

Distinguishing different axion models with low energy couplings

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[KC, Im, Seong and Kim, JHEP 08 \(2021\) 058 \(arXiv:2106.05816\)](#)

Axions or axion-like particles (ALPs) are light pseudo-scalar bosons postulated in many well-motivated models for BSM physics:

Strong CP problem,

Dark matter,

Inflation,

Baryogenesis,

String theory, ...

Axions have a finite field range (periodic):

$$a(x) \cong a(x) + 2\pi f_a \quad (f_a = \text{axion decay constant})$$

Axions can be naturally light as their interactions are constrained by an approximate PQ-symmetry non-linearly realized in low energy effective theory:

$$U(1)_{\text{PQ}} : \quad a(x) \rightarrow a(x) + \text{constant}$$

Classification of axions by the origin of the field variable

Field-theoretic axions from the phase of complex scalar fields:

$$\sigma(x) = \rho(x)e^{ia(x)/f_a} \quad \left(\frac{f_a}{\sqrt{2}} = \langle \rho \rangle \right)$$

PQWW, KSVZ, DFSZ, ...

String-theoretic axions from p-form gauge fields in string theory:
(or in any theory with extra-dim)

Witten '84, Svrcek & Witten '06, ...

$$a(x) = \int_{\Sigma_p} C_{[m_1, \dots, m_p]}^{(p)}(x, y) dy^{m_1} \dots dy^{m_p} \quad (\Sigma_p = p\text{-cycle in extra-dim})$$

or $\partial_\mu a = \epsilon_{\mu\nu\rho\sigma}(\partial^\nu B^{\rho\sigma} + \dots)$ ($B_{\mu\nu} = 2\text{-form gauge field in 4D spacetime}$)

Classification by low energy couplings

For model-independent discussion of axion couplings, it is convenient to use a field basis for which all fields other than the axion are PQ-invariant:

Georgi, Kaplan, Randall '85

$$U(1)_{\text{PQ}} : a(x) \rightarrow a(x) + \text{constant}, \quad \Phi = (\psi, H, \dots) \rightarrow \Phi$$

Axion couplings to the SM at high scales around f_a :

PQ-invariant derivative couplings to matter fields described by real-valued \mathbf{c}_ψ & \mathbf{c}_H

$$\mathcal{L}_{\text{eff}} = \frac{1}{2} f_a^2 (\partial_\mu \theta)^2 + \partial_\mu \theta \left(\sum_\psi c_\psi \bar{\psi} \sigma^\mu \psi + i c_H (H^\dagger D^\mu H - D^\mu H^\dagger H) \right) \quad (\text{PQ-invariant})$$

$$+ \sum_{F^A=G,W,B} c_A \frac{g_A^2}{32\pi^2} \theta(x) F^{A\mu\nu} \tilde{F}_{\mu\nu}^A - V_0(\theta(x)) - m_H^2(\theta(x)) |H|^2 + \dots \quad (\text{PQ-breaking})$$

Bare axion potential

Couplings to the SM gauge fields normalized in such a way that $\mathbf{c}_G, \mathbf{c}_W, \mathbf{c}_B$ are rational numbers or integers.

Non-derivative coupling to $|H|^2$ (axion-dependent Higgs mass)

$$\theta(x) = \frac{a(x)}{f_a} \cong \theta(x) + 2\pi \quad \Rightarrow \quad \text{Axion couplings} \propto \frac{1}{f_a}$$

QCD axion postulated to solve the strong CP problem

$U(1)_{\text{PQ}}$ is broken dominantly by the coupling to the gluon anomaly:

Peccei & Quinn

$$\frac{c_G}{32\pi^2} \frac{a(x)}{f_a} G^{\mu\nu} \tilde{G}_{\mu\nu}$$

PQ-breaking couplings other than the anomalous couplings to the SM gauge fields are highly suppressed as

$$\frac{\partial}{\partial\theta} \left(V_0(\theta), \frac{1}{16\pi^2} m_H^2(\theta) \Lambda^2, \dots \right) \lesssim 10^{-10} f_\pi^2 m_\pi^2$$

$$\rightarrow V_{\text{axion}}(\theta) = -\frac{f_\pi^2 m_\pi^2}{m_u + m_d} \sqrt{m_u^2 + m_d^2 + 2m_u m_d \cos(c_G \theta)} + \mathcal{O}(10^{-10} f_\pi^2 m_\pi^2)$$

P- and CP-invariant QCD: $|\theta_{\text{QCD}}| \equiv |c_G \langle \frac{a}{f_a} \rangle| \lesssim 10^{-10}$

$$m_a \simeq 5.7 c_G \left(\frac{10^{12} \text{GeV}}{f_a} \right) \mu\text{eV}$$

Ultra-light ALP which may constitute (part of) dark matter

(e.g. fuzzy dark matter with $m_a \sim 10^{-20} - 10^{-21} \text{ eV}$)

No coupling to the gluon anomaly ($\mathbf{c}_G = 0$) , but may have non-zero coupling to the EW anomalies

$$\frac{1}{32\pi^2} \frac{a(x)}{f_a} \left(c_B B^{\mu\nu} \tilde{B}_{\mu\nu} + c_W W^{\mu\nu} \tilde{W}_{\mu\nu} \right)$$

All other PQ-breakings are so tiny, yielding

$$m_a \ll m_{a\text{QCD}} \sim \frac{m_\pi f_\pi}{f_a}$$

Low energy axion couplings:

At $\mu = \mathcal{O}(1)$ GeV

$$\frac{1}{32\pi^2} \frac{a(x)}{f_a} \left(\overset{\text{photon}}{c_\gamma F^{\mu\nu} \tilde{F}_{\mu\nu}} + \overset{\text{gluons}}{c_G G^{\alpha\mu\nu} G_{\mu\nu}^\alpha} \right) + \sum_{\Psi=u,d,e} \overset{\text{light quarks and electron}}{\frac{\partial_\mu a}{2f_a} C_\Psi \bar{\Psi} \gamma^\mu \gamma_5 \Psi}$$

$$c_\gamma = c_W + c_B, \quad C_\Psi = C_\Psi^0 + \Delta C_\Psi$$

Rational numbers
which are typically
of $\mathcal{O}(1)$.

Tree level values

Radiative corrections

$$C_u^0 = c_{q_1}(f_a) + c_{u_1^c}(f_a) + c_H(f_a),$$

$$C_d^0 = c_{q_1}(f_a) + c_{d_1^c}(f_a) - c_H(f_a),$$

$$C_e^0 = c_{\ell_1}(f_a) + c_{e_1^c}(f_a) - c_H(f_a).$$

Measurable axion couplings in low energy experiments:

$$\frac{1}{2}g_{a\gamma}a(x)\vec{E}\cdot\vec{B} + \partial_\mu a(x)\left(\frac{g_{ap}}{2m_p}\bar{p}\gamma^\mu\gamma_5p + \frac{g_{an}}{2m_n}\bar{n}\gamma^\mu\gamma_5n + \frac{g_{ae}}{2m_e}\bar{e}\gamma^\mu\gamma_5e\right)$$

$$g_{a\gamma} \simeq \frac{\alpha_{\text{em}}}{2\pi} \frac{1}{f_a} \left(c_\gamma - \frac{2}{3} \left(\frac{m_u + 4m_d}{m_u + m_d} \right) c_G \right)$$

$$g_{ap} \simeq \frac{m_p}{f_a} \left(C_u\Delta u + C_d\Delta d - \left(\frac{m_d}{m_u + m_d}\Delta u + \frac{m_u}{m_u + m_d}\Delta d \right) c_G \right)$$

$$\simeq \frac{m_p}{f_a} \left(0.88C_u - 0.39C_d - 0.47c_G \right)$$

$$g_{an} \simeq \frac{m_n}{f_a} \left(C_d\Delta u + C_u\Delta d - \left(\frac{m_u}{m_u + m_d}\Delta u + \frac{m_d}{m_u + m_d}\Delta d \right) c_G \right)$$

$$\simeq \frac{m_n}{f_a} \left(-0.39C_u + 0.88C_d - 0.02c_G \right)$$

The parts depending on \mathbf{C}_G are mostly from the axion-meson mixing.

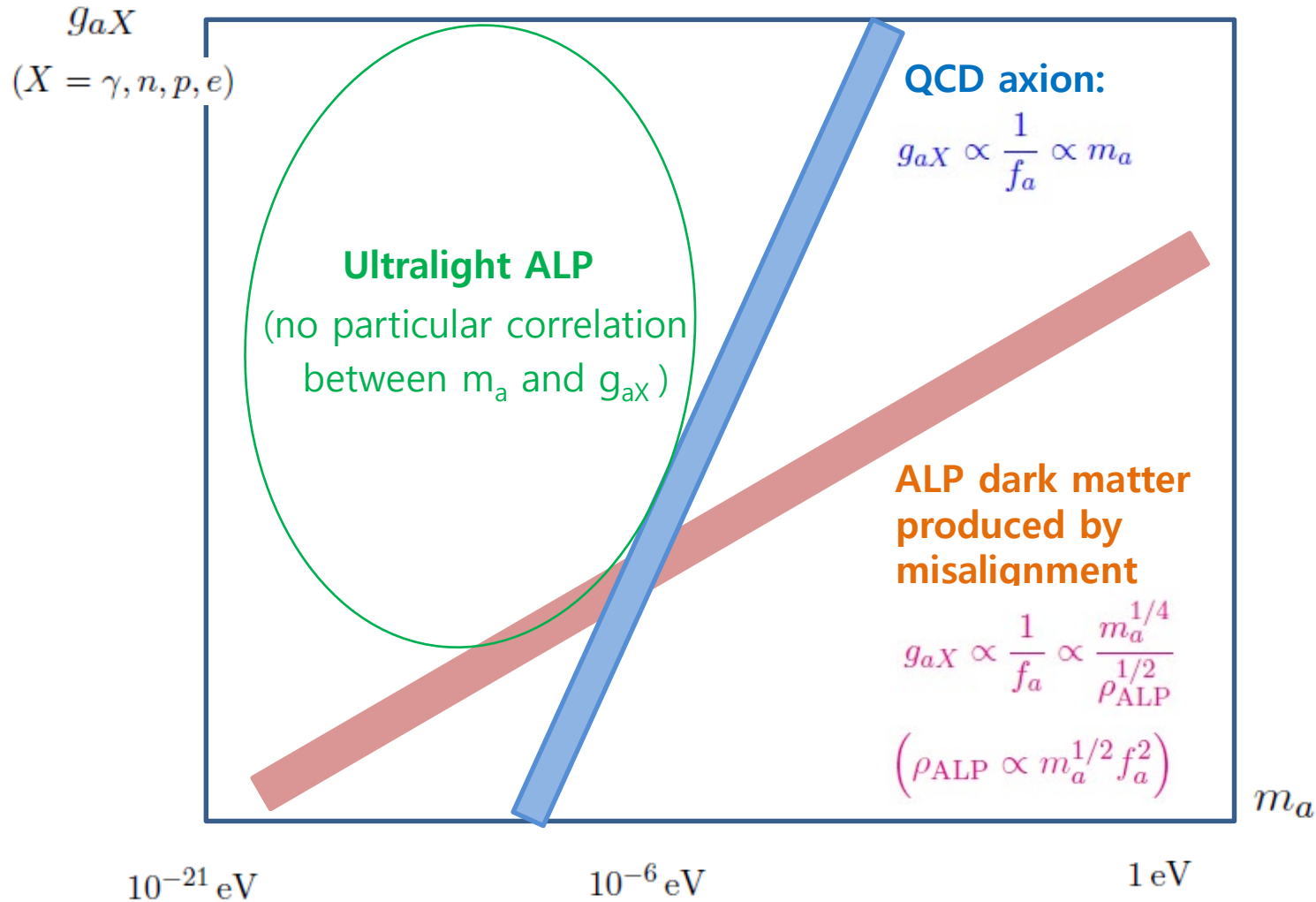
$$g_{ae} \simeq \frac{m_e}{f_a} C_e$$

$$\Delta u = 0.897(27), \quad \Delta d = -0.376(27) \quad \text{for } C_{u,d} \text{ at } \mu = 2 \text{ GeV}$$

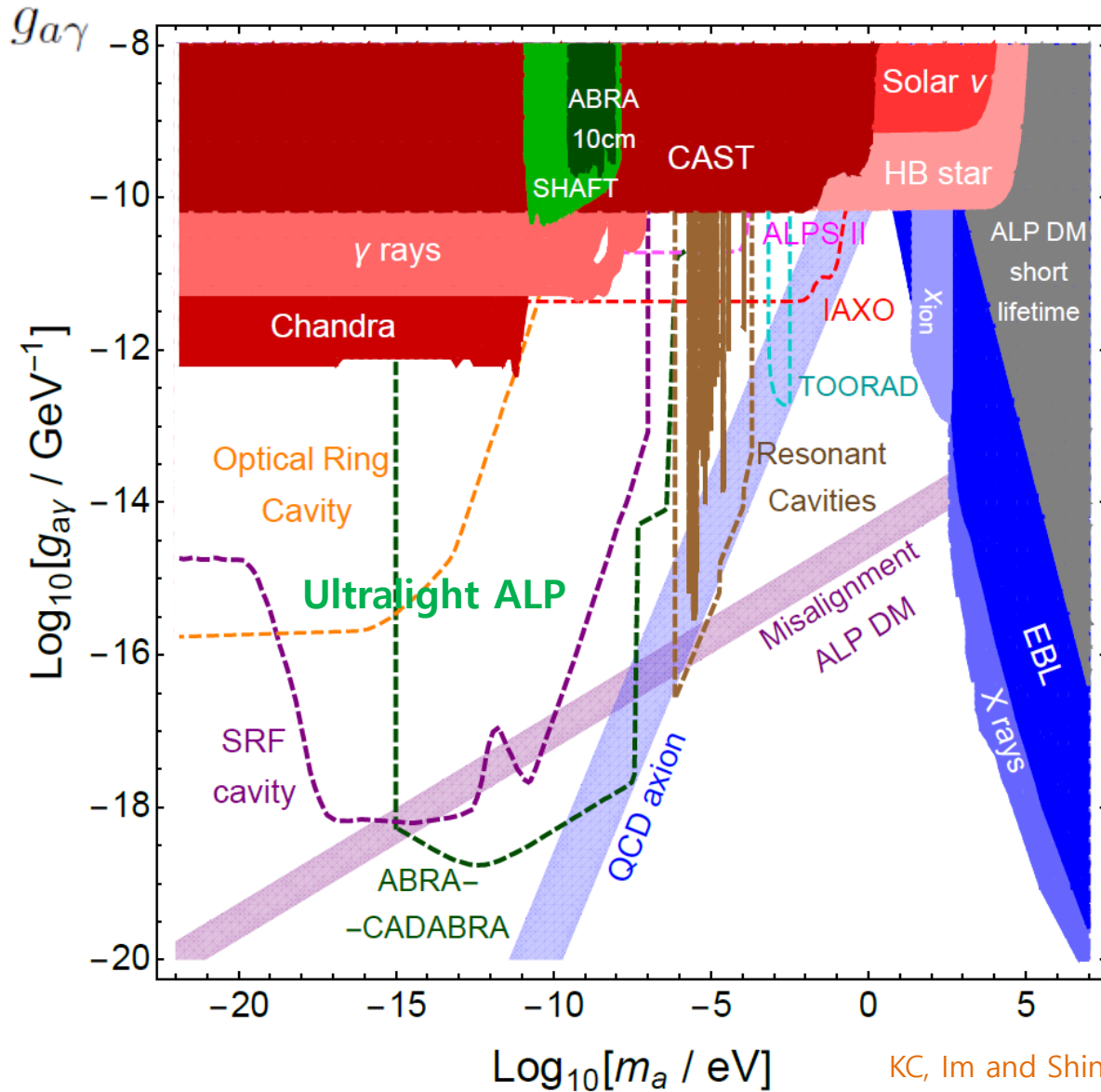
$$\left(s^\mu \Delta q = \langle p | \bar{q} \gamma^\mu \gamma_5 q | p \rangle \quad (q = u, d) \right)$$

Cortona et al, arXiv:1511.02867

Typical parameter space for theoretically well-motivated light axions which may have good potential to be experimentally detected:



Observational bounds and the sensitivities of planned experiments



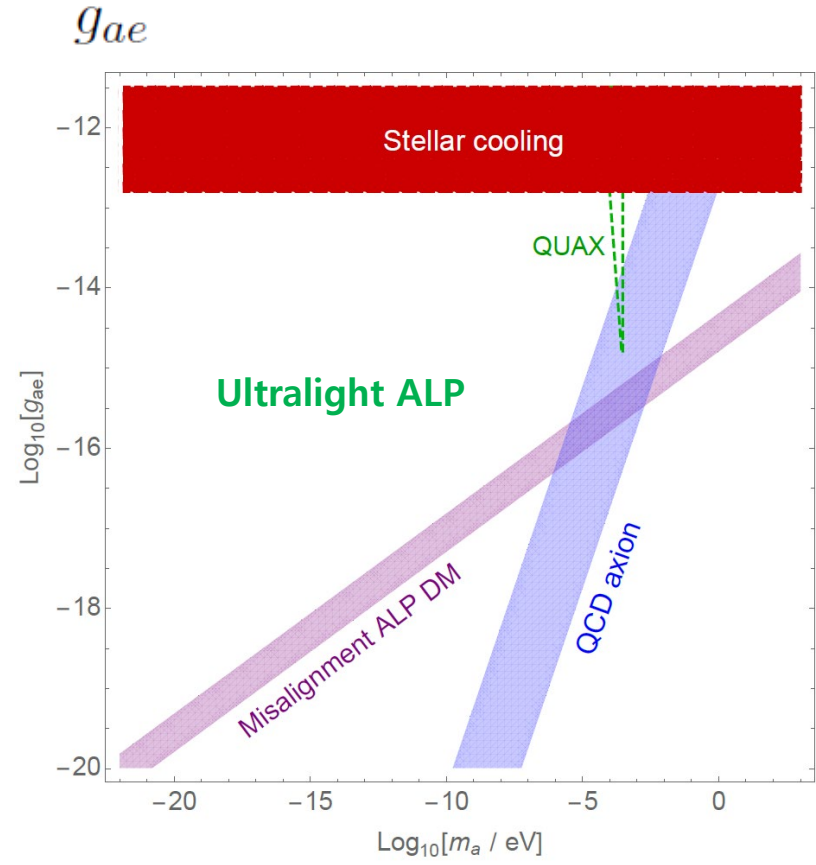
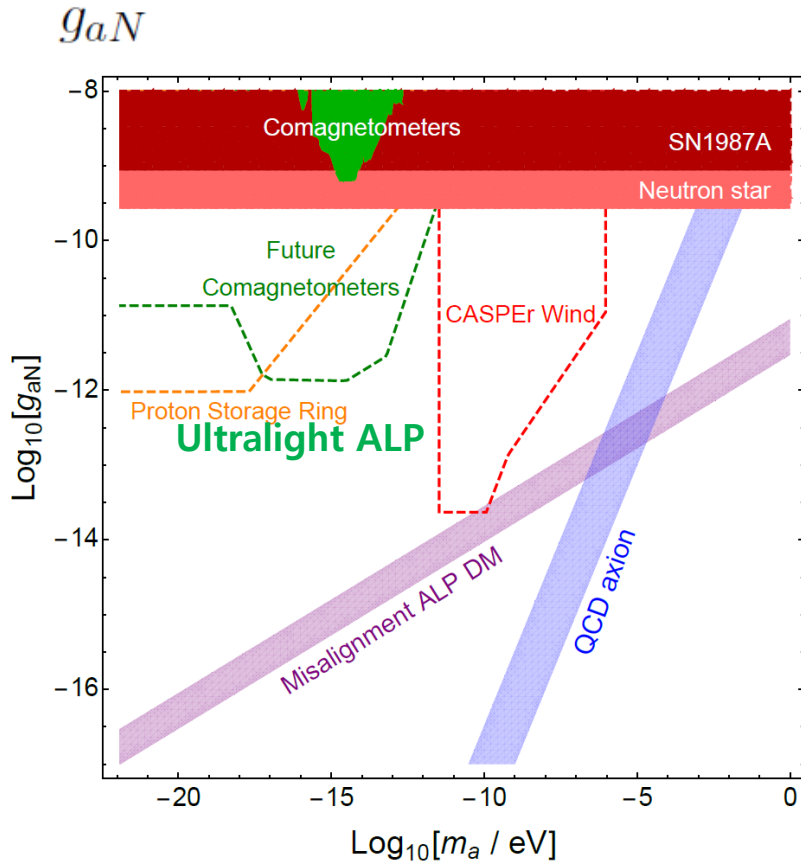
Resonant cavity:
Sikivie '83,
Semertzidis et al '09

ABRACADABRA:
Kahn et al '16

Optical ring cavity:
Obata et al '18

TOORAD:
Marsh et al '19

SRF cavity:
Berlin et al '20



Once an axion is discovered by any means, so its mass is known, we might be able to experimentally determine all of g_{aX} ($X = \gamma, n, p, e$) in some future.

Couplings of field-theoretic axions

UV completion within 4D EFT with linearly realized PQWW, KSVZ, DFSZ, ..

$$U(1)_{\text{PQ}} : \Phi \rightarrow e^{iX_{\Phi}\alpha}\Phi \quad (X_{\Phi} = \text{quantized PQ-charges})$$

Couplings of field-theoretic axions (in the unit of $1/f_a$) are determined mostly by the quantized PQ-charges:

* Couplings to the photon and gluons:

$$c_{\gamma} = \sum_{\psi} X_{\psi} Q_{\text{em}}^2(\psi), \quad c_G = \sum_{\psi} X_{\psi} Q_{\text{color}}^2(\psi) \quad (\psi = \text{chiral fermions})$$

* Tree-level couplings to the light quarks and electron:

$$C_u^0 = X_Q + X_{u^c} + \dots$$

$$C_d^0 = X_Q + X_{d^c} + \dots,$$

$$C_e^0 = X_L + X_{e^c} + \dots$$

Minimal **DFSZ**: SM fields have flavor-universal non-zero $U(1)_{PQ}$ charges as

Dine-Fischler-Srednicki-Zhitnitsky

$$c_{H_u} = c_{H_d} = -\frac{1}{2}, \quad c_\psi = \frac{1}{4} \quad (\psi \in \text{SM}), \quad c_G = c_W = \frac{3}{5}c_B = 3$$

$$\rightarrow C_u^0 = \cos^2 \beta, \quad C_d^0 = C_e^0 = \sin^2 \beta \quad \left(\tan \beta = H_u/H_d \right)$$

Minimal **KSVZ**: All SM fields are neutral under $U(1)_{PQ}$, but there exist exotic PQ-charged quark Q & Q^c

Kim-Shifman-Vainshtein-Zakharov

$$c_Q = c_{Q^c} = -\frac{1}{2}, \quad c_\psi = 0 \quad (\psi \in \text{SM}), \quad c_G = 1, \quad c_W = 0, \quad c_B = 6Y_Q^2$$

$$\rightarrow C_\Psi^0 = 0 \quad (\Psi = u, d, e)$$

Generic field-theoretic axions may be categorized into two classes:

DFSZ-type : SM fields are PQ-charged

$$\rightarrow c_G \text{ and/or } c_{W,B} = \mathcal{O}(1)$$

$$C_{\Psi}^0 = \mathcal{O}(1) \quad (\Psi = u, d, e)$$

KSVZ-type: SM fields are all PQ-neutral, but there exist some exotic PQ-charged fermions charged under the SM gauge group

$$\rightarrow c_G \text{ and/or } c_{W,B} = \mathcal{O}(1)$$

$$C_{\Psi}^0 = 0 \quad (\Psi = u, d, e)$$

Couplings of string-theoretic axions

$$a(x) = \int_{\Sigma_p} C_{[m_1, \dots, m_p]}^{(p)}(x, y) dy^{m_1} \dots dy^{m_p}$$

$$\text{or } \partial_\mu a = \epsilon_{\mu\nu\rho\sigma} (\partial^\nu B^{\rho\sigma} + \dots)$$

For each string-theoretic axion, there exists a modulus partner which forms an N=1 chiral superfield together with the axion:

$$T = \tau(x) + i \frac{a(x)}{f_a}$$

τ = Modulus partner of an angular axion field $\frac{a(x)}{f_a} \cong \frac{a(x)}{f_a} + 2\pi$

= Euclidean action (\mathbf{S}_{ins}) of the brane instanton which is a low energy consequence of the brane which couples to the underlying p-form gauge field

$$\propto \text{Vol}(\Sigma_p) \text{ and/or } \frac{1}{g_{\text{st}}^n} \quad (g_{\text{st}} = \text{string coupling})$$

4D effective SUGRA

Kahler potential: $K = K_0(T + T^*) + Z_I(T + T^*)\Phi_I^*\Phi_I + \dots$,

Gauge kinetic function: $\mathcal{F}_A = \frac{c_A}{8\pi^2}T + \dots$ ($c_A =$ rational numbers of order unity)

$$\left(-\frac{1}{4}\text{Re}(\mathcal{F}_A)F^{A\mu\nu}F_{\mu\nu}^A + \frac{1}{4}\text{Im}(\mathcal{F}_A)F^{A\mu\nu}\tilde{F}_{\mu\nu}^A \right)$$

Couplings of string-theoretic axions (in the unit of $1/f_a$) at $\mu \sim f_a \sim M_{\text{st}}$ are determined by the modulus-dependence of the 4D gauge couplings and the matter wavefunction coefficients:

$$\frac{1}{2}\partial_\mu a \partial^\mu a + \frac{\partial_\mu a}{f_a} \left[ic_\phi(\phi^* D^\mu \phi - \phi D^\mu \phi^*) + c_\psi \bar{\psi} \bar{\sigma}^\mu \psi \right] + \frac{c_A}{32\pi^2} \frac{a(x)}{f_a} F^{A\mu\nu} \tilde{F}_{\mu\nu}^A$$

$$\frac{1}{2}f_a^2 = M_P^2 \frac{\partial^2 K_0}{\partial T \partial T^*}$$

$$c_{\phi_I} = \frac{\partial \ln Z_I}{\partial T}, \quad c_{\psi_I} = \frac{\partial \ln(e^{-K_0/2} Z_I)}{\partial T}, \quad c_A = 8\pi^2 \frac{\partial}{\partial T} \mathcal{F}_A,$$

Moduli vacuum value:

* Axion weak gravity conjecture: $\tau = S_{\text{ins}} \lesssim \mathcal{O}\left(\frac{M_P}{f_a}\right)$ Arkani-Hamed et al '07

* Axion potential induced by the brane-instanton: Dine et al '85;
Blumenhagen et al '09; ...

$$\delta V = e^{-\tau} \Lambda^2 M_P^2 \cos\left(\frac{a}{f_a}\right) \quad \text{with} \quad \Lambda^2 \sim m_{3/2} M_P \text{ or } m_{3/2}^2$$

For QCD axion, $\delta V \lesssim 10^{-10} m_\pi^2 f_\pi^2 \rightarrow \tau \gtrsim \ln(10^{10} \Lambda^2 / f_\pi m_\pi)$

For ultra-light ALP, $\delta V \lesssim m_a^2 f_a^2 \rightarrow \tau \gtrsim 2 \ln(\Lambda / m_a)$

(Note that this lower bound is meaningful only for $m_a \ll m_{3/2}$.)

* To avoid a too large $\frac{1}{g_{\text{GUT}}^2} = \text{Re}(\mathcal{F}_A) = c_A \frac{\tau}{8\pi^2} + \dots \rightarrow \tau \lesssim \mathcal{O}\left(\frac{8\pi^2}{g_{\text{GUT}}^2}\right)$

$$\rightarrow \tau = \mathcal{O}\left(\frac{8\pi^2}{g_{\text{GUT}}^2}\right)$$

For $\tau = S_{\text{ins}} = \mathcal{O}\left(\frac{8\pi^2}{g_{\text{GUT}}^2}\right) \gg 1$,

$$\frac{\partial K_0}{\partial \tau} \sim \tau \frac{\partial^2 K_0}{\partial \tau^2} \sim \tau \frac{f_a^2}{M_P^2} \lesssim \mathcal{O}\left(\frac{g_{\text{GUT}}^2}{8\pi^2}\right),$$

$$\frac{\partial \ln Z_I}{\partial \tau} \sim \frac{1}{\tau} = \mathcal{O}\left(\frac{g_{\text{GUT}}^2}{8\pi^2}\right),$$

$$\rightarrow c_\psi \sim c_\phi = \mathcal{O}\left(\frac{1}{\tau}\right) = \mathcal{O}\left(\frac{g_{\text{GUT}}^2}{8\pi^2}\right)$$

$$\rightarrow C_\Psi^0 = \mathcal{O}\left(\frac{1}{\tau}\right) = \mathcal{O}\left(\frac{g_{\text{GUT}}^2}{8\pi^2}\right) \quad (\Psi = u, d, e)$$

for string-theoretic axions

Overall strength of the axion couplings is determined by f_a .

Is there any range of f_a favored by string-theoretic axions?

In string compactifications without a big hierarchy between the string scale and the 4D Planck scale, KC & Kim, '85
Svrcek & Witten '06

$$f_a = \mathcal{O} \left(\frac{g_{\text{GUT}}^2}{8\pi^2} M_P \right) \sim 10^{16} - 10^{17} \text{ GeV}$$

However, in string compactifications with a large compactification volume and/or a large warp factor which would generate a big scale hierarchy, f_a of string-theoretic axions can in principle be anywhere in the range

$$10^7 - 10^8 \text{ GeV} \lesssim f_a \lesssim \mathcal{O} \left(\frac{g_{\text{GUT}}^2}{8\pi^2} M_P \right)$$

Astrophysics

Weak Gravity Conjecture

Burgess, Ibanez, Quevedo '99,
Conlon '06,
Cicoli et al '12;
...

The same is also true for the f_a of generic field-theoretic axions.

Would it be possible to distinguish these three types of axions by experimentally measurable g_{aX} ($X = \gamma, p, n, e$) ?

$$\frac{1}{2}g_{a\gamma}a(x)\vec{E} \cdot \vec{B} + \partial_\mu a(x) \left(\frac{g_{ap}}{2m_p} \bar{p}\gamma^\mu \gamma_5 p + \frac{g_{an}}{2m_n} \bar{n}\gamma^\mu \gamma_5 n + \frac{g_{ae}}{2m_e} \bar{e}\gamma^\mu \gamma_5 e \right)$$

Both KSVZ-type and string-theoretic axions have $C_\Psi^0/c_A \ll 1$, therefore one needs to study radiative corrections ΔC_Ψ ($\Psi = u, d, e$) to see if they have a distinguishable pattern of g_{aX} ($X = \gamma, p, n, e$).

RG evolution of axion couplings

High scale axion couplings at $\mu \sim f_a$:

$$\frac{\partial_\mu a}{f_a} \left[i c_\phi (\phi^* D^\mu \phi - \phi D^\mu \phi^*) + c_\psi \bar{\psi} \bar{\sigma}^\mu \psi \right] + c_A \frac{g_A^2}{32\pi^2} \frac{a(x)}{f_a} F^{A\mu\nu} \tilde{F}_{\mu\nu}^A$$

$$\phi = \{H_u, H_d, \dots\} \quad \psi = \{Q_i, u_i^c, d_i^c, L_i, e_i^c\} \quad F_{\mu\nu}^A = (G_{\mu\nu}^\alpha, W_{\mu\nu}^i, B_{\mu\nu})$$

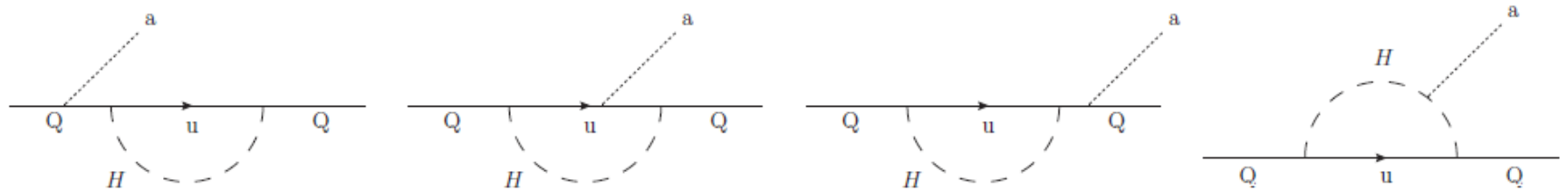
$c_A = \{c_G, c_W, c_B\}$: rational numbers

$$c_\phi(f_a) = c_\phi^0, \quad c_\psi(f_a) = c_\psi^0$$

RG evolution of axion couplings

Yukawa-induced 1-loop running:

KC, Im, Park, Yun, 1708.00021



$$\left. \frac{dc_F}{d \ln \mu} \right|_{1\text{-loop}} = \frac{\xi_y}{16\pi^2} \sum_{f,\alpha} \left(\frac{1}{2} \{c_F, \mathbf{y}_{fF\alpha}^\dagger \mathbf{y}_{fF\alpha}\} + \mathbf{y}_{fF\alpha}^\dagger \mathbf{c}_f^T \mathbf{y}_{fF\alpha} + c_{H\alpha} \mathbf{y}_{fF\alpha}^\dagger \mathbf{y}_{fF\alpha} \right),$$

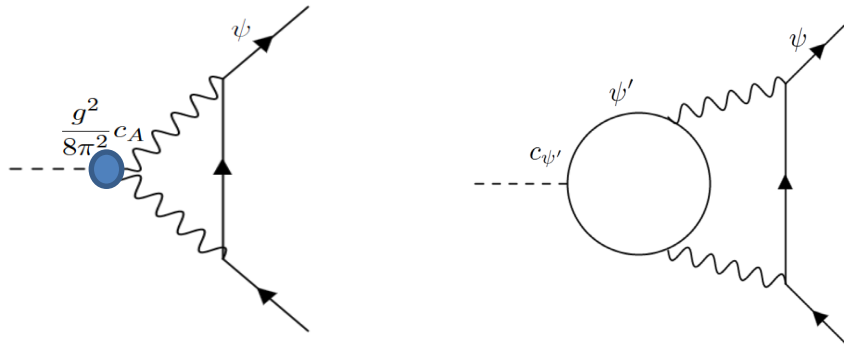
$$\left. \frac{dc_f^T}{d \ln \mu} \right|_{1\text{-loop}} = \frac{\xi_y}{16\pi^2} \sum_{F,\alpha} \left(\frac{1}{2} \{c_f^T, \mathbf{y}_{fF\alpha} \mathbf{y}_{fF\alpha}^\dagger\} + \mathbf{y}_{fF\alpha} c_F \mathbf{y}_{fF\alpha}^\dagger + c_{H\alpha} \mathbf{y}_{fF\alpha} \mathbf{y}_{fF\alpha}^\dagger \right),$$

$$\left. \frac{dc_{H\alpha}}{d \ln \mu} \right|_{1\text{-loop}} = \frac{1}{8\pi^2} \sum_{f,F} \left(c_{H\alpha} \text{tr}(\mathbf{y}_{fF\alpha}^\dagger \mathbf{y}_{fF\alpha}) + \text{tr}(\mathbf{y}_{fF\alpha} c_F \mathbf{y}_{fF\alpha}^\dagger) + \text{tr}(\mathbf{y}_{fF\alpha}^\dagger \mathbf{c}_f^T \mathbf{y}_{fF\alpha}) \right),$$

$$F_i = \{q_i, \ell_i\} \quad f_i = \{u_i^c, d_i^c, e_i^c\}$$

$$\text{non-SUSY : } \xi_y = 1 \quad \text{SUSY : } \xi_y = 2$$

Gauge-induced 2-loop (computationally 1-loop) running:



KC, Im, Shin, 2012.05029
 Baur et al, 2012.12272
 KC, Im, Seong, Kim, arXiv:2106.05816

$$\left. \frac{d\mathbf{c}_\psi}{d \ln \mu} \right|_{2\text{-loop}} = -\xi_g \sum_A \frac{3}{2} \left(\frac{g_A^2}{8\pi^2} \right)^2 \mathbb{C}_A(\psi) \left(c_A - 2 \sum_{\psi'} \text{tr}(\mathbf{c}_{\psi'}) \mathbb{T}_A(\psi') \right) \mathbf{1},$$

$$\left. \frac{dc_{H_\alpha}}{d \ln \mu} \right|_{2\text{-loop}} = -\xi_H \sum_A \frac{3}{2} \left(\frac{g_A^2}{8\pi^2} \right)^2 \mathbb{C}_A(H_\alpha) \left(c_A - 2 \sum_{\psi'} \text{tr}(\mathbf{c}_{\psi'}) \mathbb{T}_A(\psi') \right),$$

$\mathbb{C}_A(\Phi)$ = quadratic Casimir

$\mathbb{T}_A(\Phi)$ = Dynkin index

non-SUSY : $\xi_g = 1, \xi_H = 0$

SUSY : $\xi_g = \xi_H = \frac{2}{3}$

$$\begin{aligned}
\text{MSSM: } \frac{d\mathbf{c}_Q}{d \ln \mu} &= \frac{\xi_y}{16\pi^2} \left(\frac{1}{2} \{ \mathbf{c}_Q, \tilde{\mathbf{y}}_u^\dagger \tilde{\mathbf{y}}_u + \tilde{\mathbf{y}}_d^\dagger \tilde{\mathbf{y}}_d \} + \tilde{\mathbf{y}}_u^\dagger \mathbf{c}_{u^c}^T \tilde{\mathbf{y}}_u + \tilde{\mathbf{y}}_d^\dagger \mathbf{c}_{d^c}^T \tilde{\mathbf{y}}_d + c_{H_u} \tilde{\mathbf{y}}_u^\dagger \tilde{\mathbf{y}}_u + c_{H_d} \tilde{\mathbf{y}}_d^\dagger \tilde{\mathbf{y}}_d \right) \\
&\quad - \xi_g \left(\frac{\alpha_s^2}{2\pi^2} \tilde{c}_G + \frac{9\alpha_2^2}{32\pi^2} \tilde{c}_W + \frac{\alpha_1^2}{96\pi^2} \tilde{c}_B \right) \mathbb{1}, \\
\frac{d\mathbf{c}_{u^c}^T}{d \ln \mu} &= \frac{\xi_y}{8\pi^2} \left(\frac{1}{2} \{ \mathbf{c}_{u^c}^T, \tilde{\mathbf{y}}_u \tilde{\mathbf{y}}_u^\dagger \} + \tilde{\mathbf{y}}_u \mathbf{c}_Q \tilde{\mathbf{y}}_u^\dagger + c_{H_u} \tilde{\mathbf{y}}_u \tilde{\mathbf{y}}_u^\dagger \right) - \xi_g \left(\frac{\alpha_s^2}{2\pi^2} \tilde{c}_G + \frac{\alpha_1^2}{6\pi^2} \tilde{c}_B \right) \mathbb{1}, \\
\frac{d\mathbf{c}_{d^c}^T}{d \ln \mu} &= \frac{\xi_y}{8\pi^2} \left(\frac{1}{2} \{ \mathbf{c}_{d^c}^T, \tilde{\mathbf{y}}_d \tilde{\mathbf{y}}_d^\dagger \} + \tilde{\mathbf{y}}_d \mathbf{c}_Q \tilde{\mathbf{y}}_d^\dagger + c_{H_d} \tilde{\mathbf{y}}_d \tilde{\mathbf{y}}_d^\dagger \right) - \xi_g \left(\frac{\alpha_s^2}{2\pi^2} \tilde{c}_G + \frac{\alpha_1^2}{24\pi^2} \tilde{c}_B \right) \mathbb{1}, \\
\frac{d\mathbf{c}_L}{d \ln \mu} &= \frac{\xi_y}{16\pi^2} \left(\frac{1}{2} \{ \mathbf{c}_L, \tilde{\mathbf{y}}_e^\dagger \tilde{\mathbf{y}}_e \} + \tilde{\mathbf{y}}_e^\dagger \mathbf{c}_{e^c}^T \tilde{\mathbf{y}}_e + c_{H_e} \tilde{\mathbf{y}}_e^\dagger \tilde{\mathbf{y}}_e \right) - \xi_g \left(\frac{9\alpha_2^2}{32\pi^2} \tilde{c}_W + \frac{3\alpha_1^2}{32\pi^2} \tilde{c}_B \right) \mathbb{1}, \\
\frac{d\mathbf{c}_{e^c}^T}{d \ln \mu} &= \frac{\xi_y}{8\pi^2} \left(\frac{1}{2} \{ \mathbf{c}_{e^c}^T, \tilde{\mathbf{y}}_e \tilde{\mathbf{y}}_e^\dagger \} + \tilde{\mathbf{y}}_e \mathbf{c}_L \tilde{\mathbf{y}}_e^\dagger + c_{H_e} \tilde{\mathbf{y}}_e \tilde{\mathbf{y}}_e^\dagger \right) - \xi_g \frac{3\alpha_1^2}{8\pi^2} \tilde{c}_B \mathbb{1}, \\
\frac{dc_{H_u}}{d \ln \mu} &= \frac{3}{8\pi^2} \left(c_{H_u} \text{tr}(\tilde{\mathbf{y}}_u^\dagger \tilde{\mathbf{y}}_u) + \text{tr}(\tilde{\mathbf{y}}_u \mathbf{c}_Q \tilde{\mathbf{y}}_u^\dagger) + \text{tr}(\tilde{\mathbf{y}}_u^\dagger \mathbf{c}_{u^c}^T \tilde{\mathbf{y}}_u) \right) - \xi_H \left(\frac{9\alpha_2^2}{32\pi^2} \tilde{c}_W + \frac{3\alpha_1^2}{32\pi^2} \tilde{c}_B \right), \\
\frac{dc_{H_d}}{d \ln \mu} &= \frac{3}{8\pi^2} \left(c_{H_d} \text{tr}(\tilde{\mathbf{y}}_d^\dagger \tilde{\mathbf{y}}_d) + \text{tr}(\tilde{\mathbf{y}}_d \mathbf{c}_Q \tilde{\mathbf{y}}_d^\dagger) + \text{tr}(\tilde{\mathbf{y}}_d^\dagger \mathbf{c}_{d^c}^T \tilde{\mathbf{y}}_d) \right) \\
&\quad + \frac{1}{8\pi^2} \left(c_{H_d} \text{tr}(\tilde{\mathbf{y}}_e^\dagger \tilde{\mathbf{y}}_e) + \text{tr}(\tilde{\mathbf{y}}_e \mathbf{c}_L \tilde{\mathbf{y}}_e^\dagger) + \text{tr}(\tilde{\mathbf{y}}_e^\dagger \mathbf{c}_{e^c}^T \tilde{\mathbf{y}}_e) \right) - \xi_H \left(\frac{9\alpha_2^2}{32\pi^2} \tilde{c}_W + \frac{3\alpha_1^2}{32\pi^2} \tilde{c}_B \right),
\end{aligned}$$

$$\tilde{c}_G = c_G - \text{tr}(2\mathbf{c}_Q + \mathbf{c}_{u^c} + \mathbf{c}_{d^c}),$$

$$\tilde{c}_W = c_W - \text{tr}(3\mathbf{c}_Q + \mathbf{c}_L) - \frac{3}{2} \xi_H (c_{H_u} + c_{H_d}),$$

$$\tilde{c}_B = c_B - \text{tr} \left(\frac{1}{3} (\mathbf{c}_Q + 8\mathbf{c}_{u^c} + 2\mathbf{c}_{d^c}) + \mathbf{c}_L + 2\mathbf{c}_{e^c} \right) - \frac{3}{2} \xi_H (c_{H_u} + c_{H_d})$$

One can systematically integrate the RG evolution, while including the relevant threshold corrections to compute ΔC_Ψ ($\Psi = u, d, e$) at $\mu = 2 \text{ GeV}$, and then use

$$\frac{1}{2}g_{a\gamma}a(x)\vec{E} \cdot \vec{B} + \partial_\mu a(x) \left(\frac{g_{ap}}{2m_p} \bar{p}\gamma^\mu \gamma_5 p + \frac{g_{an}}{2m_n} \bar{n}\gamma^\mu \gamma_5 n + \frac{g_{ae}}{2m_e} \bar{e}\gamma^\mu \gamma_5 e \right)$$

$$g_{a\gamma} \simeq \frac{\alpha_{\text{em}}}{2\pi} \frac{1}{f_a} \left(c_\gamma - \frac{2}{3} \left(\frac{m_u + 4m_d}{m_u + m_d} \right) c_G \right)$$

$$\begin{aligned} g_{ap} &\simeq \frac{m_p}{f_a} \left(C_u \Delta u + C_d \Delta d - \left(\frac{m_d}{m_u + m_d} \Delta u + \frac{m_u}{m_u + m_d} \Delta d \right) c_G \right) \\ &\simeq \frac{m_p}{f_a} \left(0.88 C_u - 0.39 C_d - 0.47 c_G \right) \end{aligned}$$

$$\begin{aligned} g_{an} &\simeq \frac{m_n}{f_a} \left(C_d \Delta u + C_u \Delta d - \left(\frac{m_u}{m_u + m_d} \Delta u + \frac{m_d}{m_u + m_d} \Delta d \right) c_G \right) \\ &\simeq \frac{m_n}{f_a} \left(-0.39 C_u + 0.88 C_d - 0.02 c_G \right) \end{aligned}$$

$$g_{ae} \simeq \frac{m_e}{f_a} C_e$$

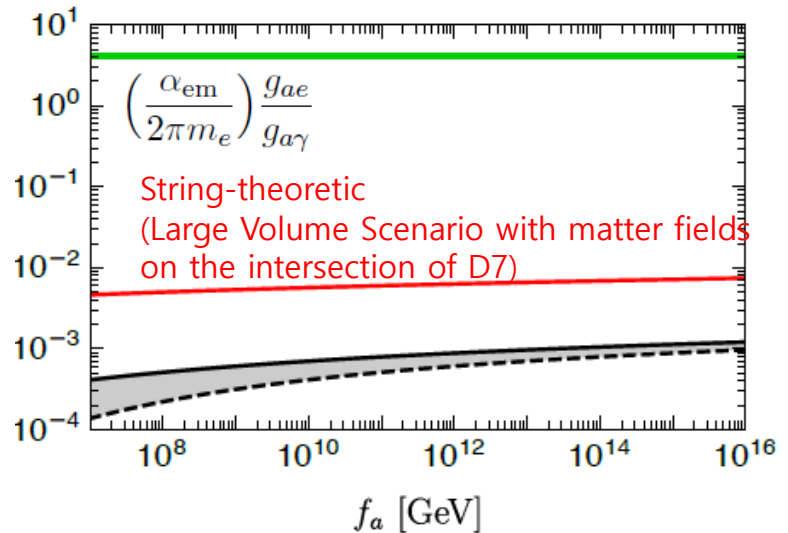
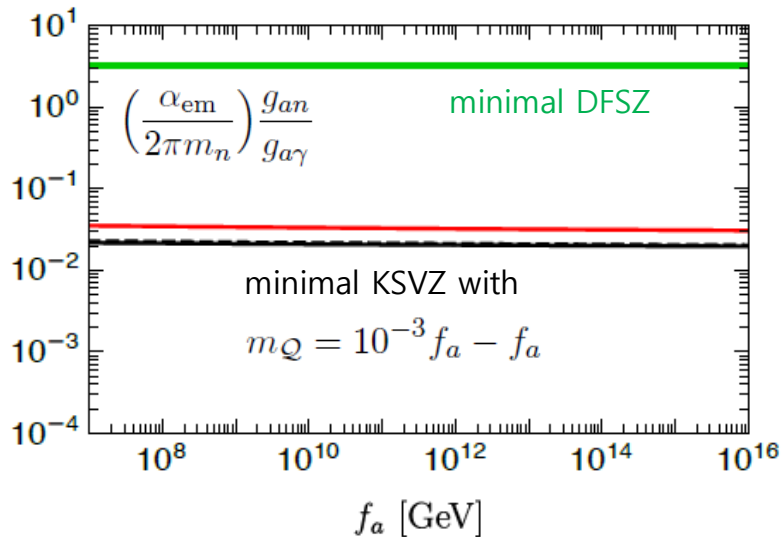
accidental cancellation for the contribution from \mathbf{c}_G

$g_{aX}/g_{a\gamma}$ ($X = p, n, e$) including radiative corrections

KC, Im, Seong, Kim, arXiv:2106.05816

QCD axion ($c_G \neq 0$)

- * All three-types of QCD axions have a similar value of $g_{ap}/g_{a\gamma}$ because of the contribution from the axion-meson mixing induced by c_G .
- * As the axion-meson mixing contribution to g_{an} & g_{ae} are small, these three-type of QCD axions have distinguishable pattern of $g_{an}/g_{a\gamma}$ and $g_{ae}/g_{a\gamma}$.



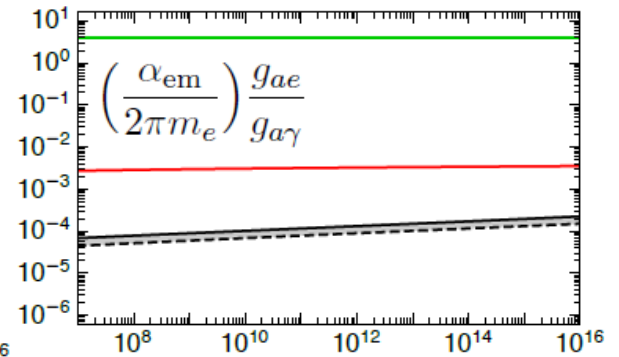
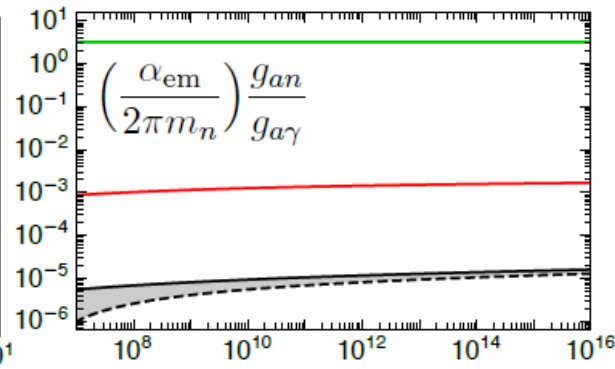
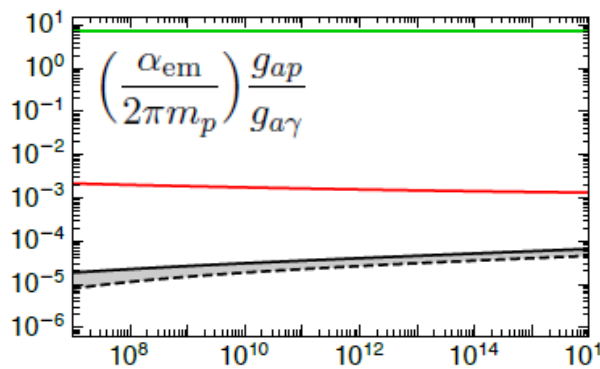
$$\left(\frac{g_{an}}{g_{a\gamma}}\right)_{\text{DFSZ}} \gg \left(\frac{g_{an}}{g_{a\gamma}}\right)_{\text{string}} \simeq \left(\frac{g_{an}}{g_{a\gamma}}\right)_{\text{KSVZ}}$$

$$\left(\frac{g_{ae}}{g_{a\gamma}}\right)_{\text{DFSZ}} \gg \left(\frac{g_{ae}}{g_{a\gamma}}\right)_{\text{string}} \gg \left(\frac{g_{ae}}{g_{a\gamma}}\right)_{\text{KSVZ}}$$

Ultra-light ALP ($c_G = 0$, c_B or $c_W = 1$)

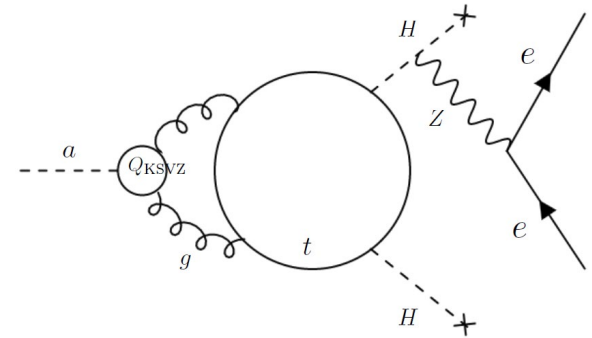
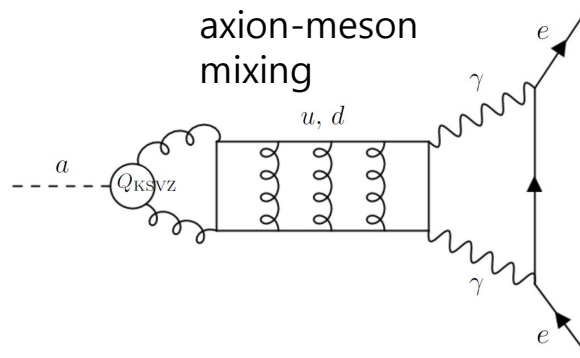
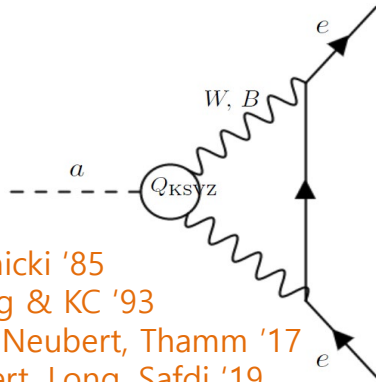
Different type of ALPs have clearly distinguishable pattern for all three coupling ratios:

$$\left(\frac{g_{aX}}{g_{a\gamma}}\right)_{\text{DFSZ}} \gg \left(\frac{g_{aX}}{g_{a\gamma}}\right)_{\text{string}} \gg \left(\frac{g_{aX}}{g_{a\gamma}}\right)_{\text{KSVZ}} \quad (X = p, n, e)$$



Coupling of KSVZ-type QCD axion to the electron

Srednicki '85
 Chang & KC '93
 Baur, Neubert, Thamm '17
 Dessert, Long, Safdi '19



Two-loop RG-running from the exotic heavy quark mass to m_e

$$\begin{aligned}
 (g_{ae})_{\text{KSVZ}} = & \frac{m_e}{f_a} \left[\frac{3\alpha_{\text{em}}^2}{4\pi^2} \left(\frac{3c_W}{8s_W^4} + \frac{5c_B}{8c_W^4} \right) \ln \frac{M_{Q_{\text{KSVZ}}}}{m_e} - \frac{\alpha_{\text{em}}^2}{2\pi^2} \left(\frac{4m_d + m_u}{m_u + m_d} \right) c_G \ln \frac{4\pi f_\pi}{m_e} \right. \\
 & \left. + \mathcal{O}\left(\frac{y_t^2 \alpha_s^2}{8\pi^4}\right) c_G \ln \frac{M_{Q_{\text{KSVZ}}}}{m_t} \right]
 \end{aligned}$$

↖ higher-loop ↖ axion-meson-mixing

$$= 10^{-4} \frac{m_e}{f_a} \left[1.3 c_B + 5.4 c_W - 0.3 c_G + 8.3 c_G \right] \quad \text{for } f_a \gtrsim m_{Q_{\text{KSVZ}}} = 10^{10} \text{ GeV}$$

The higher loop contribution which was not considered before gives the dominant contribution.

Conclusion

- Considering the low energy couplings and the origin of axion field variable, axions can be categorized into three types:

DFSZ-type, KSVZ-type, String-theoretic,

having parametrically different ratios between the couplings to gauge fields and the couplings to matter fields at the UV scale.

- If an axion is discovered, so its mass is identified, we might be able to measure multiple number of low energy axion couplings and determine the coupling ratios $\mathbf{g_{ax} / g_{ay}}$ ($\mathbf{X=p,n,e}$) which would give us information on the underlying axion model.

Compared to the other two types, DFSZ-type has a clearly distinct pattern of low energy couplings.

However, to discriminate string-theoretic from KSVZ-type, we need an analysis taking into account the radiative corrections to axion couplings.

- For **QCD axions**, $g_{ae}/g_{a\gamma}$ of KSVZ-type and string-theoretic axions differ by about an order of magnitude, while $g_{aN}/g_{a\gamma}$ are similar to each other.

For **ultra-light ALPs** without the coupling to gluons, different types of ALPs have clearly distinguishable pattern for all three coupling ratios $g_{aX}/g_{a\gamma}$ ($X=p,n,e$).

- The coupling of KSVZ-type QCD axion to the electron is dominated by the higher-loop contribution involving the top quark and Higgs boson, which was ignored in the previous studies.

Thank you for your attention.