#### **QCD Phase Diagram and the Equation of State of Strong-Interaction Matter**

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Faculty of Physics

with a lot of input from 'HotQCD,

in particular, David Clarke Jishnu Goswami Anirban Lahiri Mugdha Sarkar Sipaz Sharma

- QCD phase diagram constraining the location of a critical endpoint
- Equation of state in the BES energy range
- Fluctuations and correlations of conserved charges



#### **Exploring the phase diagram of strong-interaction matter**



# Simulating strongly interacting matter on a discrete space-time grid (lattice QCD)



#### Phases of strong-interaction matter

determination of  $T_c^0$  puts an upper limit on  $\,T^{CEP}_{}$ 



Random Matrix<br/>ModelA. Halasz, A.D. Jackson, R.E. Shrock, M.A. Stephanov,<br/>J.J.M. Verbaarschot, Phys. Rev. D58 (1998) 096007OCD motivatedM. Stephanov, Phys. Rev. D73 (2006) 094508

NJL M. Buballa, S. Carignano, Phys. Lett. B791 (2019) 361

for  $m_\ell 
ightarrow 0$  increasingly unlikely

F. Cuteri et al, arXiv:2107.12739 Sipaz Sharma et al, arXiv:2111.12599

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#### **Critical behavior in QCD**



#### **Fluctuation observables in QCD**

- chiral condensate: 
$$\langle \bar{\psi}\psi \rangle_l = \frac{\partial P/T}{\partial m_l/T}$$
,  $\langle \bar{\psi}\psi \rangle_l = (\langle \bar{\psi}\psi \rangle_u + \langle \bar{\psi}\psi \rangle_d)/2$   
- chiral order parameter:  $M = \frac{2}{f_K^4} \left[m_s \langle \bar{\psi}\psi \rangle_l - m_l \langle \bar{\psi}\psi \rangle_s\right]$   
 $m_l = (m_u + m_d)/2$ 

– chiral susceptibility: 
$$\chi_M = m_s \left( \frac{\partial M}{\partial m_u} + \frac{\partial M}{\partial m_d} \right)$$
 magnetic

– mixed chiral susceptibility: 
$$\chi_t = T rac{\partial M}{\partial T}$$
 mixed

– conserved charge fluctuations: 
$$\chi_{x} = T^{4} \frac{\partial^{4} P/T^{4}}{\partial \mu_{X}^{4}}\Big|_{\mu_{X}=0}$$
 thermal

$$X=B,\ S,\ldots$$



F. Karsch, RHIC-BES seminar, October 2022

#### The Chiral PHASE TRANSITION in (2+1)-flavor QCD



#### **Pseudo-critical and critical temperatures**



F. Karsch, RHIC-BES seminar, October 2022

#### **Curvature of the pseudo-critical line**

 $-\mu_B$ -dependent shift of maxima in susceptibilities

$$rac{\partial \chi_M(T,\mu_B)}{\partial T} = 0$$
 or  $rac{\partial^2 M(T,\mu_B)}{\partial T^2} = 0$  or  $rac{\partial^3 M(T,\mu_B)}{\partial T \partial \mu_B^2} = 0$ 

- Taylor series, e.g.

$$\chi_M(T,\mu_B) = \chi_M(T_{pc},0) + \frac{\partial \chi_M}{\partial T}(T-T_{pc}) + \frac{1}{2} \frac{\partial^2 \chi_M}{\partial (\mu_B/T)^2} \Big|_{\mu_B=0} \left(\frac{\mu_B}{T}\right)^2 + \dots$$

$$\kappa_2 \simeq \frac{T^2 \partial^2 \chi_M/\partial \mu_B^2}{2T \partial \chi_M/\partial T} \Big|_{\mu_B=0}$$

$$T_{pc}(\mu_B) = T_{pc} \left(1 - \kappa_2 \left(\frac{\mu_B}{T}\right)^2 - \kappa_4 \left(\frac{\mu_B}{T}\right)^4 + \dots\right)$$

#### Curvature of the pseudo-critical line – towards the chiral limit –

– universality relations also relate derivatives with respect to T and  $\mu_B$ 

$$rac{1}{VT^3} \ln Z(V,T,ec{\mu}) \sim -h^{(2-lpha)/eta\delta} f_f(t/h^{1/eta\delta})$$

$$h \sim \frac{m_l}{T_c} , \ t \sim \frac{T - T_c}{T_c} + \kappa_2 \left(\frac{\mu_B}{T}\right)^2 \iff \begin{bmatrix} \frac{\partial^2}{\partial (\mu_B/T)^2} &\simeq \frac{\partial}{\partial T} \end{bmatrix}$$
FK et al., arXiv:1009.5211

#### Curvature of the pseudo-critical line – towards the chiral limit –



$$t\sim rac{T-T_c}{T_c}+\kappa_2(H)\left(rac{\mu_B}{T}
ight)^2 \ , \ H=m_l/m_s$$

curvature of crossover line only mildly dependent on H

#### **Phases of strong-interaction matter**



# Summary: Phases of strong-interaction matter determination of $T_c^0$ puts an upper limit on $T^{CEP}$



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#### QCD thermodynamics at non-zero net baryon-density – Taylor expansion –

Taylor expansion of the QCD pressure:  $rac{P}{T^4} = rac{1}{VT^3} \ln Z(T,V,\mu_B,\mu_Q,\mu_S)$ 

$$\boxed{\frac{P}{T^4} = \sum_{i,j,k=0}^{\infty} \frac{1}{i!j!k!} \chi^{BQS}_{ijk}(T) \left(\frac{\mu_B}{T}\right)^i \left(\frac{\mu_Q}{T}\right)^j \left(\frac{\mu_S}{T}\right)^k}$$

cumulants of net-charge fluctuations and correlations:

$$\chi^{BQS}_{ijk} = \left. \frac{\partial^{i+j+k} P/T^4}{\partial \hat{\mu}^i_B \partial \hat{\mu}^j_Q \partial \hat{\mu}^k_S} \right|_{\mu_{B,Q,S}=0} \quad , \quad \hat{\mu}_X \equiv \frac{\mu_X}{T}$$

#### QCD thermodynamics at non-zero net baryon-density – Taylor expansion –

Taylor expansion of the QCD pressure:  $rac{P}{T^4} = rac{1}{VT^3} \ln Z(T,V,\mu_B,\mu_Q,\mu_S)$ 

II) 
$$n_S = 0: \hat{\mu}_S \to \hat{\mu}_S(\hat{\mu}_B) = s_1\hat{\mu}_B + s_3\hat{\mu}_B^3 + \dots$$
  
 $n_Q/n_B = r: \hat{\mu}_Q \to \hat{\mu}_Q(\hat{\mu}_B) = q_1\hat{\mu}_B + q_3\hat{\mu}_B^3 + \dots$ 

isospin symmetric matter  $\iff \mu_Q = 0 \iff n_Q/n_B = 0.5$ conditions met in heavy ion collisions differ  $n_Q/n_B = 0.4$ 

#### QCD thermodynamics at non-zero net baryon-density – Taylor expansion –

Taylor expansion of the QCD pressure:  $rac{P}{T^4} = rac{1}{VT^3} \ln Z(T,V,\mu_B,\mu_Q,\mu_S)$ 

Taylor expansion to 8th order in  $\mu_B/T$ 

D. Bollweg et al (HotQCD), arXiv:2202.09184 and QM 2022

#### HotQCD data collection for (2+1)-flavor QCD-EoS

#### EoS:2017: arXiv:1701.04325

$N_{ au}=6$					$N_{ au}=8$				$N_{ au} = 12$				
$\beta$	$m_l$	T[Me	eV] #coi	nf. $\beta$	$m_l$	T[MeV	] #conf.	$\beta$	$m_l$	T[Me]	V] #con	nf.	
5.980	0.00435	135.2	29 812	00 6.24	5  0.00307	134.64	180320	6.640	0.00196	134.9	4 58	34	
6.010	0.00416	139.'	71 1207	$90 \parallel 6.28$	5 0.00293	140.45	172110	6.680	0.00187	140.4	4 58	33	
6.045	0.00397	145.0	05  1207	$70 \  6.31$	5 0.00281	144.95	138150	6.712	0.00181	144.9	138	46	
6.080	0.00387	150.	59 793	$90 \  6.35$	4 0.00270	151.00	107510	6.754	0.00173	151.1	0 142	00	
6.120	0.00359	157.	17 661	$80 \  6.39$	0 0.00257	156.78	135730	6.794	0.00167	157.1	3 154'	76	
6.150	0.00345	162.2	28 796	$60 \  6.42$	3 0.00248	162.25	115850	6.825	0.00161	161.9	4 167	72	
6.170	0.00336	165.9	98 497	$60 \  6.44$	5  0.00241	165.98	120270	6.850	0.00157	165.9	11195	42	
6.200	0.00324	171.	15  1227	$00 \parallel 6.47$	4 0.00234	171.02	139980	6.880	0.00153	170.7	7 212	20	
6.225	0.00314	175.'	76 1227	$30 \parallel 6.50$	0 0.00228	175.64	133070	6.910	0.00148	175.7	6 123	03	
EoS 20	022:			time	s new		tim	es nev	v confs			а	ll new
EoS 20 arXiv:2	022: 2202.09	9184	N	time: conf: $\tau = 8$	s new s.				v confs		$\overline{N_{ au}}$	a = 16	ll new
EoS 20 arXiv:2	022: 2202.09	9184 	$\frac{N}{m_l}$	times confs r = 8 T[MeV]	s new s. #conf.	β	$\frac{1}{m_l}$	= 12 T[MeV]	v confs #conf.	$\beta$	$\frac{N_{ au}}{m_l}$	$\frac{16}{T[MeV]}$	Il new
EoS 2( arXiv:2	022: 2202.09	9184 β 6.175	$\frac{N}{0.003307}$	time: conf: <u>r = 8</u> T[MeV] 125.28	5 new 5. #conf. 2,200,000	β	$\frac{1}{m_l}$	<b>es nev</b> = 12 T[MeV]	#conf.	β	$rac{N_{ au}}{m_l}$	a = 16 T[MeV]	ll new
EoS 20 arXiv:2	022: 2202.09	<b>β</b> 6.175 6.245	N m <sub>l</sub> 0.003307 0.00307	times confs <u>r = 8</u> T[MeV] 125.28 134.84	5 new 5. #conf. 2,200,000 1,275,380	β 6.640	$\frac{N_{\tau}}{m_l}$ 0.00196	<b>ES NEV</b> = 12 T[MeV] 135.24	#conf. 330,447	β 6.935	$\frac{N_{\tau}}{m_l}$ 0.00145	<b>= 16</b> T[MeV] 135.80	<b>II NEW</b> #conf. 17671
EoS 2( arXiv:2	022: 2202.09	<b>β</b> 6.175 6.245 6.285	<b>N</b> <b>m</b> <sub>l</sub> 0.003307 0.00307 0.00293	time conf $\tau = 8$ T[MeV] 125.28 134.84 140.62	mew #conf. 2,200,000 1,275,380 1,598,555	β 6.640 6.680	$\frac{N_{\tau}}{m_{l}} = \frac{N_{\tau}}{m_{l}}$ 0.00196 0.00187	<b>EX NEW</b> <b>= 12</b> T[MeV] 135.24 140.80	#conf. 330,447 441,115	β 6.935 6.973	$N_{ au}$ $m_l$ 0.00145 0.00139	<b>= 16</b> T[MeV] 135.80 140.86	#conf. 17671 23855
EoS 20 arXiv:2	022: 2202.09	<ul> <li>β</li> <li>6.175</li> <li>6.245</li> <li>6.285</li> <li>6.315</li> </ul>	N m <sub>l</sub> 0.003307 0.00307 0.00293 0.00281	time conf <u>r = 8</u> T[MeV] 125.28 134.84 140.62 145.11	#conf. 2,200,000 1,275,380 1,598,555 1,559,003	β 6.640 6.680 6.712	$\frac{N_{\tau}}{m_l} = \frac{0.00196}{0.00187}$	<b>EXAMPLE</b> 12 T[MeV] 135.24 140.80 145.40	#conf. 330,447 441,115 416,703	β 6.935 6.973 7.010	$rac{N_{ au}}{m_l}$ 0.00145 0.00139 0.00132	<b>= 16</b> T[MeV] 135.80 140.86 145.95	#conf. 17671 23855 26122
EoS 20 arXiv:2	022: 2202.09	<ul> <li>β</li> <li>6.175</li> <li>6.245</li> <li>6.285</li> <li>6.315</li> <li>6.354</li> </ul>	N m <sub>l</sub> 0.003307 0.00293 0.00281 0.00270	time conf $\tau = 8$ T[MeV] 125.28 134.84 140.62 145.11 151.14	#conf. 2,200,000 1,275,380 1,598,555 1,559,003 1,286,603	β 6.640 6.680 6.712 6.754	$\frac{N_{\tau}}{m_l} = \frac{m_l}{0.00196}$ 0.00187 0.00181 0.00173	<b>EXAMPLE 12 135.24 140.80 145.40 151.62</b>	#conf. 330,447 441,115 416,703 323,738	β 6.935 6.973 7.010 7.054	$N_{ au}$ $m_l$ 0.00145 0.00139 0.00132 0.00129	<b>= 16</b> T[MeV] 135.80 140.86 145.95 152.19	#conf. #7671 23855 26122 26965
EoS 20 arXiv:2	022: 2202.09	<ul> <li>β</li> <li>6.175</li> <li>6.245</li> <li>6.285</li> <li>6.315</li> <li>6.354</li> <li>6.390</li> <li>6.492</li> </ul>	$\begin{array}{r} N \\ \hline m_l \\ 0.003307 \\ 0.00307 \\ 0.00293 \\ 0.00281 \\ 0.00270 \\ 0.00257 \\ 0.00257 \end{array}$	$timesconf\tau = 8T[MeV]125.28134.84140.62145.11151.14156.92$	mew #conf. 2,200,000 1,275,380 1,598,555 1,559,003 1,286,603 1,602,684 1,602,684	β 6.640 6.680 6.712 6.754 6.794 6.794	$\frac{N_{\tau}}{m_l}$ 0.00196 0.00187 0.00181 0.00173 0.00167 0.00167	ES NEW ES NEW = 12 T[MeV] 135.24 140.80 145.40 151.62 157.75	#conf. 330,447 441,115 416,703 323,738 299,029	β 6.935 6.973 7.010 7.054 7.095	$egin{array}{c} N_{ au} \ m_l \ 0.00145 \ 0.00139 \ 0.00132 \ 0.00129 \ 0.00124 \ 0.$	<b>= 16</b> T[MeV] 135.80 140.86 145.95 152.19 158.21	#conf. #conf. 17671 23855 26122 26965 21656
EoS 20 arXiv:2	022: 2202.09	<ul> <li>β</li> <li>6.175</li> <li>6.245</li> <li>6.285</li> <li>6.315</li> <li>6.354</li> <li>6.390</li> <li>6.423</li> <li>6.423</li> </ul>	$\begin{array}{c} N\\ \hline m_l\\ 0.003307\\ 0.00307\\ 0.00293\\ 0.00281\\ 0.00270\\ 0.00257\\ 0.00248\\ 0.00248\\ 0.00241 \end{array}$	$timeconf\tau = 8T[MeV]125.28134.84140.62145.11151.14156.92162.39$	mew #conf. 2,200,000 1,275,380 1,598,555 1,559,003 1,286,603 1,602,684 1,437,436 1,437,436	β 6.640 6.680 6.712 6.754 6.754 6.825 6.825	$\frac{N_{\tau}}{m_l} = \frac{m_l}{0.00196}$ 0.00187 0.00181 0.00173 0.00167 0.00161 0.00157	<b>EXAMPLE 12</b> T[MeV] 135.24 140.80 145.40 151.62 157.75 162.65	#conf. 330,447 441,115 416,703 323,738 299,029 214,671	β 6.935 6.973 7.010 7.054 7.095 7.130	$\frac{N_{\tau}}{m_l}$ 0.00145 0.00139 0.00132 0.00129 0.00124 0.00119 0.00116	<b>= 16</b> T[MeV] 135.80 140.86 145.95 152.19 158.21 163.50	#conf. #conf. 17671 23855 26122 26965 21656 18173
EoS 20 arXiv:2	022: 2202.09	β           6.175           6.245           6.315           6.354           6.390           6.423           6.445           6.454	$\begin{array}{r} N \\ \hline m_l \\ 0.003307 \\ 0.00307 \\ 0.00293 \\ 0.00281 \\ 0.00270 \\ 0.00257 \\ 0.00248 \\ 0.00241 \\ 0.00224 \end{array}$	times configure 6 configure	#conf. 2,200,000 1,275,380 1,598,555 1,559,003 1,286,603 1,602,684 1,437,436 1,186,523 272,644	β 6.640 6.680 6.712 6.754 6.794 6.825 6.850 6.850 6.820	$\frac{N_{\tau}}{m_l}$ 0.00196 0.00187 0.00181 0.00173 0.00167 0.00161 0.00157 0.00157	= 12 $T[MeV]$ $135.24$ $140.80$ $145.40$ $151.62$ $157.75$ $162.65$ $166.69$ $171.65$	#conf. 330,447 441,115 416,703 323,738 299,029 214,671 156,111 144,622	β 6.935 6.973 7.010 7.054 7.095 7.130 7.156 7.189	$\frac{N_{\tau}}{m_l}$ 0.00145 0.00139 0.00132 0.00129 0.00124 0.00119 0.00116 0.00112	<b>= 16</b> T[MeV] 135.80 140.86 145.95 152.19 158.21 163.50 167.53 179.60	#conf. #conf. 17671 23855 26122 26965 21656 18173 19926 17162
EoS 20 arXiv:2	022: 2202.09	β           6.175           6.245           6.285           6.315           6.354           6.390           6.423           6.445           6.474           6.500	$\begin{array}{r} N \\ \hline m_l \\ 0.003307 \\ 0.00307 \\ 0.00293 \\ 0.00281 \\ 0.00270 \\ 0.00257 \\ 0.00248 \\ 0.00241 \\ 0.00234 \\ 0.00234 \\ 0.00228 \end{array}$	times configure for the conf	mew #conf. 2,200,000 1,275,380 1,598,555 1,559,003 1,286,603 1,602,684 1,437,436 1,186,523 373,644 204,211	β 6.640 6.680 6.712 6.754 6.754 6.825 6.850 6.880 6.910	$\frac{N_{\tau}}{m_l}$ 0.00196 0.00187 0.00187 0.00181 0.00167 0.00167 0.00167 0.00157 0.00153 0.00148	<b>EXAMPLE EXAMPLE EXAMP</b>	#conf. 330,447 441,115 416,703 323,738 299,029 214,671 156,111 144,633 121,248	β 6.935 6.973 7.010 7.054 7.095 7.130 7.156 7.188 7.220	$\frac{N_{\tau}}{m_l}$ 0.00145 0.00139 0.00132 0.00129 0.00124 0.00119 0.00116 0.00113 0.00110	<b>= 16</b> T[MeV] 135.80 140.86 145.95 152.19 158.21 163.50 167.53 172.60 177.80	#conf. #conf. 17671 23855 26122 26965 21656 18173 19926 17163 2282

#### Up to 8<sup>th</sup> order Taylor expansion for pressure



F. Karsch, RHIC-BES seminar, October 2022

#### Up to 8<sup>th</sup> order cumulants are used frequently – imag. chem. pot. extrapolations –



S. Borsanyi et al. , JHEP 10 (2018) 205, arXiv:1805.04445

#### Equation of state of (2+1)-flavor QCD: $\mu_B/T>0$

$$\frac{\Delta p(T,\mu_B)}{T^4} = \frac{p(T,\mu_B) - p(T,0)}{T^4} = P_2(T) \left(\frac{\mu_B}{T}\right)^2 + P_4(T) \left(\frac{\mu_B}{T}\right)^4 + P_6(T) \left(\frac{\mu_B}{T}\right)^6 + \dots$$

#### EoS 2017:



#### EoS 2022: Taylor series to 8<sup>th</sup> order



#### net baryon-number density 0.7 $n_{\rm S} = 0$ , $n_{\rm O} / n_{\rm B} = 0.5$ $\mu_{\rm B} / T = 3$ ·O(μ<sub>B</sub> / T)<sup>6</sup> 0.6 $\mu_{\rm B}$ / T = 2.5 $O(\mu_{\rm B} / T)^4$ 0.5 O(µ<sub>B</sub> / T)<sup>8</sup> 0.5 20.4 - 0.4 - 0.3 $\mu_{\rm B}$ / T = 2 $\mu_{\rm B}$ / T = 1.5 $\mu_{\rm B}$ / T = 1.0 0.2 0.1 T [MeV] 0 140 160 180 220 240 260 200 280

range of reliability of 8<sup>th</sup> order results depends on T-region

 $T \geq 200 \; {
m MeV} \;\; \Rightarrow \; \mu_B/T \geq 3$ 

 $T \leq 200 \; {
m MeV} \;\; \Rightarrow \; \mu_B/T \simeq 2.5$ 

...but limited by statistics

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D. Bollweg et al (HotQCD), arXiv:2202.09184 HotQCD in preparation

#### EoS 2022: Taylor series to 8<sup>th</sup> order



### EoS 2022: Taylor series to 8<sup>th</sup> order eliminating the chemical potential



4<sup>th</sup> order Taylor series is a good approximation in this parameter range

#### closed expression for EoS

$$\hat{\mu}_B(T, n_B) = y(T, n_B)^{1/3} - \frac{N_1(T)}{3N_3(T)} y(T, n_B)^{-1/3} , \qquad \hat{\mu}_B \equiv \mu_B/T$$

$$y(T, n_B) = \frac{N_1(T)}{N_3(T)} \left( \frac{\hat{n}_B}{2N_1(T)} + \sqrt{\frac{N_1(T)}{27N_3(T)} + \left(\frac{\hat{n}_B}{2N_1(T)}\right)^2} \right) , \ \hat{n}_B \equiv n_B/T^3$$

F. Karsch, RHIC-BES seminar, October 2022

#### Pressure and energy density on lines of constant s/n<sub>B</sub>



 – 4<sup>th</sup> order Taylor series suffices to describe EoS in almost the entire parameter range covered by BES (collider mode)

HotQCD, in preparation



#### **Comparing Taylor series and Pade resummation**

**Taylor series** 

$$\begin{split} \frac{\Delta p}{T^4} &\equiv \frac{p(T,\mu_B)}{T^4} - \frac{p(T,0)}{T^4} = \sum_{k=1}^{\infty} P_{2k}(T) \left(\frac{\mu_B}{T}\right)^{2k} \\ &= \frac{P_2^2}{P_4} \left(\bar{x}^2 + \bar{x}^4 + c_{6,2}\bar{x}^6 + c_{8,2}\bar{x}^8 + \ldots\right) \quad \begin{array}{l} \text{use } P_2 > 0, \ P_4 > 0 \\ &\bar{x} \equiv \sqrt{P_4/P_2} \ (\mu_B/T) \\ &c_{6,2} = \frac{P_6P_2}{P_4^2} \ , \ c_{8,2} = \frac{P_8P_2^2}{P_4^3} \end{split}$$

**Pade approximation** 

$$P_{[4,4]} = rac{(1-c_{6,2})ar{x}^2 + (1-2c_{6,2}+c_{8,2})ar{x}^4}{(1-c_{6,2}) + (c_{8,2}-c_{6,2})ar{x}^2 + (c_{6,2}^2-c_{8,2})ar{x}^4}$$

pressure: 
$$\frac{p(T, \mu_B)}{T^4}_{[nm]} = \frac{P_2^2}{P_4} P_{[nm]} \quad \text{(similar for n_B/T^3, e/T^4....)} \quad \text{OR ....}$$
  
energy density: 
$$\left(\frac{\Delta \epsilon}{T^4}\right)_{[nm]} = 3 \left(\frac{\Delta p}{T^4}\right)_{[nm]} + T \frac{\mathrm{d}}{\mathrm{d}T} \left(\frac{P_2^2}{P_4} P_{[nm]}\right) \quad \text{(similar for n_B/T^3, e/T^4....)}$$



0.14

0.12

0.1

0.08

0.06

0.04

0.02

0

0

n<sub>B</sub> / T<sup>3</sup>

T = 135 MeV

dPd[4,4]/dµB Taylor :  $O(\mu_p^7)$ 

0.5

Pd[3,4]

 $\mu_{\rm B}$  / T

1

1.5

#### **Comparing Taylor series and Pade resummation**

agreement between Taylor series and Pade approximants in a larger  $\mu_B$  range at higher temperature; qualitatively similar  $\mu_B$  dependence **HotQCD** in preparation

0.4

0.35

0.3

0.25

0.2

0.15

0.1

0.05

2.5

2

0

0

n<sub>B</sub> / T<sup>3</sup>

T = 155 MeV

Taylor :  $O(\mu_p^7)$ 

0.5

Pd[3,4] dPd[4,4]/dµB

 $\mu_{\rm B}$  / T

1

1.5

& arXiv:2202.09184 29

2.5

F. Karsch, RHIC-BES seminar, October 2022



#### **Comparing Taylor series and Pade resummation**

agreement between Taylor series and Pade approximants in a larger  $\mu_B$  range at higher temperature; qualitatively similar  $\mu_B$  dependence HotQCD in preparation

& arXiv:2202.09184

#### **EoS 2022: Taylor series versus Pade**







energy 8<sup>th</sup> order and Pade-deriv.



#### **Poles of [n,n] Pade approximants in QCD**

$$\hat{\mu}_{B,c}^{\pm} = \pm r_{c,4} e^{\pm i \Theta_{c,4}} , \ r_{c,4} = \sqrt{\frac{12 \tilde{\chi}_0^{B,2}}{\tilde{\chi}_0^{B,4}}} \left| \frac{1 - c_{6,2}}{c_{6,2}^2 - c_{8,2}} \right|^{1/4} , \ c_{2k,2} = \frac{2 \tilde{\chi}_0^{B,2}}{(2k)! \tilde{\chi}_0^{B,2k}} \left( \frac{12 \tilde{\chi}_0^{B,2}}{\tilde{\chi}_0^{B,4}} \right)^{k-1}$$

complex poles move to real axis as temperature decreases

distance of complex poles from the origin is given by the Mercer-Roberts estimator for the radius of convergence



within current errors poles on the real axis (critical point) are possible only for  $T < 135 {
m MeV} \;,\; \mu_B/T > 2.5$ 

higher statistics will sharpen the constraint

# Higher order cumulants – Taylor expansion of QCD EoS and the HRG –

Taylor expansion of the QCD pressure:  $\frac{P}{T^4} = \frac{1}{VT^3} \ln Z(T, V, \mu_B, \mu_Q, \mu_S)$ 

$$egin{aligned} rac{P}{T^4} = \sum_{i,j,k=0}^\infty rac{1}{i!j!k!} \chi^{BQS}_{ijk}(T) \left(rac{\mu_B}{T}
ight)^i \left(rac{\mu_Q}{T}
ight)^j \left(rac{\mu_S}{T}
ight)^k \end{aligned}$$

cumulants of net-charge fluctuations and correlations:

$$\chi^{BQS}_{ijk} = \left. \frac{\partial^{i+j+k} P/T^4}{\partial \hat{\mu}^i_B \partial \hat{\mu}^j_Q \partial \hat{\mu}^k_S} \right|_{\mu_{B,Q,S}=0} \quad , \quad \hat{\mu}_X \equiv \frac{\mu_X}{T}$$

cumulants at vanishing chemical potential provide information on the equation of state as well as freeze-out conditions at small non-zero chemical potential Ratio of baryon number – strangeness correlation and net strangeness fluctuations



D. Bollweg et al. (HotQCD), arXiv:2107.10011

PDG-HRG: uses experimentally known hadron spectrum listed by the Particle Data Group QM-HRG: uses additional hadrons predicted to exist in Quark Model calculations

#### Ratio of baryon number – strangeness correlation and net strangeness fluctuations BS ratios probe flavor-correlations

$$-rac{\chi^{BS}_{11}}{\chi^{S}_{2}}=rac{1}{3}+rac{2}{3}rac{\chi^{us}_{11}}{\chi^{s}_{2}}$$



conserved charge i quark number fluctuations:

$$\chi_{11}^{BS} = -\frac{1}{3}\chi_{11}^{us} - \frac{1}{3}\chi_{11}^{ds} - \frac{1}{3}\chi_2^s$$

 ${\cal O}(g^6 \ln g^2)$ 

J.-P. Blaizot, E. Iancu, A. Rebhan, Phys. Lett. B 523 (2001) 143

O(10) stronger flavor correlations at  $T_{pc}$  than at  $2T_{pc}$ 

#### Baryon number – strangeness chemical potentials at freeze-out from strange baryon yields BS ratios probe strangeness content in an HRG



STAR multi-strange baryon yields are consistent with freeze-out at  $T_{pc}$  and a  $\mu_S/\mu_B$  that reflects contributions from additional strange baryons

at 
$$T_{pc}$$
 QCD:  $\frac{\mu_S}{\mu_B} \simeq 0.24$   
PDG-HRG:  $\frac{\mu_S}{\mu_B} \simeq 0.21$ 

#### baryon number – strangeness correlation from HIC



$$egin{aligned} &-rac{\chi_{BS}}{T^2}>rac{1}{VT^3}[2\langle\Lambda+\Sigma^0
angle+4\langle\Sigma^+
angle\ &+8\langle\Xi
angle+6\langle\Omega^-
angle]=97.4\pm5.8. \end{aligned}$$

$$egin{aligned} &rac{\chi_S}{T^2}\simeq rac{1}{VT^3}[(\langle K^+
angle+\langle K^0
angle+\langle\Lambda+\Sigma^0
angle+\langle\Sigma^+
angle\ &+\langle\Sigma^-
angle+4\langle\Xi^-
angle+4\langle\Xi^0
angle+9\langle\Omega^-
angle+ ext{antiparticles})\ &-(\Gamma_{\phi o K^+}+\Gamma_{\phi o K^-}+\Gamma_{\phi o K^0}+\Gamma_{\phi o ar K^0})\langle\phi
angle]=(504\pm24). \end{aligned}$$

P. Braun-Munzinger, A. Kalweit, K. Redlich, J. Stachel, Phys. Lett. B747 (2015) 292, arXiv:1412.8614 for Skellam distributions mean values and second order cumulants as well as correlations are related:

$$egin{aligned} &rac{\chi_N}{T^2} = rac{1}{VT^3} (\langle N_q 
angle + \langle N_{-q} 
angle) \ &rac{\chi_N}{T^2} = rac{1}{VT^3} \sum_{n=1}^{|q|} n^2 (\langle N_n 
angle + \langle N_{-n} 
angle \end{aligned}$$

$$rac{Q_{NM}}{T^2} = rac{1}{VT^3} \sum_{n=-q_N}^{q_N} \sum_{m=-q_M}^{q_M} nm \langle N_{n,m} 
angle 
onumber \ |q| = 1 \ (B), \ 2 \ (Q), \ 3 \ (S)$$

### getting control over B rather than P fluctuations is important!!!

$$\implies -\frac{\chi_{BS}}{\chi_S} \ge 0.193(22)$$

### Ratios of second order cumulants at $\mu_B = 0$ – observables at the LHC ? –



$$\chi^{BQ}_{11}/\chi^{BS}_{11}=-0.368(24)$$

large deviations from PDGHRG still 25% deviations from QMHRG2020 sensitive test of QCD



 $\chi^{BS}_{11}/\chi^{QS}_{11}=0.623(40)$ 

at ALICE freeze-out temperature == QCD crossover temperature  $T_{fo} = 156.5(1.5) \mathrm{MeV}$ 

#### Ratios of 4<sup>th</sup> and 2<sup>nd</sup> order cumulants

- large deviations from Skellam -





- ratios of 4<sup>th</sup> and 2<sup>nd</sup> order cumulants differ from non-inter. HRG for T>145 MeV
- they change by ~(20-40)% in the crossover region



#### Conclusions

ES-II range



### What we learned so far about the CEP in QCD from lattice QCD calculations:

I) the critical temperature is below  $T_c = 132^{+3}_{-6}~{
m MeV}$ 

II) the corresponding critical chemical potential is likely to be above 400 MeV

 Taylor expansions need to be resummed in order to reach CEP

– no CEP for  $\mu_B/T \leq 2.5$ 

- CEP not in the BES-II range (in collider mode)
- EoS (pressure & number density) well controlled for

 $\mu_B/T \leq 2.0 \; orall T > 135 \; {
m MeV}$  (larger range for higher T)

– reliable  $\mu_B$  - range is smaller for higher order cumulants, given only an 8<sup>th</sup> order Taylor series for the pressure