

Study of Giant resonances in Exotic Nuclei by Means of Reactions in Inverse Kinematics

Muhsin N. Harakeh

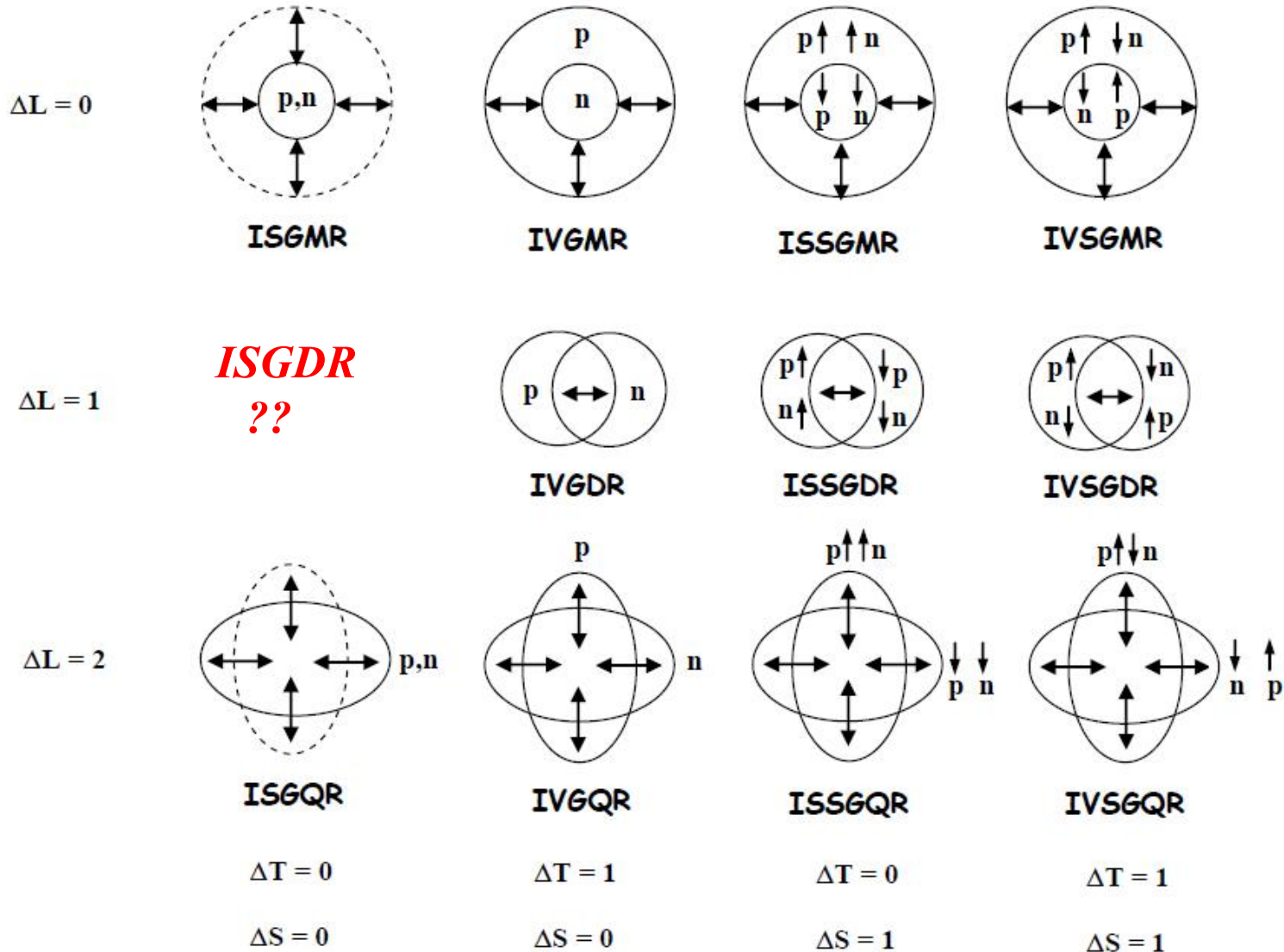
University of Groningen, the Netherlands

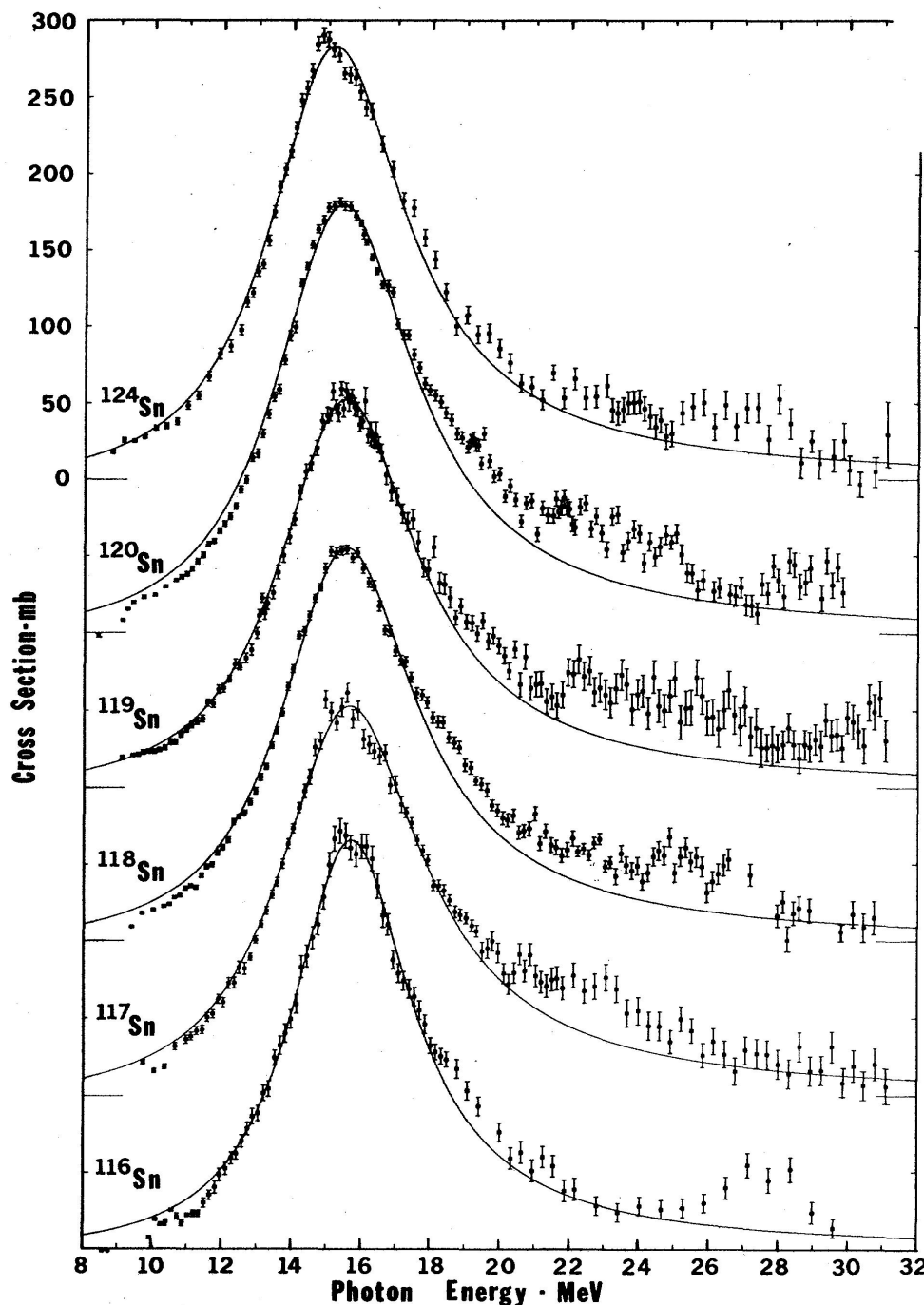
**3rd NUclear physics School for Young Scientists
(NUSYS)**

**Fudan University, Shanghai
6-12 August 2023**

Giant Resonances in hydrodynamic models

Coherent vibrations of nucleonic fluids (p & n; \uparrow & \downarrow) in a nucleus





Measurement of the giant dipole resonance with mono-energetic photons

B.L. Berman and S.C. Fultz
Rev. Mod. Phys. 47 (1975) 713

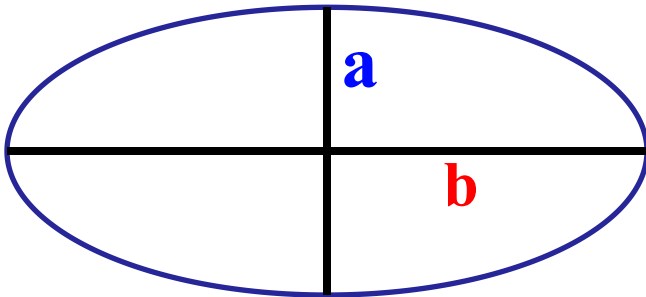
Nucleus	Centroid (MeV)	Width (MeV)
¹¹⁶ Sn	15.68	4.19
¹¹⁷ Sn	15.66	5.02
¹¹⁸ Sn	15.59	4.77
¹¹⁹ Sn	15.53	4.81
¹²⁰ Sn	15.40	4.89
¹²⁴ Sn	15.19	4.81

Photo-neutron cross section in deformed nuclei:

Deformed Nucleus

$$R(\theta, \phi) = R_0(1 + \beta_2 Y_{20}(\theta, \phi))$$

$$\beta_2(^{150}\text{Nd}) = 0.285(3)$$



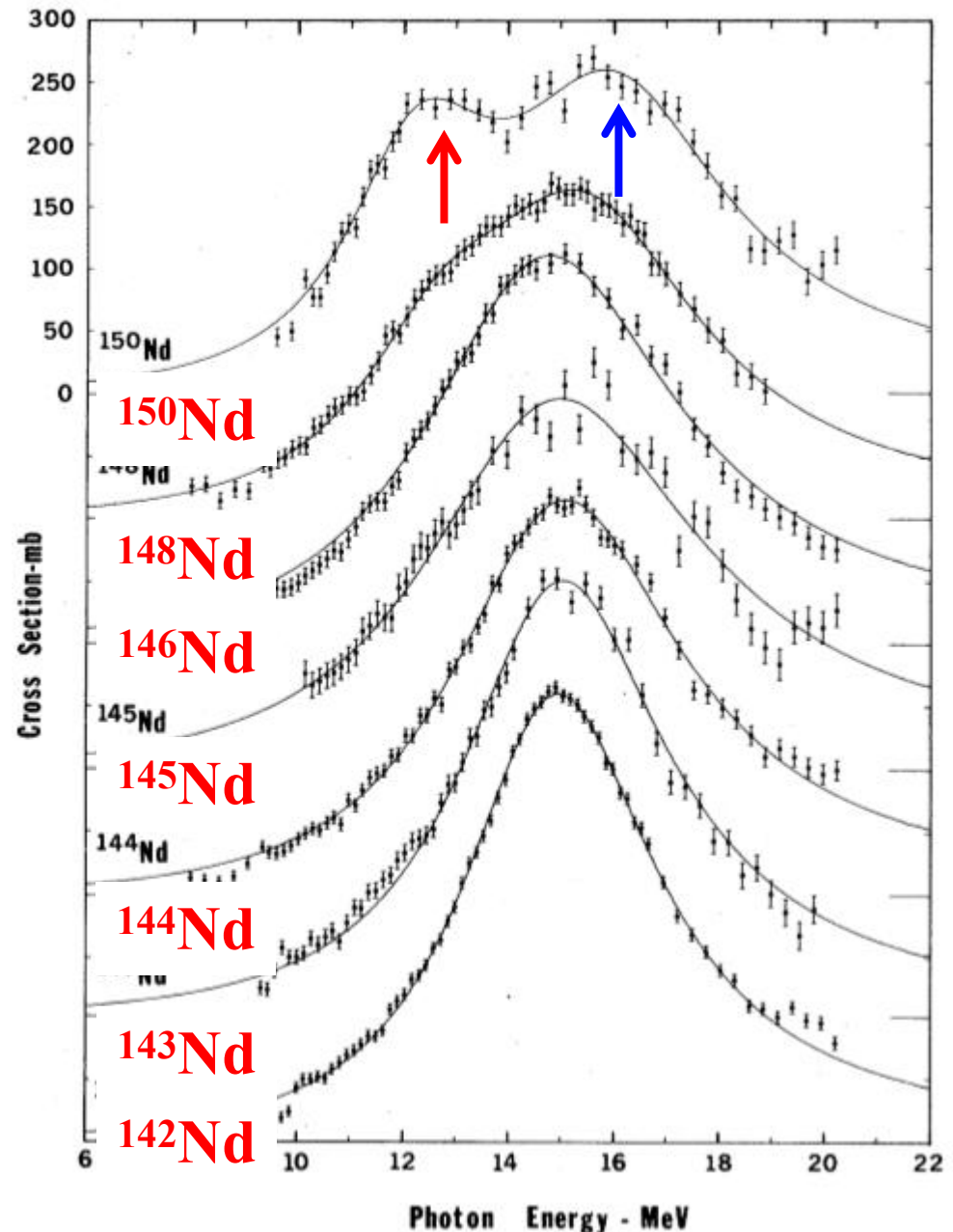
Excitation energies:

$$E_2/E_1 = 0.911\eta + 0.089$$

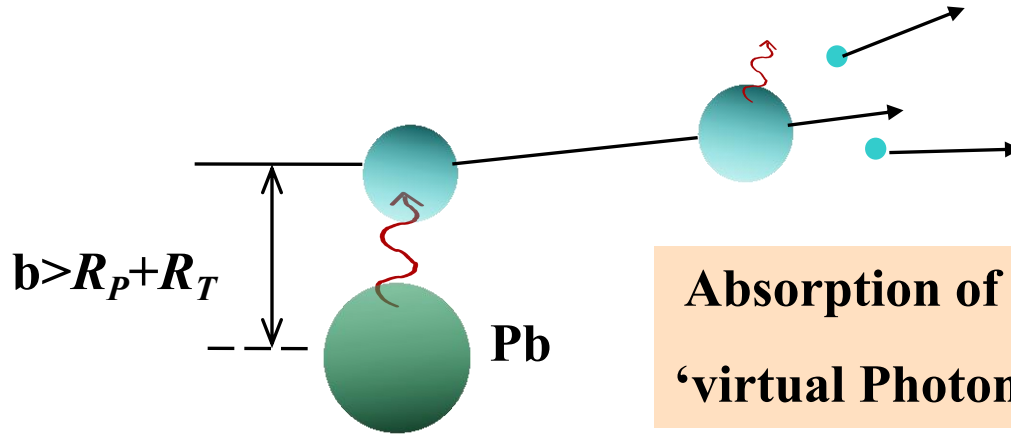
Where $\eta = b/a$

$$S_1/S_2 = 1/2$$

B. L. Berman and S. C. Fultz,
Rev. Mod. Phys. 47, 713 (1975)



Experimental Tool: Electromagnetic excitation at high energies

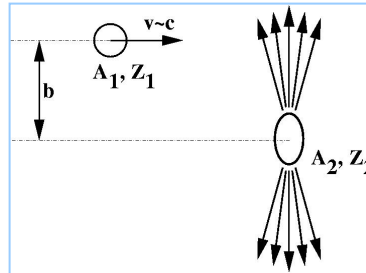


Absorption of
'virtual Photons'

$$\sigma_{e.m.} \sim Z^2$$

adiabatic cut-off:

$$E_{\max} = \frac{\hbar}{\tau} = \frac{\hbar c \gamma \beta}{b}$$

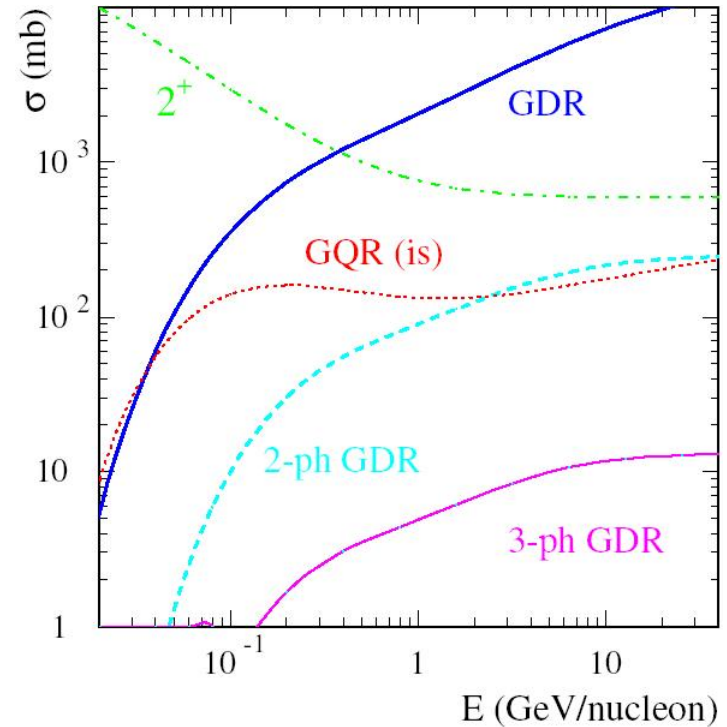


High velocities $v/c \approx 0.6-0.9$
 \Rightarrow High-frequency Fourier components

$$E_{\gamma, \max} \approx 25 \text{ MeV (@ 1 GeV/u)}$$

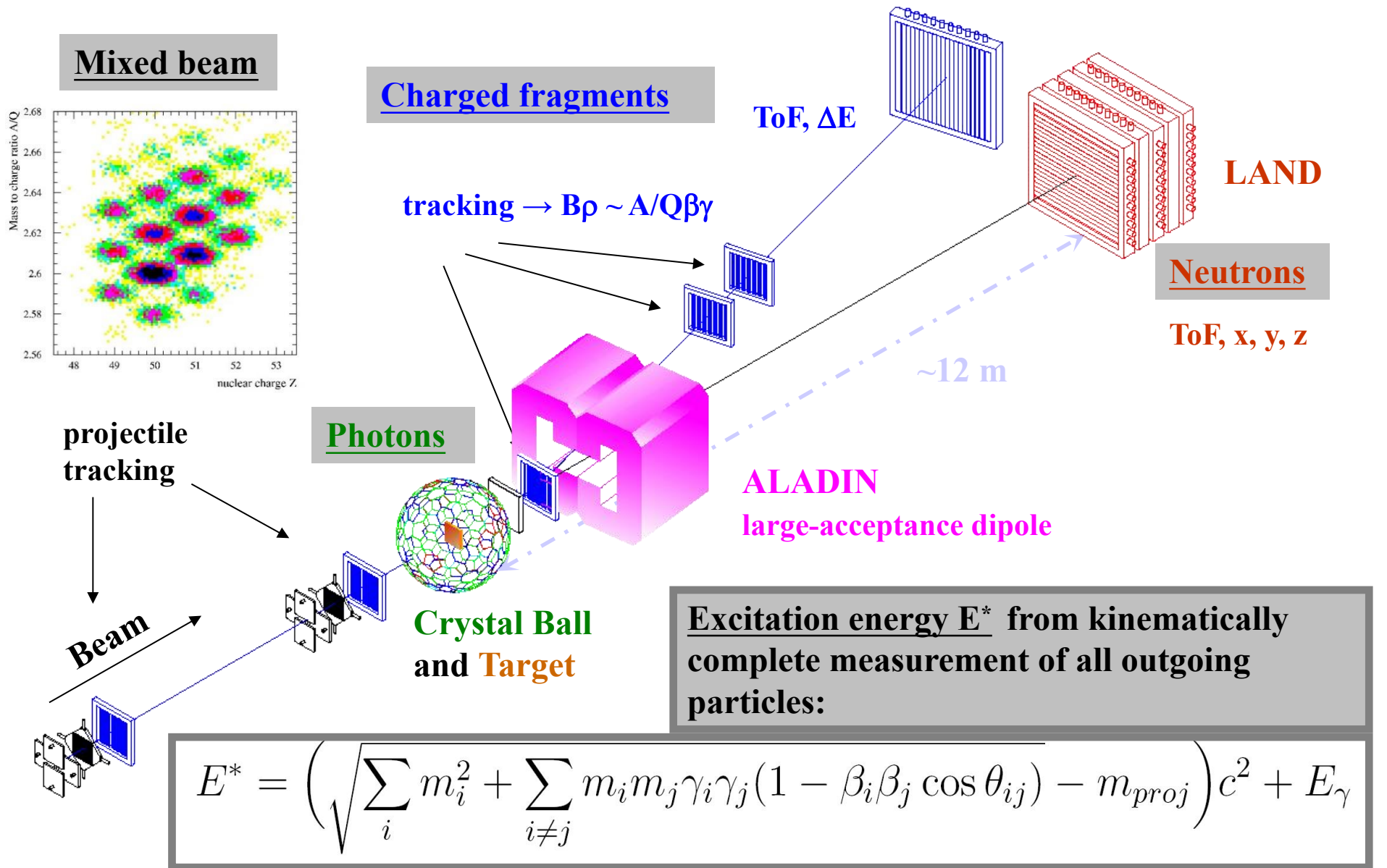
Semi-classical theory:

$$d\sigma_{e.m.} / dE = N_{\gamma}(E) s_{\gamma}(E)$$



Determination of 'photon energy' (excitation energy) via a kinematically complete measurement of the momenta of all outgoing particles (invariant mass)

Experimental Scheme: The LAND reaction setup @GSI

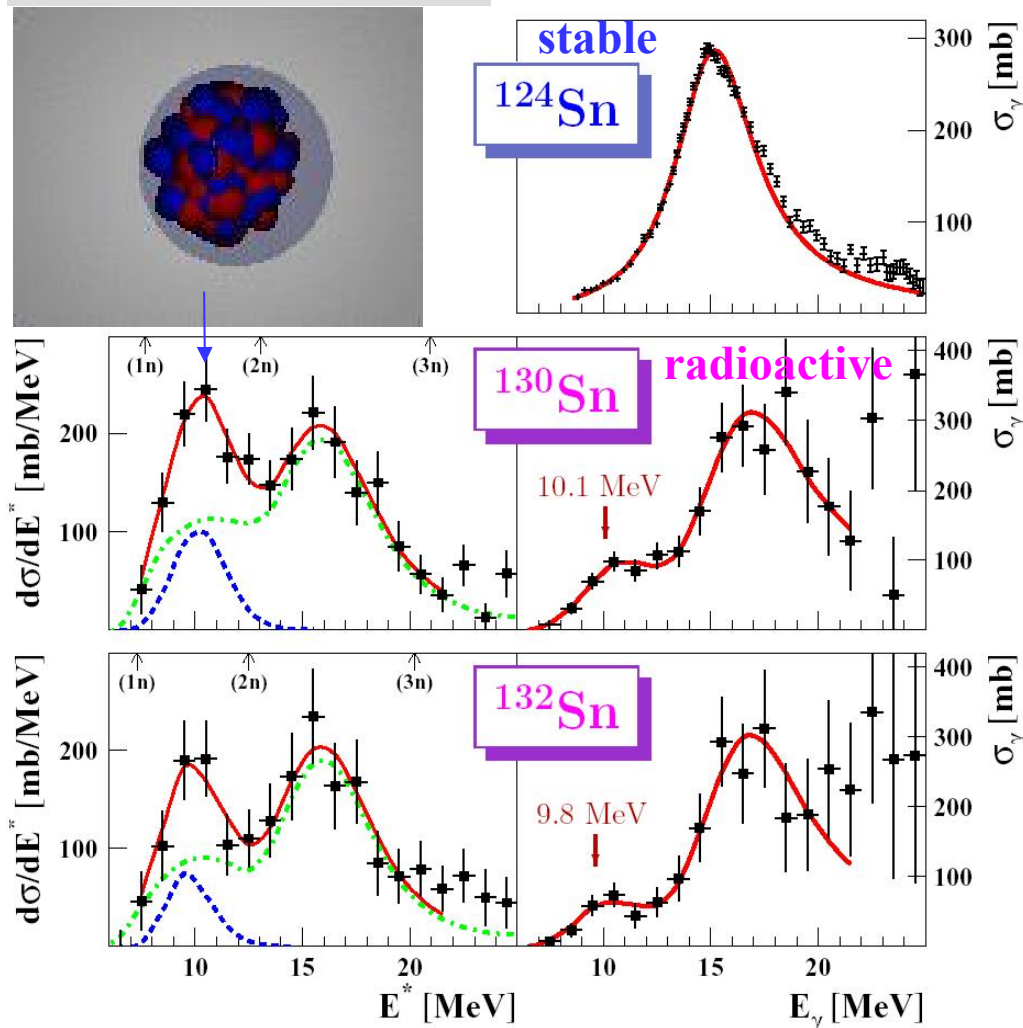


Dipole strength distributions in neutron-rich Sn isotopes

Electromagnetic-excitation
cross section

Photo-neutron cross section

$E(^A\text{Sn}) \sim 500$ MeV/nucleon



A	PDR		GDR		
	E_{centr} [MeV]	sum-rule fraction [%]	E_{centr} [MeV]	Γ [MeV]	sum-rule fraction [%]
^{124}Sn	-	-	15.3	4.8	116
^{130}Sn	10.1 (0.7)	7.0 (3.0)	15.9 (0.5)	4.8 (1.8)	145 (19)
^{132}Sn	9.8 (0.7)	4.0 (3.1)	16.1 (0.8)	4.7 (2.2)	125 (32)

Virtual photon spectra decrease as a function of virtual photon energy \Rightarrow correct for that!

PDR

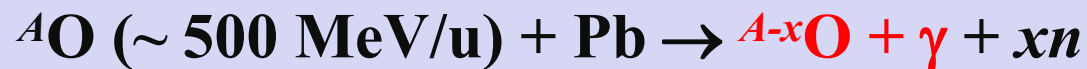
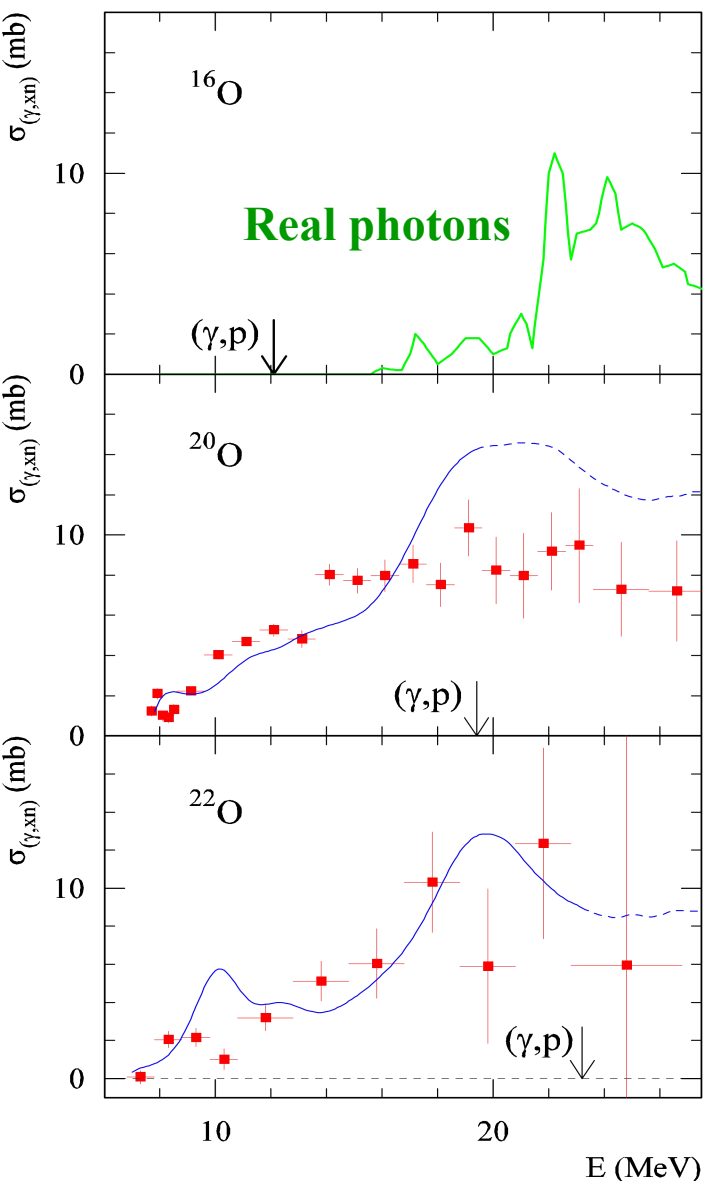
- located at 10 MeV
- exhausts a few % TRK sum rule
- in agreement with theory

GDR

- no deviation from systematics

P. Adrich *et al.*, PRL 95 (2005) 132501

Dipole Strength Distribution of n-Rich Nuclei



$N-Z=0$

\Rightarrow Photo-neutron cross sections from virtual photons

$N-Z=4$

\Rightarrow Low-lying dipole strength

\Rightarrow Fragmentation of GDR strength

? Collective soft mode ?

$N-Z=6$

— Large-scale shell model calculation

H. Sagawa, T. Suzuki,

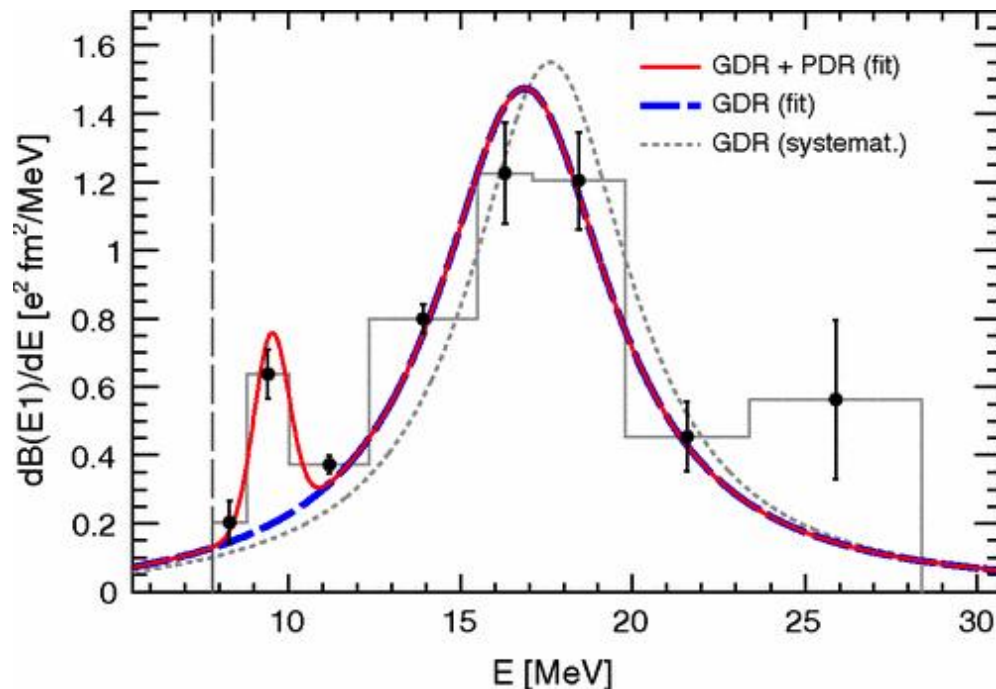
Phys. Rev. C 59 (1999) 3116

Data: LAND-FRS@GSI

A. Leistenschneider *et al.*, Phys. Rev. Lett. 86 (2001) 5442

Dipole strength distributions in ^{68}Ni

^{68}Ni incident energy 600 MeV/u

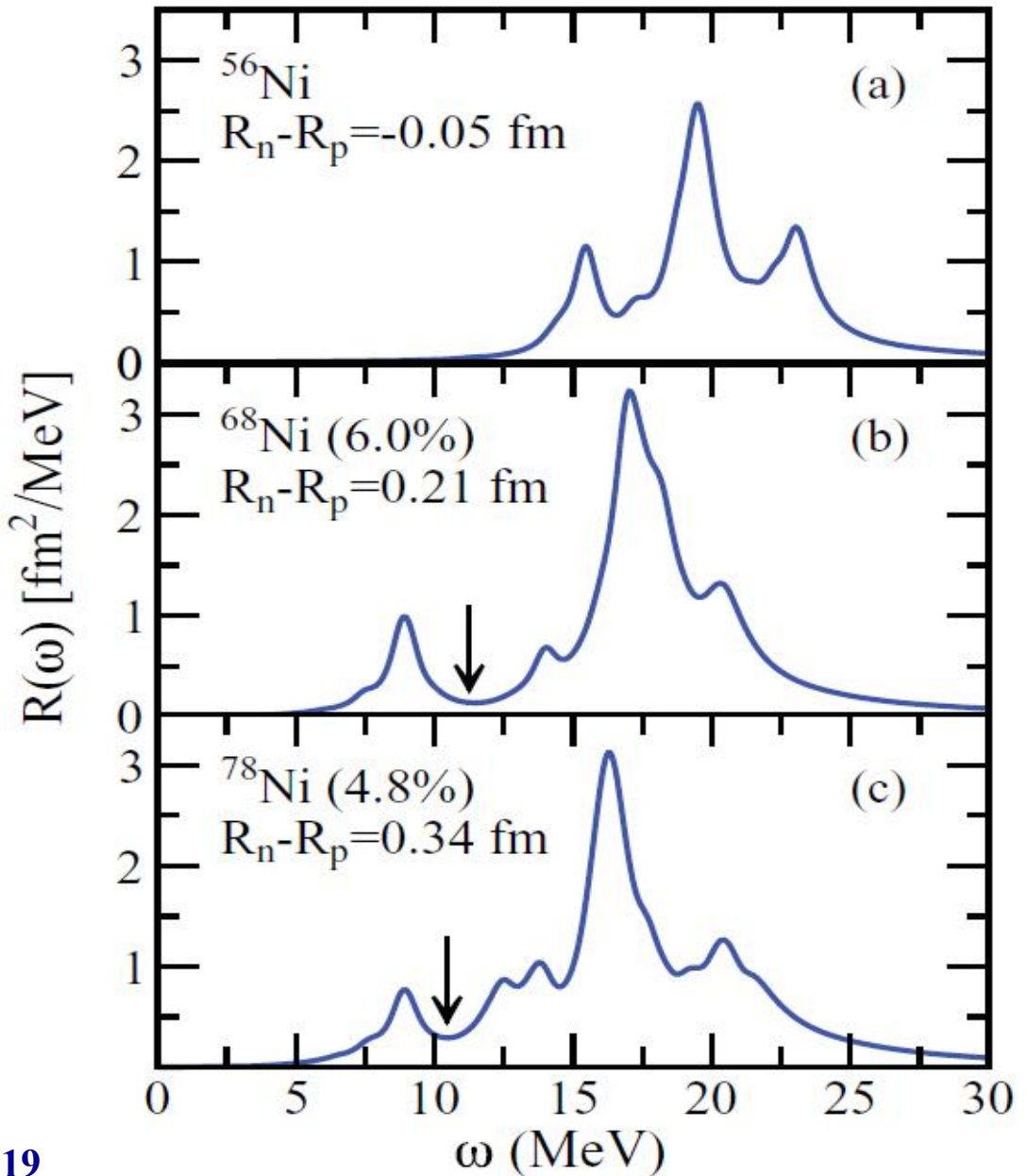


E1 strength distribution (histogram and black data points) with IVGDR+PDR fit function (solid red line). The IVGDR contribution (dashed blue line) and the IVGDR from systematics (dotted gray line) are shown for reference. The neutron threshold is indicated by the dashed vertical line at 7.792 MeV.

	E (MeV)	Γ (MeV)	%EWSR
IVGDR	17.1(2)	6.1(5)	98(7)
PDR	9.55(17)	0.51(13)	2.8(5)

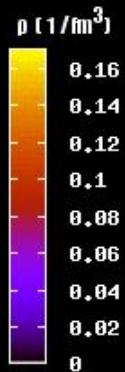
D. Rossi *et al.*, Phys. Rev. Lett. 111 (2013) 242503

Distribution of isovector dipole strength for the three closed-(sub)shell nickel isotopes ^{56}Ni , ^{68}Ni , and ^{78}Ni calculated in HF-plus-RPA using the FSUGold interaction parameter set.



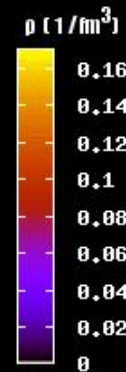
J. Piekarewicz, PRC 83 (2011) 034319

$L=0$



ISGMR

$L=1$

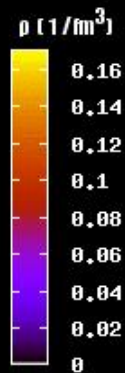


ISGDR

Macroscopic
hydrodynamic
description of
 $L=0-3$
isoscalar giant
multipole
resonances

ISGMR & ISGDR compression modes

$L=2$



ISGQR

$L=3$



ISGOR

Drawing by
Masatoshi Itoh

Compression Modes of Nuclei

Compression modes: **ISGMR, ISGDR**

In Constrained and Scaling Models:

$$E_{ISGMR} = \hbar \sqrt{\frac{K_A}{m \langle r^2 \rangle}}$$

$$E_{ISGDR} = \hbar \sqrt{\frac{7}{3} \frac{K_A + \frac{27}{25} \varepsilon_F}{m \langle r^2 \rangle}}$$

ISGMR (T=0, L=0)



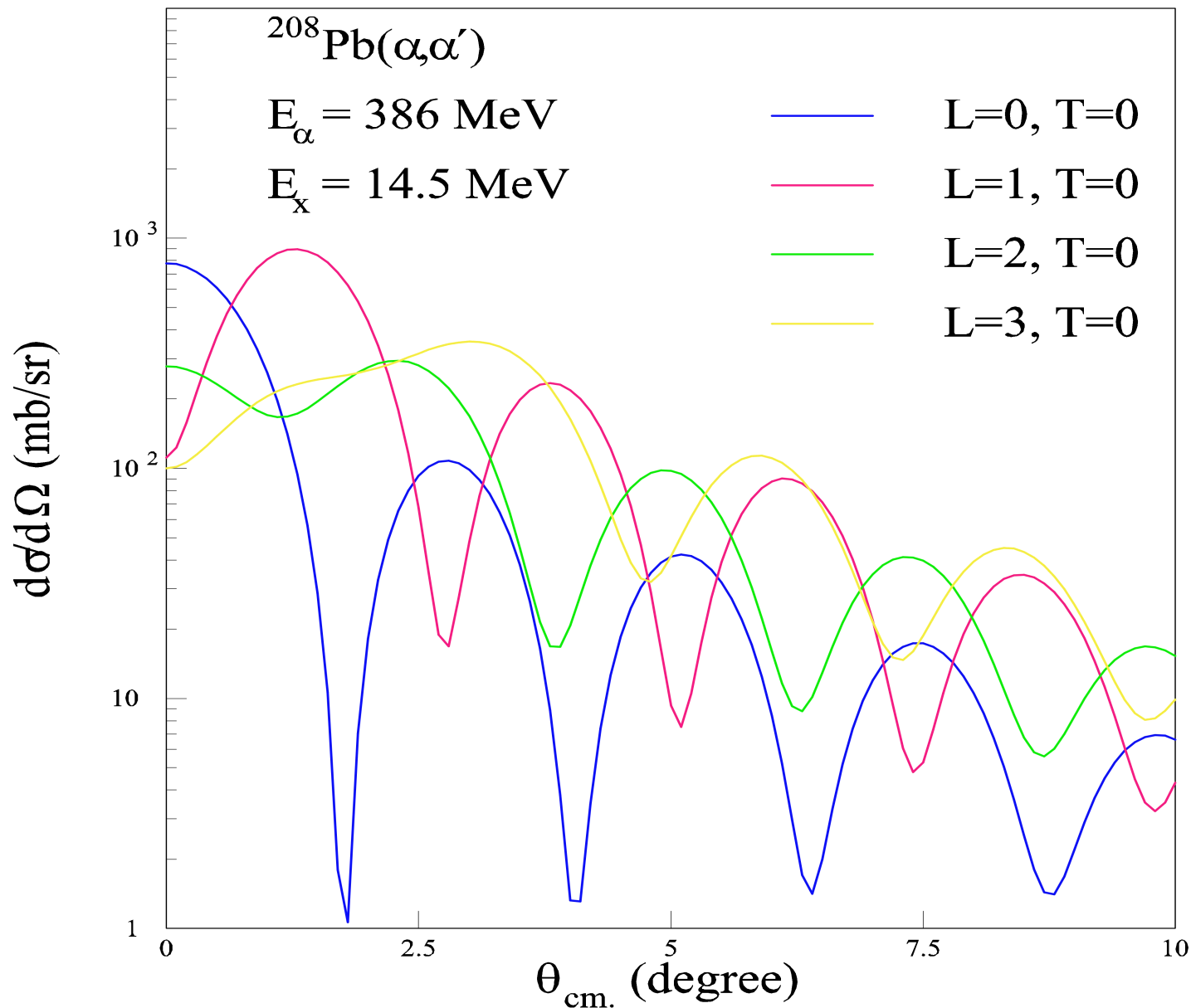
ISGDR (T=0, L=1)



ε_F is the Fermi energy and K_A is the nucleus incompressibility

→ ISGMR, ISGDR \Rightarrow Incompressibility, symmetry energy

$$K_A = K_{vol} + K_{surf} A^{-1/3} + K_{sym} ((N-Z)/A)^2 + K_{Coul} Z^2 A^{-4/3}$$



ISGMR, ISGDR

ISGQR, HEOR

100 % EWSR

**At $E_x = 14.5$
MeV**

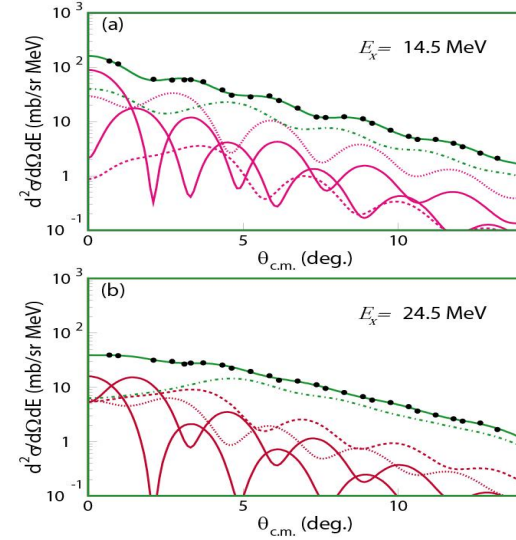
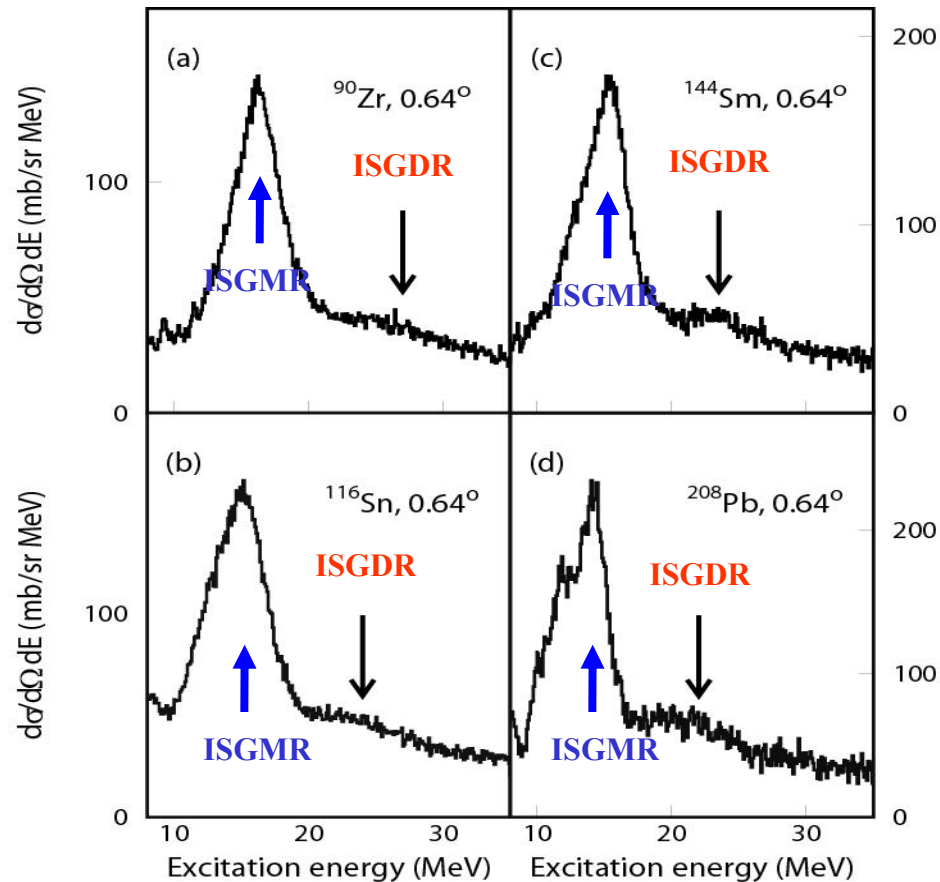
Experiments at RCNP, Osaka University

- (α, α') reaction at ~ 400 MeV
 - High-resolution spectrometer “Grand Raiden”



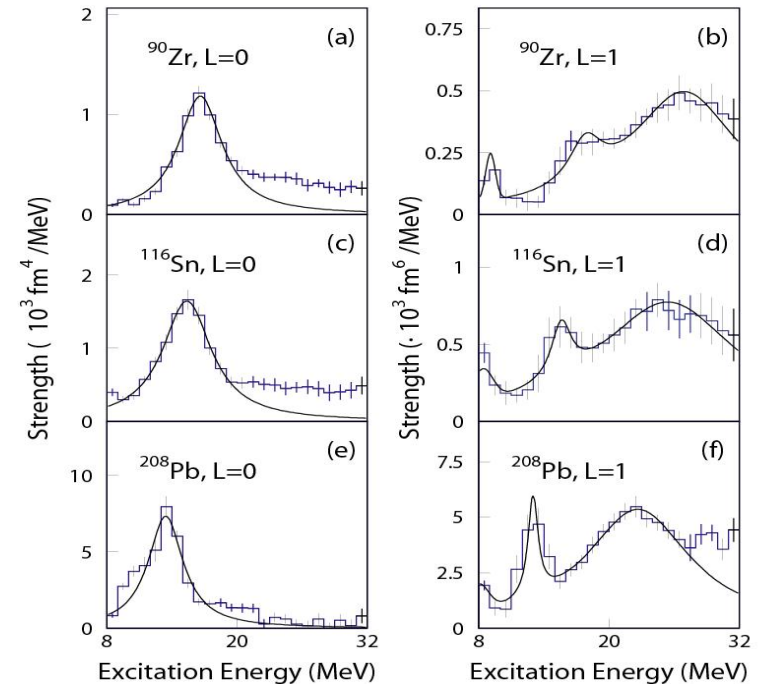
Uchida *et al.*,
Phys. Lett. B557 (2003) 12
Phys. Rev. C69 (2004) 051301

(α, α') spectra at 386 MeV



^{116}Sn

MDA results for $L=0$ and $L=1$



From GMR data on ^{208}Pb and ^{90}Zr ,

$$K_{\infty} = 240 \pm 10 (\pm 20) \text{ MeV}$$

[See, *e.g.*, G. Colò *et al.*, Phys. Rev. C 70 (2004) 024307]

**This number is consistent
with both ISGMR and ISGDR Data
and
with non-relativistic and relativistic calculations.**

Z.Z. Li, Y.F. Niu, and G. Colò, Phys. Rev. Lett. accepted
QRPA+QPVC for ^{208}Pb , $^{112-124}\text{Sn}$ and ^{48}Ca
 $K_{\infty} = 226 \text{ MeV}$ and 229 MeV

$$K_A = K_{vol} + K_{surf}A^{-1/3} + K_{sym}((N-Z)/A)^2 + K_{coul}Z^2A^{-4/3}$$

$$K_A \sim K_{vol}(1 + cA^{-1/3}) + K_{\tau}((N - Z)/A)^2 + K_{Coul}Z^2A^{-4/3}$$

$$K_A - K_{Coul}Z^2A^{-4/3} \sim K_{vol}(1 + cA^{-1/3}) + K_{\tau}((N - Z)/A)^2$$

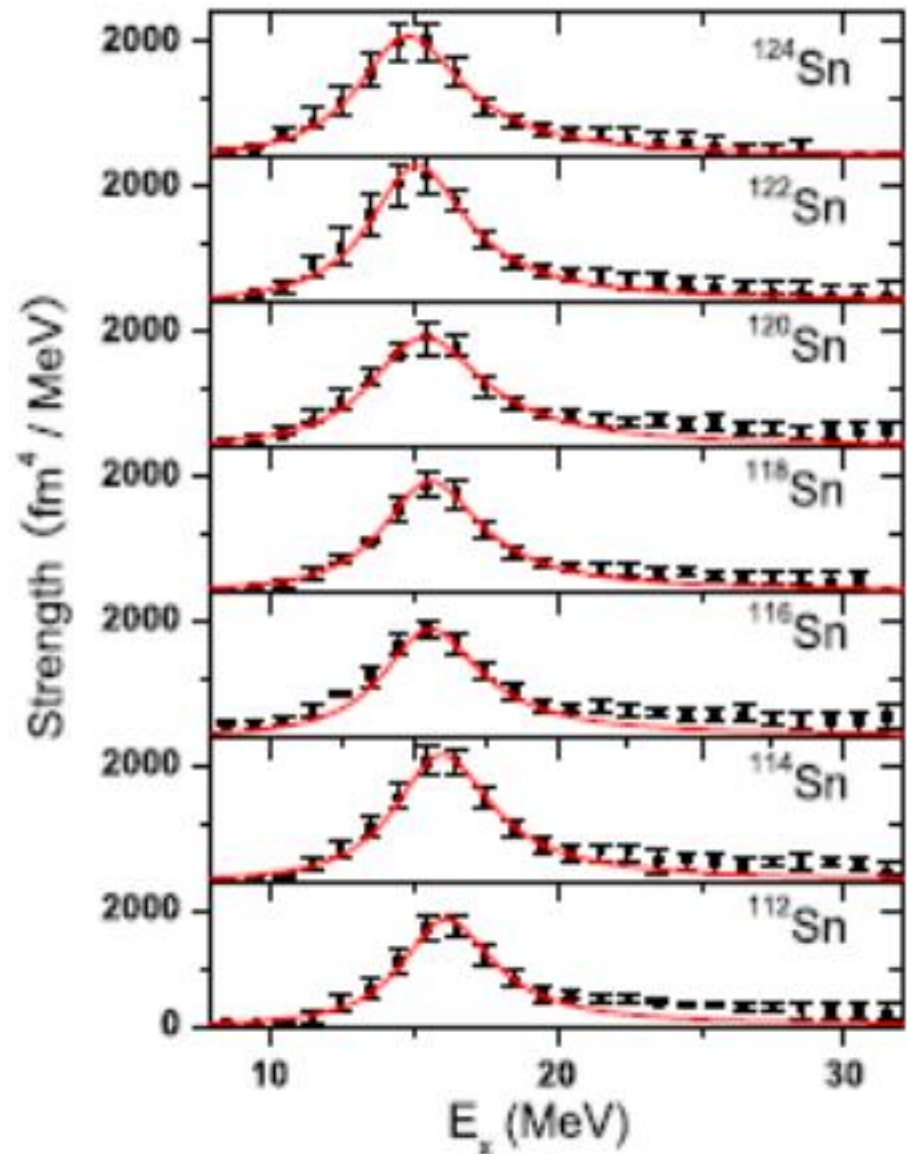
$$\sim \text{Constant} + K_{\tau}((N - Z)/A)^2$$

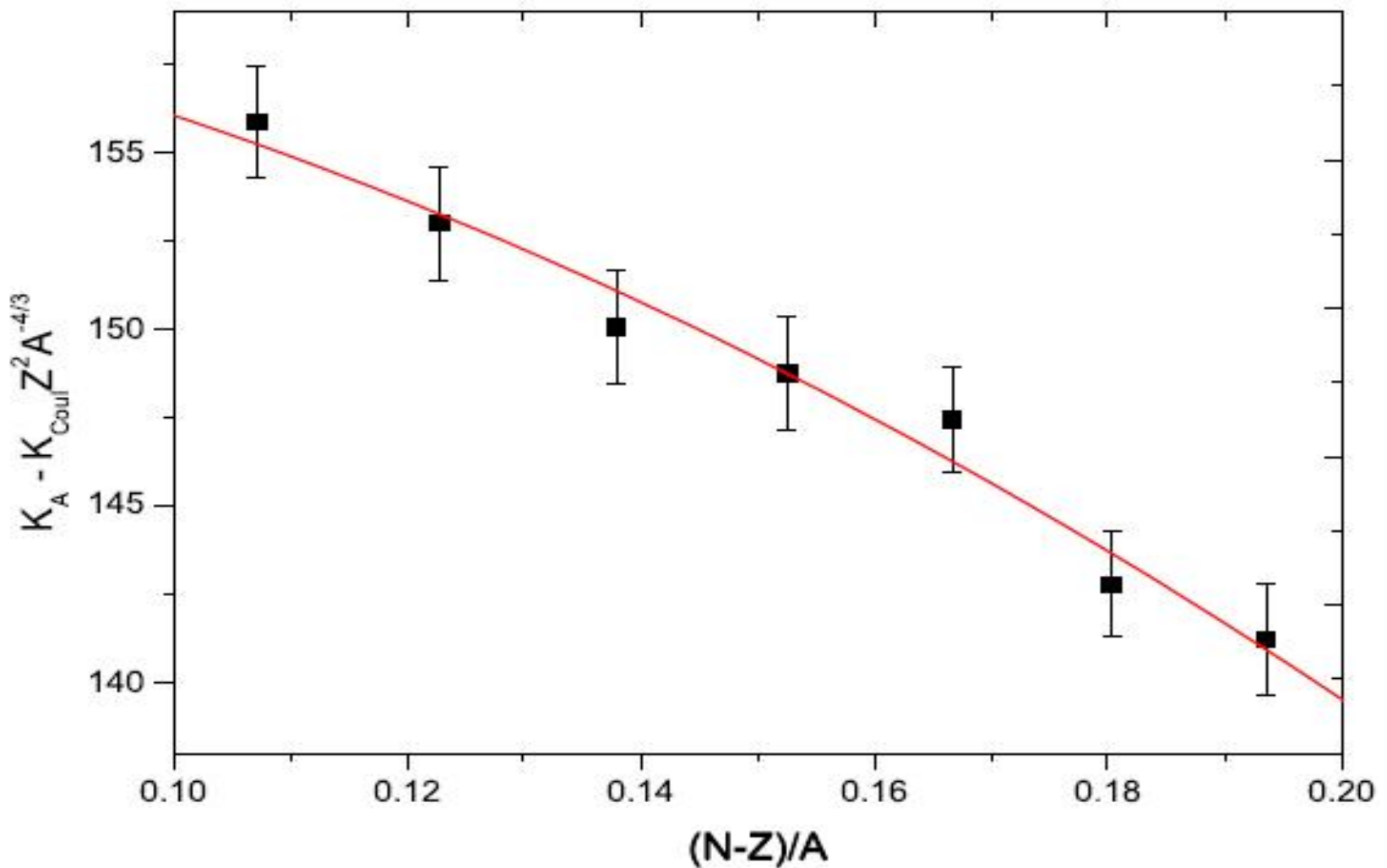
We use $K_{Coul} = -5.2$ MeV (from Sagawa)

$$(N - Z)/A$$

$$^{112}\text{Sn} - ^{124}\text{Sn}: \mathbf{0.107 - 0.194}$$

Isoscalar GMR strength distribution in Sn-isotopes obtained by Multipole Decomposition Analysis of singles spectra obtained in $^A\text{Sn}(\alpha, \alpha')$ measurements at incident energy 400 MeV and angles from 0° to 9°





Sn isotopes $\Rightarrow K_\tau = -550 \pm 100 \text{ MeV}$

Outlook

Radioactive ion beams will be available at energies where it will be possible to study excitation of ISGMR and ISGDR

RIBF/RIKEN, GSI/FAIR, GANIL/SPIRAL2, FRIB/NSCL

E.g., Determine ISGMR and ISGDR in unstable Sn nuclei.

$A = 106$ to 134 possible

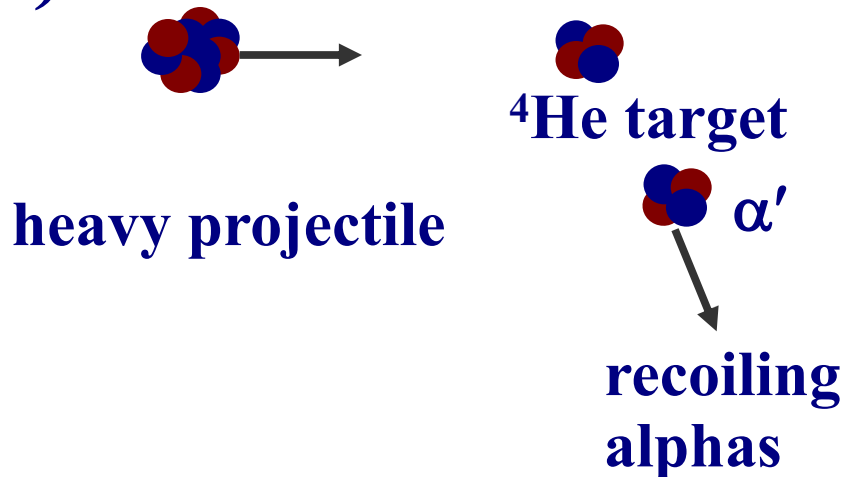
\Rightarrow A more precise determination of K_τ

Nuclear structure studies with reactions in inverse kinematics

Possible at GSI/FAIR, RIKEN/RIBF, GANIL and
NSCL/FRIB

(beam energies of 50-100 MeV/u are needed!)

(α, α')



Approach
measure the recoiling alphas

The diagram shows a cluster of four red and blue spheres with an arrow pointing to the right, labeled 'heavy ejectile'.

Inconvenience:
difficulty to detect the low-
energy alphas

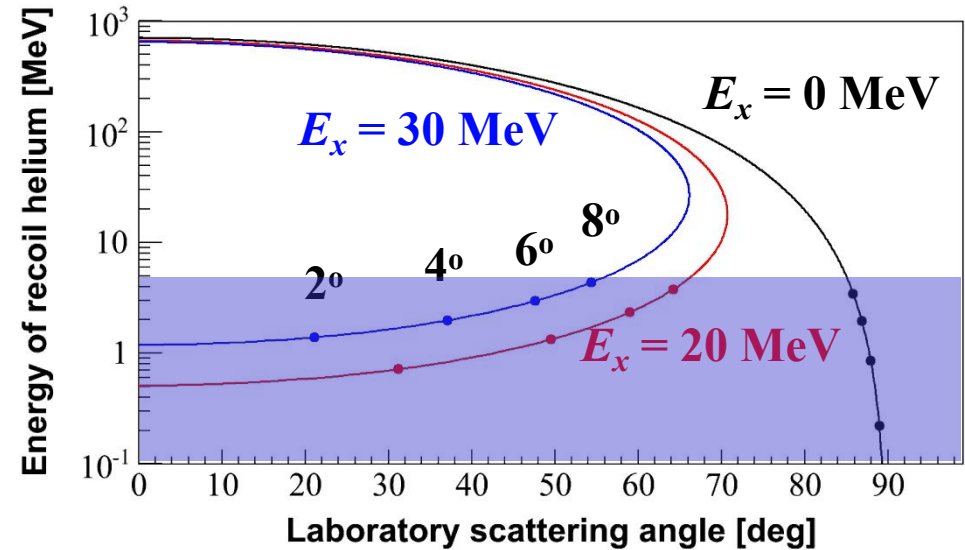
Nuclear structure studies with reactions in inverse kinematics

Challenges with exotic beams

- Inverse kinematics

$^{56}\text{Ni}(\alpha, \alpha')^{56}\text{Ni}^*$
 $\alpha = \text{Target}$
 $^{56}\text{Ni} = \text{Projectile}$

Secondary Beam
 ^{56}Ni at 50 MeV/u



- Intensity of exotic beams is very low ($\sim 10^4 - 10^5$ pps)
- To get reasonable yields thick target is needed
- Very low energy (\sim sub MeV) recoil particle will not come out of the thick target

Active target

A gas detector where the target gas also acts as a detector

Advantages

- **Good angular coverage**
- **Effective target thickness can be increased without much loss of resolution**
- **Detection of very low-energy recoil particle is possible**

Disadvantages

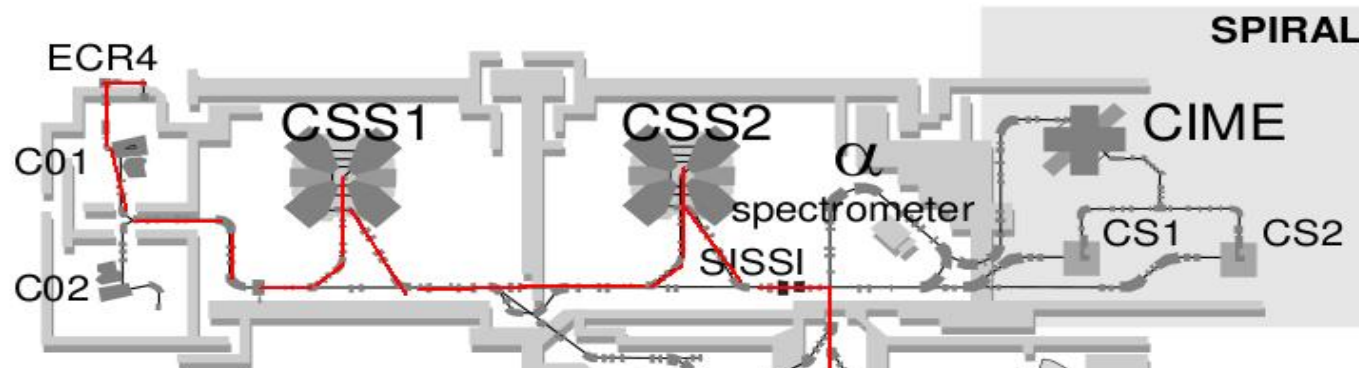
- **Low beam intensities ($\sim 10^4 - 10^5$ pps)**
- **Very small recoil energies for low q**

MAYA active-target detector at GANIL

Primary Beam
 ^{58}Ni at 75 MeV/u

Primary Target: ^9Be
(thickness 525.6 μm)

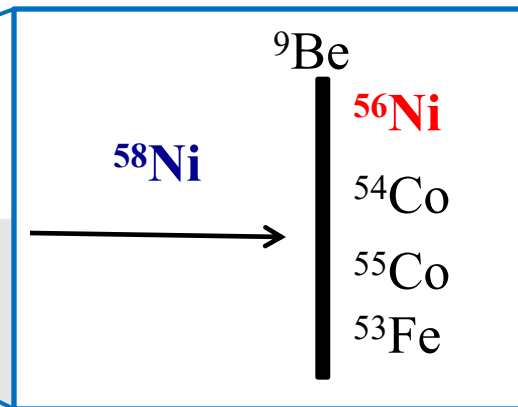
Secondary Beam
 ^{56}Ni at 50 MeV/u



GANIL Facility

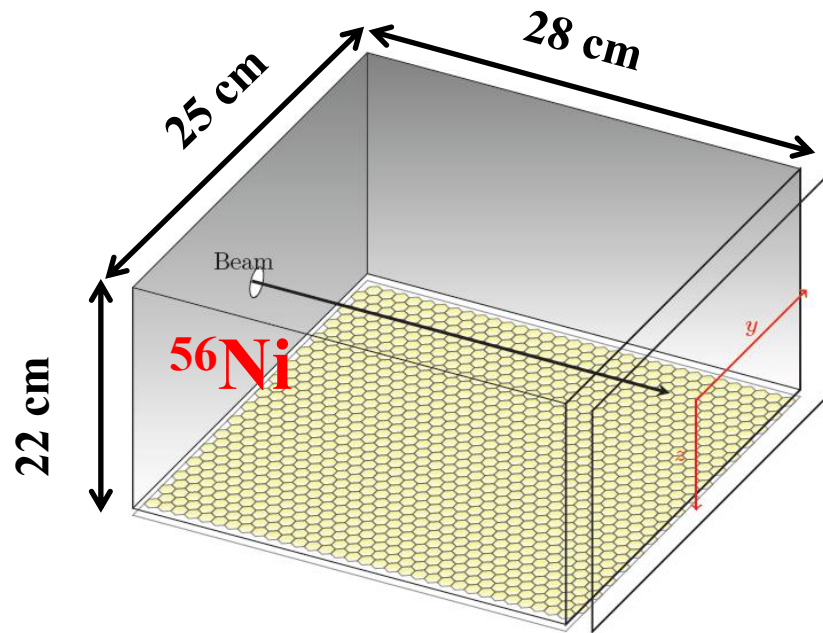
Primary target ^9Be

95-97% ^{56}Ni
 $\sim 10^4$ pps
 $^{56}\text{Ni}(\alpha, \alpha')^{56}\text{Ni}^*$

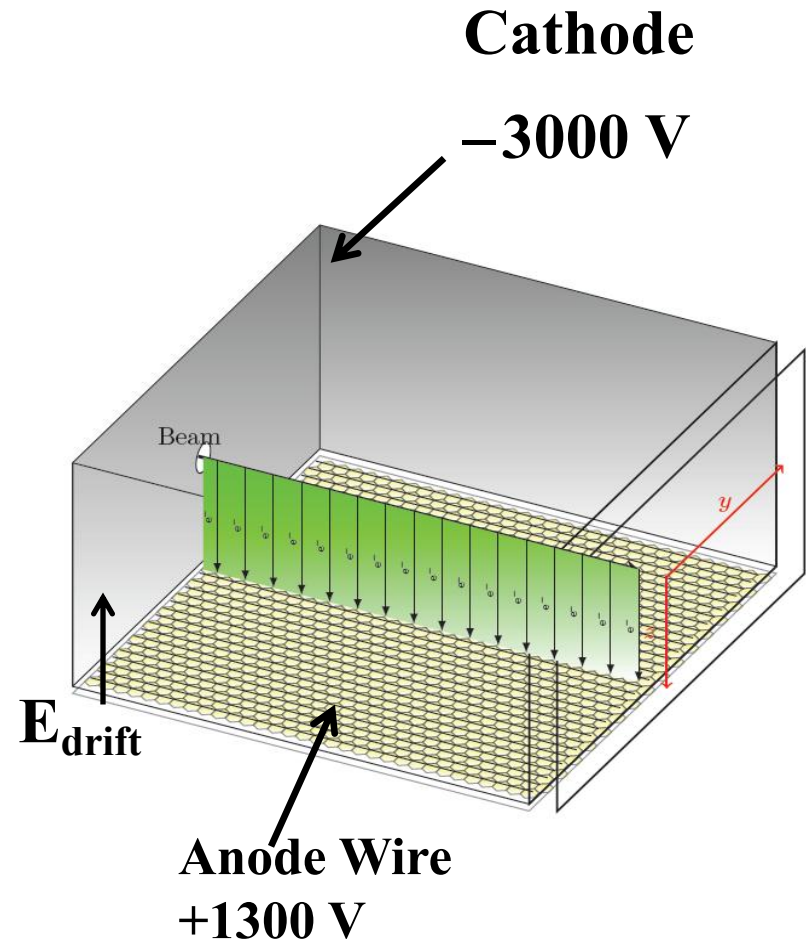


MAYA setup

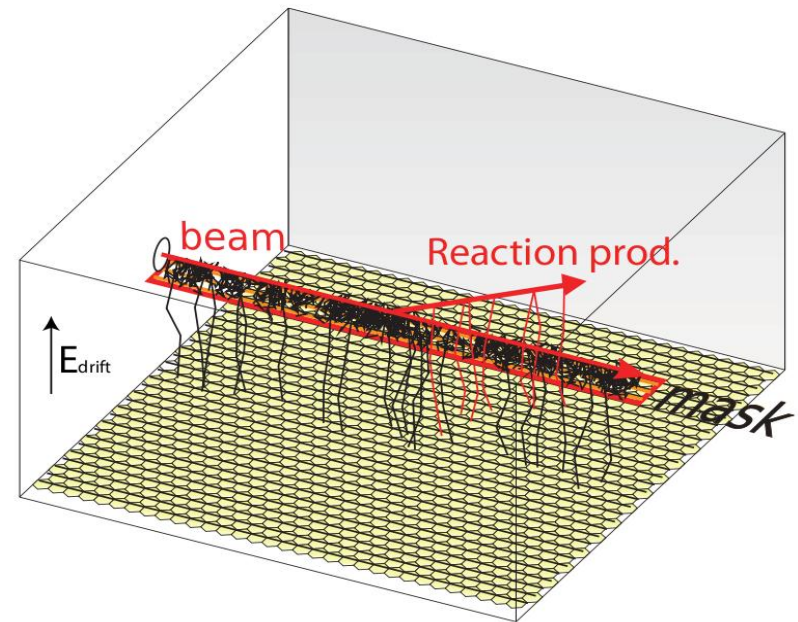
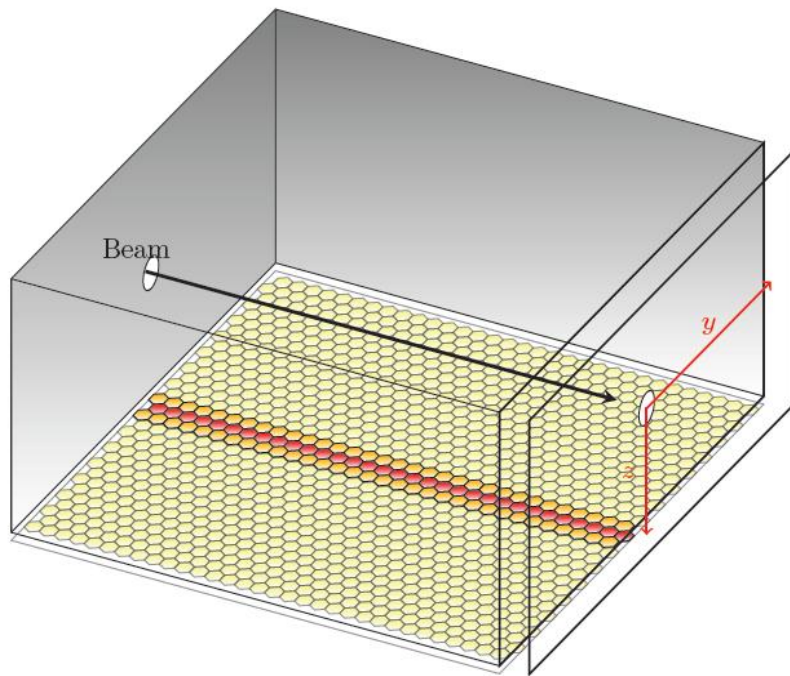
MAYA setup (I)



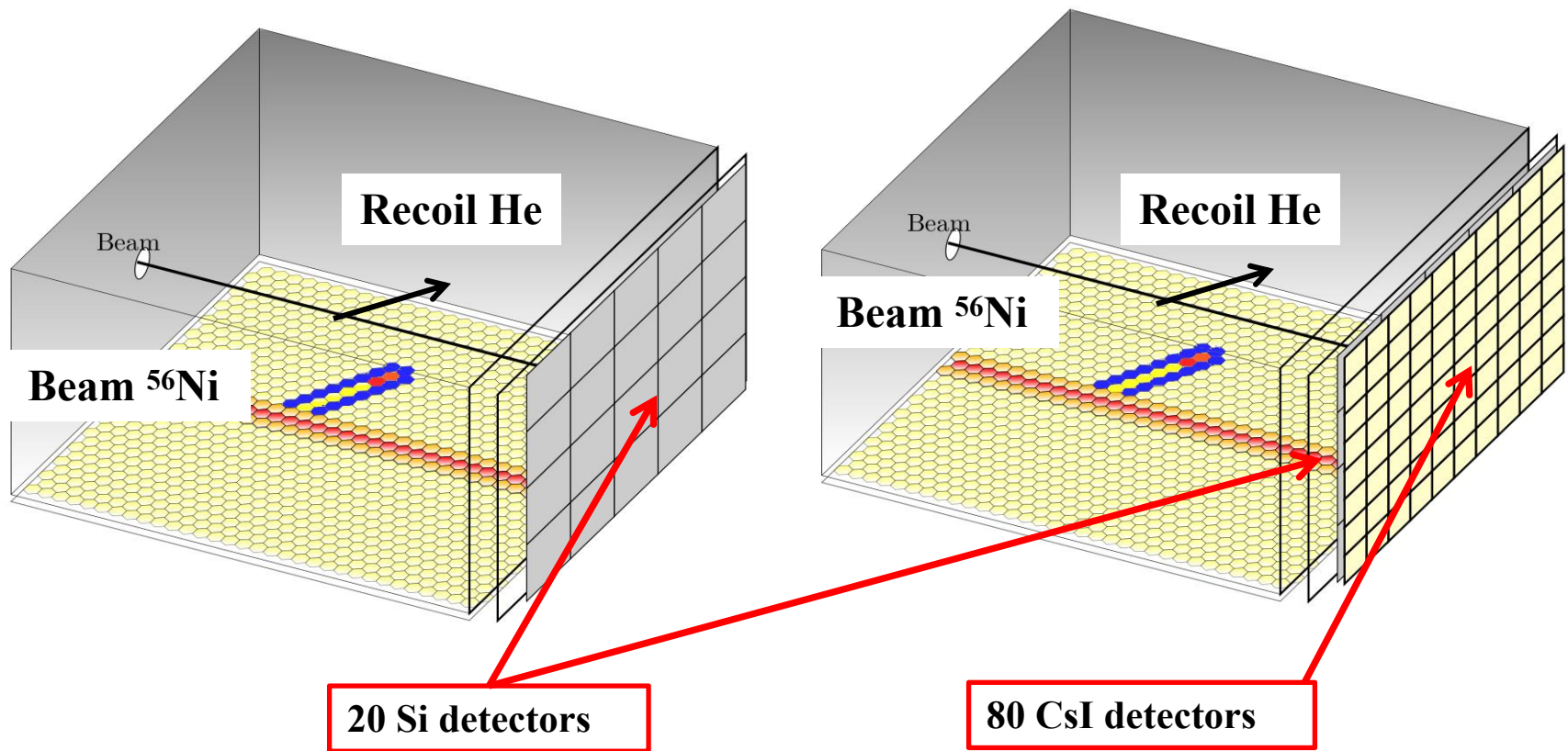
500 mbar
95% He and 5% CF_4

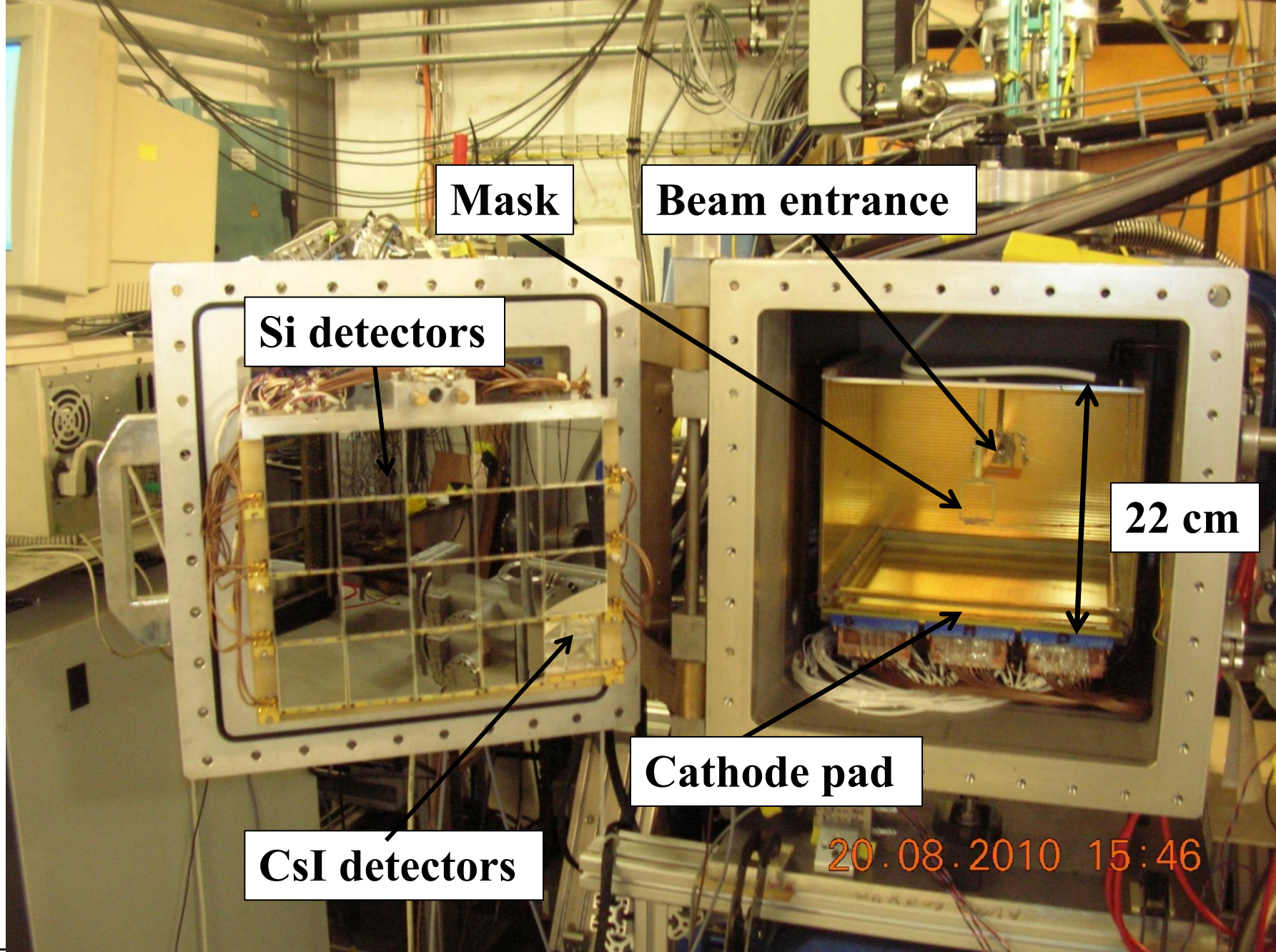


MAYA setup (II): Electrostatic mask

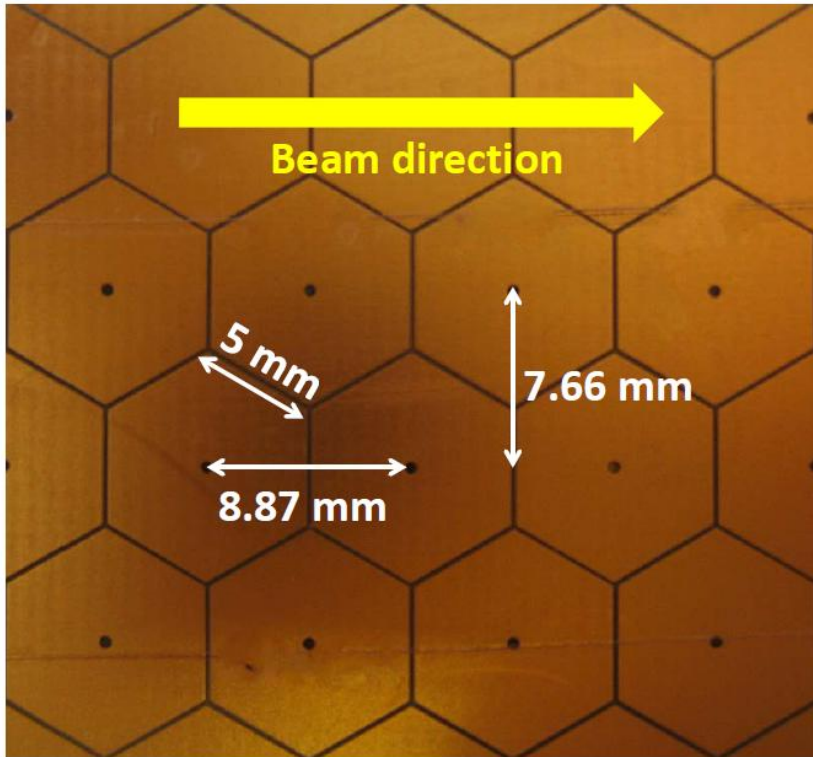


MAYA setup (III): Si/CsI telescope

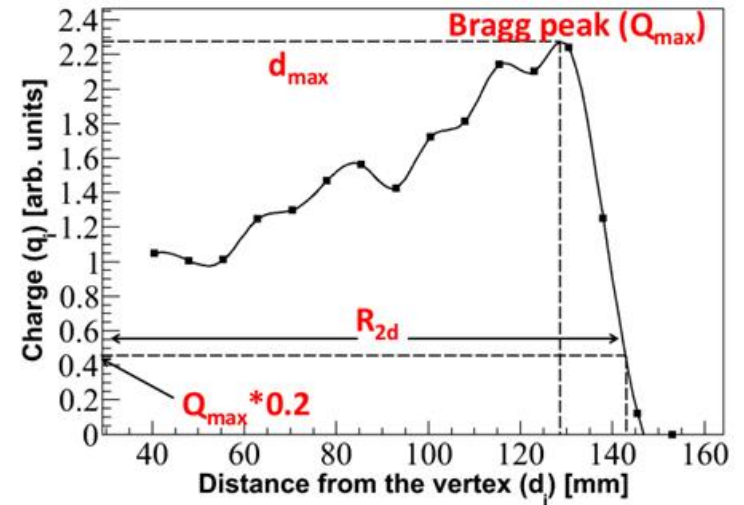
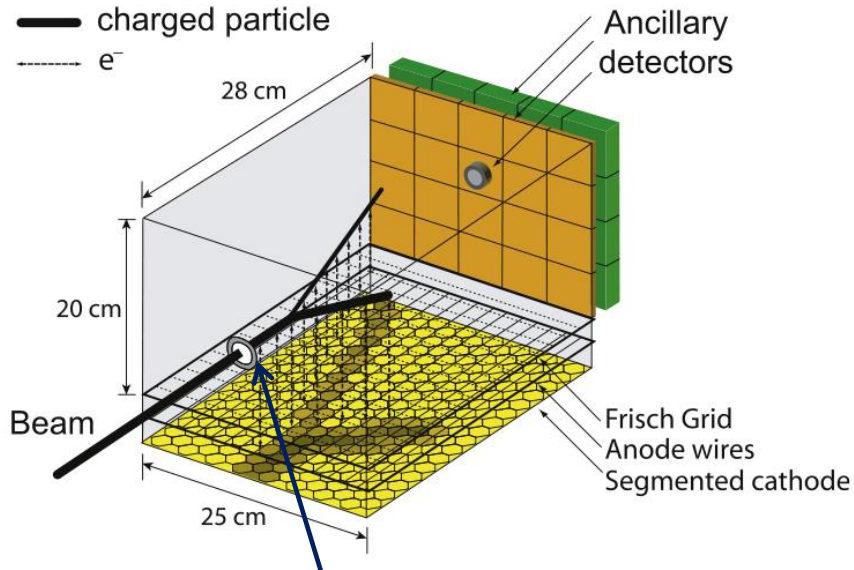




Cathode pad in MAYA



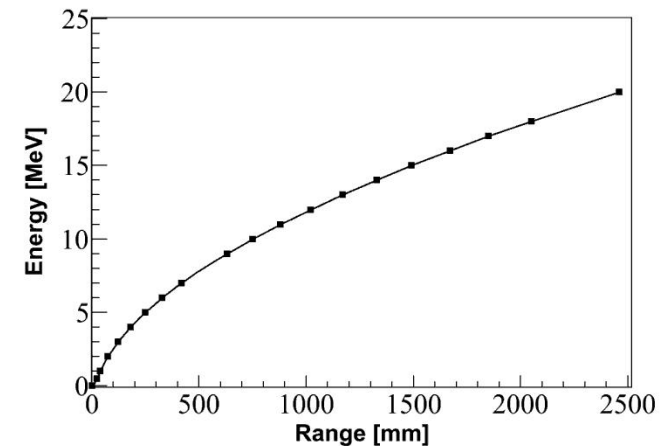
Basics of kinematics reconstruction inside MAYA



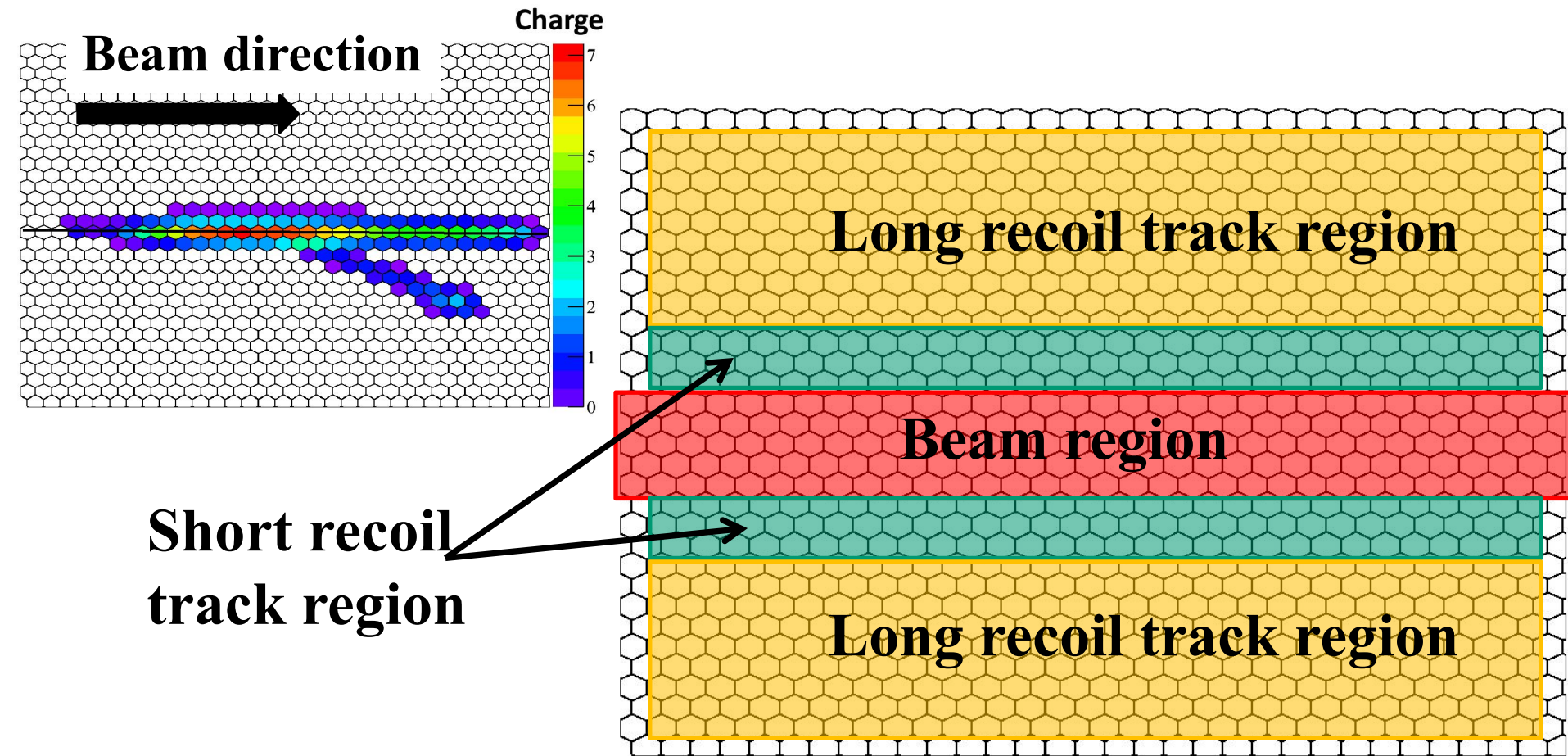
Timing information from Amplification wires

Range \rightarrow Energy
(SRIM)

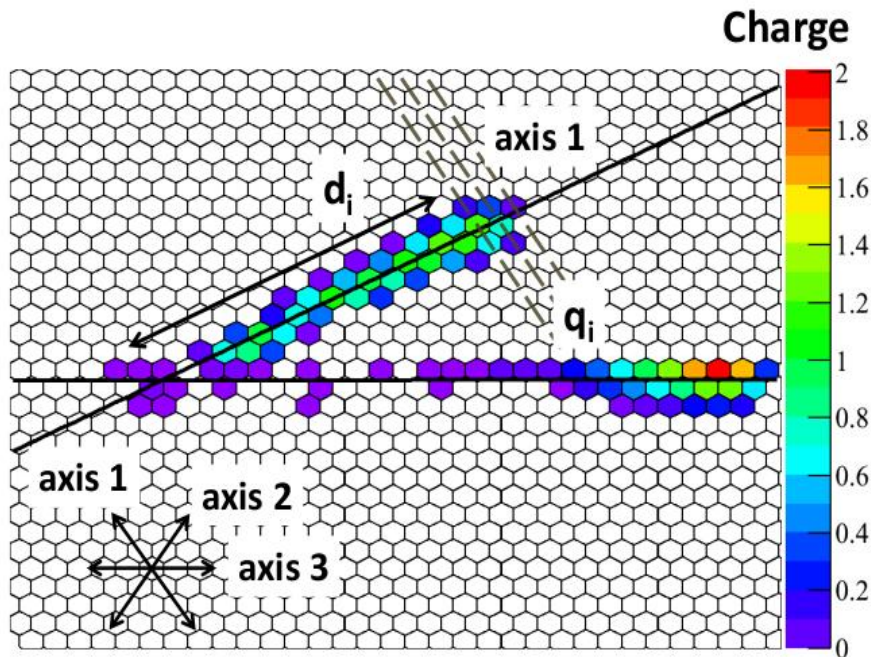
$$R_{2d} \rightarrow R_{3d}, \theta_{2d} \rightarrow \theta_{3d}$$



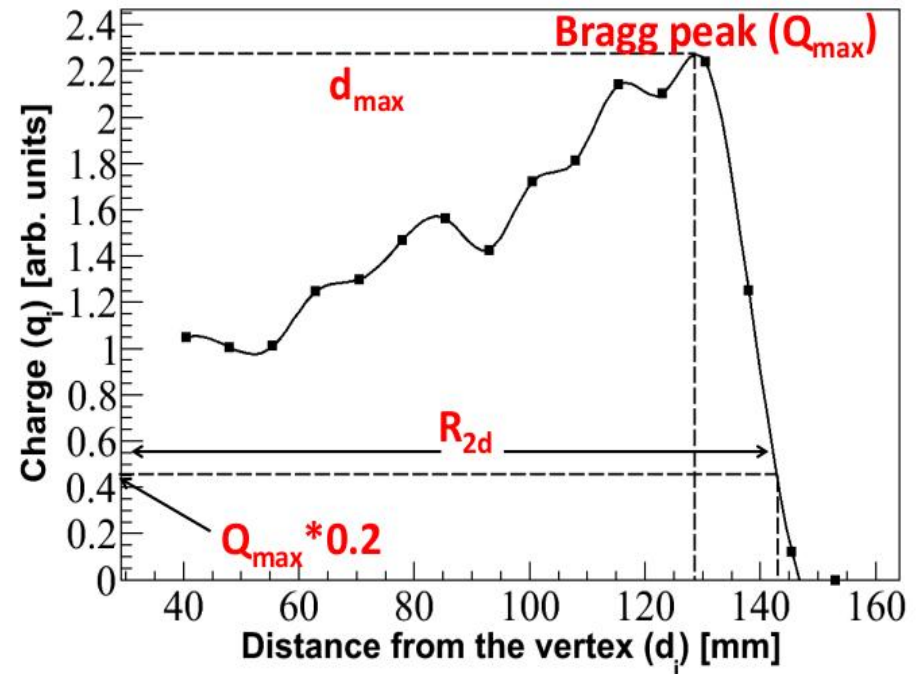
Track reconstruction: Different regions on the cathode pad



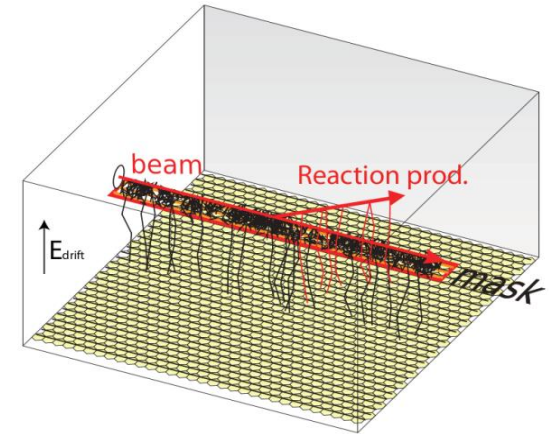
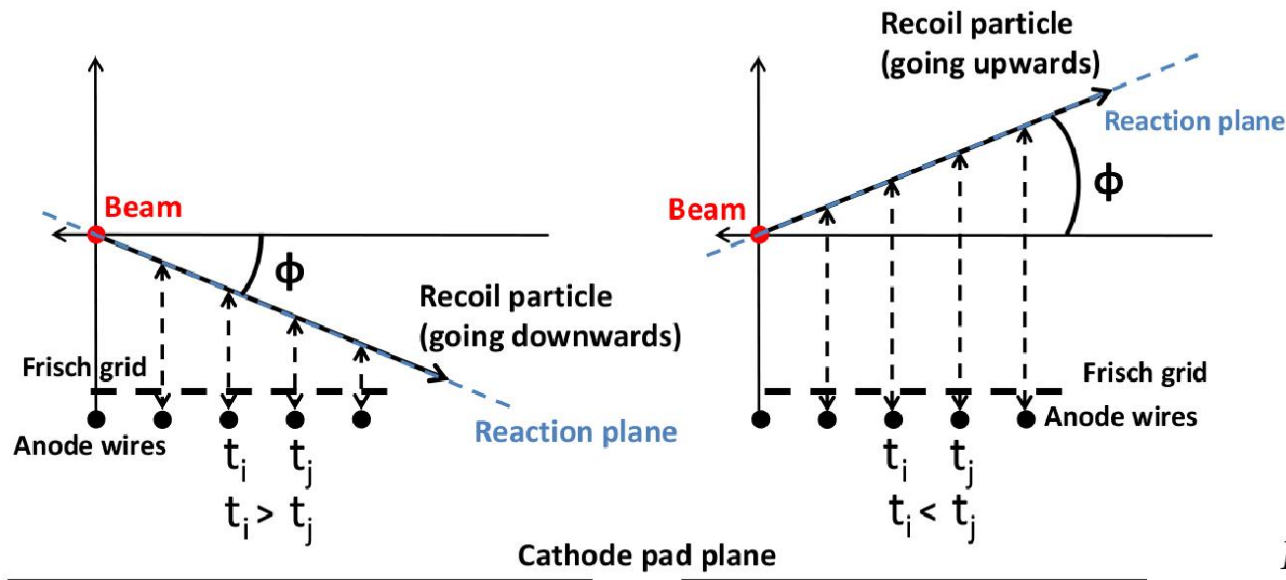
Charge projection along the symmetry axis of a hexagon



Beam direction



3rd dimension from timing information of the anode wires



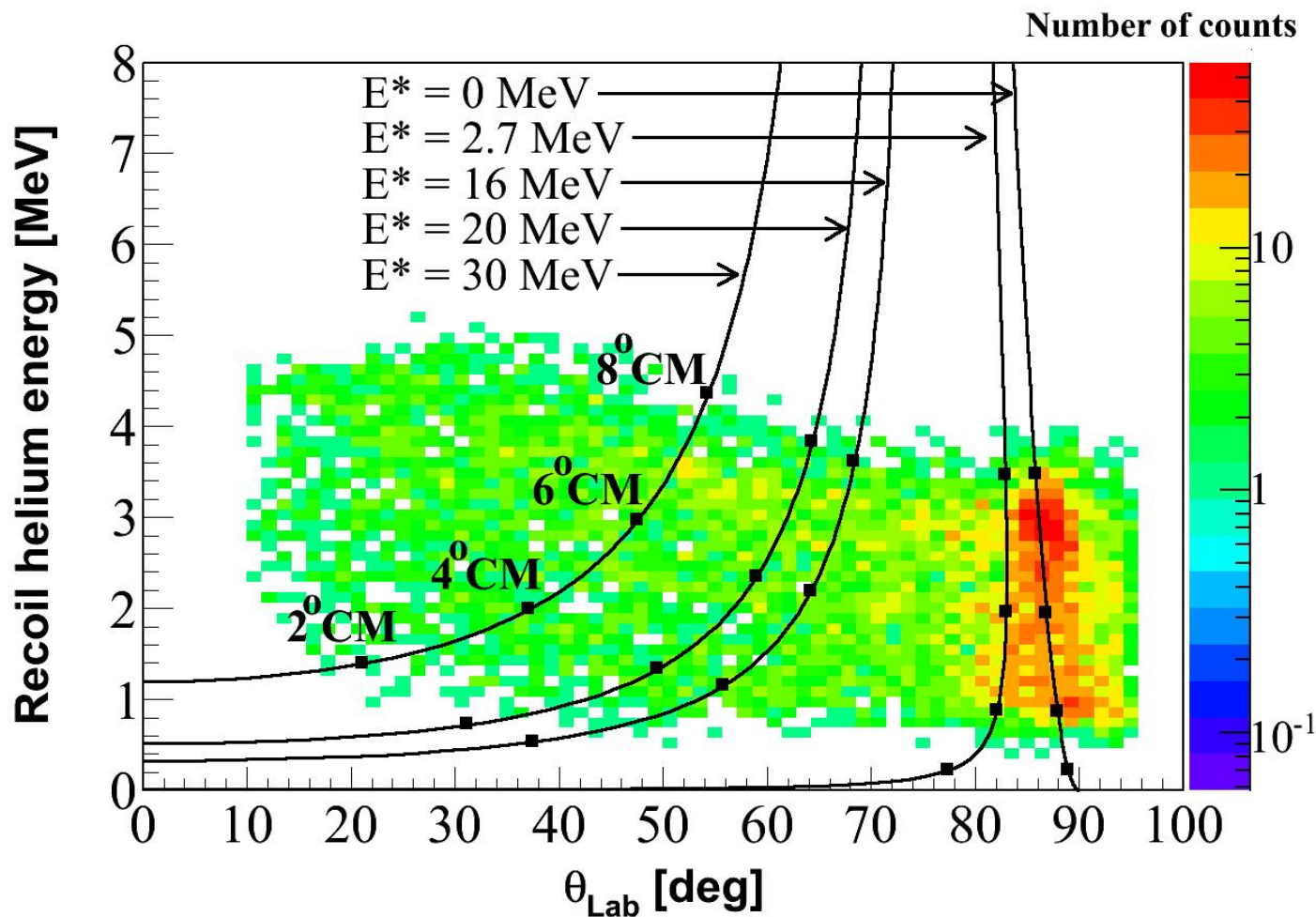
$$R = R_{2d} \sqrt{1 + \sin^2 \theta_{2d} \tan^2 \phi}$$

Range \rightarrow Energy

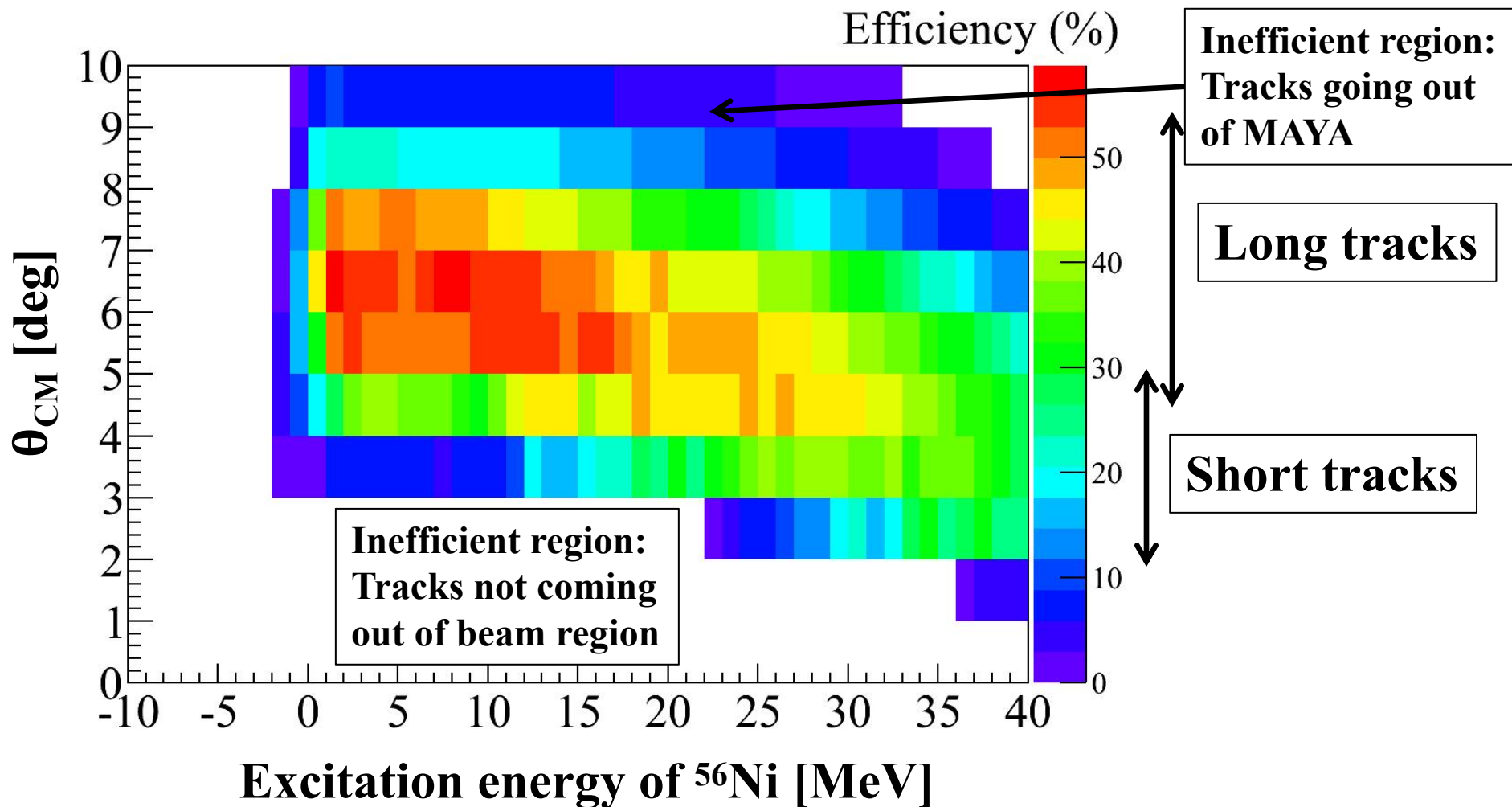
$$\cos \theta = \frac{\cos \theta_{2d}}{\sqrt{1 + \sin^2 \theta_{2d} \tan^2 \phi}}$$

Data

Kinematics plot

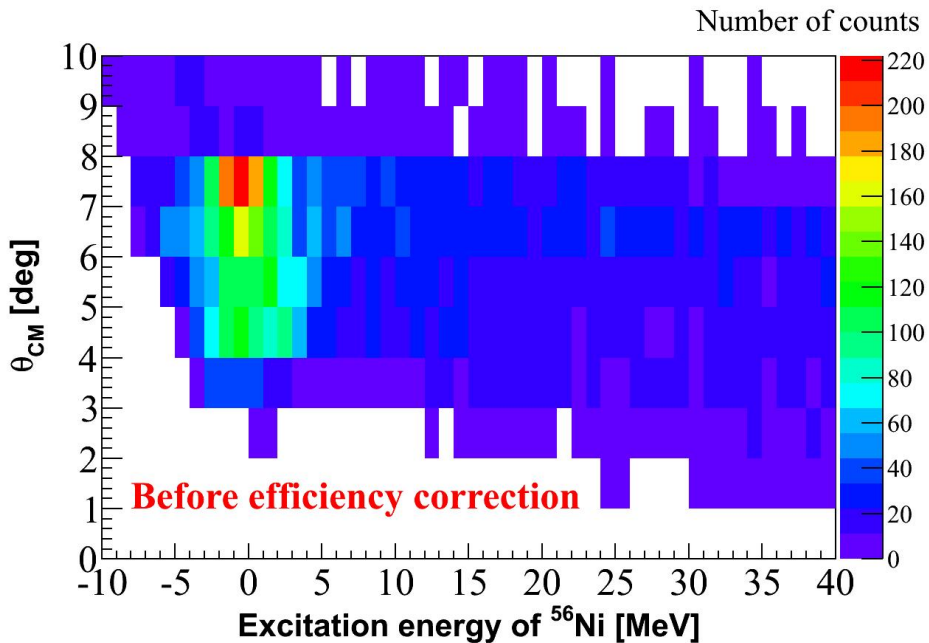


Efficiency (reconstruction and geometric) plot from simulation

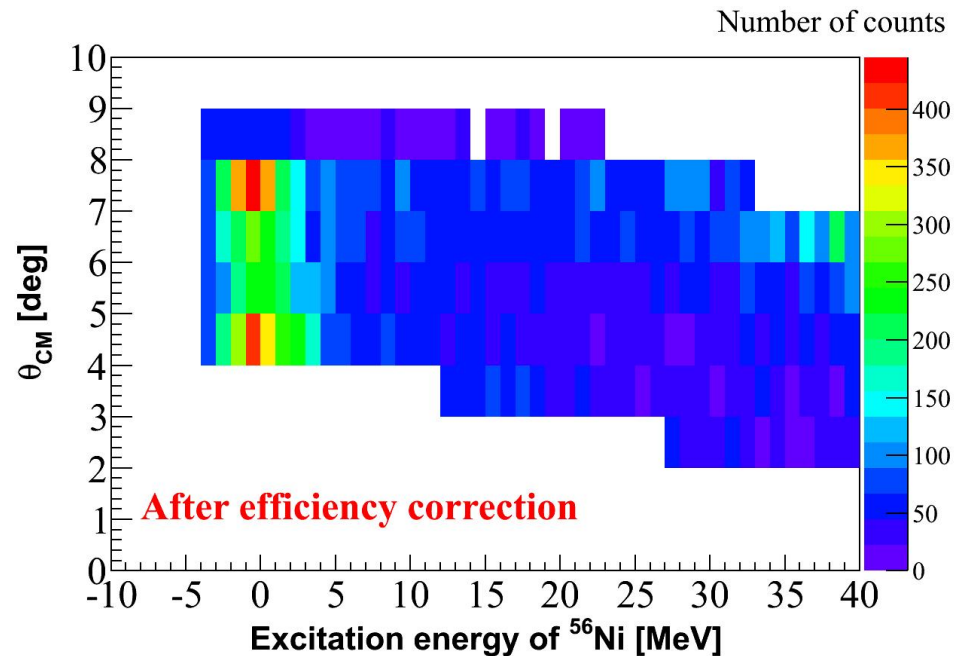


Efficiency correction

Data (Not efficiency corrected)

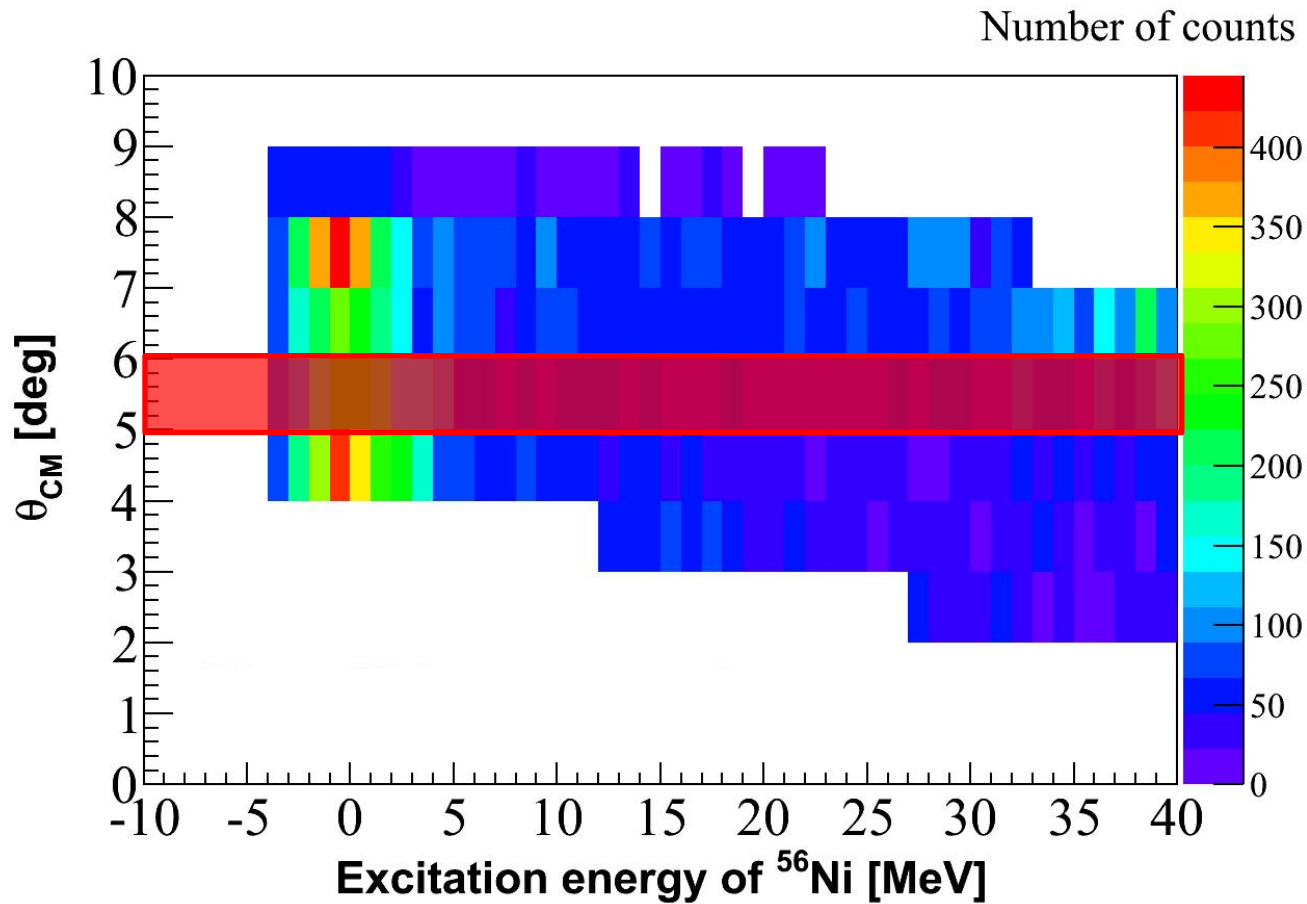


Data (Efficiency corrected)

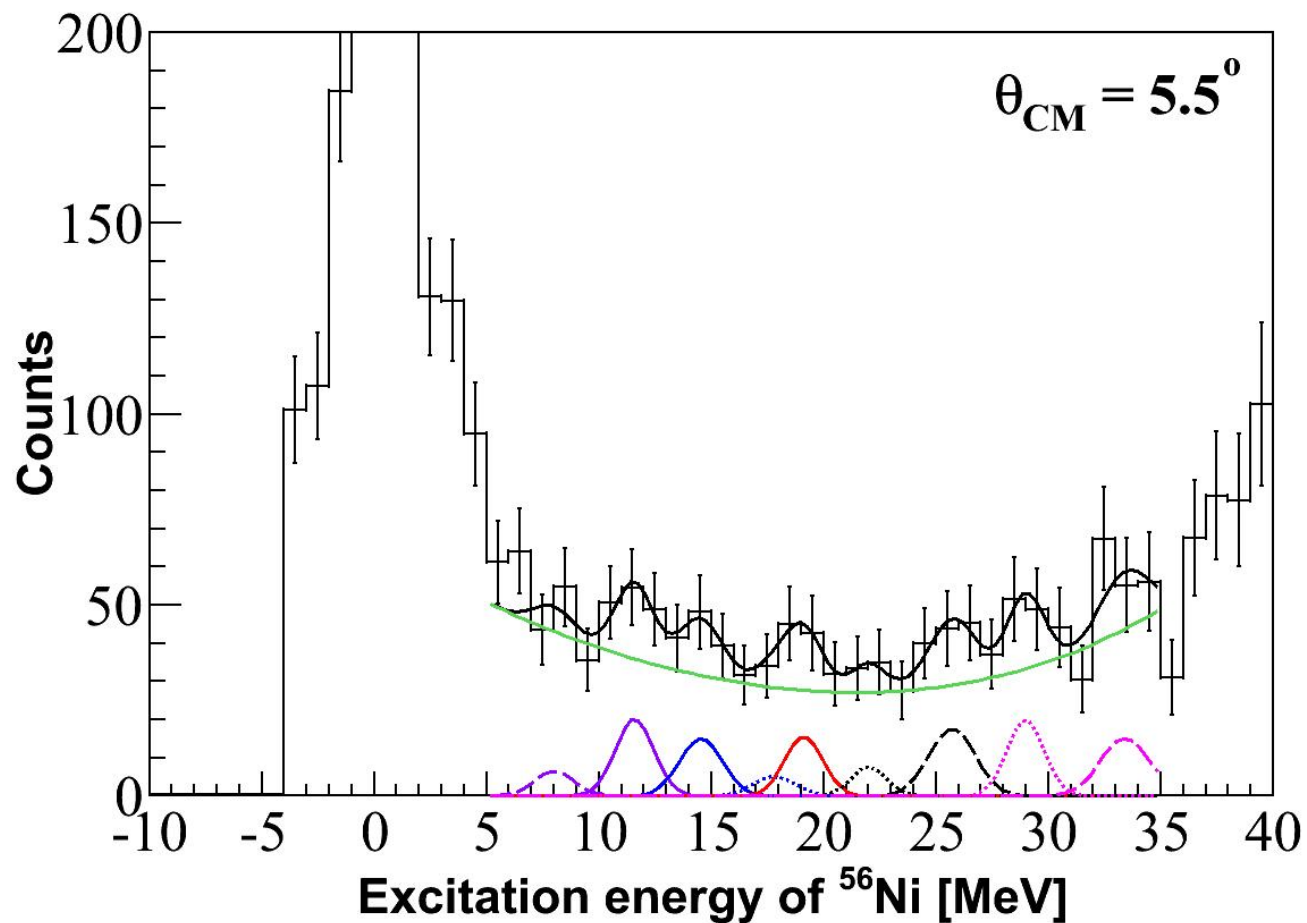


Data (Efficiency corrected)

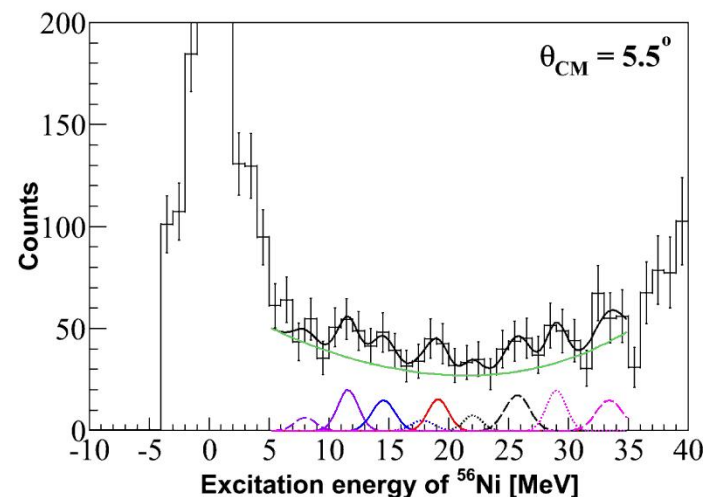
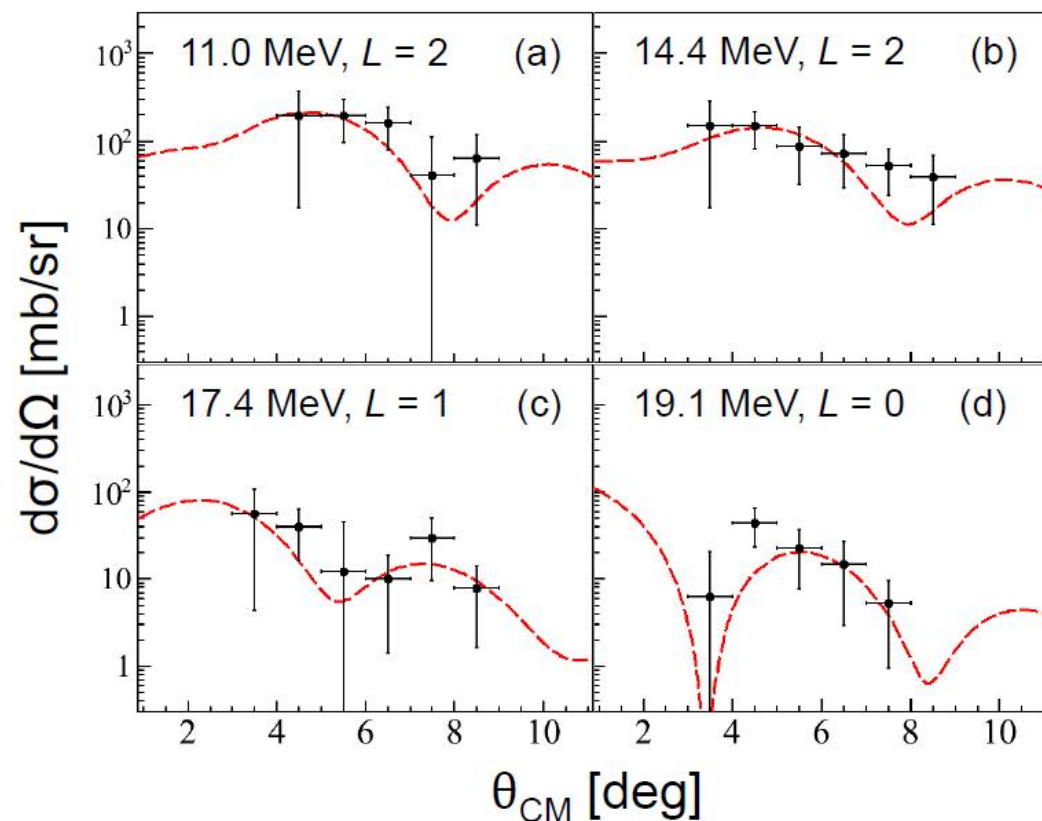
Peak fitting method



Peak fitting method

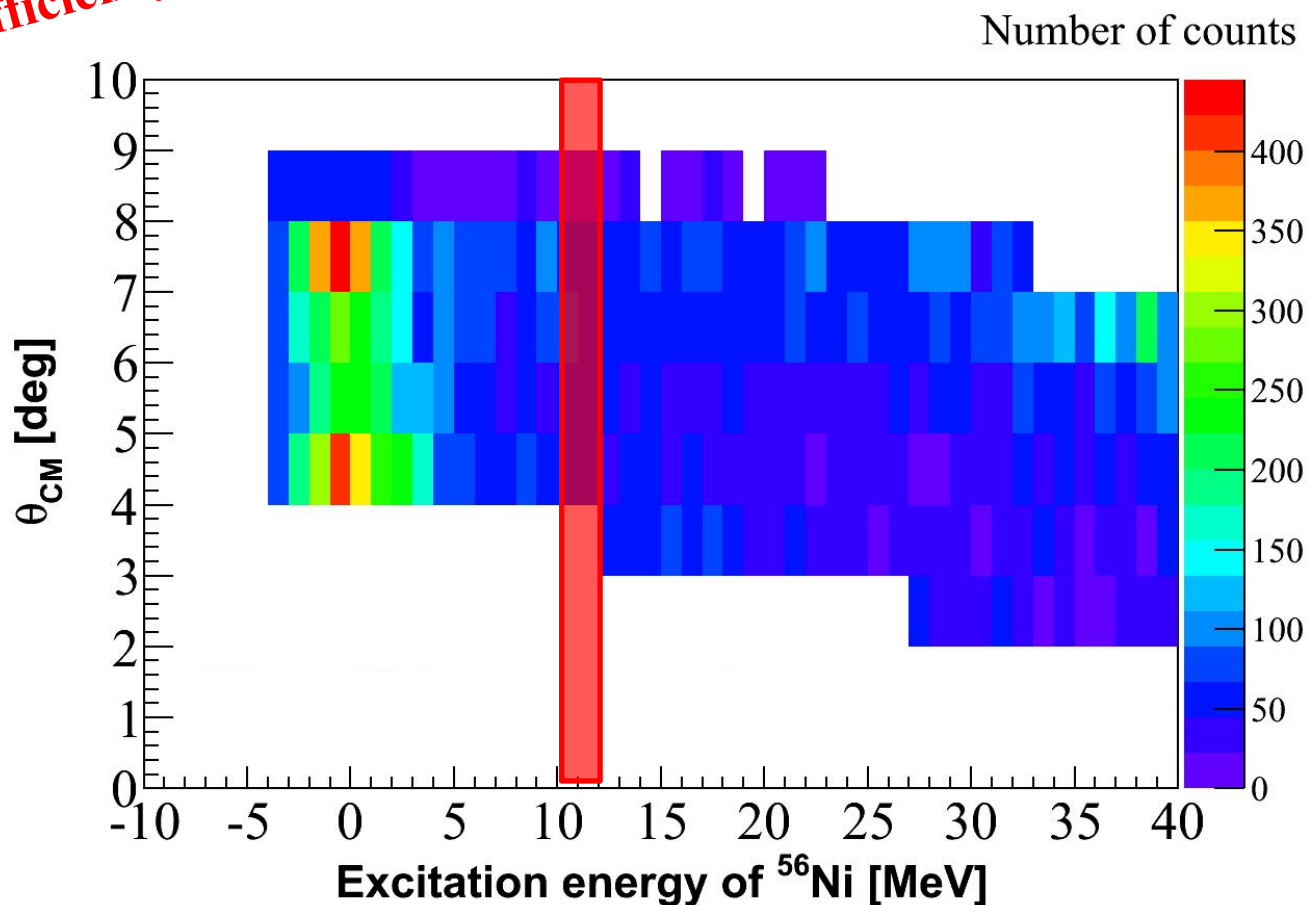


Angular distribution of the Gaussian peaks



Multipole Decomposition Analysis (MDA)

Data (Efficiency corrected)

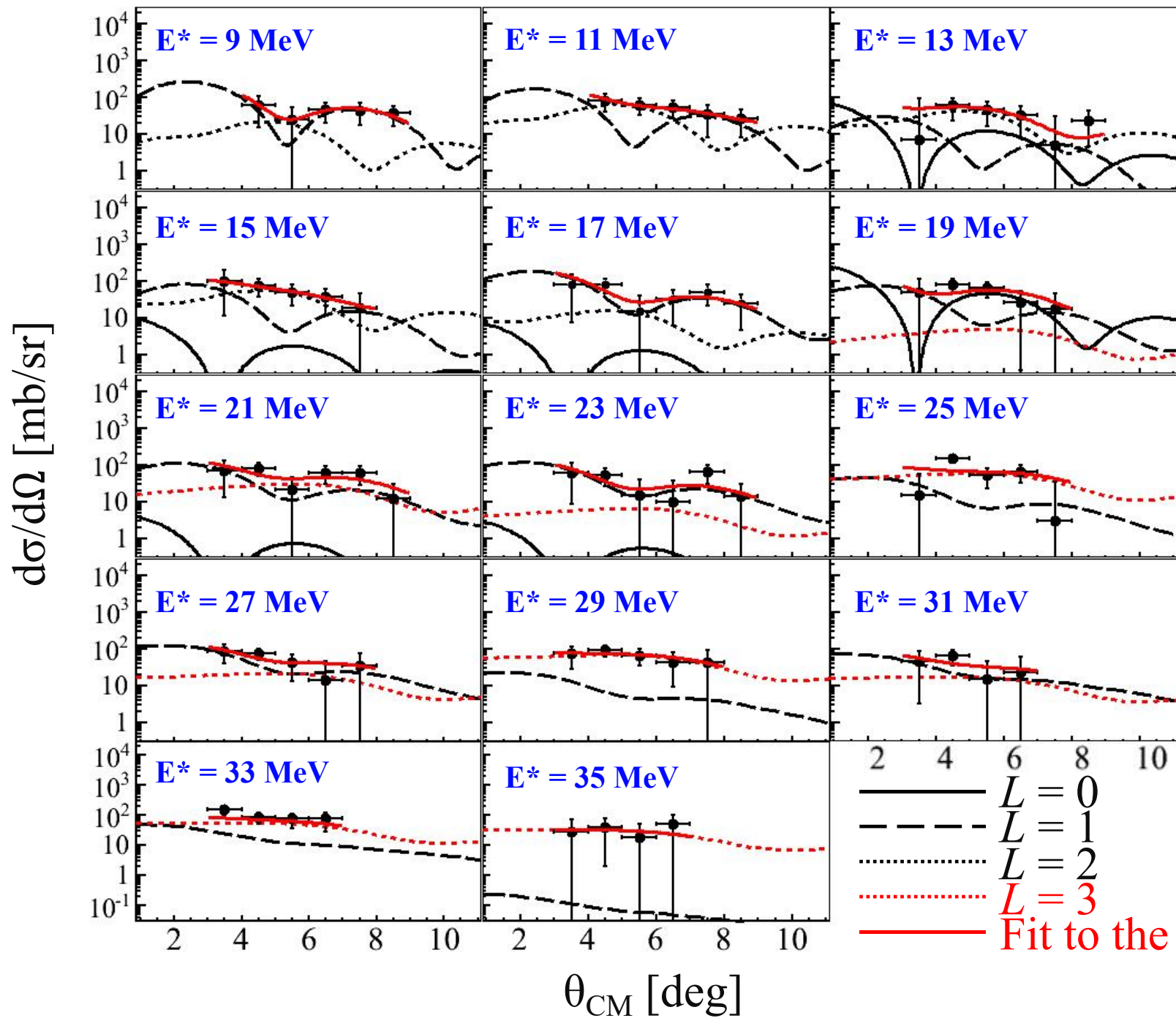


$$\left. \frac{d^2\sigma}{d\Omega dE}(\theta_{CM}, E^*) \right|_{exp} = \sum_{L=0}^{L=3} a_L(E^*) \left. \frac{d^2\sigma_L}{d\Omega dE}(\theta_{CM}, E^*) \right|_{theory}$$

Coefficient represents the sum-rule strength for L and E^*

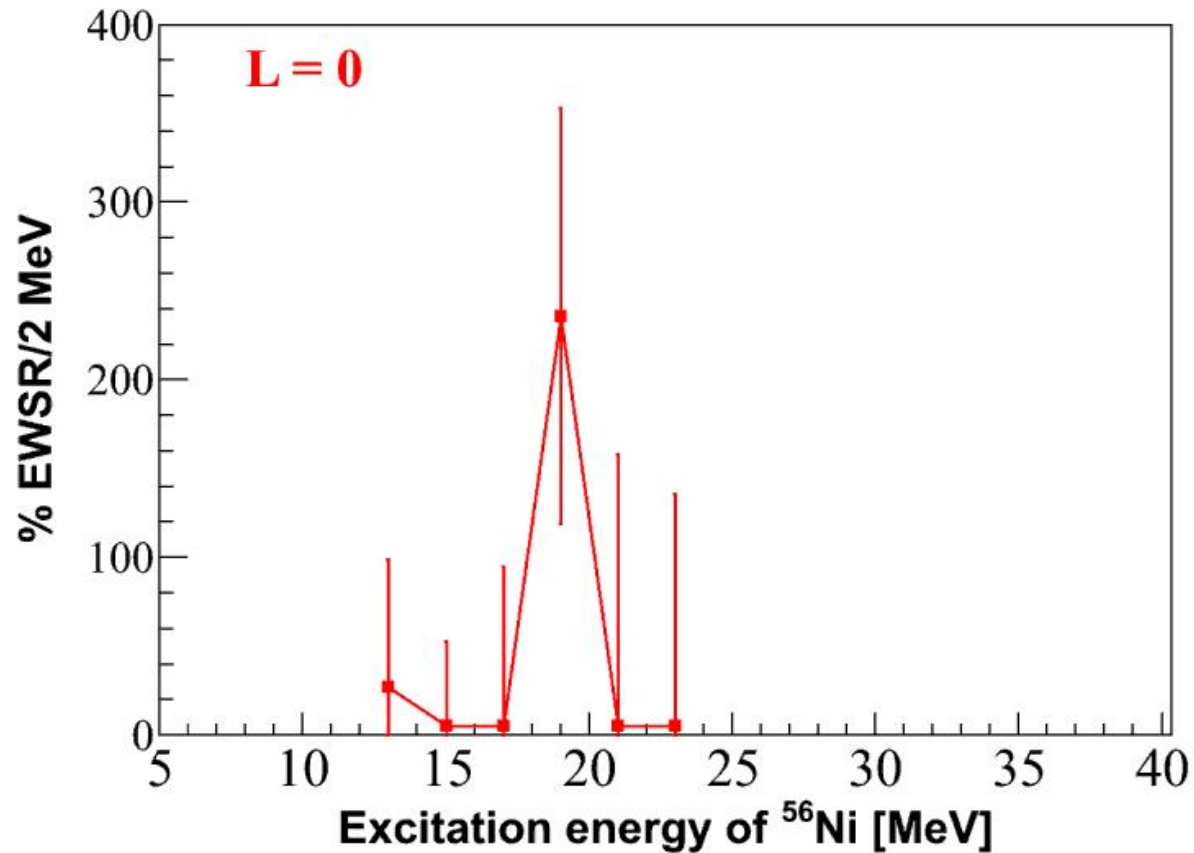
Experimental angular distribution for a given energy interval

Angular distribution obtained from DWBA calculations for a given energy interval and multipolarity



$L = 0, T = 0$ (ISGMR)

EWSR obtained from MDA analysis



$L = 0, T = 0$ (ISGMR)

Reaction	Gaussian fitting		MDA	
	E^* [MeV]	FWHM [MeV]	E^* [MeV]	Width (rms) [MeV]
$^{56}\text{Ni}(\alpha, \alpha')^{56}\text{Ni}^*$ (this work)	19.1 ± 0.5	2.0 ± 0.3	18.4 ± 1.8	2.0 ± 1.2
$^{56}\text{Ni}(d, d')^{56}\text{Ni}^*$	19.5 ± 0.3	5.2	19.3 ± 0.5	2.3
$^{58}\text{Ni}(\alpha, \alpha')^{58}\text{Ni}^*$	18.43 ± 0.15	7.41 ± 0.13	$19.2^{+0.44}_{-0.19}$	$4.89^{+1.05}_{-0.31}$
$^{58}\text{Ni}(\alpha, \alpha')^{58}\text{Ni}^*$	-	-	$19.9^{+0.7}_{-0.8}$	-
$^{60}\text{Ni}(\alpha, \alpha')^{60}\text{Ni}^*$	17.62 ± 0.15	7.55 ± 0.13	$18.04^{+0.35}_{-0.23}$	$4.5^{+0.97}_{-0.22}$
$^{68}\text{Ni}(\alpha, \alpha')^{68}\text{Ni}^*$	21.1 ± 1.9	1.3 ± 1.0	23.4	6.5

S. Bagchi *et al.*, Phys. Lett. B751 (2015) 371; C. Monrozeau *et al.*, Phys. Rev. Lett. 100 (2008) 042501;
M. Vandebrouck *et al.*, Phys. Rev. Lett. 113 (2014) 032504; Y.-W. Lui *et al.*, Phys. Rev. C 73 (2006)
014314.

Summary & conclusion:

- Inelastic α -particle scattering is suitable for isoscalar excitations.
- First measurement of the isoscalar responses of ^{56}Ni via inelastic α -particle scattering.
- Active target techniques can be used to study the properties of unstable nuclei.
- $L = 0$ mode:
Centroid: 19.1 ± 0.5 MeV (Peak-fitting method)
 18.4 ± 1.8 MeV (MDA)
- Low-lying component of the ISGDR.

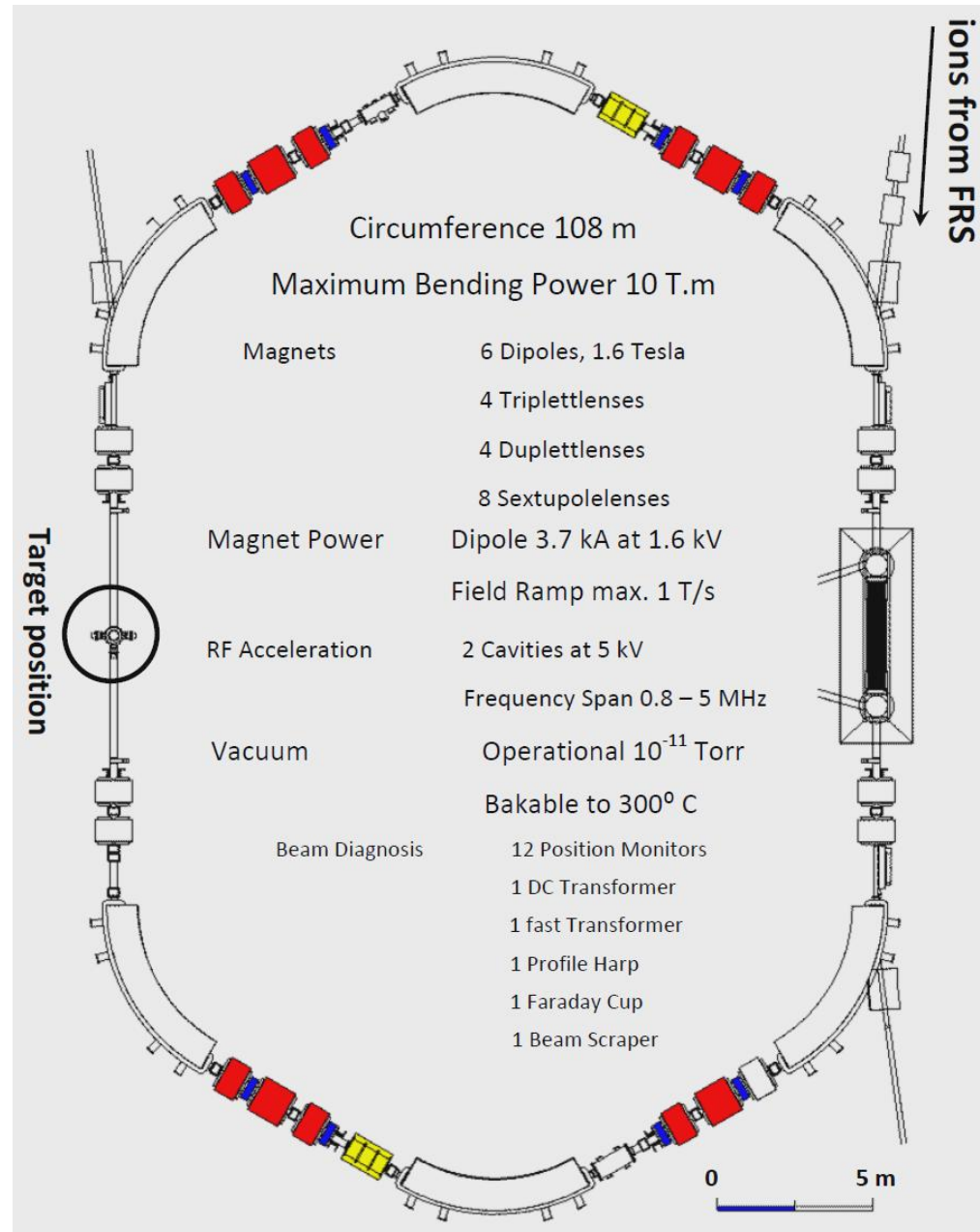
GSI Storage Ring

Experimental Storage Ring

Luminosity:

$$10^{26} - 10^{27} \text{ cm}^{-2}\text{s}^{-1}$$

EPJ Web Conf. 66 (2014) 03093



Advantages and disadvantages of storage-ring experiments

Advantages:

Large intensities in the ring

Little energy loss in the target

No target window (no background)

High resolution of the beam (electron cooling)

Forward focusing for high-energy particles

Low-energy threshold

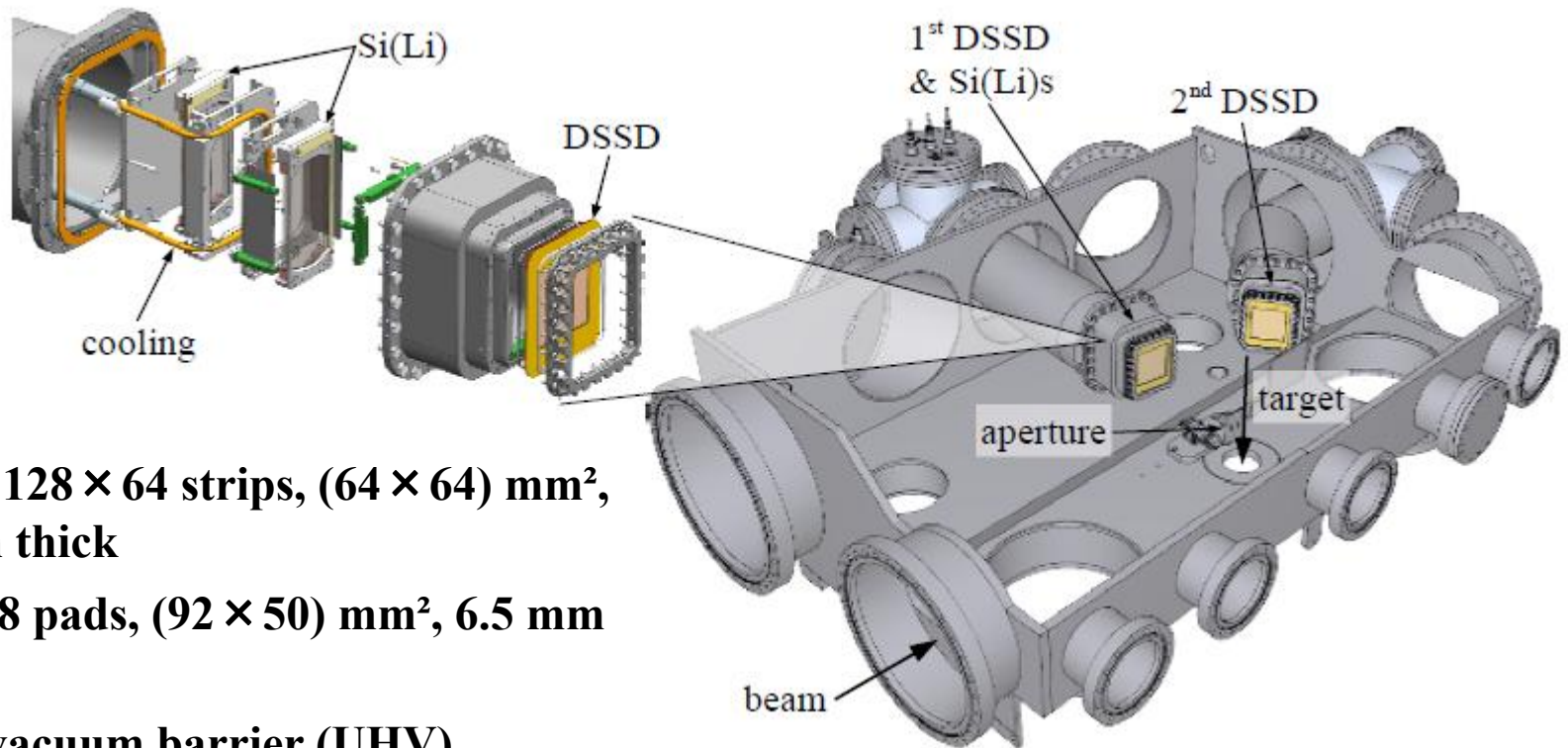
Disadvantages:

Ultra high vacuum

Very small recoil energies for low q

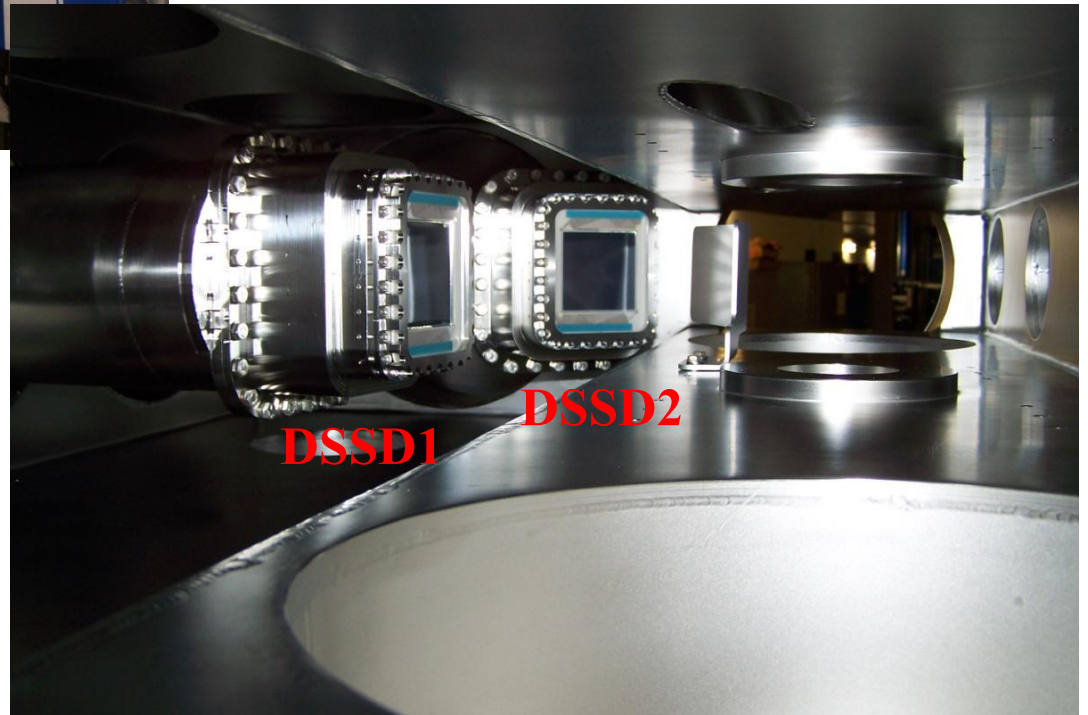
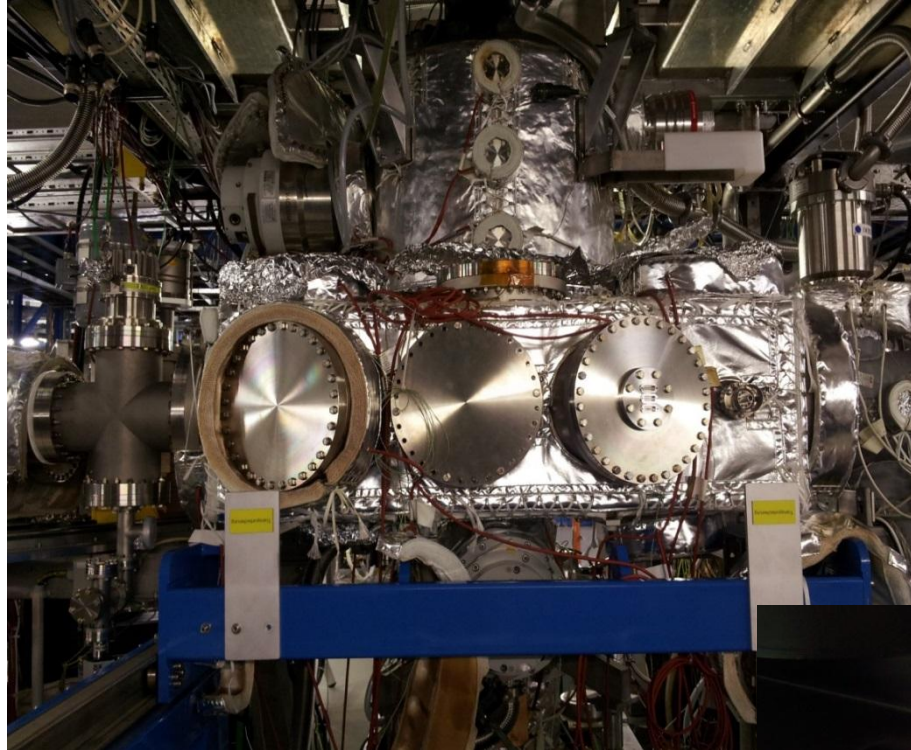
Thin targets

Prototype of EXL (EXotic nuclei studied in Light-ion induced reactions at the (N)ESR storage ring) detector at ESR



- DSSD: 128×64 strips, (64×64) mm², 285 μ m thick
- Si(Li): 8 pads, (92×50) mm², 6.5 mm thick
- active vacuum barrier (UHV)
- moveable aperture to improve angular resolution

$73^\circ < \theta_{\text{lab}} < 88^\circ$ for DSSD1, and
 $27^\circ < \theta_{\text{lab}} < 38^\circ$ for DSSD2





Date 02.Dec.'93 Time 23:01:30

Ref.Lev
-75.00 dBm

Res.Bw 300.0 Hz [3dB]

Vid.Bw 10 kHz

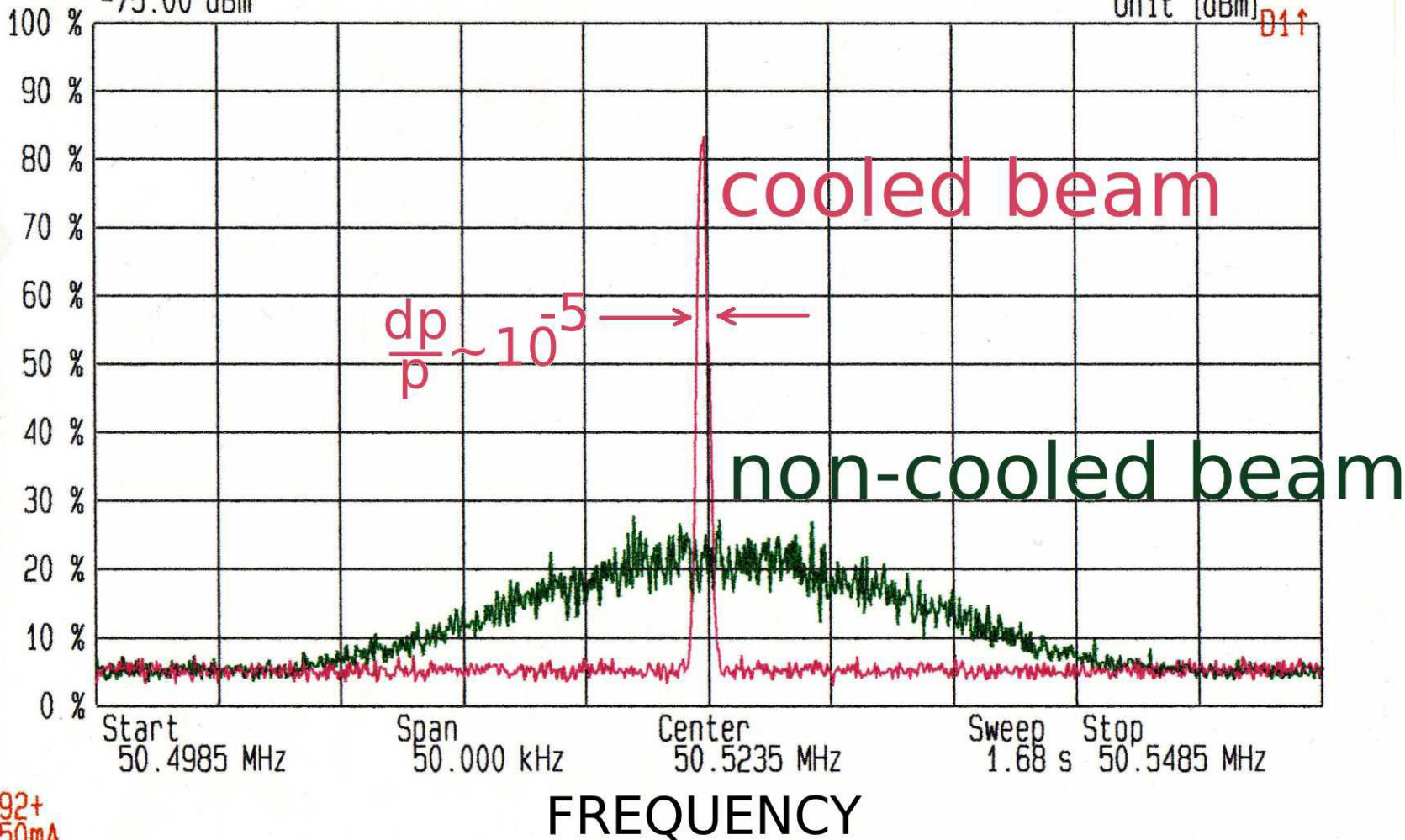
CF.Stp 5.000 kHz

RF.Att 10 dB

Unit [dBm]

D1↑

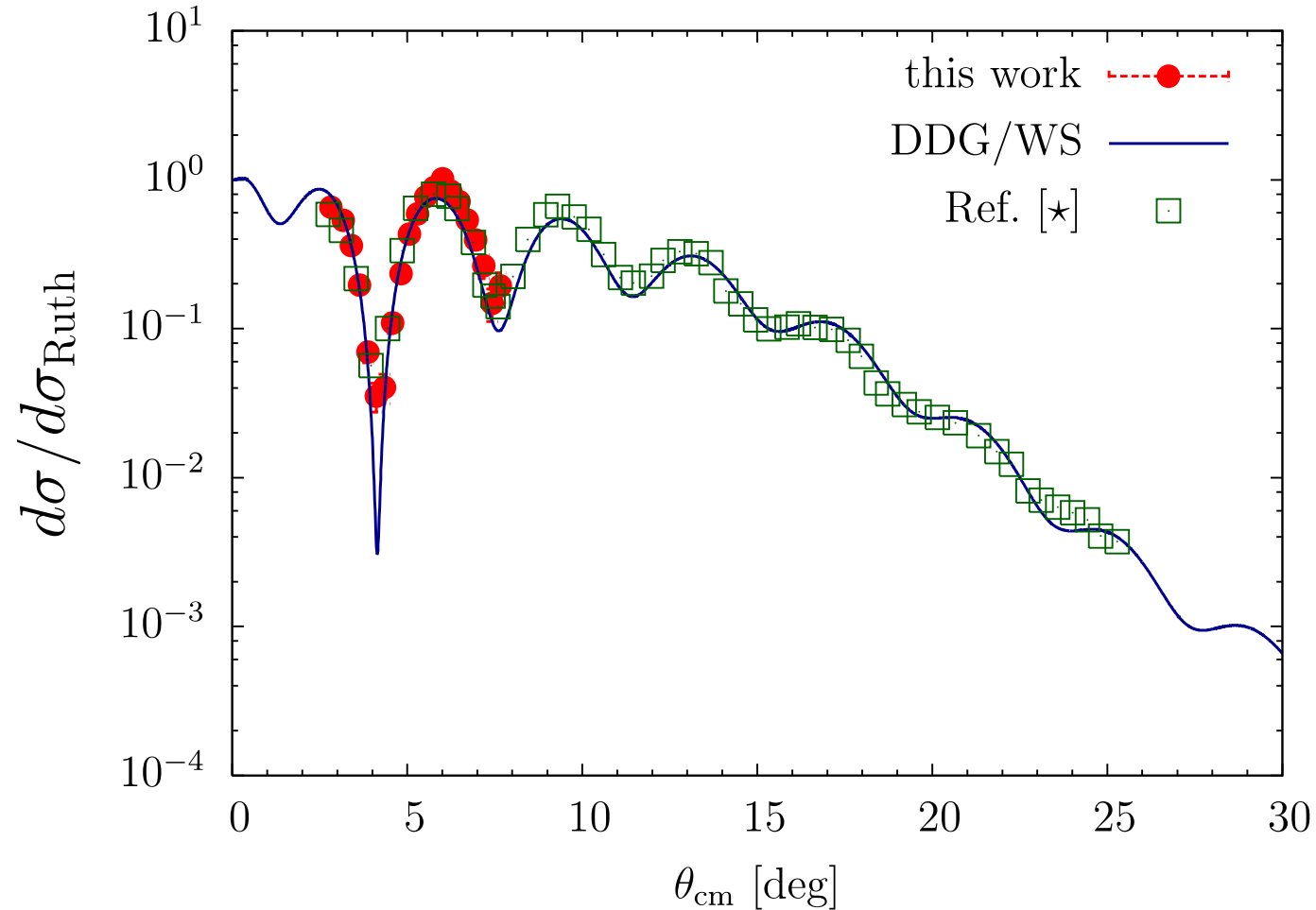
INTENSITY



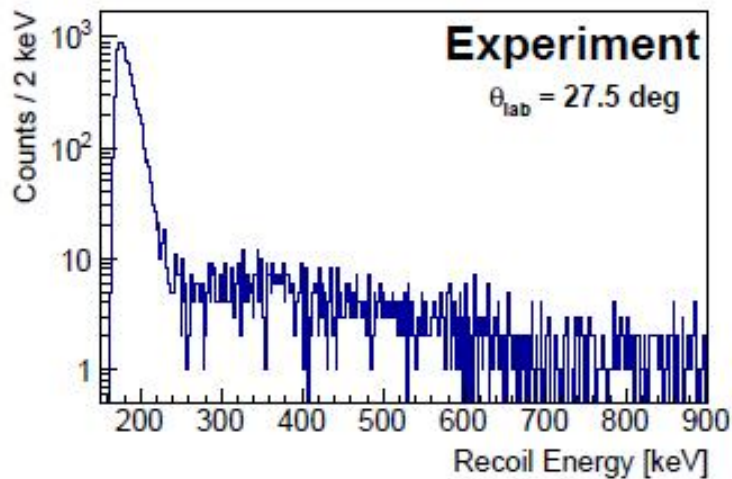
J92+
250mA

Elastic alpha scattering at 100 MeV/nucleon

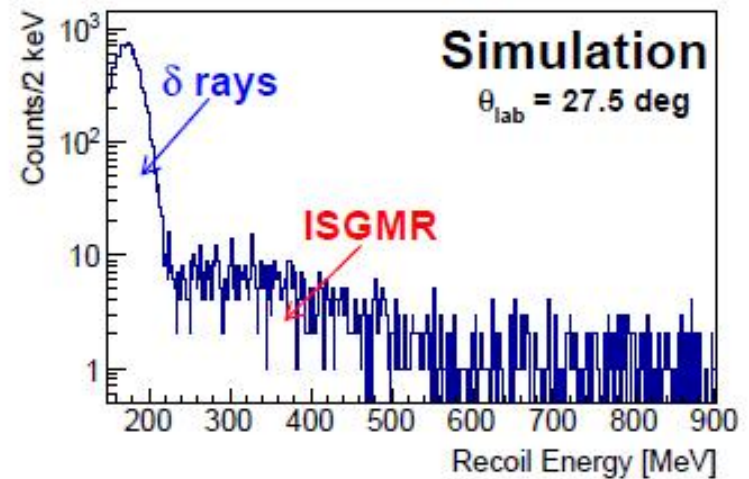
Ph.D. work of J.C. Zamora



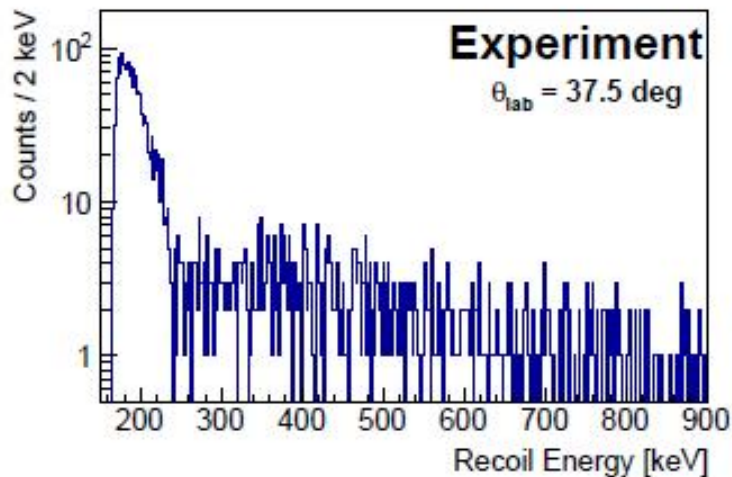
* B.K. Nayak *et al.*, Phys. Lett. B 637 (2006) 43



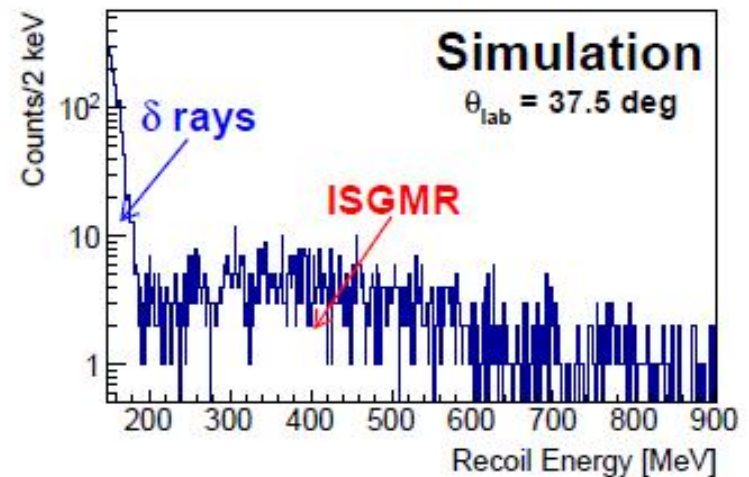
(a) Experiment, strip number 0



(b) Simulation, strip number 0.

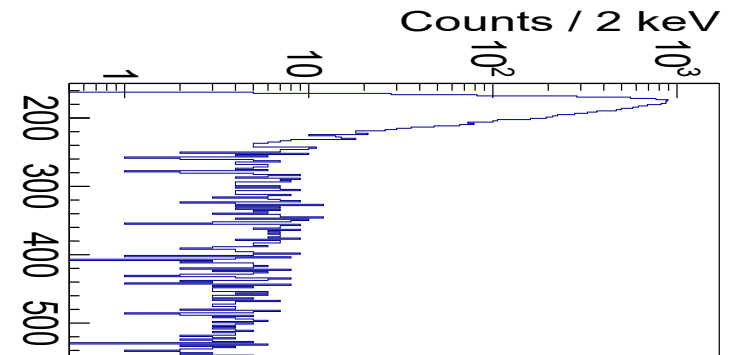
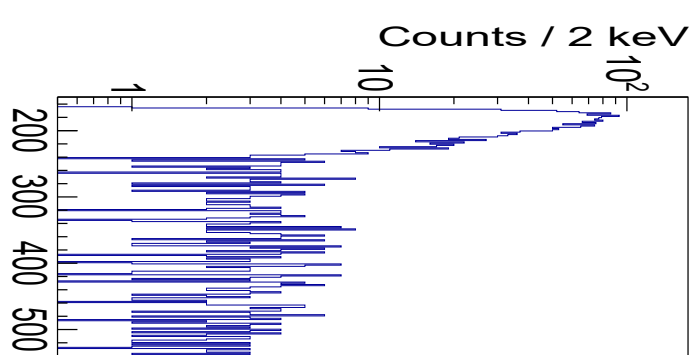


(c) Experiment, strip number 31



(d) Simulation, strip number 31

Inelastic alpha scattering (100 MeV/nucleon, PhD J.C. Zamora)



	Centroid [MeV]	Width _{rms} [MeV]
this work	19(1)	5(1)
PRC 61, 067307 (2000)	$20.3^{+1.7}_{-0.1}$	$4.3^{+0.7}_{-0.2}$
PRC 73, 014314 (2006)	$19.2^{+0.4}_{-0.2}$	$4.9^{+1.1}_{-0.3}$
Phys. Lett. B 637, 43 (2006)	$19.9^{+0.7}_{-0.8}$	—

this work

19(1)

5(1)

PRC 61, 067307 (2000)

$20.3^{+1.7}_{-0.1}$

$4.3^{+0.7}_{-0.2}$

PRC 73, 014314 (2006)

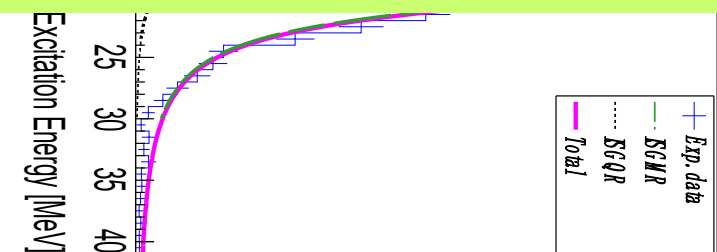
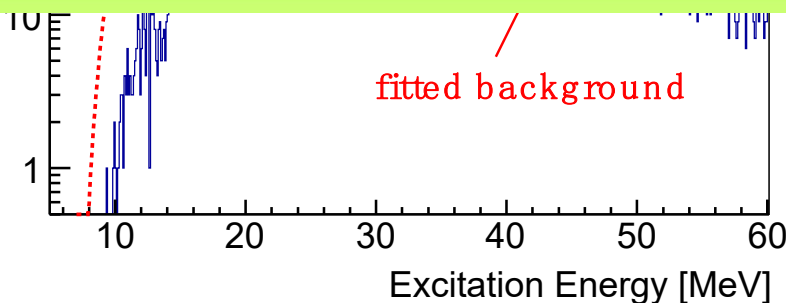
$19.2^{+0.4}_{-0.2}$

$4.9^{+1.1}_{-0.3}$

Phys. Lett. B 637, 43 (2006)

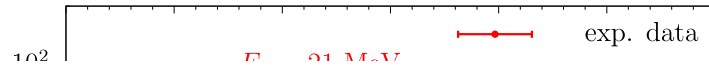
$19.9^{+0.7}_{-0.8}$

—

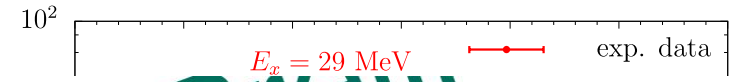


Inelastic alpha scattering (100 MeV/nucleon) from ^{58}Ni

J.C. Zamora, PhD work [J.C. Zamora *et al.*, Phys. Lett. B763 (2016) 16-19]



(centroid)



m_1/m_0 [MeV] EWSR [%]

21^{+5}_{-3}	86^{+17}_{-9}
$20.3^{+1.7}_{-0.1}$	74^{+22}_{-12}
$19.9^{+0.7}_{-0.8}$	92^{+4}_{-3}
20.8	108

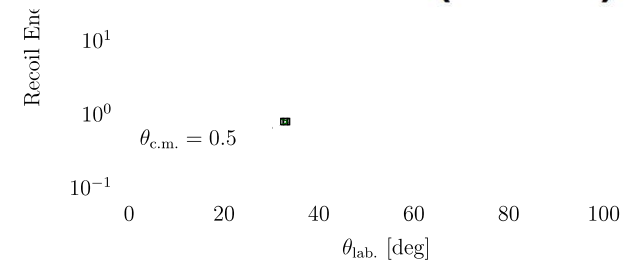
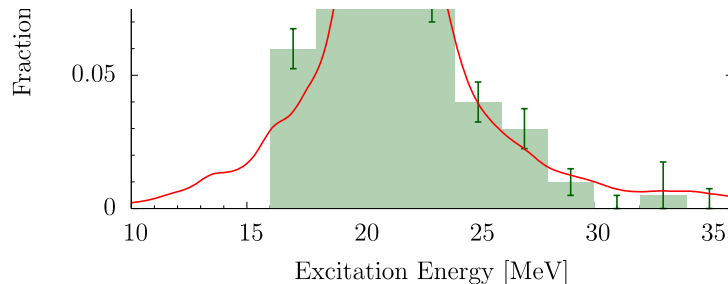
this work

PRC **61**, 067307 (2000)

Phys. Lett. B **637**, 43 (2006)

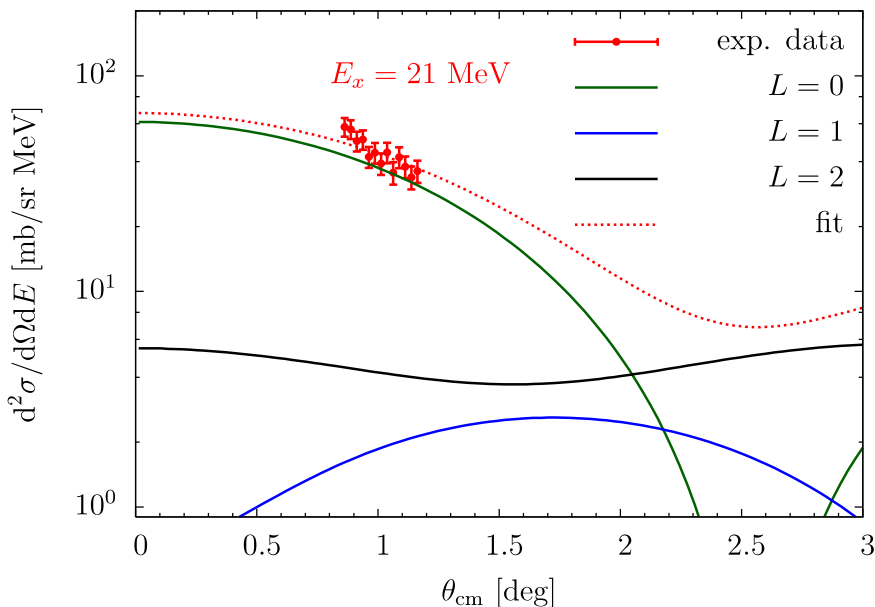
RPA calculation *

* G. Colò *et al*, Comput. Phys. Commun. 184 (2013)

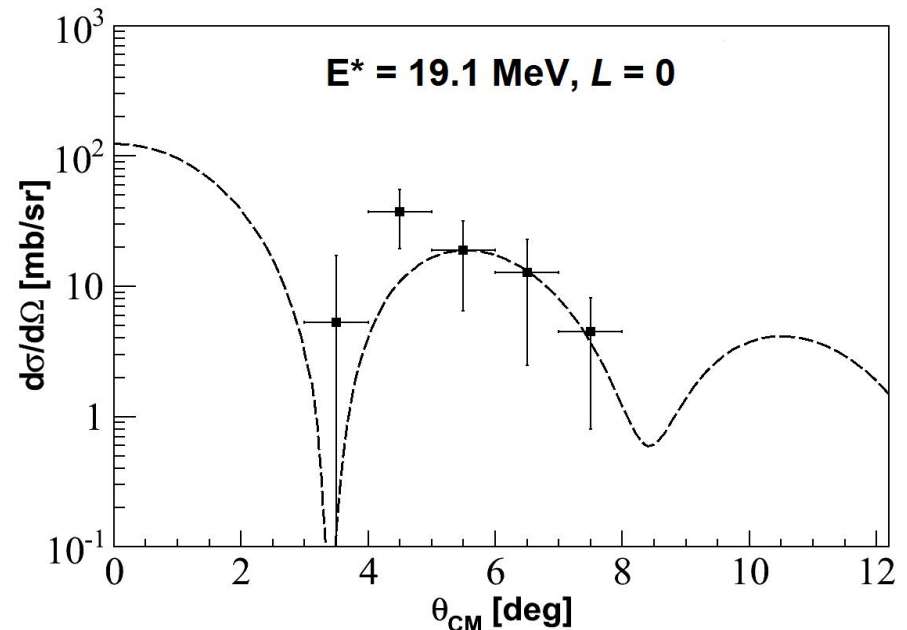


Monopole mode in ^{58}Ni and ^{56}Ni

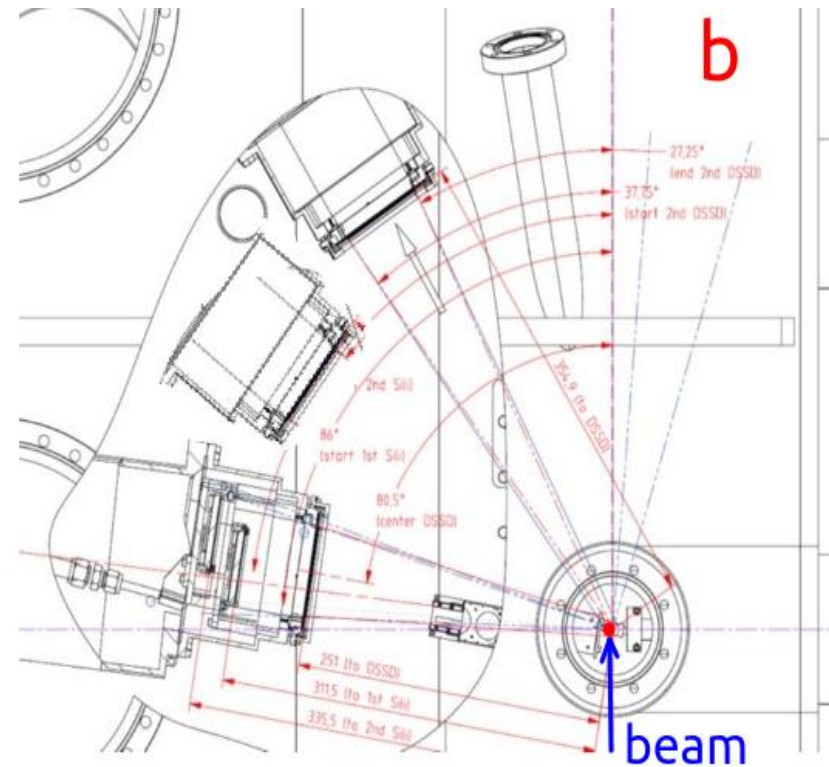
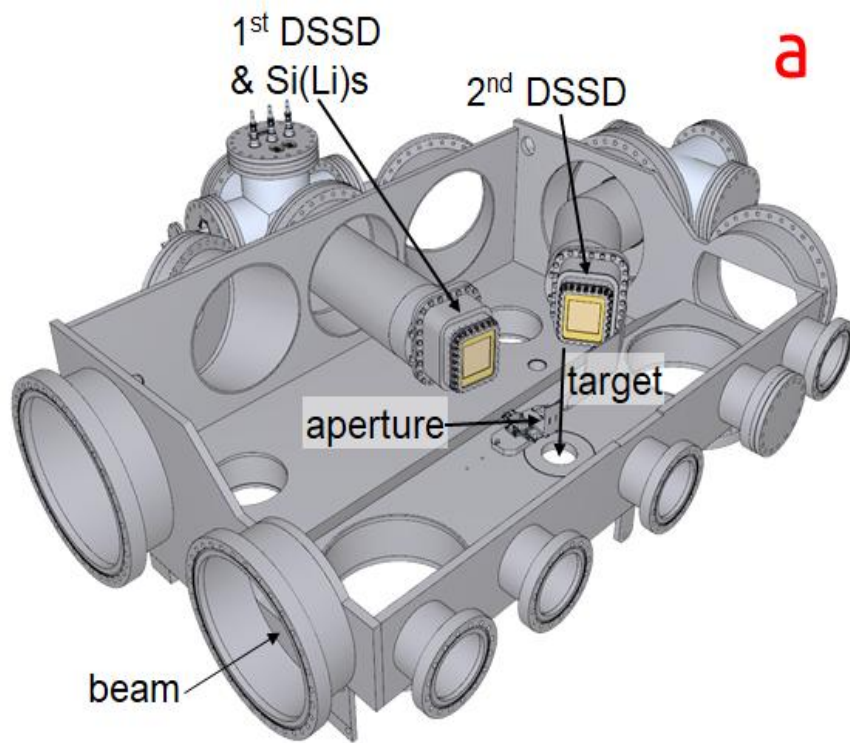
Ring vs. Active target



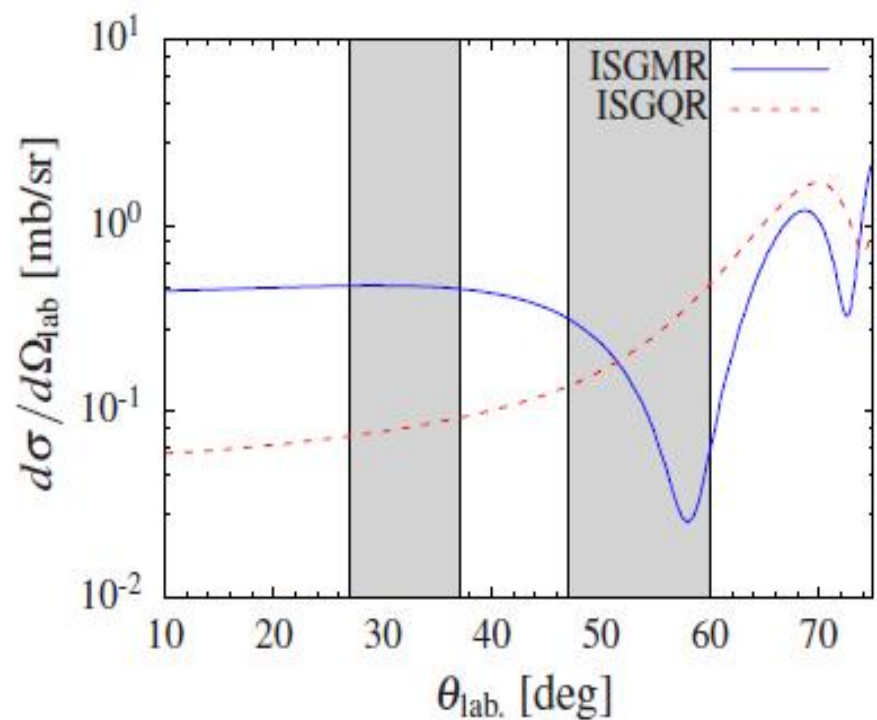
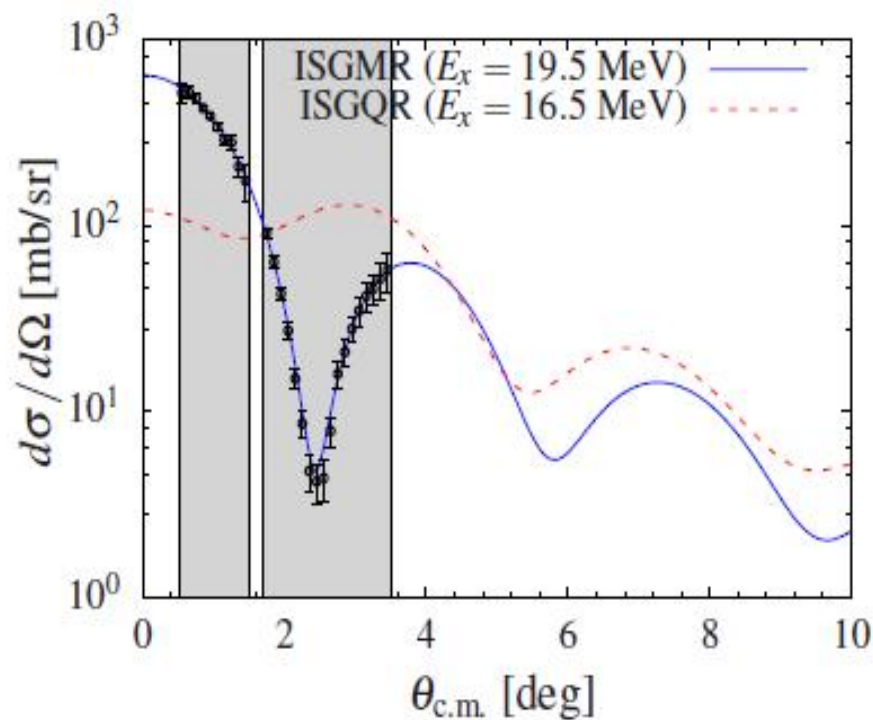
^{58}Ni



^{56}Ni



- a) Drawing of the UHV-compatible scattering chamber. The detectors were centred at 80° and 32° with respect to the beam direction.
- b) **Sketch of the upgraded experimental setup for the proposed experiment. A new detector will be installed in between the two existing pockets covering angles from 47° to 60° .**



DWBA for the ISGMR and the ISGQR in ^{56}Ni at 100 MeV/u for inelastic scattering off ^4He assuming exhaustion of 100% of the respective EWSR.

Angular distributions in centre-of-mass frame (left) and in laboratory frame (right). The shadowed regions mark the angular range covered by the detector set-up in the proposed experiment. The data points (Monte-Carlo simulation) show the expected statistical uncertainty for the ISGMR with the requested beam time (14 days).

E605: ISGDR in ^{56}Ni

Linda Achouri	Hidetoshi Akimune	Soumya Bagchi
Konstanze Boretzky	Haifa Bouzomita	Lucia Caceres
Franck Delaunay	Manuel Caamaño	Beatriz Fernandez
Mamoru Fujiwara	Umesh Garg	Julien Gibelin
Geoff Grinyer	Muhsin N. Harakeh	Nasser Kalantar-Nayestanaki
Omar Kamalou	Elias Khan	Attila Krasznahorkay
Gregoire Lhoutellier	Sergei Lukyanov	Katarzyna Mazurek
Mohammad Ali Najafi	Julien Pancin	Yuri Penionzkhevich
Luc Perrot	Riccardo Raabe	Catherine E. Rigollet
Thomas Roger	Sara Sambi	Hervé Savajols
Christelle Stodel	Laura Suen	Jean Charles Thomas
Marine Vandebrouck	Jarno van de Walle	

The EXL-E105 Collaboration



*S. Bagchi¹, S. Bönig², M. Castlós³, I. Dillmann⁴, C. Dimopoulou⁴, P. Egelhof⁴, V. Eremin⁵, H. Geissel⁴, R. Gernhäuser⁶, M.N. Harakeh¹, A.-L. Hartig², S. Ilieva², N. Kalantar-Nayestanaki¹, O. Kiselev⁴, H. Kollmus⁴, C. Kozhuharov⁴, A. Krasznahorkay³, T. Kröll², M. Kuilman¹, S. Litvinov⁴, Yu.A. Litvinov⁴, M. Mahjour-Shafiei¹, M. Mutterer⁴, D. Nagae⁸, M.A. Najafi¹, C. Nociforo⁴, F. Nolden⁴, U. Popp⁴, C. Rigollet¹, S. Roy¹, C. Scheidenberger⁴, **M. von Schmid²**, M. Steck⁴, B. Streicher^{2,4}, L. Stuhl³, M. Takechi⁴, M. Thürauf², T. Uesaka⁹, H. Weick⁴, J.S. Winfield⁴, D. Winters⁴, P.J. Woods¹⁰, T. Yamaguchi¹¹, K. Yue^{4,7}, **J.C. Zamora²**, J. Zenihiro⁹*

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² Technische Universität Darmstadt

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⁴ GSI, Darmstadt

⁵ Ioffe Physico-Technical Institute, St. Petersburg

⁶ Technische Universität München

⁷ Institute of Modern Physics, Lanzhou

⁸ University of Tsukuba

⁹ RIKEN Nishina Centre

¹⁰ The University of Edinburgh

¹¹ Saitama University

Thank you for your attention

For the equation of state of symmetric nuclear matter at saturation nuclear density:

$$\left[\frac{d(E/A)}{d\rho} \right]_{\rho=\rho_0} = 0$$

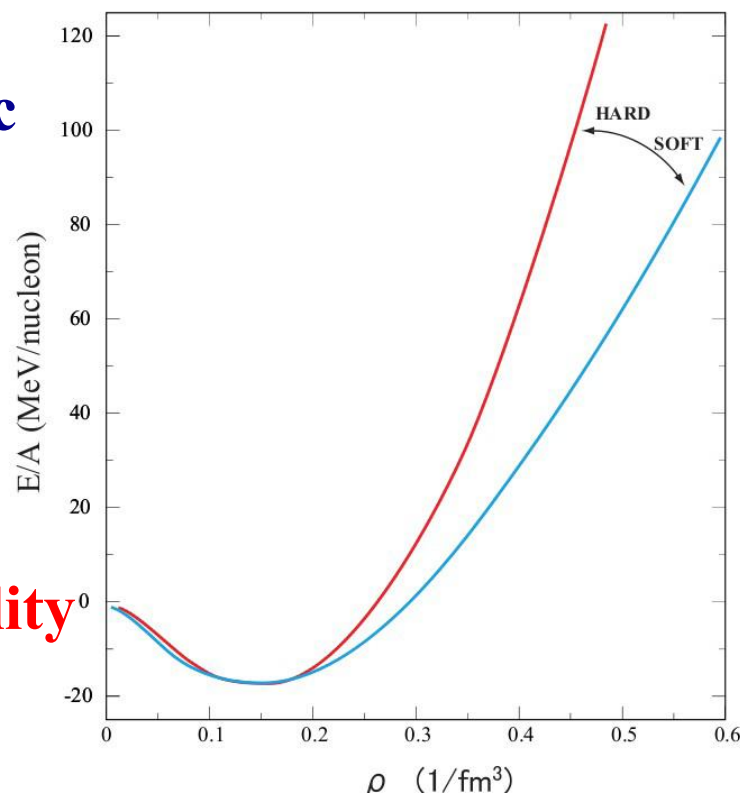
and one can derive the incompressibility of nuclear matter:

$$K_{nm} = \left[9\rho^2 \frac{d^2(E/A)}{d\rho^2} \right]_{\rho=\rho_0}$$

E/A : binding energy per nucleon

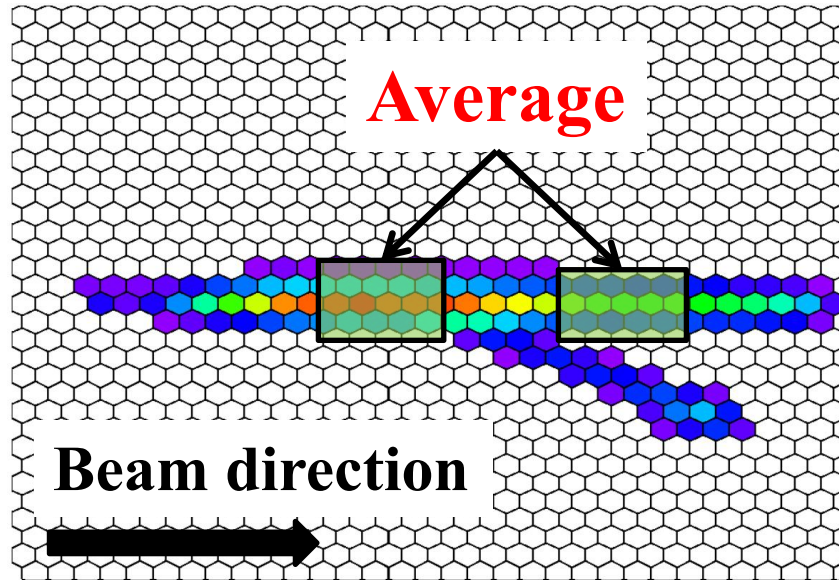
ρ : nuclear density

ρ_0 : nuclear density at saturation

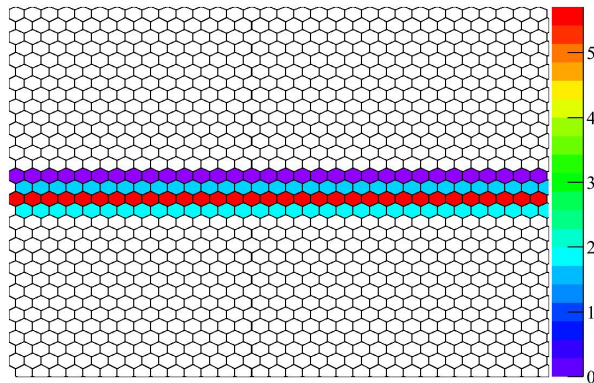


J.P. Blaizot, Phys. Rep. 64 (1980) 171

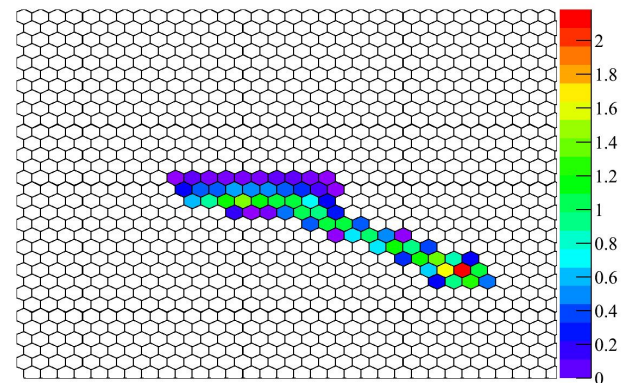
Beam subtraction



Extrapolated beam

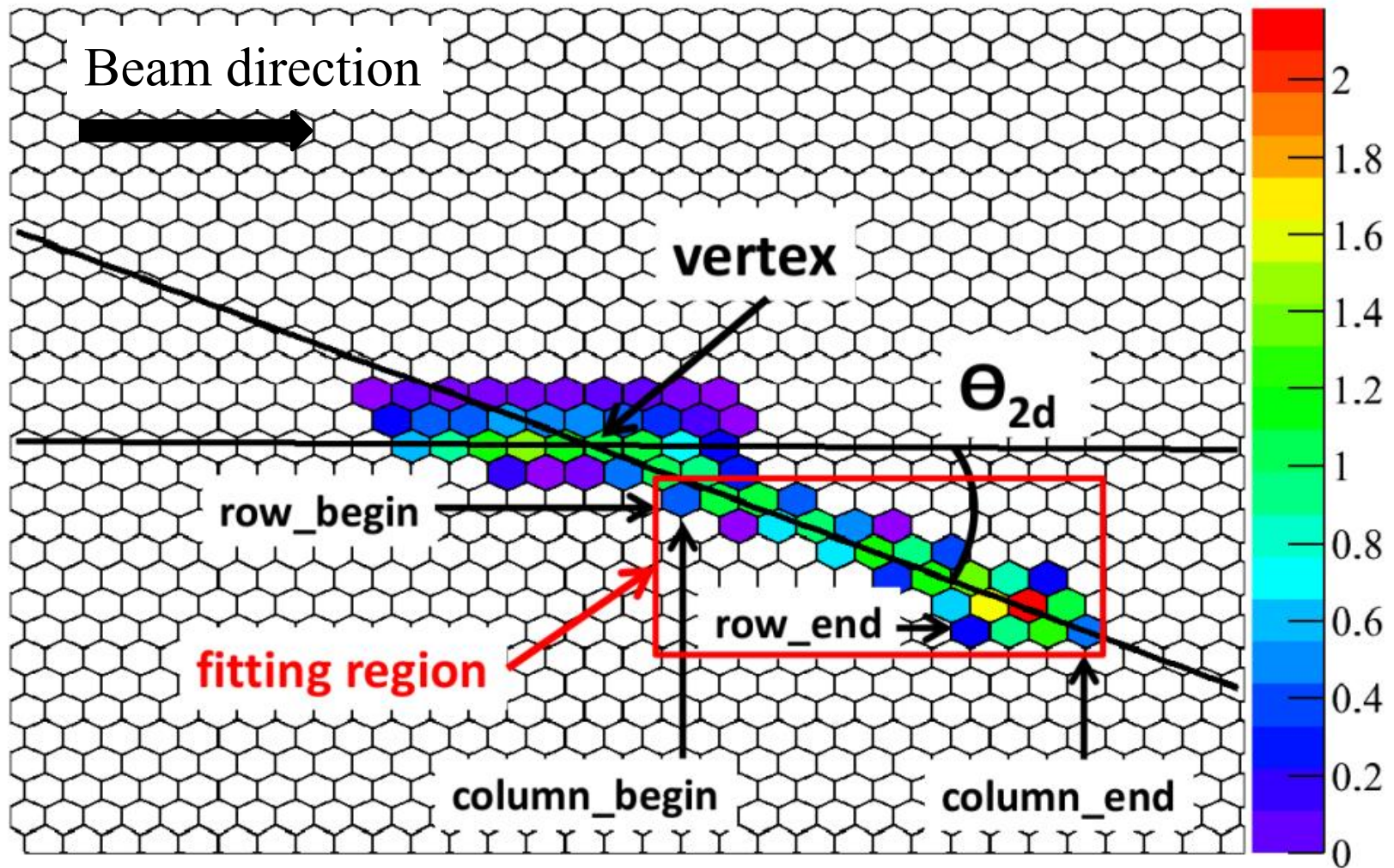


Beam subtracted



Long recoil track fitting

Charge



Short recoil track fitting

