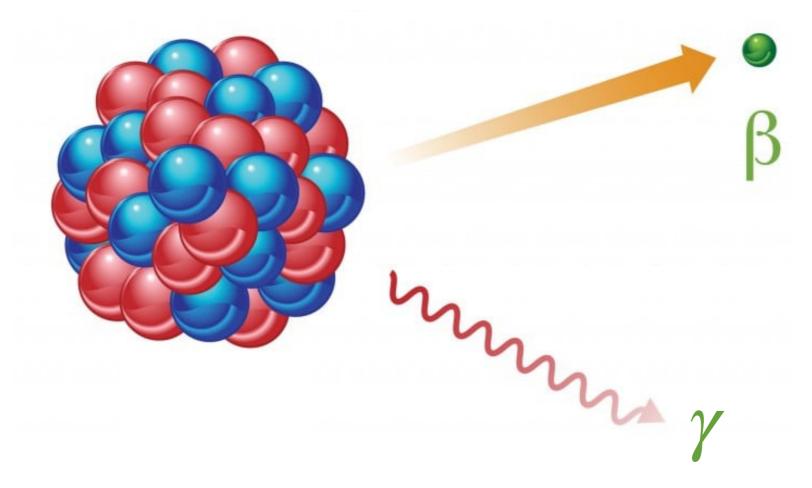
#### Searching for axion dark matter via nuclear decay anomalies



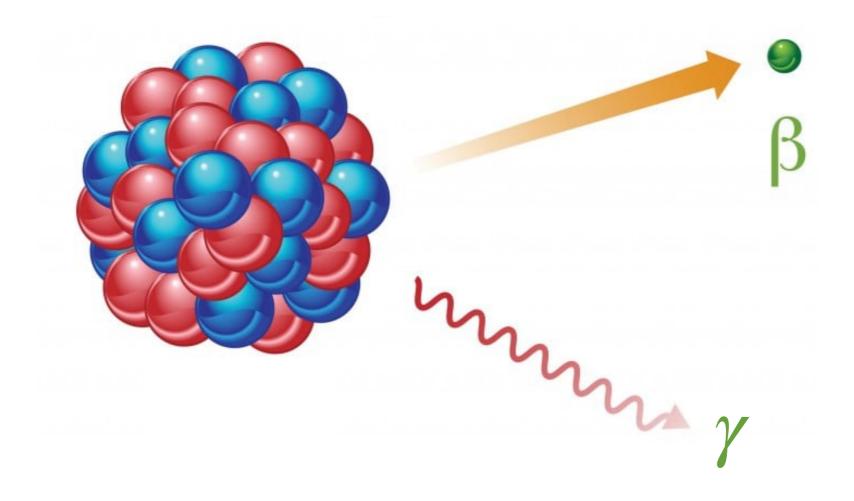
Nick Houston, Beijing University of Technology

In collaboration with Xin Zhang (NAOC) and Tianjun Li (ITP-CAS)

Based on 2212.XXXX

Axion 2022, Nov 2022, <u>nhouston@bjut.edu.cn</u>

## The big picture



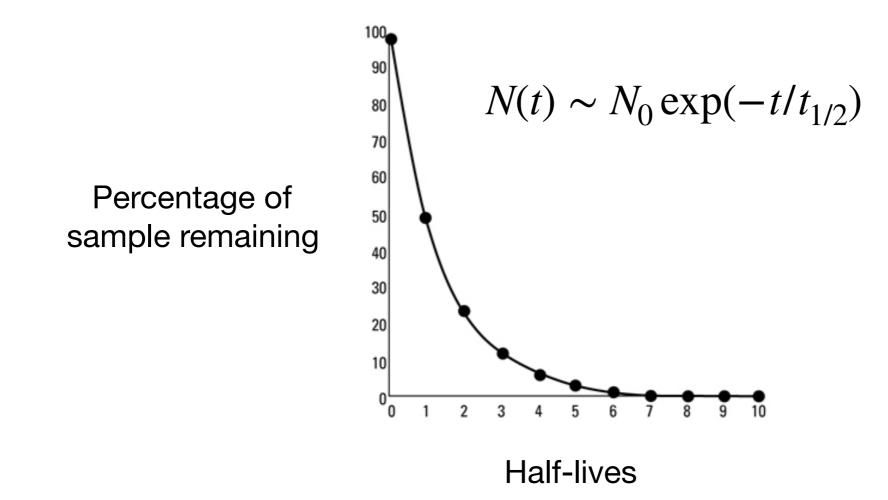
Fundamentally, we believe that nuclear decay is **random** and **spontaneous** 

However, we also expect QCD axion DM will lead to an oscillating  $\theta$ -angle

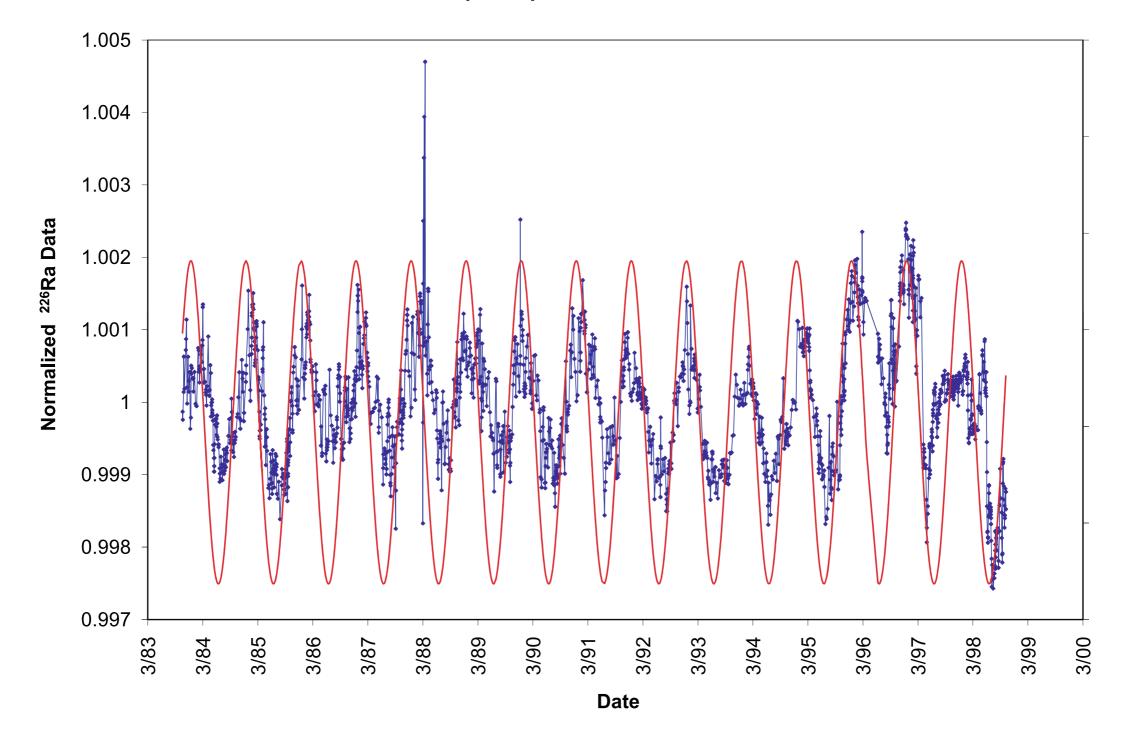
As  $\theta$  modifies nuclear physics, this can lead to non-random decay behaviour

# What is nuclear decay?

- "Nuclear decay is the process by which an unstable atom loses energy by emitting radiation, generally changing the number of protons and neutrons in the nucleus"
- We can only predict how often this will happen on average



• This is well established science, why should we question this?

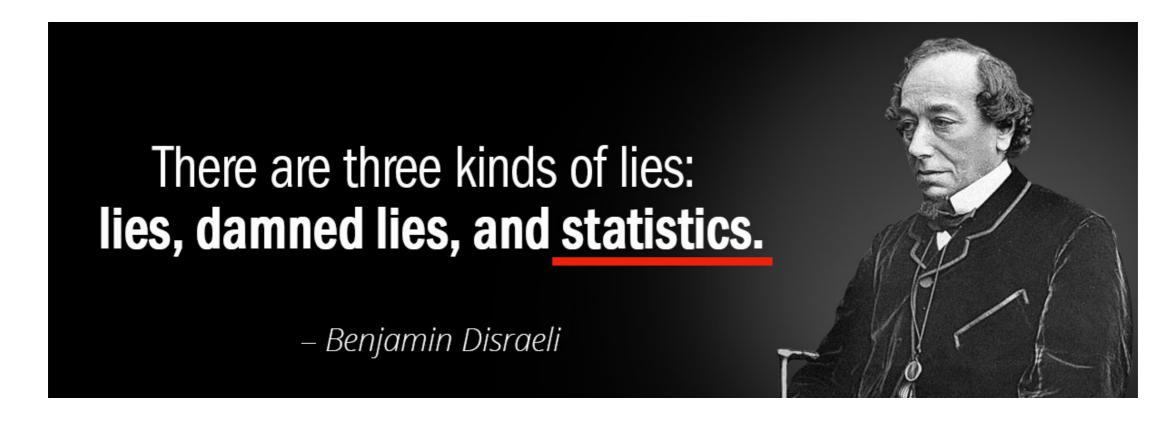


#### Normalized <sup>226</sup>Ra (PTB) Data with Earth-Sun Distance

• Should we believe this?

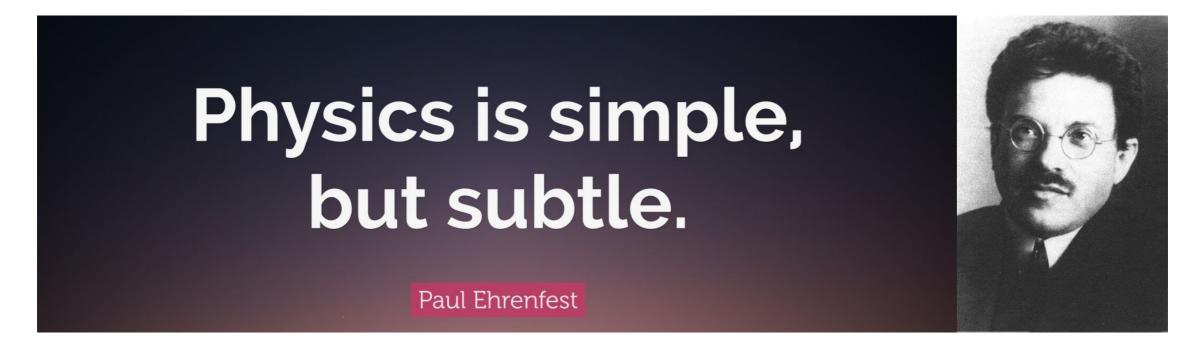
"Time-dependent nuclear decay parameters: New evidence for new forces?", *Space Sci.Rev.* 145 (2009) 285-335 "Anomalies in Radioactive Decay Rates: A Bibliography of Measurements and Theory", arxiv: 2111.03149

## **Reasons to be skeptical: 1**



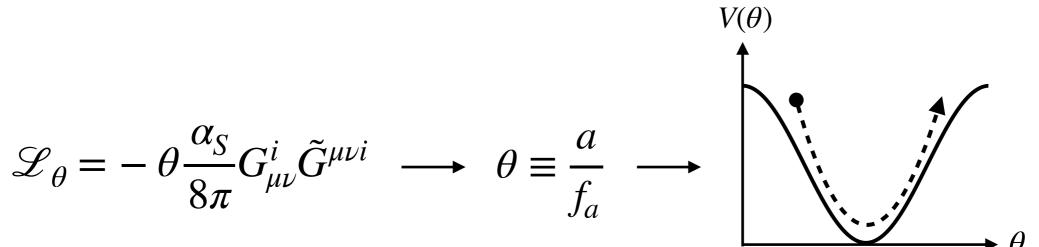
- The data analysis here is quite subtle
- Is it possible these anomalies are due to incorrect statistics?

### **Reasons to be skeptical: 2**



- Can we explain these anomalies without rewriting the laws of physics?
- Did seasonal variations in temperature influence the experiment?
- Is there any possible explanation in terms of fundamental physics?

### **Recall: the misalignment mechanism**



• For QCD axions, with initial condition  $\theta_{a,i}$  we typically have

$$\Omega_a h^2 \sim 2 \times 10^4 \left( \frac{f_a}{10^{16} \text{ GeV}} \right)^{7/6} \langle \theta_{a,i}^2 \rangle, \quad \theta \simeq \sqrt{\frac{2\rho_{DM}}{m_a^2 f_a^2}} \cos(\omega t + \overrightarrow{p} \cdot \overrightarrow{x} + \phi)$$

• Many aspects of nuclear physics depend on  $\theta$ , for example:

$$d_n = \frac{g_{\pi NN}}{4\pi} \left(\frac{e}{m_p f_\pi}\right) \ln\left(\frac{m_\rho}{m_\pi}\right) \left(\frac{m_u m_d}{m_u + m_d}\right) \theta, \, m_n - m_p \simeq \left(1.29 + 0.37 \,\theta^2\right) \,\text{MeV}$$

• By modifying nuclear binding energies,  $\theta$  changes decay rates

### **Tritium decay**

• For simple nuclei this is calculable, let's consider tritium decay:

$${}^{3}H \rightarrow {}^{3}He + e^{-} + \bar{\nu}_{e}, \ t_{1/2} \simeq 12.3$$
 years,  $Q = 18.6$  keV

$$\Gamma^{\beta}(^{3}\mathrm{H}) = \frac{1}{2\pi^{3}} m_{e} (G_{\beta} m_{e}^{2})^{2} (B_{F}(^{3}\mathrm{H}) + B_{GT}(^{3}\mathrm{H})) I^{\beta}(^{3}\mathrm{H})$$

$$B_F(^{3}\mathrm{H}) = \frac{1}{2} \left| {}^{_{3}\mathrm{He}}\langle (1/2)^{+} \parallel \sum_{n} \tau_n^{+} \parallel (1/2)^{+} \rangle_{^{3}\mathrm{H}} \right|^{^{2}}, B_{GT}(^{^{3}}\mathrm{H}) = g_A^2 \frac{1}{2} \left| {}^{_{3}\mathrm{He}}\langle (1/2)^{+} \parallel \sum_{n} \tau_n^{+} \sigma_n \parallel (1/2)^{+} \rangle_{^{3}\mathrm{H}} \right|^{^{2}}$$

$$I^{\beta}(^{3}\mathrm{H}) = \frac{1}{m_{e}^{5}} \int_{m_{e}}^{E_{i}-E_{f}} F_{0}(Z+1, E_{e}) p_{e} E_{e} (E_{i} - E_{f} - E_{e})^{2} dE_{e}$$

• Where does  $\theta$ -dependence enter?

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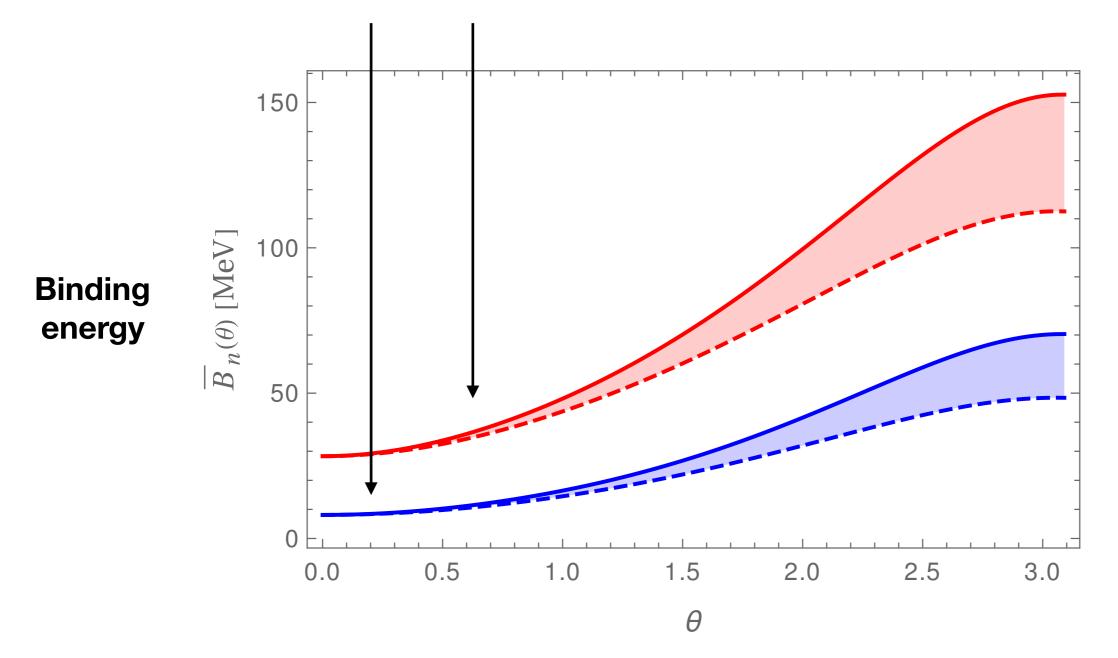
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• Where does  $\theta$ -dependence enter?

- $\theta$  changes the decay rate by modifying  ${}^{3}H/{}^{3}He$  binding energies
- Fortunately for 3 and 4 nucleon systems this is already calculated



 $\theta$ -dependence of light nuclei and nucleosynthesis, 2006.12321

### **Tritium decay**

• So, let's add a perturbation  $\delta E(\theta)$  to  $E_i - E_f$ : (Using Primakoff-Rosen approximation for  $F_0$ )

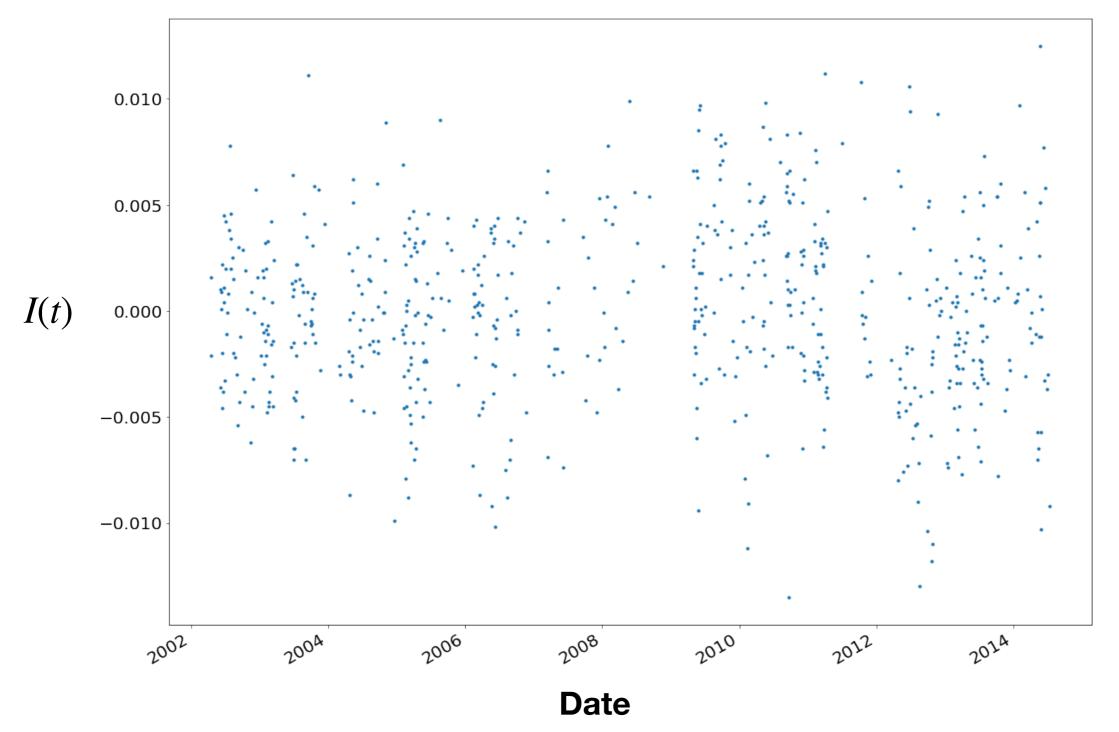
$$\frac{\delta\Gamma^{\beta}}{\Gamma^{\beta}} = 1 - \frac{5\delta E(\theta) \left( E_f^2 - 2E_f(E_i + m_e) + E_i^2 + 2E_i m_e + 3m_e^2 \right)}{(E_f - E_i + m_e) \left( 3m_e(E_i - E_f) + (E_f - E_i)^2 + 6m_e^2 \right)} + \mathcal{O}(\delta E^2)$$

• From the previous slide, we know how  $\delta E$  depends on  $\theta$ , and so

$$\delta E \simeq \mu \text{eV} \left( \frac{\rho_{DM}}{0.4 \text{GeV/cm}^3} \right) \left( \frac{10^{16} \text{GeV}}{f_a} \right)^2 \left( \frac{10^{-22} \text{eV}}{m_a} \right)^2 \cos(\omega t)$$

• So, now all we need is some tritium...

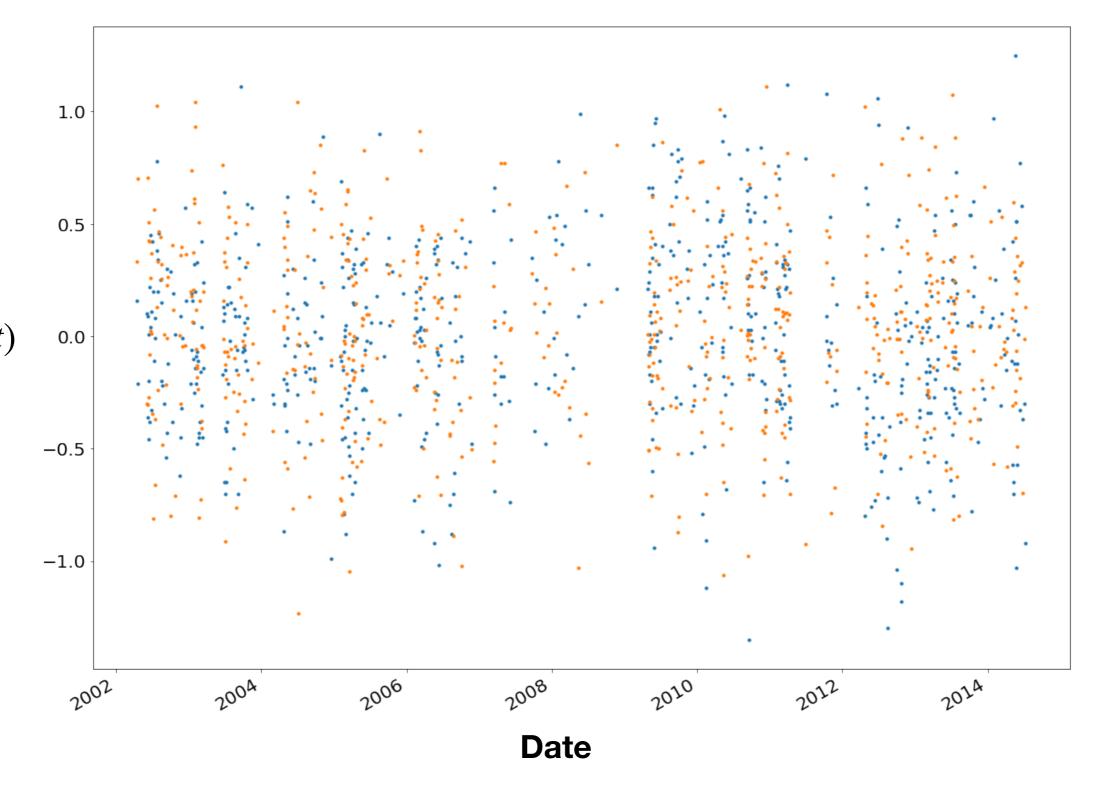
#### **Tritium decay data**



 $I(t) \equiv \frac{N(t) - \langle N \rangle}{\langle N \rangle}$ 

Data is from the European Union's Joint Research Centre, at the Directorate for Nuclear Safety and Security in Belgium

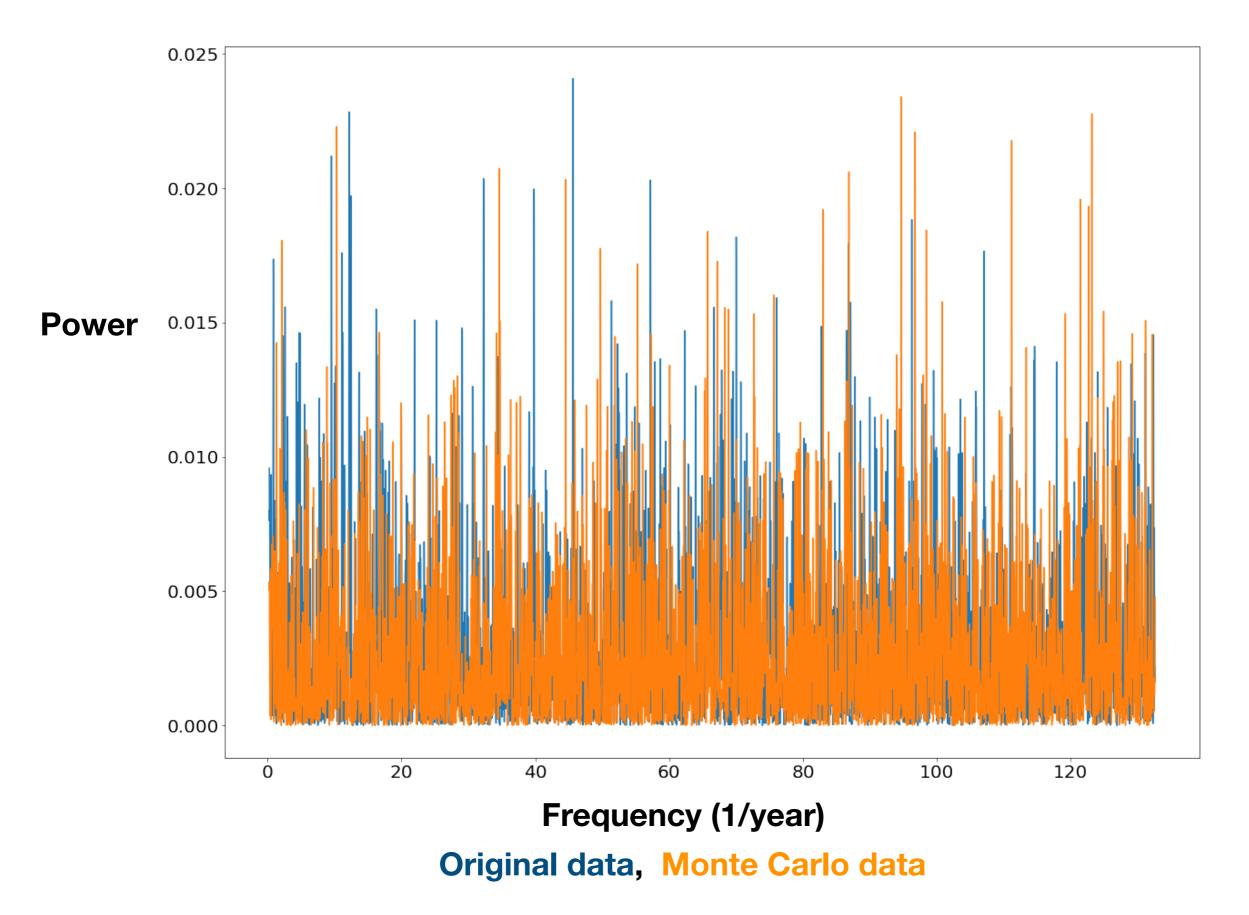
- Let's compare the real data to Monte Carlo simulations:
- 1. Generate N datasets with randomly generated I(t)
- 2. For each dataset, convert to frequency space
- 3. Construct the CDF at each frequency
- 4. Find the 95 % CL limit
- 5. Compare to the real power at that frequency
- For example:



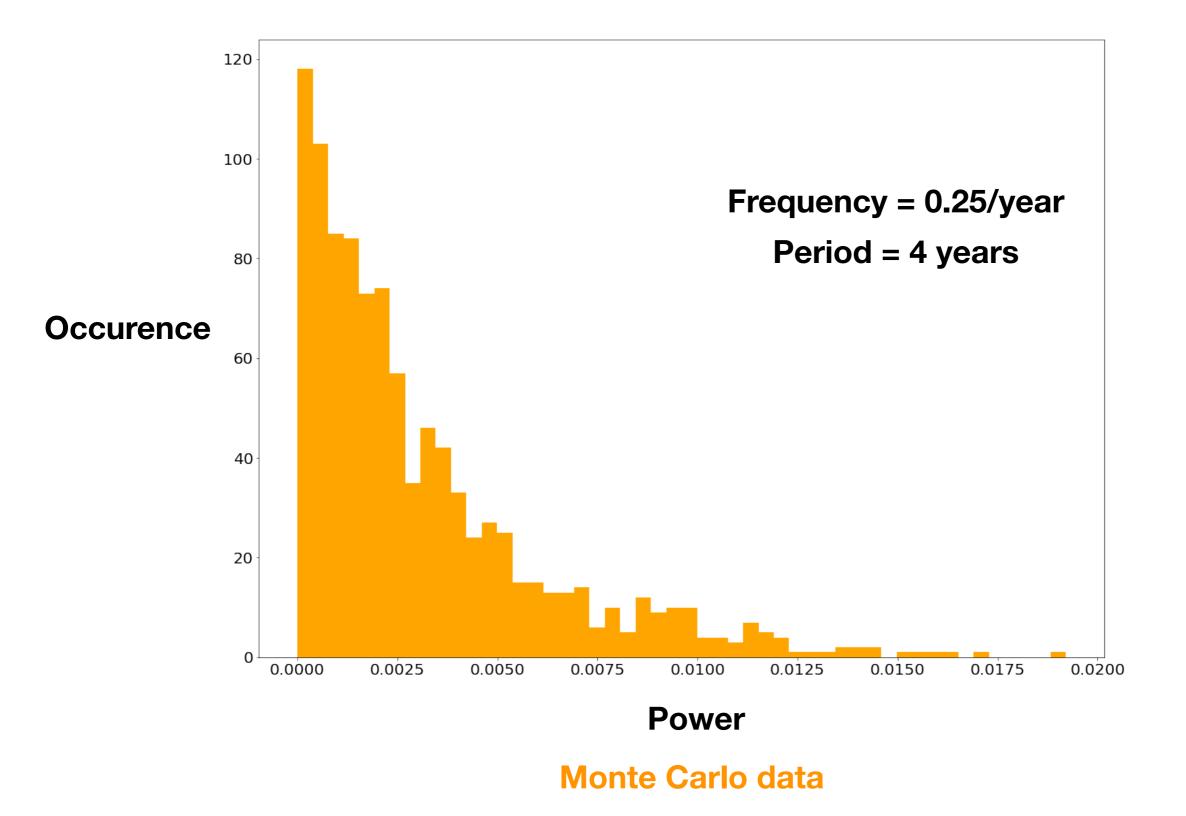
**Original data, Monte Carlo data** 

I(t)

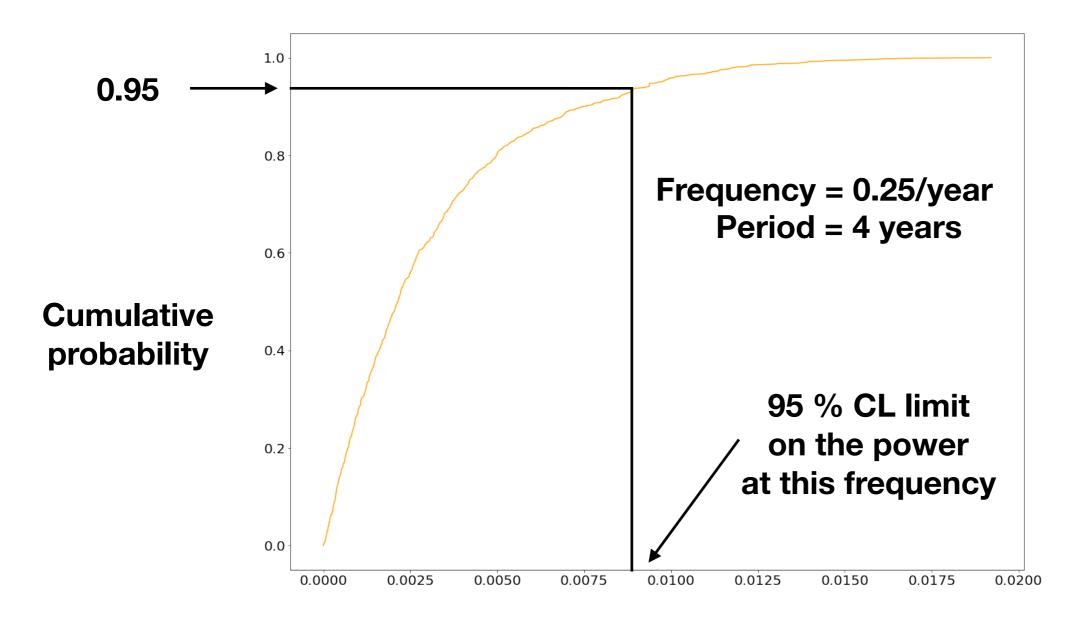
#### Lomb-Scargle periodogram



• Repeating this process N times allows us to estimate the probability distribution function (PDF) of power at each frequency

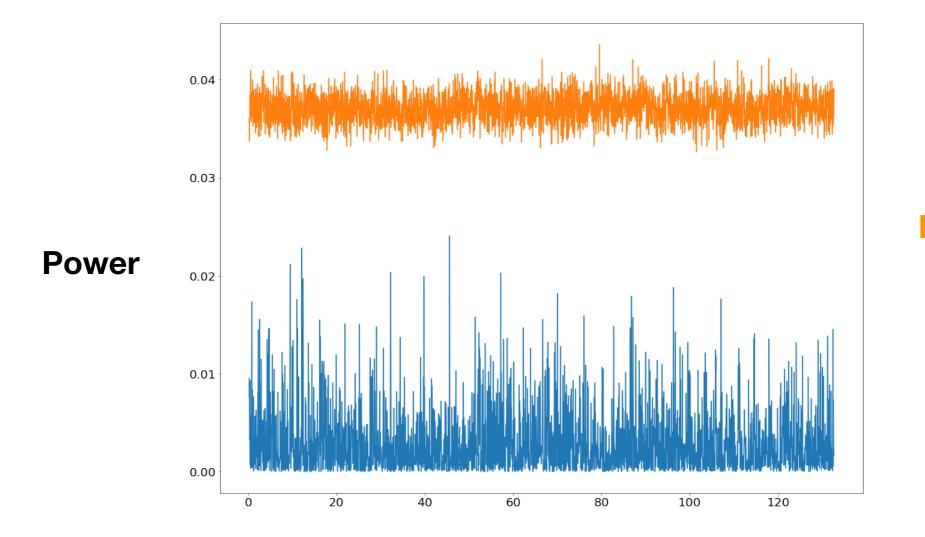


 This PDF integrates to give a cumulative probability distribution (CDF):



Power

• Repeating this at each frequency:



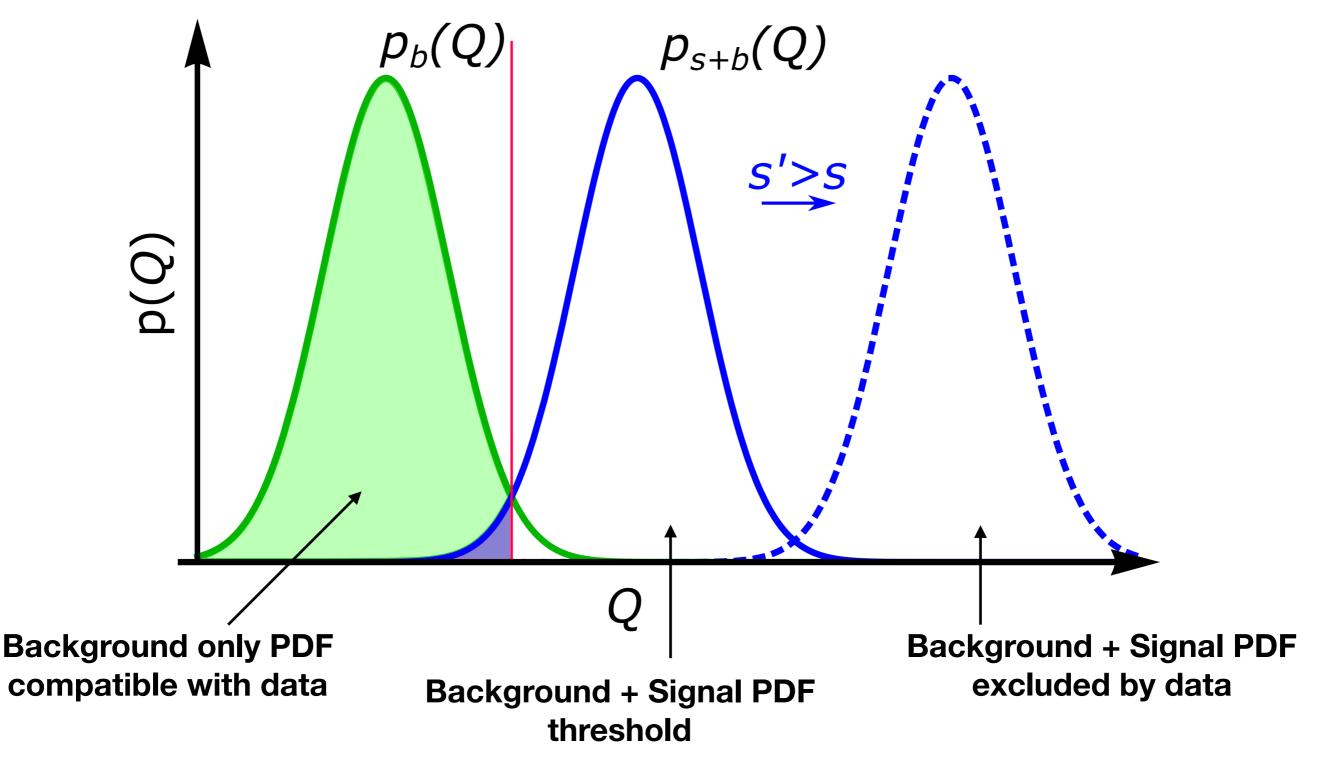
#### Original data, Monte Carlo limit

#### Frequency (1/year)

 We can see that the real data points (blue) are all below the 95 % CL limit (orange), and hence well-modelled by random noise

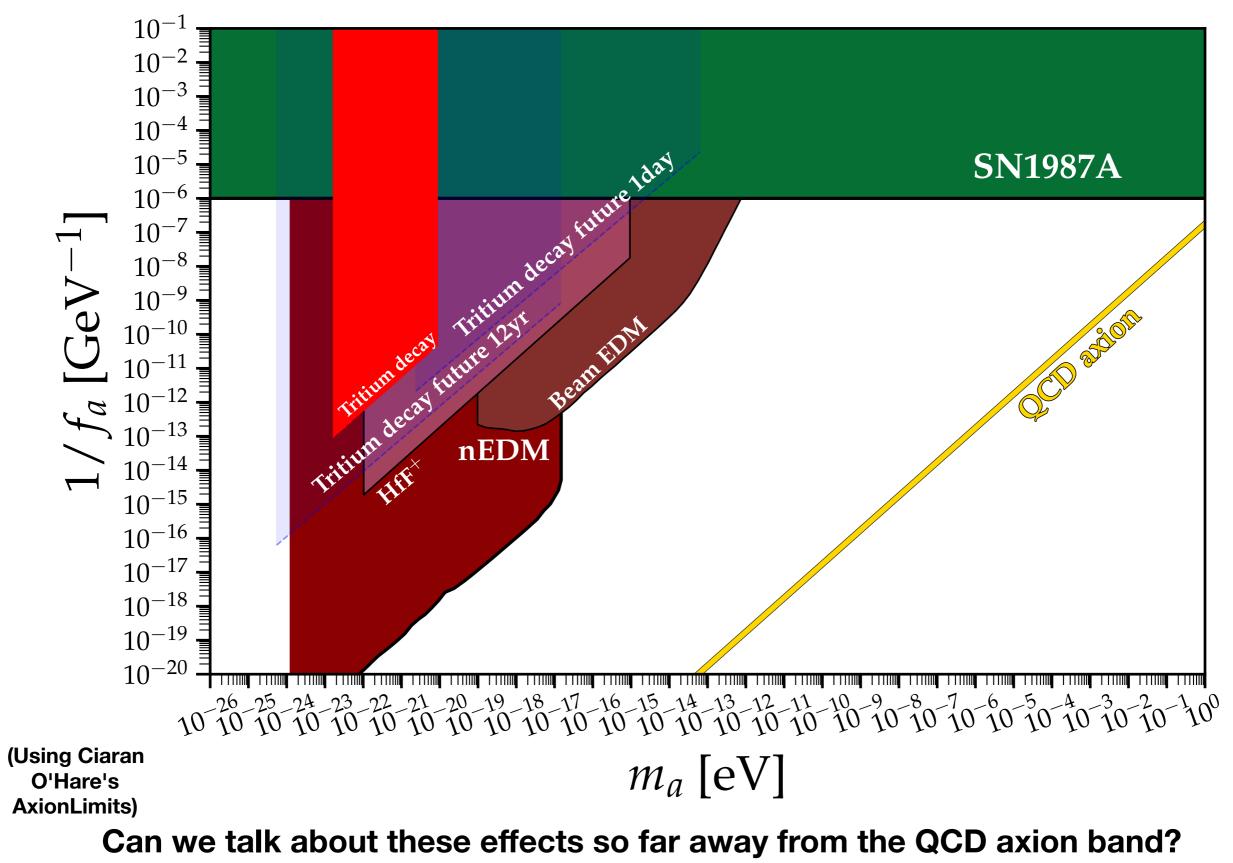
#### No evidence of non-random behaviour!

• Repeating this with an injected axion signal:

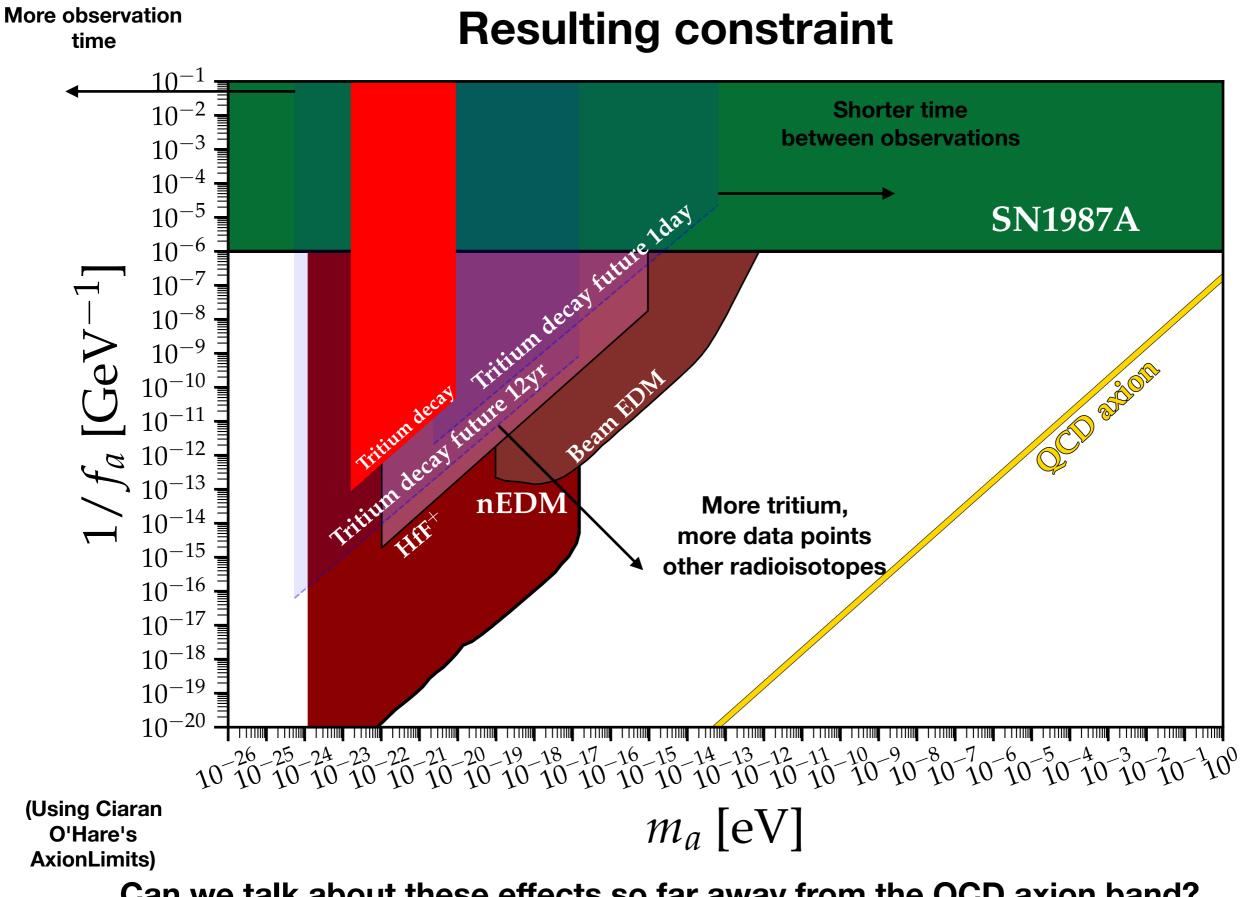


Varying the axion coupling allows us to find the threshold values

#### **Resulting constraint**



See Luca Di Luzio's talk from yesterday



Can we talk about these effects so far away from the QCD axion band? See Luca Di Luzio's talk from yesterday

## **Discussion and conclusions**

- We have examined reports of non-random behaviour in nuclear decay
- In 12 years of tritium decay data we find no evidence of this phenomenon
- We used the data to place constraints on axion DM
- Is nuclear decay random? Yes, probably...

More details in a paper soon to appear online!

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Thanks for listening!