mathbf{B} \rightarrow 2m\omega \sim 2|\mathbf{p}|\omega \rightarrow 2\mu\omega$$

CVE can be " derived" from CME expect $T^2/6$.

- Why have we gotten $T^2/6$?
 - Holographic models: it is the gravitational anomaly
 - Curved space QED: No...
- Will the $T^2/6$ Non-renormalizable?
No. D.F. Hou, H. Liu, H.C. Ren, PRD 2012; S. Golkar, D.T. Son, JHEP 2015

Chiral kinetic equation

$$\sqrt{G} \partial_t f + \sqrt{G} \dot{\mathbf{x}} \cdot \nabla_x f + \sqrt{G} \dot{\mathbf{p}} \cdot \nabla_p f = C[f].$$

- Particle's effective velocity:

$$\sqrt{G} \dot{\mathbf{x}} = \frac{\partial \varepsilon}{\partial \mathbf{p}} + \hbar \left(\frac{\partial \varepsilon}{\partial \mathbf{p}} \cdot \boldsymbol{\Omega} \right) \mathbf{B} + \hbar \mathbf{E} \times \boldsymbol{\Omega},$$

- Effective force:

$$\sqrt{G} \dot{\mathbf{p}} = \mathbf{E} + \frac{\partial \varepsilon}{\partial \mathbf{p}} \times \mathbf{B} + \hbar (\mathbf{E} \cdot \mathbf{B}) \boldsymbol{\Omega},$$

- Berry curvature

$$\sqrt{G} = 1 + \hbar \mathbf{B} \cdot \boldsymbol{\Omega}, \quad \boldsymbol{\Omega} = \frac{\mathbf{p}}{2|\mathbf{p}|^3},$$

Classical kinetic theory for Weyl fermions

- In a non-interacting system of fermions without collisions

$$\frac{df(t, \mathbf{x}, \mathbf{p})}{dt} = 0$$

$$\rightarrow [\sqrt{G}\partial_t + \sqrt{G}\dot{\mathbf{x}} \cdot \nabla_{\mathbf{x}} + \sqrt{G}\dot{\mathbf{p}} \cdot \nabla_{\mathbf{p}}] f(t, \mathbf{x}, \mathbf{p}) = 0.$$

$$\left\{ \nabla_{\mathbf{p}} \cdot \Omega_{\mathbf{p}} = 2\pi\delta^3(\mathbf{p}), \right\}$$

- Integral over momentum, we get

$$\partial_t \rho + \nabla \cdot \mathbf{J} = \frac{1}{2\pi} (\mathbf{E} \cdot \mathbf{B}) f(\mathbf{p} = 0)$$

$$\rho = \int \frac{d^3 \mathbf{p}}{(2\pi)^3} \sqrt{G} f(t, \mathbf{x}, \mathbf{p}),$$

$$\mathbf{J} = \int \frac{d^3 \mathbf{p}}{(2\pi)^3} \sqrt{G} \dot{\mathbf{x}} f(t, \mathbf{x}, \mathbf{p}),$$



Chiral anomaly

$$\boxed{\frac{1}{4\pi^2} (\mathbf{E} \cdot \mathbf{B})}$$

Number density for right handed fermions

Right handed fermions current

Non-trivial Lorentz symmetry

- Quantum field theory

$$j^\mu = \bar{\psi} \sigma^\mu \psi \rightarrow \Lambda_\nu^\mu j^\nu$$

- Lorentz transformation

$$x^{\mu'} = \Lambda_\nu^\mu x^\nu, \quad p^{\mu'} = \Lambda_\nu^\mu p^\nu,$$

$$\left\{ f'(x', p', t') = f(x', p', t') + \hbar N^\mu (\partial_\mu^x + F_{\nu\mu} \partial_p^\nu) f, \right\}$$

Infinitesimal
Lorentz
Transform

$$\left\{ \begin{array}{lcl} \delta x & = & \hbar \frac{\beta \times \hat{p}}{2|p|}, \\ \delta p & = & \hbar \frac{\beta \times \hat{p}}{2|p|} \times B \end{array} \right\}$$

Chen, Son, Stephanov, PRL, (2015);
Y. Hidaka, SP, D.L. Yang, PRD (2016)

1. Definition of gauge invariant covariant Wigner function

(1a) Gauge invariant covariant Wigner function

How to define “position” and “momentum”?

- In a “classical” many body system, we use distribution function $f(x,p)$.

- In Quantum mechanics, we know,

$$[x, p_x] = i\hbar.$$

- If we still want to use the distribution function $f(t,x,p)$ for a quantum many body system, how can we measure the x and p at the same time t ?

(For simplicity, we use the natural unit $\hbar = c = k_B = 1$.)

“position” and “momentum”

- Assuming we have two points, x_1^μ and x_2^μ with $\mu = 0, 1, 2, 3$.
(For simplicity, we neglect the upper index μ here)
- We introduce another two variables,

$$x = \frac{x_1 + x_2}{2}, \text{ and } y = x_1 - x_2$$

$$\partial_x = \partial_{x_1} + \partial_{x_2}$$

$$\partial_y = \frac{1}{2}(\partial_{x_1} - \partial_{x_2}) \quad \longrightarrow \quad p = -i\partial_y$$



$$[x, p] = \left[\frac{x_1 + x_2}{2}, -i\frac{1}{2}(\partial_{x_1} - \partial_{x_2}) \right] = 0$$

Wigner-Weyl transformation

$$O(x_1, x_2) = \psi^*(x_1)\psi(x_2) \rightarrow O(x, p) = \int d^3y e^{-ip \cdot y} \psi^*\left(x + \frac{y}{2}\right) \psi\left(x - \frac{y}{2}\right)$$

Definition of Wigner function in Quantum Mechanics

$$W(\mathbf{x}, \mathbf{p}) = \int d^3\mathbf{y} \exp\left(\frac{i}{\hbar}\mathbf{p} \cdot \mathbf{y}\right) \left\langle \mathbf{x} - \frac{\mathbf{y}}{2} \middle| \hat{\rho} \middle| \mathbf{x} + \frac{\mathbf{y}}{2} \right\rangle,$$

where ρ is the density matrix in QM.

$$\int d^3\mathbf{p} W(\mathbf{x}, \mathbf{p}) = |\psi(\mathbf{x})|^2$$

$$\int d^3\mathbf{x} W(\mathbf{x}, \mathbf{p}) = |\psi(\mathbf{p})|^2$$

$$\int d^3\mathbf{x} \int d^3\mathbf{p} W(\mathbf{x}, \mathbf{p}) = \text{Tr } \hat{\rho} = 1$$

Wigner function plays a role as the “probability” in phase space (\mathbf{x}, \mathbf{p}) .

Note that, Wigner function itself could be negative sometimes.

Covariant Wigner function in quantum field theory

- (Covariant) Wigner operator

$$\hat{W}_{\alpha\beta} = \int \frac{d^4y}{(2\pi)^4} e^{-ip\cdot y} \bar{\psi}_\beta(x_1) U(x_1, x_2) \psi_\alpha(x_2)$$

with gauge link

$$U(x_1, x_2) = \exp \left[-iQ \int_{x_2}^{x_1} dz \cdot A(z) \right]$$

- Wigner function:

$$W(x, p) = \langle : \hat{W}(x, p) : \rangle$$

W operator in thermal ensemble average and normal ordering of the operators.

Vasak, Gyulassy, Elze, Ann. Phys. (N.Y.) 173, 462 (1987);

Elze, Heinz, Phys. Rep. 183, 81 (1989).

1. Definition of gauge invariant covariant Wigner function

(1b) Vector, chiral currents

Decomposition of Wigner function

- For massive spinors, Wigner function is a 4x4 matrix.

$$W = \mathcal{F} + i\mathcal{P}\gamma^5 + \mathcal{V}^\mu\gamma_\mu + \mathcal{A}^\mu\gamma^5\gamma_\mu + \frac{1}{2}\mathcal{S}^{\mu\nu}\sigma_{\mu\nu},$$

$$\mathcal{F} = \text{Tr } W(x, p) \sim \langle \bar{\psi}\psi \rangle \quad \sigma_{\mu\nu} = i[\gamma_\mu, \gamma_\nu]/2$$

$$\mathcal{P} = \text{Tr } \gamma^5 W(x, p) \sim \langle \bar{\psi}\gamma^5\psi \rangle$$

$$\boxed{\mathcal{V}^\mu = \text{Tr } \gamma^\mu W(x, p) \sim \langle \bar{\psi}\gamma^\mu\psi \rangle}$$

Vector current

$$\boxed{\mathcal{A}^\mu = \text{Tr } \gamma^\mu\gamma^5 W(x, p) \sim \langle \bar{\psi}\gamma^5\gamma^\mu\psi \rangle}$$

Chiral current

(Axial vector current)

$$\boxed{\mathcal{S}^{\mu\nu} = \text{Tr } \sigma^{\mu\nu} W(x, p) \sim \langle \bar{\psi}\sigma^{\mu\nu}\psi \rangle}$$

$$W^\dagger = \gamma^0 W \gamma^0$$

All the coefficients, F, P, V, A, S are real!

$$\mathcal{V}^{\mu\dagger} = \text{Tr } W^\dagger (\gamma^\mu)^\dagger = \text{Tr } [(\gamma^0 W \gamma^0)(\gamma^0 \gamma^\mu \gamma^0)] = \mathcal{V}^\mu$$

Vector and chiral currents

$$\mathcal{V}^\mu = \text{Tr } \gamma^\mu W(x, p) = \int \frac{d^4y}{(2\pi)^4} e^{-ip \cdot y} \bar{\psi} \left(x + \frac{y}{2} \right) \gamma^\mu \exp \left[-i \int_{x-y/2}^{x+y/2} dz \cdot A(z) \right] \psi \left(x - \frac{y}{2} \right)$$

$$\mathcal{A}^\mu = \text{Tr } \gamma^\mu \gamma^5 W(x, p) = \int \frac{d^4y}{(2\pi)^4} e^{-ip \cdot y} \bar{\psi} \left(x + \frac{y}{2} \right) \gamma^\mu \gamma^5 U \exp \left[-i \int_{x-y/2}^{x+y/2} dz \cdot A(z) \right] \psi \left(x - \frac{y}{2} \right)$$

- Our definition means the vector and chiral currents are **regularized**. If we integrate over momentum,

$$j_5^\mu = \int d^4p \mathcal{A}^\mu(x, p) \sim \lim_{y \rightarrow 0} \bar{\psi} \left(x + \frac{y}{2} \right) \gamma^\mu \gamma^5 U \exp \left[-i \int_{x-y/2}^{x+y/2} dz \cdot A(z) \right] \psi \left(x - \frac{y}{2} \right)$$

which is the chiral current in position space in **Peskin's textbook QFT Sec. 19**,

$$j^{\mu 5} = \text{symm lim}_{\epsilon \rightarrow 0} \left\{ \bar{\psi}(x + \frac{\epsilon}{2}) \gamma^\mu \gamma^5 \exp \left[-ie \int_{x-\epsilon/2}^{x+\epsilon/2} dz \cdot A(z) \right] \psi(x - \frac{\epsilon}{2}) \right\}. \quad (19.22)$$

That regularization and gauge link are the key to get **chiral anomaly**!

An example: 1+1 dim chiral anomaly in operator level

$$\partial_\mu j^{\mu 5} = \text{symm lim}_{\epsilon \rightarrow 0} \left\{ \bar{\psi}(x + \frac{\epsilon}{2}) [-ie\gamma^\mu \epsilon^\nu (\partial_\mu A_\nu - \partial_\nu A_\mu)] \gamma^5 \psi(x - \frac{\epsilon}{2}) \right\}.$$

Coming from gauge link

- In 1+1 dim, we can have

$$\overline{\psi}(x + \frac{\epsilon}{2}) \Gamma \psi(x - \frac{\epsilon}{2}) = \frac{-i}{2\pi} \text{tr} \left[\frac{\gamma^\alpha \epsilon_\alpha}{\epsilon^2} \Gamma \right]$$

$$\begin{aligned} \overline{\psi}(y) \overline{\psi}(z) &= \int \frac{d^2 k}{(2\pi)^2} e^{-ik \cdot (y-z)} \frac{i k}{k^2} \\ &= -\partial \left(\frac{i}{4\pi} \log(y-z)^2 \right) \\ &= \frac{-i}{2\pi} \frac{\gamma^\alpha (y-z)_\alpha}{(y-z)^2}. \end{aligned}$$

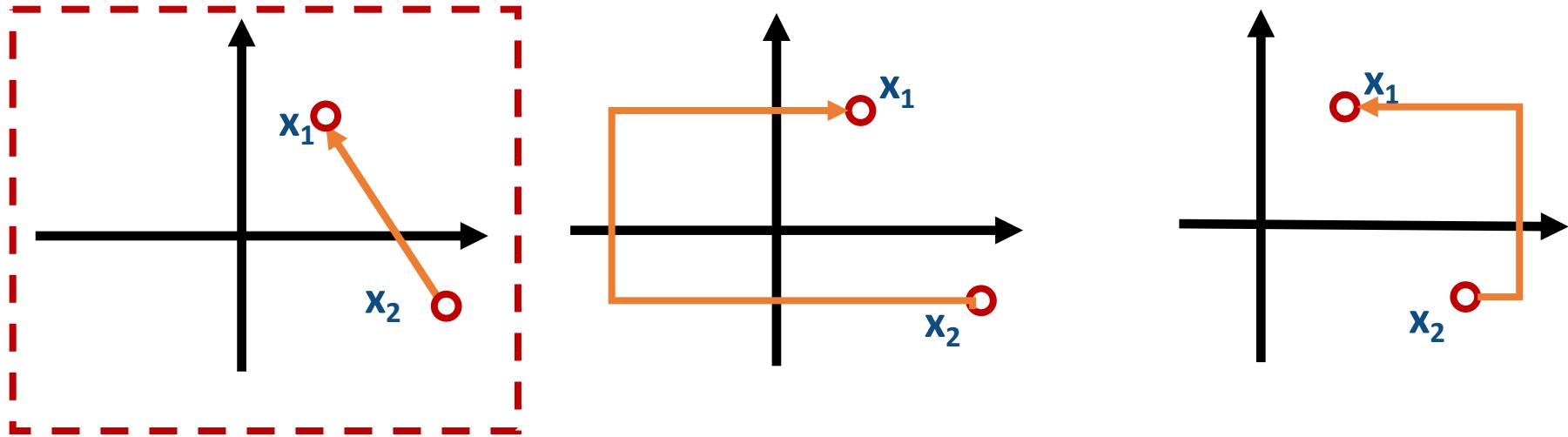
Regularization also plays a crucial role.

$$\begin{aligned} \partial_\mu j^{\mu 5} &= \text{symm lim}_{\epsilon \rightarrow 0} \left\{ \frac{-i}{2\pi} \text{tr} \left[\frac{\gamma^\alpha \epsilon_\alpha}{\epsilon^2} \gamma^\mu \gamma^5 \right] \cdot (-ie\epsilon^\nu F_{\mu\nu}) \right\}. \\ &= \frac{e}{2\pi} \text{symm lim}_{\epsilon \rightarrow 0} \left\{ 2 \frac{\epsilon^\mu \epsilon^\nu}{\epsilon^2} \right\} \epsilon^{\mu\alpha} F_{\nu\alpha} \\ &= \frac{e}{2\pi} \epsilon^{\mu\nu} F_{\mu\nu} \quad \text{chiral anomaly in 1+1 dim} \end{aligned}$$

1. Definition of gauge invariant covariant Wigner function

(1c) The choice of gauge link

The choice of gauge link



A compact form for Wigner function (I)

- For a function $A(a+x)$ with a small x , we can expand it as

$$A(a+x) = A(a) + A'(a)x + \frac{1}{2}A''(a)x^2 + \dots \frac{1}{n!} \frac{d^{(n)}A}{dx^n} x^n = e^{x \cdot \partial} A(a)$$

- Similarly, we have

$$\bar{\psi}_\beta \left(x + \frac{y}{2} \right) \psi_\alpha \left(x - \frac{y}{2} \right) \rightarrow \bar{\psi}_\beta(x) e^{+\frac{1}{2}y \cdot \overleftarrow{\partial}} \times e^{-\frac{1}{2}y \cdot \partial} \psi_\alpha(x) = \bar{\psi}_\beta(x) \exp \left[-\frac{1}{2}y \cdot (\partial - \overleftarrow{\partial}) \right] \psi_\alpha(x)$$

- Without gauge link, the Wigner function can be formally written as

$$\begin{aligned} \hat{W}_{\beta\alpha} &= \int \frac{d^4y}{(2\pi)^4} e^{-ip \cdot y} \bar{\psi}_\beta \left(x + \frac{y}{2} \right) \psi_\alpha \left(x - \frac{y}{2} \right) = \int \frac{d^4y}{(2\pi)^4} \bar{\psi}_\beta(x) \exp [iy \cdot (p - \hat{p})] \psi_\alpha(x) \\ &\sim \bar{\psi}_\beta(x) \underbrace{\delta^{(4)}(p^\mu - \hat{p}_c^\mu)}_{\text{--- --- --- ---}} \psi_\alpha(x) \end{aligned}$$

$$\hat{p}_c^\mu \equiv \frac{1}{2}i(\partial^\mu - \overleftarrow{\partial}^\mu) \quad \text{Momentum}$$

Elze, Gyulassy, Vasak, Nucl.Phys.B 276 (1980), 706

A compact form for Wigner function (II)

- With the straight line type gauge link, we find

$$\begin{aligned}\hat{W}_{\alpha\beta} &= \int \frac{d^4y}{(2\pi)^4} e^{-ip\cdot y} \bar{\psi}_\beta \left(x + \frac{y}{2} \right) \exp \left[-i \int_{x_2}^{x_1} dz \cdot A(z) \right] \psi_\alpha \left(x - \frac{y}{2} \right) \\ &\rightarrow \int \frac{d^4y}{(2\pi)^4} e^{-ip\cdot y} \bar{\psi}_\beta(x) \exp \left[-\frac{1}{2} y \cdot (D - \overleftarrow{D}) \right] \psi_\alpha(x) \\ &\sim \bar{\psi}_\beta(x) \delta^{(4)}(p^\mu - \hat{p}^\mu) \psi_\alpha(x) \quad D_\mu = \partial_\mu + iA_\mu\end{aligned}$$

$$\boxed{\hat{p}^\mu = \frac{i}{2}(D^\mu - \overleftarrow{D}) = \hat{p}_c^\mu - A^\mu}$$

Kinetic momentum

canonical momentum

The straight line type gauge link corresponds the momentum in $W(x,p)$ is the kinetic one.

Elze, Gyulassy, Vasak, Nucl.Phys.B 276 (1980), 706

2. Master equation for covariant Wigner function

(2a) Closed-Time-Path formalism

References for Closed-Time-Path formalism

PHYSICS REPORTS (Review Section of Physics Letters) 118, nos. 1 & 2 (1985) 1–131. North-Holland, Amsterdam

Review:

K.-c. Chou, Z.-b. Su, B.-l. Hao, L. Yu,
Phys. Rept. 118 (1985)

周光召、苏肇冰、郝柏林、于
渌，统一描述平衡与非平衡体系
的格林函数理论研究

- 1999, 中国科学院自然科学奖
一等奖
- 2000, 国家自然科学奖二等奖

EQUILIBRIUM AND NONEQUILIBRIUM FORMALISMS MADE UNIFIED

Kuang-chao CHOU, Zhao-bin SU,* Bai-lin HAO and Lu YU

Institute of Theoretical Physics, Academia Sinica, P.O. Box 2735, Beijing, China

Received 5 June 1984

Contents:

1. Introduction	3	5.2. General considerations concerning multi-point functions	62
1.1. Why closed time-path?	3	5.3. Plausible generalization of FDT	67
1.2. Few historical remarks	4	6. Path integral representation and symmetry breaking	70
1.3. Outline of the paper	5	6.1. Initial correlations	71
1.4. Notations	6	6.2. Order parameter and stability of state	76
2. Basic properties of CTPGF	7	6.3. Ward-Takahashi identity and Goldstone theorem	79
2.1. Two-point functions	7	6.4. Functional description of fluctuation	82
2.2. Generating functionals	12	7. Practical calculation scheme using CTPGF	89
2.3. Single time and physical representations	18	7.1. Coupled equations of order parameter and elementary excitations	90
2.4. Normalization and causality	24	7.2. Loop expansion for vertex functional	92
2.5. Lehmann spectral representation	28	7.3. Generalization of Bogoliubov-de Gennes equation	96
3. Quasiumiform systems	31	7.4. Calculation of free energy	99
3.1. The Dyson equation	31	8. Quenched random systems	103
3.2. Systems near thermoequilibrium	34	8.1. Dynamic formulation	104
3.3. Transport equation	39	8.2. Infinite-ranged Ising spin glass	109
3.4. Multi-time-scale perturbation	44	8.3. Disordered electron system	114
3.5. Time dependent Ginzburg-Landau equation	46	9. Connection with other formalisms	119
4. Time reversal symmetry and nonequilibrium stationary state (NESS)	48	9.1. Imaginary versus real time technique	119
4.1. Time inversion and stationarity	49	9.2. Quantum versus fluctuation field theory	123
4.2. Potential condition and generalized FDT	52	9.3. A plausible microscopic derivation of MSR field theory	125
4.3. Generalized Onsager reciprocity relations	53	10. Concluding remarks	127
4.4. Symmetry decomposition of the inverse relaxation matrix	55	Note added in proof	128
5. Theory of nonlinear response	58	References	128
5.1. General expressions for nonlinear response	58		

Schrodinger, Heisenberg's and interaction picture

- Schrodinger picture:

States are time dependent, operators do not depend on t.

- Heisenberg's picture:

States do not depend on t, operators are time dependent.

$$\hat{O}(t) = e^{iH(t-t_0)} \hat{O}(t_0) e^{-iH(t-t_0)}$$

- Interaction picture: $H = H_0 + H_{\text{int}}$

$$\hat{O}_I(t) = e^{iH_0(t-t_0)} \hat{O}(t_0) e^{-iH_0(t-t_0)}$$

$$\begin{aligned}\hat{O}(t) &= e^{iH(t-t_0)} e^{-iH_0(t-t_0)} \hat{O}_I(t) e^{iH_0(t-t_0)} e^{-iH(t-t_0)} \\ &\equiv U^\dagger(t, t_0) \hat{O}_I(t) U(t, t_0)\end{aligned}$$

$$U(t, t_0) = e^{iH_0(t-t_0)} e^{-iH(t-t_0)}$$

Interaction picture for perturbative QFT

- Let us consider the scalar field. Note that, scalar field is operator.

$$\phi(t, \mathbf{x}) = U^\dagger(t, t_0) \phi_I(t, \mathbf{x}) U(t, t_0)$$

- By solving the eq.

$$i\partial_t U(t, t_0) = H_I(t) U(t, t_0)$$

One can get

$$U(t, t_0) = T \left\{ \exp \left[-i \int_{t_0}^t dt' H_I[\phi_I(t')] \right] \right\}$$

T: time ordering

$$U(t_1, t_2) U(t_2, t_3) = U(t_1, t_3);$$

$$U(t_1, t_3) [U(t_2, t_3)]^\dagger = U(t_1, t_2).$$

Green function in zero temperature

- The two point function in zero temperature is defined as

$$G(x_1, x_2) = \langle \Omega | T\phi(x_1)\phi(x_2) | \Omega \rangle$$

T is time ordering and $|\Omega\rangle$ is the ground state of H.

- Usually, we express the scalar fields and ground state by them in free cases.

$$\phi(t, \mathbf{x}) = U^\dagger(t, t_0)\phi_I(t, \mathbf{x})U(t, t_0)$$

$$|\Omega\rangle \propto \lim_{\mathcal{T} \rightarrow \infty(1-i\epsilon)} e^{-iH\mathcal{T}} |0\rangle \propto \lim_{\mathcal{T} \rightarrow \infty(1-i\epsilon)} U(t_0, -\mathcal{T}) |0\rangle$$

$$G(x_1, x_2) = \lim_{\mathcal{T} \rightarrow \infty(1-i\epsilon)} \langle 0 | U^\dagger(-\mathcal{T}, \mathcal{T}) [T\phi_I(x_1)\phi_I(x_2)U(-\mathcal{T}, \mathcal{T})] | 0 \rangle$$

In zero temperature case, $U^+(t, t_0) = U^{-1}(t, t_0)$ is the quantum fluctuations.
Note that $U|0\rangle$ only gives a trivial phase factor. As a consequence, we do not need to consider the disconnected Feynman diagrams.

Green correlation function at finite temperature (I)

- In non-equilibrium quantum field theory, the two-point correlation function is defined through the density operator at an equilibrium state (at time t),

$$G(x_1, x_2) = \langle T\phi(x_1)\phi(x_2) \rangle = \frac{1}{\text{Tr}\hat{\rho}(t)} \text{Tr}[\hat{\rho}(t)T\phi(x_1)\phi(x_2)]$$

$\phi(x)$: field operator in Heisenberg's picture

$\rho(t)$: density operator at the initial time t

T : (standard) time ordering

$$\hat{\rho}(t) = \exp[\beta(\hat{H} - \mu\hat{N})] \Big|_{\beta=-it}$$

- We can use the similar interaction picture for $G(x_1, x_2)$. But, note that, we also need to express the density matrix $\rho(t)$ in interaction picture.

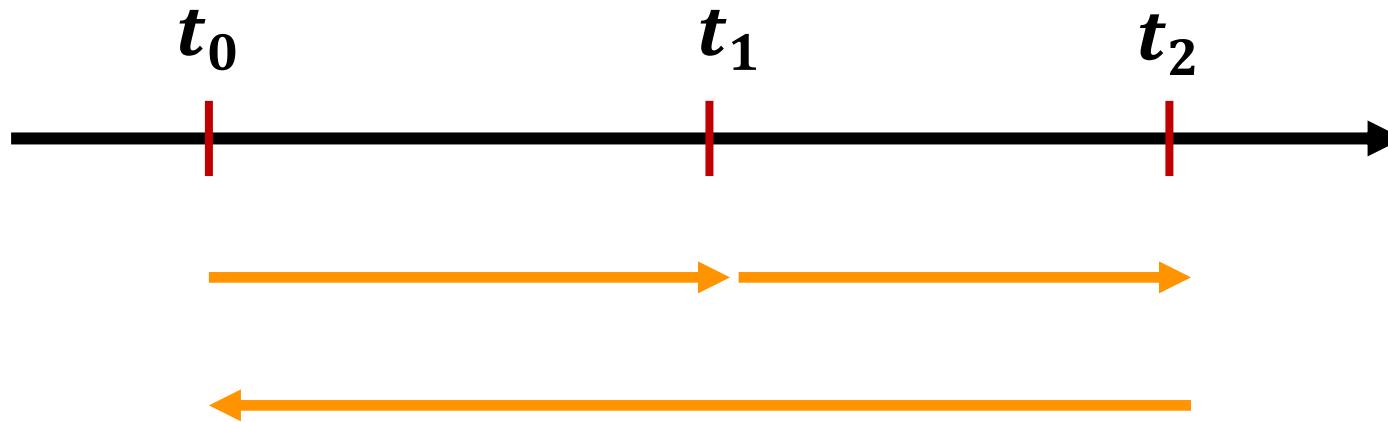
J. Schwinger, J.Math. Phys. 2(1961) 407

Keldysh, J. Exptl. Theoret. Phys. 47, (1964) 1515-1527

Green correlation function at finite temperature (II)

- We assume that the interaction H_I start from t_0 .

$$G(x_1, x_2) = \frac{1}{\text{Tr} \hat{\rho}(t_0)} \text{Tr} \{ \hat{\rho}(t_0) U(t_0, t) T[U(t, t_1) \phi_I(x_1) U(t_1, t_2) \phi_I(x_2) U(t_2, t)] U(t, t_0) \}$$

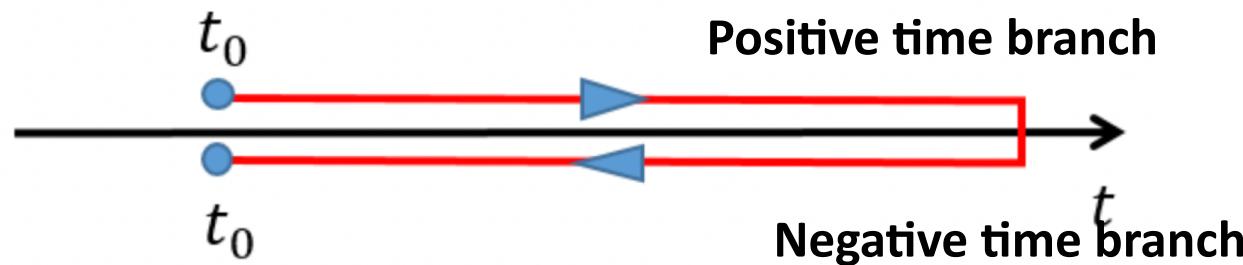


- Note that the $U(t_0, t)$ and $U(t, t_0)$ are not no longer trivial phase factors. It is difficult to handle them.

Schwinger-Keldysh contour

$$G(x_1, x_2) = \frac{1}{\text{Tr} \hat{\rho}(t_0)} \text{Tr} \{ \hat{\rho}(t_0) T[\phi_I(x_1) \phi_I(x_2) U_{CTP}(t_0)] \}$$

- **Schwinger-Keldysh contour**



$$\begin{aligned} U_{CTP}(t_0) &\equiv T_P \left[\exp \left(-i \int_{CTP} dt H_I(t) \right) \right] \\ &= T_P \left[\exp \left(-i \int_{t_0}^{\infty} dt_+ H_I(t_+) + i \int_{t_0}^{\infty} dt_- H_I(t_-) \right) \right] \end{aligned}$$

J. Schwinger, J.Math. Phys.2(1961) 407

Keldysh, J. Exptl. Theoret. Phys. 47, (1964) 1515-1527

Calzetta, Hu, PRD, 37 (1988), 10

Jordan, PRD, 33 (1986), 15

Two point Green functions

$$G(x_1, x_2) = \begin{pmatrix} G^{++}(x_1, x_2) & G^{+-}(x_1, x_2) \\ G^{-+}(x_1, x_2) & G^{--}(x_1, x_2) \end{pmatrix} = \begin{pmatrix} G^F(x_1, x_2) & \pm G^<(x_1, x_2) \\ G^>(x_1, x_2) & G^{\bar{F}}(x_1, x_2) \end{pmatrix}$$

$G^{ij}(x_1, x_2)$:

i=+,- : first time argument x_1^0 lives on the time branch i +: positive time branch;
j=+,- : second time argument x_2^0 lives on the time branch j -: negative time branch

- For spin-1/2 fermions, the definitions of the various two-point correlation functions are

$$G_{\alpha\beta}^F(x_1, x_2) = \langle T\psi_\alpha(x_1)\bar{\psi}_\beta(x_2) \rangle, \quad G^F + G^{\bar{F}} = G^> - G^<,$$

$$G_{\alpha\beta}^{\bar{F}}(x_1, x_2) = \langle T_A\psi_\alpha(x_1)\bar{\psi}_\beta(x_2) \rangle, \quad -iG^R = G^F + G^< = -G^{\bar{F}} + G^>,$$

$$G_{\alpha\beta}^<(x_1, x_2) = \langle \bar{\psi}_\beta(x_2)\psi_\alpha(x_1) \rangle, \quad -iG^A = G^F - G^> = -G^{\bar{F}} - G^<,$$

$$G_{\alpha\beta}^>(x_1, x_2) = \langle \psi_\alpha(x_1)\bar{\psi}_\beta(x_2) \rangle,$$

R: Retarded
A: Advanced

2. Master equation for covariant Wigner function

(2b) From Dirac equation to master equations for chiral fermions

Dirac equation for chiral fermions

- Let us start from the Lagrangian for massless fermions in a background electromagnetic field,

$$\mathcal{L} = \bar{\psi} i\gamma \cdot D\psi = \chi_R^\dagger i\sigma \cdot D\chi_R + \chi_L^\dagger i\bar{\sigma} \cdot D\chi_L,$$

$$\psi = (\chi_L, \chi_R)^T \quad \text{L, R: left or right handed Pauli spinors}$$

$$\sigma^\mu = (1, \boldsymbol{\sigma}), \bar{\sigma}^\mu = (1, -\boldsymbol{\sigma}), \text{ and } D_\mu = \hbar\partial_\mu + iA_\mu$$

- We concentrate on right handed fermions and suppress the subscript R

$$i\sigma \cdot D\chi_R = 0, \quad -\chi_R^\dagger i\sigma \cdot \overleftarrow{D}^\dagger = 0,$$

Blaizot, Iancu, Phys. Rept. 359 (2002) 355–528

Equation for $S^<$

- We define $S^>$ and $S^<$ for right-handed fermions

$$\begin{aligned} S_{ab}^>(x_1, x_2) &= \left\langle \chi_a(x_1) \chi_b^\dagger(x_2) \right\rangle, \\ S_{ab}^<(x_1, x_2) &= \left\langle \chi_b^\dagger(x_2) \chi_a(x_1) \right\rangle, \end{aligned}$$

$S = S^<.$

$$\begin{array}{ccc} i\sigma \cdot D \chi_R = 0, & \text{orange arrow} & i\sigma \cdot D_{x_1} S(x_1, x_2) = 0, \\ -\chi_R^\dagger i\sigma \cdot \overleftarrow{D}^\dagger = 0, & & -S(x_1, x_2) i\sigma \cdot \overleftarrow{D}_{x_2}^\dagger = 0, \end{array}$$

Gauge invariant S

- We can define the **gauge invariant** two-point function by including gauge links,

$$\tilde{S}(x, y) = U(x, x_1) S(x_1, x_2) U(x_2, x), \quad x = \frac{1}{2}(x_1 + x_2),$$

$$U(x_1, x_2) = \mathcal{P} \exp \left[-i \frac{1}{\hbar} \int_{x_2}^{x_1} dz \cdot A(z) \right], \quad y = x_1 - x_2,$$

- If electromagnetic fields are classical background ones, $U(x_1, x_2)$ is just a phase factor

$$\tilde{S}(x_1, x_2) = S(x_1, x_2) U(x_2, x_1) = U(x_2, x_1) S(x_1, x_2).$$

→ $S(x_1, x_2) = U(x_2, x_1)^\dagger \tilde{S}(x_1, x_2) = U(x_1, x_2) \tilde{S}(x_1, x_2)$

Equation for gauge invariant S (I)

$$\begin{aligned} i\sigma \cdot D_{x_1} S(x_1, x_2) &= i\sigma \cdot [\hbar\partial_{x_1} + ieA(x_1)]S(x_1, x_2) \\ &= i\sigma \cdot D_{x_1} U(x_1, x_2) \tilde{S}(x_1, x_2) \\ &= i\sigma \cdot [D_{x_1} U(x_1, x_2)] \tilde{S}(x_1, x_2) + U(x_1, x_2) i\hbar\sigma \cdot \partial_{x_1} \tilde{S}(x_1, x_2), \end{aligned}$$

$$\begin{aligned} D_{x_1, \mu} U(x_1, x_2) &= -U(x_1, x_2) i(x_1^\nu - x_2^\nu) \int_0^1 ds s F_{\mu\nu}[z(s)] \\ &= -U(x_1, x_2) iy^\nu \int_0^1 ds s F_{\mu\nu}[z(s)], \quad \text{See next page for Details.} \end{aligned}$$

$$z(s) = x + (s - 1/2)y \text{ with } z(0) = x_2 \text{ and } z(1) = x_1.$$

- Also see Eq. (3.16) of Elze, Gyulassy, Vasak, Nucl. Phys. B276 (1986) 706

Get rid of $\partial_{x_1} U(x_1, x_2)$

$$\int_{x_2}^{x_1} dz \cdot A(z) = y^\nu \int_0^1 ds A_\nu \left(x + (s - \frac{1}{2})y \right)$$

$$\begin{aligned} \partial_\mu^{x_1} \int_{x_2}^{x_1} dz \cdot A(z) &= \left(\frac{1}{2} \partial_\mu^x + \partial_y \right) \left[y^\nu \int_0^1 ds A_\nu \left(x + (s - \frac{1}{2})y \right) \right] \\ &= \int_0^1 ds A_\mu(z(s)) + \frac{1}{2} y^\nu \int_0^1 ds \partial_\mu^x A_\nu(z(s)) + y^\nu \int_0^1 ds \partial_\mu^y A_\nu(z(s)) \\ &= \int_0^1 ds A_\mu(z(s)) + y^\nu \int_0^1 ds s [\partial_\mu A_\nu(z(s))] \\ &= \int_0^1 ds A_\mu(z(s)) + y^\nu \int_0^1 ds s [F_{\mu\nu}(z(s)) + \partial_\nu A_\mu(z(s))] \\ &= \int_0^1 ds A_\mu(z(s)) + y^\nu \int_0^1 ds s [F_{\mu\nu}(z(s))] + \int_0^1 ds s \frac{d}{ds} A_\mu(z(s)) \\ &= y^\nu \int_0^1 ds s [F_{\mu\nu}(z(s))] + A_\mu(x_1) \end{aligned}$$

$$y^\nu \partial_\nu^z A_\mu(z(s)) = (s - \frac{1}{2}) y^\nu \partial_\nu^y A_\mu(z) = \frac{d}{ds} A_\mu(z(s))$$

$$\frac{d}{ds} A_\mu(z(s)) = \frac{\partial z^\alpha}{\partial s} \partial_\alpha^z A_\mu = y^\alpha \partial_\alpha^z A_\mu$$

Equation for gauge invariant S (II)

$$\sigma^\mu \left\{ \frac{1}{2} i\hbar \frac{\partial}{\partial x^\mu} + i\hbar \frac{\partial}{\partial y^\mu} + y^\nu \int_0^1 ds s F_{\mu\nu}[z(s)] \right\} \tilde{S}(x, y) = 0.$$

- We can expand $F_{\mu\nu}$ in powers of y ,

$$\begin{aligned} \int_0^1 ds s F_{\mu\nu}[z(s)] &= \sum_{n=0} \frac{1}{n!} \int_0^1 ds s \left(s - \frac{1}{2} \right)^n (y \cdot \partial_x)^n F_{\mu\nu}(x) \\ &= \sum_{n=0} \frac{[(-1)^n + 3 + 2n]}{4(n+2)!} \left(\frac{1}{2} y \cdot \partial_x \right)^n F_{\mu\nu}(x). \end{aligned}$$

Wigner transformation to gauge invariant S

- Wigner transformation and its inverse transformation

$$\tilde{S}(x, p) = \int d^4y \exp\left(\frac{i}{\hbar} p \cdot y\right) \tilde{S}(x, y),$$

$$\tilde{S}(x, y) = \int \frac{d^4p}{(2\pi\hbar)^4} \exp\left(-\frac{i}{\hbar} p \cdot y\right) \tilde{S}(x, p).$$

For simplicity, we define $| S(x, p) \equiv S^<(x, p) \equiv \tilde{S}(x, p).$

- Then, we get

$$\left\{ \begin{array}{l} \sigma \cdot \left(\frac{1}{2} i\hbar \nabla + \Pi \right) S(x, p) = 0. \\ \left(-\frac{1}{2} i\hbar \nabla + \Pi \right) S(x, p) \cdot \sigma = 0. \end{array} \right.$$

$$\begin{aligned} \nabla_\mu &= \partial_\mu^x - j_0(\Delta) F_{\mu\nu}(x) \partial_p^\nu, \\ \Pi_\mu &= p_\mu - \frac{1}{2} \hbar j_1(\Delta) F_{\mu\nu}(x) \partial_p^\nu, \\ \Delta &= (1/2) \hbar \partial_x \cdot \partial_p \end{aligned}$$

$$j_0(z) = \sin z / z \text{ and } j_1(z) = (\sin z - z \cos z) / z^2$$

Constant EM fields limits

$$\begin{aligned}\nabla^\mu &= \sum_{n=0}^{\infty} \hbar^{2n} \nabla_{(2n)}^\mu, & \nabla_\mu^{(0)} &= \partial_\mu^x - F_{\mu\nu} \partial_p^\nu, \quad \Pi_\mu^{(0)} = p_\mu. \\ \Pi^\mu &= \sum_{n=0}^{\infty} \hbar^{2n} \Pi_{(2n)}^\mu. & \nabla_{(2n)}^\mu &= (-1)^{n+1} \frac{1}{(2n+1)!} \left(\frac{1}{2} \hbar \partial_x \cdot \partial_p \right)^{2n} F_{\mu\nu}(x) \partial_p^\nu, \\ && \Pi_{(2n)}^\mu &= (-1)^n \frac{n}{(2n+1)!} \left(\frac{1}{2} \hbar \partial_x \cdot \partial_p \right)^{2n-1} F_{\mu\nu}(x) \partial_p^\nu.\end{aligned}$$

- In constant background fields, we find

$$\begin{aligned}\sigma^\mu \left(\frac{1}{2} i \hbar \nabla_\mu^{(0)} + p_\mu \right) S(x, p) &= 0, \\ \left(-\frac{1}{2} i \hbar \nabla_\mu^{(0)} + p_\mu \right) S(x, p) \sigma^\mu &= 0.\end{aligned}$$

Decomposition of S

Since $S(x,p)$ is a 2×2 matrix, we can use the $\mathbf{1}$ and Pauli matrix to decompose it,

$$\left\{ \begin{array}{l} S_R(x,p) = \bar{\sigma}^\mu \mathcal{J}_\mu^+, \\ S_L(x,p) = \sigma^\mu \mathcal{J}_\mu^-, \end{array} \right\} \text{Tr } (\sigma^\mu \bar{\sigma}^\nu) = 2\eta^{\mu\nu} \xrightarrow{\text{orange arrow}} \left\{ \begin{array}{l} \mathcal{J}_\mu^+ = \frac{1}{2} \text{Tr } (\sigma_\mu S_R), \\ \mathcal{J}_\mu^- = \frac{1}{2} \text{Tr } (\bar{\sigma}_\mu S_L), \end{array} \right\}$$

$\sigma^\mu = (1, \boldsymbol{\sigma}), \bar{\sigma}^\mu = (1, -\boldsymbol{\sigma}),$ \pm represent right/left-handed components of Wigner functions

which are connected to the V and A in Wigner functions

$$\left\{ \quad \mathcal{V}_\mu = \mathcal{J}_\mu^+ + \mathcal{J}_\mu^-, \quad \mathcal{A}_\mu = \mathcal{J}_\mu^+ - \mathcal{J}_\mu^-. \quad \right\}$$

Master eq. for massless Wigner function

By using identities

$$\begin{aligned}\sigma^\mu \bar{\sigma}^\nu &= \eta^{\mu\nu} - \frac{1}{2}i\epsilon^{\mu\nu\lambda\rho}\sigma_\lambda \bar{\sigma}_\rho, \\ \bar{\sigma}^\nu \sigma^\mu &= \eta^{\mu\nu} + \frac{1}{2}i\epsilon^{\mu\nu\lambda\rho}\sigma_\lambda \bar{\sigma}_\rho,\end{aligned}$$

and insert the decomposition of S into the eqs.

$$\begin{aligned}\sigma^\mu \left(\frac{1}{2}i\hbar\nabla_\mu^{(0)} + p_\mu \right) S(x, p) &= 0, \\ \left(-\frac{1}{2}i\hbar\nabla_\mu^{(0)} + p_\mu \right) S(x, p)\sigma^\mu &= 0.\end{aligned}$$

We can get the master equation for massless Wigner function

$$\begin{aligned}\Pi^\mu \mathcal{J}_\mu^s(x, p) &= 0, \\ \nabla^\mu \mathcal{J}_\mu^s(x, p) &= 0, \\ 2s(\Pi^\mu \mathcal{J}_s^\nu - \Pi^\nu \mathcal{J}_s^\mu) &= -\hbar\epsilon^{\mu\nu\rho\sigma}\nabla_\rho \mathcal{J}_\sigma^s,\end{aligned}$$

$s = \pm$ is the chirality index for right-handed (+) and left-handed (-) fermions

Summary for Master eq.

$$\begin{aligned}\Pi^\mu \mathcal{J}_\mu^s(x, p) &= 0, \\ \nabla^\mu \mathcal{J}_\mu^s(x, p) &= 0, \\ 2s(\Pi^\mu \mathcal{J}_s^\nu - \Pi^\nu \mathcal{J}_s^\mu) &= -\hbar \epsilon^{\mu\nu\rho\sigma} \nabla_\rho \mathcal{J}_\sigma^s,\end{aligned}$$

$s = \pm$ is the chirality index for right-handed (+) and left-handed (-) fermions

- Assumption:
 - Constant homogenous EM fields $\partial_\alpha F^{\mu\nu} = 0$
 - No-interactions between particles

2. Master equation for covariant Wigner function

(2c) Master equations in general case (massive case)

General cases

- If we start from a 4x4 Wigner function, by implementing the Dirac equations, we can get, in a classical EM fields,

$$(\gamma \cdot K - m)W = 0$$

$$\begin{aligned} K^\mu &= \Pi^\mu + \frac{1}{2}i\nabla^\mu & \nabla_\mu &= \partial_\mu^x - j_0(\Delta)F_{\mu\nu}(x)\partial_p^\nu, \\ && \Pi_\mu &= p_\mu - \frac{1}{2}\hbar j_1(\Delta)F_{\mu\nu}(x)\partial_p^\nu, \end{aligned}$$

- Next, we need to insert the decomposition of W into the main equations

$$W = \mathcal{F} + i\mathcal{P}\gamma^5 + \mathcal{V}^\mu\gamma_\mu + \mathcal{A}^\mu\gamma^5\gamma_\mu + \frac{1}{2}\mathcal{S}^{\mu\nu}\sigma_{\mu\nu},$$

Vasak, Gyulassy, Elze, Ann. Phys. (N.Y.) 173, 462 (1987);
Elze, Heinz, Phys. Rep. 183, 81 (1989).

Master equations in massive case (I)

- Inserting the decomposition of W into the main equations, yields

$$\begin{aligned} K^\mu &= \Pi^\mu + \frac{1}{2}i\nabla^\mu & K \cdot \mathcal{V} - m\mathcal{F} &= 0, \\ && iK \cdot \mathcal{A} + m\mathcal{P} &= 0, \\ && K_\mu \mathcal{F} - iK^\nu \mathcal{S}_{\mu\nu} - m\mathcal{V}_\mu &= 0, \\ && iK_\mu \mathcal{P} + \frac{1}{2}\epsilon_{\mu\nu\rho\sigma}K^\nu \mathcal{S}^{\rho\sigma} - m\mathcal{A}_\mu &= 0, \\ && i(K_\mu \mathcal{V}_\nu - K_\nu \mathcal{V}_\mu) - \epsilon_{\mu\nu\rho\sigma}K^\rho \mathcal{A}^\sigma + m\mathcal{S}_{\mu\nu} &= 0. \end{aligned}$$

Master equations in massive case (II)

- Since the coefficients V, A, F, P, S are real, we can get

Real part of main eq. gives

$$\left. \begin{aligned} \Pi \cdot V &= mF, \\ \nabla \cdot A &= 2mP, \\ \Pi_\mu F - iK^\nu S_{\mu\nu} &= mV_\mu, \\ -\nabla_\mu P + \frac{1}{2}\epsilon_{\mu\nu\rho\sigma}\Pi^\nu S^{\rho\sigma} &= 2mA_\mu, \\ \frac{1}{2}(\nabla_\mu V_\nu - \nabla_\nu V_\mu) + \epsilon_{\mu\nu\rho\sigma}\Pi^\rho A^\sigma &= mS_{\mu\nu}. \end{aligned} \right\}$$

Imaginary part of main eq. gives

$$\left. \begin{aligned} \nabla \cdot V &= 0, \\ \Pi \cdot A &= 0, \\ \frac{1}{2}\nabla_\mu F - \Pi^\nu S_{\mu\nu} &= 0, \\ \Pi_\mu P + \frac{1}{4}\epsilon_{\mu\nu\rho\sigma}\nabla^\nu S^{\rho\sigma} &= 0, \\ (\Pi_\mu V_\nu - \Pi_\nu V_\mu) - \frac{1}{2}\epsilon_{\mu\nu\rho\sigma}\nabla^\rho A^\sigma &= 0. \end{aligned} \right\}$$

Master eqs. in massless limit (I)

- In massless limit, V, A are decoupled with others.

$$\nabla \cdot \mathcal{V} = 0,$$

$$\Pi \cdot \mathcal{A} = 0,$$

$$\Pi \cdot \mathcal{V} = 0,$$

$$\nabla \cdot \mathcal{A} = 0,$$

$$(\Pi_\mu \mathcal{V}_\nu - \Pi_\nu \mathcal{V}_\mu) - \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} \nabla^\rho \mathcal{A}^\sigma = 0.$$

$$\epsilon^{\mu\nu\alpha\beta} \times \xrightarrow{\quad} \frac{1}{2} (\nabla_\mu \mathcal{V}_\nu - \nabla_\nu \mathcal{V}_\mu) + \epsilon_{\mu\nu\rho\sigma} \Pi^\rho \mathcal{A}^\sigma = 0.$$

$$\frac{1}{2} \nabla_\mu \mathcal{F} - \Pi^\nu \mathcal{S}_{\mu\nu} = 0,$$

$$\Pi_\mu \mathcal{P} + \frac{1}{4} \epsilon_{\mu\nu\rho\sigma} \nabla^\nu \mathcal{S}^{\rho\sigma} = 0,$$

$$\Pi_\mu \mathcal{F} - i K^\nu \mathcal{S}_{\mu\nu} = 0,$$

$$-\nabla_\mu \mathcal{P} + \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} \Pi^\nu \mathcal{S}^{\rho\sigma} = 0.$$

Master eqs. in massless limit (II)

- Eqs. for V and A are similar

$$\left\{ \begin{array}{l} (\Pi_\mu \mathcal{V}_\nu - \Pi_\nu \mathcal{V}_\mu) - \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} \nabla^\rho \mathcal{A}^\sigma = 0, \\ (\Pi_\mu \mathcal{A}_\nu - \Pi_\nu \mathcal{A}_\mu) - \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} \nabla^\rho \mathcal{V}^\sigma = 0, \\ \nabla \cdot \mathcal{V} = 0, \\ \Pi \cdot \mathcal{A} = 0, \\ \Pi \cdot \mathcal{V} = 0, \\ \nabla \cdot \mathcal{A} = 0, \end{array} \right\}$$

Master equations for chiral fermions

- Naturally, we introduce the Wigner function for right and left handed fermions,

$$\mathcal{J}_\mu^s = \mathcal{V}_\mu + s \mathcal{A}_\mu, \quad s = \pm$$

+: right handed
-: left handed

$$\left. \begin{aligned} \Pi^\mu \mathcal{J}_\mu^s(x, p) &= 0, \\ \nabla^\mu \mathcal{J}_\mu^s(x, p) &= 0, \\ 2s (\Pi^\mu \mathcal{J}_s^\nu - \Pi^\nu \mathcal{J}_s^\mu) &= -\hbar \epsilon^{\mu\nu\rho\sigma} \nabla_\rho \mathcal{J}_\sigma^s, \end{aligned} \right\}$$

3. Equal-time formulism for Wigner function

Equal-time Wigner function

- Equal-time Wigner function

I. Bialynicki-Birula, P. Gornicki and J. Rafelski, Phys. Rev. D44, 1825 (1991).

C. Best, P. Gornicki and W. Greiner, Ann. Phys. (N.Y.) 225, 169(1993)

Zhuang, Heinz, Phys. Rev. D 57 (1998) 6525

Ochs, Heinz, Annals Phys. 266 (1998) 351

- Application to Schwinger pair production:

Hebenstreit, Alkofer, Gies, Phys.Rev.D82:105026,2010

Sheng, Fang, Wang, Rischke, Phys.Rev.D 99 (2019) 5, 056004

- Application to QKT:

Chen, Wang, Zhuang, arXiv:2101.07596.

Definition of equal time Wigner function

- Equal time Wigner function is integration of covariant Wigner function over \mathbf{p}^0 ,

$$\begin{aligned}W_{\alpha\beta}(t, \mathbf{x}, \mathbf{p}) &= \int dp^0 W_{\alpha\beta}(x, p) \\&= \int \frac{d^3y}{(2\pi)^3} e^{i\mathbf{p}\cdot\mathbf{y}} \left\langle \bar{\psi}_\beta \left(t, \mathbf{x} + \frac{\mathbf{y}}{2}\right) \exp \left[iQ \int_{\mathbf{x}-\mathbf{y}/2}^{\mathbf{x}+\mathbf{y}/2} d\mathbf{s} \cdot \mathbf{A}(t, \mathbf{s}) \right] \psi_\alpha \left(t, \mathbf{x} - \frac{\mathbf{y}}{2}\right) \right\rangle\end{aligned}$$

- Then, the two spinors are set to be at equal time.

$$\hat{W}_{\alpha\beta} = \int \frac{d^4y}{(2\pi)^4} e^{-ip\cdot y} \bar{\psi}_\beta(x_1) U(x_1, x_2) \psi_\alpha(x_2)$$

Equal time formalism Vs covariant formalism

- Covariant Wigner function:
 - Lorentz covariance
 - Might be hard to be simulated directly.
- Equal time Wigner function:
 - Easy to set up as an initial problem
 - No Lorentz covariance
 - e.g. Schwinger pair production

Zhuang, Heinz, Phys. Rev. D 57 (1998) 6525

Dirac equations and gauge fixing

- Next, we will derive the master eq. for equal time Wigner function.
- We start from the Dirac equation,

$$(i\gamma \cdot D - m + \mu\gamma^0)\psi = 0,$$
$$\bar{\psi}[-i\gamma \cdot (\overleftarrow{\partial} - ieA) - m + \mu\gamma^0] = 0.$$

→ $i(\partial_t + ieA_t)\psi = [-i\gamma^0\gamma \cdot (\partial - ie\mathbf{A}) + m\gamma^0 - \mu]\psi,$

$$-i(\partial_t - ieA_t)\bar{\psi} = \bar{\psi}[i\gamma \cdot (\overleftarrow{\partial} + ie\mathbf{A})\gamma^0 + m\gamma^0 - \mu].$$

- For simplicity, we choose the gauge fixing condition

$$A^t = 0. \quad \mathbf{E} = F^{i0} = \partial^i A^0 - \partial^0 A^i = -\partial_t \mathbf{A},$$

$$\mathbf{B} = -\frac{1}{2}\epsilon^{ijk}F^{jk} = -\epsilon^{ijk}\partial^j A^k = \nabla \times \mathbf{A}.$$

Equations for equal time Wigner function (I)

- For simplicity, we set EM fields are homogenous in space.

$$\begin{aligned} D_t W_{\alpha\beta} &= [\partial_t + e\mathbf{E}(t) \cdot \nabla_{\mathbf{p}}] W_{\alpha\beta} \\ &= \int \frac{d^3y}{(2\pi)^3} \exp \left[i\mathbf{y} \cdot \mathbf{p} + ie \int_{\mathbf{x}_2}^{\mathbf{x}_1} d\mathbf{s} \cdot \mathbf{A}(t, \mathbf{s}) \right] \left\langle \partial_t \left[\bar{\psi}_\beta \left(t, \mathbf{x} + \frac{\mathbf{y}}{2} \right) \psi_\alpha \left(t, \mathbf{x} - \frac{\mathbf{y}}{2} \right) \right] \right\rangle \\ &= -im [W, \gamma^0]_{\alpha\beta} \\ &\quad - \int \frac{d^3y}{(2\pi)^3} \exp \left[i\mathbf{y} \cdot \mathbf{p} + ie \int_{\mathbf{x}_2}^{\mathbf{x}_1} d\mathbf{s} \cdot \mathbf{A}(t, \mathbf{s}) \right] \\ &\quad \times \left\langle [\bar{\psi}(\boldsymbol{\gamma} \cdot (\overleftrightarrow{\partial} + ie\mathbf{A})\gamma^0)]_\beta \psi_\alpha + \bar{\psi}_\beta [\gamma^0 \boldsymbol{\gamma} \cdot (\partial_x - ie\mathbf{A})\psi]_\alpha \right\rangle \end{aligned}$$

- Integration by part and after a long calculation, we get,

$$[\partial_t + e\mathbf{E}(t) \cdot \nabla_{\mathbf{p}}] W = \frac{1}{2} (\nabla_x + e\mathbf{B} \times \nabla_{\mathbf{p}}) [W, \gamma^0 \boldsymbol{\gamma}] - i\mathbf{p} \cdot \{W, \gamma^0 \boldsymbol{\gamma}\} + im [W, \gamma^0].$$

Equation for equal time Wigner function (II)

- We summary the master eqs. here,

$$D_t W = \frac{1}{2} \mathbf{D}_x \cdot [W, \gamma^0 \boldsymbol{\gamma}] - i \boldsymbol{\Pi} \cdot \{W, \gamma^0 \boldsymbol{\gamma}\} + im [W, \gamma^0],$$

$$D_t = \partial_t + e \mathbf{E}(t) \cdot \nabla_{\mathbf{p}},$$

$$\mathbf{D}_x = \nabla_x + e \mathbf{B} \times \nabla_{\mathbf{p}},$$

$$\boldsymbol{\Pi} = \mathbf{p}.$$

- Similarly, we decompose the Wigner function as

$$W = \mathcal{F} + i \mathcal{P} \gamma^5 + \mathcal{V}^\mu \gamma_\mu + \mathcal{A}^\mu \gamma^5 \gamma_\mu + \frac{1}{2} \mathcal{S}^{\mu\nu} \sigma_{\mu\nu},$$

$$\mathbf{T} = (\mathcal{S}^{10}, \mathcal{S}^{20}, \mathcal{S}^{30}),$$

$$\mathbf{S} = (\mathcal{S}^{23}, \mathcal{S}^{31}, \mathcal{S}^{12}).$$

Master equation for equal time Wigner function

$$D_t \mathcal{F} = 2\boldsymbol{\Pi} \cdot \mathbf{T},$$

$$D_t \mathcal{P} = -2\boldsymbol{\Pi} \cdot \mathbf{S} + 2m\mathcal{A}^0,$$

$$D_t \mathcal{V}^0 = -\mathbf{D}_x \cdot \mathbf{V},$$

$$D_t \mathbf{V} = -\mathbf{D}_x V^0 + 2\boldsymbol{\Pi} \times \mathbf{A} - 2m\mathbf{T},$$

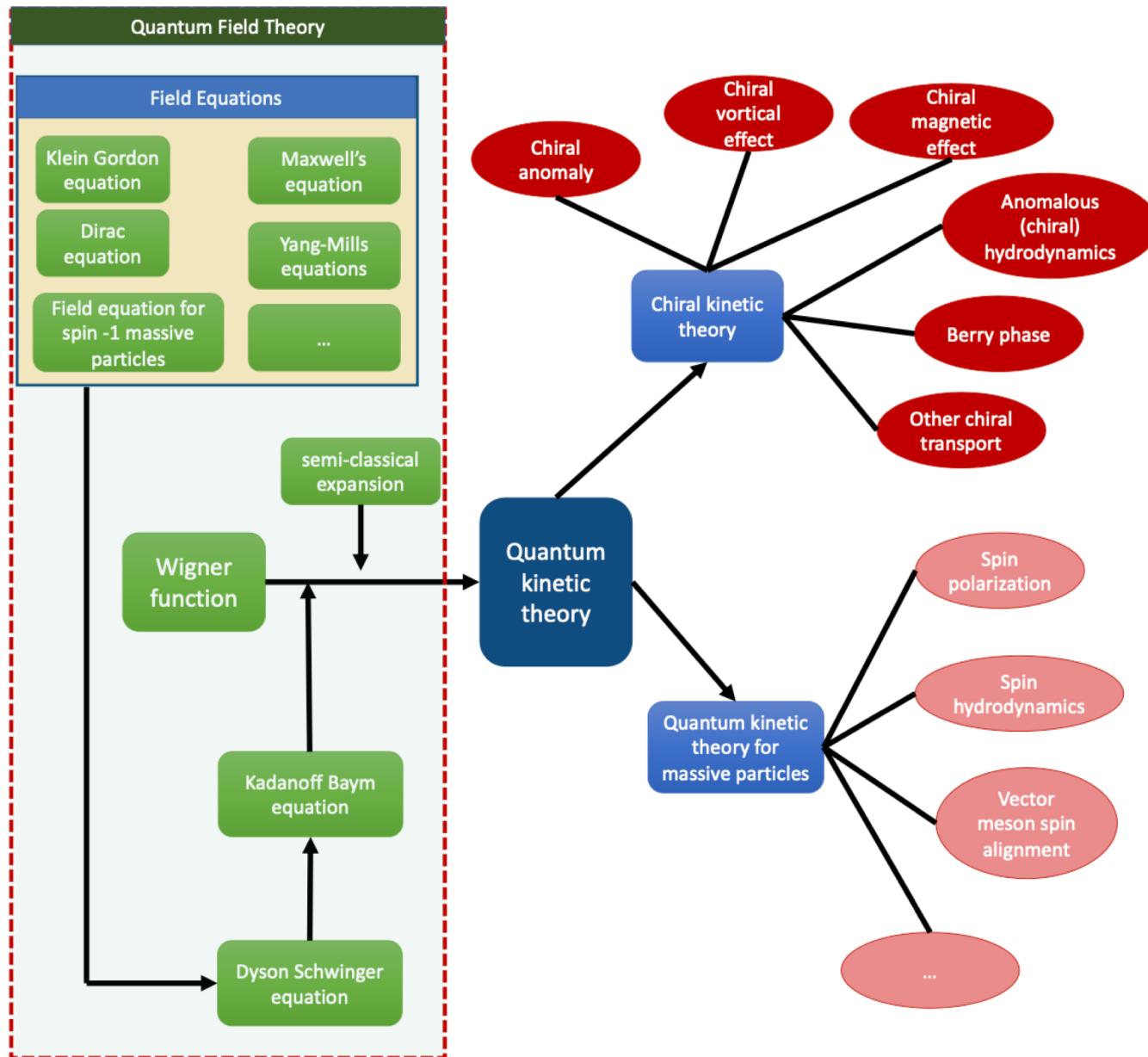
$$D_t \mathcal{A}^0 = -\mathbf{D}_x \cdot \mathbf{A} - 2m\mathcal{P},$$

$$D_t \mathbf{A} = -\mathbf{D}_x \mathcal{A}^0 + 2\boldsymbol{\Pi} \times \mathbf{V},$$

$$\mathbf{D}_t \mathbf{T} = \mathbf{D}_x \times \mathbf{S} - 2\boldsymbol{\Pi} \mathcal{F} + 2m\mathbf{V},$$

$$\mathbf{D}_t \mathbf{S} = \mathbf{D}_x \times \mathbf{T} + 2\boldsymbol{\Pi} \mathcal{P}.$$

Summary

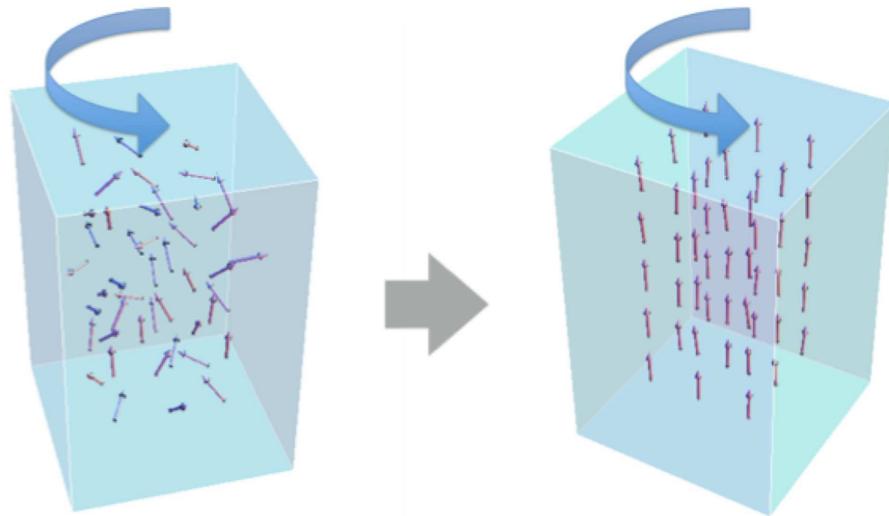


Thank you for your time!

欢迎批评指正！

Backup

Barnet effects and Einstein-de Hass effects



Barnett effect:

Rotation \Rightarrow Magnetization

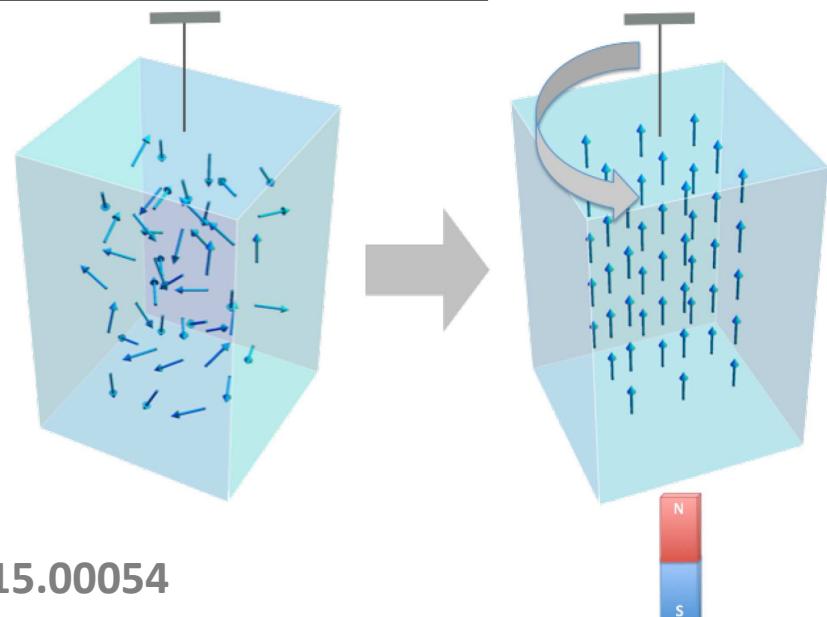
Barnett, Magnetization by rotation, Phys Rev. (1915) 6:239–70.

Einstein-de Haas effect:

Magnetization \Rightarrow Rotation

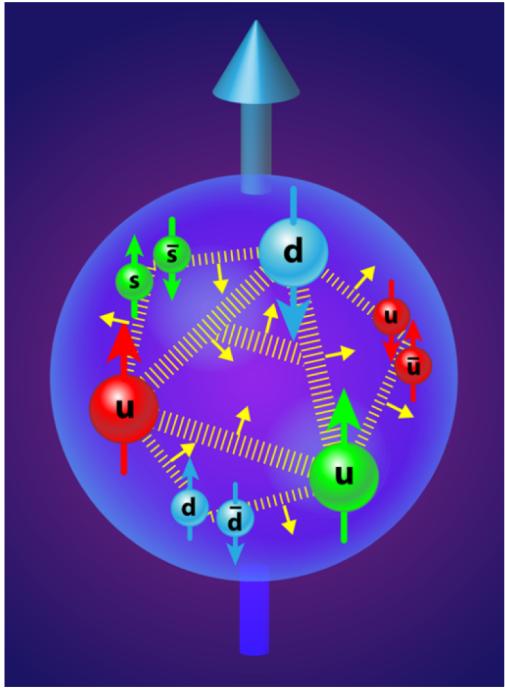
Einstein, de Haas, Experimental proof of the existence of Ampere's molecular currents.

Verh Dtsch Phys Ges. (1915) 17:152.



Figures: copy from paper doi: 10.3389/fphy.2015.00054

Connection to “spin physics” (QCD)



- Proton spin problem:
(slides from Hatta's talk)

$$\frac{1}{2} = \frac{1}{2} \Delta \Sigma + \Delta G + L^q + L^g$$

↑ ↑ ↑
Quarks' Gluons' Orbital angular
helicity helicity Momentum (OAM)

Total angular momentum conservation

- Nöther's theorem :

$$x^\mu \rightarrow x'^\mu = \Lambda_\nu^\mu x^\nu = (\delta_\nu^\mu + \epsilon_\nu^\mu) x^\nu, \quad A^\mu(x) \rightarrow A'^\mu(x) = \Lambda_\nu^\mu A^\nu(\Lambda^{-1}x),$$
$$\psi(x) \rightarrow \psi'(x) = \Lambda_{\frac{1}{2}} \psi(\Lambda^{-1}x),$$



$$\partial_\lambda (J_A^{\lambda\mu\nu} + J_\psi^{\lambda\mu\nu}) = 0$$

- Nöther current

Gauge part $J_A^{\lambda\mu\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_\lambda A^\alpha)} \Delta A^{\mu\nu\alpha} = -F_\alpha^\lambda (x^\mu \partial^\nu - x^\nu \partial^\mu) A^\alpha - F^{\lambda\mu} A^\nu + F^{\lambda\nu} A^\mu.$

Fermionic part $J_\psi^{\lambda\mu\nu} = \frac{\partial \mathcal{L}}{\partial(\partial_\lambda \psi)} \Delta \psi^{\mu\nu} = \bar{\psi} i \gamma^\lambda (x^\mu \partial^\nu - x^\nu \partial^\mu - i \Sigma^{\mu\nu}) \psi$

- How to define the orbital and spin parts?

Table of two different forms

- Canonical (Jaffe-Manohar) decomposition

$$\mathbf{J} = \underbrace{-\frac{1}{2}\bar{\psi}\gamma_5\gamma\psi}_{\frac{1}{2}\Delta\Sigma} + \underbrace{\mathbf{E} \times \mathbf{A}}_{\Delta G} - \underbrace{i\psi^\dagger(\mathbf{x} \times \nabla)\psi}_{L_{\text{can}}^q} + \underbrace{\mathbf{E}(\mathbf{x} \times \nabla)\mathbf{A}}_{L_{\text{can}}^g}$$

$$T_{\psi,\text{can}}^{\mu\nu} = \bar{\psi}i\gamma^\mu \cancel{\partial}^\nu \psi - g^{\mu\nu}\bar{\psi}(i\gamma^\alpha D_\alpha - m)\psi$$

non-symmetric
Not gauge invariant

- Belinfante (Ji) decomposition

$$\mathbf{J} = \underbrace{-\frac{1}{2}\bar{\psi}\gamma_5\gamma\psi}_{\frac{1}{2}\Delta\Sigma} - \underbrace{i\psi^\dagger(\mathbf{x} \times \mathbf{D})\psi}_{L_{\text{Ji}}^q} + \underbrace{\mathbf{x} \times (\mathbf{E} \times \mathbf{B})}_{J_{\text{Ji}}^g}$$

$$T_{\psi,\text{Bel}}^{\mu\nu} \equiv \bar{\psi}i\gamma^\mu \cancel{\partial}^\nu \psi + \frac{1}{4}\epsilon^{\mu\nu\lambda\rho}\partial_\lambda(\bar{\psi}\gamma_5\gamma_\rho\psi)$$

Connected by
pseudo gauge
transformation

Symmetric
Gauge invariant

GLW decomposition

- Another Pseudo gauge transformation

$$T_{\text{can}}^{\mu\nu} = T_{\text{GLW}}^{\mu\nu} + \frac{1}{2}\partial_\lambda \left(\Phi_{\text{can}}^{\lambda,\mu\nu} + \Phi_{\text{can}}^{\mu,\nu\lambda} + \Phi_{\text{can}}^{\nu,\mu\lambda} \right)$$

$$S_{\text{can}}^{\lambda,\mu\nu} = S_{\text{GLW}}^{\lambda,\mu\nu} - \Phi_{\text{can}}^{\lambda,\mu\nu} \quad \Phi_{\text{can}}^{\lambda,\mu\nu} \equiv S_{\text{GLW}}^{\mu,\lambda\nu} - S_{\text{GLW}}^{\nu,\lambda\mu}$$

$$\partial_\lambda S_{\text{can}}^{\lambda,\mu\nu}(x) = T_{\text{can}}^{\nu\mu} - T_{\text{can}}^{\mu\nu} = -\partial_\lambda S_{\text{GLW}}^{\mu,\lambda\nu}(x) + \partial_\lambda S_{\text{GLW}}^{\nu,\lambda\mu}(x).$$

$$T_{\text{Bel}}^{\mu\nu} = T_{\text{GLW}}^{\mu\nu} - \frac{1}{2}\partial_\lambda \left(S_{\text{GLW}}^{\nu,\lambda\mu} + S_{\text{GLW}}^{\mu,\lambda\nu} \right)$$

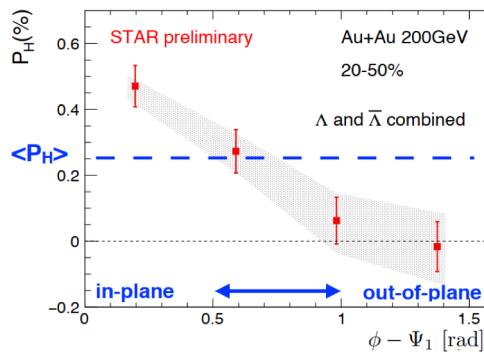
Textbook written by de Groot, van Leeuwen, and van Weert

Review: *W. Florkowski, R. Ryblewski and Avdhesh Kumar, Prog. Part. Nucl. Phys. 108 (2019) 103709*

- Microscopic kinetic theory: GLW is the classical one.

$$T_{\text{GLW}}^{\mu\nu}(x) = \frac{1}{m} \text{tr}_4 \int d^4k k^\mu k^\nu \mathcal{W}(x, k) = \frac{1}{m} \int d^4k k^\mu k^\nu \mathcal{F}(x, k).$$

Polarization induced by thermal vorticity



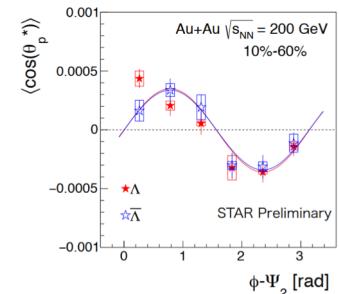
Thermal vorticity

$$\omega_{\rho\sigma}^{\text{th}} = \frac{1}{2} \left[\partial_\rho \left(\frac{u_\sigma}{T} \right) - \partial_\sigma \left(\frac{u_\rho}{T} \right) \right]$$

Distribution function: f_0

$$S^\mu(p) = \frac{1}{8m_\Lambda} \epsilon^{\mu\nu\rho\sigma} p_\nu \frac{\int d\Sigma_\lambda p^\lambda f_0 (1 - f_0) \omega_{\rho\sigma}^{\text{th}}}{\int d\Sigma_\lambda p^\lambda f_0}$$

Freezeout surface



Karpenko, F. Becattini, Eur. Phys. J. C 77 (2017) 213

R.-H. Fang, L.-G. Pang, Q. Wang, X.-N. Wang, Phys. Rev. C94, 024904 (2016)

Polarization and axial current

- Recalling the original equations

$$\mathcal{S}^\mu(\mathbf{p}) = \frac{\int d\Sigma \cdot p \mathcal{J}_5^\mu(p, X)}{2m_\Lambda \int d\Sigma \cdot \mathcal{N}(p, X)},$$

- For massless fermions, the left and right handed currents read

$$\begin{aligned}\mathcal{J}_\lambda^\mu(p, X) = & 2\pi \text{sign}(u \cdot p) \left\{ p^\mu + \lambda \frac{\hbar}{2} \delta(p^2) [u^\mu(p \cdot \omega) - \omega^\mu(u \cdot p) \right. \\ & \left. - 2S_{(u)}^{\mu\nu} \tilde{E}_\nu] \partial_{u \cdot p} + \lambda \frac{\hbar}{4} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} \partial_\nu^p \delta(p^2) \right\} f_\lambda^{(0)},\end{aligned}$$

$$S_{(u)}^{\mu\nu} = \epsilon^{\mu\nu\alpha\beta} p_\alpha u_\beta / (2u \cdot p),$$

$\lambda = \pm$
+: right
-: left

$$\tilde{E}_\nu = E_\nu + T \partial_\nu \frac{\mu_\lambda}{T} + \frac{(u \cdot p)}{T} \partial_\nu T - p^\sigma [\partial_{<\sigma} u_{\nu>} + \frac{1}{3} \Delta_{\sigma\nu} (\partial \cdot u) + u_\nu D u_\sigma].$$

$$f_\lambda^{(0)} = 1/(e^{(u \cdot p - \mu_\lambda)/T} + 1),$$

Y. Hidaka, SP, and D.L. Yang, Phys. Rev. D97, 016004 (2018)

