## 强耦合暗物质的相变与引力波

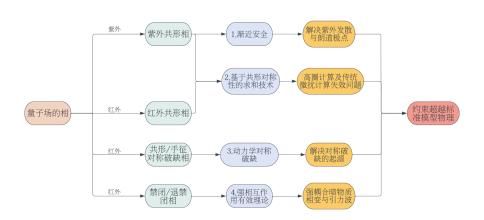
王志伟

电子科技大学

4月23日, 2023

国科大高能核物理课题组前沿讲座

## A Landscape of Phases in QFT and its Relation to BSM Physics



#### Motivations and what we do

- (Dark) composite dynamics: non perturbative physics, dynamical symmetry breaking, UV completion, naturalness
- (Dark) composite dynamics face challenges to be explored both theoretically and via experiments and thus any extra test is important
- We unify first principle lattice simulations and gravitational wave astronomy to constrain the dark sector

## What composes the strongly coupled sector?

- Dark Yang-Mills theories
- Pure gluons ⇒ confinement-deconfinement phase transition
- Gluons + Fermions
  - Fermions in fundamental representation ⇒ chiral phase transition
  - Fermions in adjoint rep. ⇒ confinement & chiral phase transition
  - Fermions in 2-index symmetric rep. ⇒ confinement & chiral phase transition
- Gluons + Fermions + Scalars (not explored yet)

## How to describe the strongly coupled sector?

#### Pure gluons

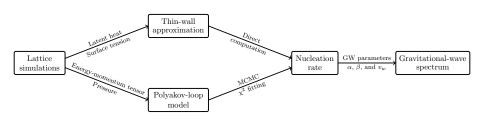
- Polyakov loop model (Huang, Reichert, Sannino and Z-WW, PRD 104 (2021) 035005;
   Kang, Zhu, Matsuzaki, JHEP 09 (2021) 060)
- Matrix Model (Halverson, Long, Maiti, Nelson, Salinas, JHEP 05 (2021) 154)
- Holographic QCD model (Ares, Henriksson, Hindmarsh, Hoyos, Jokela, PRD 105 (2022) 066020; Ares, Henriksson, Hindmarsh, Hoyos, Jokela, PRL 128 (2022) 131101)

#### Gluons + Fermions

- Polyakov loop improved Nambu-Jona-Lasinio model (Reichert, Sannino, Z-WW and Zhang, JHEP 01 (2022) 003; Helmboldt, Kubo, Woude, PRD 100 (2019) 055025)
- linear sigma model (Helmboldt, Kubo, Woude, PRD 100 (2019) 055025)
- Polyakov Quark Meson model (Schaefer, Pawlowski, Wambach, PRD 76 (2007) 074023)

## Procedure of pure gluon case

(Huang, Reichert, Sannino and Z-W W, PRD 104 (2021) 035005



## Polyakov Loop Model for Pure Gluons: I

- Pisarski first proposed the Polyakov-loop Model as an effective field theory to describe the confinement-deconfinement phase transition of SU(N) gauge theory (Pisarski, PRD 62 (2000) 111501).
- In a local SU(N) gauge theory, a global center symmetry Z(N) is used to distinguish confinement phase (unbroken phase) and deconfinement phase (broken phase)
- $\bullet$  An order parameter for the Z(N) symmetry is constructed using the Polyakov Loop (thermal Wilson line)  $_{\rm (Polyakov,\,PLB\,72\,(1978)\,477)}$

$$\mathbf{L}(\vec{x}) = \mathcal{P} \exp \left[ i \int_0^{1/T} A_4(\vec{x}, \tau) \, \mathrm{d}\tau \right]$$

The symbol  ${\cal P}$  denotes path ordering and  $A_4$  is the Euclidean temporal component of the gauge field

 $\bullet$  The Polyakov Loop transforms like an adjoint field under local SU(N) gauge transformations

## Polyakov Loop Model for Pure Gluons: II

ullet Convenient to define the trace of the Polyakov loop as an order parameter for the Z(N) symmetry

$$\ell\left(\vec{x}\right) = \frac{1}{N} \mathrm{Tr}_c[\mathbf{L}] \,,$$

where  ${\rm Tr}_c$  denotes the trace in the colour space.

 $\bullet$  Under a global Z(N) transformation, the Polyakov loop  $\ell$  transforms as a field with charge one

$$\ell \to e^{i\phi}\ell, \qquad \phi = \frac{2\pi j}{N}, \qquad j = 0, 1, \dots, (N-1)$$

• The expectation value of  $\ell$  i.e.  $< \ell >$  has the important property:

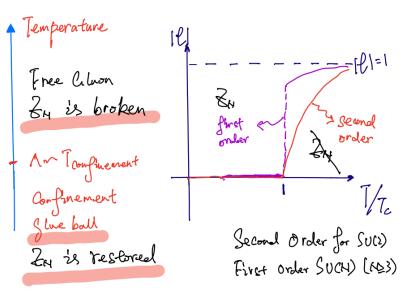
$$\langle \ell \rangle = 0 \quad (T < T_c, \text{ Confined}); \qquad \langle \ell \rangle > 0 \quad (T > T_c, \text{ Deconfined})$$

ullet At very high temperature, the vacua exhibit a N-fold degeneracy:

$$\langle \ell \rangle = \exp\left(i\frac{2\pi j}{N}\right)\ell_0, \qquad j = 0, 1, \dots, (N-1)$$

where  $\ell_0$  is defined to be real and  $\ell_0 \to 1$  as  $T \to \infty$ 

## Summary of Pure Gluon Facts



## Effective Potential of the Polyakov Loop Model: I

• The simplest effective potential preserving the  $Z_N$  symmetry in the polynomial form is given by (Pisarski, PRD 62 (2000) 111501)

$$V_{\rm PLM}^{(\rm poly)} = T^4 \left( -\frac{b_2(T)}{2} |\ell|^2 + b_4 |\ell|^4 + \dots - b_3 (\ell^N + \ell^{*N}) \right)$$
 where  $b_2(T) = a_0 + a_1 \left( \frac{T_0}{T} \right) + a_2 \left( \frac{T_0}{T} \right)^2 + a_3 \left( \frac{T_0}{T} \right)^3 + a_4 \left( \frac{T_0}{T} \right)^4$ 

- "..." represent any required lower dimension operator than  $\ell^N$  i.e.  $(\ell\ell^*)^k = |\ell|^{2k} \text{with } 2k < N.$
- $\bullet$  For the SU(3) case, there is also an alternative logarithmic form

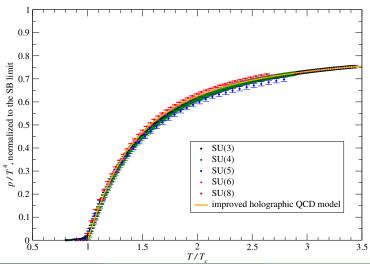
$$V_{\text{PLM}}^{(3\text{log})} = T^4 \left( -\frac{a(T)}{2} |\ell|^2 + b(T) \ln(1 - 6|\ell|^2 + 4(\ell^{*3} + \ell^3) - 3|\ell|^4) \right)$$
$$a(T) = a_0 + a_1 \left( \frac{T_0}{T} \right) + a_2 \left( \frac{T_0}{T} \right)^2 + a_3 \left( \frac{T_0}{T} \right)^3, \quad b(T) = b_3 \left( \frac{T_0}{T} \right)^3$$

• The  $a_i,\,b_i$  coefficients in  $V_{\rm PLM}^{({
m poly})}$  and  $V_{\rm PLM}^{(3{
m log})}$  are determined by fitting the lattice results

## Fitting the Coefficients Using the Lattice Results: I

Marco Panero, Phys.Rev.Lett. 103 (2009) 232001

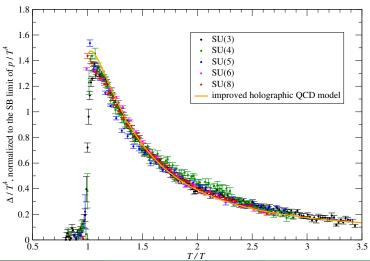




## Fitting the Coefficients Using the Lattice Results: II

Marco Panero, Phys.Rev.Lett. 103 (2009) 232001

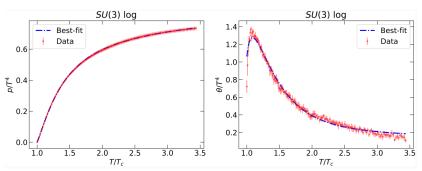
#### Trace of the energy-momentum tensor



## Fitting the Coefficients Using the Lattice Results: III

(Huang, Reichert, Sannino and Z-W W, PRD 104 (2021) 035005

Fitted to lattice data of pressure and the trace of energy momentum tensor.



## Fitting the Coefficients Using the Lattice Results: IV

(Huang, Reichert, Sannino and Z-W W, PRD 104 (2021) 035005

表: The parameters for the best-fit points.

N	3	3 log	4	5	6	8
$a_0$	3.72	4.26	9.51	14.3	16.6	28.7
$a_1$	-5.73	-6.53	-8.79	-14.2	-47.4	-69.8
$a_2$	8.49	22.8	10.1	6.40	108	134
$a_3$	-9.29	-4.10	-12.2	1.74	-147	-180
$a_4$	0.27		0.489	-10.1	51.9	56.1
$b_3$	2.40	-1.77		-5.61		
$b_4$	4.53		-2.46	-10.5	-54.8	-90.5
$b_6$			3.23		97.3	157
$b_8$					-43.5	-68.9

#### **Include Fermions**

(K. Fukushima, PLB 591 (2004) 277; Ratti, Thaler Weise, PRD 73 (2006) 014019)Reichert, Sannino, Z-WW and Zhang, JHEP 01 (2022) 003, arXiv:2109.11552.

- The Polyakov-loop-Nambu-Jona-Lasinio (PNJL) model is used to describe phase-transition dynamics in dark gauge-fermion sectors
- The finite-temperature grand potential of the PNJL models can be generically written as

$$V_{\rm PNJL} = V_{\rm PLM}[\ell, \ell^*] + V_{\rm cond} \left[ \langle \bar{\psi}\psi \rangle \right] + V_{\rm zero} \left[ \langle \bar{\psi}\psi \rangle \right] + V_{\rm medium} \left[ \langle \bar{\psi}\psi \rangle, \ell, \ell^* \right]$$

- ullet  $V_{\mathrm{PLM}}[\ell,\ell^*]$  is the Polyakov loop model potential (discussed above)
- ullet  $V_{
  m cond}ig[\langlear\psi\psi
  angleig]$  represents the condensate energy
- ullet  $V_{
  m zero}ig[\langlear\psi\psi
  angleig]$  denotes the fermion zero-point energy
- The medium potential  $V_{
  m medium}\left[\langle \bar{\psi}\psi \rangle,\ell,\ell^* \right]$  encodes the interactions between the chiral and gauge sector which arises from an integration over the quark fields coupled to a background gauge field

#### The PNJL model Lagrangian (Fukushima, Skokov, PPNP 96 (2017) 154)

The PNJL Lagrangian can be generically written as:

$$\mathcal{L}_{\mathrm{PNJL}} = \mathcal{L}_{\mathsf{pure-gauge}} + \mathcal{L}_{\mathrm{4F}} + \mathcal{L}_{\mathrm{6F}} + \mathcal{L}_{k}$$

- Without losing generality, we consider below massless 3-flavour case in fundamental representation of SU(3) gauge symmetry
- Here,  $\mathcal{L}_{4F}$  is the four-quark interaction which reads:

$$\mathcal{L}_{4F} = G_S \sum_{a=0}^{8} [(\bar{\psi}\lambda^a\psi)^2 + (\bar{\psi}i\gamma^5\lambda^a\psi)^2], \qquad \psi = (u, d, s)^T$$

• Six-fermion interaction  $\mathcal{L}_{6\mathrm{F}}$  denotes the Kobayashi-Maskawa-'t Hooft (KMT) term breaking  $U(1)_A$  down to  $Z_3$  (generically  $Z_{N_f}$  for  $N_f$  flavours)

$$\mathcal{L}_{6F} = G_D[\det(\bar{\psi}_{Li}\psi_{Rj}) + \det(\bar{\psi}_{Ri}\psi_{Lj})]$$

ullet In  $\mathcal{L}_{\mathrm{4F}}$ , the condensate energy then comes from the combination

$$(\bar{\psi}\lambda^0\psi)^2 + (\bar{\psi}\lambda^3\psi)^2 + (\bar{\psi}\lambda^8\psi)^2 = 2(\bar{u}u)^2 + 2(\bar{d}d)^2 + 2(\bar{s}s)^2$$

• We use the trick is to rewrite  $(\bar{u}u)^2$  as

$$(\bar{u}u)^{2} = [(\bar{u}u - \langle \bar{u}u \rangle) + \langle \bar{u}u \rangle]^{2} = (\bar{u}u - \langle \bar{u}u \rangle)^{2} + 2\langle \bar{u}u \rangle (\bar{u}u - \langle \bar{u}u \rangle) + \langle \bar{u}u \rangle^{2}$$
  

$$\simeq -\langle \bar{u}u \rangle^{2} + 2\langle \bar{u}u \rangle \bar{u}u,$$

where the  $(\bar{u}u - \langle \bar{u}u \rangle)^2$  term is dropped in the spirit of the mean-field approximation.

- The  $2\langle \bar{u}u\rangle \bar{u}u$  term contributes to the constituent quark mass of u
- The  $-\langle \bar{u}u \rangle^2$  term leads to a contribution to the condensate energy
- Similar procedures can be applied to  $(\bar{d}d)^2$  and  $(\bar{s}s)^2$ , and to  $\mathcal{L}_{6\mathrm{F}}$  gives  $\langle \bar{u}u \rangle^3$  and we obtain the total condensate energy:

$$V_{\rm cond} = 6G_S \sigma^2 + \frac{1}{2}G_D \sigma^3$$
,  $\sigma \equiv \langle \bar{u}u \rangle = \langle \bar{d}d \rangle = \langle \bar{s}s \rangle = \frac{1}{3}\langle \bar{\psi}\psi \rangle$ 

王志伟 (电子科技大学) 强耦合暗物质的相变与引力波 4月23日、2023 17/42

## The Constituent Quark Mass and Zero Point Energy: I

(Fukushima, Skokov, PPNP 96 (2017) 154)

- In  $\mathcal{L}_{6\mathrm{F}}$ , there is also  $\langle \bar{u}u \rangle^2 \bar{u}u$  term contributes to the constituent quark mass of u
- ullet The total constituent quark mass from  $\mathcal{L}_{\mathrm{4F}}$  and  $\mathcal{L}_{\mathrm{6F}}$  is:

$$M = -4G_S\sigma - \frac{1}{4}G_D\sigma^2$$

• The expression for the zero-point energy is given by:

$$V_{\text{zero}}[\langle \bar{\psi}\psi \rangle] = -\text{dim}(\mathbf{R}) \, 2N_f \int \frac{\mathrm{d}^3 p}{(2\pi)^3} E_p \,, \qquad E_p = \sqrt{\vec{p}^2 + M^2}$$

 $E_p$  is the energy of a free quark with constituent mass M and three-momentum  $\vec{p}$ 

- The above integration diverges and a regularization is required. We choose a sharp three-momentum cutoff  $\Lambda$  entering the expression for observables and thus also a parameter of the theory.
- Parameters:  $G_S, G_D, \Lambda$ ; Observables:  $M, f_{\pi}, m_{\sigma}$

(Fukushima, Skokov, PPNP 96 (2017) 154)

• The integration can be carried analytically and the result is:

$$\begin{split} V_{\mathsf{Zero}}\big[\langle\bar{\psi}\psi\rangle\big] &= -\frac{\dim(\mathbf{R})N_f\Lambda^4}{8\pi^2} \bigg[(2+\xi^2)\sqrt{1+\xi^2} \\ &\quad + \frac{\xi^4}{2}\ln\frac{\sqrt{1+\xi^2}-1}{\sqrt{1+\xi^2}+1}\bigg], \end{split}$$

in which  $\xi \equiv \frac{M}{\Lambda}$ .

## Medium Potential: Finite Temperature Contribution

- In the standard NJL model, the medium effect (finite temperature contribution) is implemented by the grand canonical partition function
- In the PNJL model, we can simply do the following replacement to include the contribution from Polyakov loop

$$\begin{split} V_{\text{medium}} &= -2N_c T \sum_{u,d,s} \int \frac{d^3 p}{(2\pi)^3} \left( \ln \left[ 1 + e^{-\beta(E-\mu)} \right] + \ln \left[ 1 + e^{-\beta(E+\mu)} \right] \right) \\ &\rightarrow -2T \sum_{u,d,s} \int \frac{d^3 p}{(2\pi)^3} \operatorname{Tr}_c \left\{ \left( \ln \left[ 1 + \mathbf{L} \, e^{-\beta(E-\mu)} \right] + \ln \left[ 1 + \mathbf{L}^{\dagger} e^{-\beta(E+\mu)} \right] \right) \right\} \end{split}$$

L is the Polyakov loop:

$$\mathbf{L}(\vec{x}) = \mathcal{P} \exp \left[ i \int_0^{1/T} A_4(\vec{x}, \tau) \, \mathrm{d}\tau \right]$$

• In this work, we consider chemical potential  $\mu = 0$ .



# Second Part: Bubble Nucleation and Gravitatioanl Wave

#### **Bubble Nucleation: Generic Discussion**

- In a first-order phase transition, the transition occurs via bubble nucleation and it is essential to compute the nucleation rate
- The tunnelling rate due to thermal fluctuations from the metastable vacuum to the stable one is suppressed by the three-dimensional Euclidean action  $S_3(T)$

$$\Gamma(T) = T^4 \left(\frac{S_3(T)}{2\pi T}\right)^{3/2} e^{-S_3(T)/T}$$

The generic three-dimensional Euclidean action reads

$$S_3(T) = 4\pi \int_0^\infty \! \mathrm{d}r \, r^2 \! \left\lceil \frac{1}{2} \! \left( \frac{\mathrm{d}\rho}{\mathrm{d}r} \right)^2 + V_{\mathrm{eff}}(\rho,T) \right\rceil \, , \label{eq:S3T}$$

where  $\rho$  denotes a generic scalar field with mass dimension one,  $[\rho]=1$ 

#### **Bubble Nucleation: Confinement Phase Transition**

(Huang, Reichert, Sannino and Z-WW, PRD 104 (2021) 035005

- Confinement phase transition occurs for pure gluon and adjoint fermions
- $[\ell] = 0$  dimensionless while  $[\rho] = 1$ , we rewrite  $\rho$  as  $\rho = \ell T$  and convert the radius into a dimensionless quantity r' = r T:

$$S_3(T) = 4\pi T \int_0^\infty \!\!\mathrm{d}r'\,r'^2 \!\left[ \frac{1}{2} \left( \frac{\mathrm{d}\ell}{\mathrm{d}r'} \right)^2 + V'_{\mathsf{eff}}(\ell,T) \right] \,, \label{eq:S3}$$

which has the same form as the above generic equation.

• The bubble profile (instanton solution) is obtained by solving the E.O.M. of the  $S_3(T)$ 

$$\frac{\mathrm{d}^2 \ell(r')}{\mathrm{d}r'^2} + \frac{2}{r'} \frac{\mathrm{d}\ell(r')}{\mathrm{d}r'} - \frac{\partial V'_{\mathsf{eff}}(\ell, T)}{\partial \ell} = 0$$

The boundary conditions (deconfinement → confinement) are

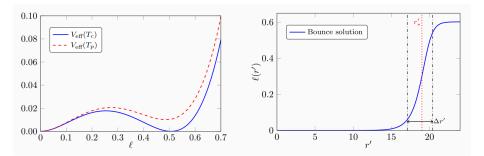
$$\frac{\mathrm{d}\ell(r'=0,T)}{\mathrm{d}r'} = 0, \qquad \lim_{r'\to 0} \ell(r',T) = 0$$

• We used the method of overshooting/undershooting (Python package)

王志伟 (电子科技大学) 强耦合暗物质的相变与引力波 4月23日, 2023 23/42

#### **Bubble Profile of Confinement Phase Transition**

(Huang, Reichert, Sannino and Z-W W, PRD 104 (2021) 035005



#### **Bubble Nucleation: Chiral Phase Transition**

(Reichert, Sannino, Z-W W and Zhang, JHEP **01** (2022) 003, arXiv:2109.11552)

- Chiral phase transition occurs when including fermions
- $\bullet$   $\bar{\sigma}$  is classically nonpropagating in PNJL and it's kinetic term is induced only via quantum fluctuations
- ullet We thus include its wave-function renormalization  $Z_{\sigma}$  with

$$Z_{\sigma}^{-1} = -\frac{\mathrm{d}\Gamma_{\sigma\sigma}(q^0,\mathbf{q},\bar{\sigma})}{\mathrm{d}\mathbf{q}^2}\bigg|_{q^0=0,\mathbf{q}^2=0}, \qquad \Gamma_{\sigma\sigma} = -i\sum 2\,\mathrm{point}\,1\mathrm{PI}\,\sigma\sigma\,\mathrm{graph}$$

The three-dimensional Euclidean action and E.O.M. are modified to:

$$\begin{split} S_3(T) &= 4\pi \int_0^\infty \!\! \mathrm{d} r \, r^2 \! \left[ \frac{Z_\sigma^{-1}}{2} \! \left( \frac{\mathrm{d} \bar{\sigma}}{\mathrm{d} r} \right)^2 + V_{\mathrm{eff}}(\bar{\sigma}, T) \right] \\ \frac{\mathrm{d}^2 \bar{\sigma}}{\mathrm{d} r^2} &+ \frac{2}{r} \frac{\mathrm{d} \bar{\sigma}}{\mathrm{d} r} - \frac{1}{2} \frac{\partial \log Z_\sigma}{\partial \bar{\sigma}} \left( \frac{\mathrm{d} \bar{\sigma}}{\mathrm{d} r} \right)^2 = Z_\sigma \frac{\partial V_{\mathrm{eff}}}{\partial \bar{\sigma}} \end{split}$$

• The associated boundary conditions:

$$\frac{\mathrm{d}\bar{\sigma}(r=0,T)}{\mathrm{d}r} = 0, \qquad \lim_{r \to \infty} \bar{\sigma}(r,T) = 0$$

#### Gravitational Wave Parameters: Inverse Duration Time

- The phase-transition temperature  $T_*$  is often identified with the nucleation temperature  $T_n$  defined as the temperature where the rate of bubble nucleation per Hubble volume and time is order one:  $\Gamma/H^4 \sim \mathcal{O}(1)$
- More accurately, we can use percolation temperature  $T_p$ : the temperature at which 34% of false vacuum is converted
- For sufficiently fast phase transitions, the decay rate is approximated by:

$$\Gamma(T) \approx \Gamma(t_*) e^{\beta(t-t_*)}$$

The inverse duration time then follows as

$$\beta = -\frac{\mathrm{d}}{\mathrm{d}t} \frac{S_3(T)}{T} \bigg|_{t=t_*}$$

• The dimensionless version  $\tilde{\beta}$  is defined relative to the Hubble parameter  $H_*$  at the characteristic time  $t_*$ 

$$\tilde{\beta} = \frac{\beta}{H_*} = T \frac{\mathrm{d}}{\mathrm{d}T} \frac{S_3(T)}{T} \bigg|_{T=T_*},$$

where we used that dT/dt = -H(T)T.

## Gravitational Wave Parameters: Strength Parameter I

(Huang, Reichert, Sannino and Z-W W, PRD 104 (2021) 035005

Reichert, Sannino, Z-W W and Zhang, JHEP 01 (2022)003, arXiv:2109.11552.)

• We define the strength parameter  $\alpha$  from the trace of the energy-momentum tensor  $\theta$  weighted by the enthalpy

$$\alpha = \frac{1}{3} \frac{\Delta \theta}{w_{+}} = \frac{1}{3} \frac{\Delta e - 3\Delta p}{w_{+}}, \qquad \Delta X = X^{(+)} - X^{(-)}, \text{ for } X = (\theta, e, p)$$

(+) denotes the meta-stable phase (outside of the bubble) while (-) denotes the stable phase (inside of the bubble).

• The relations between enthalpy w, pressure p, and energy e are given by

$$w = \frac{\partial p}{\partial \ln T} \,, \qquad \qquad e = \frac{\partial p}{\partial \ln T} - p \,,$$

which are extracted from the effective potential with

$$p^{(\pm)} = -V_{\rm eff}^{(\pm)}$$



## Gravitational Wave Parameters: Strength Parameter II

(Huang, Reichert, Sannino and Z-W W, PRD 104 (2021) 035005

Reichert, Sannino, Z-WW and Zhang, JHEP 01 (2022) 003, arXiv:2109.11552.)

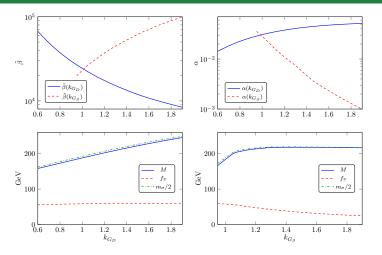
ullet  $\alpha$  is thus given by

$$\alpha = \frac{1}{3} \frac{4\Delta V_{\rm eff} - T \frac{\partial \Delta V_{\rm eff}}{\partial T}}{-T \frac{\partial V_{\rm eff}^{(+)}}{\partial T}} \,, \label{eq:alpha}$$

- For confinement phase transition:  $\alpha \approx 1/3$  ( $\Delta V_{\rm eff}$  is negligible since  $e_+ \gg p_+$  and  $e_- \sim p_- \sim 0$  in PLM potential )
- For chiral phase transition: we find smaller values,  $\alpha \sim \mathcal{O}(10^{-2})$ , due to the fact that more relativistic d.o.f.s participate in the phase transition
- ullet Relativistic SM d.o.f.s do not contribute to our definition of lpha since they are fully decoupled from the phase transition but these d.o.f.s will play a role to dilute the GW signals

## GW parameters $\alpha$ , $\beta$ and PNJL observables

(Reichert, Sannino, Z-W W and Zhang, JHEP 01 (2022) 003, arXiv:2109.11552.)



 $\blacksquare$ : The GW parameters  $\tilde{\beta}$ ,  $\alpha$  with the observables M,  $f_{\pi}$ , and  $m_{\sigma}$  as a function of  $G_S = k_{G_S} \cdot 4.6 \, \text{GeV}^{-2}$  and  $G_D = k_{G_D} \cdot (-743 \, \text{GeV}^{-5})$ . We use  $T_c = 100 \, \text{GeV}$ , the ratio  $\Lambda/T_0 = 3.54$ . Below  $k_{G_S,\text{crit}} = 0.882$ , no chiral symmetry breaking occurs.

## Gravitational-wave spectrum

(Huang, Reichert, Sannino and Z-W W, PRD  $\mathbf{104}$  (2021) 035005)

- Contributions from bubble collision and turbulence are subleading
- The GW spectrum from sound waves is given by

$$h^2\Omega_{\rm GW}(f) = h^2\Omega_{\rm GW}^{\rm peak} \left(\frac{f}{f_{\rm peak}}\right)^3 \left[\frac{4}{7} + \frac{3}{7} \left(\frac{f}{f_{\rm peak}}\right)^2\right]^{-\frac{\iota}{2}}$$

The peak frequency

$$f_{\rm peak} \simeq 1.9 \cdot 10^{-5} \, {\rm Hz} \left(\frac{g_*}{100}\right)^{\frac{1}{6}} \left(\frac{T}{100 \, {\rm GeV}}\right) \left(\frac{\tilde{\beta}}{v_w}\right)$$

The peak amplitude

$$h^2 \Omega_{\rm GW}^{\rm peak} \simeq 2.65 \cdot 10^{-6} \left(\frac{v_w}{\tilde{\beta}}\right) \left(\frac{\kappa_{sw} \, \alpha}{1+\alpha}\right)^2 \left(\frac{100}{g_*}\right)^{\frac{1}{3}} \Omega_{\rm dark}^2$$

• The factor  $\Omega^2_{\rm dark}$  accounts for the dilution of the GWs by the non-participating SM d.o.f.

$$\Omega_{\mathsf{dark}} = rac{
ho_{\mathsf{rad},\mathsf{dark}}}{
ho_{\mathsf{rad},\mathsf{tot}}} = rac{g_{*,\mathsf{dark}}}{g_{*,\mathsf{dark}} + g_{*,\mathsf{SM}}}$$

## The Efficiency Factor $\kappa$

• The efficiency factor for the sound waves  $\kappa_{\rm sw}$  consist of  $\kappa_v$  as well as an additional suppression due to the length of the sound-wave period  $\tau_{\rm sw}$ 

$$\kappa_{\rm SW} = \sqrt{\tau_{\rm SW}} \, \kappa_v$$

•  $\tau_{\text{SW}}$  is dimensionless and measured in units of the Hubble time (H.-K. Guo, Sinha, Vagie and White, JCAP **01** (2021) 001)

$$\tau_{\rm sw} = 1 - 1/\sqrt{1 + 2\frac{(8\pi)^{\frac{1}{3}}v_w}{\tilde{\beta}\,\bar{U}_f}} \, \Rightarrow \tau_{\rm sw} \sim \frac{(8\pi)^{\frac{1}{3}}v_w}{\tilde{\beta}\,\bar{U}_f} \ \, {\rm for}\, \beta >> 1$$

where  $ar{U}_f$  is the root-mean-square fluid velocity

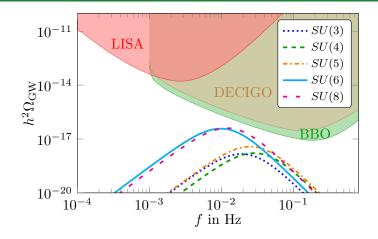
$$\bar{U}_f^2 = \frac{3}{v_w(1+\alpha)} \int_{c_s}^{v_w} d\xi \, \xi^2 \frac{v(\xi)^2}{1-v(\xi)^2} \simeq \frac{3}{4} \frac{\alpha}{1+\alpha} \kappa_v$$

- $\tau_{\rm sw}$  is suppressed for large  $\beta$  occurring often in strongly coupled sectors
- ullet  $\kappa_v$  was numerically fitted to simulation results depends  $\alpha$  and  $v_w$ . At the Chapman-Jouguet detonation velocity it reads

$$\kappa_v(v_w = v_J) = \frac{\sqrt{\alpha}}{0.135 + \sqrt{0.98 + \alpha}}$$

## GW Signatures for Arbitrary N in the Pure Gluon Case

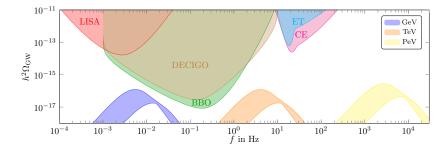
(Huang, Reichert, Sannino and Z-WW, PRD 104 (2021) 035005)



 $\blacksquare$ : The dependence of the GW spectrum on the number of dark colours is shown for the values N=3,4,5,6,8. All spectra are plotted with the bubble wall velocity set to the Chapman-Jouguet detonation velocity and with  $\mathsf{Tc}=1~\mathrm{GeV}$ .

## A Landscape of GW Signatures with Pure Gluon

(Huang, Reichert, Sannino and Z-W W, PRD 104 (2021) 035005)



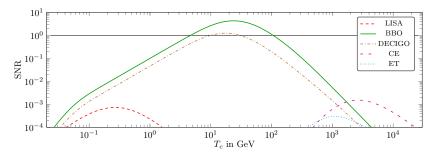
 $\blacksquare$ : We display the GW spectrum of the SU(6) phase transition for different confinement scales including  $T_c=1$  GeV, 1 TeV, and 1 PeV. We compare it to the power-law integrated sensitivity curves of LISA, BBO, DECIGO, CE, and ET.

## Signal to Noise Ratio

(Huang, Reichert, Sannino and Z-WW, PRD 104 (2021) 035005)

$${\rm SNR} = \sqrt{\frac{3 {\rm year}}{s}} \int_{f_{\rm min}}^{f_{\rm max}} {\rm d}f \left(\frac{h^2 \Omega_{\rm GW}}{h^2 \Omega_{\rm det}}\right)^2$$

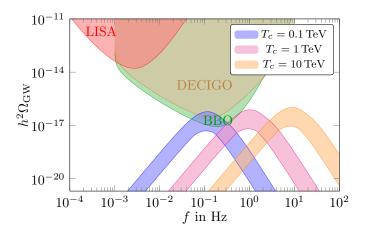
 $h^2\Omega_{\rm GW}$  is the GW spectrum while  $\Omega_{\rm det}$  is the sensitivity curve of the detector.



王志伟(电子科技大学) 强耦合暗物质的相变与引力波 4月23日,2023 34/4

## Landscape of GW spectrum with three Dirac fermions

(Reichert, Sannino, Z-W W and Zhang, JHEP 01 (2022) 003, arXiv:2109.11552.)



 $\blacksquare$ : Gravitational-wave spectrum with three Dirac fermions in the fundamental representation for different critical temperatures. The band comes from varying wall velocity  $c_s \le v_w \le 1$ .

## Representation Matters

(Reichert, Sannino, Z-W W and Zhang, JHEP 01 (2022) 003, arXiv:2109.11552.)

Rep.	flavour	chiral PT	confdeconf.
Fund.	3	1st	X
adjoint	1	2nd	1st
2-index Sym.	1	2nd	1st

表: Representations versus different phase transitions.

• Need small  $N_f$  to remain below the conformal Banks-Zaks window  $(N_f \le 2$  for adjoint and  $N_f \le 3$  for 2-index symmetric under SU(3)).

## Signal to Noise Ratio for Different Representations

(Reichert, Sannino, Z-W W and Zhang, JHEP 01 (2022) 003, arXiv:2109.11552.)

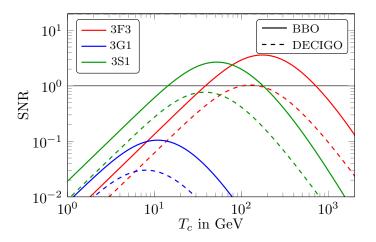


图: Signal-to-noise ratio as a function of the critical temperature for the best-case scenarios of each model at BBO and DECIGO with an observation time of 3 years.

## **Future Thinking**

• The three-dimensional Euclidean action  $S_3$  can be written as a function of the latent heat L and the surface tension  $\sigma$ 

$$S_3 = \frac{16\pi}{3} \frac{\sigma(T_c)^3}{L(T_c)^2} \frac{T_c^2}{(T_c - T)^2} \,,$$

• The ratio  $S_3(T_p)/T_p$  is typically a number  $\mathcal{O}(150)$  for phase transitions around the electroweak scale and the inverse duration  $\tilde{\beta}$  follows as

$$\tilde{\beta} = T \frac{\mathrm{d}}{\mathrm{d}T} \frac{S_3(T)}{T} \bigg|_{T=T_p} \approx \mathcal{O}(10^3) \frac{T_c^{1/2} L}{\sigma^{3/2}}.$$

- $\tilde{\beta}$  stems from the competition between the surface tension and latent heat.  $L \sim N^2$  while  $\sigma$  can be either  $\sigma \sim N$  or  $\sigma \sim N^2$  with limited data up to SU(8)
- How to construct models with smaller latent heat and larger surface tension to enhance the gravitational wave signals?

## Thank you for your attention!

## **About Center Symmetry and Confinement**

 The standard physical interpretation is that it is related to the free energy of adding an external static color source in the fundamental representation.

$$\ell\left(\vec{x}\right) = \exp\left(-F\beta\right)$$

 In the confinement phase, Polyakov loop is zero corresponds to infinity free energy to add a color source and the same time center symmetry is unbroken.

## Center Symmetry Z(N) at Nonzero Temperature

 $\bullet$  The boundary conditions in imaginary time  $\tau$  the fields must satisfy are:

$$A_{\mu}\left(\vec{x},\beta\right) = +A_{\mu}\left(\vec{x},0\right), \qquad q\left(\vec{x},\beta\right) = -q\left(\vec{x},0\right),$$

where gluons as bosons must be periodic in  $\tau$  while quarks as fermions must be anti-periodic.

• 't Hooft first noticed that one can consider more general gauge transformations which are only periodic up to  $\Omega_c$ 

$$\Omega(\vec{x}, \beta) = \Omega_c, \quad \Omega(\vec{x}, 0) = 1 \qquad \left(\text{here, } \Omega_c = e^{i\phi}I, \ \phi = \frac{2\pi j}{N}\right).$$

 Color adjoint fields are invariant under this transformation, while those in the fundamental representation are not:

$$\begin{split} A^{\Omega}\left(\vec{x},\beta\right) &= \Omega_{c}^{\dagger}A_{\mu}\left(\vec{x},\beta\right)\Omega_{c} = A_{\mu}\left(\vec{x},\beta\right) = +A_{\mu}\left(\vec{x},0\right)\,,\\ q^{\Omega}\left(\vec{x},\beta\right) &= \Omega_{c}^{\dagger}q\left(\vec{x},\beta\right) = e^{-i\phi}q\left(\vec{x},\beta\right) \neq -q\left(\vec{x},0\right)\,. \end{split}$$

 $\bullet$  Thermal Wilson line transforms like an adjoint field under local SU(N) gauge transformations:

$$L(x) o \Omega^\dagger \left( ec{x}, eta 
ight) L(ec{x}) \Omega^\dagger \left( ec{x}, 0 
ight)$$
 .

## About Thin Wall Approximation

The three-dimensional Euclidean action  $S_3$  can be written as a function of the latent heat L and the surface tension  $\sigma$ 

$$S_3 = \frac{16\pi}{3} \frac{\sigma(T_c)^3}{L(T_c)^2} \frac{T_c^2}{(T_c - T)^2} \,,$$

The ratio  $S_3(T_p)/T_p$  is typically a number  $\mathcal{O}(150)$  for phase transitions around the electroweak scale. From this we infer that

$$T_c - T_p \approx \sqrt{\frac{16\pi\sigma^3 T_c}{3L^2 \cdot \mathcal{O}(150)}}$$
,

and the inverse duration  $\tilde{\beta}$  follows as

$$\tilde{\beta} = T \frac{\mathrm{d}}{\mathrm{d}T} \frac{S_3(T)}{T} \bigg|_{T=T_p} \approx \mathcal{O}(10^3) \frac{T_c^{1/2} L}{\sigma^{3/2}}.$$

 $\tilde{\beta}$  stems from the competition between the surface tension and latent heat.

42/42