The results of Performance evaluation

Performance evaluation of 20cm MCP-PMTs

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- **2** The method of Performance evaluation
- **3** The results of Performance evaluation
- **4** Summary



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Jinping Neutrino Neutrino Observatory





- Jinping Mountain Tunnel in Xichang City, Sichuan Province (2400 m of rock);
- 2 500-ton liquid scintillator detector in construction;
- **3** ν deposits energy in LS and emits photons.
- How to observe these photons?

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Jinping Neutrino Neutrino Observatory





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How to observe these photons?

Photomultiplier Tube (PMT)



Photoelectric effect and electron multiplication

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The requirement of PMT



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The requirement of PMT





- 1 nearly 4,000 PMTs: to control the cost of PMTs
- Wigh Photon detection efficiency (PDE): up to 30% for energy measurement
- 3 High time resolution: nanosecond timing resolution, for position measurement
- @ Reduce background impacts: low dark noise (fake signals), high signal-to-noise ratio

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Keys: All parameters mentioned above, particularly PDE and Transit time spread (TTS).

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Classical PMT: multi-stage dynode

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Ø20 cm MCP-PMT



Classical PMT: multi-stage dynode



MCP: small, fast time response

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Ø20 cm MCP-PMT



Classical PMT: multi-stage dynode



MCP: small, fast time response



- Photons hit the photocathode and generate electrons through the photoelectric effect
- @ Electrons are accelerated into the MCP and amplified by a factor of 1×10^7 to form observable pulse signals

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Ø20 cm MCP-PMT



Classical PMT: multi-stage dynode



MCP: small, fast time response



- Photons hit the photocathode and generate electrons through the photoelectric effect
- @ Electrons are accelerated into the MCP and amplified by a factor of 1×10^7 to form observable pulse signals
- JNE uses the \emptyset 20 cm MCP-PMT from NNVT
- To improve PDE, the MCP has an **ALD coating** on upper surface

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Experimental Setup



- 1 Measurements are conducted in a dark box.
- 2 FADC: 10-bit, 1 GHz sampling rate.
- PiL040XSM picosecond laser: 405 nm, splited into four channels.
- Each pulse collects 10 us of data, with approximately 5 % containing photon information.
- Seach testing run lasts for 20 hours, with data acquisition, storage, and analysis managed by self-innovate software for fully automatical execution.

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Testing Procedure

- **1** Nine MCP-PMTs were tested, with a dynode PMT from Hamamatsu as a reference.
- ② The four PMTs were moved in a cycle within the dark box to eliminate the influence of different light intensities in different compartments.

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Data Processing



- laser signals at 240 ns
- detected light signals at 440 ns
- negative pulse

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Data Processing		
954	FWWW	 laser signals at 240 ns detected light signals at 440 ns negative pulse
Voltage 950 – PMT waveform Trigger waveforr	threshold N n(scaled) I I I I I I I I I I I I I I I I I I I	Photon Counting Capability
§ 948 − × t _{uig} − − × − × − − × − × − − × − × − − × − − × − × − − × − × − − × − × − × − − × −× −	t_{10} t_p μ_p $0.9V_p$ risetime falltime integration window	Charge distributionPhoton detection efficiency(PDE)
200 250 400	450 500 55	Time Resolution Capability

Transit time spread(TTS)

٠ Calculate the baseline and subtract it.

t/ns

- Charge: the result of integrating the pulse, which is proportional to the pulse height.
- TT: the difference between the 10% rise time of the pulse and the laser trigger time.

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Charge Distribution Relative Photon Detection Efficiency Energy Resolution Time Resolution Capability Afterpulse

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Charge Distribution



- Gaussian fit to the peak (light red)
- quadratic fit to the valley (light green).

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Charge Distribution



- Gaussian fit to the peak (light red)
- quadratic fit to the valley (light green).
- Peak-to-Valley ratio reaches 6, better than similar products (usually 3-4).
- a long tail on the right-hand side, approximately 40%.
- The relative standard deviation:

 $\nu = \sqrt{\text{Var[Charge]}}/\text{E[Charge]}$

the smaller ν , the better energy resolution

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- The long tail structure: significantly different from that of Classical dynode PMTs.
- Two amplification modes:
 - tail: large and broad
 - 2 main peak: small and narrow

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Two Amplification Modes





- Triple Gamma Fit

 $\alpha_n = 17.499 \pm 0.803$

 $\beta = 0.058 \pm 0.003$

 $\alpha_{i} = 4.696 \pm 0.124$ $\beta = 0.595 \pm 0.007$

 $\alpha_{\rm c} = 22.500 \pm 0.853$

 $\beta_{-} = 0.022 \pm 0.001$ $\gamma^2/ndf = 124.0 / 112$

••• peak: p = 0.519 ± 0.010

•••• tail: $p = 0.440 \pm 0.011$ ---- small gamma

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Two Amplification Modes





main peak (channel mode)

- small and narrow
- 2 electrons enter MCP channels directly
- **3** easy to count photons

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Two Amplification Modes





main peak (channel mode)

- small and narrow
- 2 electrons enter MCP channels directly
- **(3)** easy to count photons

tail (surface mode)

- large and broad
- 2 surface secondary emission electrons enter MCP
- **③** hard to count photons

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Two Amplification Modes



main peak (channel mode)

- small and narrow
- 2 electrons enter MCP channels directly
- easy to count photons

tail (surface mode)

- large and broad
- 2 surface secondary emission electrons enter MCP
- **8** hard to count photons
- "Tail structure" reduces the energy resolution and needs to be studied!

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Relative Photon Detection Efficiency (ϵ^0)





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Relative Photon Detection Efficiency (ϵ^0)



Characterizing the probability of photons converting into observable signals

- The PDE of reference PMT is : ϵ_0
- The PDE of tested PMT is : ϵ_k
- Testing each PMT on each channel
- The relative relationships: $\epsilon_k^0 = \frac{\epsilon_k}{\epsilon_0}$
- General linear model of Binomial exponential family distribution is used for ϵ^0 .

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Relative Photon Detection Efficiency (ϵ^0)



Characterizing the probability of photons converting into observable signals

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tested $\epsilon^0 = 1.7$, significantly improving photon counting and energy resolution

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Energy Resolution



• Energy resolution is determined by $\frac{\sqrt{1+\nu^2}}{\epsilon^0}$ which is the smaller, the better.

- The thicker the ALD coating, the bigger the tail, the larger the u, the higher ϵ^0
- There is a balance between PDE and energy resolution, and finally T₄ nm is chosen.

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Transit Time Spread(TTS)



- Gaussian fitting of the Transit Time (TT) to get TTS of 1.7 ns, a significant improvement compared to the previous generation Ø50 cm PMT (which is 10 ns).
- 2 There is a certain correlation between transit time and charge size to be guantified.

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Afterpulse



lonization of gas inside the MCP-PMT generates ions that can strike the photocathode and produce additional electrons, with a delay time on the order of 100 ns.

- The timing of afterpulses is related to the properties of the ions $\binom{Z}{M}X$ with a relationship of $\sqrt{\frac{M}{Z}}$
- Afterpulses are concentrated at 300 ns, 550 ns, 1200 ns, and 1700 ns, with a ratio of approximately $1:\sqrt{3}:\sqrt{16}:\sqrt{32}$
- Possible ion components include H⁺, He⁺, O⁺, or CH₄⁺, O₂⁺. $\langle \Box \rangle \langle \Box \rangle \langle$

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Summary

- The new Ø20 cm MCP-PMT from NNVT used in the Jinping Neutrino Experiment demonstrates superior performance in terms of PDE and Peak-to-Valley ratio, with other terms comparable to similar international products.
- The long tail in the charge distribution is countered by the high PDE, resulting in an overall boost in energy resolution
- The ongoing research on new waveform analysis methods holds promise for mitigating the impact of the tail component in the charge spectrum.

Paper Information

Performance evaluation of the 8-inch MCP-PMT for Jinping Neutrino Experiment. https://doi.org/10.1016/j.nima.2023.168506

- 1 Dark count rate
- 2 Afterpulse

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Gain and single PE resolution



- The gain of the main peak G_1 : $\frac{C_1}{e \times 50\Omega}$
- The gain of the total charge $G: G = \frac{\mu_C}{e \times 50\Omega}$
- The main peak resolution Res₁: $\frac{\sigma_{C_1}}{C_1}$
- The total charge resolution Res: $G_{e \times 50\Omega}^{\mu_C}$
- G is about 2 times G_1 for the MCP-PMTs
- Mean of Res_1 , Res : about 0.25, 0.69

Peak-to-valley ratio and some time characteristics

- **1** The mean P/V ratio of MCP-PMTs is about 5.8 while the reference PMT is about 2.4.
- @ Estimated mean and deviation of rise time, fall time, and FWHM are 3.71 ± 0.15 ns, 15.6 ± 1.8 ns, and 9.07 ± 0.63 ns for 9 MCP-PMTs.

 Energy resolution boost Relative PDE TTS



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Energy resolution boost



- the number of expected photons N on a PMT: $\pi(\mu_N)$
- Energy *E* of an event is proportional to $N = k\eta E$
- The output charge distribution *C* is a hierarchical model

$$E[C] = \mu_N \mu_C \tag{1}$$

(日)

$$Var[C] = \mu_C^2 \mu_N + \mu_N \sigma_C^2$$
 (2)

N is estimated as $\hat{N} = \frac{C}{\mu_c}$ and E is estimated as $\hat{E} = \frac{\hat{N}}{k}$. The reconstructed energy resolution

$$\frac{\sqrt{\operatorname{Var}[\hat{E}]}}{\operatorname{E}[\hat{E}]} = \frac{\sqrt{\mu_c^2 \mu_N + \mu_N \sigma_c^2}}{\mu_N \mu_c} = \frac{\sqrt{1 + (\frac{\sigma_c}{\mu_c})^2}}{\sqrt{\mu_N}} = \frac{\sqrt{1 + (\frac{\sigma_c}{\mu_c})^2}}{\sqrt{\operatorname{kn}E}}$$

Dark count rate

$$\mathrm{DCR/kHz} = \frac{N_{\mathrm{noise}}}{N_{\mathrm{trig}}} \frac{1}{T_{\mathrm{DCR}}/\mathrm{ns}} \times 10^{6} \tag{3}$$

- [-200, -150] ns relative to main pulse
- $T_{\rm DCR}$ is 50 ns

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Relative PDE

1 total light intensity I_n for nth run, jth splitter ratio α_j , kth PMT PDE η_k

- 2 Expected photon number $p_{njk} = I_n \alpha_j \eta_k$
- 3 Observed trigger rate $R_{njk} = 1 e^{-p_{njk}}$

Oth PMT is the only one reference PMT. $\alpha_j^0 = \frac{\alpha_j}{\alpha_0}$, $\eta_k^0 = \frac{\eta_k}{\eta_0}$, $l_n^0 = l_n \alpha_0 \eta_0$, $i \equiv njk$

$$\log(p_i) = \log(I_0\alpha_0\eta_0) + \log(I_n^0) + \log(\alpha_j^0) + \log(\eta_k^0)$$
(4)

$$R_{i} = 1 - e^{-e^{\log(l_{0}\alpha_{0}\eta_{0}) + \log(l_{n}^{0}) + \log(\alpha_{j}^{0}) + \log(\eta_{k}^{0})}}$$
(5)

The trigger number N_{trig_i} of kth PMT in nth run with jth splitter obey Binomial distribution $B(R_i, N_{t_i})$, in which N_{t_i} is total number of waveforms.

$$\mathcal{L} = \prod_{i} R_{i}^{N_{\mathrm{trig}_{i}}} (1 - R_{i})^{N_{t_{i}} - N_{\mathrm{trig}_{i}}}$$
(6)

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General linear model of Binomial exponential family distribution with Cloglog link function

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• Cathode, focus dynode, MCP

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• CST studio: electric field and trajectory simulation



- The drift times of the electrons at the top of PMT with 0 eV and 3 eV are respectively about 21 ns and 18 ns
- The electrons hitting on the surface of MCP generate the secondary electrons (including a single elastic scattering electron)



- Multiple secondary electrons with different kinetic energy may cause two or more pulses
- Elastic scattering electrons: The sharp difference between them at about 40 ns after the main peak, twice times drift time of electrons from the cathode to the MCP

•
$$B + N_t G(\mu_{\rm TT}, \sigma_{\rm TT}^2) + N_K G(\mu_K, \sigma_K^2) + H(\mu_{\rm TT} + 2\sigma_{\rm TT}) \left(b_S + N_S e^{-\frac{t - (\mu_{\rm TT} + 2\sigma_{\rm TT})}{\tau_S}} \right)$$

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- TTS is defined as FWHM: $2\sqrt{2\ln(2)}\sigma_{TT}$
- The long tail in charge distribution

After pulse categories



- Pre-pulses: photons hitting on the MCP or the first dynode directly rather than the photocathode; 10 ns scale
- After-pulses: the ionization of gaseous impurities between the cathode and first dynode or MCP when photoelectrons go through;100 ns scale
- After-pulses: the relation between time and ions ($_{\rm M}^{\rm Z}{\rm X}$) is $\sqrt{\frac{M}{Z}}$
- Search window: <-10 ns; >200 ns
- The peak position t_p and equivalent charge $C_{\rm equ}$ of the after-pulse and pre-pulse are calculated

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Parameterization



- The relative t: the difference between t_p of pre/after-pulse and t_r^{10} of main pulse
- around 300 ns, 550 ns, 1200 ns and 1700 ns, $1:\sqrt{3}:\sqrt{16}:\sqrt{32}$
- Assumption: H^+ , He^+ or other unknown ions, O^+ or CH_4^+ , and O_2^+ or other unknown ions
- Substracting dark noise rate $N_{\rm DCR} = N_{\rm trig} \cdot {\rm DCR} \cdot T_{\rm bin}$, in which $N_{\rm trig}$ is the number of triggered waveforms

•
$$\sum_{i=1}^{4} A_i G(t_i, \sigma_i^2)$$

• $[-150, -10] \operatorname{ns} [200, 9800] \operatorname{ns}$

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