

Locating a Super PeVatron at Cygnus Region

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Outline

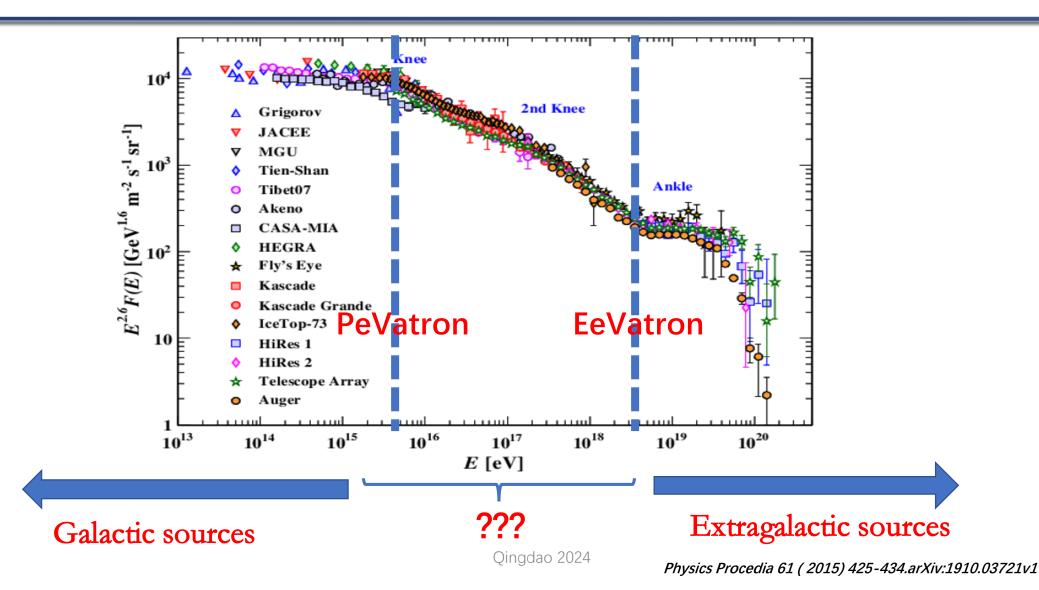


- Origin of cosmic rays
- LHAASO detector

- Results from Cygnus region
- Prospects

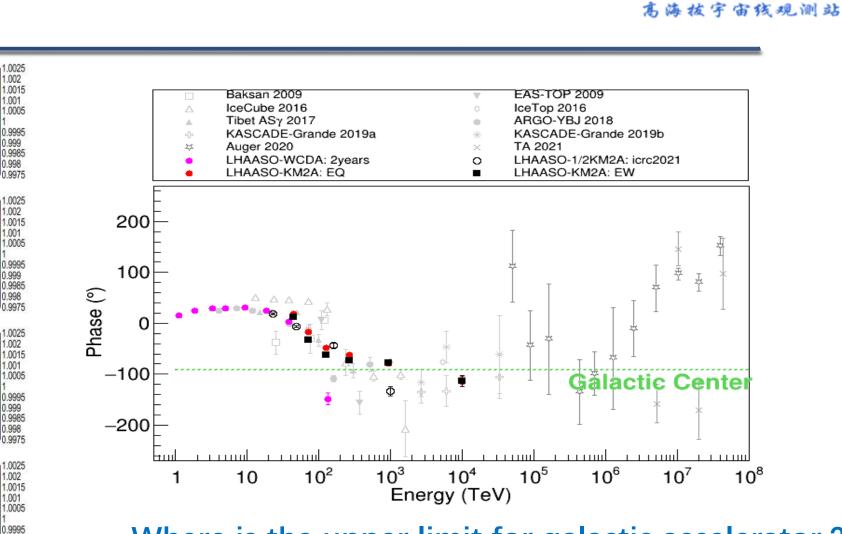
Where dose the highest energy particles come from ?





3

Where dose the highest energy particles come from ?



Where is the upper limit for galactic accelerator?

Qingdao 2024

ICRC2023 WeiGao

https://iopscience.iop.org/article/10.3847/1538-4357/aac6cc

150 200 Right Ascension (deg)

150 200 Right Ascension (deg)

150 200 Right Ascension (deg)

150 200 Right Ascension (deg) 250

250

250

250

300

350

0.999 0.9985 0.998 0.998

350

39 TeV

50

50

50

50

520 Te\

160 TeV

71 TeV

100

100

100

Declination (deg)

clination (deg)

Ő

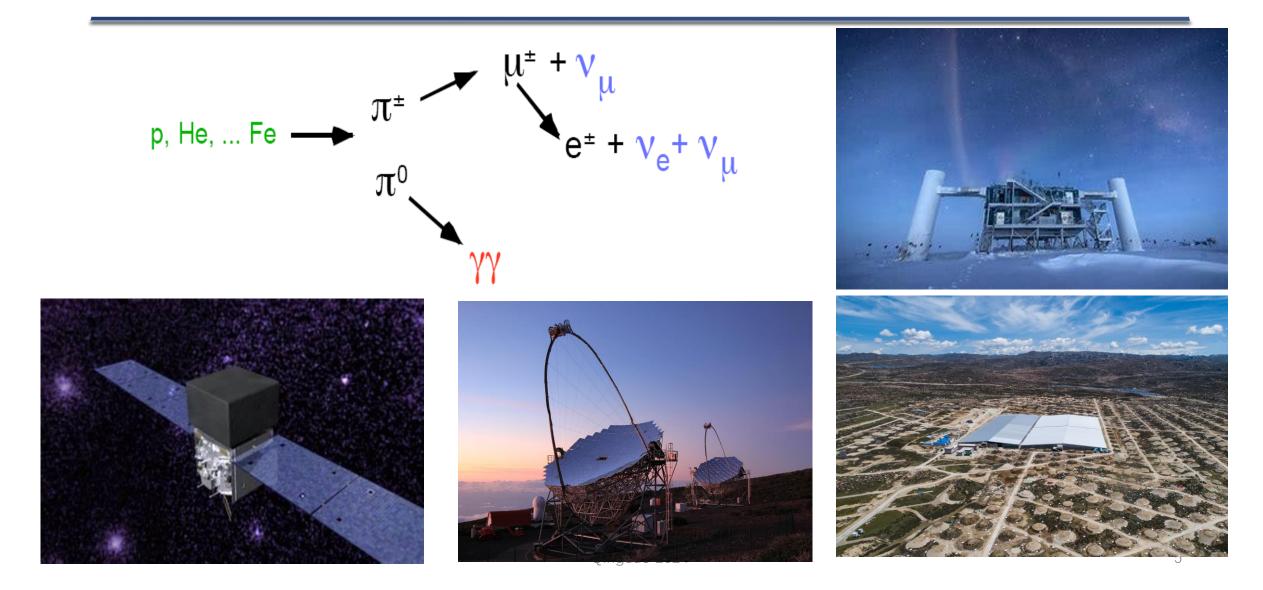
ination (deg)

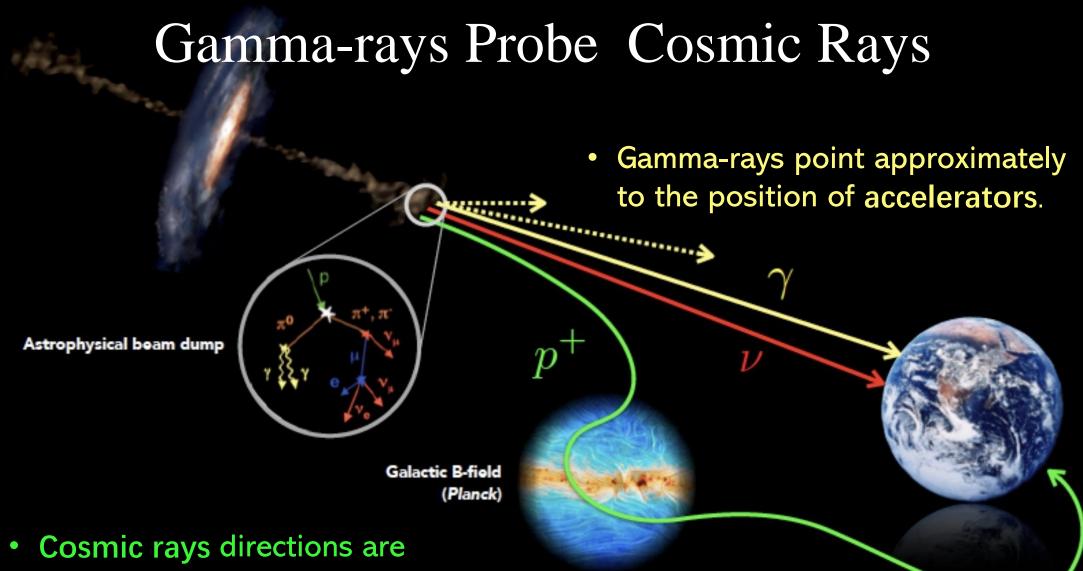
Dec

Declination (deg)

How to trace the cosmic ray accelerator?

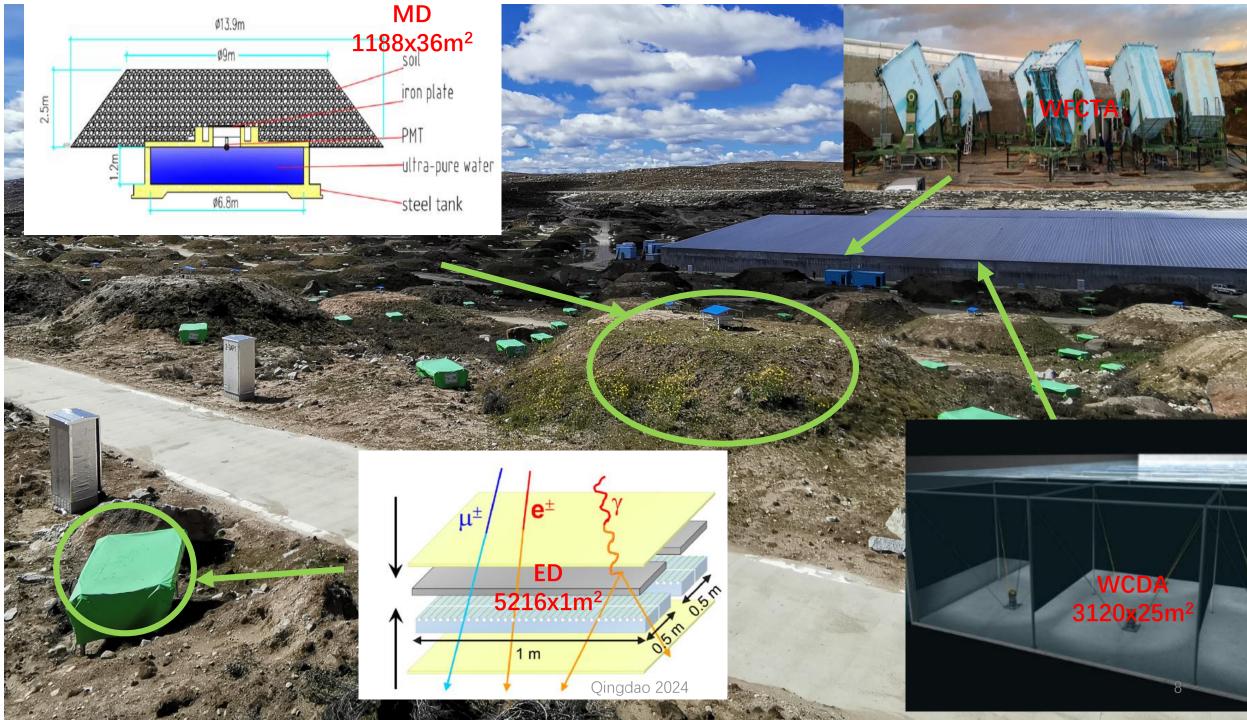






randomized by Magnetic Fields in the Universe.

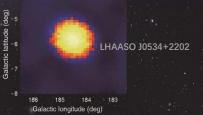
Location: 29°21′ 27.6″*N* 100°08′ 19.6″*E* Altitude: 4410*m a.s.l*



The 1st CR-Source Candidate by LHAASO



Cygnus Bubble, Science Bulletin, arXiv:2310.10100 LHAASO J2018+5157 LHAASO J2018+3651 LHAASO J2018+3651

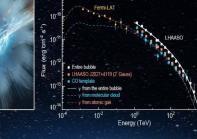


PeVatrons, Nature 594:33-36 (2021)

LHAASO J1849-0003 LHAASO J1908+0621 LHAASO J1849-0003 LHAASO J1843-0338

EHAASO J1825-1326

LHAASO discovers giant ultra-high-energy gamma-ray bubble, identifying the first Super PeVatron

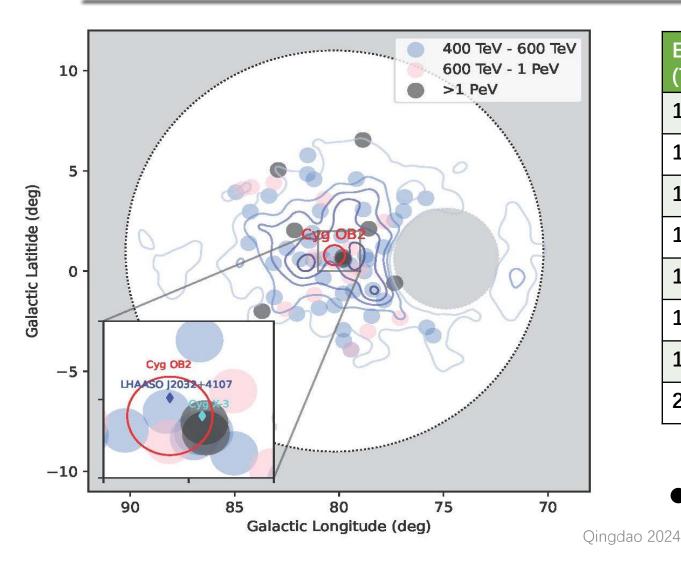


- Discovering a giant ultra-high-energy gamma-ray bubble in the star-forming region of Cygnus, which extends to more than one thousand light-years
- Positioning the source of cosmic ray with energy above 10 PeV for the first time





A Bubble of UHE γ 's centered at a complex core



Energy (TeV)	Ne	Nu	Theta (deg)	Dr (m)
1087	5904	13	19.4	143
1188	5480	14	34.4	73
1208	6939	13	14.2	131
1350	6938	8	27.1	43
1379	6469	9	17.4	52
1421	6258	7	12.7	57
1784	6665	13	18.0	41
2481	13815	29	33.0	99

• The UHE photos are dispersed distributed, and are not correlated with any small scale sources. 10



Clear correlation with gas distribution indicating a

The signal is elongated along the disk and extends to

hadronic origin of photons in the Bubble

Association with HI gas distribution over $\sim 200 \text{ pc}$

- The significance map is smoothed with a Gaussian kernel=1.0°
- The contour is from HI4PI 21-cm line survey
- Galactic Latitude [degree] 120Galactic Latitude [degree] Significance [a 10 Galactic Longitude [degree] Galactic Longitude [degree] 25-100TeV 1-20TeV 85 80 90 75 70 Qingdao 2024 11 Galactic Longitude [degree]

10°

Association with molecular cloud

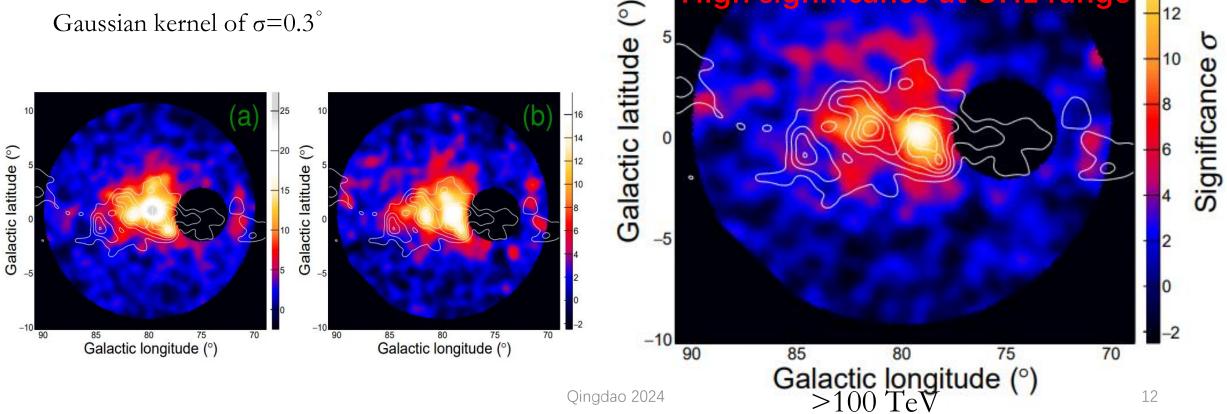


16

14

12

- The contour is from CfA galactic CO survey
- The significance map is smoothed with a Gaussian kernel of $\sigma = 0.3^{\circ}$



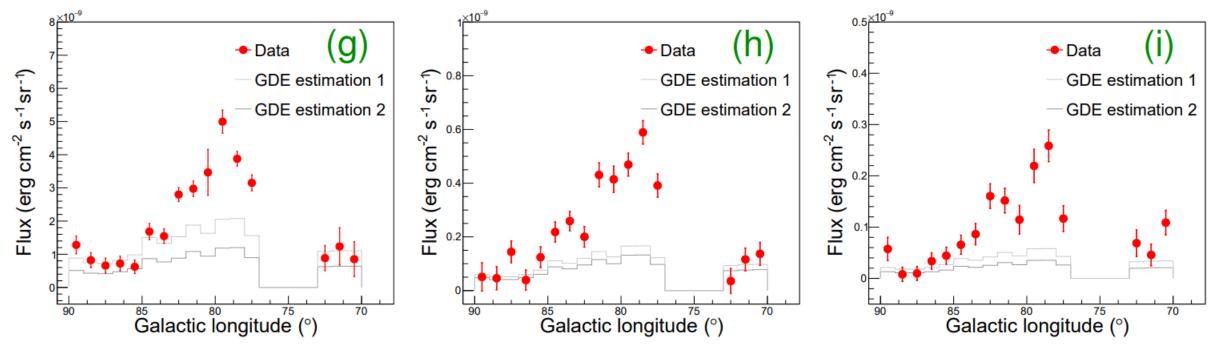
10

High significance at UHE range

Radial profile of the extended emission



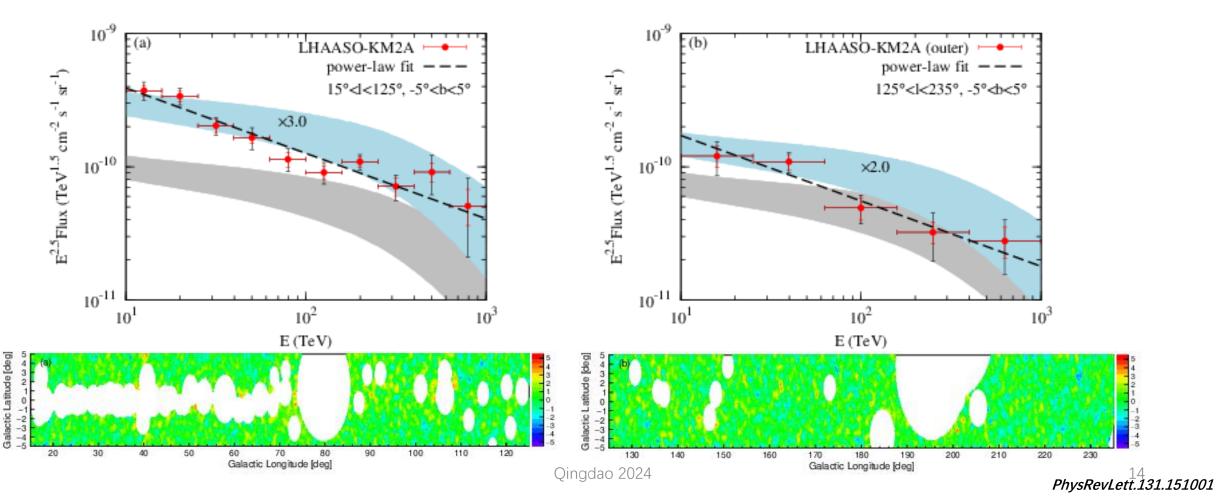
- $F_{\gamma} = w_{CR} \times N_{HI+H2}$
- A very sharp distribution of gamma ray emission towards the center agree with CR propagation scenario and rule out a significant contribution from GDE.



Galactic Diffuse Emission



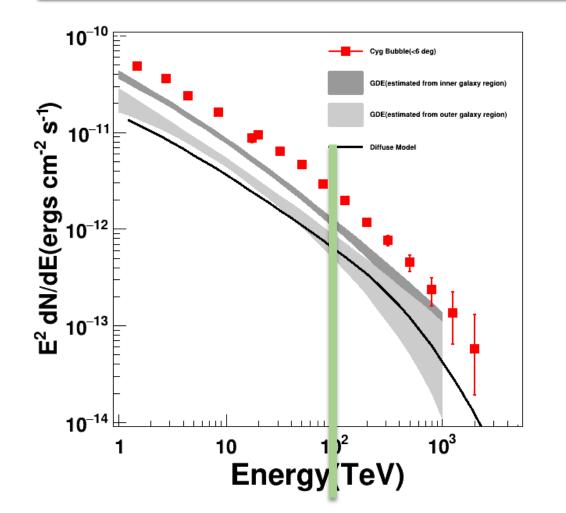
Inner Galactic Region



Outer Galactic Region

Energy spectrum





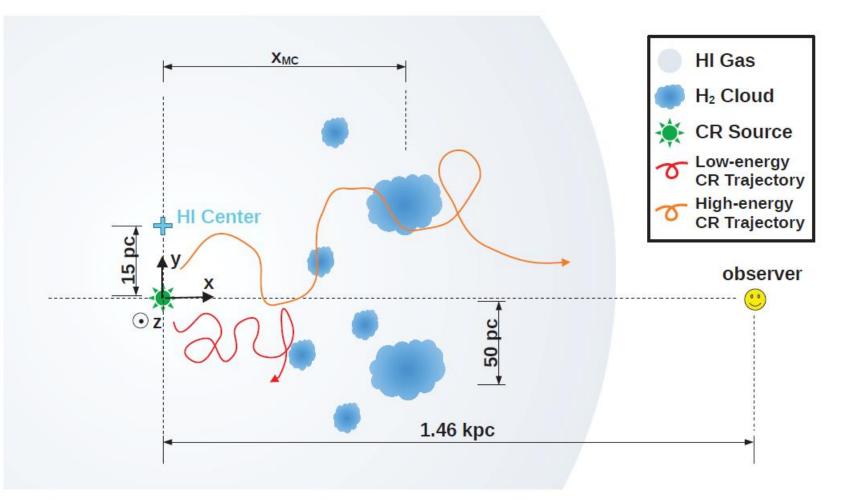
Energy Bin	Non	Nb			
400TeV-630TeV	42	6.8			
630TeV-1PeV	14	1.9			
1PeV-1.6PeV	6	0.6			
1.6PeV-2.5PeV	2	0.2			
Almost background free					

The spectrum can extend beyond PeV without sharp cut-off, which shows a slightly softening feature.

It is definitely a PeVatron, or even a Super-PeVatron

Interpretation

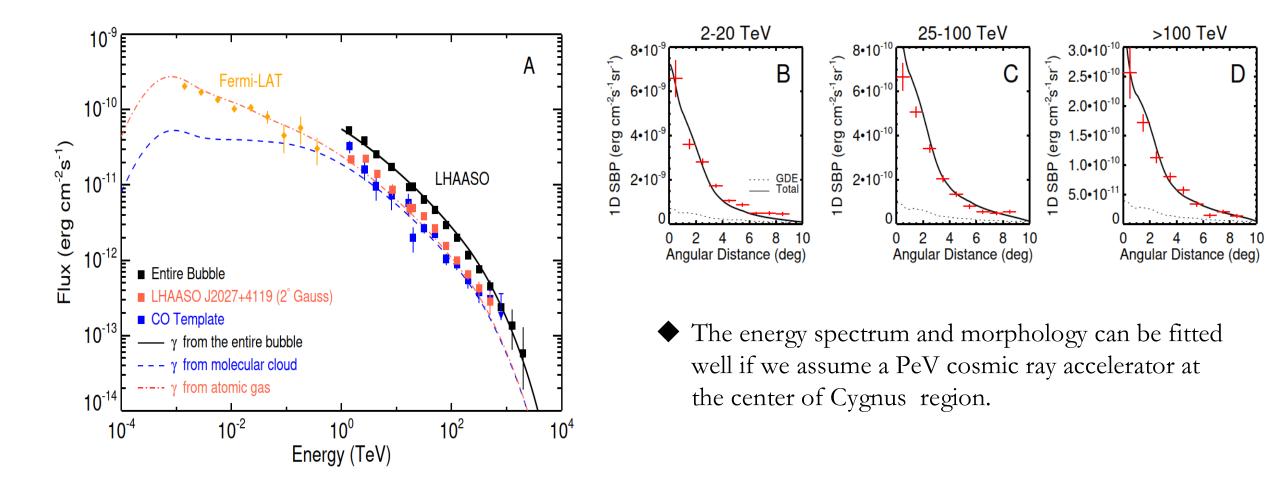
- Have the highest energy cosmic rays have escaped from the accelerator?
- What we observed is the projection of gamma rays produced is a 3D space.





Fitting results









- A giant ultra-high energy gamma-ray bubble is discovered at Cygnus Region.
- The morphology and spectrum indicates a cosmic ray accelerator located at the center accelerating cosmic rays at least to 10 PeV.
- This is the first time locating a Super PeVatron in galaxy.



Thanks very much !

Any question?

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