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Radiative Capture $d(\alpha, \gamma)^6Li$ Reaction in Cluster Effective Field Theory

In the standard Big Bang Nucleosynthesis (BBN) framework, the primitive 6Li abundance is mainly determined by two nuclear reactions: the $d(\alpha, \gamma){}^6Li$ reaction, where deuteron reacts with an alpha particle to produce 6Li . This reaction leads to the formation of 6Li during the primordial nucleosynthesis process. Conversely, ${}^6Li(p, \alpha){}^3He$, can destroy 6Li and reduces the abundance of 6Li in the early universe. In this study, we focused on the radiative capture process of the deuteron on alpha leading to the formation of 6Li through the cluster Effective Field Theory. The contribution of the Coulomb interaction between particles was taken into account nonperturbatively. In the first step, the asymptotic normalization coefficients (ANCs) of the bound state of 6Li was calculated based on the elastic scattering process of the alpha-deuteron system. In the following, the calculation of the transition amplitude and astrophysical S-factor were outlined. To evaluate the accuracy of the approach, we compare the astrophysical S-factor of the dominant electromagnetic transitions (E1 and E2) up to next-to-leading order with the experimental data.

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