

Learning from the Large Electron-Positron Collider LEP



 L = 26.659 km
 \times 3

 Max SR Power 18 MW
 \times 2.8

 Max RF Voltage 3.7 MV
 \times 2.6

 Lumi / x-ing 2.5e11 cm⁻²s⁻¹
 \times 3

 Ecms 92 - 209 GeV
 \times 1.8

 Intens 4×10¹¹ e+, e- / bunch
 \times 0.4



LEP : tunneling 13/9/1983 - 8/2/1988; installation largely in 1988 + octant test Pilot run, first Z's, low L, superconducting final focus magnets off : August 1989 Operation : 1990 - 2000 after tough discussions stopped and dismantled for LHC then busy with LHC + CLIC, LHeC, ELFE renewed e+e- ring interest more recent



LEP performance workshops



initiated by Steve Myers, critical review to further improve LEP, held during the winter stops



Photo courtesy John Jowett LEP Performance workshop #1, Chamonix, January 13-19, 1991

numerous detailed improvements, new optics every year





Discussed in Chamonix meetings, well documented in proceedings

Had disappeared, restored in 2020 following my request inspired by the Jan'20 IAS MDI workshop

1st Workshop on LEP Performance, Chamonix 1991: 2nd Workshop on LEP Performance, Chamonix 1992: 3rd Workshop on LEP performance, Chamonix 1993: 4th Workshop on LEP Performance, Chamonix 1994: 5th Workshop on LEP Performance, Chamonix 1995: 6th LEP Performance Workshop, Chamonix 1996: 7th LEP Performance Workshop, Chamonix 1997: 8th LEP Performance Workshop, Chamonix 1998: 9th LEP-SPS Performance Workshop, Chamonix 1999: 10th Workshop on LEP-SPS Performance, Chamonix 2000:

https://cds.cern.ch/record/256125 https://cds.cern.ch/record/260389 https://cds.cern.ch/record/248984 https://cds.cern.ch/record/265955 https://cds.cern.ch/record/277821 https://cds.cern.ch/record/289995 https://cds.cern.ch/record/312024 https://cds.cern.ch/record/330057 https://cds.cern.ch/record/359023 https://cds.cern.ch/record/394989

Lesson #1 :

Very dynamic, very complex, changing all the time, orbit, (vertical) emittance, major beam-beam tune shift $(\xi_y = 0.08/IP)$ and (vertical) tails; core/halo see different machine Requiering continous efforts and follow up LEP optics changed a lot : 60/60 ('89-'91), 90/90 ('92), 90/60 ('93/97), 102/90 ('98-'00) Collimation and operational procedures improved, including As a result : LEP2 backgrounds comparable to LEP1



IPs at

IP1'

Collimator

COLH.QL8.R1"

COLH.QS10.L2"

COLH.QS5.L2"

COLV.QS5.L2"

COLZ.QS4.L2"

COLH. 084.L2"

COLZ.QS2.L2"

COLV.0S2.L2"

COLH.QS1B.L2

TP2"

IP3'

IP5

TP6

COLV.QS5.R6"

98.700

33.2

0

. 00

. 386

11.6

30.0

COLH.QF23.L2" -637.260

COLH.QD20.L2" -522.170

COLH.OS17.L2" -419.630

COLH.QS15.L2" -356.600

Detailed info, example of my records



Page:

Mecintosh HD:Heimut:fortran:MAD:worktwiss.out Friday, 18 July 1997 / 9:27

emittance in x- 45,000 in y- 4.500 nm

dist to IP

175.458

-220.540

-129.440

-98,660

-66.110

61.970

-21.320

-15,150

-8.750

.000

[m]

.000

twiss=n0520p97v2 Ebeam=45.60 GeV assumed from twiss file name

[m]

25.3

57.3

115.7

25.4

67.3

9.1

31.3

36.6

20.9

132.0

33.1

63.8

2.0

131.1

146.9

beta disp number of

B2 BI bends

782

808

808

12

12

24

24

24

24

24

24

24

24 24

24 24

[m]

.00 612

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. 60 791 24

.84 798

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.00

1.11

.000 3332.360 6664.720 9997.080 13329.400 16661.800 19994.200 23326.500 EndLep

sawtooth sigmabeam opening nsig

[mm]

23.3

32.4

16.0

24.7

34.9

9.3

17.2

12.2

23.1

14.0

11.6

16.1

35.2

14.5

12.5

12.5

12.5

12.5

14.5

14.5

30.0

30.0

14.5

30.0

30.0

14.5

[mm]

1.066

1.605

2.537

1.245

1.931

2.738

1.187

.640

.406

.771

.969

386

.536

. 300

2.429

opening

disp

.029

.027

.030

.037

.037

.030

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.029

.029

.026

.029

.037

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.029

.026

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.187

.187

.029

.027

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.034

deltaE/E= 1.000E-03

[mm]

00

.00

-.74

-.41

-.58

-.66

.00

.00

.00

.00

.00

.00

.00

sawtooth, sign is for positrons, opposite for electrons, collimator position increased by ABS(SAWT)

Page: 1

Macintosh HD:Helmut:fortran:MAD:worktwiss.out Friday, 18 July 1997 / 9:27

COLH.QS6.R6"	129.400	26.7	.00	0	0	.00	1.096	15.9	14.5	
COLH.QS10.R6"	220.500	11.3	.00	0	0	.00	,714	10.4	14.5	
COLH.QS15.R6"	356,600	107.1	.89	6	0	, 60	2.370	30,2	12.5	.034
COLH.QS17.R6"	419.600	79.0	.93	10	0	. 62	2.102	26.9	12.5	.029
COLH.QD20.R6"	522.200	26.4	.60	17	0	. 40	1.245	16.0	12.5	.027
COLH.QF23.R6"	637.300	115.5	1.11	26	0	.71	2,535	32.4	12.5	.029
IP7"	.000	25.2	.00	208	0	.00	1.066			
COLH.QF23.L8"	-637.200	115.8	1.11	390	0	.06	2.537	31.8	12.5	.029
COLH.QD20.L8"	-522.200	26.4	.60	399	0	.02	1.245	15.6	12.5	.026
COLH.QS17.L8"	-419.600	62.1	. 79	406	0	.02	1.850	23.1	12.5	.029
COLH,QS15,L8"	-356.600	136.6	.88	410	0	.02	2.632	32.9	12.5	.037
COLH,QS10.L8"	-220.500	17.3	.00	416	0	,00	. 883	12.8	14.5	
COLH.QS6.L8	-108.000	16.8	.00	416	0	.00	.871	12.6	14.5	
COLV.QS5.L8"	-79,200	43.6	.00	416	0	.00	. 443	13.3	30.0	
COLZ.QS3A.L8"	-66.100	73.8	.00	416	0	.00	.576	17.3	30.0	
COLH.QS3B.L8"	-56.300	46.9	.00	416	0	.00	1,452	21.1	14.5	
COLZ.QS2.L8*	-21.300	28.2	.00	416	0	.00	. 356	10.7	30.0	
COLV.QS2.L8"	-15.100	63.5	.00	416	0	, 00	. 535	16.0	30.0	
COLH.QS1B.L8"	-8.500	120.2	.00	416	0	. 00	2.326	33.7	14.5	
IP8"	.000	2.0	.00	416	0	.00	. 300			
COLH.QS1B.R8"	8.600	120.1	.00	416	0	.00	2.325	33.7	14.5	
COLV.QS2.R8"	15.200	63.9	.00	416	0	.00	.536	16.1	30.0	
COLZ.QS2.R8"	21.300	28.8	.00	416	0	.00	.360	10.8	30.0	
COLH, QS3B. R8"	56.300	46.9	.00	416	0	.00	1.452	21.1	14.5	
COLZ.QS3A.R8"	66.100	70.8	.00	416	0	.00	. 564	16.9	30.0	
COLV.QS5.R8"	79.200	42.6	.00	416	0	.00	. 438	13.1	30.0	
COLH.QS6.R8"	108.000	16.8	.00	416	0	,00	.870	12.6	14.5	
COLH.QS10.R8"	220.500	17.3	.00	416	0	.00	.883	12.8	14.5	
COLH.QS15.R8"	356.600	136.8	, 88	422	0	.00	2.633	32,9	12.5	.037
COLH.QS17.R8"	419.600	62.2	.79	426	0	01	1.850	23.1	12.5	.029
COLH.QD20.R8"	522.200	26.4	.60	433	0	01	1.245	15.6	12.5	.026
COLH.QF23.R8"	637.300	115.6	1.11	442	0	04	2.536	31.7	12.5	.029
COLH.QL8.L1	-175.500	57,2	.00	612	12	.00	1.605	23.3	14.5	
end of LEP NB	2,NBI= 6	12 12								
there were in	total NB2	tot= 16	40 star	ndard	bends	called	B2L or B2M or	B2R or I	B2S	
there were in	total NBI	tot=	24 inje	ection	bends	called	BI			
there were in	total NBW	tot=	32 weal	k bend	s calle	ed BW1 (or BW2 or BW3 o	or BW4		

Also kept: full set of LEP mad8 optics files my logging of LEP snapshots avery 15 min, used 2019 to re-analyze lumi calibration, σ_Z , # Neutrinos, see <u>doi</u>

FILL 6811	TIME 11.3794	Ie+ 1.60178	Ie- 1.63328	EWIG .813	.000	e+y .000	e-x .000	e-y .000	L 3.96	.3 ALEPH 5 5.047	0PAL 1.795	DELP 5.3	HI 10
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681	1 03-04	-00 10:4	5:33 adju	ıst		45.6	20		5	g0520b	٥ <u>99</u> ر	71	8
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131.1 .00 8.750 .00 2.429 35.2 14.5 COLH.OS1B.R2' 15.150 63.8 .00 .00 .536 16.1 30.0 COLV. 0S2. R2" COLZ.QS2.R2" 33.1 .00 .00 11.6 21,310 .386 30.0 .00 .00 COLH.QS4.R2" 61.970 20.9 .969 14.0 14.5 132.0 .00 .00 .771 23.1 COLZ.0S4.R2* 66.110 30.0 .00 .00 .406 12.2 30.0 COLV.QS5.R2" 98.660 36.6 .00 COLH.OS6.R2" 129.440 31.3 .00 0 0 1.187 17.2 14.5 .00 .00 .640 9.3 14.5 COT.H.QS10.R2 220.540 9.1 0 .QS15.R2" 356,600 146.9 .94 . 64 2.738 34.9 12.5 6 0 . 84 10 17 1.931 $12.5 \\ 12.5$ CULH.0S17.R2" 419,630 67.3 24.7 .40 .60 1,245 16.0 COLH.QD20.R2" 522.170 26.4 1.11 12.5 115.7 26 2.537 32.4 COLH.QF23.R2" 637.260 .00 1.066 25.3 208 .000 .00 115.6 .06 12.5 COLH.OF23.L4" -637.260 1.11 390 2.536 31.8 .02 1.245 15.6 COLH.0D20.L4" -522.170 26.4 399 12.5 .60 62.2 .79 .02 23.1 12.5 COLH.0S17.L4" -419.640 406 1.850 .88 410 .02 2.633 32.9 12.5 COLH.QS15.L4" -356,600 136.8 COLH.QS10.L4 -220.490 17.3 .00 416 .00 ,883 12.8 14.5 .00 416 , 00 .870 12.6 14.5 COLH.OS6.L4" -108.00016.8 COLV.QS5.L4" -79.18042.7 .00 416 .00 .438 13.1 30.0 COLZ.QS3A.L4" -66.110 70.8 .00 416 .00 .565 16.9 30.0 COLH.QS3B.L4" -56.280 46.9 .00 416 .00 1.452 21.1 14.5 .00 .360 COLZ.0S2.L4" -21.32028.B 416 10.8 30.0 .00 .00 .536 COLV.OS2.L4" -15.15063.9 416 16.1 30.0 .00 .00 2.325 33.7 COLH.QS1B.L4 -8.550 120.1 416 14.5 .00 416 ,00 .300 2.0 TP4 .000 COLH.QS1B.R4 8.520 120.2 .00 416 .00 2.326 33.7 14.5 15.120 63.5 .00 416 .00 16.0 .535 30.0 COLV. 0S2. R4 21.320 28.2 .00 416 .00 .356 10.7 30.0 COLZ.0S2.R4' COLH.QS3B.R4" 56.320 46.9 .00 416 . 00 1.452 21.1 14.5 COLZ.QS3A.R4" 66.120 73.8 .00 416 , 00 .576 17.3 30.0 COLV. 055. R4* 79.220 43.6 .00 416 .00 .443 13.3 30.0 COLH, OS6, R4" 108.020 16.8 .00 416 ,00 .871 12.6 14.5 .883 CCT 4.0510.R4" 220.520 17.3 .00 416 .00 12.8 14.5 2.632 . 88 .00 .QS15.R4" 356.620 136.6 422 32.9 12.5 .79 426 -.01 1.850 23.1 12.5 COLH.QS17.R4" 419.620 62.1 433 26.4 . 60 -.01 1.245 12.5 COLH. OD20. R4" 15.6 522.120 COLH. OF23. R4" 115.8 1.11 442 -.04 2.537 31.7 12.5 637.220 12.5 COLH.QL13.L5" -299.100622 -.04 1.987 24.9 87.3 .13 25.2 .00 624 .00 1.066 .000 .00 13.3 12.5 COLH. IP5" . 200 25.2 624 .00 1.066 50.5 .00 .00 COLV1.QL8.R5" 168.400 624 .477 11.9 25.0 COLV2.QL8.R5" 176.100 50.2 .00 624 .00 .520 13.0 25.0 COLV.QLS .R5" .00 624 .00 .305 7.6 25.0 195.400 20.5 COLH.QL13.R5" 87.5 .13 626 -.04 1.989 24.9 12.5 299.200 COLV, QD20, R5" 519.300 153.5 .00 641 .00 .831 20.8 25.0 COLV.QD30.R5" 914.300 149.1 .00 671 .00 .819 20.5 25.0 .00 .801 2.535 COLV.QD40.R5" 1309.300 142.6 .00 701 20.0 25.0 1.11 806 COLH.QF23.L6" -637.300 115.532.4 12.5 COLH.QD20.L6" -522.200 26.4 .60 815 -.38 $1.245 \\ 2.102$ 16.0 12.5 COLH.QS17.L6" -419.600 .93 822 -. 61 26.9 12.5 COLH.OS15.L6" -356.600 107.1 .89 826 -.59 2.370 30.2 12.5 ,00 .00 10.4 COLH.OS10.L6" -220.500 11.3 832 .714 14.5 832 .00 1.096 COLH.OS6.L6" -129.50026.7 .00 15.9 14.5 33.2 .00 COLV.QS5.L6" .00 832 .386 11.6 30.0 -98.700 .00 .00 .758 COLZ.QS4.L6" 66.100 127.7 832 22.7 30.0 COLH.QS4.L6" 20.9 .00 .00 .970 14.1 62.000 832 14.5 36.2 .00 .00 ,403 12.1 COLZ.QS2.L6" -21.300 832 30.0 COLV.0S2.16" -15.300 67.9 .00 832 .00 .553 16.6 30.0 COLH.QS1B.L6* -8.700 124.6 832 2.368 34.3 14.5 .000 2.0 .00 0 .00 . 300 2.368 .553 .403 COLH.OS1B.R6 8.700 124.6 ,00 0 34.3 14.5 15.300 .00 0 0 .00 16.6 COLV.OS2.R6" 67.9 30.0 ,00 .00 21,300 36.2 0 30.0 COLZ. OS2. R6' 62.000 .00 0 .00 .970 COLH.OS4.R6 20.9 0 14.1 14.5 127.7 COLZ.QS4.R6" 66.100 .00 0 0 .00 .758 22.7 30.0

4



High energy e+e- rings "touchy and potentially non-reproducible"



LHC (proton, ion, cold machine) much more stable and reproducible than LEP

Synchrotron Radiation, Beamstrahlung, HOM source of major local heating

Weak magnetic fields as needed to limit SR Spurious fields : earth field and leakage-(earth) currents, magnetic materials Flat beams — by design very small V/H emittance ratio, can easily be spoiled by coupling and spurious vertical dispersion ; dispersion free steering / golden orbit need to re-align LEP every year

large size and number of components : hard to exclude polarity / cables / poor contact issues

Lifetime/background with strong beam-beam very tune sensitive $\sim 10^{-4}$

A major source of coupling observed in the early LEP octant test : identified as caused by Ni-component in glue used for Pb-shielding around pipe minimized by de-magnetization + choice of coupling friendly optics choice of integer part of tunes and H/V difference important

Very small momentum-compaction

- tiny changes in circumference change energy (earth tides, rain / water level)



Peak performance





Performance increases steadily (slowly) over many years

arguably more than in pp machines where the beam brightness is made by the injectors **key role in IR design / MDI**

minimum β^* and maximum tune shift were limited in LEP by the need for stable low background running conditions





betatron injection/accumulation at 20 GeV, later 22 GeV, synchrotron injection

followed by ramp & squeeze with coarse collimation, physics with tight collimation

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FCC-ee, CEPC instead plan to work with top-up injection

advantage :

no loss in physics time for injection, ramp squeeze

extra challenges :

- need for more aperture to efficiently capture beams
- background spikes by losses and lager amplitude (halo) from injection
- continuously running at top maximum intensity and power

LEP beams were typically dumped after some hours when the luminosity had fallen (at constant §y linearly with current) backgrounds and beam sizes decreased, stability increased end of fill very useful for tuning improve golden orbits more luminosity / less background



Beam-gas, thermal photons, off momentum background





Figs. from my contribution to Landolt-Börnstein New Series I/21C

0.9

 $k = E_{\gamma} / E_{b}$

0.8

0.7

- Beamstrahlung ~ negligible @ LEP at high energy elastic scattering small inelastic generator of off-momentum tail well visible in LEP generally not a major problem (<< 1 electron lost at IR / crossing) thanks to
- excellent vacuum
- powerful momentum collimation

both in dedicated collimation section + local each IR

9



LEP, example of background particle tracking





Illustration of beam particle tracking through the LEP lattice over 1000 meters up to an experimental region (cs coordinates). The distance X from the nominal orbit is given in cm units.

The tracks are for particles that are lost within ± 9 m from the interaction point. The 12 σ beam envelope is shown as broken line.

The physical aperture limitation given by the beam pipes is shaded.

The position of collimators (called COLH.QS15, COLH.QS17..) as used in LEP physics runs is shown as vertical straight lines.

Codes : <u>MAD8</u>, <u>Turtle</u>, <u>DIMAD</u>, <u>EGS</u> + "own generators" beam gas, <u>thermal</u>, <u>SR</u>, <u>radiative Bbhabha</u>



LEP movable collimators, essential for background



11







Collimating high energy e+, e– will generate muons, roughly at the 10⁻⁴ – 10⁻⁵ level



G4GammaConversionToMuons



G4AnnihiToMuPair

implemented in GEANT4 in 2002

Came as a bad surprise to SLC, hard to avoid in linear colliders

Carefully studied for CLIC, hard (long magnetised shielding) to reduce

CLIC Muon Sweeper Design, Aloev, H.B. et al., and Belgin Pilicer thesis

Not an issue for LEP-MDI since losses were collimated far from the experiments

In CEPC, FCC-Z we expect to lose several 10¹¹ e+, e- per second generating millions of muons / second —> minimize collimation of e+, e- in line of sight to experiments



Signal exchange, monitoring, logging and status displays



For good performance, LEP beams required continuous monitoring and tuning including hundreds of orbit corrections during a fill

Also here MDI essential

pioneered for SPS ppbar,

also important for LHC \rightarrow

Using normalised background signals

- 5: maximum tolerable
 - and upper limit to declare stable beam
- 1: and below meaning very good

allow for several background signals / experiment

BKG1 more sensitive to photon

BKG2 more beam loss or more beam 2



LEP Run 8984 -** STABLE	CERN SI data o BEAMS	02-11 F:02-11 **	-00 08 -00 08	: 00: 26 : 00: 17
E = 105.000 Ge Beams I(t) uA tau(t) h	V/c Bea e+ 0.0 0.00	am In '	Coast: e- 0,0 0,00	0.5 h
LUMINOSITIES L(t) cm-2%s-1 /L(t) nb-1 Bkg 1 Bkg 2	L3 48.5 78.1 0.67 0.72	ALEPH 43.6 77.5 3.50 1.19	0PAL 42.7 78.1 4.71 1.03	DELPHI 47.7 79.4 2.37 4.88
COMMENTS 02- COLLIMATORS A PS: Thanks a 1	11-00 T PHYS ot for	07:49 ICS SET all th	TINGS ese le:	otons
dumping LEP be Will go to max negative frequ	am at imum e lency s	approx. nergy w hift	8:00 H ith a) !



Optimize collisions (1/2)



LEP beam separated during injection

ramp & squeeze

using electrostatic separators



Collisions optimised by separation scans based on luminosity

avoid partial separation :

reduces luminosity, can trigger coherent beam-beam, flip-flop, increase halo

Optimize collisions (2/2)

Beam Lifetime

Losses add \rightarrow inverse lifetimes add $1/\tau_{tot} = \sum \tau_i$

Example LEP2 fill 5259 4/10/1998, Eb = 94.5 GeV

 τ_{th} = 60 h predicted for thermal photons

 τ_{bg} = 80 h beam gas, 0.6 nTorr

34 h as measured before collisions

Colliding :

Lifetime dropped from 34 h to 5 h

and slowly increased to 7.5 h towards the end of the fill matching well with the expectation for a collision cross section of

 $\sigma_{\rm bb} = 0.215 \ {\rm barn} \quad \tau_{\rm bb} = 0.44 \ / \ \xi_{\rm y}$

ref: R. Kleiss, H.B. <u>BBBrem</u>

Lifetimes in LEP well accounted for by 3 loss processes Thermal photon, beam-gas, radiative Bhabha (when colliding) with occasionally (LEP1, high ξ_y) additional losses and background spikes related to non-Gaussian tails and coherent instabilities

Distribution of arrival times relative to the maximum drift time for the hit wire of SR photons in the OPAL vertex drift chamber

X-Ray - Fluorescence and Specular Reflection

Fluorescence was known and mitigated for absorbers by surface coating in LEP Fluorescence + Auger cascade are implemented in GEANT4 make sure materials well defined and these processes activated

X-Ray specular reflection came as a surprise in LEP mitigated with far coll. like COLH.QS6 at 120 m

currently working hard to (finally) implement this in GEANT4 presented at annual G4 meeting end Sept. meanwhile merged into G4 developer repository checks ongoing for upcoming public Geant4 Release 11.2, also of more general interest including x-ray mirrors, space science

Important to take these properly into account

Requires good knowledge and optimization of materials, geometry, surfaces (roughness) and tight control of orbit / alignment tolerances Ideally addressed by dedicated work with benchmarking measurents + simulation

Tails from : beam-beam, high chromaticity, particle scatteringBackground spikes, enhanced synchrotron radiation from quadruples

H.B. I. Reichel, G. Roy, Transverse beam tails due to inelastic scattering in LEP, <u>PRSTAB</u>, <u>3:091001</u>, 2000; I. Reichel, <u>CERN-Thesis-98-017</u> H.B. "Beam lifetime and beam tails in LEP." <u>Ref [2]</u>

LEP, as example of an IR optimized for SR

Eb = 45 GeV to 104 GeVlast bend 260 m, from IP 10 × weakerMachine induced backgrounds, MIB in LEP~ 100 collimators to reduce MIB8.5 m COLH.QS1 instrumented as compact Si-tungsten calorimeter - loss and IR rate monitorflat, symmetric machine, no crossing angle, few (4-12) bunchesSynchrotron radiation - no direct and single reflected radiation to experiments in IP regionOff-momentumbeam-gas and thermal photon

[1] Very High-Energy e+e- colliding beams.., B. Richter, <u>NIM 136:47</u> 1976

[2] LEP design report, Vol II, <u>CERN-LEP-81-01</u> 1984; Vol III LEP2, <u>CERN-AC/96-01</u>, 1996

[3] Test of EW theory at the Z resonance, H.B., J.Steinberger, Annu.Rev.Nucl.Part.Sci.41 (1991) 55

[4] Study of beam induced particle backgrounds.., G.v. Holtey, A.Ball et al., <u>NIM A403</u>, 1998

[5] Accelerator Physics at LEP, D. Brandt, H.B., M. Lamont, S. Myers, J. Wenninger, Rept.Prog.Phys.63, 2000

[6] Ultimate performance of the LEP RF system, P. Brown et al. PAC-2001-MPPH123

[7] A retrospective on LEP, H. B., J. Jowett, ICFA Beam Dyn.Newslett.48:143-152, 2009

Pictures & Anecdotes :

Running the LEP Machine, Steve Myers, Mike Lamont, John Poole, H.B., <u>The Aleph Experience, CERN 2005</u> The Greatest Lepton Collider, Steve Myers, <u>Colloquium for the 30th anniversary of the start of LEP</u>, 2019

https://home.cern/news/press-release/cern/lep-story https://cerncourier.com/a/the-greatest-lepton-collider/

LEP worked very well, in many respects and particular energy precision better than anticipated encouraging as demonstrator for CEPC / FCC feasibility

Much of the work is on details

- **MDI IR design particularly important**
- minimize SR and losses towards IR
- last upstream bend aim for Ecr < 100 keV, keep bend SR and loss collimation far from IP

Importance of excellent continuous monitoring + frequent correction + alignment close collaboration with experiments, minimize vertical dispersion + coupling;

LEP had a octant test 2 years and modest pilot run 1 year before start of proper operation A larger (inner) r = 78 mm Al beam pipe at the IR for the first year, later r = 53 mm Be

Goals very ambitious — need for testing & ramp-up strategy

Backup

Energy calibration

Figure 29. Magnetic field measured in a LEP dipole by an NMR probe over 10 h. For convenience, the magnetic field has been converted to an equivalent beam energy in MeV. Large short-term fluctuations and a slow rise in field are clearly visible. Between midnight and 4:30 am the field is stable while the fluctuations disappear.

Figures from [5]

Amon the key persons on machine side :

J. Wenninger, Ralph Assmann,

Bernd Dehning (Polarimeter)

Figure 30. Energy variation of the LEP beams during a full moon day. The curve is the energy change predicted from the horizontal strain induced by the Earth tides.

Figure 31. Evolution of the LEP circumference (corrected for tidal changes) as a function of the day in 1999. A drift of over to 2 mm is observed during the LEP run. In the summer months the circumference increases gradually. Following periods of heavy rainfall, indicated by the arrows, the circumference shrinks for some time before expanding again.

LEP peak performance parameters

Table 1: LEP beam parameters corresponding to the best performances at three differentenergies. The luminosities and beam-beam tune shifts are averaged over a time interval of 15minutes. For each beam energy, the first line corresponds to the horizontal, the second line to
the vertical plane.

E _b (GeV)	$N_b \ (\times 10^{11})$	k_b	\mathcal{L} (cm ⁻¹ s ⁻²)	Q_s	Q	β* (m)	<i>е</i> (nm)	σ (μm)	ξ
45.6	1.18	8	1.51×10^{31}	0.065	90.31	2.0	19.3	197	0.030
					76.17	0.05	0.23	3.4	0.044
65	2.20	4	2.11×10^{31}	0.076	90.26	2.5	24.3	247	0.029
					76.17	0.05	0.16	2.8	0.051
97.8	4.01	4	9.73×10^{31}	0.116	98.34	1.5	21.1	178	0.043
					96.18	0.05	0.22	3.3	0.079

Table 2: Overview of LEP (instantaneous) peak performance 1989-1999. $\int \mathcal{L} dt$ is the luminosity integrated per experiment over each year. The design luminosity at 45 GeV was 17×10^{30} cm⁻²s⁻¹.

Year	$\int \mathcal{L} dt $ (pb ⁻¹)	E_b (GeV/c ²)	k_b	$2k_bI_b$ (mA)	\mathcal{L} (10 ³⁰ cm ⁻² s ⁻¹)	ξ_{y}
1989	1.74	45.6	4	2.6	4.3	0.017
1990	8.6	45.6	4	3.6	7	0.020
1991	18.9	45.6	4	3.7	10	0.27
1992	28.6	45.6	4/8	5.0	11.5	0.027
1993	40.0	45.6	8	5.5	19	0.040
1994	64.5	45.6	8	5.5	23.1	0.047
1995	46.1	45.6	8/12	8.4	34.1	0.030
1996	24.7	80.5 to 86	4	4.2	35.6	0.040
1997	73.4	90 to 92	4	5.2	47.0	0.055
1998	199.7	94.5	4	6.1	100	0.075
1999	253	98 to 101	4	6.2	100	0.083
2000	233.4	102 - 104	4	5.2	60	0.055

from <u>Ref 7</u>

LEP, LHC built in the same tunnel, 26658.9 m circumference
LEP as single ring, single beam pipe
LHC two pipes in twin magnets separated by 19.4 cm
FCC-ee two rings separated by 30 cm

8 straight sections, ± 284 m around IPs

4 used as interaction regions

distance IP to 1st superconducting Quadrupole (centre) L* = 3.7 m for LEP 2.8 m FCC-ee 23 m for LHC

2 rings : allow for many bunches without parasitic collisions
disadvantage : less evident to find collisions, need to frequently re-steer to centre collisions —>

 $\int L dt / year$ increase by ~ 10 × over 10 years

2e-62

6.5e-08

6.5e-08

30

164	BW4.QS12.R2 RBEND	272.1	11.55	0.0003768	72.37	0.7767	30652.0	0.0109	33.8668	1.2293	0.0379	0.04989
172	B2L.QS12.R2 RBEND	287.3	11.55	0.003768	723.7	7.767	3065.2	0.1088	88.0931	1.9827	0.0637	4.989
174	B2R.QS13.R2 RBEND	299.2	11.55	0.003768	723.7	7.767	3065.2	0.1088 1	63.5957	2.7019	0.0636	4.989

Quads, at 1 sigmax, horizontal

iele	Element	S	\mathbf{L}	betx	sigx	divx	K1L	k0	х	Angle	Ecrit	ngam	Power
		m	m	m	mm	mrad	m-2	m-1	mm		keV		kW
2	QS0.R2	5.7	2	27.8	1.115	0.04003	-0.327	0.0003474	-0.0524	0.0006948	770.7	1.432	0.9798
10	QS1B.R2	11.2	2	226	3.176	0.01405	0.06314	0.0001918	-0.1377	0.0003836	425.5	0.7907	0.2987
12	QS1A.R2	13.7	2	278	3.523	0.01267	0.06314	0.0002129	-0.1509	0.0004259	472.4	0.8778	0.3681
20	QS2.R2	18	1.6	276	3.507	0.01272	0.01788	6.006e-05	-0.1471	9.61e-05	133.2	0.1981	0.023423
36	QS3.R2	59	2	39.4	1.326	0.03366	0.01879	2.45e-05	-0.02171	4.9e-05	54.35	0.101	0.004873