Pseudo-Nambu-Goldstone Dark Matter from Non-Abelian Gauge Symmetry

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Collaborators:

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Naoki Yamatsu (National Taiwan U.)

Based on Phys. Rev. D 106 (2022) 11, 115033 [2210.08696]

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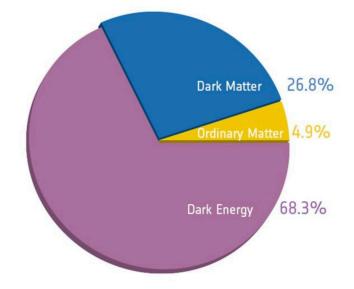
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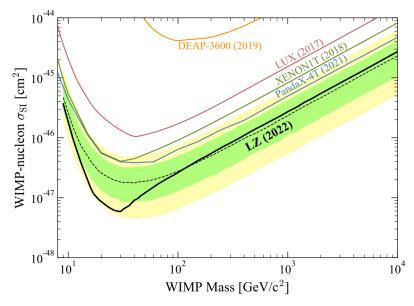
C.W. Chang, K. Tsumura, **YU**, N Yamatsu, "Pseudo-Nambu-Goldstone Dark Matter in SU(7) Grand Unification" [2311.13753]

- Unknown matter (Dark Matter) accounts for 26.8% of the total in the universe
- An attractive candidate for a DM is ...

- thermally produced in the early universe
- severely constrained by the direct detection



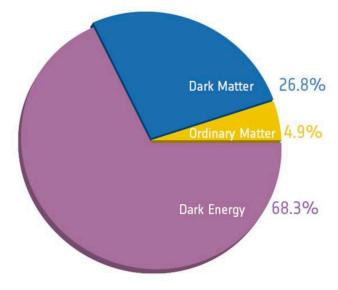
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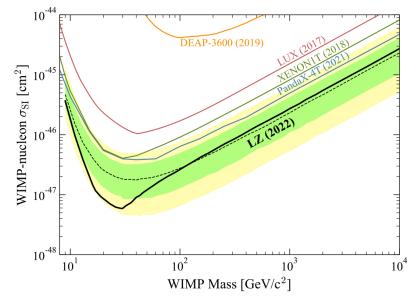
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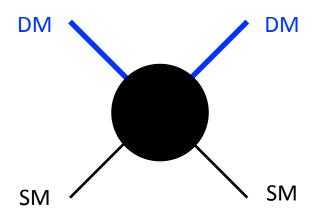
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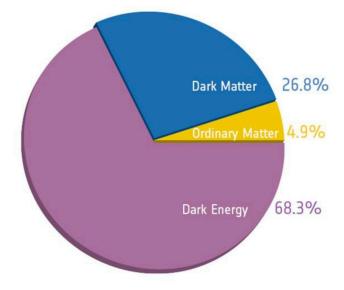


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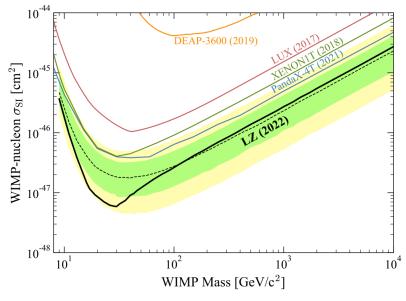
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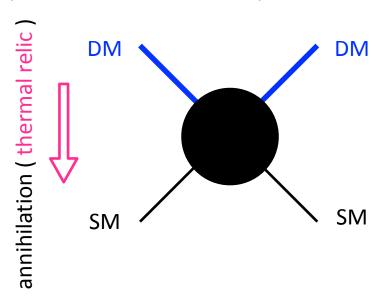
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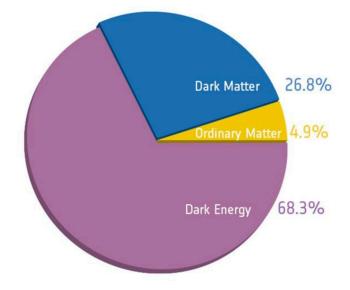


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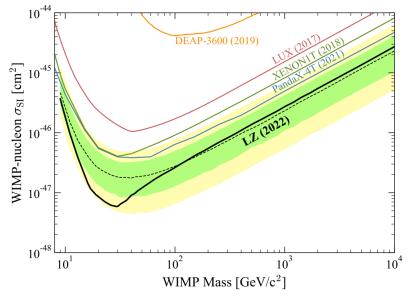
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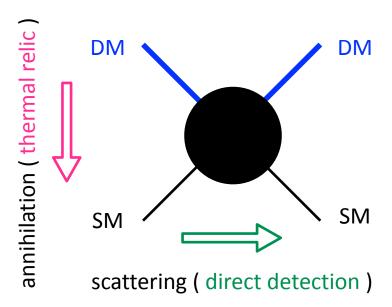
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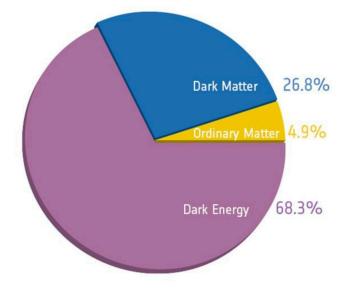


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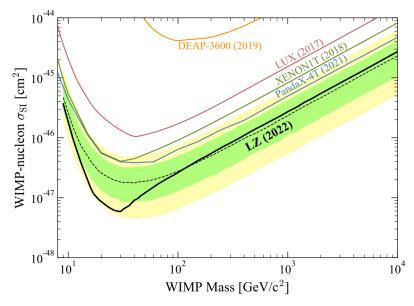
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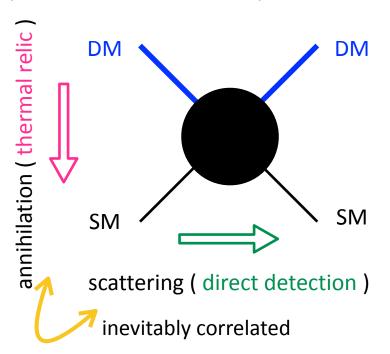
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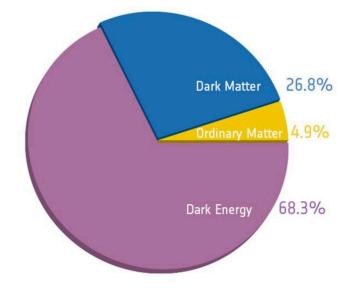


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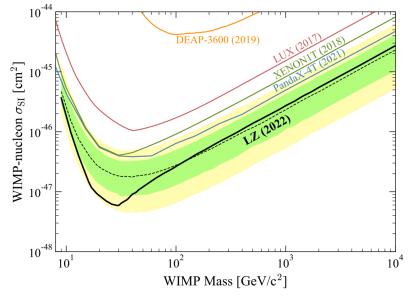
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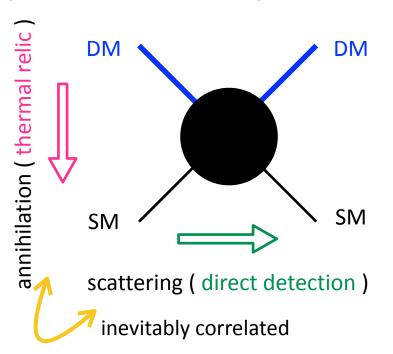


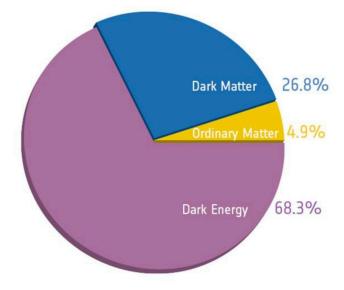
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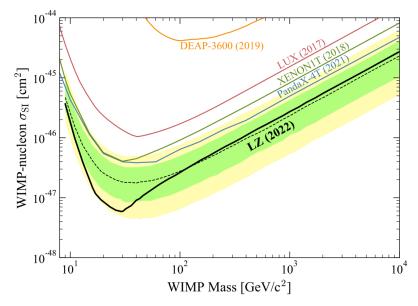
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Why is WIMP so constrained by direct detection?





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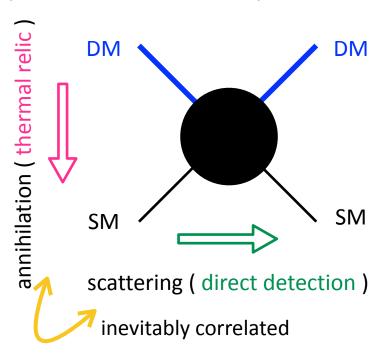
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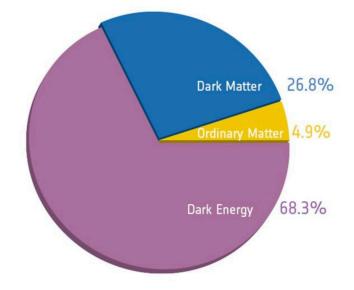
Stronger DM-SM interaction helps DM to stay longer in thermal bath, leading to $~\Omega h^2 \simeq 0.12~$, but also increases DM-nucleon scattering.

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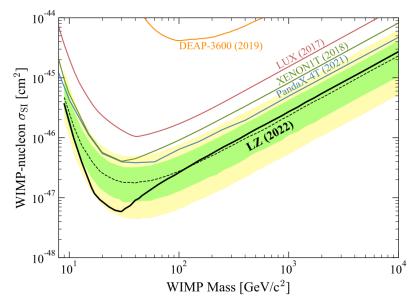
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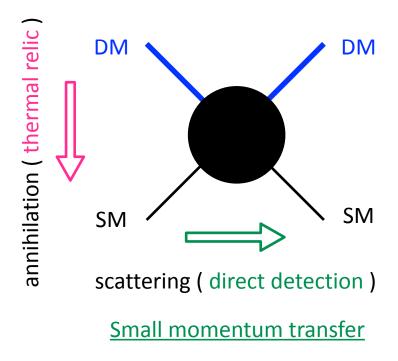
To realize viable WIMP model, we must address this dilemma.

What type of WIMP model would solve this dilemma?

C. Gross, O. Lebedev, and T, Toma Phys. Rev. Lett. 119 (2017) 19, 191801, [1708.02253]

pseudo Nambu-Goldstone Boson Dark Matter (pNGB-DM)

DM communicates with SM particles via derivative interaction



symmetry :
$$G_{\rm SM} \times U(1)_{\rm global}$$
 new fields : complex $S \in \mathbf{1}_0$
$$V(H,S) = -\frac{\mu_H^2}{2}|H|^2 - \frac{\mu_S^2}{2}|S|^2 + \frac{\lambda_H}{2}|H|^4 + \lambda_{HS}|H|^2|S|^2 + \frac{\lambda_S}{2}|S|^4$$
 Origin for pNGB mass $-\frac{\mu_S'^2}{4}S^2 + {\rm h.c.}$

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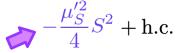
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Problems: $\mu_S'''S^3$, $\mu_S'''|S|^2S$, ... are dropped by hands

Solutions: gauged $U(1)_{B-L}$ model

Y. Abe, T. Toma, and K. Tsumura, JHEP 05 (2020) 057, [2001.03954]

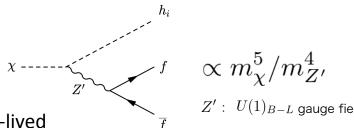
Y. Abe, T. Toma, K. Tsumura, and N. Yamatsu, Phys.Rev.D 104 (2021) 3, 035011 [2104.13523]

 $v_{B-L} \simeq 10^{15} {\rm GeV}$

" hierarchy problem "

pNGB DM decays

Higher $U(1)_{B-L}$ breaking scale required to make DM long-lived



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Hint: custodial symmetry

$$V_{\mathrm{SM}}(H) = -\mu_H^2 H^\dagger H + \lambda (H^\dagger H)^2$$
 $H = \frac{1}{\sqrt{2}} \begin{pmatrix} \phi_1 + i\phi_2 \\ \phi_3 + i\phi_4 \end{pmatrix}$ $G_{\mathrm{SM}} = SU(2)_{L} \times U(1)_{Y}$ invariant

symmetry : $G_{\mathrm{SM}} imes U(1)_{\mathrm{global}}$ new fields : complex $S \in \mathbf{1}_0$ $V(H,S) = -\frac{\mu_H^2}{2}|H|^2 - \frac{\mu_S^2}{2}|S|^2 + \frac{\lambda_H}{2}|H|^4 + \lambda_{HS}|H|^2|S|^2 + \frac{\lambda_S}{2}|S|^4$ Origin for pNGB mass $-\frac{\mu_S'^2}{4}S^2 + \mathrm{h.c.}$

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Accidental global symmetry : $G_{
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 W_{μ}^{\pm} and Z_{μ} form a $SU(2)_V$ triplet.

Accidental global symmetry : $G_{
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Our Model



UV completion of pNGB-DM model & predict stable DM



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We consider $G_{\mathrm{SM}} imes SU(2)_D^{\mathrm{gauge}}$ symmetry and introduce $\Phi \in \mathbf{2}$, $\Delta \in \mathbf{3}$ under $SU(2)_D^{\mathrm{gauge}}$ $\Sigma = \left(\tilde{\Phi}, \Phi \right)$

	$SU(2)_L$	$U(1)_Y$	$SU(2)_D^{ m gauge}$
H	2	1/2	1
Φ	1	0	2
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$$V(H,\Phi,\Delta)$$
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$$= -\mu_H^2 H^{\dagger} H - \frac{1}{2} \mu_{\Phi}^2 \text{Tr} \left[\Sigma^{\dagger} \Sigma \right] - \frac{1}{2} \mu_{\Delta}^2 \text{Tr} \left[\Delta^2 \right]$$

Mass terms

$$+ \, \lambda_H \left(H^\dagger H \right)^2 + \frac{\lambda_\Phi}{4} \left(\mathrm{Tr} \left[\Sigma^\dagger \Sigma \right] \right)^2 + \frac{\lambda_\Delta}{4} \left(\mathrm{Tr} \left[\Delta^2 \right] \right)^2$$

4-point self-int.

$$+ \lambda_{H\Phi} \left(H^{\dagger} H \right) \operatorname{Tr} \left[\Sigma^{\dagger} \Sigma \right] + \lambda_{H\Delta} \left(H^{\dagger} H \right) \operatorname{Tr} \left[\Delta^{2} \right] + \frac{\lambda_{\Phi\Delta}}{2} \operatorname{Tr} \left[\Sigma^{\dagger} \Sigma \right] \operatorname{Tr} \left[\Delta^{2} \right]$$

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$$-\sqrt{2}\kappa \mathrm{Tr}\left[\sigma_3 \Sigma^\dagger \Delta \Sigma\right]$$

Invariant under

global "Dark custodial symmetry"

$$\Delta \;
ightarrow \; U_{m L}^{
m dark} \Delta \; \; U_{m L}^{
m dark \, \dagger} \; ($$
 $^{
m dark \, }$ $^{
m dark \, }$

$$\Sigma \;
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$$-\sqrt{2}\kappa {
m Tr} \left[\sigma_3 \Sigma^\dagger \Delta \Sigma
ight]$$

Explicitly breaks

global "Dark custodial symmetry"

$$\Delta \;
ightarrow \; U_{m L}^{
m dark} \Delta \; U_{m L}^{
m dark} ^{\dagger}$$
 ($H
ightarrow H$)

$$\Sigma \rightarrow U_L^{ ext{dark}} \Sigma U_R^{ ext{dark}\dagger}$$

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$$V(H,\Phi,\Delta)$$
 $\Sigma=(ilde{\Phi},\Phi)$

$$= -\mu_H^2 H^{\dagger} H - \frac{1}{2} \mu_{\Phi}^2 \text{Tr} \left[\Sigma^{\dagger} \Sigma \right] - \frac{1}{2} \mu_{\Delta}^2 \text{Tr} \left[\Delta^2 \right]$$

$$+ \lambda_H \left(H^{\dagger} H \right)^2 + \frac{\lambda_{\Phi}}{4} \left(\text{Tr} \left[\Sigma^{\dagger} \Sigma \right] \right)^2 + \frac{\lambda_{\Delta}}{4} \left(\text{Tr} \left[\Delta^2 \right] \right)^2$$

$$+ \lambda_{H\Phi} \left(H^{\dagger} H \right) \operatorname{Tr} \left[\Sigma^{\dagger} \Sigma \right] + \lambda_{H\Delta} \left(H^{\dagger} H \right) \operatorname{Tr} \left[\Delta^{2} \right] + \frac{\lambda_{\Phi\Delta}}{2} \operatorname{Tr} \left[\Sigma^{\dagger} \Sigma \right] \operatorname{Tr} \left[\Delta^{2} \right]$$

$$-\sqrt{2}\kappa {
m Tr} \left[\sigma_3 \Sigma^\dagger \Delta \Sigma
ight]$$

Explicitly breaks

global "Dark custodial symmetry"

Even after $\ \langle \Phi \rangle \neq 0 \ \ \& \ \ \langle \Delta \rangle \neq 0$, the exact $\ U(1)_{
m global}$ remains unbroken

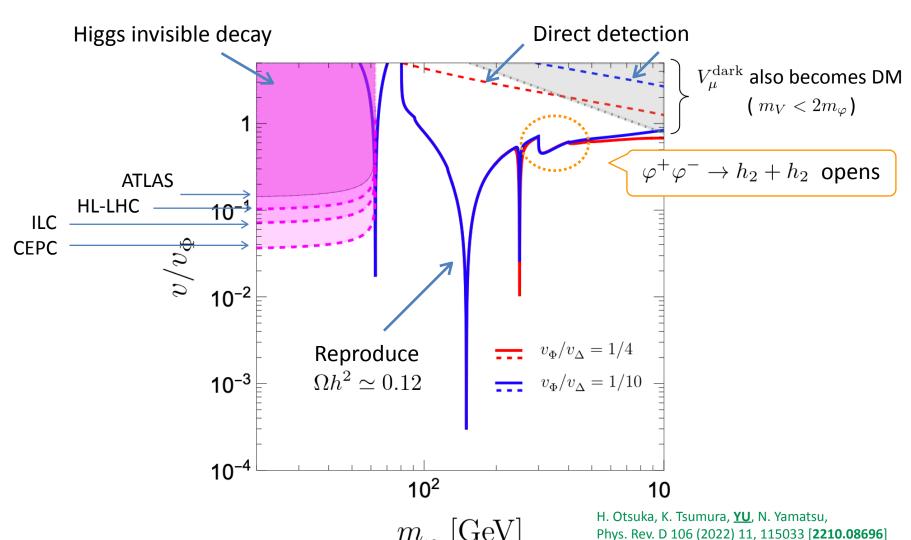


Benchmark

scalar mass: $(m_{h_1}, m_{h_2}, m_{h_3}) = (125, 300, 500) \,\text{GeV}$

mixing angle : $(\sin \alpha_x, \sin \alpha_y, \sin \alpha_z) = (0.06, 0.05, 0.1)$

$$\begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \alpha_x & \sin \alpha_x \\ 0 & -\sin \alpha_x & \cos \alpha_x \end{pmatrix} \begin{pmatrix} \cos \alpha_y & 0 & \sin \alpha_y \\ 0 & 1 & 0 \\ -\sin \alpha_y & 0 & \cos \alpha_y \end{pmatrix} \begin{pmatrix} \cos \alpha_z & \sin \alpha_z & 0 \\ -\sin \alpha_z & \cos \alpha_z & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} h \\ \phi_3 \\ \eta_3 \end{pmatrix}$$



Summary

- In the original abelian pNGB DM model, particular soft-breaking terms are included, and their origins are not addressed.
- UV completed models are proposed, but all of them predict decaying DM. In order to make DM long-lived, we must introduce large hierarchy in symmetry breaking scales.
- We construct pNGB-DM model with non-abelian gauge symmetry. Unbroken dark custodial symmetry ensure stability of pNGB DM. We don't need to introduce large hierarchy.

Back Up

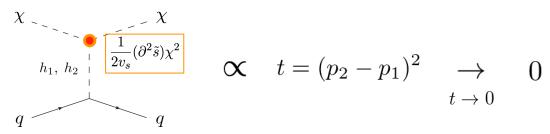
But, this is not the end of the story ...

Three-point breaking term may spoil the cancellation

Two-point breaking term

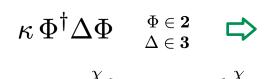
T. Abe and Y. Hamada, [2205.11919]

$$\mu_\chi^2 \left(\phi^\dagger T^3 \phi \right)$$
 \Longrightarrow Origin of pNGBs' mass

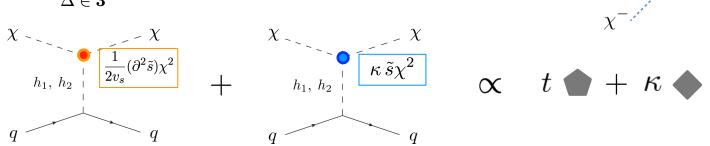


• Three-point breaking term

Our Model



Origin of pNGBs' mass & interactions



We must make sure *DM-nucleon scattering is suppressed enough*

Soft breaking terms

Soft-breaking = Quadratic

$$V_{\rm SM}(H) + \lambda_{HS}|H|^2|S|^2 - \frac{\mu_S^2}{2}|S|^2 + \frac{\lambda_S}{2}|S|^4 - \left(\frac{\mu_S'^2}{4}S^2 + \text{h.c.}\right)$$

C. Gross, O. Lebedev, and T, Toma, Phys. Rev. Lett. 119 (2017) 19, 191801, [1708.02253]

Soft-breaking = Quadratic + tadpole

$$V_{\mathrm{SM}}(H) + \lambda_{HS}|H|^2|S|^2 - rac{\mu_S^2}{2}|S|^2 + rac{\lambda_S}{2}|S|^4 - \left(rac{\mu_S'^2}{4}S^2 + \mathrm{h.c.}
ight) + \left(aS + \mathrm{h.c.}
ight)$$

V. Barger, P. Langacker, M. McCaskey, M. Ramsey-Musolf and G. Shaughnessy, Phys. Rev. D 79 (2009), 015018, [0811.0393] G. C. Cho, C. Idegawa and E. Senaha, Phys. Lett. B 823 (2021), 136787, [2105.11830]

$$\begin{array}{c} \chi \\ \hline \\ h_1, \ h_2 \end{array} \rangle \propto \left\{ \left(-\frac{m_{h_1}^2}{t - m_{h_1}^2} + \frac{m_{h_2}^2}{t - m_{h_2}^2} \right) \right. \simeq 0 \quad \text{@ } t \to 0 \\ + \left. \frac{\sqrt{2}a}{v_S} \left(-\frac{1}{t - m_{h_1}^2} + \frac{1}{t - m_{h_2}^2} \right) \right. \rangle \\ \simeq 0 \quad \text{@ } m_{h_1} \simeq m_{h_2} \end{array}$$

Based on Idegawa-san's slide