



南京理工大学  
NANJING UNIVERSITY OF SCIENCE & TECHNOLOGY

# **Higgs** as a portal and **BSM Exotic Higgs** decays: Testing EWPT

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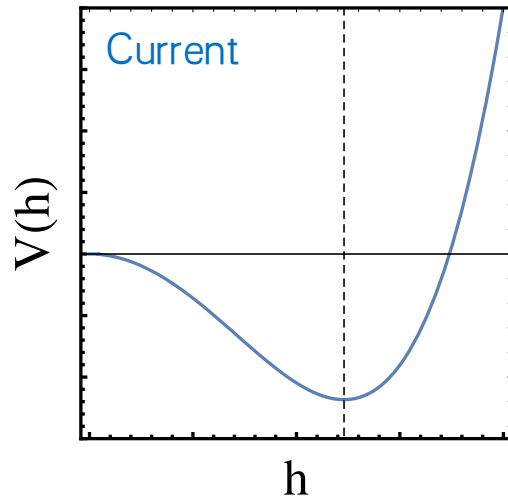
Nanjing University of Science and Technology

JHEP 04(2021) 015, Phys.Rev.D 105 (2022) 11, 115040

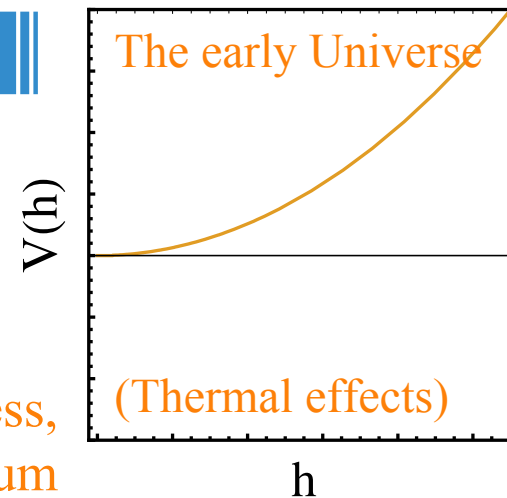
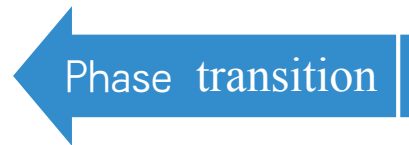
Higgs 2023

# Phase transition in electroweak theory

EW symmetry restoration in the early Universe



$W$  &  $Z$  bosons are massive;  
Photon is massless,  
Mexican-hat like



$SU(2)_L$  &  $U(1)_Y$  bosons are massless,  
True vacuum

# What is the pattern of EW phase transition (PT)?

It could be –

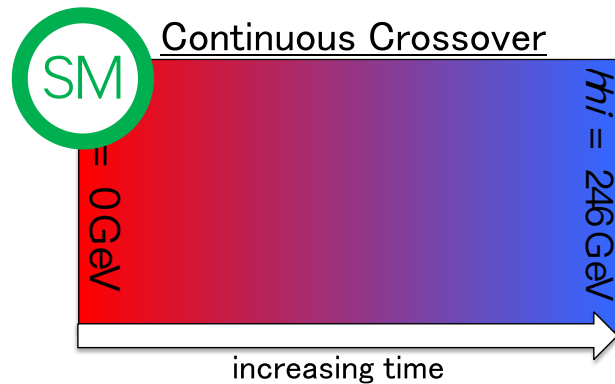
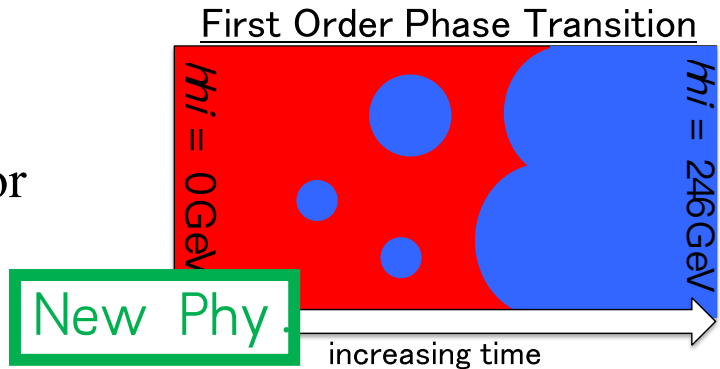


Figure from L.-T. Wang's talk in IHEP workshop

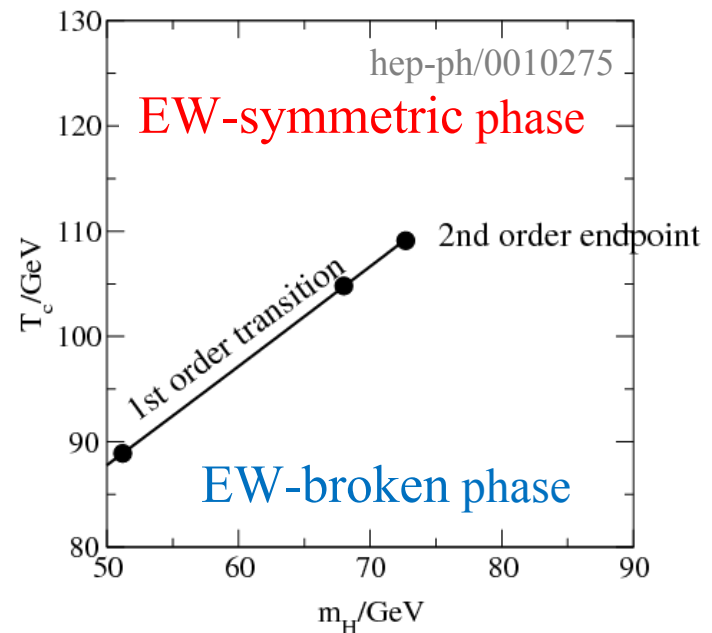
or



Lattice calculation shows the phase diagram ==>

Thus in the SM it is a crossover, since  $M_h = 125 \text{ GeV} > 75 \text{ GeV}$ ;

However, a 1<sup>st</sup>-order EWPT is more interesting.  
(Needs **new physics**)

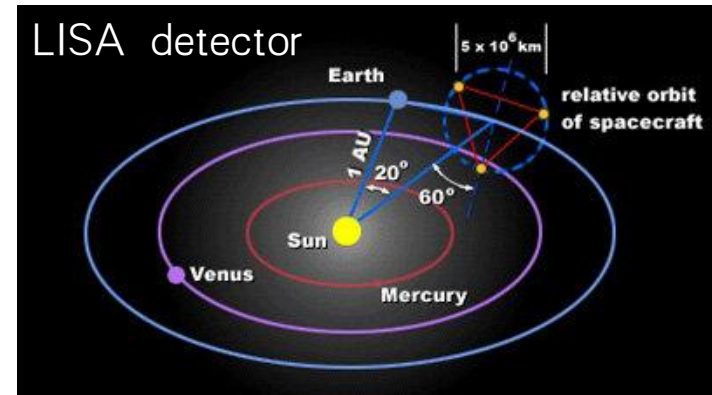


# Why is a 1<sup>st</sup>-order EWPT interesting?

- It's the essential ingredient of the **EW baryogenesis**, to explain **Baryon Asymmetry**
- Acting as the background of very rich **dark matter** mechanisms
- Sources of the stochastic GWs:

- Collision of the bubbles
- Sound waves in plasma
- Turbulence in plasma

EWPT GWs typically peak in mHz.



# How to achieve a 1<sup>st</sup>-order EWPT?

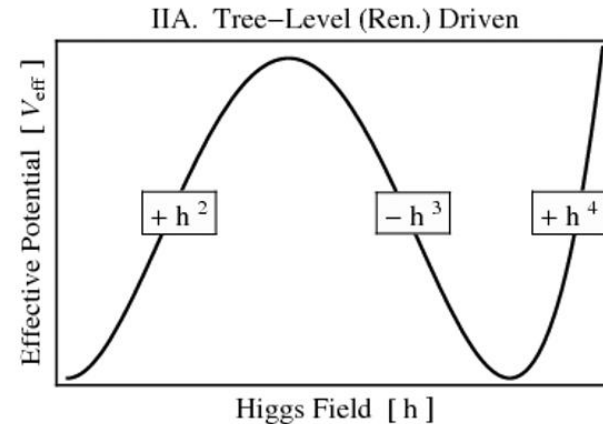
Adding a barrier for the Higgs potential via new physics!

The decay between two vacua separated by a barrier.

The VEV of the Higgs field *jumps*.

Getting a barrier via the help of additional scalar field(s):

- **SM + real singlet (xSM);**
- 2HDM;
- Georgi-Machacek model;
- .....



We choose the **xSM** as the benchmark model.

- It's simple, but has captured the most important feature of EWPT;
- It can be treated as the prototype of many new physics EWPT models.

## EWPT in the xSM (SM + real singlet)

We choose the **xSM** as the benchmark model.

It's simple, but has captured the most important feature of EWPT.

The scalar potential of the xSM

$$V = -\mu^2 |H|^2 + \lambda |H|^4 + \frac{a_1}{2} |H|^2 S + \frac{a_2}{2} |H|^2 S^2 \\ + b_1 S + \frac{b_2}{2} S^2 + \frac{b_3}{3} S^3 + \frac{b_4}{4} S^4$$

8 input parameters:

1 unphysical, 2 fixed by Higgs mass & VEV; 5 *free* parameters.

Expansion around the VEV

$$H = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v + h \end{pmatrix}, \quad S = v_s + s,$$

$$h_1 = s \cos \theta + h \sin \theta,$$

$$h_2 = -s \sin \theta + h \cos \theta,$$

Mass eigenstates & the mixing angle.

$h_1$  is **singlet-like**,  $h_2$  is **SM-like**

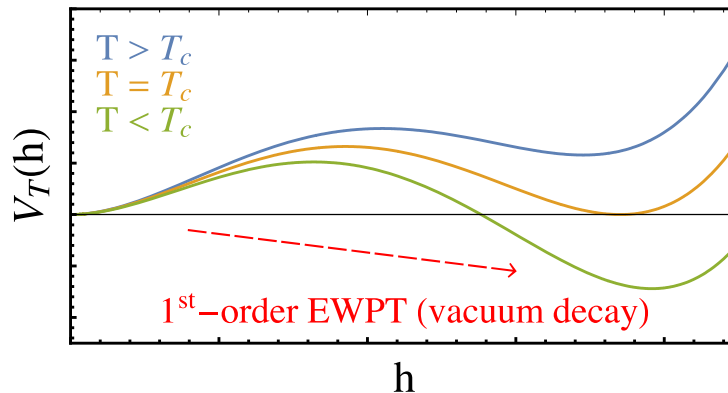
# 1<sup>st</sup>-order EWPT in the xSM

At finite temperature:

$$V = -(\mu^2 - c_H T^2)|H|^2 + \lambda|H|^4 + \frac{a_1}{2}|H|^2 S + \frac{a_2}{2}|H|^2 S^2 \\ + (b_1 + m_1 T^2)S + \frac{b_2 + c_S T^2}{2}S^2 + \frac{b_3}{3}S^3 + \frac{b_4}{4}S^4$$

$$c_H = \frac{3g^2 + g'^2}{16} + \frac{y_t^2}{4} + \frac{\lambda}{2} + \frac{a_2}{24}, \quad c_S = \frac{a_2}{6} + \frac{b_4}{4}, \quad m_1 = \frac{a_1 + b_3}{12}$$

An Illustration --

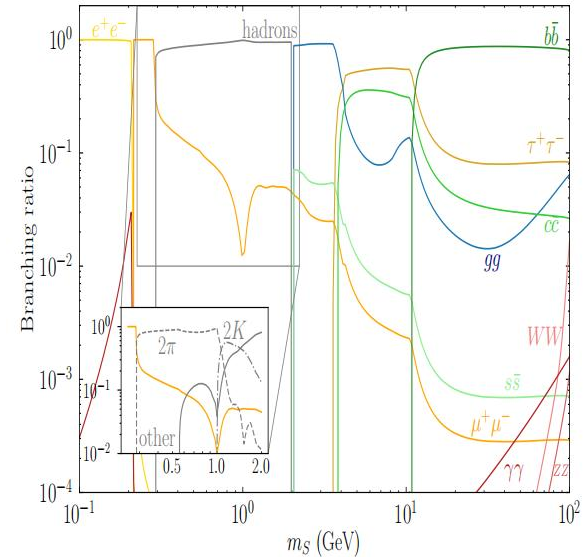
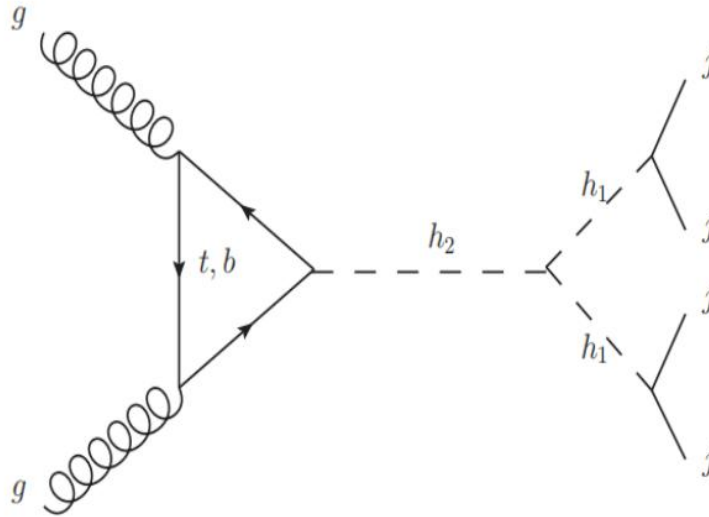


Question:

Can collider experiments probe the 1<sup>st</sup>-order EWPT parameter space?

# Higgs Exotic Decays

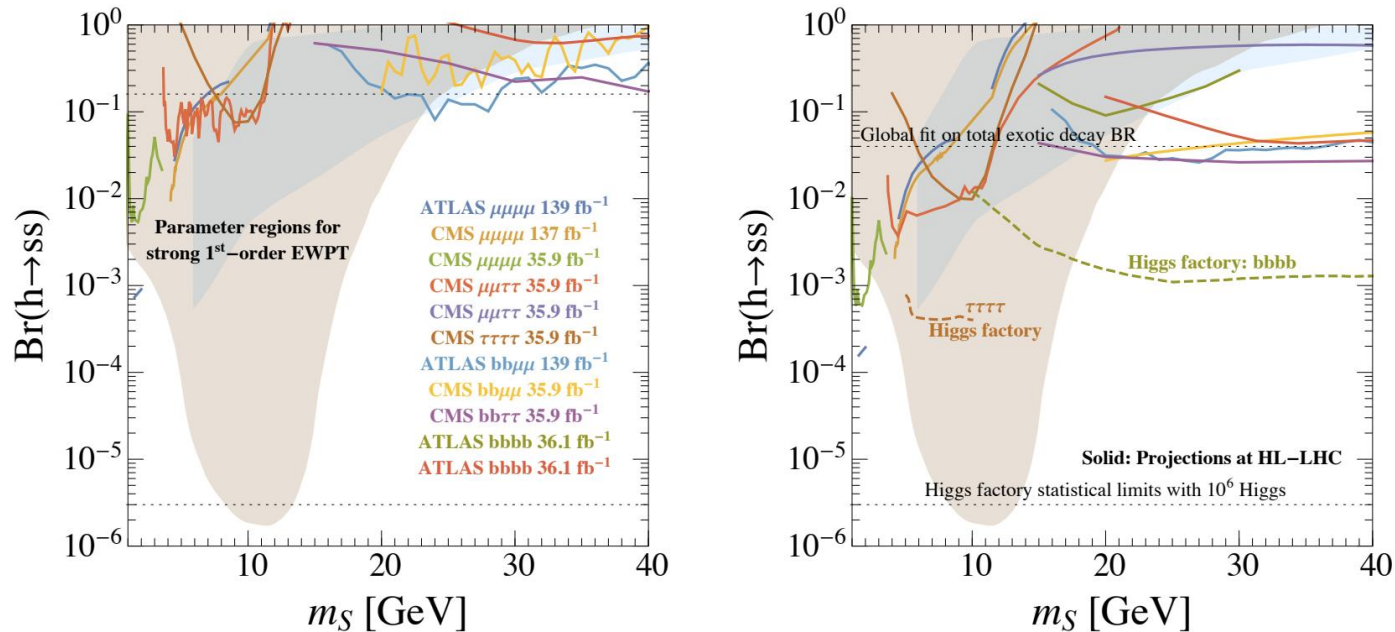
$$gg \rightarrow h_{2(SM)} \rightarrow h_1 h_1 \rightarrow 4j$$



1st-order EWPT leads to **large  $BR(h_2 \rightarrow h_1 h_1)$** .  
 $h_1$  decays into jets dominantly.



# Current Bounds



- BSM Higgs Exotic Decays can be bounded by **Prompt** Searches at Colliders
- Mapped into the Parameter Space of the **1<sup>st</sup>-order EWPT**
- **1<sup>st</sup>-order EWPT** with low  $Br(h_2 \rightarrow h_1 h_1)$  can not be reached even for Higgs factory
- **Can we reach even lower  $Br(h_2 \rightarrow h_1 h_1)$ ?**

Carena, Kozaczuk, Liu, Ou, J.Ramsey-Musolf et al, 2203.08206

Kozaczuk, J.Ramsey-Musolf and Shelton, 1911.10210

Carena, Liu, Wang, 1911.10206

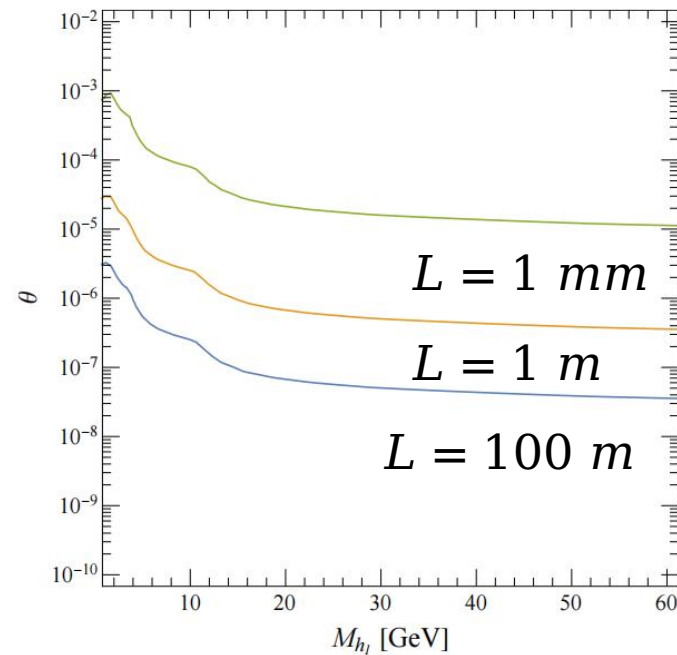
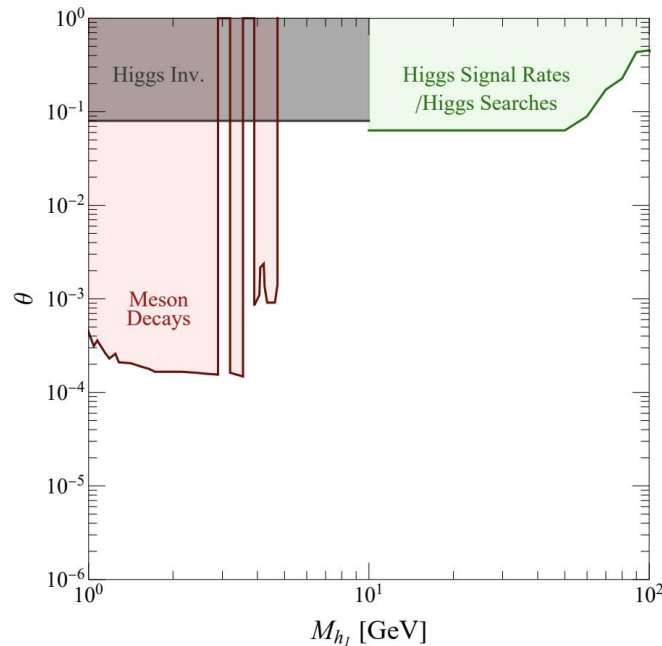
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[illegible]

- **Light weakly-coupled** particles as LLPs, are strongly motivated, including the **light scalar** responsible for 1st-order EWPT in the xSM.

# Current Limits on Higgs Mixing

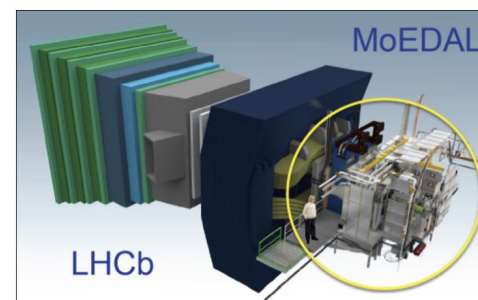
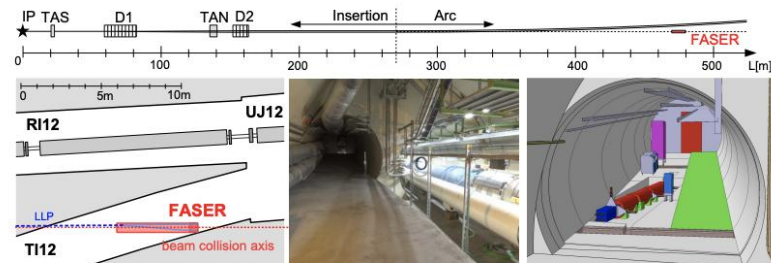
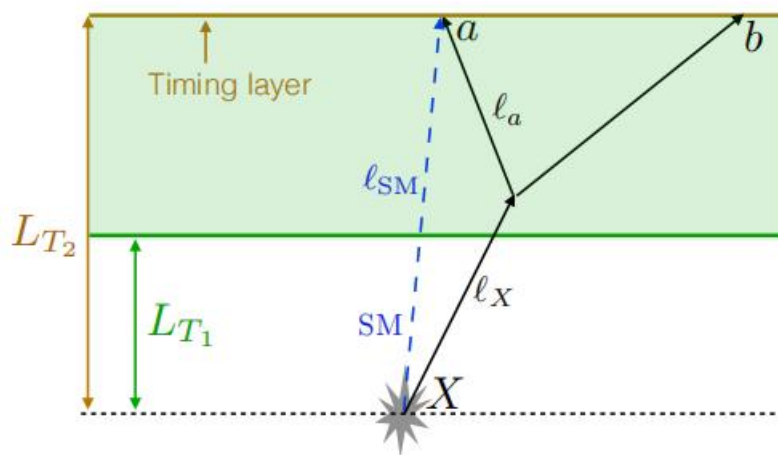
The current limits on  $(M_{h_1}, \theta)$



- For  $M_{h_1} < 10$  GeV, the current limits from rare meson decays at the LHCb, leads to  $h_1$  as a **long-lived particle** (LLP).

# Detectors for LLPs

CMS-Timing, FASER, MoEDAL-MAPP are to be operated at Run 3.

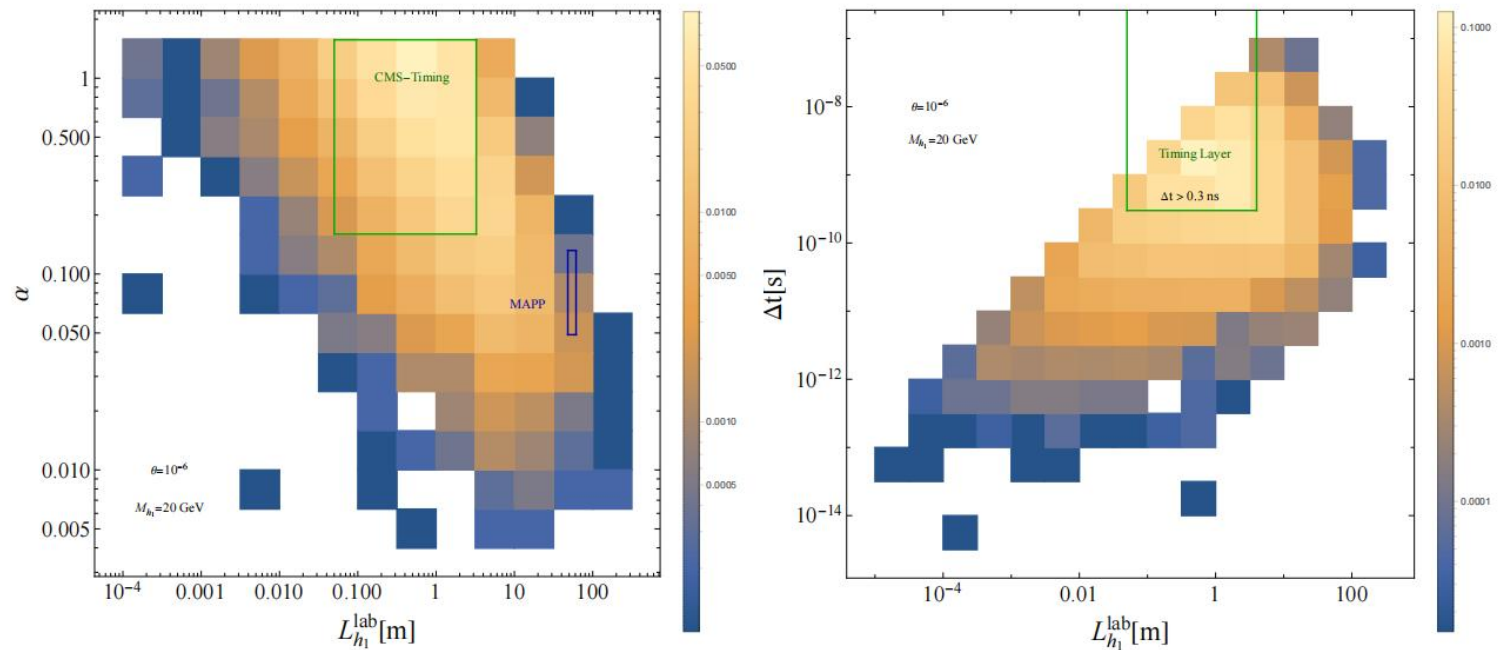


Many others, e.g. MATHUSLA and CODEX-b are in discussions.

CMS-Timing detector using the **time-delayed leptons/jets** as signals, while the other detectors using displaced vertex.

# Detector efficiency

Detector efficiency is a function of geometrical coverage,



and resolution in time for timing detector.

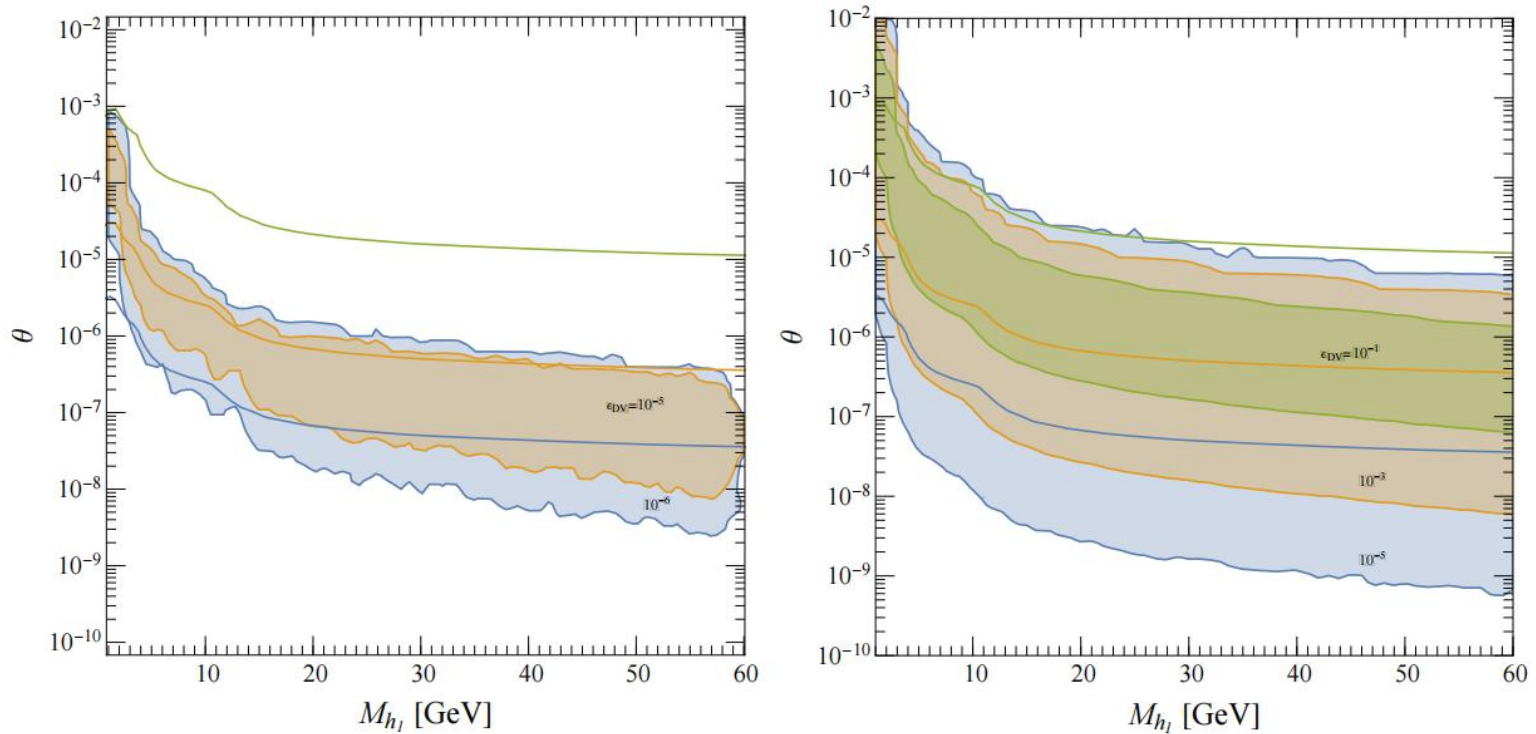
CMS-Timing has large coverage, and good resolution.

MAPP has small coverage,

while negligible for FASER.

# Detector efficiency

Detector efficiency is a function of geometrical coverage,



CMS-Timing has efficiency up to  $10^{-1}$ .

MAPP has  $10^{-4}$ ,

while negligible for FASER.

# 1<sup>st</sup>-order EWPT parameter space

From J. Kozaczuk, M. Ramsey-Musolf, J. Shelton, 1911.10210

The Higgs mixing is small

Approximate  **$Z_2$  symmetry**

EWPT can happen as **two-step transition**

$$(h = 0, s \simeq 0) \rightarrow (h = 0, s \neq 0) \rightarrow (h \neq 0, s \simeq 0),$$

$$a_2 \gtrsim \frac{m_{h_1}^2}{4v^2} \frac{\Delta}{1 - \Delta},$$
$$|b_3| > \sqrt{\frac{9}{4} b_4 (2m_{h_1}^2 - a_2 v^2 + 2T_{\text{EW}}^2 \beta)},$$
$$b_4 \gtrsim \frac{m_{h_1}^4 \Delta}{4\lambda v^4 (1 - \Delta)},$$

# Connects to 1st-order EWPT

Connection between the number of events and 1st-order EWPT

$$N_{signal}$$

$$= \sigma_{pp \rightarrow h_2} \times L \times BR_{h_2 \rightarrow h_1 h_1}(a_2, M_{h_1}) \\ \times BR_{h_1 \rightarrow jj}^2(M_{h_1}) \times \epsilon_{kin}(M_{h_1}) \times \epsilon_{geo}(M_{h_1}, \theta)$$

$$BR_{h_2 \rightarrow h_1 h_1}(a_2, M_{h_1}) = \frac{\Gamma_{h_2 \rightarrow h_1 h_1}}{\Gamma_{h_2}^{SM} + \Gamma_{h_2 \rightarrow h_1 h_1}}$$

$$\Gamma_{h_2 \rightarrow h_1 h_1} \propto (a_2 v)^2$$

- **LLP** events are sensitive to  $|H|^2 S^2$  couplings to trigger **1st-order EWPT**.

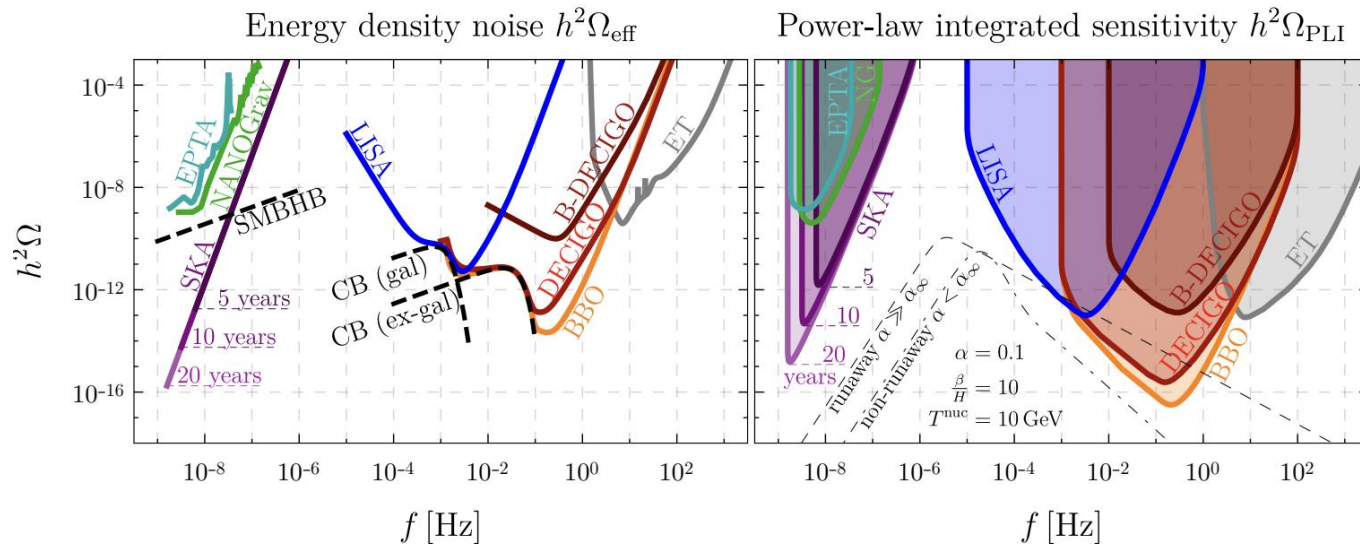
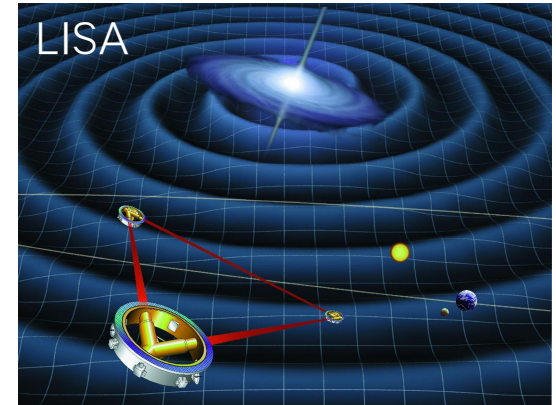


# GW Signal

**1<sup>st</sup>-order EWPT** can lead to **Primordial GW signal**

For the LISA detector, signal-to-noise ratio (SNR):

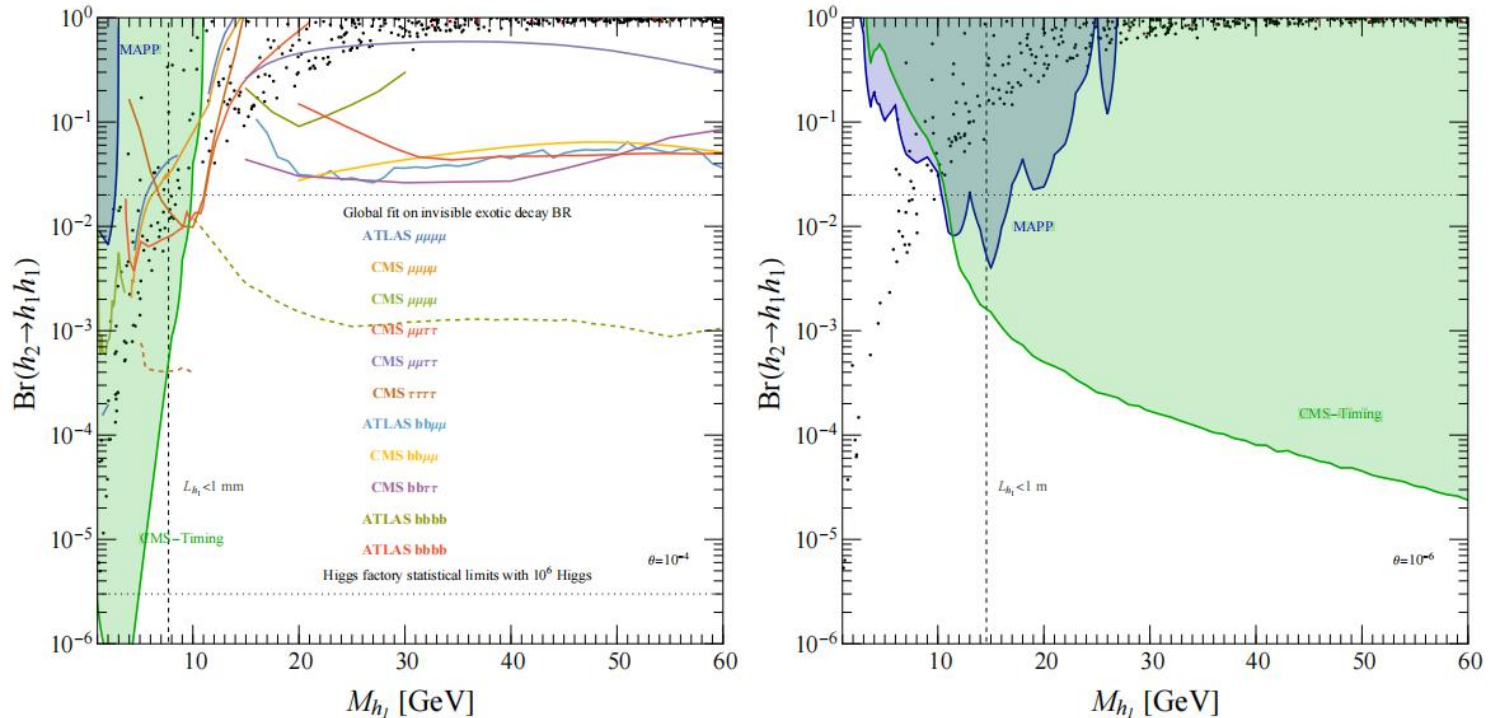
$$\text{SNR} = \sqrt{\mathcal{T} \int_{f_{\min}}^{f_{\max}} df \left( \frac{\Omega_{\text{GW}}(f)}{\Omega_{\text{LISA}}(f)} \right)^2}$$



Breitbach, Kopp, Madge, Opferkuch, Schwaller, 1811.11175

# Sensitivity

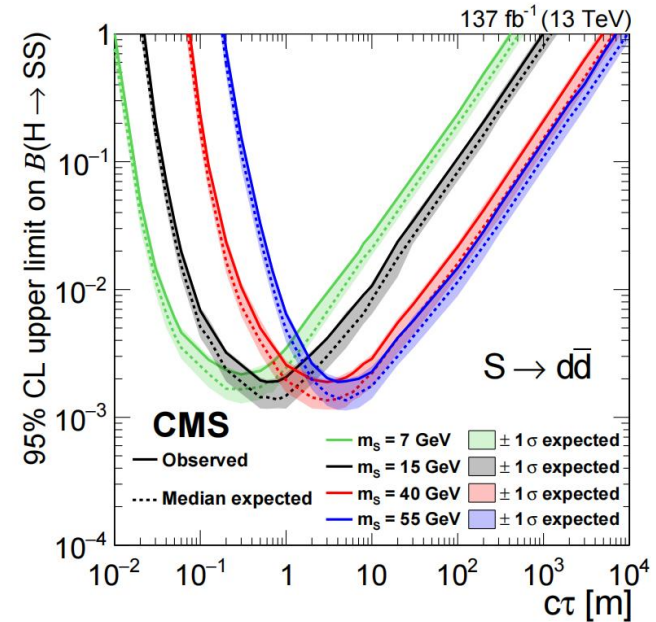
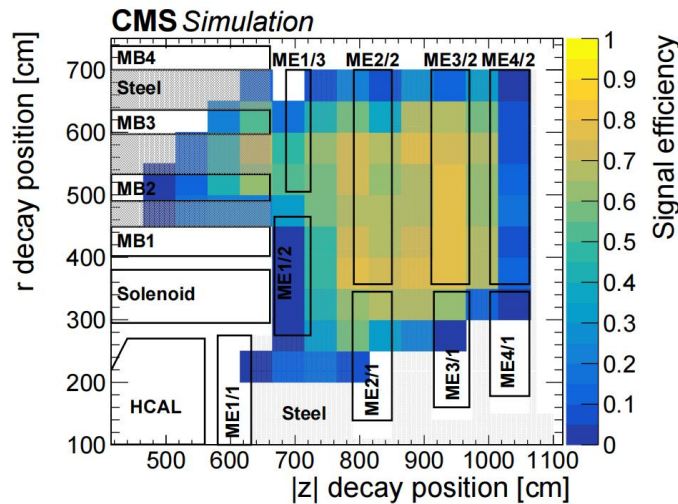
Fixed  $\theta = 10^{-4}$  (left),  $10^{-6}$  (right).



- No FASER sensitivity, No Enough Higgs from Meson decay
- **CMS-Timing** can probe large parameter space where the searches for **promptly Higgs exotic decays** can not reach.
- **MAPP** can only probe small parameter space
- No **GW signals!**
- **W.L., A.Y. and H. S., Phys.Rev.D 105 (2022) 11, 115040**

# An Existing Search at the CMS

CMS-EXO-20-015



- Search for LLPs in the CMS endcap muon detectors
- Can be **recast** to test **1<sup>st</sup>-order EWPT**, **to be done**

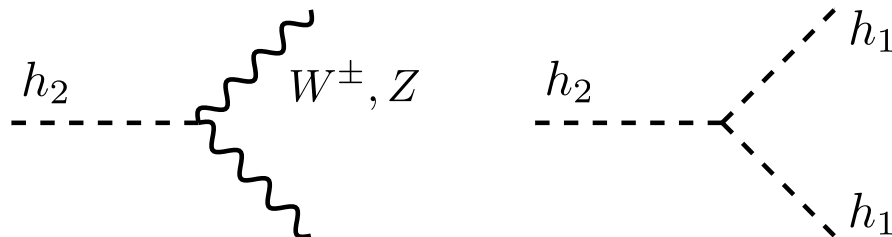
CMS Collaboration, 2107.04838

# BSM Scalar Decays

## Feature of the xSM

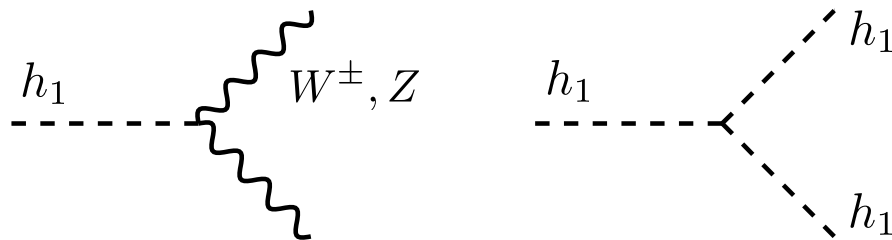
Two neutral scalars:  $h_1$  (Higgs-like) and  $h_2$  (singlet-like, TeV), with mixing angle  $\theta$ ;

$$\begin{aligned} g_{h_2 V V} &= g_{h V V}^{\text{SM}} \sin \theta \\ g_{h_2 f \bar{f}} &= g_{h f \bar{f}}^{\text{SM}} \sin \theta \\ \lambda_{h_2 h_1 h_1} &\propto \sin \theta \end{aligned}$$



**Direct searches** at the  $pp$  colliders

$$\begin{aligned} g_{h_1 V V} &= g_{h V V}^{\text{SM}} \cos \theta \\ g_{h_1 f \bar{f}} &= g_{h f \bar{f}}^{\text{SM}} \cos \theta \\ \lambda_{h_1 h_1 h_1} &= \lambda_{h h h}^{\text{SM}} f(\theta) \end{aligned}$$



**Indirect searches** at the  $e^+e^-$  colliders

# Muon collider!

## Precision and Energy Frontier!

A high-energy muon collider is able to execute both the

- **direct search**
- **indirect search**

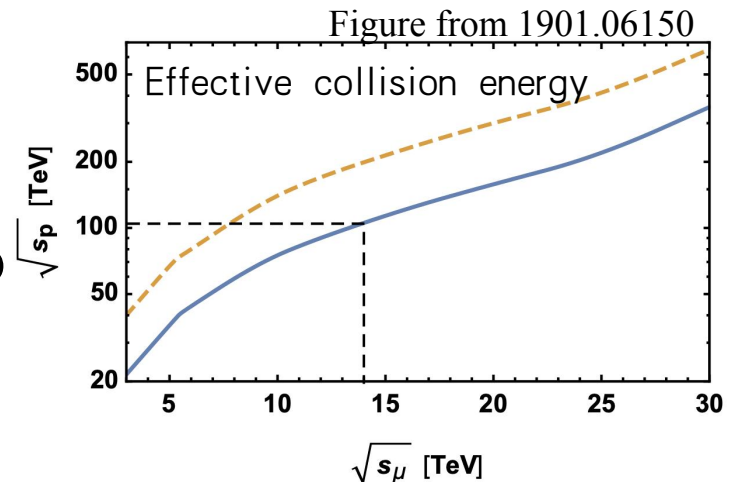
strategies for EWPT in xSM!

Compared to the  $e^+e^-$  machine:

- Synchrotron radiation is **suppressed by  $10^9$**  since  $M_\mu \gg M_e$ , hence the collision energy can reach O(10) TeV;
- Also **very clean**, as long as the beam-induced-background is controllable (main challenge).

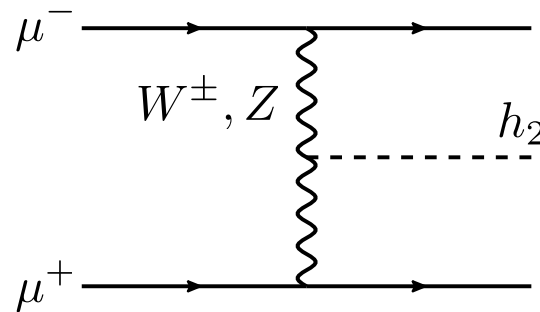
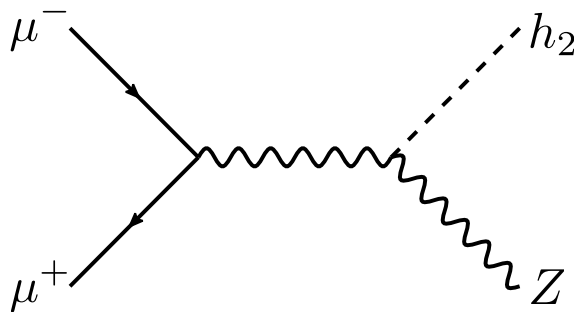
Compared to the  $pp$  machine:

- The **entire collision energy** can be used to probe hard process;
- Much **cleaner** due to the small QCD background.



# Muon collider: direct search

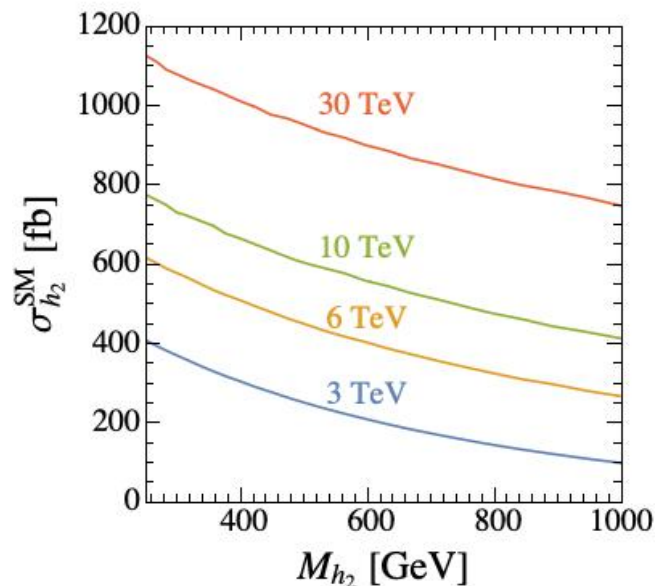
Producing the  $h_2$  at a muon collider



$Zh_2$  associated production & Vector Boson Fusion (VBF).

At a multi-TeV collider, the dominant channel is VBF, in which  $W^+W^-$  fusion dominates (90%);

$\sigma^{\text{SM}}(h_2)$ : rate obtained by assuming a Higgs-like coupling for the  $h_2$ .

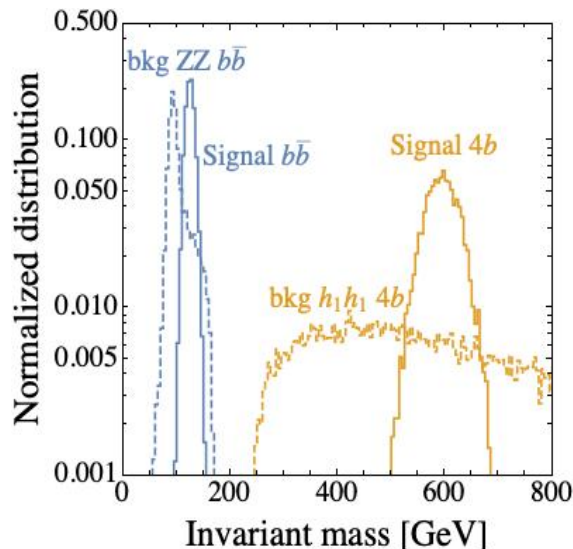


# Muon collider: direct search

The  $h_2 \rightarrow h_1 h_1 \rightarrow bbbb$  channel:

Main background:

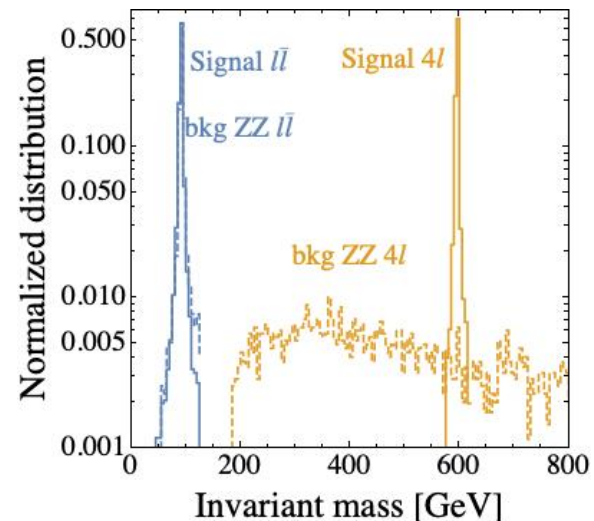
- Vector Boson Scattering  $ZZ \rightarrow bbbb$
- $h_1 h_1 \rightarrow bbbb$ .



The  $h_2 \rightarrow ZZ \rightarrow l^+ l^- l^+ l^-$  channel:

Main background:

- Vector Boson Scattering  $ZZ \rightarrow l^+ l^- l^+ l^-$ .



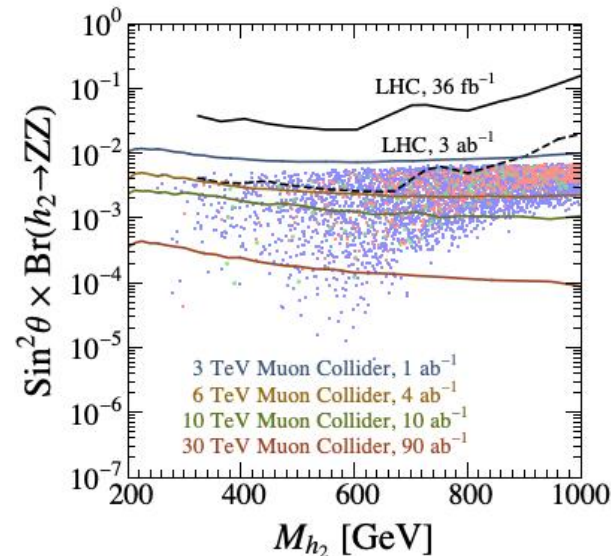
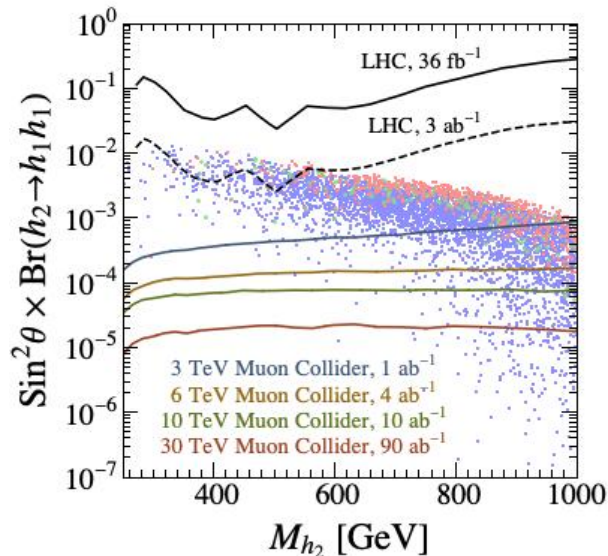
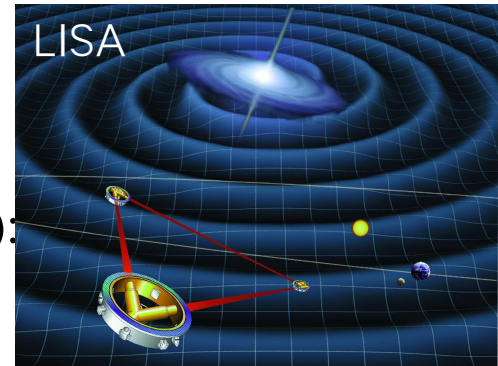


# Muon collider: direct search

The collider search and gravitational wave detection are complementary!

For the LISA detector, signal-to-noise ratio (SNR):

$$\text{SNR} = \sqrt{\mathcal{T} \int_{f_{\min}}^{f_{\max}} df \left( \frac{\Omega_{\text{GW}}(f)}{\Omega_{\text{LISA}}(f)} \right)^2}$$



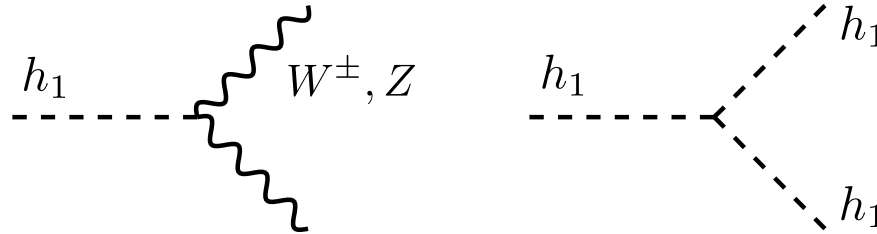
The diHiggs & diboson channels are complementary as well

**W.L., K.X., JHEP 04(2021) 015**



# Muon collider: indirect search

The gauge boson coupling & triple Higgs coupling. Making use of the results in [Han, Liu, Low and Wang, 2008.12204]:



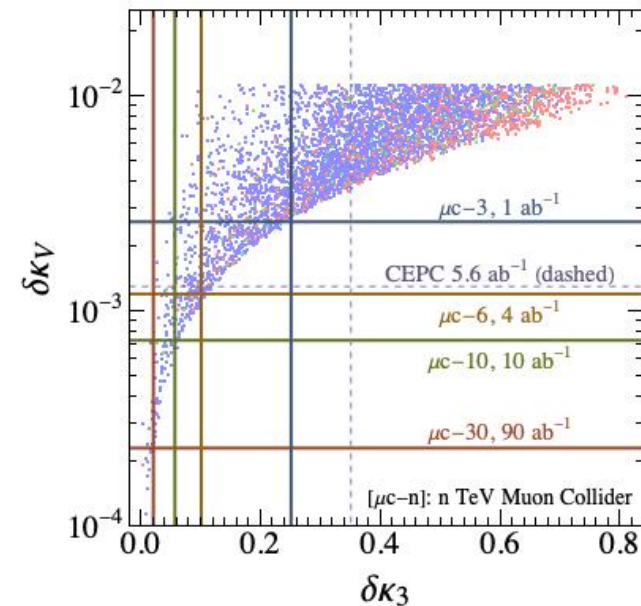
Defining deviations

$$\delta\kappa_V = \left| \frac{g_{h_1 VV}}{g_{h_1 VV}^{\text{SM}}} - 1 \right|,$$

$$\delta\kappa_3 = \frac{\lambda_{h_1 h_1 h_1}}{\lambda_{h_1 h_1 h_1}^{\text{SM}}} - 1$$

We can obtain the projections.

**W.L., K.X., JHEP 04(2021) 015**



# Conclusion

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**1<sup>st</sup>-order EW phase transition** is interesting:

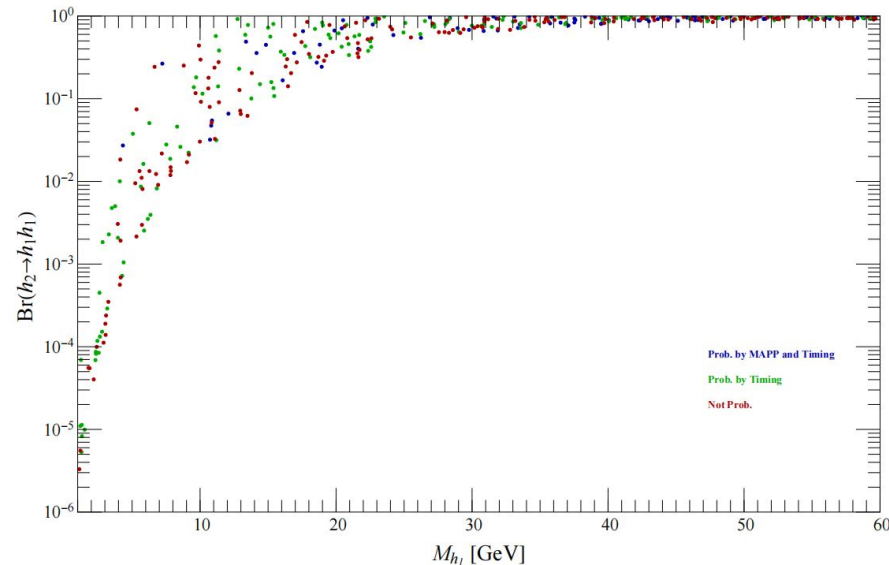
- Theoretically, it is the essential ingredient of EW baryogenesis, and can trigger very rich dark matter mechanisms;
- Experimentally, it yields detectable gravitational waves.

For **light, weakly coupled scalar** corresponding to **1<sup>st</sup>-order EWPT**, we search for **Higgs Exotic decays to LLP signatures**, but **no detectable gravitational waves**.

For **heavy, strongly coupled scalar** corresponding to **1<sup>st</sup>-order EWPT**, we search for **BSM Scalar Decays to direct and indirect signals** at the muon colliders, **and complementary GW signals**.

# Signatures at Colliders

Running  $\theta$



Green points are probed by CMS-Timing, but not by MAPP. CMS-Timing can probe **a lot more** 1st-order EWPT points. There are still appreciable points not probed by any of LLP detectors.

# Muon collider: direct search

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Main background:

✓ Vector Boson Scattering  $ZZ \rightarrow bbbb$  ( $llll$ ) and  $h_1 h_1 \rightarrow bbbb$ .

Kinematic Cuts:

**Cut I:**  $p_T > 30 \text{ GeV}$ ,  $|\eta| < 2.43$ ,  $M_{recoil} > 200 \text{ GeV}$ , (Cut I)

**Cut II:** minimizing  $\chi^2 = (M_{12} - M_h)^2 + (M_{34} - M_h)^2$

$|M_{12} - M_h| < 15(10) \text{ GeV}$ ,  $|M_{34} - M_h| < 15(10) \text{ GeV}$

**Cut III:**  $|M_{1234} - M_{h_2}| < 30(20) \text{ GeV}$ ,

$\Delta E/E = 10\%$ ,  $\epsilon_{b-tag} = 70\%$

# Muon collider: direct search

Decay of  $h_2$  to SM particles ( $X$  = vector boson or fermion)

$$\Gamma(h_2 \rightarrow XX) = \sin^2 \theta \times \Gamma^{\text{SM}}(h_2 \rightarrow XX),$$

$$\Gamma(h_2 \rightarrow h_1 h_1) \propto \lambda_{h_2 h_1 h_1}^2$$

Dominant channels: di-boson ( $W^+W^-$ ,  $ZZ$ ),  $tt$ , and  $h_1 h_1$ .

The  $h_1 h_1$  channel can reach a branching ratio of 80%;

For heavy  $h_2$ , the  $VV$  channel dominates;

We choose

- $h_2 \rightarrow ZZ \rightarrow l^+l^-l^+l^-$
  - $h_2 \rightarrow h_1 h_1 \rightarrow bbbb$
- for a detailed simulation.

