

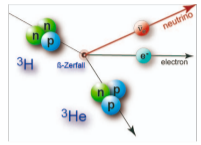
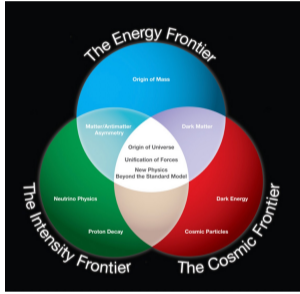
Advances in
modeling neutrinoless double-beta
decay in atomic nuclei with operators
from chiral effective field theory

Jiangming Yao (尧江明)

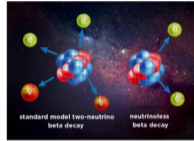
School of Physics and Astronomy, Sun Yat-sen University

第八届手征有效场论研讨会，河南开封，2023年10月28日

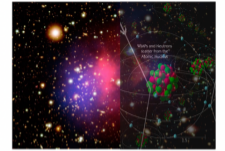
- 1 Neutrinoless double-beta decay as a probe into new physics
- 2 Status of modern studies on the nuclear matrix elements of $0\nu\beta\beta$ decay
- 3 Recent studies with operators from (chiral) effective field theory
- 4 Summary and perspectives



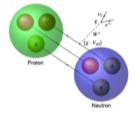
Single-beta decay



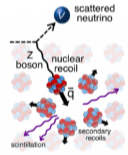
Neutrinoless double beta decay



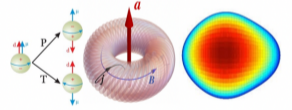
Dark matter direct detection



Superallowed Fermi transitions



Neutrino scattering



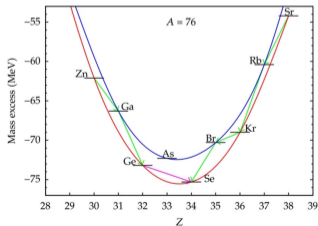
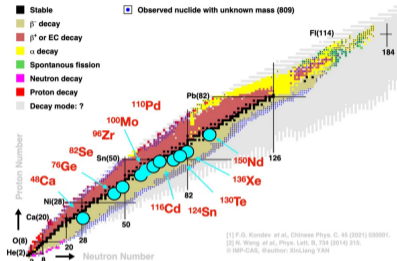
Symmetry-violating moments

- Three frontiers: HE, LS, HP
- Atomic nuclei: low-energy probes

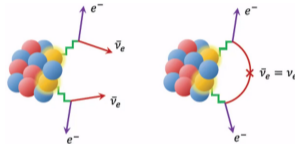
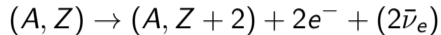
- $\beta(\beta)$ decay, WIMP-nucleus scattering, etc.
- All about Nuclear Matrix Elements (NME)

A special decay mode: $0\nu\beta\beta$ decay

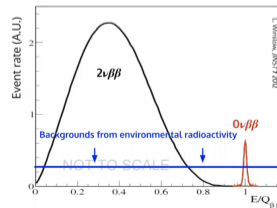
Nuclear Chart: decay mode of the ground state nuclide(NUBASE2020)



- The two modes of $\beta^-\beta^-$ decay:



- Kinetic energy spectrum of electrons



Neutrino oscillation

- From mass to flavor states

$$|\nu_\alpha\rangle = \sum_{j=1}^{N=3} U_{\alpha j}^* |\nu_j\rangle.$$

- $\Delta m_{ij}^2 (\neq 0)$, and $\theta_{ij} (\neq 0)$.

Open questions

- The nature of neutrinos.
- Neutrino mass m_j and its origin.

The observation of $0\nu\beta\beta$ decay would provide answers.

If $0\nu\beta\beta$ decay is driven by exchanging light massive Majorana neutrinos:

$$\langle m_{\beta\beta} \rangle \equiv \left| \sum_{j=1}^3 U_{ej}^2 m_j \right| = \left[\frac{m_e^2}{g_A^4 G_{0\nu} T_{1/2}^{0\nu} |M^{0\nu}|^2} \right]^{1/2}$$

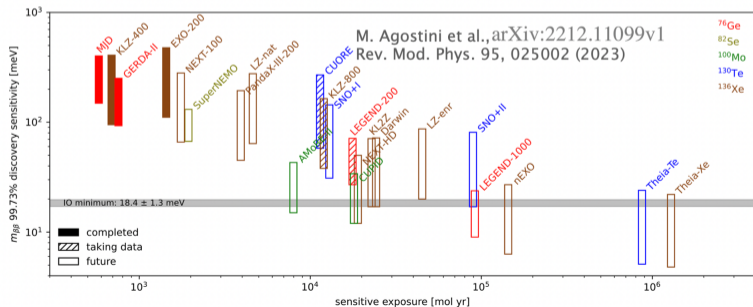
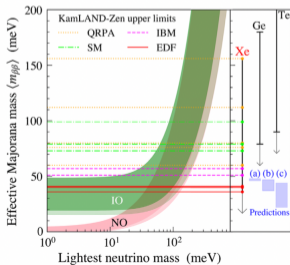
- U_{ej} : elements of the PMNS matrix
- $G_{0\nu}$: phase-space factor
- $M^{0\nu}$: the nuclear matrix element

$$M^{0\nu} = \langle \Psi_F | \hat{O}^{0\nu} | \Psi_I \rangle$$

- Transition operator: $\hat{O}^{0\nu}$
- Nuclear many-body wfs: $|\Psi_{I/F}\rangle$

Constraints on neutrino mass from $0\nu\beta\beta$ decay

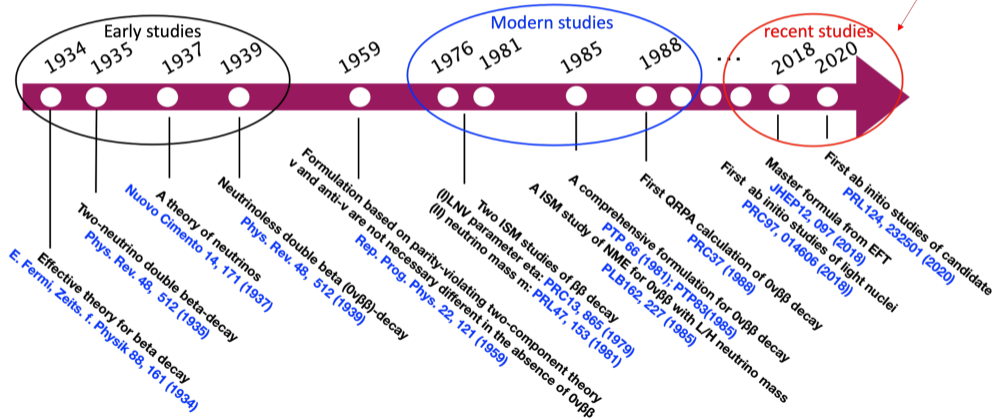
Isotope	$G_{0\nu}$ [10^{-14} yr^{-1}]	$M^{0\nu}$ [min, max]	$T_{1/2}^{0\nu}$ [yr]	$\langle m_{\beta\beta} \rangle$ [meV]	Experiments References
^{48}Ca	2.48	[0.85, 2.94]	$> 5.8 \cdot 10^{22}$	[2841, 9828]	CANDLES: PRC78, 058501 (2008)
^{76}Ge	0.24	[2.38, 6.64]	$> 1.8 \cdot 10^{26}$	[73, 180]	GERDA: PRL125, 252502(2020)
^{82}Se	1.01	[2.72, 5.30]	$> 4.6 \cdot 10^{24}$	[277, 540]	CUPID-0: PRL129, 111801 (2023)
^{96}Zr	2.06	[2.86, 6.47]	$> 9.2 \cdot 10^{21}$	[3557, 8047]	NPA847, 168 (2010)
^{100}Mo	1.59	[3.84, 6.59]	$> 1.5 \cdot 10^{24}$	[310, 540]	CUPID-Mo: PRL126, 181802(2021)
^{116}Cd	0.48	[3.29, 5.52]	$> 2.2 \cdot 10^{23}$	[1766, 2963]	PRD 98, 092007 (2018)
^{130}Te	1.42	[1.37, 6.41]	$> 2.2 \cdot 10^{25}$	[90, 305]	CUORE: Nature 604, 53(2022)
^{136}Xe	1.46	[1.11, 4.77]	$> 2.3 \cdot 10^{26}$	[36, 156]	KamLAND-Zen: PRL130, 051801(2023)
^{150}Nd	6.30	[1.71, 5.60]	$> 2.0 \cdot 10^{22}$	[1593, 5219]	NEMO-3: PRD 94, 072003 (2016)

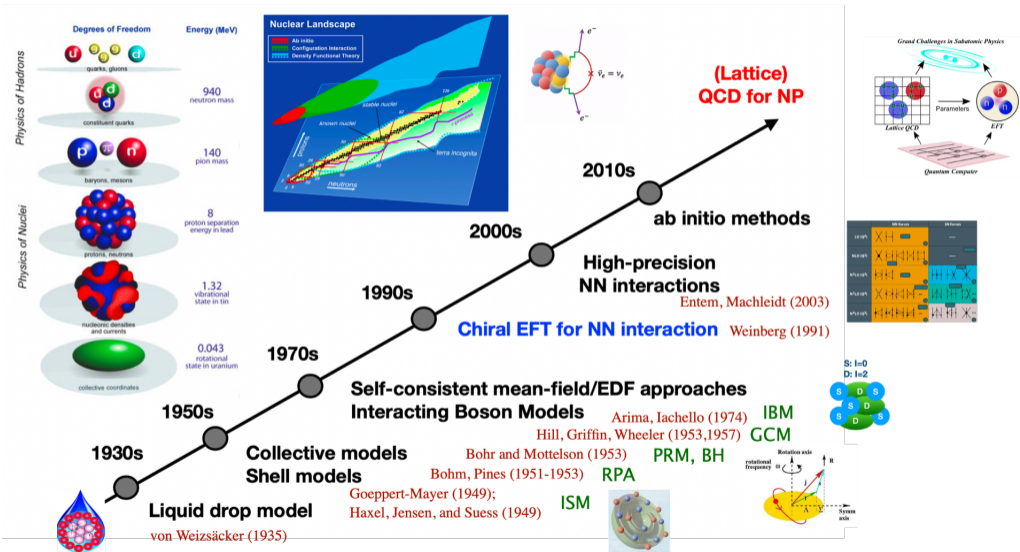


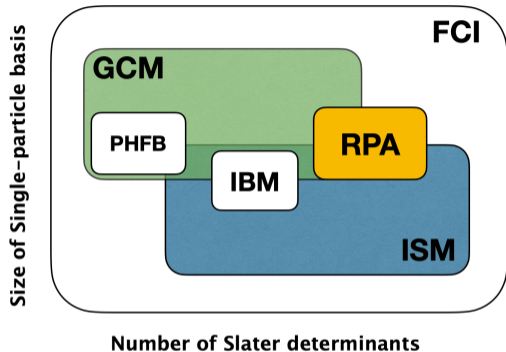
Brief history on modeling the $\beta(\beta\beta)$ decay rate



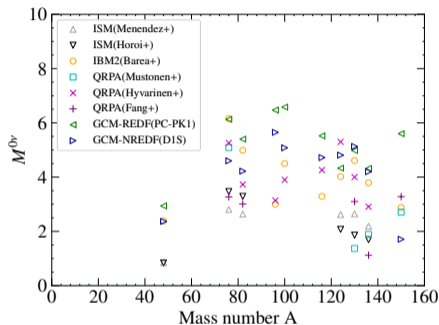
Fermi Goepfert-Mayer Majorana Furry Primakoff Vergados Haxton 吴慧芳 Engel







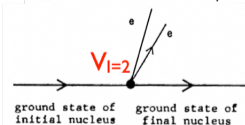
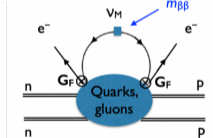
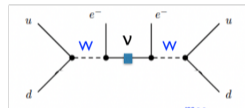
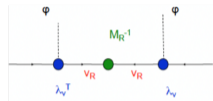
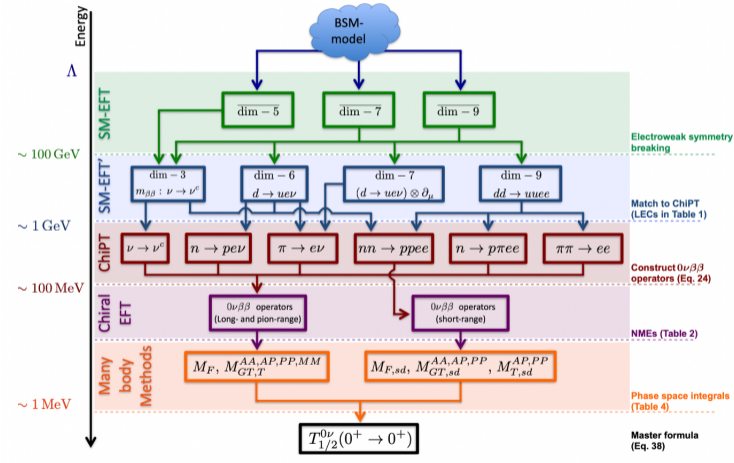
JMY, J. Meng, Y.F. Niu, P. Ring, PPNP 126, 103965 (2022)



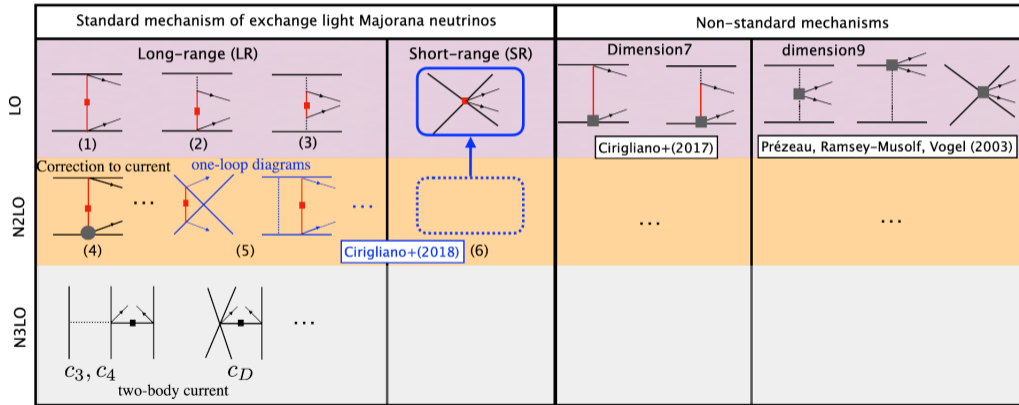
- ISM predicts small NMEs, while IBM and EDF predict large NMEs. Discrepancy is about a factor of THREE or even larger.
- Different models are not equivalent! Different schemes (model spaces and interactions): compare apples to oranges?
- Efforts in resolving the discrepancy: Challenging or even impossible?

EFT: a model-independent analysis of operators at different energy scales






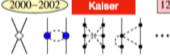
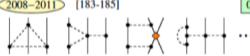
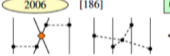
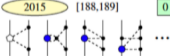


Cirigliano (2018)



- At $E \sim 100$ MeV: operators are expressed in terms of nucleons, pions, and leptons.



- Non-relativistic chiral 2N+3N interactions (Weinberg power counting and others)

	NN	3N	4N
LO $\mathcal{O}(Q^0/\Lambda^0)$	 1990 Weinberg 2	—	—
NLO $\mathcal{O}(Q^2/\Lambda^2)$	 1992 Ordonez, van Kolck 7	 1992, 1994 [166-169]	—
N ² LO $\mathcal{O}(Q^3/\Lambda^3)$	 1992 Ordonez, van Kolck 0	 1994 Weinberg, van Kolck, Epelbaum 2	—
N ³ LO $\mathcal{O}(Q^4/\Lambda^4)$	 2000-2002 Kaiser 12	 2008-2011 [183-185] 0	 2006 [186] 0
N ⁴ LO $\mathcal{O}(Q^5/\Lambda^5)$	 2015 [188, 189] 0	 2011- [190-192] ?	 ?

K. Hebeler, Phys. Rep. 890, 1 (2020)

- Relativistic chiral 2N interaction (up to N²LO, different PC from the NR case)

J.-X. Lu et al., PRL128, 142002 (2022)

Ab initio methods for the lightest candidate ^{48}Ca

- Multi-reference in-medium generator coordinate method (IM-GCM)

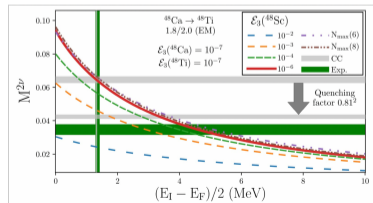
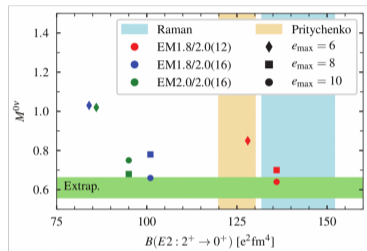
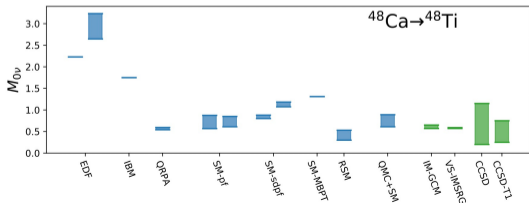
JMY et al., PRL124, 232501 (2020)

- IMSRG+ISM (VS-IMSRG)

A. Belley et al., PRL126, 042502 (2021)

- Coupled-cluster with singlets, doublets, and partial triplets (CCSDT1)

S. Novario et al., PRL126, 182502 (2021)



Featured in Physics

Editors' Suggestion


Open Access

New Leading Contribution to Neutrinoless Double- β Decay


Vincenzo Cirigliano, Wouter Dekens, Jordy de Vries, Michael L. Graesser, Emanuele Mereghetti, Saori Pastore, and Udirajara van Kolck
Phys. Rev. Lett. **120**, 202001 – Published 16 May 2018

PhysiCS See Synopsis: [A Missing Piece in the Neutrinoless Beta-Dec](#)

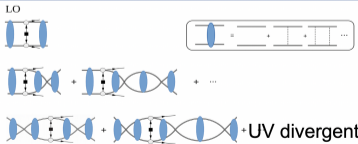
Nuclear force



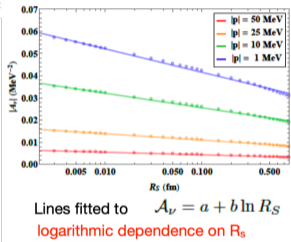
Transition operator



LO



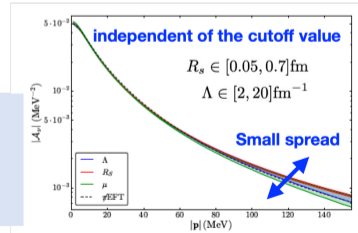
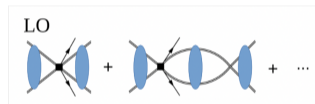
+UV divergent



- The transition amplitude is regulator-dependent!
- Needs a counter term at LO in order to ensure renormalizability.

Introducing a contact transition operator

$$V_{\nu,S} = -2g_\nu^{NN} \tau(1) + \tau(2) + \dots$$



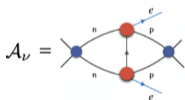
Toward Complete Leading-Order Predictions for Neutrinoless Double β Decay

Vincenzo Cirigliano, Wouter Dekens, Jordy de Vries, Martin Hoferichter, and Emar Mereghetti

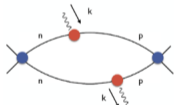
Phys. Rev. Lett. **126**, 172002 (2021) – Published 30 April 2021

- **Cottingham formula** W.N. Cottingham, Ann. Phys. 25, 424 (1963)

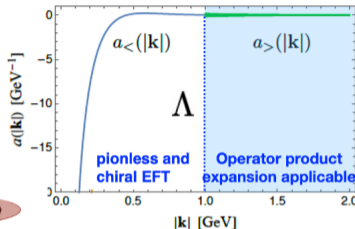
$$\mathcal{A}_\nu \propto \int \frac{d^4k}{(2\pi)^4} \frac{g_{\mu\nu}}{k^2 + i\epsilon} \int d^4x e^{ik \cdot x} \langle pp | T \{ j_w^\mu(x) j_w^\nu(0) \} | nn \rangle$$



$$\propto \int \frac{d^4k}{(2\pi)^4} \frac{1}{k^2 + i\epsilon}$$



forward Compton amplitude



- **Synthetic datum**

$$\mathcal{A}_\nu(|\mathbf{p}|, |\mathbf{p}'|) \times e^{-i(\delta_1 s_0(|\mathbf{p}|) + \delta_1 s_0(|\mathbf{p}'|))} = - \left(2.271 - 0.075 \tilde{C}_1(4 M_\pi) \right) \times 10^{-2} \text{ MeV}^{-2}$$

$$\boxed{|\mathbf{p}| = 25 \text{ MeV} \quad (|\mathbf{p}'| = 30 \text{ MeV})} \quad = -1.95(5) \tilde{c}_1 \times 10^{-2} \text{ MeV}^{-2},$$

Uncertainty from the estimate of the **inelastic** contributions

The transition amplitude is observable and thus scheme independent.

A recent study in the relativistic chiral EFT shows that

- the $nn \rightarrow ppe^-e^-$ transition amplitude \mathcal{A}_ν is regulator-independent, thus no need to introduce the contact transition operator.
- The predicted $\mathcal{A}_\nu = 0.02085\text{MeV}^{-2}$, about 10% larger than the value by Cirigliano (2021).
- The discrepancy could be attributed to the different power counting: the LO of relativistic chiral EFT contains partial N2LO contribution of non-relativistic EFT.

Y.L. Yang and P. W. Zhao, arXiv:2308.03356v1 (2023)

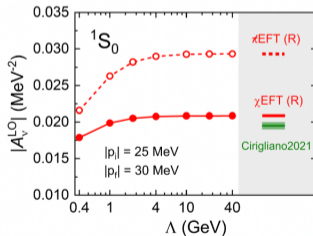
Bethe-Salpeter equation

$$T(\vec{p}', \vec{p}) = V(\vec{p}', \vec{p}) + \int \frac{d^3 p''}{(2\pi)^3} V(\vec{p}', \vec{p}'') \frac{M_N^2}{E_{p''}} \frac{1}{p^2 - p''^2 + i\epsilon} T(\vec{p}'', \vec{p})$$

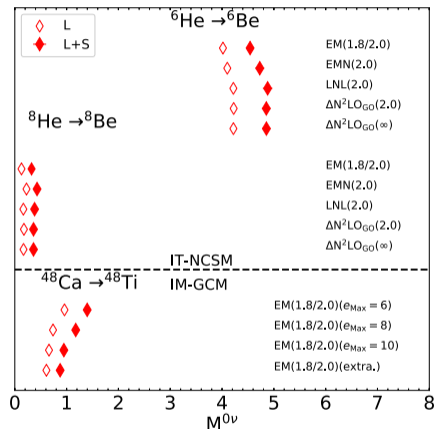
$$E_{p''} \equiv \sqrt{M_N^2 + p''^2}$$

Lippmann-Schwinger equation

$$\widehat{T}(\vec{p}', \vec{p}) = \widehat{V}(\vec{p}', \vec{p}) + \int d^3 p'' \widehat{V}(\vec{p}', \vec{p}'') \frac{M_N}{p^2 - p''^2 + i\epsilon} \widehat{T}(\vec{p}'', \vec{p})$$



- The LEC g_{ν}^{NN} consistent with the employed chiral interaction (EM1.8/2.0) is determined based on the synthetic data.
- The contact term turns out to enhance (instead of quench) the NME for ^{48}Ca by 43(7)%, thus the half-life $T_{1/2}^{0\nu\beta\beta}$ is only half of the previously expected value.
- The uncertainty (7%) is due to the synthetic data which can be reduced by using an accurate value of the LEC (g_{ν}^{NN}).



R. Wirth, JMY, H. Hergert, PRL127, 242502 (2021)

With both the long- and short-range transition operators, the VS-IMSRG method is applied to study the NMEs of heavier candidates:

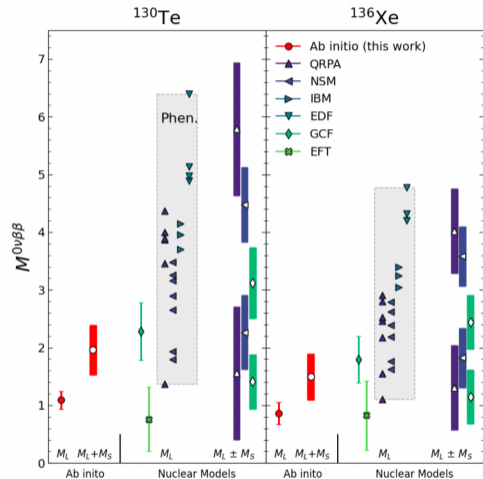
- For ^{130}Te , $M_{L+S}^{0\nu} \in [1.52, 2.40]$
- For ^{136}Xe , $M_{L+S}^{0\nu} \in [1.08, 1.90]$

Sources of uncertainty: Nuclear interaction, reference-state, basis extrapolation, closure approximation, and the LEC for the short-range transition operators.

A more comprehensive uncertainty analysis in order

including the errors from

- many-body truncations
- chiral expansion orders



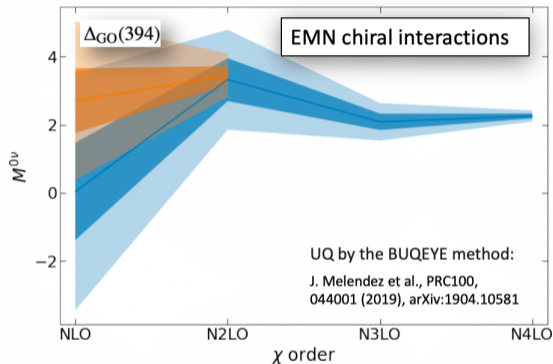
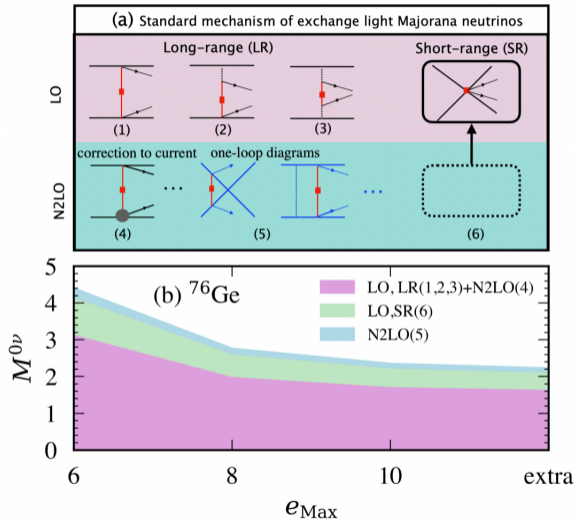
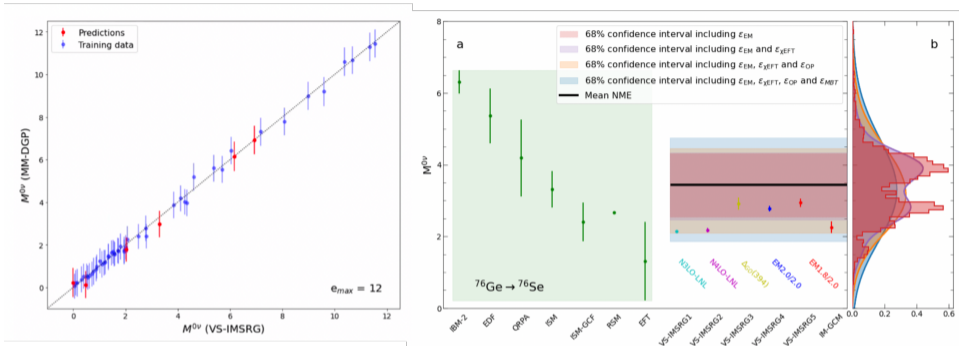


Table 1 | The recommended value for the total NME of $0\nu\beta\beta$ decay in ^{76}Ge , together with the uncertainties from different sources.

$M^{0\nu}$	ϵ_{LEC}	$\epsilon_{\chi\text{EFT}}$	ϵ_{MBT}	ϵ_{OP}	ϵ_{EM}
$3.44^{+1.33}_{-1.56}$	0.9	0.3	0.8	0.5	<0.06



- Emulator, 8188 samples of chiral interactions, phase shift, $M^{0\nu} = 3.44^{+1.33}_{-1.56}$.
- Current upper limit for the effective neutrino mass $\langle m_{\beta\beta} \rangle = 141^{+117}_{-39}$ meV.
- The next-generation ton-scale Germanium experiment ($\sim 1.3 \times 10^{28}$ yr):
 $m_{\beta\beta} = 17^{+14}_{-5}$ meV, covering almost the entire range of IO hierarchy.

- $0\nu\beta\beta$ decay: lepton-number-violation process, a complementary way to determine the absolute mass scale of neutrinos.
- Next-generation experiments: tonne-scale detectors with a half-life sensitivity up to 10^{28} years.
- Large uncertainty in NMEs: systematical uncertainty, impacting extracted neutrino mass, attracting a lot of efforts from nuclear community.
- Ab initio studies of NMEs: remarkable progress, disclosing non-trivial contributions from high-energy light neutrinos. The NMEs for heavier candidate nuclei (^{48}Ca , ^{76}Ge , ^{82}Se , ^{130}Te , ^{136}Xe) have been computed. Convergence w.r.t. the chiral expansion order is rather rapid.

Next

- Considering higher-order nuclear interactions, reducing many-body truncation errors, and finding more constraints to shrink the uncertainty.
- Contributions from other mechanisms.

Collaborators

- **SYSU**

Chenrong Ding, Changfeng Jiao, Gang Li,
Xin Zhang

- **PKU**

Lingshuang Song, Jie Meng, Peter Ring

- **LZU**: Yifei Niu

- **SCU**: Chunlin Bai

- **IMP, CAS**: Dongliang Fang

- **SWU**: Longjun Wang

- **MSU**: Scott Bogner, Heiko Hergert, Roland Wirth

- **UNC**: Jonathan Engel, A. M. Romero

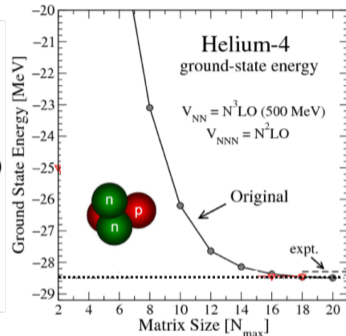
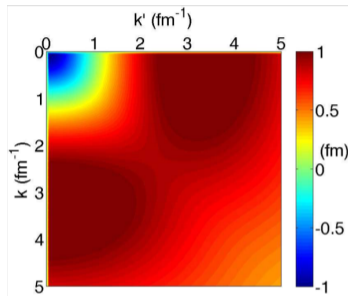
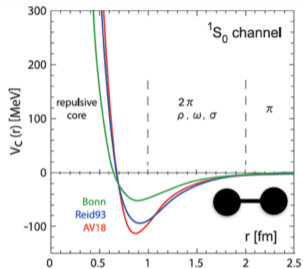
- **TRIUMF**: Antonie Belly, Jason Holt

- **TU Darmstadt**: Takayuki Miyagi

- **Notre-Dame U**: Ragnar Stroberg

- **UAM**: Benjamin Bally, Tomas Rodriguez

Thank you for your attention!



- Repulsive core & strong tensor force: **low and high k modes strongly coupled.**
- non-perturbative, poorly convergence in basis expansion methods.

S. Bogner et al., PPNP (2010)

- Apply unitary transformations to Hamiltonian

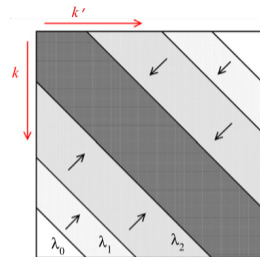
$$H_s = U_s H U_s^\dagger \equiv T_{\text{rel}} + V_s$$

from which one finds the flow equation

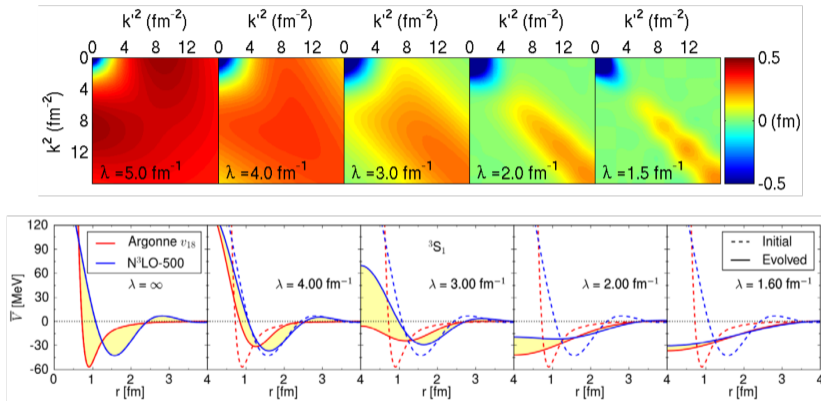
$$\frac{dH_s}{ds} = [\eta_s, H_s], \quad \eta_s = [T_{\text{rel}}, H_s]$$

Evolution of the potential

$$\frac{dV_s(k, k')}{ds} = -(k^2 - k'^2)V_s(k, k') + \frac{2}{\pi} \int_0^\infty q^2 dq (k^2 + k'^2 - 2q^2)V_s(k, q)V_s(q, k')$$



The flow parameter s is usually replaced with $\lambda = s^{-1/4}$ in units of fm^{-1} . *S. K. Bogner et al. (2007)*



Local projection of AV18 and N³LO(500 MeV) potentials $V(r)$.

- The hard core "disappears" in the softened interactions

