

The QCD phase transition intrigued by turbulence

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Outline

1. Introduction

2. Framework

3. Algorithm

4. Results

5. Conclusion and Future



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$$\begin{split} \delta N &= N - \langle N \rangle \\ C_1 &= \langle N \rangle, \\ C_2 &= \langle (\delta N)^2 \rangle = \mu_2, \\ C_3 &= \langle (\delta N)^3 \rangle = \mu_3, \\ C_4 &= \langle (\delta N)^4 \rangle - 3 \langle (\delta N)^2 \rangle^2 \\ &= \mu_4 - 3\mu_2^2, \end{split}$$



$${\cal L}=ar\psi(i\hbar\gamma^\mu\partial_\mu-m_0)\psi+G[(ar\psi\psi)^2+(ar\psi i\gamma^5\psi)^2]$$



Our motivation is to study the evolution of the QCD system when there is a fluctuation around the phase transition line. For example, this fluctuation is caused by turbulence.

The isotropic turbulence can be given by the following

$$u^i = \sum_n a(n_1, n_2, n_3) \sin 2\pi (n_1 x + n_2 y + n_3 z + \epsilon),$$

Where $u^{\mu} = (u_0, u_x, u_y, u_z)$, $u^i = (u_x, u_y, u_z)$, $u^0 = \sqrt{1 + (u^i)^2}$, and $a = \sqrt{C^2 k^2 e^{-2(k/k_0)^2}}$, $k = 2\pi n$ is wave vector and $k_0 = 10$, $k = |\mathbf{k}|$, ϵ is the random number in [0, 1] for each combination of $n = n_1, n_2, n_3$, $n_i = 0, \pm 1, \pm 2, \cdots$, i = 1, 2, 3.

1. A. Sakurai and F. Takayama. Molecular kinetic approach to the problem of compressible turbulence. 2. V.V. Aristov and O.I. Rovenskaya. Application of the Boltzmann kinetic equation to the eddy problems.

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$$\frac{\partial f}{\partial t} + \xi \frac{\partial f}{\partial \mathbf{x}} = J(f, f) = \int_{0}^{2\pi} \int_{0}^{b_m} (f' f'_* - f f_*) gb db d\varepsilon d\xi_* = -v(f)f + N(f, f),$$

The initial condition with isotropic turbulence

$$f_M(t, \mathbf{x}, \xi) = \frac{\rho_0}{(2\pi T_0)^{3/2}} \exp\left[-\frac{\mathbf{c}^2}{2T_0}\right], \quad \mathbf{c} = \xi - \mathbf{V}_0,$$
$$\mathbf{V}_0 = (u_0, v_0, w_0),$$

$$u^i = \sum_n a(n_1, n_2, n_3) \sin 2\pi (n_1 x + n_2 y + n_3 z + \epsilon),$$

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The evolution of classical system intrigued by turbulence



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The evolution of classical system intrigued by turbulence

 $\rho(x, y, z) \qquad \rho(x, y, z = 0.5)$



t

We recalculated it to a long time.

$\rho(x,y,z) \qquad \rho(x,y,z=0.5)$





There are number of points lay in the lane with inclination "-5/3", as expected for the inertial interval of 3D isotropic turbulence.

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$$u^i = \sum_n a(n_1, n_2, n_3) \sin 2\pi (n_1 x + n_2 y + n_3 z + \epsilon),$$

Let's explore how the QCD system evolves in the presence of fluctuations around the phase transition line, which are caused by turbulence.

$$\mathcal{L} = ar{\psi}(i\hbar\gamma^\mu\partial_\mu - m_0)\psi + G[(ar{\psi}\psi)^2 + (ar{\psi}i\gamma^5\psi)^2]$$

$$\mathbf{v}_q = \mathbf{p}/E_p$$
, and $E_p = \sqrt{\mathbf{p}^2 + m(x)^2}$

transport equation: $\partial_t f_q + \mathbf{v}_q \cdot \nabla_{\mathbf{x}} f_q - \nabla_{\mathbf{x}} E_p \cdot \nabla_{\mathbf{p}} f_q = C[f_q].$

Gap equation:
$$m(x) = m_0 + 4GN_cN_fm(x)\int_0^\Lambda \frac{d^3p}{(2\pi)^3} \frac{1}{E_p}(1 - f_q(x,p) - f_{\bar{q}}(x,p))$$

- 1. Che Ming Ko, et al, Nucl. Phys. A 928 (2014) 234-246.
- 2. P. Rehberg, et al, Nucl. Phys. A 653 (1999) 415-435.
- 3. WeiMin Zhang, et al, Phys.Rev.C 45 (1992) 1900-1917.
- 4. Pengfei Zhuang, et al, Quantum transport theory and nonequilibrium chiral condensates.
- 5. Abdellatif Abada, Chiral Phase Transition in an Expanding Quark-Antiquark Plasma.

For simplicity, we chose the collision term with the relaxation time model.

$$C[f_q] = -\frac{\mathbf{p} \cdot u}{E_p} \frac{f_q - f_q^{eq}}{\tau}$$

 τ is the relaxation time, which is chosen as $\tau = 0.1$ fm or use the relaxation time formula.

$$f_q^{eq} = \frac{1}{e^{(p \cdot u - \mu(x))/T(x)} + 1}$$

herein, the temperature and chemical potential are fixed by the matching conditions for number density and energy at every space-time point, i.e

$$n_{eq}(x) = n(x)$$

$$e_{eq}(x) = e(x)$$

$$\mu(x), T(x)$$

The number density $n_{eq}(x)$ and energy density $n_{eq}(x)$ are all defined at the local rest frame of the fluid.

$$egin{aligned} n_{eq}(x) &= \int rac{d^3p}{(2\pi)^3} (f_q^{eq} - f_{ar q}^{eq}) = \int rac{d^3p}{(2\pi)^3} (rac{1}{e^{(p\cdot u-\mu)/T}+1} - rac{1}{e^{(p\cdot u+\mu)/T}+1}), \ e_{eq}(x) &= \int rac{d^3p}{(2\pi)^3} E_p(f_q^{eq} + f_{ar q}^{eq}) = \int rac{d^3p}{(2\pi)^3} E_p(rac{1}{e^{(p\cdot u-\mu)/T}+1} + rac{1}{e^{(p\cdot u+\mu)/T}+1}). \end{aligned}$$

Where
$$p \cdot u = p_0 = E_p$$
, $u^\mu = (1, \mathbf{0})$

While the number density n(x) and energy density e(x) are defined by the following

$$n(x)=J^{\mu}u_{\mu}, \quad e(x)=T^{\mu
u}u_{\mu}u_{
u}.$$

The current density $J^{\mu}(x)$ and stress tensor density $T^{\mu\nu}(x)$ are defined by the following,

$$J^{\mu} = \int rac{d^3 p}{(2\pi)^3} rac{p^{\mu}}{E_p} (f_q - f_{ar q}), \quad T^{\mu
u} = \int rac{d^3 p}{(2\pi)^3} rac{p^{\mu} p^{
u}}{E_p} (f_q + f_{ar q})$$

$$n(x)=J^{\mu}u_{\mu}, \quad e(x)=T^{\mu
u}u_{\mu}u_{
u}.$$

The fluid velocity $u^\mu=\gamma(1,v_x,v_y,v_z)$ can be obtained by the energy flow, $T^{\mu
u}u_
u=eu^\mu$, then we can get

$$egin{aligned} &rac{T^{00}-T^{0i}v^i}{T^{0i}-T^{ij}v^j} = 1/v^i, \ & ext{Where } v^i = (v^x,v^y,v^z) ext{ for i=1,2,3.} \ & ext{T}^{0i}-T^{ij}v^j = T^{00}v^i - T^{0j}v^iv^j. \end{aligned}$$

Let use the perturbation method to solve velocity, first let the velocity expanded order by order like as

$$egin{aligned} v^i &= v^i_1 + v^i_2 + v^i_3 + \cdots \ v^i_1 &\sim \mathbf{O}(1), v^i_2 &\sim \mathbf{O}(2), v^i_3 &\sim \mathbf{O}(3) \end{aligned}$$

Let assume

$$egin{aligned} v_1^i &\sim \mathbf{O}(1), v_2^i &\sim \mathbf{O}(2), v_3^i &\sim \mathbf{O}(3) \ \hat{T}^{ij} &\sim T^{0i} &\sim T^{i0} &\sim \mathbf{O}(1) \end{aligned}$$

Where $\widehat{T}^{ij} = T^{ij} - T^{ii}$, Then we can now solve the velocity order by order.

$$T^{0i} - T^{ij}v^j = T^{00}v^i - T^{0j}v^iv^j$$

the first order

the second order

the third order

$$v_1^i = rac{T^{i0}}{T^{00} + T^{ii}}$$

$$v_2^i = -rac{\hat{T}^{ij}v_0^j}{T^{00}+T^{ii}}$$

$$v_3^i = -rac{\hat{T}^{ij}v_2^j}{T^{00}+T^{ii}} + rac{T^{0j}v_1^iv_1^j}{T^{00}+T^{ii}}$$

Finally, we will sum all of these above results as the velocity, $v^i = v_1^i + v_2^i + v_3^i$. The corresponding fluid velocity $u^\mu = (u_0, u^i)$ can be constructed by $u_0 = \gamma = \frac{1}{\sqrt{1-v^2}}$, $u^i = u_0 * v^i$, herein, $v^2 = \sum_i (v^i)^2$.

There is another alternative method to obtain the fluid velocity named the Eckart frame. In Eckart's frame, the fluid velocity can be more easily get by the following relations,

$$v^i = rac{J^i}{J^0}$$

$$\mathbf{C}[\mathbf{f}_q] = -\frac{\mathbf{p} \cdot u}{E_p} \frac{f_q - f_q^{eq}}{\tau}$$

Algorithm

$$\partial_t f_q + \mathbf{v}_q \cdot
abla_\mathbf{x} f_q -
abla_\mathbf{x} E_p \cdot
abla_\mathbf{p} f_q = C[f_q].$$
 $m(x) = m_0 + 4GN_cN_fm(x)\int_0^\Lambda rac{d^3p}{(2\pi)^3}rac{1}{E_p}(1 - f_q(x, p) - f_{ar q}(x, p))$

The transport equation will be split into two steps.

1. The first step is to describe the free-streaming of particles.

$$\partial_t f_q + \mathbf{v}_q \cdot \nabla_{\mathbf{x}} f_q - \nabla_{\mathbf{x}} E_p \cdot \nabla_{\mathbf{p}} f_q = 0.$$

To solve this part, we can use the QUICK algorithm or WENO method. 2. The second step is collision term

$$\partial_t f_q = C[f_{eq}].$$

Algorithm

We use the boundary conditions as follows

$$f(x + 1, y + 1, z + 1, \vec{p}) = f(x, y, z, \vec{p})$$

The grid of phase space

x, y, z are interval on [-1,1]fmpx, py, pz is interval on $[-1.5\Lambda, 1.5\Lambda]$

Np*x*, *Npy*, *Npz*=26 Nx,Ny,Nz=21 or 31

Algorithm



$$u^i = \sum_n a(n_1, n_2, n_3) \sin 2\pi (n_1 x + n_2 y + n_3 z + \epsilon),$$



1. The evolution of number density and vorticity

n(x, y, z)

t







3.0 2.5 2.0 1.5 1.0

3.5

3.0

2.5

2.0

1.5

1.0

1. The evolution of number density and vorticity

t

 $n(x, y, z) \qquad n(x, y, z = 0.0) \qquad \omega(x, y, z)$



The evolution of the system tends to a stable state.

1. The animation of number density and vorticity

n(x, y, z)





2. The evolution of mass at different points



The evolution of the system tends to a stable state.

3. The fluctuation of the number density



This means that there is a phase transition under the evolution of the system.

4. The turbulence properties



$$egin{aligned} u(t,\mathbf{x}_{ijk}) &= rac{1}{V} \sum_{n,m,l=0}^{N_p-1} u(t,\mathbf{n}) \, \cos(rac{2\pi}{L_x}nx_i) \cos(rac{2\pi}{L_y}my_j) \cos(rac{2\pi}{L_z}lz_k), \ u(t,\mathbf{n}) &= dxdydz \, \sum_{i,j,k=0}^{N-1} u(t,\mathbf{x}_{ijk}) \, \cos(rac{2\pi}{L_x}nx_i) \cos(rac{2\pi}{L_y}my_j) \cos(rac{2\pi}{L_z}lz_k). \end{aligned}$$

$$E_{nml} = (u_{x,nml}^2 + u_{y,nml}^2 + u_{z,nml}^2)/2.$$
 $q = \sqrt{n^2 + m^2 + l^2},$

5.The plot of $(T-\mu)$

near the CEP at right-hand side



far from the CEP at right-hand side



5.The plot of $(T-\mu)$

near the cross-over line at righthand side

near the cross-over line at left-hand side





5.The plot of $(T-\mu)$

near the CEP at left-hand side



far from the CEP at left-hand side





6. (e-3P)/T^4

near the CEP at right-hand side

far from the CEP at right-hand side





6. (e-3P)/T^4

near the CEP at left-hand side



far from the CEP at left-hand side



Conclusion and Future

- 1. The phase transition is intrigued by turbulence.
- 2. The evolution of system is tending to a stable state.

Thank You!