

Science and Technology Facilities Council



Neutrino Interactions and Future Experiments with GeV Neutrinos

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Neutrino Mass

Beyond Standard Model Standard Model Pontecorvo-Maki-Nakagawa-Sakata $\begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$ $\left(egin{array}{c}
u_e
ight)
u_\mu \end{array}
ight)$ PMNS matrix ± **Mass Ordering** Normal Inverted

PMNS Matrix



 θ_{12} : mixing between ν_1 and ν_2

 $\theta_{23}\text{:}$ mixing between ν_{μ} and ν_{τ}

 θ_{13} : if 0, effective 2 flavour mixing

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} \nu_1' \\ \nu_2' \\ \nu_3 \end{pmatrix} \qquad P(\nu_\alpha) + P(\nu_\beta) = 1$$

Antineutrinos

2-flavor oscillation

$$v_{\beta}$$

$$P(v_{\alpha}) + P(v_{\beta}) = 1$$

$$v_{\alpha}$$



Oscillation as a function of *time* line-in-**line** → same trivia

Antineutrinos

3-flavor oscillation





Oscillation as a function of *time* line-in-**plane** \rightarrow CP-violation possible

CP Violation

Neutrino oscillations depend on mixing parameters and mass differences.

 $\begin{aligned} c_{ij} &= \cos\theta_{ij} \\ s_{ij} &= \sin\theta_{ij} \\ \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{bmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \\ \theta_{13} \neq 0 \to \delta_{CP} \text{ can be observed} \end{aligned}$

Appearance probability
of
$$\mathbf{v}_e$$
 in a \mathbf{v}_{μ} beam $P(\nu_{\mu} \to \nu_e) \simeq \sin^2 2\theta_{13} \sin^2 \Delta_{32} \left(\sin^2 \theta_{23} - \frac{\sin 2\theta_{12} \sin 2\theta_{23}}{2 \sin \theta_{13}} \sin \delta_{CP} \sin \Delta_{21} \right)$

CP-odd term

* neglecting matter effects

CP Violation

Neutrino oscillations depend on mixing parameters and mass differences.

 $\begin{aligned} c_{ij} &= \cos\theta_{ij} \\ s_{ij} &= \sin\theta_{ij} \\ \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{bmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \\ \theta_{13} \neq \mathbf{0} \rightarrow \delta_{CP} \text{ can be observed} \end{aligned}$ Appearance probability of $\bar{\nu}_e$ in a $\bar{\nu}_\mu$ beam $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \simeq \sin^2 2\theta_{13} \sin^2 \Delta_{32} \left(\sin^2 \theta_{23} + \frac{\sin 2\theta_{12} \sin 2\theta_{23}}{2\sin \theta_{13}} \sin \delta_{CP} \sin \Delta_{21} \right)$

flip sign

 $\delta_{CP} \rightarrow CP$ violation

CP violation: electron flavor appears from muon-flavor neutrinos and antineutrinos differently.

* neglecting matter effects

Neutrinos Sources



Accelerator Neutrinos



Ad: K.-J. Plows and XL, Modeling heavy neutral leptons in accelerator beamlines, Phys.Rev.D 107, 055003 (2023).



" β decay" of energetic collision products (mostly ν_{μ} from π)



Atmospheric Neutrinos





Near Detectors





Water Cherenkov detector



Water Cherenkov detector





18 April, 2023

Source: http://www.ps.uci.edu/~tomba/sk/tscan/tl



Liquid argon Time Projection Chamber (LArTPC)

The New Impressive

Future oscillation experiments

This talk only on

* accelerator and atmospheric GeV- ν

✤ ν_{μ} flux*: ν_{μ} disappear, ν_{e} appear



Interaction inside nuclei

 $\Box v_{\mu e}$ Charged Current (CC) for v detection

 \Box GeV- ν interaction: ν **N** interaction embedded in *nuclei* (A)





Medium effects—source of systematics

- ✓ v energy reconstruction, event classification
- □ Through initial state, vertex, final state
- Fermi motion & nuclear potential
- NN correlations
 Multinucleon excitation
- Pauli-blocking
 FSI



Counting oscillated v

At *(far-)detector*, interactions *cannot* be measured with *unknown oscillated flux*

Measurement = (*flux* × *interaction*) \oplus **detector effects**





Detector

- * Near detectors for accelerator- ν experiments
- Non-accel: rely on externally constrained models
- Unconstrained flavour and/or target nuclei

No two unknowns at the same time

2

Medium

effects



□ Also *active target*

Tracking + calorimetry

Hydrogen target from CH-C subtraction: T. Cai *et al.* [MINERvA], Measurement of the axial vector form

factor from antineutrino-proton scattering, *Nature* 614, 48 (2023)

Hyper-Kamiokande

Current role in studying ν interactions \Box Largest data set

Systematic investigation cf. e.g. MINERVA, Eur. Phys. J. ST 230, 4243 (2021)



Detector



Cube assembly and fiber insertion ...





Weijun Li 利伟君 (Oxford/Warwick) January 2023, J-PARC



Detector

□ FD (Far Detector)

DUNE

- LArTPC (Liquid Argon TPC)
- ✓ Mass-scalable for tracking + calo

□ Near Detector ND-LAr

- Same technology as FD
- Near Detector ND-GAr (Gaseous)

Argon)—Reference Design

- ✤ 10-bar argon-based gas TPC
- ~100 m³ gas volume surrounded by calorimeter
- ✤ B-field provides sign selection
- ✓ Large statistics of v interactions on gas
- \checkmark 4 π acceptance, very low tracking threshold
- ✓ Arguably ultimate <u>exclusivity</u> for v interactions

DUNE, instruments 5, 31 (2021)

Exclusivity: to measure all final states (except nuclear remnant)

Vessel (200 L 10 bar) for high pressure TPC R&D @ WarTPC lab



Matt Snape (Warwick) and Philip Hamacher-Baumann (Aachen/Warwick) August 2022, Warwick

DUNE

ProtoDUNE

Detector

LArTPC Demonstrator at CERN for DUNE FD

□ Hadron beams of 0.3-7 GeV/*c*

- ✤ 4.7 mm wire spacing (same as FD)
- ✓ Versatile reconstruction in LAr





e/γ separation

DUNE

ProtoDUNE

Detector

LArTPC Demonstrator at CERN for DUNE FD

□ Hadron beams of 0.3-7 GeV/c

- ✤ 4.7 mm wire spacing (same as FD)
- Versatile reconstruction in LAr
- ✓ hAr interactions to constrain *v*-int. FSI
- ✓ *Exclusivity* + beam energy, can "see" inside argon nuclei

<u>Exclusivity</u>: to measure all final states (except nuclear remnant)



Exclusive event candidates



DUNE, JINST 15, P12004 (2020)

Kinematic fitting improves π^0 energy resolution from 13% to 10%



Kang Yang 杨康 (Oxford/Warwick)



SBND 20~30 × current world ν Ar data

***** Large statistics for v_{μ} and v_{e}



Detector



SBND

Detector

$20~30 \times \text{current world } \nu\text{Ar data}$

 3 mm wire spacing (same as MicroBooNE and ICARUS)







SBND

Detector

$20~30 \times current$ world νAr data

- 3 mm wire spacing (same as MicroBooNE and ICARUS)
- ✓ Proton tracking threshold ~ 40 MeV (277 MeV/c)
- ✓ Proton tagging at vertex

<u>Exclusivity</u>: to measure all final states (except nuclear remnant)









Detector limit can be pushed, but inside of a nucleus is never allowed...







PRISM (Precision Reaction Independent Spectrum Measurement)

Detector



Hyper-Kamiokande

□ FD: water Cherenkov

□ ND: IWCD (Intermediate Water Cherenkov Detector)

- ✤ Same technology as FD
- ✤ 50 m vertical shaft @ 750 m from beam source
 - ✓ 1°-4° off-axis (OA) angle ("PRISM Definition Part 1")



Medium effects



Hyper-Kamiokande

□ FD: water Cherenkov

□ ND: IWCD (Intermediate Water Cherenkov Detector)

- Same technology as FD
- ✤ 50 m vertical shaft @ 750 m from beam source
 - ✓ 1°-4° off-axis (OA) angle ("PRISM Definition Part 1")
- ✤ ~ 1% residual v_e/\bar{v}_e beam components
 - ✓ Large fraction at far-OA angle
 - ✓ Constrain v_e / \overline{v}_e (besides v_μ / \overline{v}_μ) cross sections on water (enabled by active γ shielding)









Medium effects



DUNE-PRISM

Medium effects

- ND-LAr & ND-GAr
- ✤ Up to 30 m off axis @ 574 m from beam source
 - ✓ 0°-3° off-axis angle
 - \checkmark E_v up to ~ 3 GeV, covering different interaction dynamics
 - ✓ **Probe energy-dependent medium effects**

DUNE, instruments 5, 31 (2021)







v_e/\bar{v}_e interactions

 $\Box \delta_{\rm CP}$ requires v_e and \bar{v}_e appearance

✓ Suppress v_e and \bar{v}_e bkg in beams

 \Box Need v_e/\bar{v}_e interaction data

- $\Box v_{\mu}-A + \text{lepton universality constrains} \\ v_{e}-A \text{ to } 1^{\text{st}} \text{ order precision}$
- □ Oscillation requires 2nd order precision
 - Higher statistics and better-understood fluxes





v_e / \bar{v}_e interactions

□ δ_{CP} requires v_e and \bar{v}_e appearance ✓ Suppress v_e and \bar{v}_e bkg in beams

□ Need $\nu_e / \bar{\nu}_e$ interaction data

- $\Box v_{\mu} A + lepton universality constrains$ $v_e - A to 1st order precision$
- □ Oscillation requires 2nd order precision
 - Higher statistics and better-understood fluxes



- Enhanced NeUtrino BEams from kaon Tagging (ENUBET)
 - v_e from e^+ tagging for $K^+ \to \pi^0 e^+ v_e$
 - ♦ ν_{μ} from μ^+ tagging
 - ✤ Flux uncertainty ~ 1%



v_e / \overline{v}_e interactions

□ δ_{CP} requires ν_e and $\bar{\nu}_e$ appearance ✓ Suppress ν_e and $\bar{\nu}_e$ bkg in beams

□ Need $\nu_e / \bar{\nu}_e$ interaction data

 $\Box v_{\mu} - A + lepton universality constrains$ $v_e - A to 1st order precision$

□ Oscillation requires 2nd order precision

 Higher statistics and better-understood fluxes
 pustopm arXiv:2203.07545



Workshop: Exploring the Physics Opportunities of nuSTORM, London, 6 April 2023

- \Box ν from STORed Muons (nuSTORM)
 - $v_{\mu}/\bar{v}_{e}/\bar{v}_{\mu}/v_{e}$ fluxes from μ^{\pm} decays
 - ✓ 1% or better flux precision





NMO with atmospheric v $\Box v$ energy & angle for *L/E*-variation



GeV-v interaction more critical and challenging

Future Oscillation Experiment	E _v /GeV	Detector Technology	Target Nuclei
IceCube Upgrade	3-10 (NMO sensitive region)	Cherenkov in ice	H ₂ O
KM3NeT/ORCA		WC	H ₂ O
Atmos-ν @JUNO		LS	CH _{1.6}



NMO with atmospheric v

 \Box ν energy & angle for *L*/*E*-variation

- No near detector
 - flux × interaction ambiguity
- □ Sensitive to new unknowns
 - E.g. unconstrained low-momentum proton production (450 MeV/c common tracker threshold)
 - Impact on very-low-threshold calo

Future Oscillation Experiment	E_{ν} /GeV	Detector Technology	Target Nuclei
IceCube Upgrade	3-10 (NMO sensitive region)	Cherenkov in ice	H ₂ O
KM3NeT/ORCA		WC	H ₂ O
Atmos-ν @JUNO		LS	CH _{1.6}

Dedicated GeV-v interaction measurements: MINERvA Medium Energy data

- \checkmark E_v peak at 6 GeV, tail up to 20 GeV
- ✓ CH and nuclear targets
- ✓ ~ 10 M-event data set



Atmospheric neutrino interaction products: big surprise (fixed) in a very popular event generator (Interesting story: <u>https://github.com/GENIE-MC/Generator/issues/226</u>)



Qiyu Yan 严启宇 (UCAS/Warwick)

Summary

Future oscillation experiments require surgical precision inside a black box



Awaiting the future

Detector Technology: neutrons

- ✓ *v* energy budget and event classification—missing piece for <u>exclusivity</u>
- Tagging and calorimetry exist
- 4-momentum determination on the verge (e.g. time of flight)

Medium effects

Analysis methods: ν -hydrogen interaction

- ✓ Complete removal of medium effects
- Established: statistical subtraction between targets
- Ideas: <u>exclusivity</u> + TKI event-by-event selection using mass-scalable H-based compounds

Model constr't

Ex situ interaction measurements: precise nuclear response

- ✓ Break flux × interaction ambiguity
- Electron scattering + <u>exclusivity</u> for initial-and final-state effects (not vertex)

BACKUP