Search for WIMP Dark Matter in PandaX-4T Experiment



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Progress of DM Direct Detection



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PandaX Collaboration



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PandaX Detectors

- Increasing the detector sensitive target volume
- Lowering radioactive background

PandaX-I

120kg

2010-2014







PandaX start

2009

🔄 PANDA X

2015-2019

PandaX-II

580kg

PandaX: A Dual-phase Xenon TPC





- **PandaX**: Particle and Astrophysical Xenon Observatory
- Pure xenon target, enhanced DM signals, achievable liquefication temperature, high light & charge yield
- Good ER/NR discrimination by S2/S1 ratio



• 3D reconstruction rejects external background

Outline



- Run1 data taking and running conditions/challenges
- Low-energy calibration, reconstruction, selection & efficiency
- Signal response model
- Background evaluation
- Statistical inference results for WIMP search
- Summary and outlook

After Commissioning

2020/11 _ 2021/04	Commissioning (Run 0) 95 days
2021/07 _ 2021/10	Tritium removal xenon distillation, gas flushing, etc
2021/11 _ 2022/05	Physics run (Run 1) 164 days
2022/09 _ 2023/12	CJPL B2 hall construction xenon recuperation, detector upgrade
Current Status	Resuming physics data-taking



- > WIMP search: Combined blind analysis of Run0 and Run1.
- Total exposure: 1.54 tonne-year

Run1 Data Taking & Challenges

- Gate -6kV, Cathode -16kV (Gate trip once)
- \succ e⁻-lifetime monitoring through α events
 - maximum reaches 1800 us
 - sensitive to operation condition
- Failure of liquid level controlling
 - liquid level sensitive to the circulation flow rate
 - monitoring through the drift time of gate events and single electron gain (SEG)
 - dividing into 6 subsets accordingly
- Additional malfunctioned PMTs (see next)
- Improvements and updates
 - Charge correction
 - Position correction



Additional "Off-PMT" Problem

400











\succ Additional non-functioning PMT channels at top, with 8 channels sharing the same negative HV due to the short of the photocathode.

- Concentrated in the same top area, affecting event reconstruction and selection (TBA, position reconstruction, etc.).
- Solution in Run2: During the detector upgrade, repair the non-functioning PMTs and distribute them as evenly as possible to reduce the coupling of adjacent channels.

Charge Spatial Uniformity Correction





- ^{83m}Kr (41.5 keV): internal conversion e⁻
- ➢ Binned map → Unbinned map
 - Perform a 9th-degree polynomial function fit $\sum_{ijk} c_{ijk} x^i y^j z^k$ where *i*, *j*, *k* = 0, ... 9



S2 also adopts z-dependent charge correction according to its e⁻ lifetime

Charge Temporal Variation Correction

- C:
- Due to the change in liquid level over time, both the light signals and ionization signals fluctuate over time.
- Method: Run-by-run correction utilizing ²²²Rn 5.6MeV α event (mono-energy, existing over time)



Ionization charge before/after correction

Position Correction



Low-energy ER/NR Calibration



- End-of-run low-E ER/NR calibration (ER: ^{220/222}Rn, NR: DD + ²⁴¹AmBe)
- Determine all selection criteria, efficiencies and charge biases (together with waveform simulation)
- > 0.3% ER leak ratio in Rn calibration \rightarrow good ER/NR separation

Quality Cuts & Efficiencies

Quality selection cuts:

- ➢ S1- and S2-related, waveform cleanness
- Relaxed version for "off-PMT" region

Total efficiencies as a function of energy: plateau ~90%

Quality

- > ROI
 - S1^c: 2-135 PE, \geq 2-hit coincidence
 - S2^{raw}: 120-20,000 PE
- Reconstruction
 - Signal classification (tagging)
 - ➤ S1-S2 pairing



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PandaX-4T Signal Response Model

- Signal Response of the PandaX-4T detector (Run0) & Run1): How deposited energy converts to detectable signals
 - Light yield & Charge yield
 - Recombination fluctuation
- Compare w/ nominal NEST
 - the mean recombination fraction $\langle r \rangle$ is adjusted by adding a 3rd-order Legendre polynomial multiplied by an exponential function P
 - the recombination fluctuation Δr is scaled by a factor λ

 $\langle r \rangle(\xi) = \langle r \rangle_0(\xi) + P_3(\xi/\xi_{\text{norm}}; p_0, p_1, p_2, p_3) \cdot e^{-\xi/\xi_{\text{norm}}}$ $\Delta r(\xi) = \Delta r_0(\xi) \cdot \lambda,$

- Run0+Run1 simultaneous fit to all ER/NR calibration data (DD + AmBe + Rn)
- Data-driven determined detector effects, corrections, etc.





P4-NEST vs. Data: Run1 DD



Comparing the observed calibration data with the signal response model, a good agreement has been achieved.



Background Budget





Tritium Level (after unblinding)

- Significant reduction from Run0 to Run1 (~8 times)
- Consistent with S1-only estimation used before unblinding
- Floating in the final PLR fit





²¹⁴Pb Background

- ²¹⁴Pb: daughter nuclei of ²²²Rn
- Select ²²²Rn alpha events as a monitor
- ²²²Rn level varies with running condition

Rn-222 level	[µBq/kg]
Run 0	7.07 ± 0.02(stat.) ± 0.23(sys.)
Run 1	8.67 ± 0.01(stat.) ± 0.27(sys.)

• ²¹⁴Pb is derived from spectrum fitting (0.2-1MeV)





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set2

set3

set4

set5

set6

⁸⁵Kr Background

- Residual impurity in product xenon
- Identified through β - γ coincidence selection



Kr/Xe [ppt]

Surface Background

- ER events, whose S2 suppressed by the TPC surface
- Estimate the radial distribution by ²¹⁰Po alpha events
- Good consistency with data outside blind region

Surface events in fiducial volume				
Run0	0.09 ± 0.06			
Run1	0.17 ± 0.11			



Neutron Background

> Combine 3 approaches to estimate (weighted average):

(n

¹³¹Xe

capture

GEANT4 + signal model



High energy γ / neutron ratio

Multi-scatter / single-scatter ratio



$$N_{
m neutron} = N_{\gamma} * r_{
m n/\gamma}$$

$$N_{\rm neutron} = N_{MS} * r_{SS/MS}$$

Unit: counts	Run0	Run1
Pure Neutron	0.44 ± 0.13	0.81 ± 0.17
Neutron X	0.19 ± 0.06	0.29 ± 0.06
Total Neutron	0.63 ± 0.18	1.10 ± 0.24

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2024/8/15

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Accidental Background

- Accidental coincidence (AC): The pairing of random isolated S1s and S2s that can mimic real physical single-scatters
- Use random pair MC and "off-window" events to determine unphysical AC background
- The AC background is underestimated (~2 times) due to an inconsistent acceptance cut on the scrambled data (found after unblinding, yet no further selection cuts are added/tuned).

Unit: counts	Run0	Run1
Accidental (updated)	11.3 ± 3.4	12.7 ± 3.8







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SI Upper Limit & Sensitivity Band

arXiv: 2408.00664

Fully blind analysis Run0+Run1:

- Scanning WIMP mass from 5 to 10000 GeV/c²
- \rightarrow No significant excess!
- +1 σ upward fluctuation: < 8GeV/c² Global significance (after LEE correction): $Z^{global} = 1.2$
- State-of-the-art: >100 GeV/c²
- Lowest upper limit: 1.6×10⁻⁴⁷ cm² at 40 GeV/c² after -1σ power-constraint



SD Upper Limit & Sensitivity Band

• Refresh spin-dependent constraints for Xe-based experiment



Summary and Outlook



- A combined blind analysis of PandaX-4T Run0+Run1 for WIMP search comes out
- No significant excess and we present the latest stringent upper limit for DM mass above 100 GeV/c²
- Studies of more physical topics are ongoing, and we keep accumulating more data & Stay tuned!

Thank you for your attention!

Backups

PMT Gain Self-calibration



• Previous

- Approach: Weekly gain calibration with LED light.
- Problem: The response to changes in gain (due to aging, etc.) is not timely enough.
- New:
 - Approach: Self-calibration via the fitted single photon peak from its own run.



PandaX-4T Signal Response Model

- > To study how deposited energy converts to the detectable signals
- Run0+Run1 simultaneous fit to ER/NR data (DD+AmBe+Rn)
- Tune light & charge yields





 $E_{\rm ee} = W_q \left(\frac{Q_{S1}^c}{q_1} + \frac{Q_{S2_{\rm b}}^c}{q_{2_{\rm b}}} \right)$

 $Q_{S2_b}^{c}/\xi$ [PE/keV]

PandaX-4T Signal Response Model

Consider multiple-scatter (MS) process



Best Fit Result (Run0+Run1)



PLR fit: with unbinned likelihood with all signal/background PDFs in (S1, S2_b)

Best fit DM counts:

- 3.7 w/ 5 GeV WIMP
- 0 w/ 40 GeV WIMP



Distribution of DM Candidates

S1

 $\log_{10} (S2_{
m b}/$

1.0

WIMP (1000 GeV)

6 8 10 12

Tritium β^-

Candidate Index

Other ER

¹²⁷Xe L-shell EC

Neutron

¹²⁴Xe LL-shell ECEC

Accidental

 $^{8}B CE \nu NS$

Surface

14 16 18 20 22

Candidate Index

 10^{2}







> 24 (12+12) below NR median events

> Uniformly distributed in the FV.

S1 [PE]

101

 10^{1}

S1 [PE]

NR median

ER 5%-95% quantiles

NR 99.5% acceptance

Total: 2490

2.5

 $\log_{10}\left(S2_{
m b}/S1
ight)$

1.0



 some upward fluctuation for DM mass < 8GeV, and some downward fluctuation for high mass DM





- Standard WIMP searches
 - Spin-independent (SI)
 - Spin-dependent (SD)
- Luminance of DM
- Several novel approaches
 - Lower threshold (S2-only)
 - Migdal effect
 - $-\chi$ -v conversion
 - Boosted mechanism



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- First experimental constraints on DM charge radius:
 4 orders of magnitude than the neutrino
- Up to 3 10 times improvement for other electromagnetic properties



X. Ning et al. **Nature** 618 (2023)

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DEAP-3600

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10

D. Huang et al. PRL 131, 191002 (2023)

10

 10^{2}

 $m_{\gamma} [MeV/c^2]$

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 10^{-41}

0.2 0.3

Dark-matter mass [GeV/c²]

0.1

 10^{3}

– Spin-independent (SI)

- Spin-dependent (SD)
- Luminance of DM
- Several novel approaches

Standard WIMP searches

- Lower threshold (S2-only)
- Migdal effect
- $-\chi$ -v conversion
- Boosted mechanism
 - Overcome detection threshold



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X. Ning et al. PRL 131, 041001 (2023)

arXiv: 2403.08361